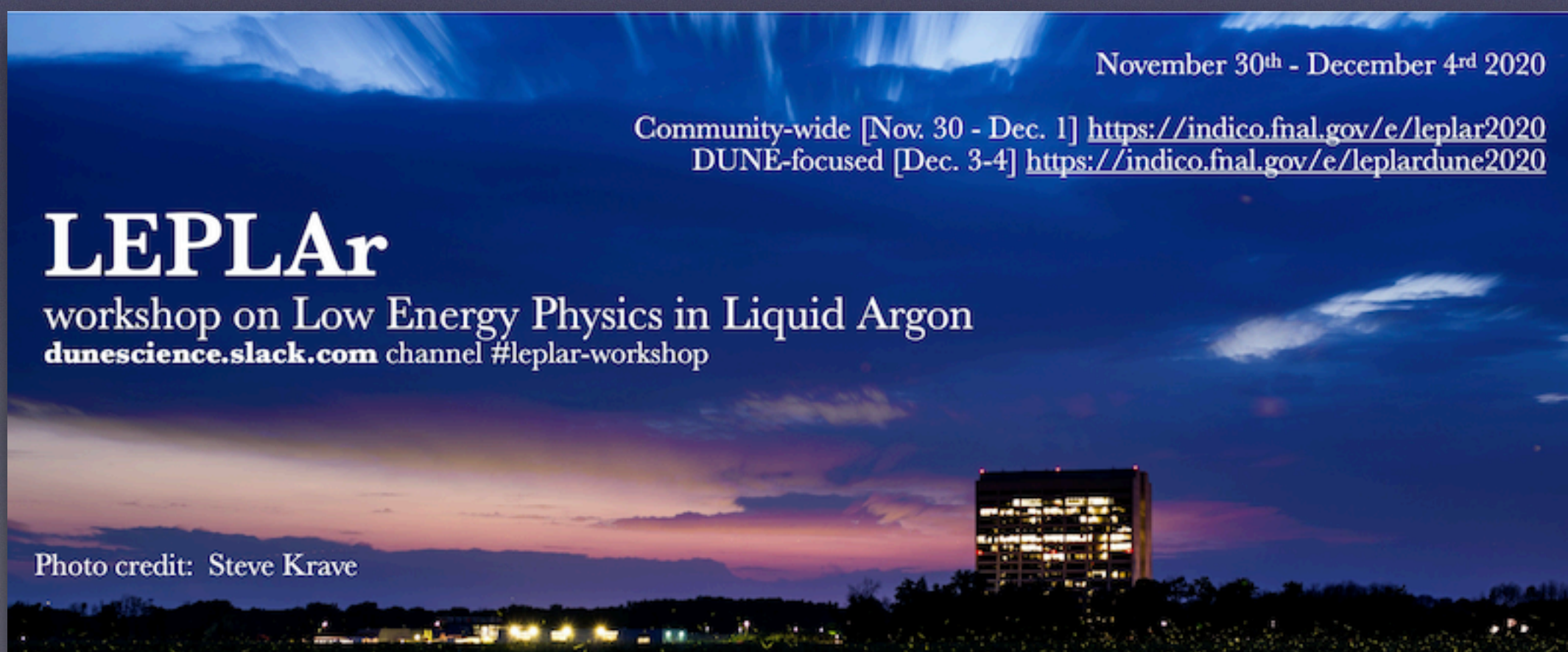


Enlarging the DUNE Physics Scope with

An Enhanced Photon Detection System in the VD layout (proposed for future DUNE FD Modules)

A banner for the LEPLAr workshop. The background is a photograph of a building at night with lights on, set against a twilight sky. The text is overlaid on the top left of the image.

November 30th - December 4rd 2020

Community-wide [Nov. 30 - Dec. 1] <https://indico.fnal.gov/e/leplar2020>
DUNE-focused [Dec. 3-4] <https://indico.fnal.gov/e/leplardune2020>

LEPLAr
workshop on Low Energy Physics in Liquid Argon
dunescience.slack.com channel #leplar-workshop

Photo credit: Steve Krave

Flavio Cavanna - FERMILAB

LArTPC technology approaches the limits of its *full reconstruction capability** when event energy falls below the ~ten(s) MeV range:

- *This is where an important part of rare UG Physics may lie*
- *This is where traditional large Volume Liq. Scintillator UG experiments successfully operate - e.g. Borexino, **or were proposed - e.g. LENA***

While developing the (new) **VD LAr detector concept**
for **DUNE FD(UG) Module ≥ 2** ,

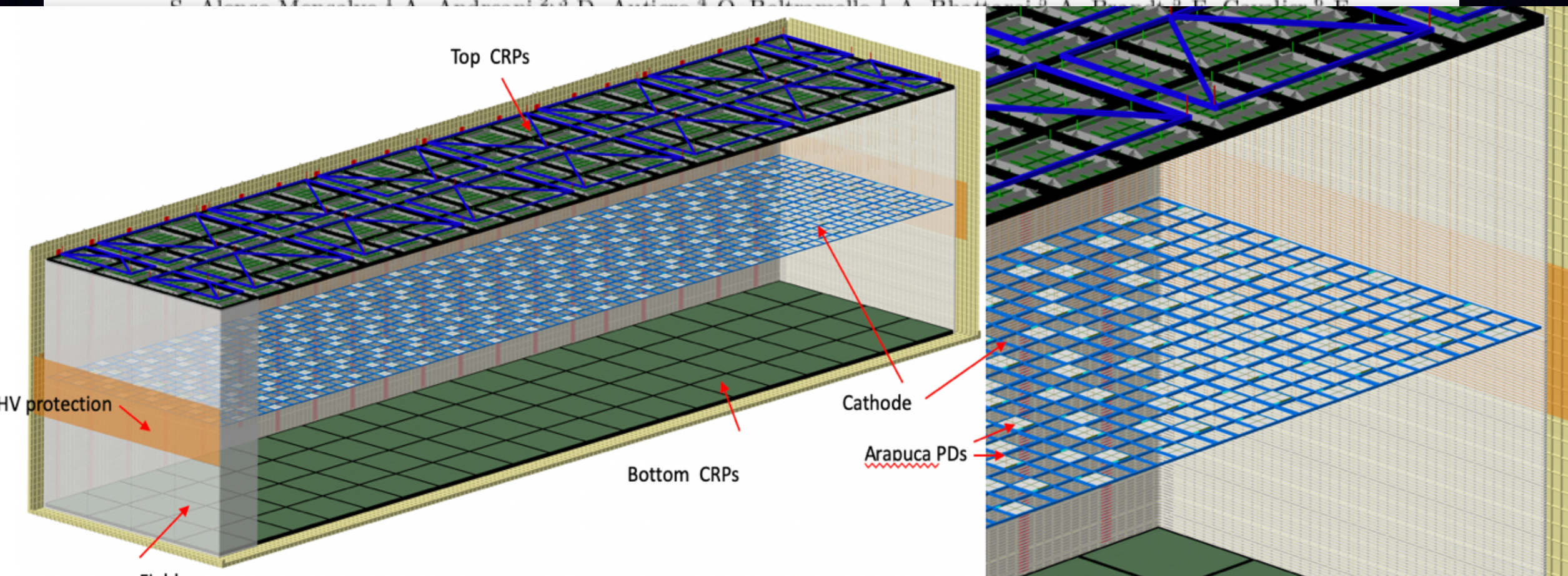
exploiting abundant LAr scintillation light (*complementary to ionization charge*)
appeared as a “natural” way to enhance/extend detection sensitivity for UG
low-energy rare events.

The key point is to extend PD Optical Coverage as close to 4π as possible: to embed it into LArTPC layout is a big technological challenge.

* complete Topology (vtx, energy, position, direction, ptcl. Id.) - En. Threshold in E_{dep} can be much lower, in the ~tens keV, depending on S/N

1
2
3
4
5
6

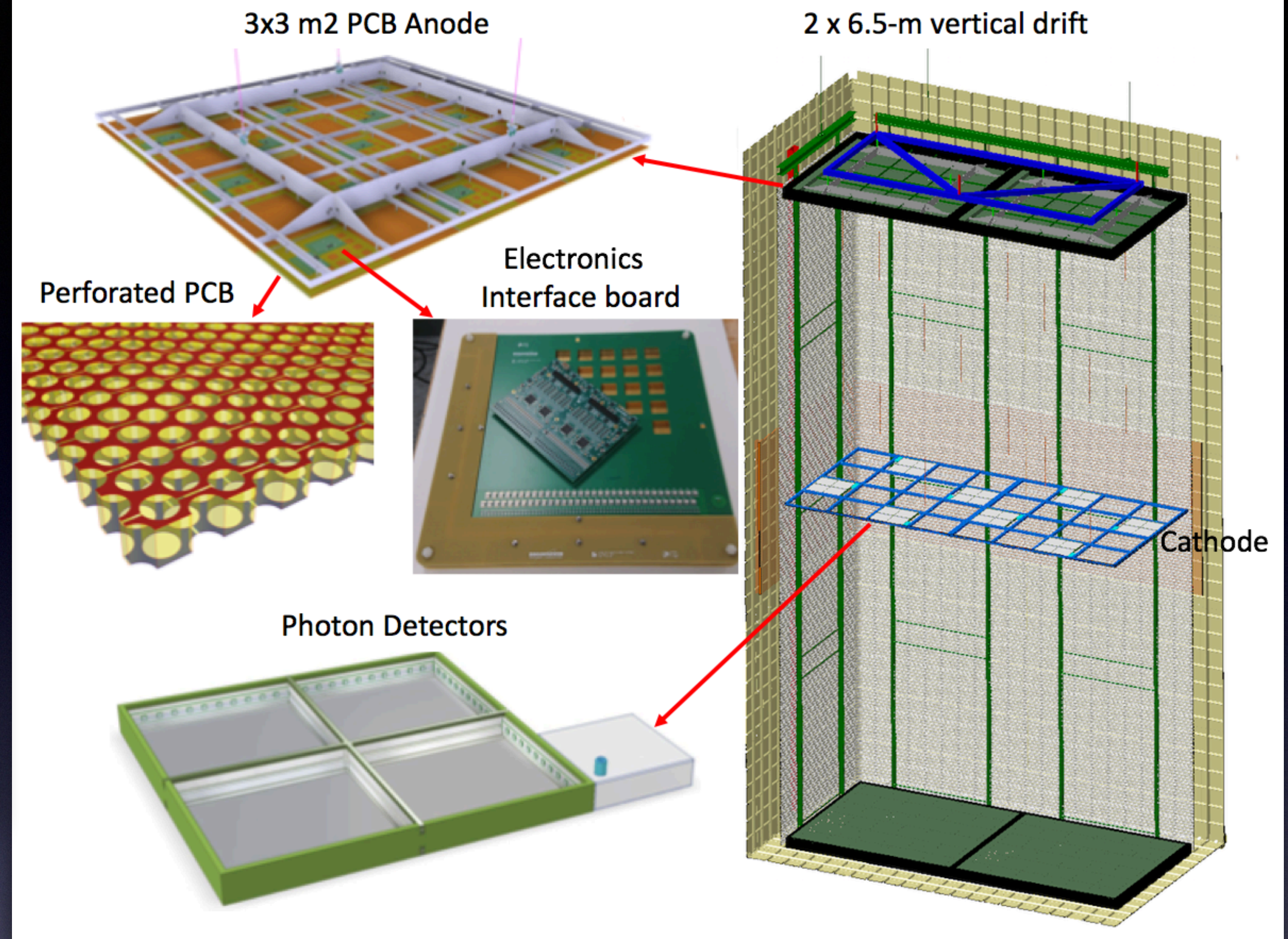
Next-generation LArTPC Detector Technology for the Deep Underground Neutrino Experiment: a Vertical Drift Single-phase Solution with Perforated PCB Anode



28
29
30

This dynamic author list will be updated regularly.

DRAFT – 0



Next-generation LArTPC concept:
a Vertical Drift Solution with PCB based
Charge Readout
complemented by a robust X-ARAPUCA
Photon Detection System

Conceptual design for a $\sim 4\pi$ PD System for the VD LAr Volume

- PDS cannot be located at the Anode Plane (as in the DUNE SP Module#1) !
- If a solution for operating a PD on HV surfaces is found:

PD Active Optical Coverage distributed onto **5 sides of the LAr Volume** (Cathode side and 4 Field Cage sides)
+
PD Passive Optical Coverage (reflector) onto Anode side (laminated on perforated PCB)
+
Xe doping (minimize Rayleigh scatter for light at far distance)

This would allow $\sim 4\pi$ coverage:

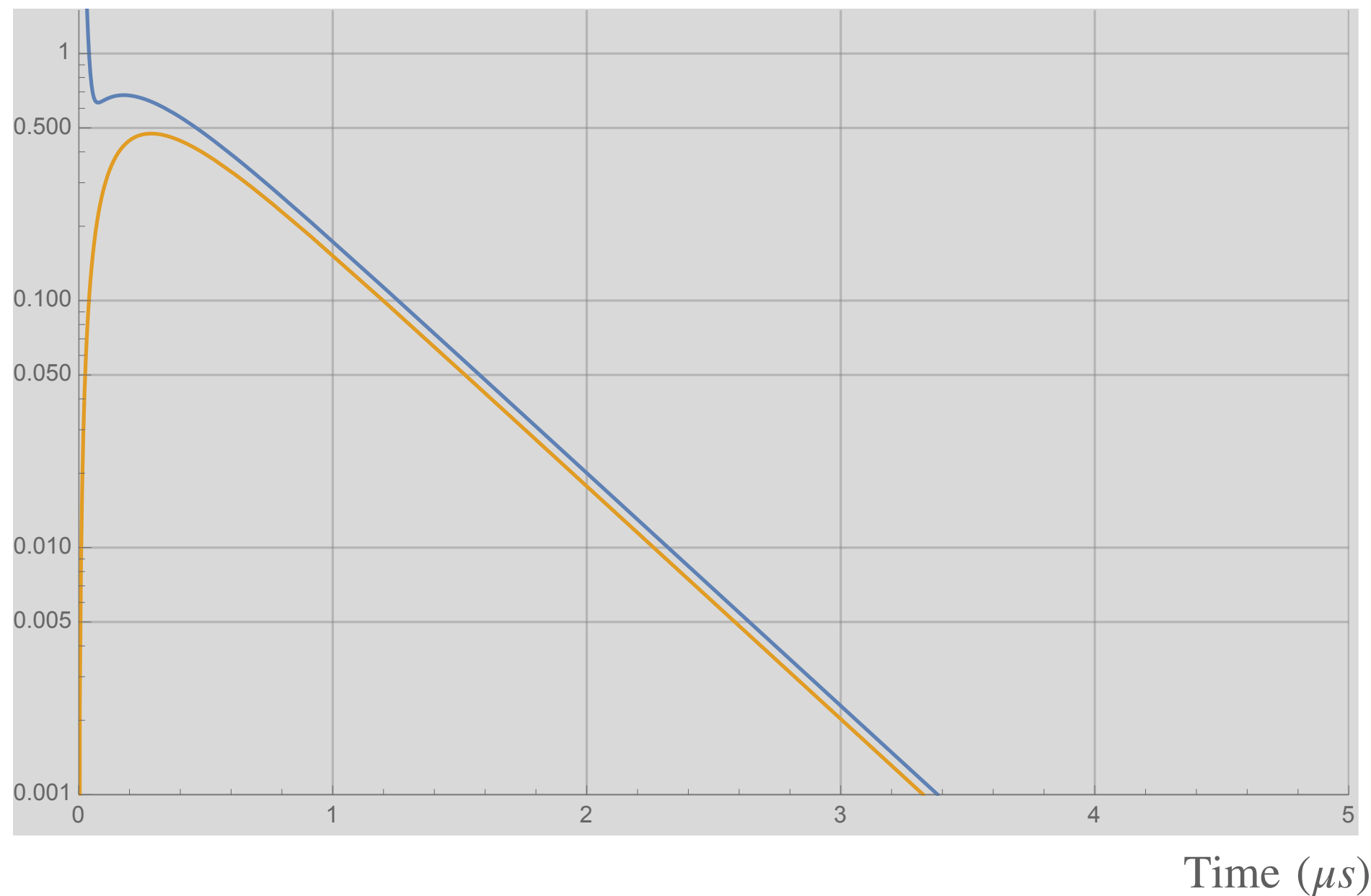
⇒ *~uniformity of response, low detection threshold, energy resolution and position resolution capability*

It would be a second detector - for Ar Light Signals - complementary to LArTPC (directionality):

- *complete exploitation of LAr features (collect all energy deposited)*
- *Guarantee highest Live Time (PD active also when LArTPC may be OFF for purity drop/maintenance, HV issues/ maintenance,..) very relevant for UG Physics*
- *Start data taking (SN observer) six months/one year before LArTPC (while LAr filling)*

Baseline assumptions here are:

- Liquid Scintillator: **Ar + Xe(10 ppm)** [residual impurity: $[N_2] < 1$ ppm, $[O_2]$ negligible $\Rightarrow \lambda_{Abs} \geq 30m$]. Photon Yield (mip, nominal EF): $Y_{ph}(Ar) \simeq 6k$ ph/MeV, $Y_{ph}(Xe) \simeq 19k$ ph/MeV ($\lambda_R(Ar) \simeq 1m$, $\lambda_R(Xe) \simeq 7m$).
- Photo-collector: **X-ARAPUCA technology** (light trapping by dichroic filters and 2-stages WLS), sensitive to both Ar VUV light (128 nm) and Xe UV light (147 and 175 nm), detection efficiency $\epsilon_D \simeq 3\%$.
- Photo-sensor: **SiPM/MPPC** Si avalanche photodiode micro-cell array, $6 \times 6mm^2$ area, single-photon sensitive, $PDE(430\text{ nm}) \simeq 40\%$, $SNR \geq 5$. Cold electronics read-out (active ganging + shape&noise filtering + digital conversion& transmission).



(Ar, Ar+ArXe+Xe) & (Xe)
detectable Light Time Distribution

Where's the challenge ??

Operating PD on HV surface

- requirements, base solutions, alternatives
 - * PD based on SiPM (low Bias V, minimal occupancy)
 - * **SiPM Bias Voltage Supply (IN), SiPM transmit Signal (OUT)**
 - * **PoF (Bias V) Receiver & PoF (Signal) Transmitter**
 - * **PoF Receiver (Bias V) & WiFi/RF (Signal) Transmitter**
 - * **SiPM Cold Electronics** (if used, it also requires Power => more from PoF receiver)

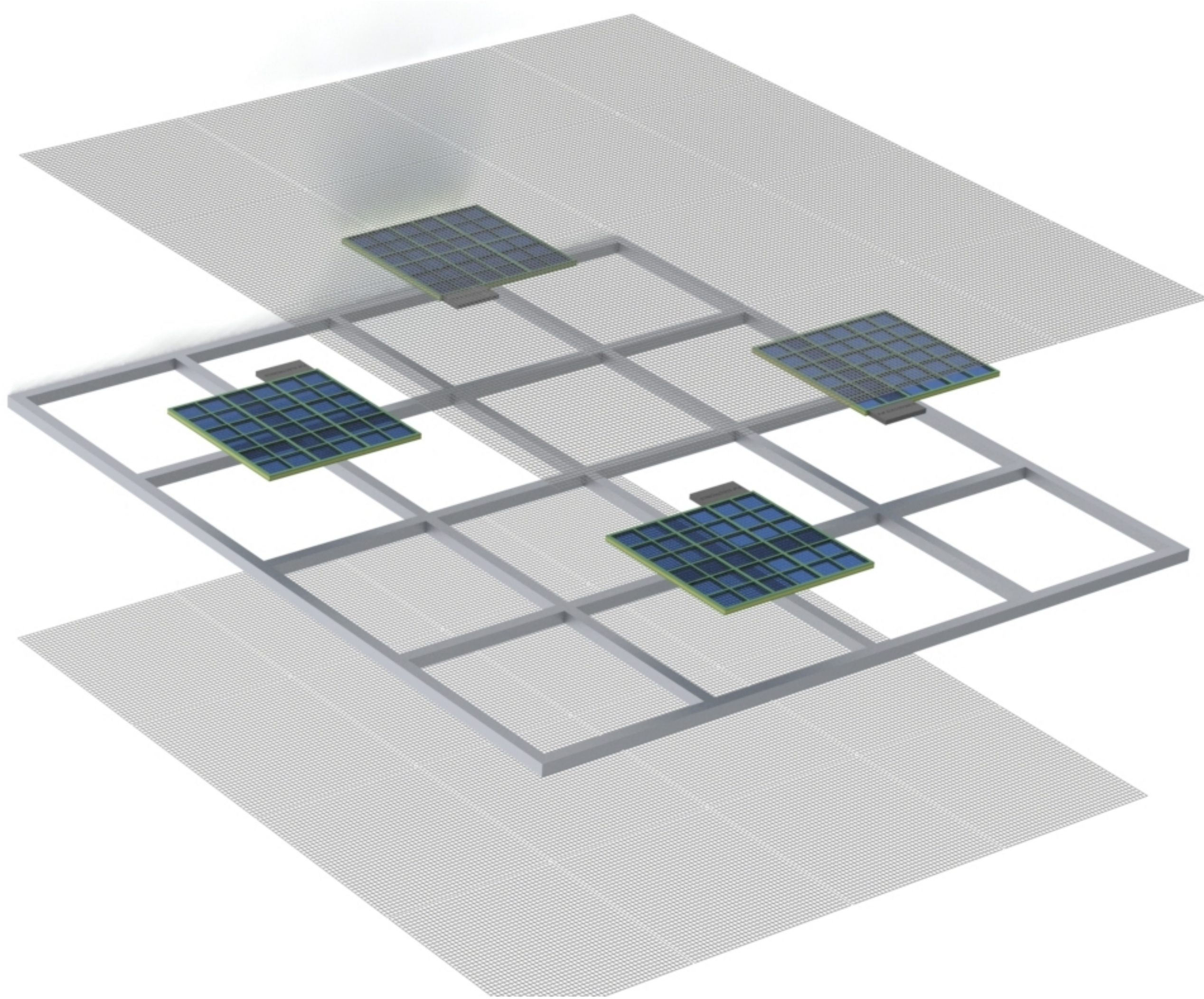


Major
R&D

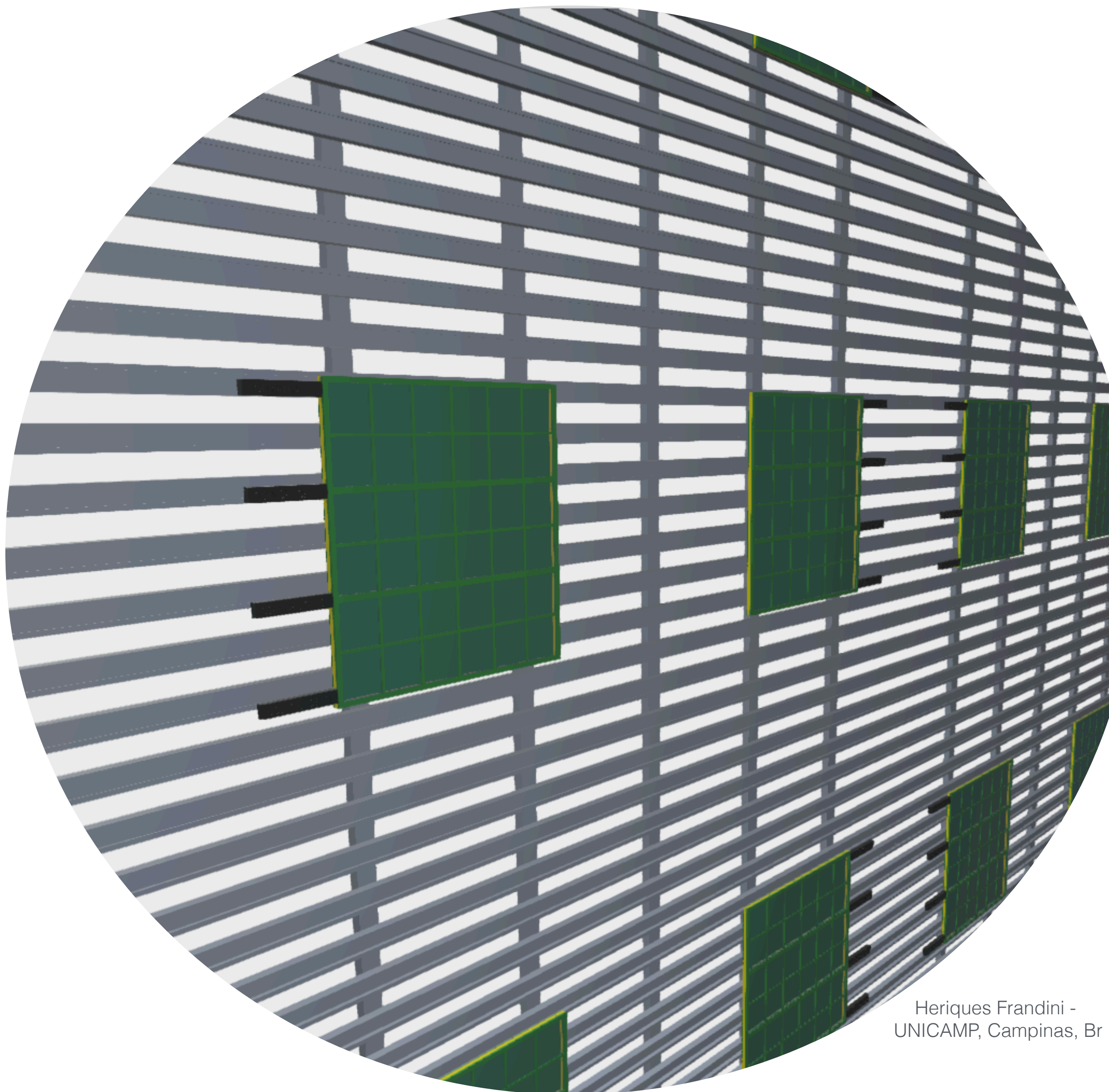
- Detector design and coverage: existing (X)ARAPUCA technology with SiPM photosensors is suitable for this application (flexible design opposite to PMT, **optimization for Xe light**)] **R&D**

- Fiber Routing (IN and OUT)] **Design Effort**

● $\sim 4\pi$ PD System design for the VD LAr Volume

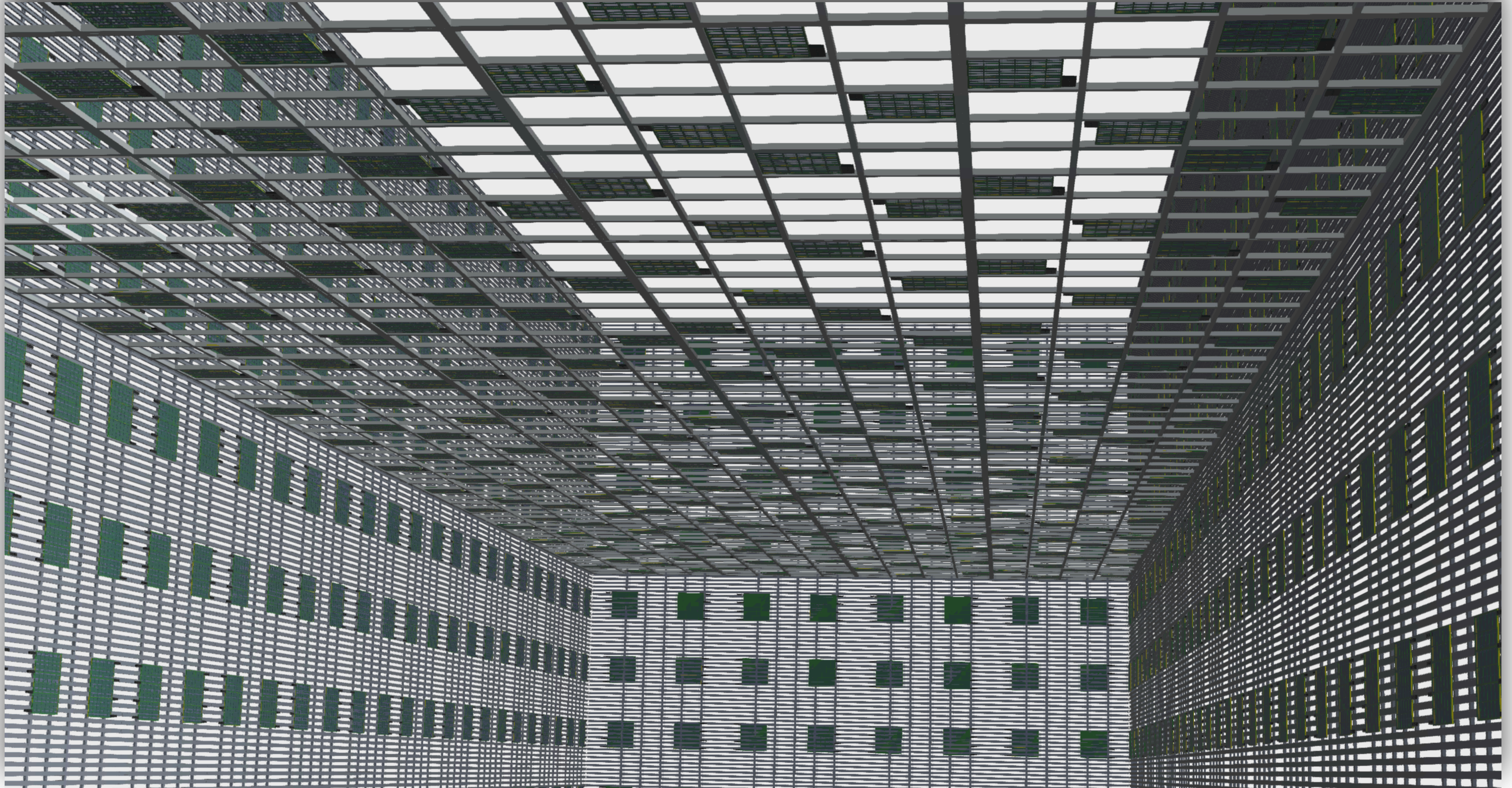


Photon Detector into the Cathode frame under conductor mesh (exploded view)



Photon Detectors hanging on Field Cage Walls

Heriques Frandini - UNICAMP, Campinas, Br



Heriques Frandini - UNICAMP, Campinas, Br

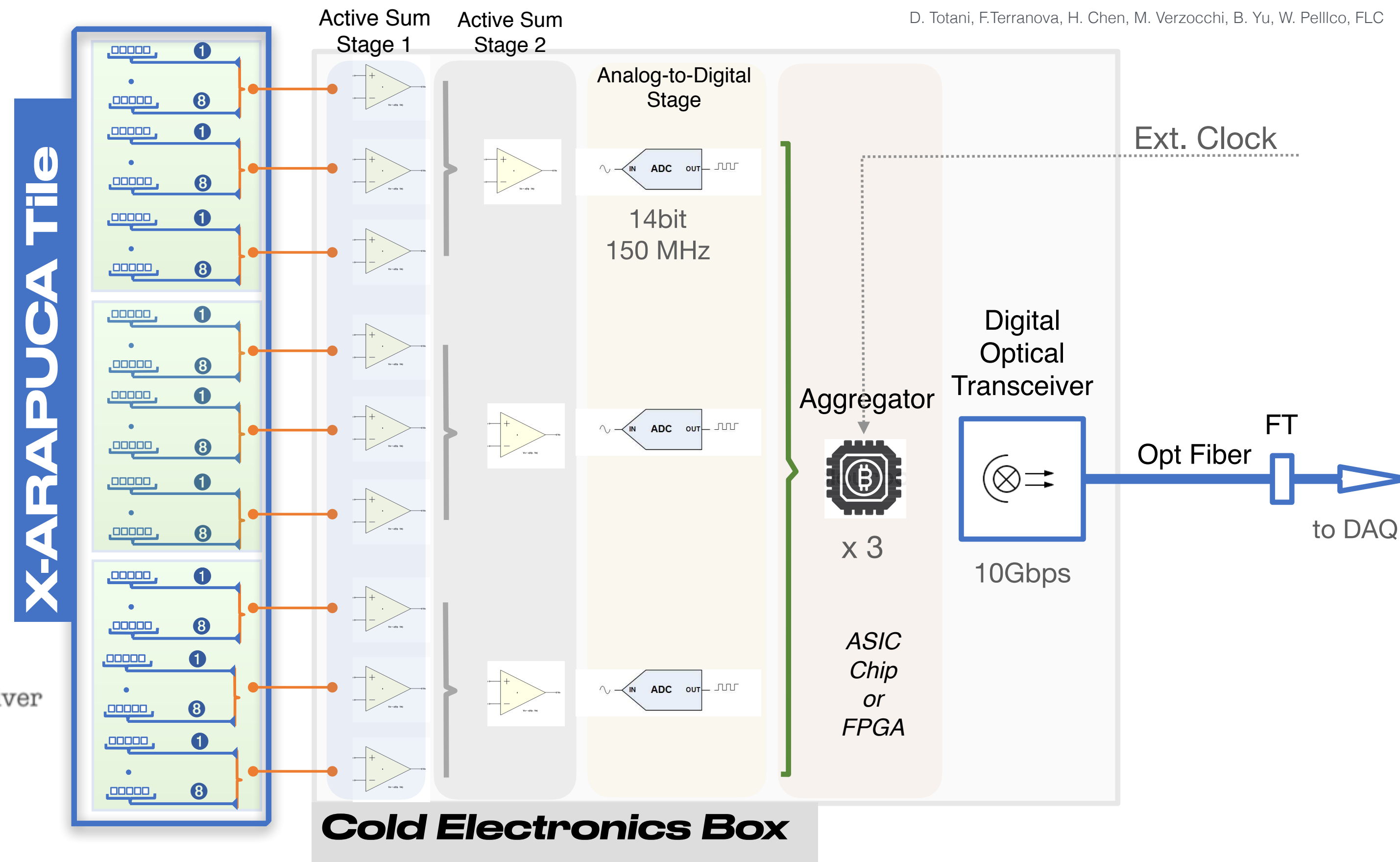
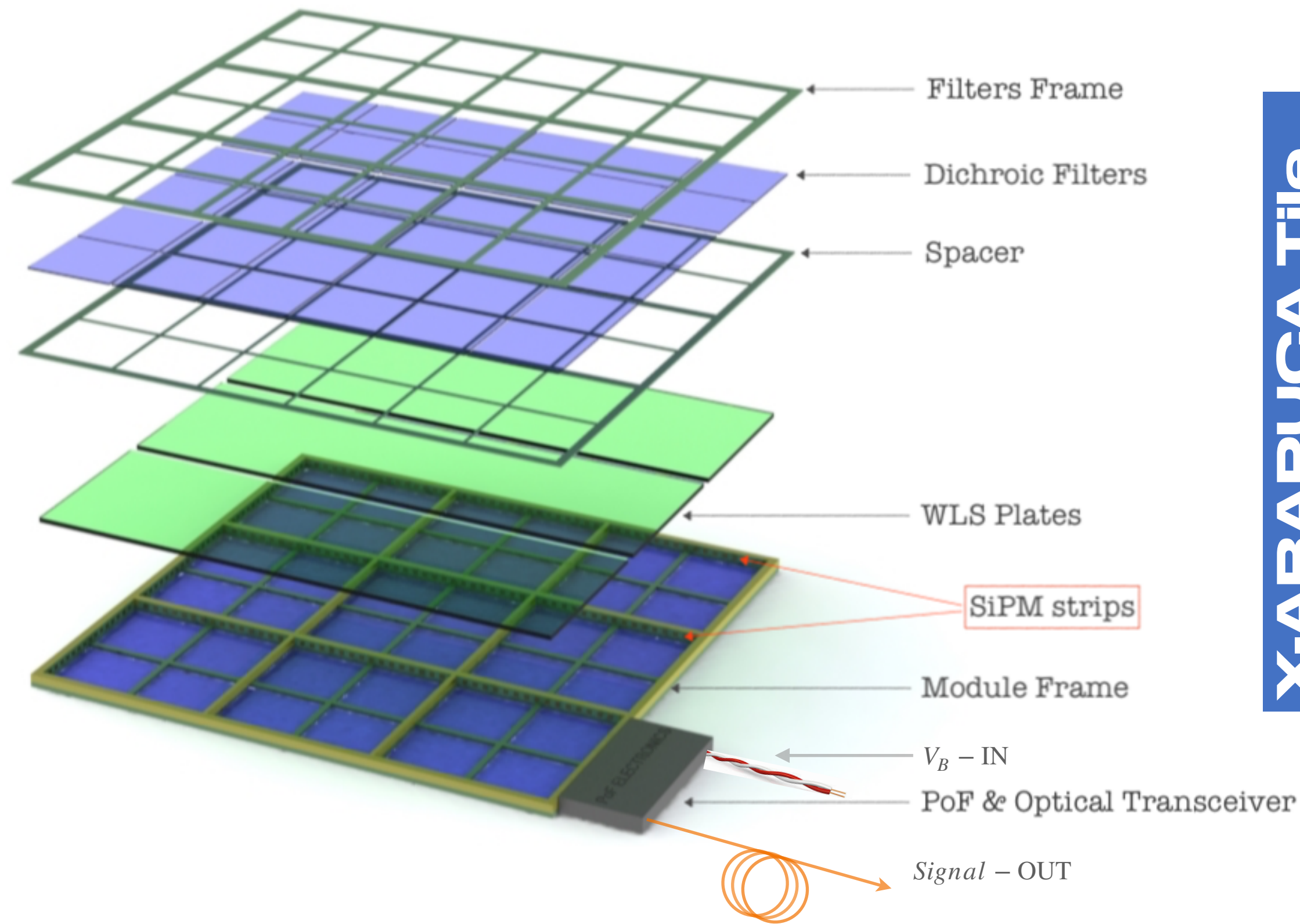
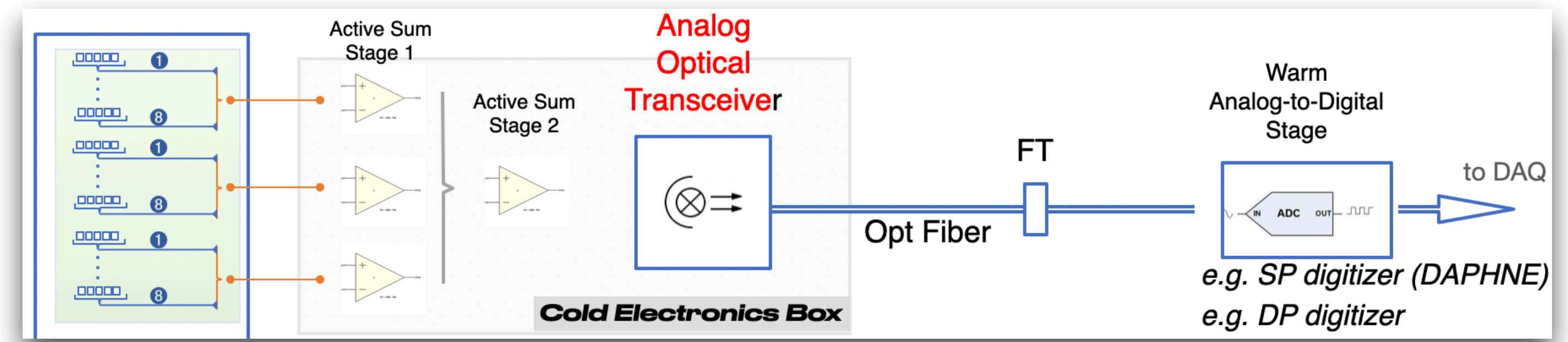


TABLE V. PD basic unit: X-ARAPUCA Tile

	Quantity	Dimensions
Area	1	$630 \times 630 \text{ mm}^2 = 0.4 \text{ m}^2$
Thickness	1	22 mm
Weight	1	~ 4.5 kg
Optical Area	2 (two-sided)	$600 \times 600 \text{ mm}^2 = 0.36 \text{ m}^2$
Sectors ("MegaCell")	3	$600 \times 200 \text{ mm}^2 = 0.12 \text{ m}^2$
Dichroic Filters	36×2	$100 \times 100 \text{ mm}^2$
WLS plates	3	$600 \times 200 \text{ mm}^2 = 0.12 \text{ m}^2$
PhotoSensors (SiPM)	360	$6 \times 6 \text{ mm}^2$
Read-out Channels	3	
SiPMs per channel	120	

OR (alternative)
depending on available solution for Optical Transceiver



● $\sim 4\pi$ PD System design for the VD LAr Volume

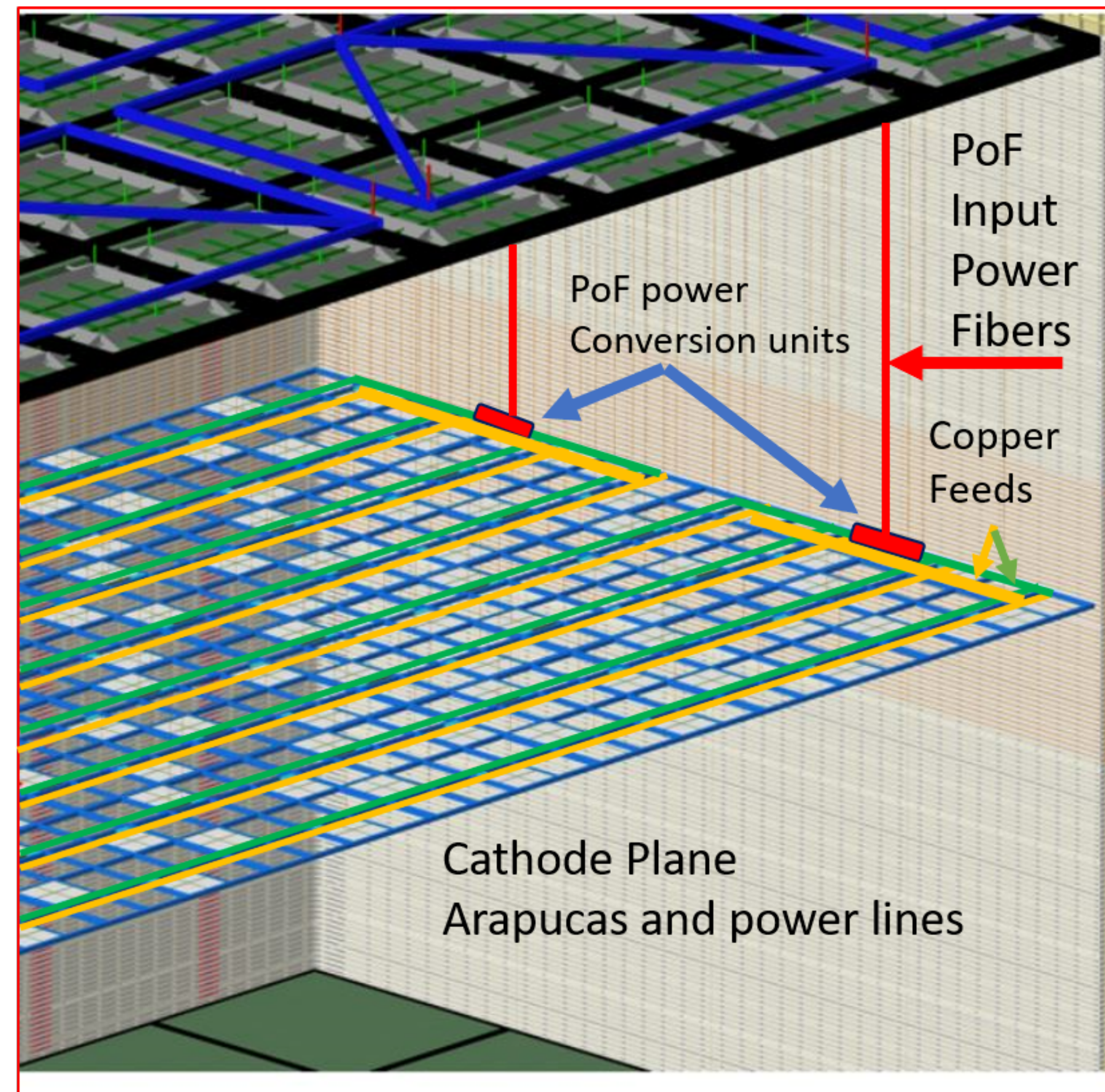
W. Pellico - FNAL

TABLE VI. VD PDS $\sim 4\pi$ -Configuration

Item	Quantity	HV Surface
X-ARAPUCA Tiles	320 double-side 768 single-side	Cathode plane Field Cage walls
Dichroic Filters	50,688	Cathode plane Field Cage walls
WLS plates	3,264	
PhotoSensors (SiPM)	115,200 207,360	
Signal Channels	960 2,268	Cathode plane Field Cage walls
Fibers (Serialized Channels)	1088	Cathode plane Field Cage walls
SiPMs per channel	120 90	
Optical Area	115 m ² + 115 m ² 277 m ²	
Active coverage	14%	Cathode plane Field Cage walls

*HD-Single
Phase Module*

1500
48,000
6000
288,000
6000
48
(anode side only)
12%



- The role of Xe doping
- $LY(x, y, z)$ - the detector Light Yield with a $\sim 4\pi$ Optical Coverage
- Energy Reconstruction and Energy Resolution for low-E UG events
- Position Reconstruction and Space Resolution for low-E UG events
- First evaluation of Trigger Efficiency in Fiducial VD Volume for low-E UG events
- Time Resolution study in progress (not reported today)

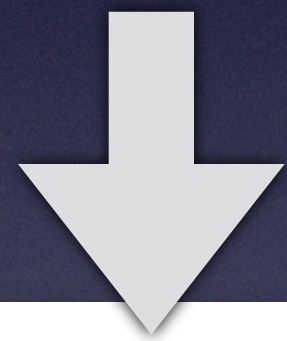
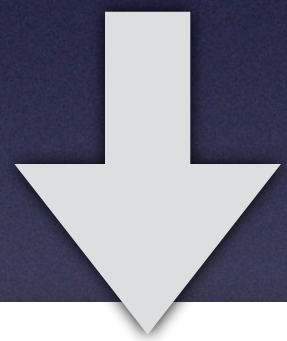


TABLE IV. Requirements and Physics purposes for the VD Photon Detector System - $\sim 4\pi$ -configuration option

Detector Requirement	Value	Physics Purpose (*)
Trigger efficiency for interactions with energy deposit $E_{dep} \geq 5$ MeV in 100% of detector fiducial volume	$\geq 99\%$	- SN burst trigger up to the Large Magellanic Cloud (50 kpc) yielding 10 interactions in 10 kt LAr - Low-energy background rejection
Spatial resolution for interactions with energy deposit $E_{dep} \geq 10$ MeV	≤ 1 m	- Background rejection for SN, solar, nucleon decay
Energy resolution for interactions with energy deposit $E_{dep} \geq 5$ MeV	$\leq 8\%$	- Identification of SN spectrum features from different SN dynamical models
Time resolution	≤ 200 ns	- SN burst triggering - Identification of SN time features due to standing accretion shock instabilities - Identification of neutrino "trapping notch" (SN dip in luminosity)

Ph. Detector Performance: Expectation from first MC simulations of Response

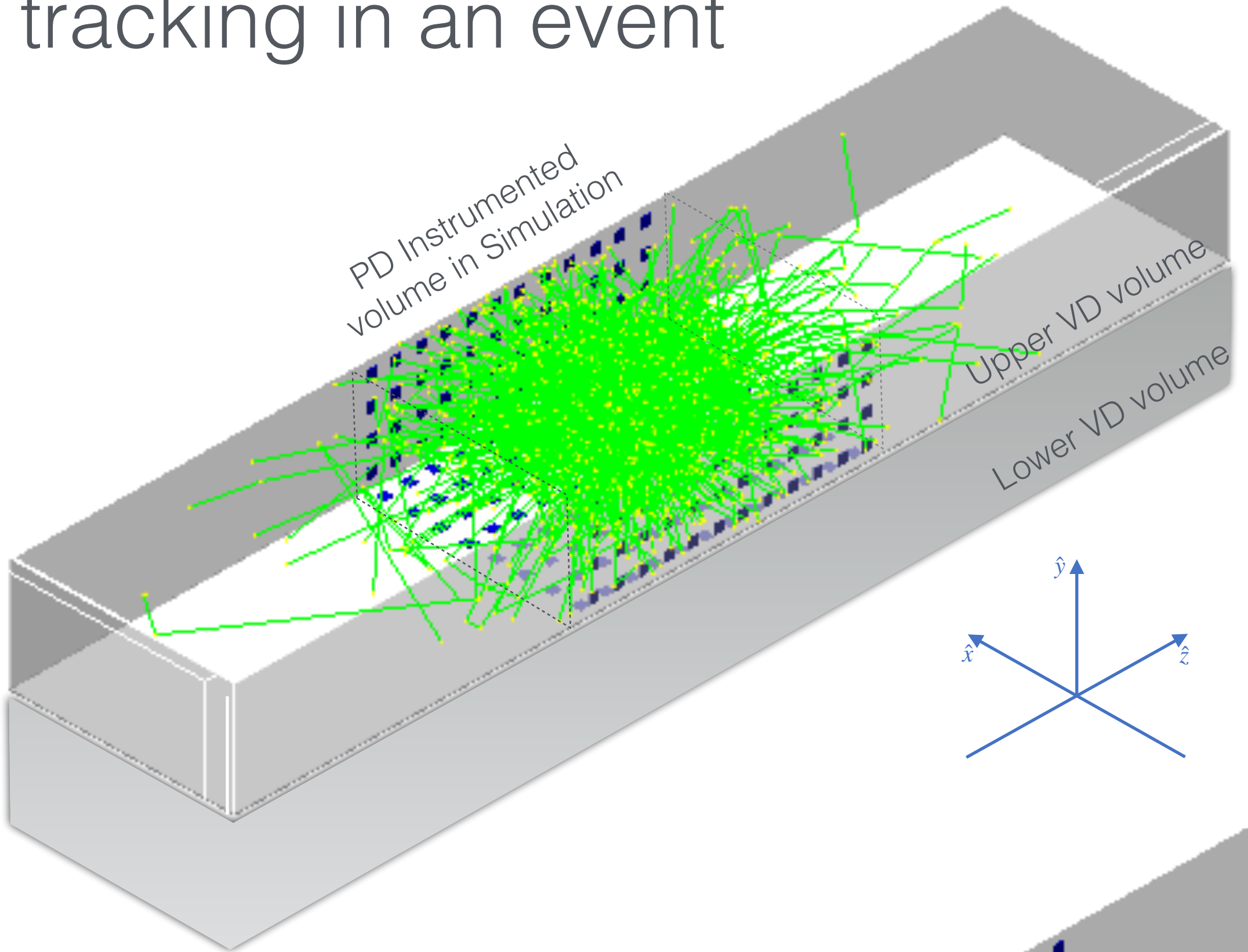
L. Paulucci, Universidade Federal do ABC, Santo André, SP, Brazil

F. Marinho, Universidade Federal de São Carlos, Brazil

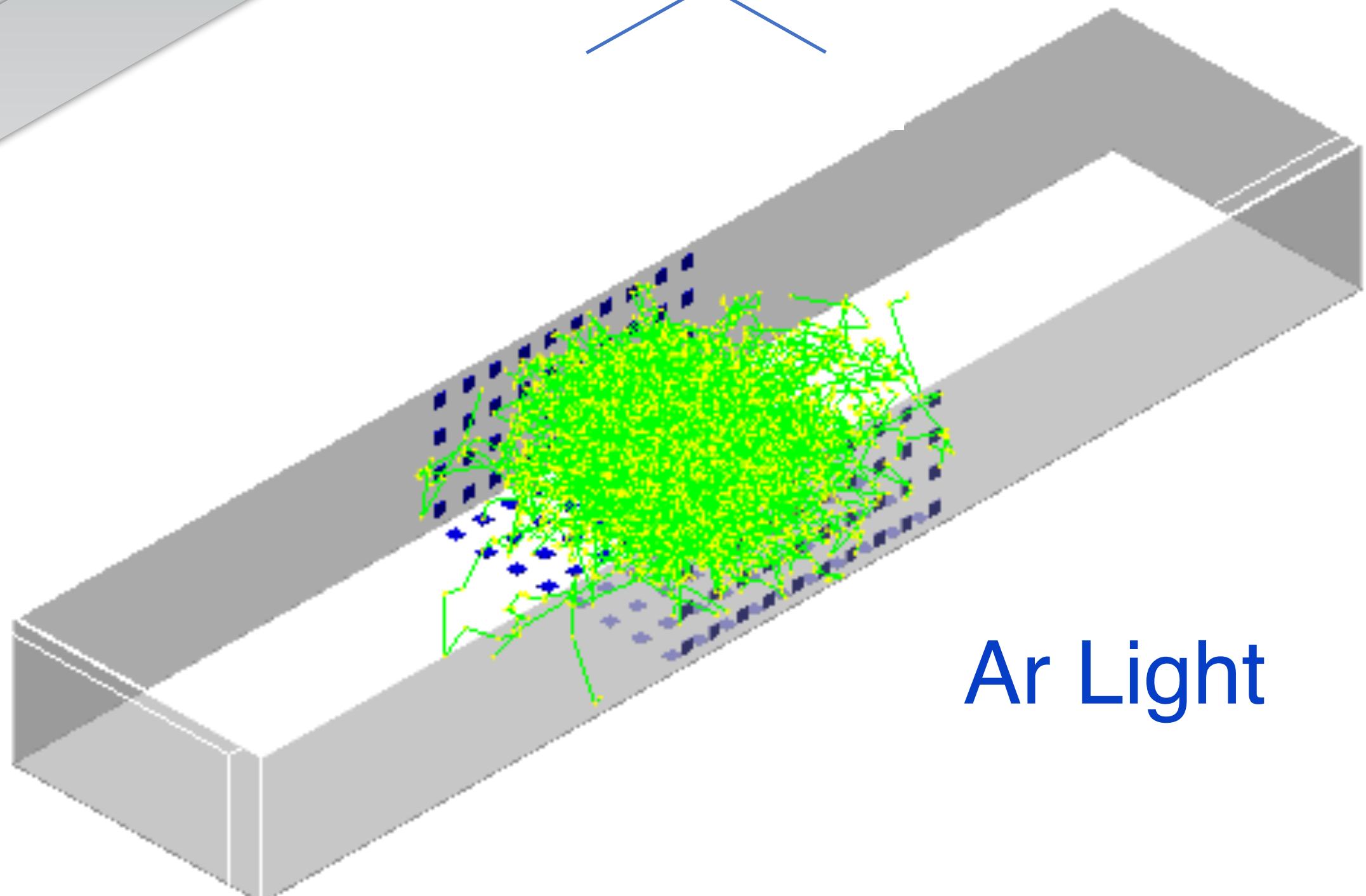
D. Totani, University California Santa Barbara, USA

* I. Gil-Botella - for fundamental advice on Low-E Physics and PD concepts

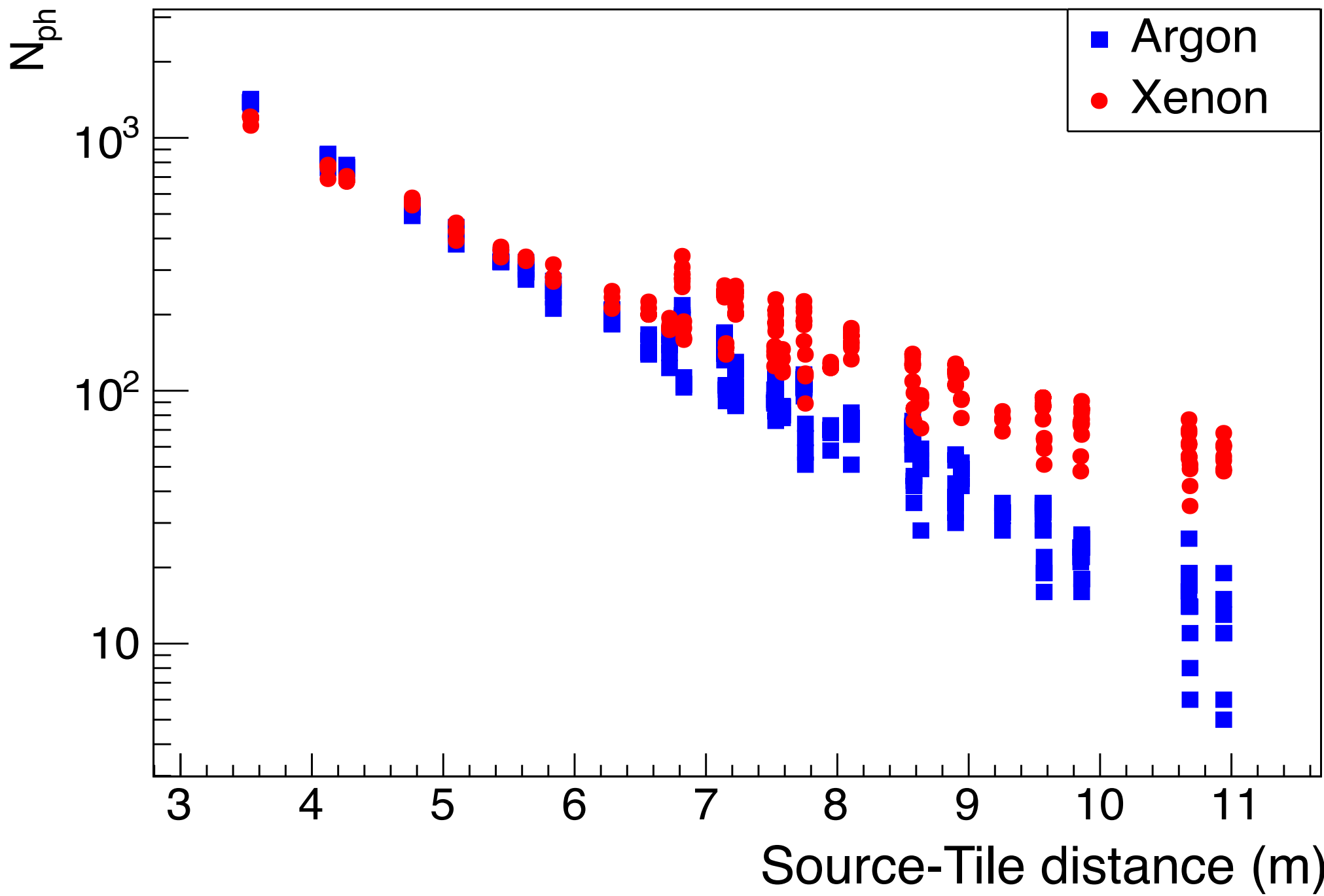
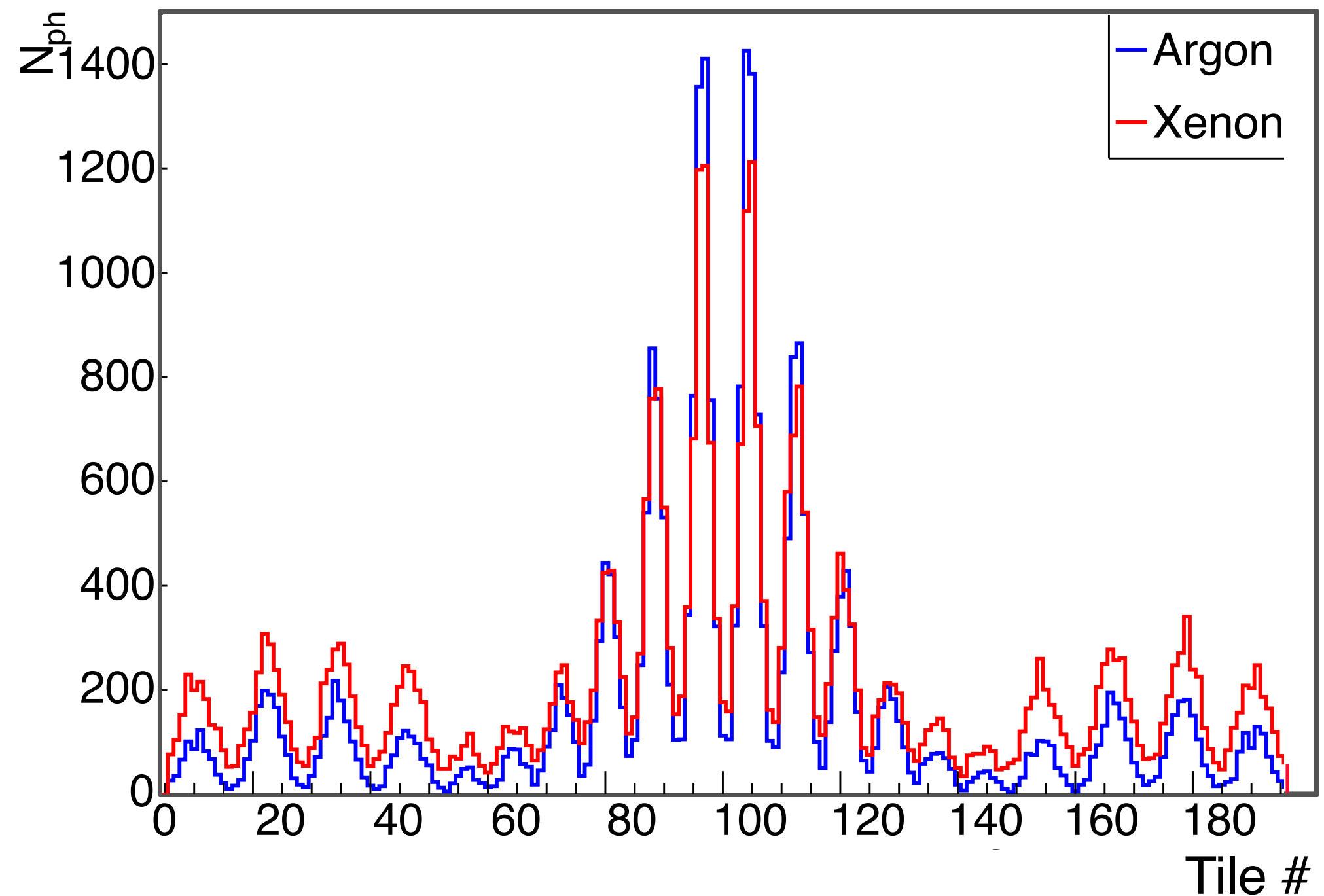
GEANT4 photon tracking in an event



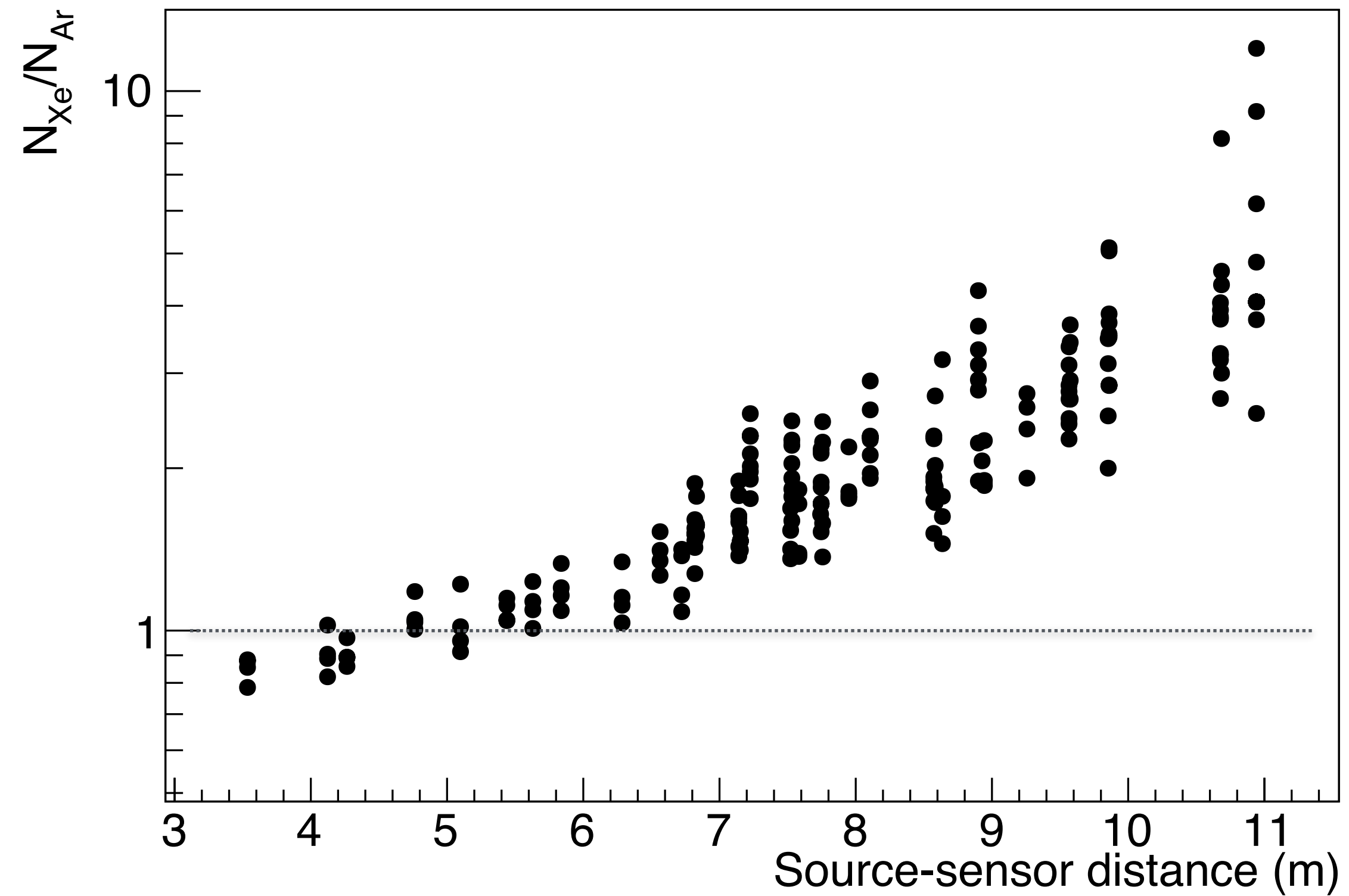
Xe Light



Ar Light



The collected light is found to be ~30% larger for Xe-doped Argon (8.2% of the total emitted photons reach the PD optical surface vs 6.3% in case of pure Ar), due to the effect of the longer Rayleigh scattering length enhancing collection probability for light emitted at longer distances from the PD-detectors



This supports the choice of Xe-doped Ar as scintillation medium.

The VD $\sim 4\pi$ -PD w/ $\sim 14\%$ Optical Coverage

Ar+ Xe(10 ppm)

$$\langle LY \rangle \simeq 60 \frac{PE}{MeV}$$

$LY \left[\frac{PE}{MeV} \right]$

Anode

\hat{y} (m)

6.5

Cathode 0

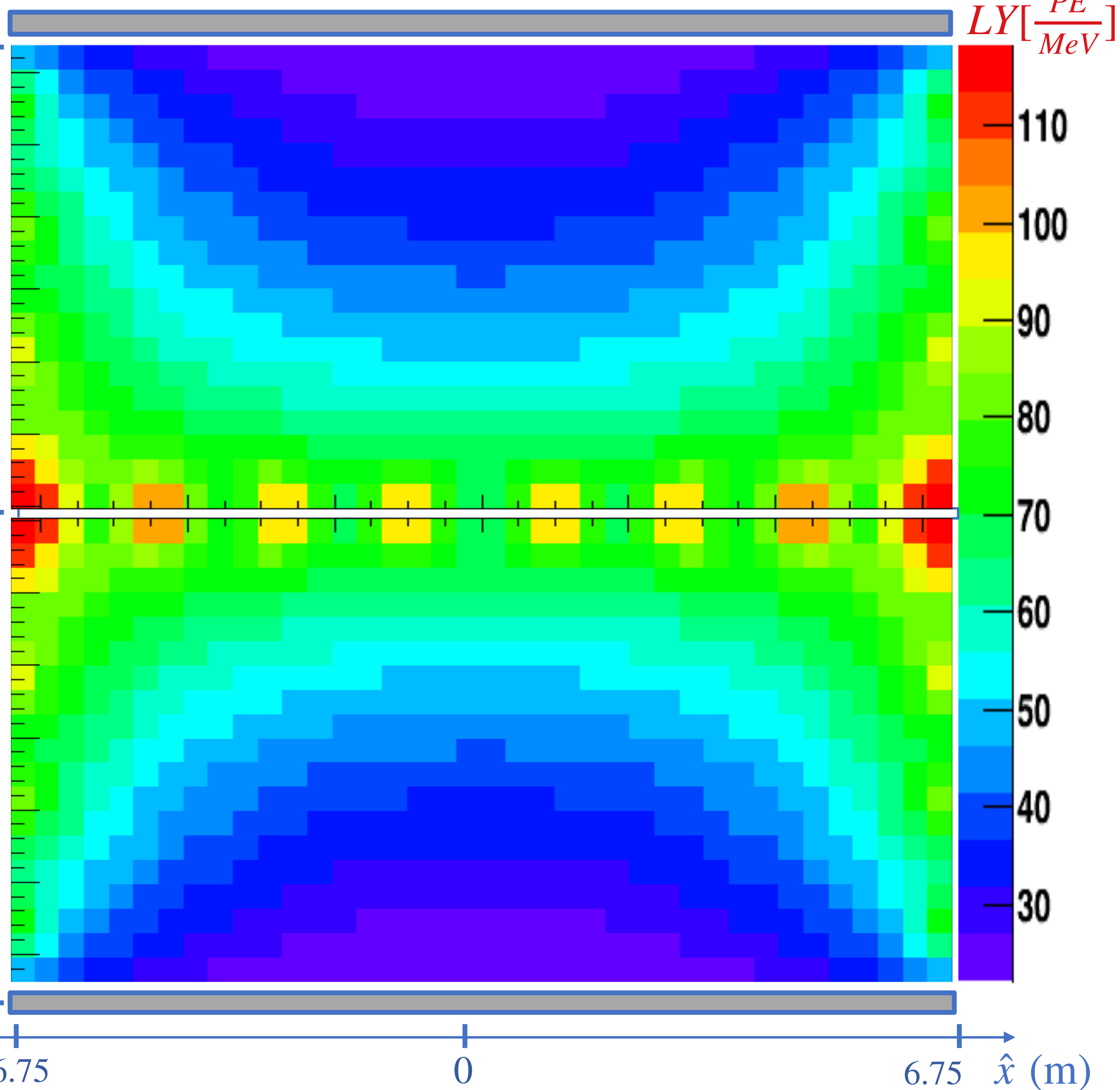
Anode -6.5

-6.75

0

6.75 \hat{x} (m)

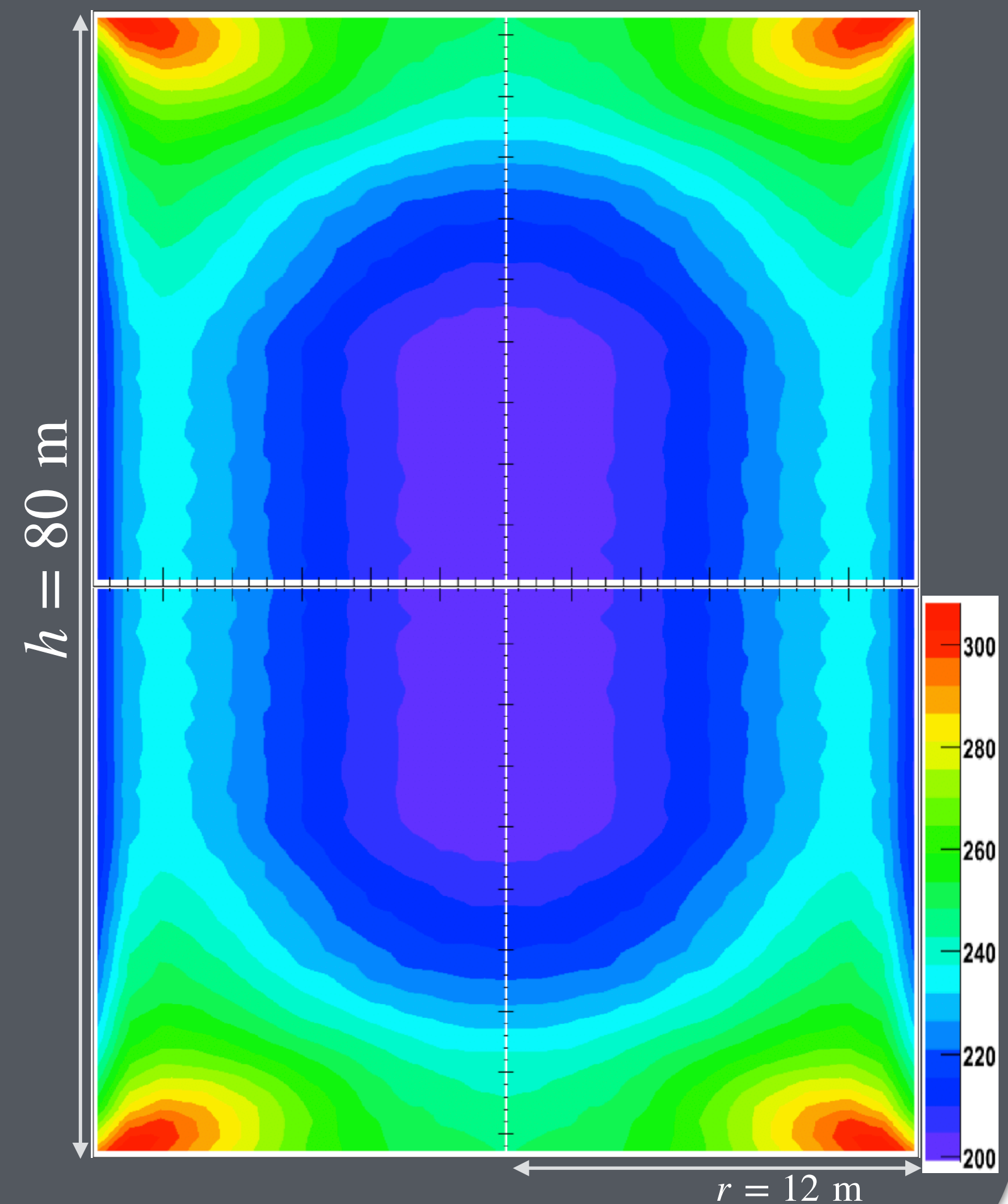
Light Yield Map in the Detector Transverse plane



LAGUNA-LBNO design study
for Solar Neutrino
and SN/DSNB Experiment

LENA - 4π LiqScint w/ 30% O.C.

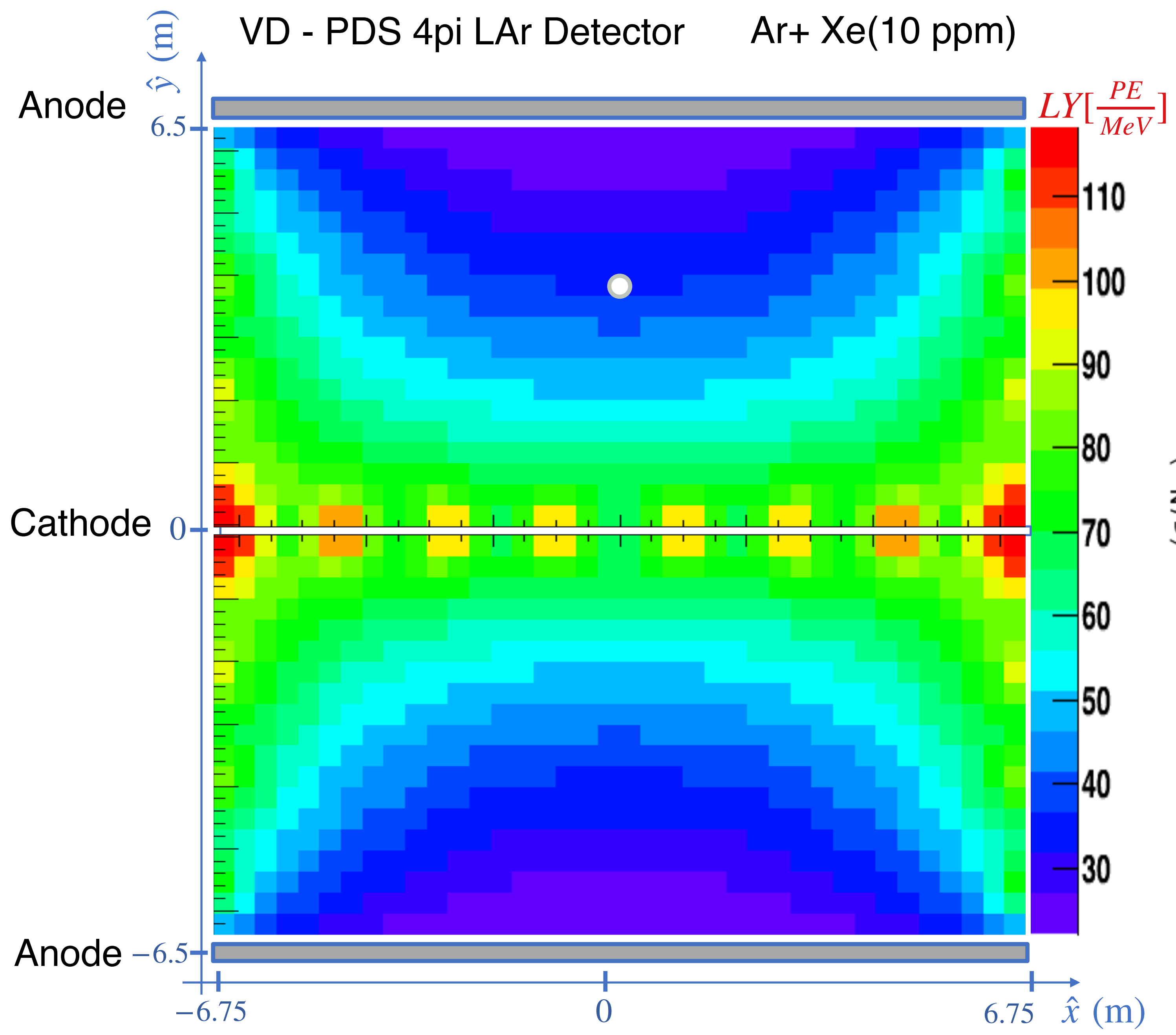
$$\langle LY \rangle = 180 - 200 \frac{PE}{MeV} \quad E_{Thr} = 0.25 \text{ MeV}$$



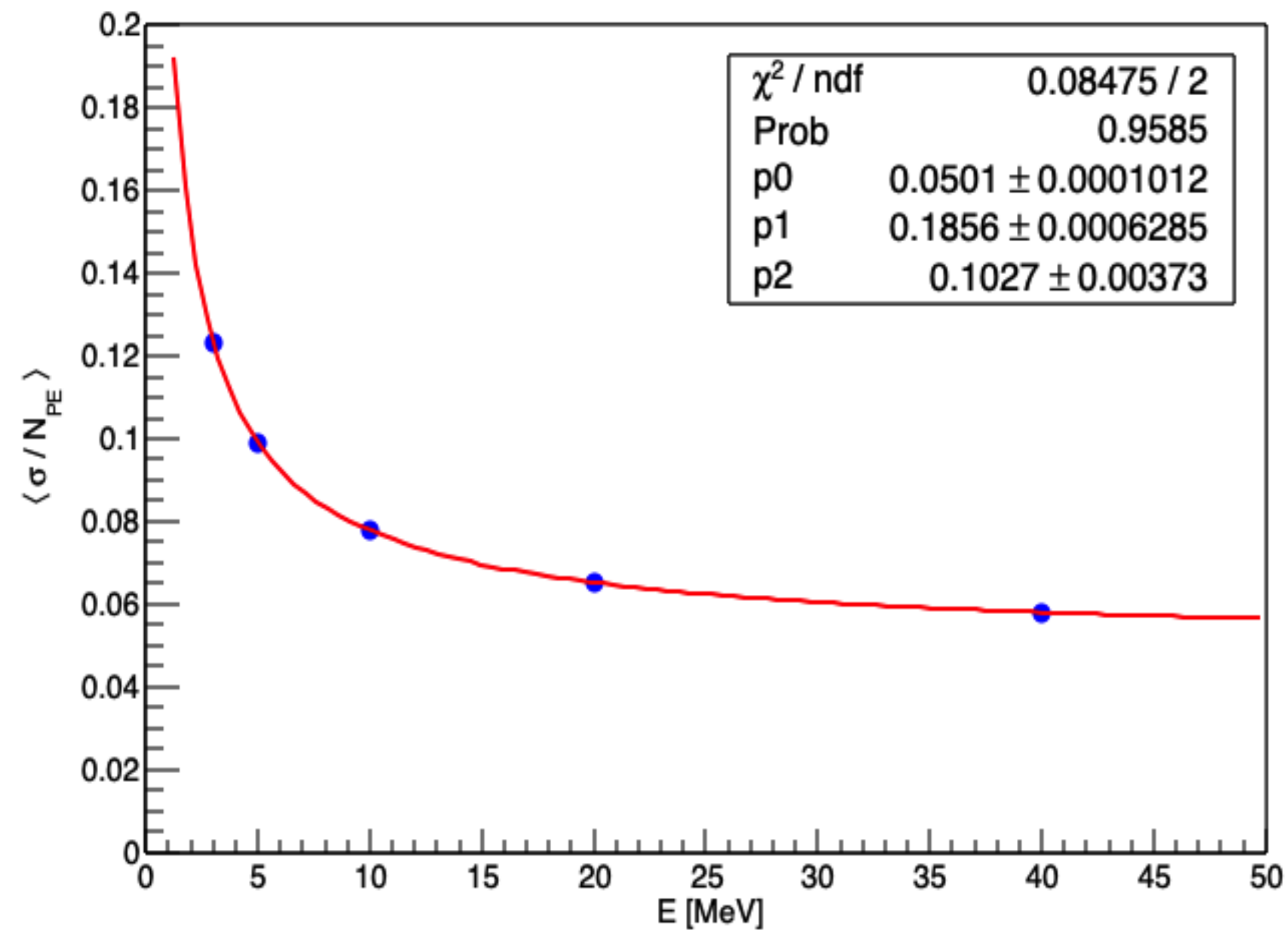
Calorimetric Energy reconstruction from detected photon counting

Energy Resolution

from statistical fluctuation (p1) on the number of detected PEs and to uncertainty on energy calibration (p0)



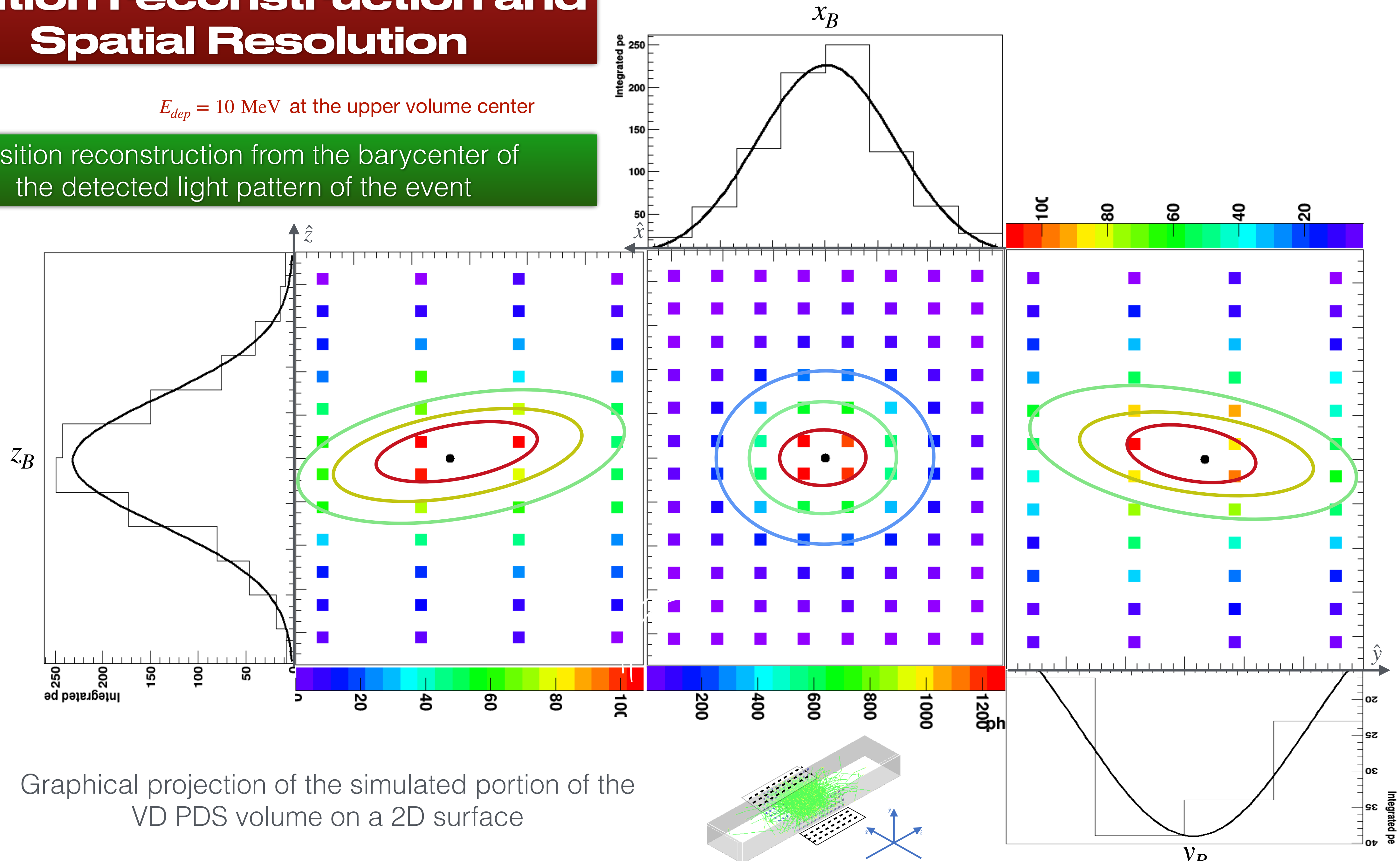
PD Resolution



Position reconstruction and Spatial Resolution

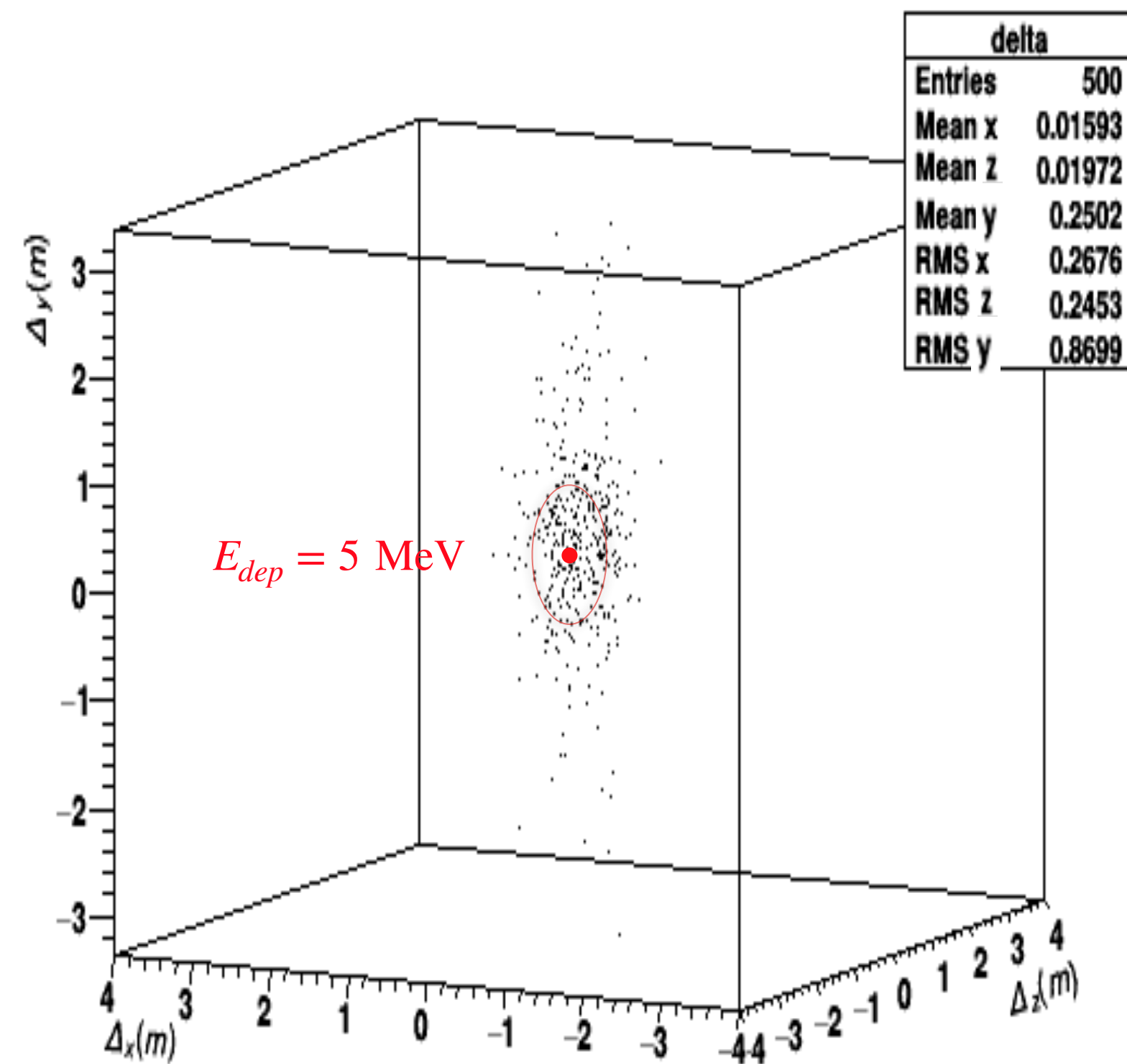
$E_{dep} = 10$ MeV at the upper volume center

position reconstruction from the barycenter of the detected light pattern of the event



Graphical projection of the simulated portion of the VD PDS volume on a 2D surface

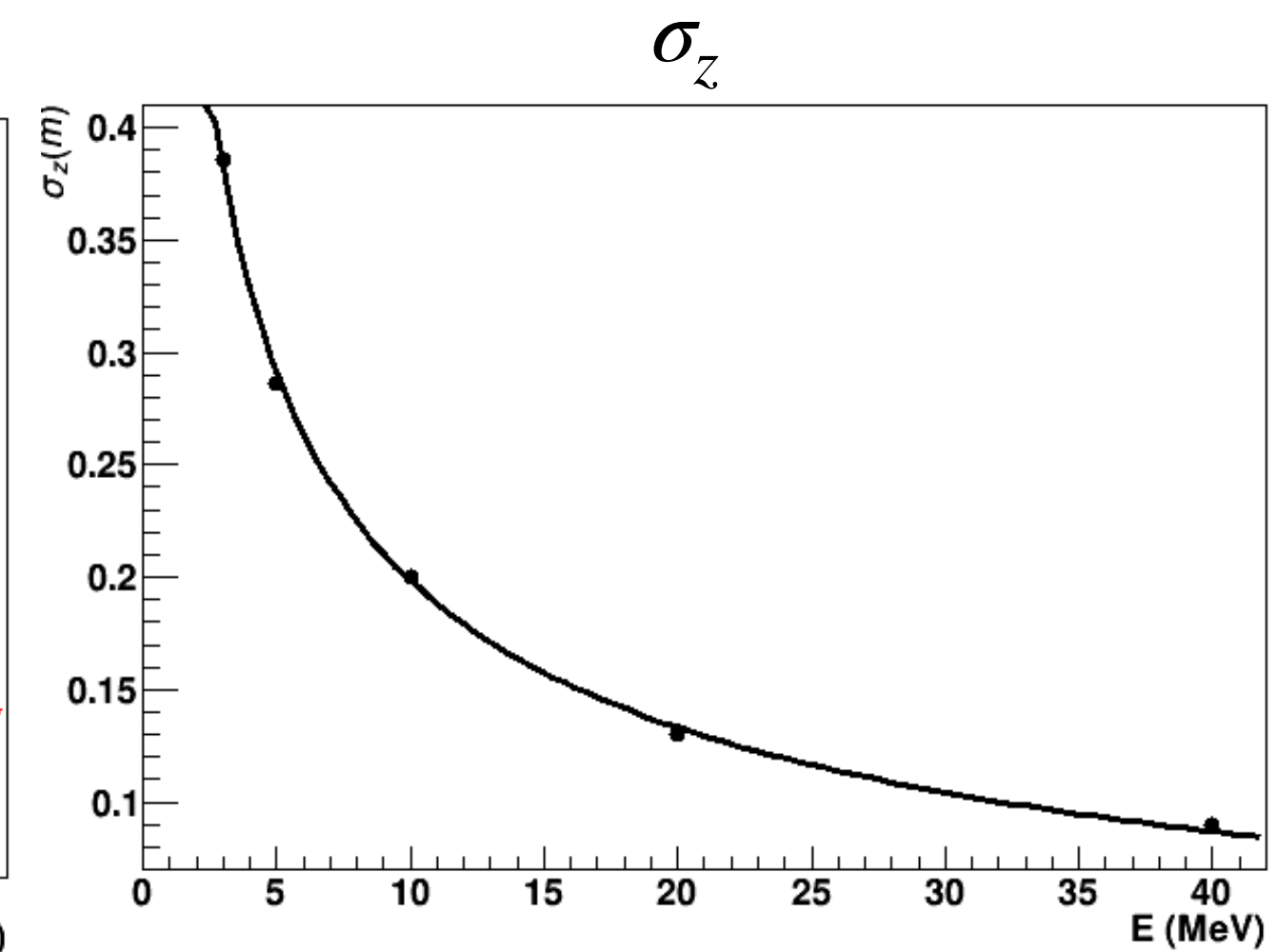
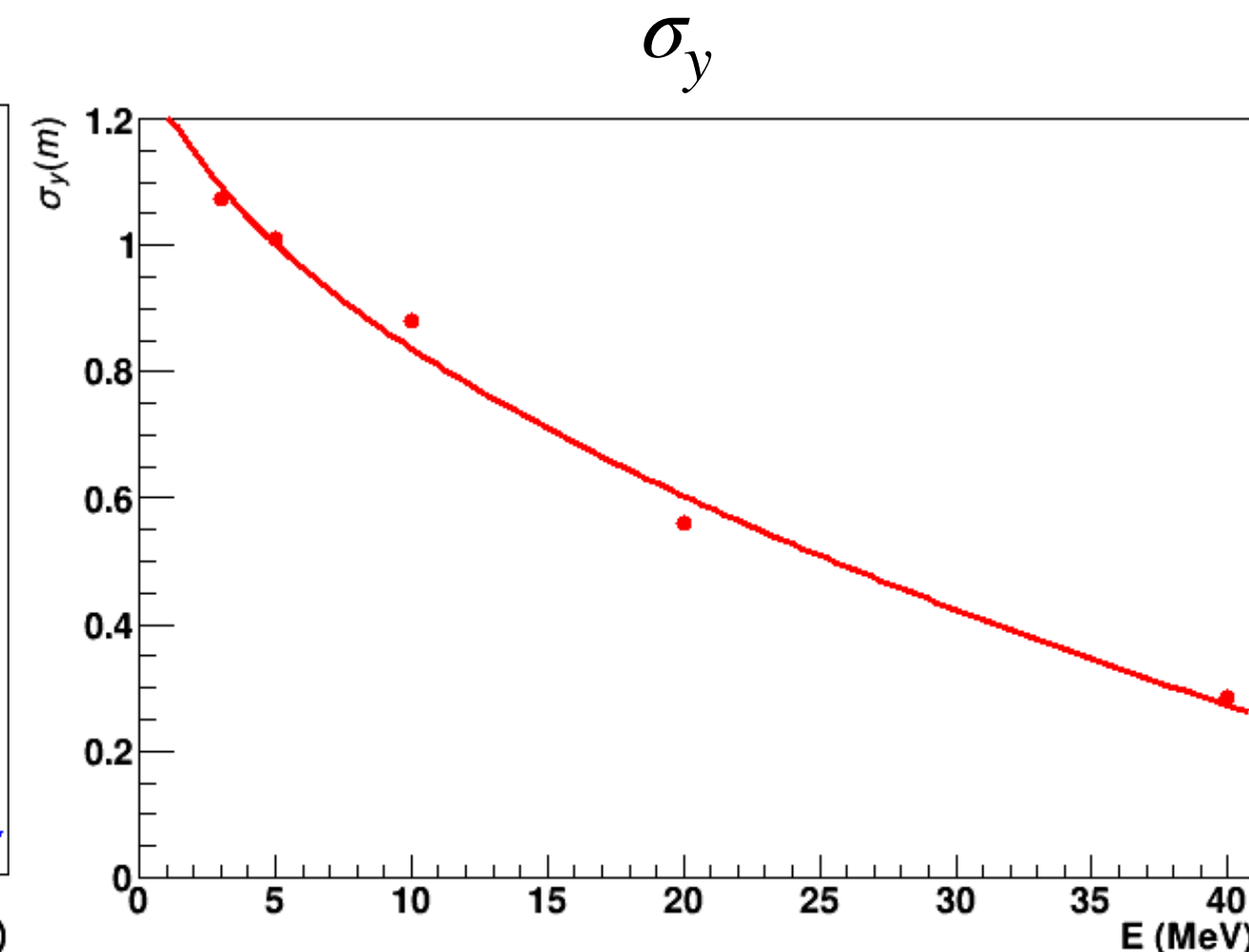
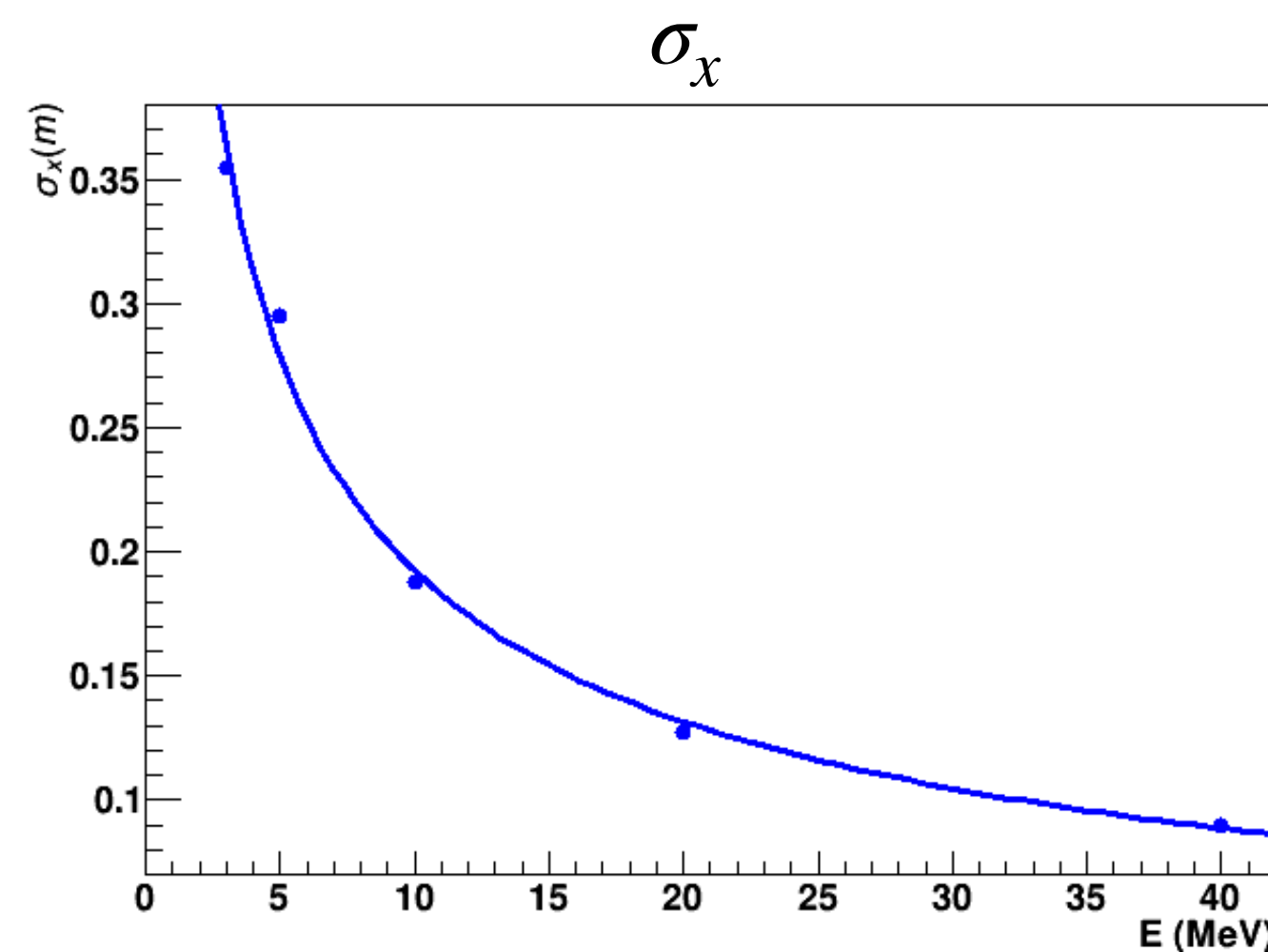
Reconstructed position in space for a sample of 5 MeV MC events generated at a fixed position



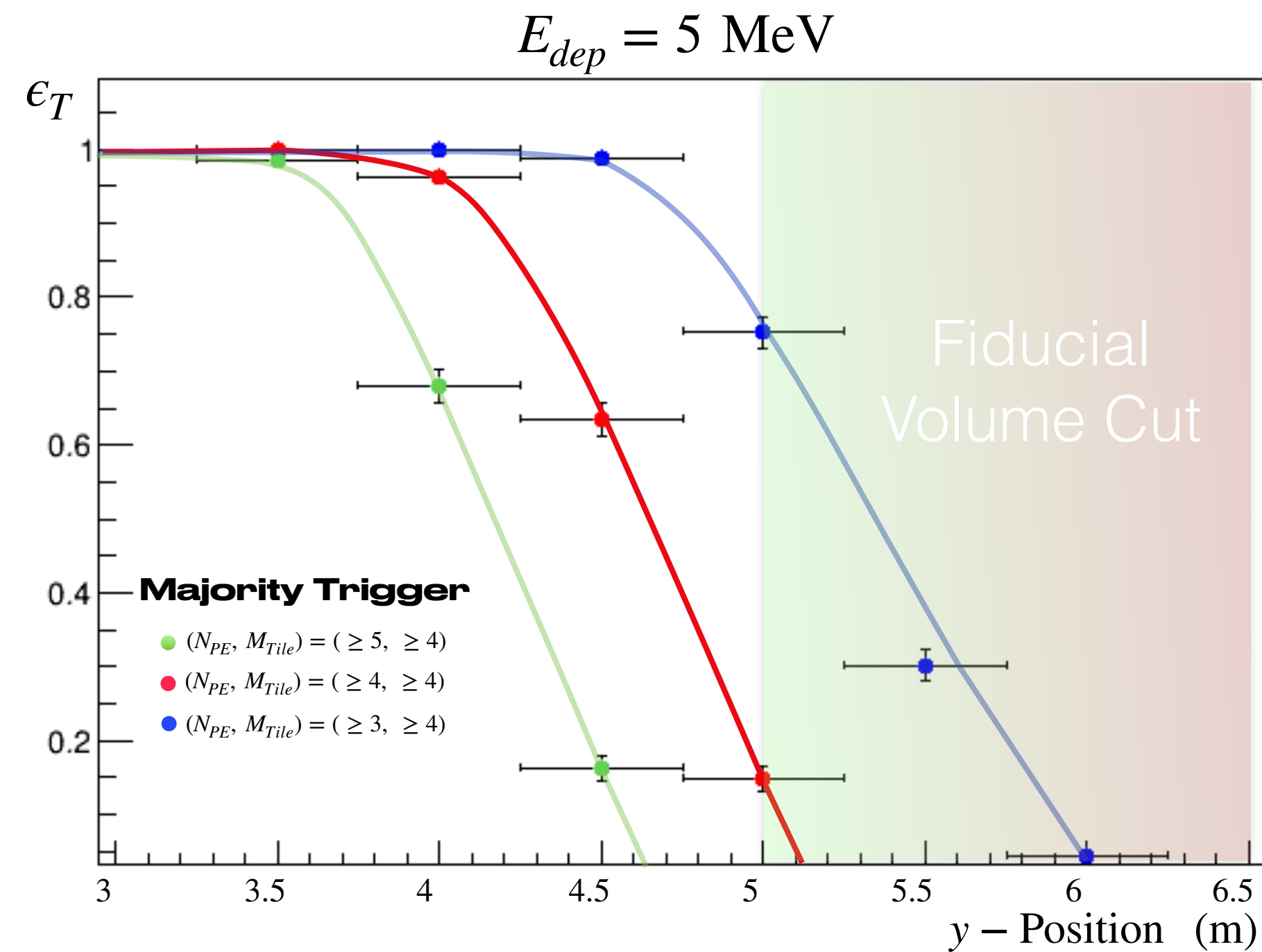
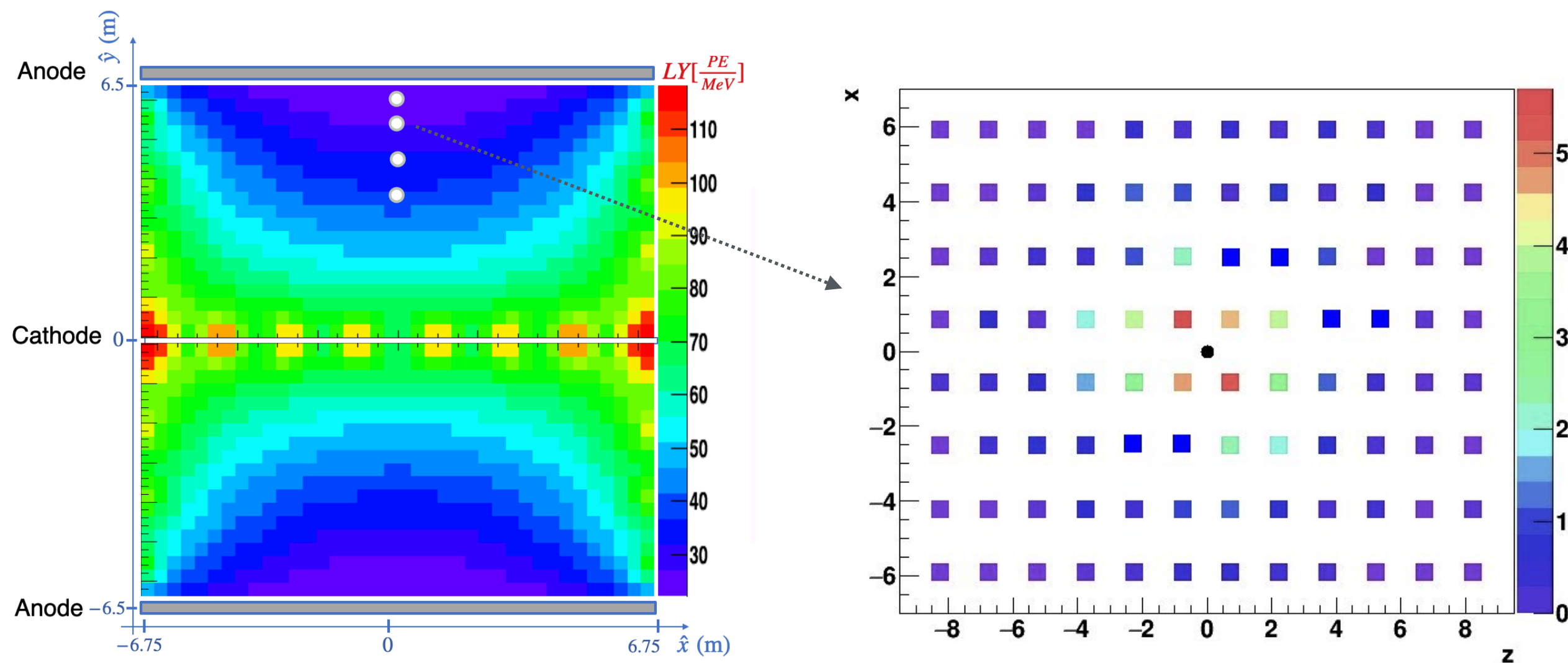
Position resolution \approx inversely proportional to the square root of the number of photons detected in the event

Good Position resolution in \hat{x} , \hat{z} ($\sigma_{x,z} \leq 30 \text{ cm}$)
worse in \hat{y} ($\sigma_z \leq 1 \text{ m}$), due to less n. of PD tiles along VD direction

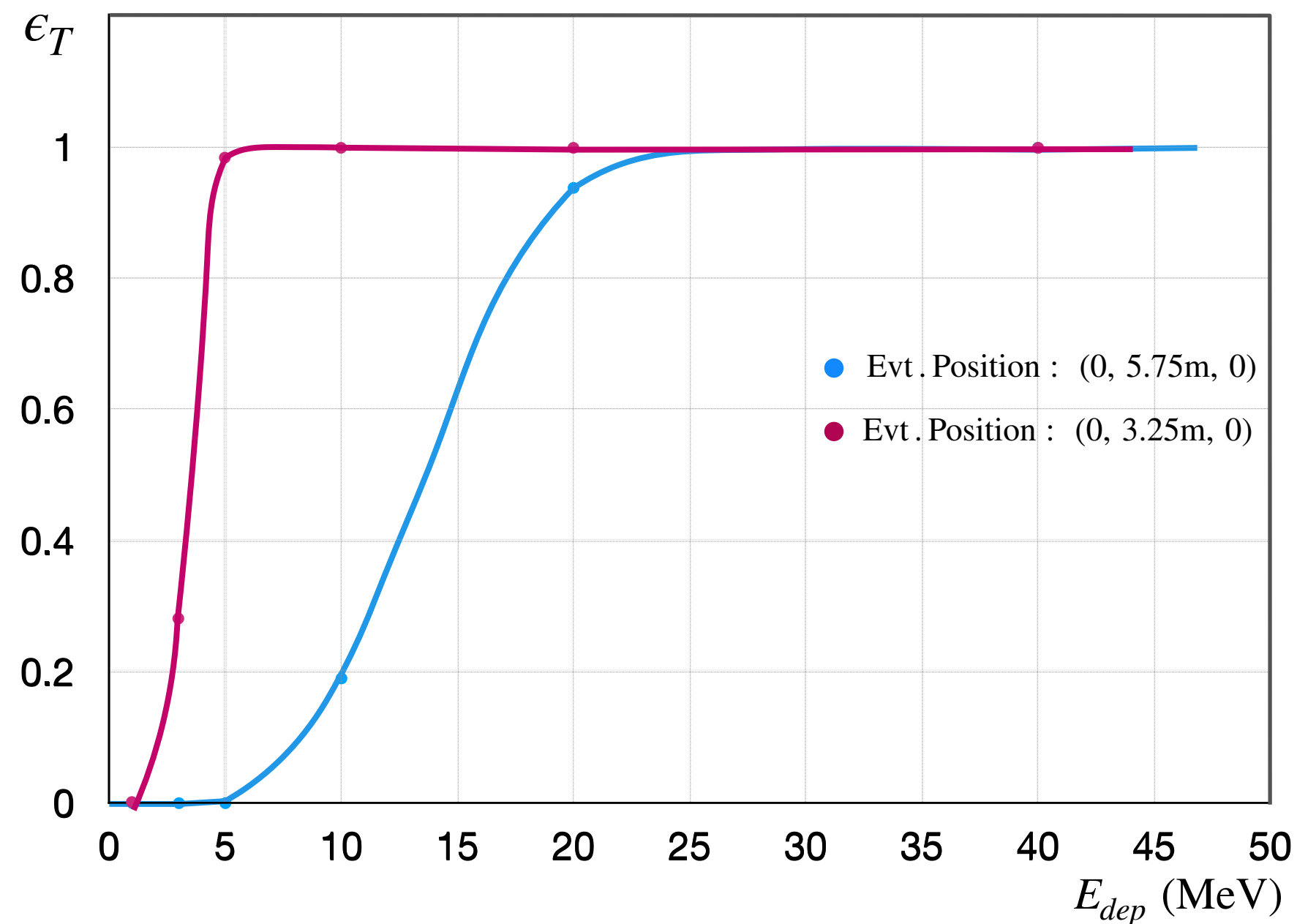
Timing information (not used here) should improve Space Resolution



(N_{PE}, M_{Tile}) – Majority Trigger condition



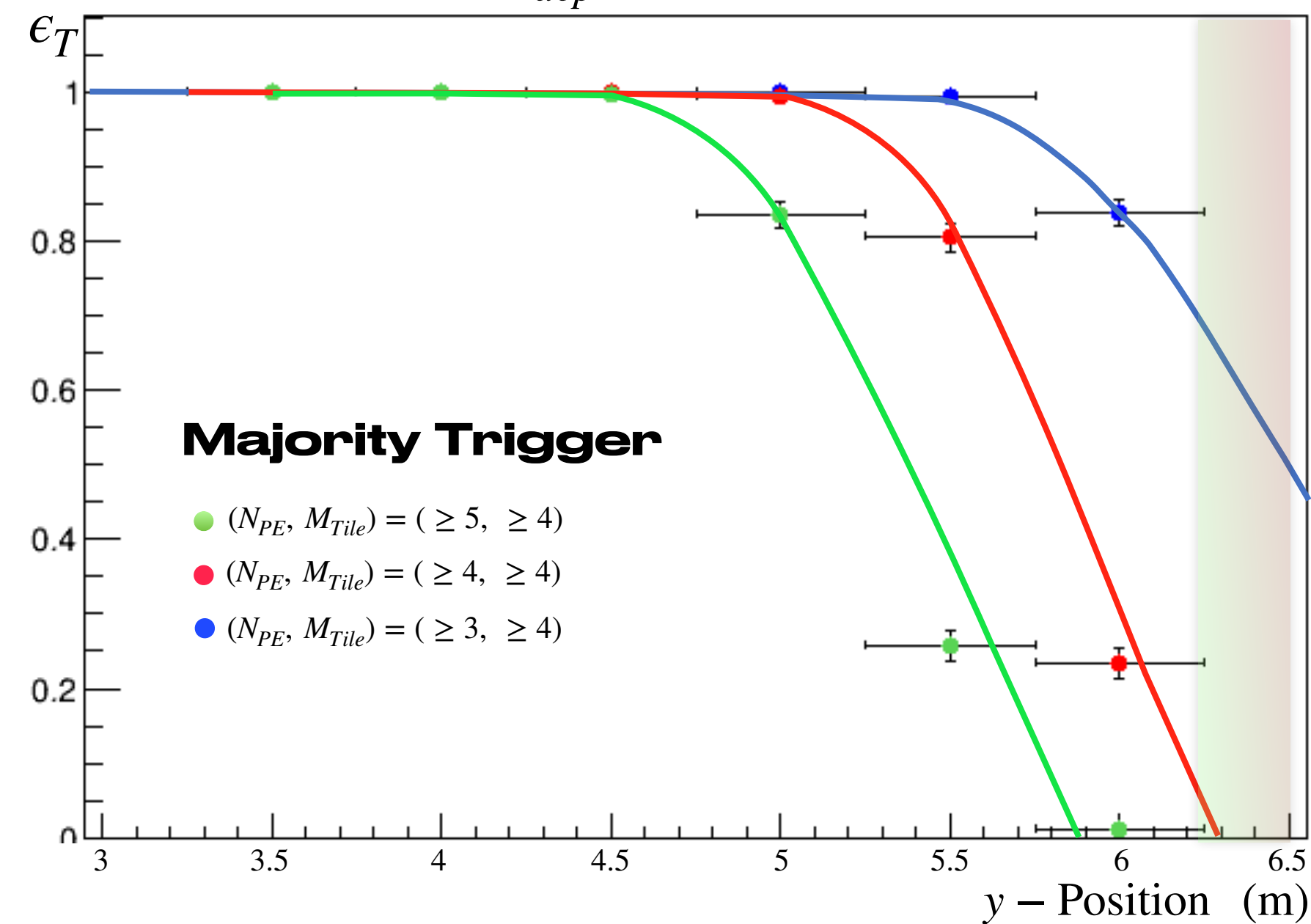
Majority Trigger $(N_{PE}, M_{Tile}) = (\geq 5, \geq 4)$



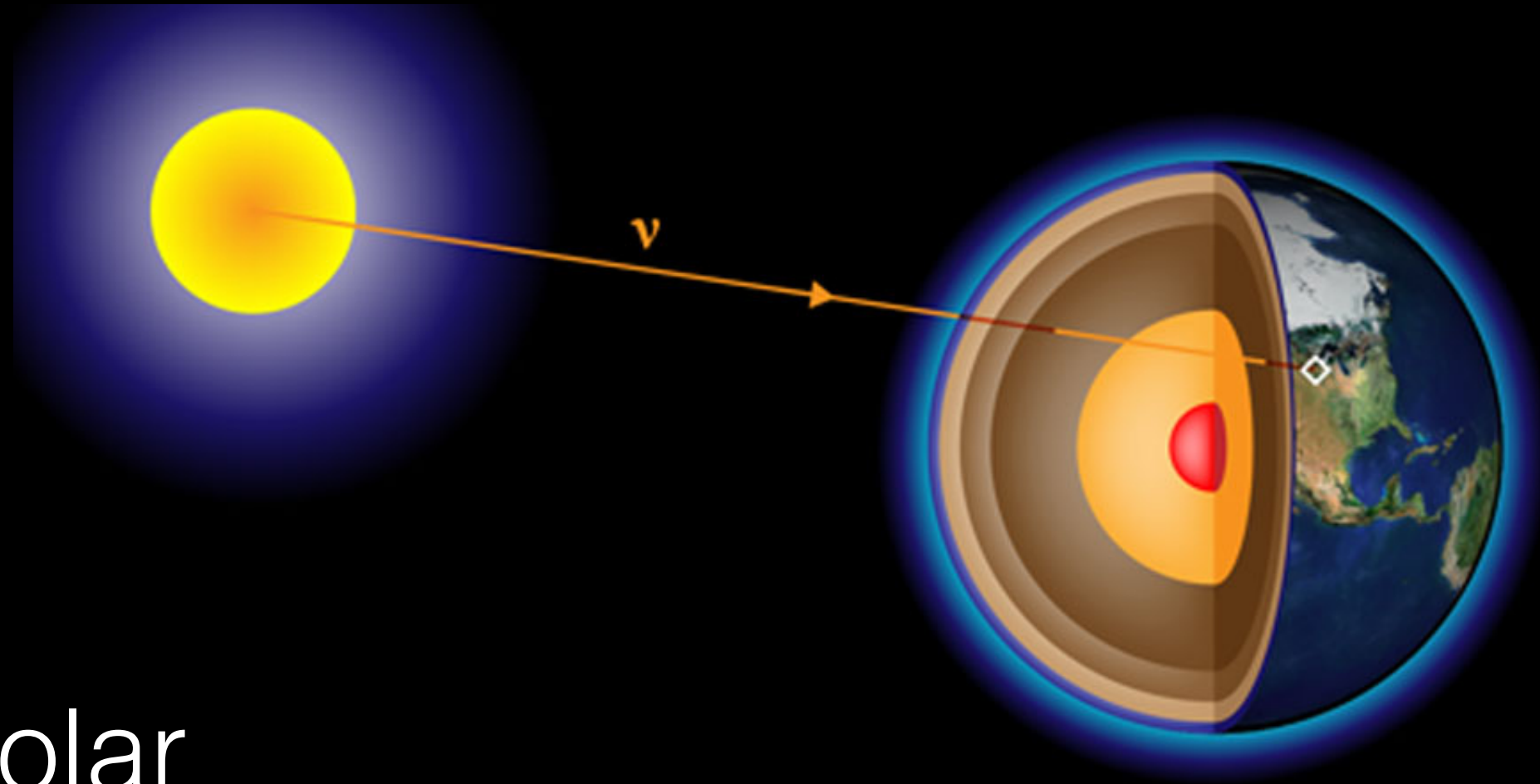
Relaxing (N,M)-Majority requirements enhance trigger efficiency, but also increase rate of false-positive triggers

Trigger Efficiency $\geq 99\%$ for interactions with $E_{dep} \geq 5 \text{ MeV}$ expected in 100% of a 10 kT Fiducial Volume

$E_{dep} = 10 \text{ MeV}$



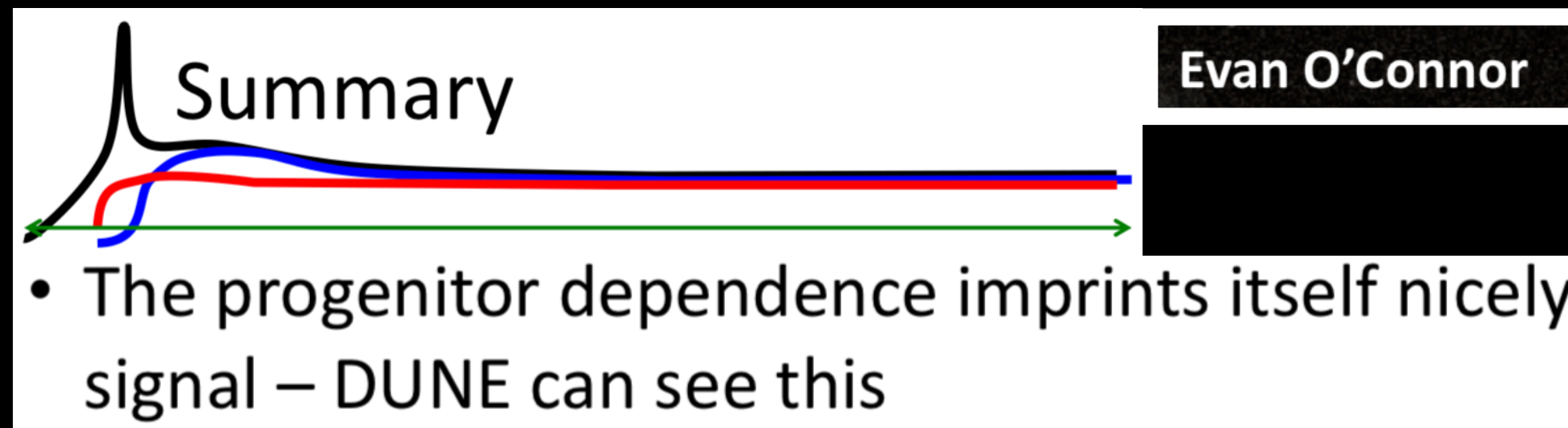
Enlarging the DUNE Physics Scope



Solar

John Beacom *Concluding Remarks*

- With new liquid-argon detectors, we can lead exciting new solar-neutrino studies, opening substantial discovery space in particle physics and astrophysics
- It is critical to DUNE's overall science program to succeed at measuring low energies well



- The progenitor dependence imprints itself nicely on the neutrino signal – DUNE can see this

- *It is is critical to lower Trigger E-threshold to extend range of SN detection (toward and beyond Galaxy edge).*
- *It is critical to guarantee good Time resolution and improve Energy resolution for SN-signatures in time & energy spectra*

SuperNova

How to **supply bias voltage** to the photo-sensors (in the range of 50 V or less) on the HV surfaces
and to **read-out the signal** out of HV surface

POF Technology for VD application

Two Parts

W. Pellico - FNAL

Warm

Cold

• **(1) Power to fiber**

- Convert electrical power to light
 - Four Laser modules to generate 48 V
 - Each are **4 watt** laser systems
 - Individual adjustable output power
 - Interlocked – to protect laser/personnel

- Transmit via fiber



Cold

• Fiber optic **Receivers**

- Four receivers tied in series \Rightarrow 48 volt for SiPM and power for LEDs for calibration
- Typical conversion efficiency 22 %
- 14 W dissipation (heat)

- SiPMs cold electronics module

- Gang some number of SiPMs
- Passive or/and Active (w/ preAmp&Shaping)

• **(2) Signal to fiber**

- Convert electrical to light
- Eleds – analog light **Transmitters**

- Transmit via fiber



Warm

- SiPMs warm electronics module

- Fiber to copper
- Signal conditioning
- Signal processing