



An Experimental Program in Neutrinos, Nucleon Decay  
and Astroparticle Physics Enabled by the Fermilab Long-  
Baseline Neutrino Facility

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Colorado State University  
For the DUNE Collaboration



**FNAL User's Meeting**

June 15 - 16, 2016

Fermi National Accelerator Laboratory

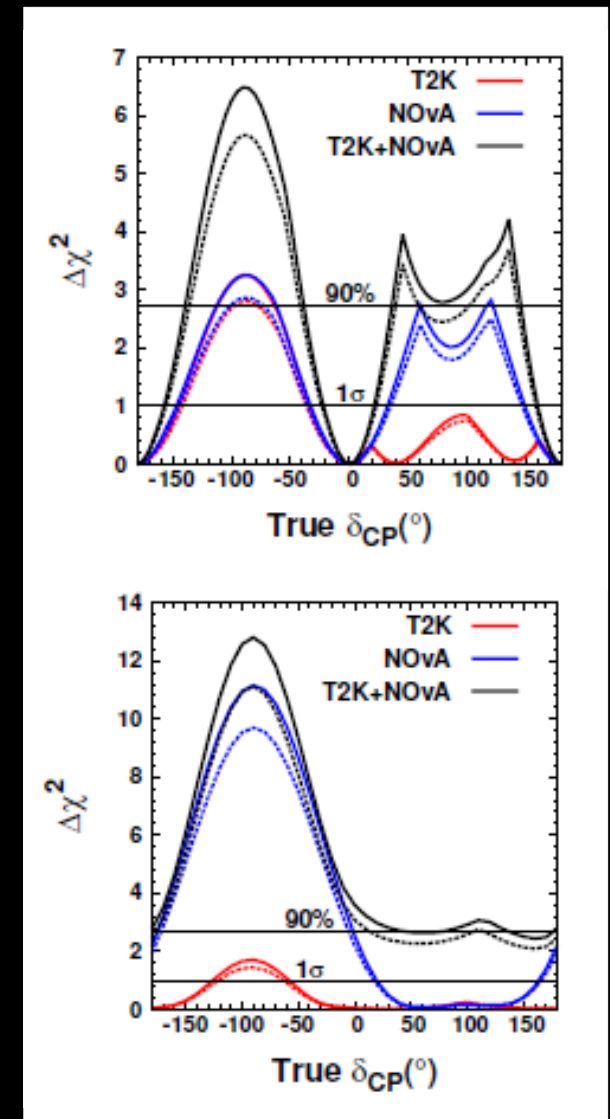
# The Deep Underground Neutrino Experiment



- September 2015 collaboration meeting at FNAL
- 886 Collaborators → 26+ countries
- 153 institutions → Members from LBNE, LBNO and more

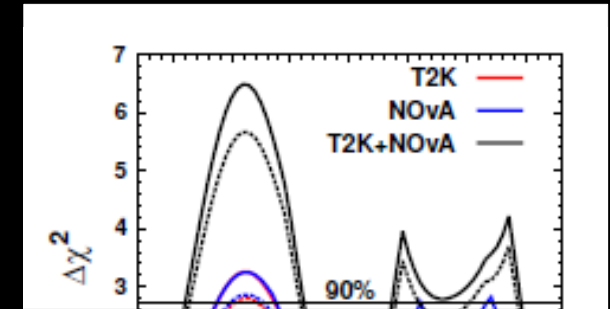
# Potential of Current Experiments

- T2K and NOvA will continue to run over next several years
  - measure  $\nu_e$  appearance and  $\nu_\mu$  disappearance
  - Run in both  $\nu$  mode and  $\bar{\nu}$  mode
  - Provide sensitivity to CPV and MH determination
  - A combined analysis has “indication” potential
- Reactor experiments
  - Continue to constrain  $\theta_{13}$  from  $\bar{\nu}_e$  disappearance
  - Constraints help T2K and NOvA
- MH determination may come from several sources like INO, PINGU, JUNO, and  $0\nu\beta\beta$
- SK will continue to asymptotically approach limits on nucleon decay, and atmospheric neutrino measurements



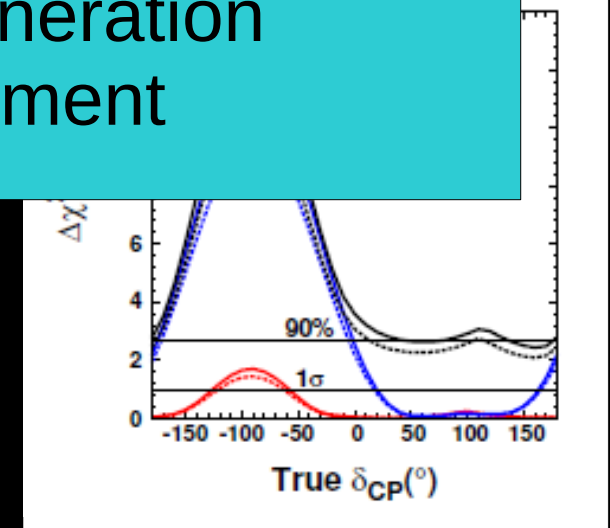
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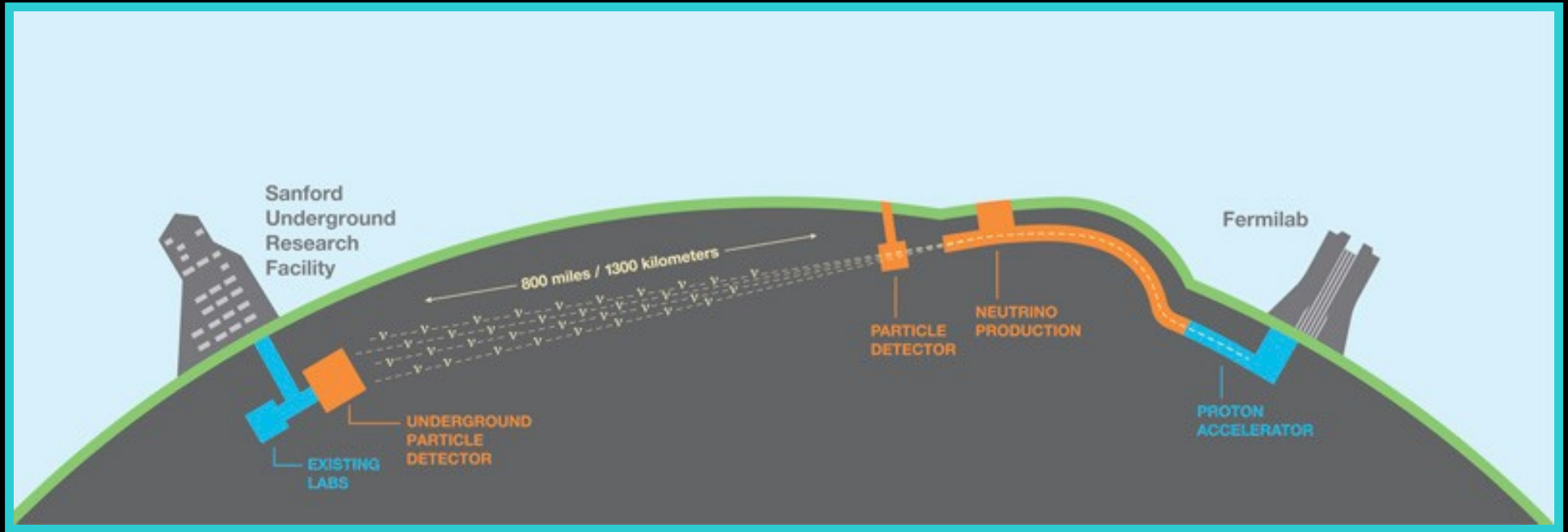


To measure  $\delta_{cp}$  and determine the MH to high precision in a single experiment will require a next generation long-baseline neutrino experiment

- MH determination may come from several sources like INO, PINGU, JUNO, and  $0\nu\beta\beta$
- SK will continue to asymptotically approach limits on nucleon decay, and atmospheric neutrino measurements

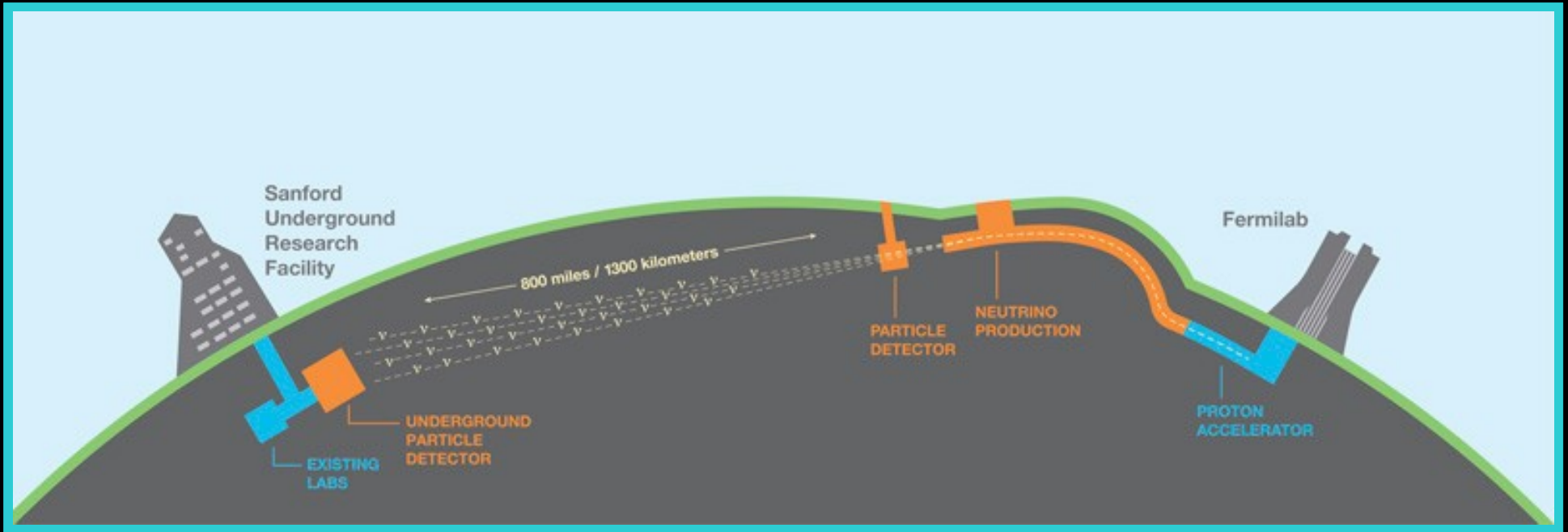


# The DUNE Experimental Setup



- DUNE is designed to provide a broad program of:
  - $\nu$  oscillation physics
  - $\nu$  interaction physics
  - Proton decay
  - Supernova physics
  - BSM physics

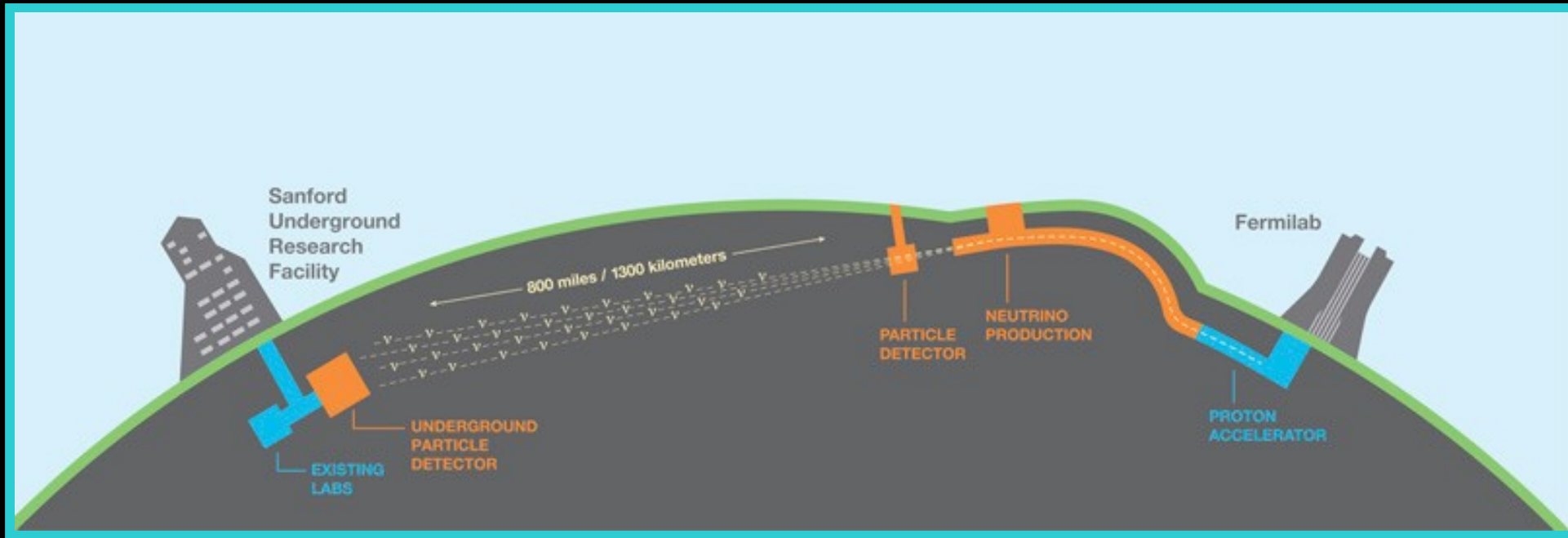
# The DUNE Experimental Setup



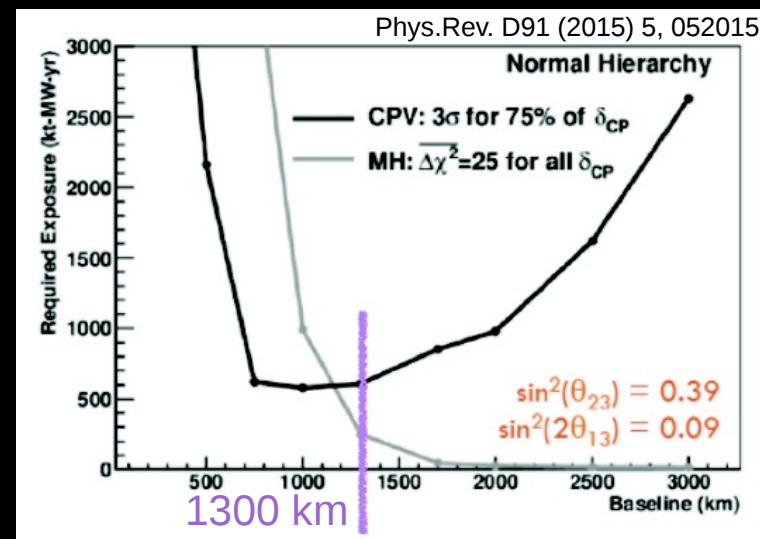
- Oscillation Physics:
  - Baseline of 1300 km
  - A megawatt class beam covering the 1<sup>st</sup> and 2<sup>nd</sup> oscillation maxima
  - A highly capable ND to constrain the FD event rate prediction



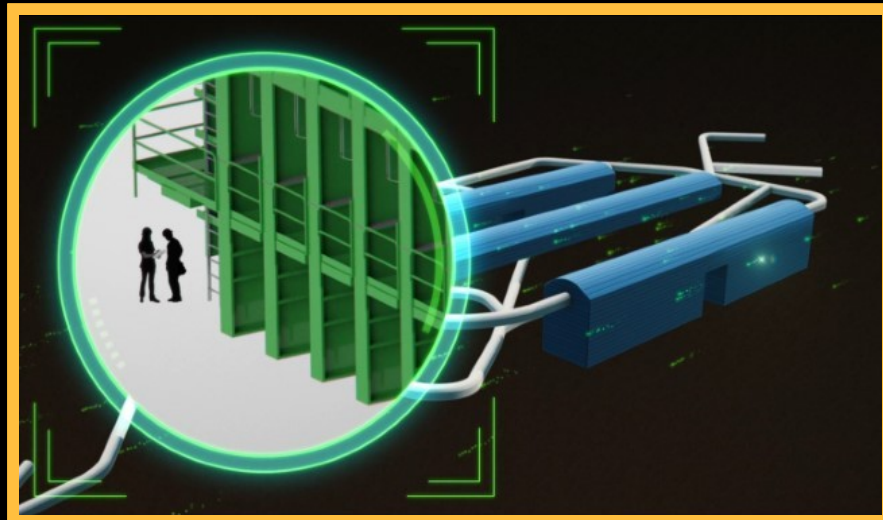
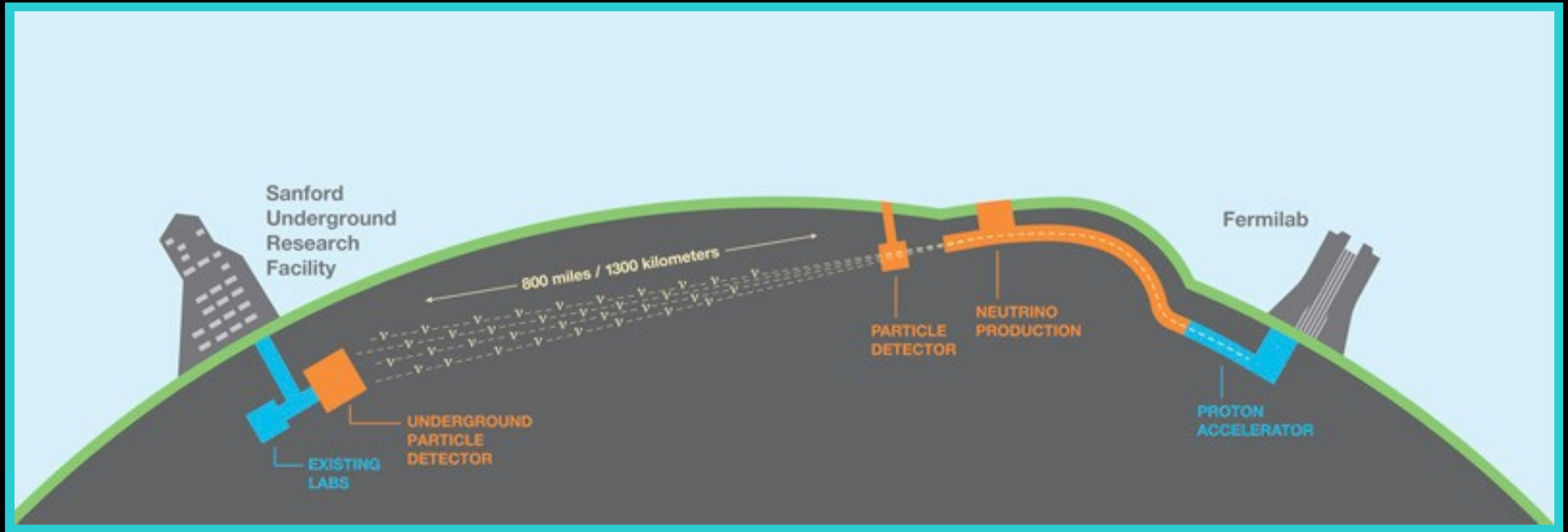
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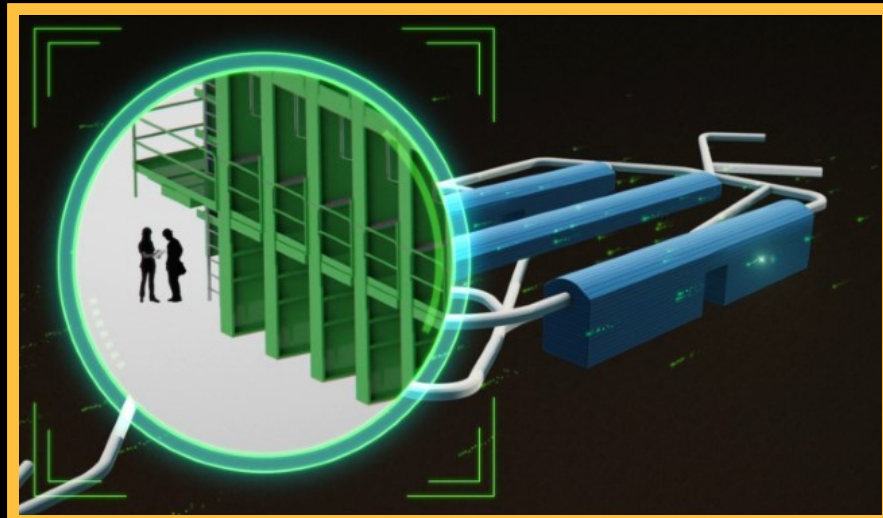
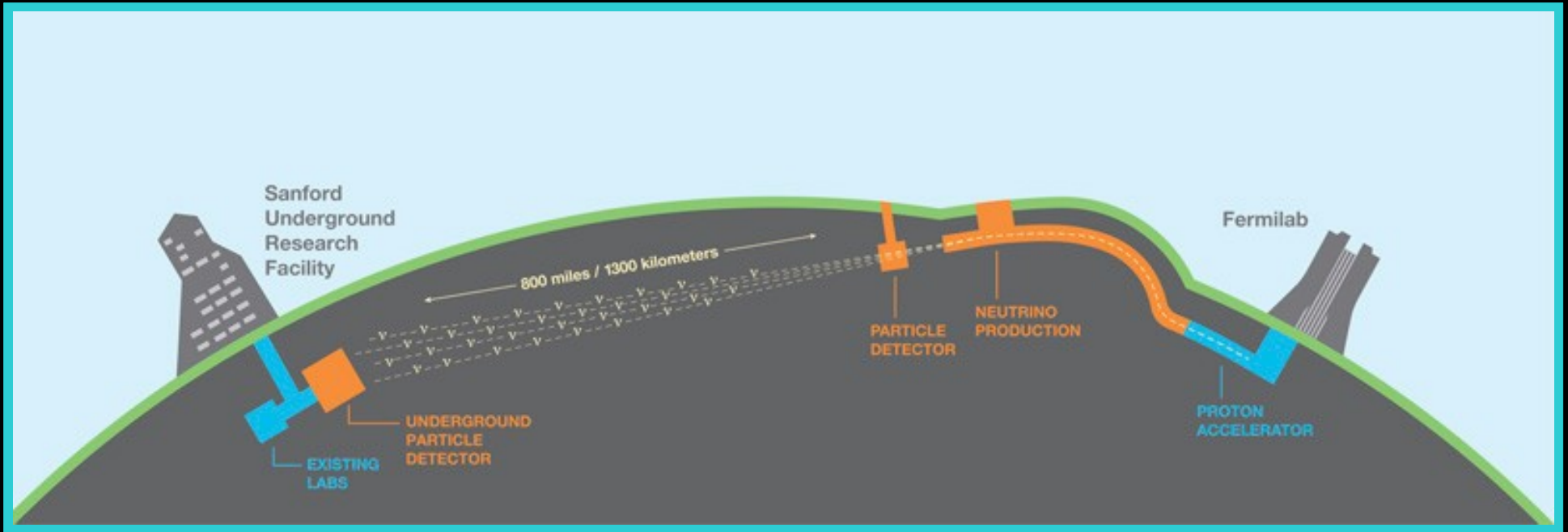
# The DUNE Experimental Setup



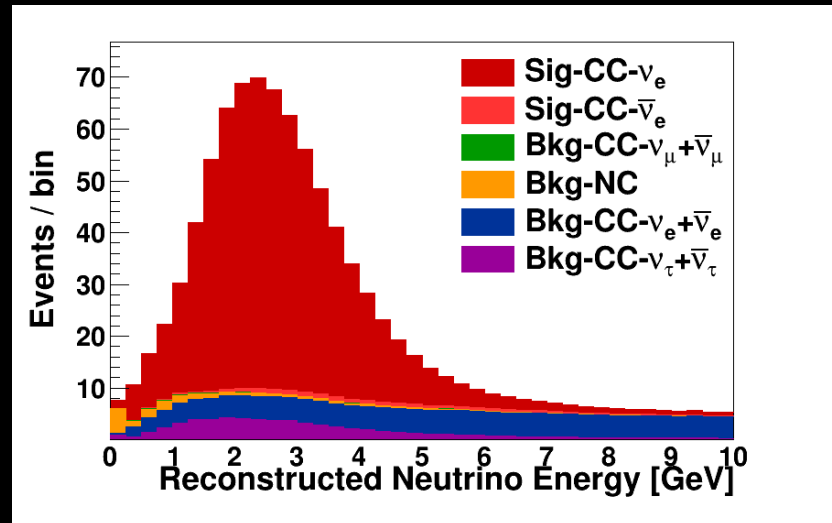
- Oscillation Physics:
  - Baseline of 1300 km
  - A large (~ 40 kt), high resolution FD deployed deep underground
  - Exposure of 6-12 yr with ~ 50% / 50%  $\nu$  /  $\bar{\nu}$  running



# The DUNE Experimental Setup



DUNE  $\nu_e$  appearance



# The DUNE Experimental Setup

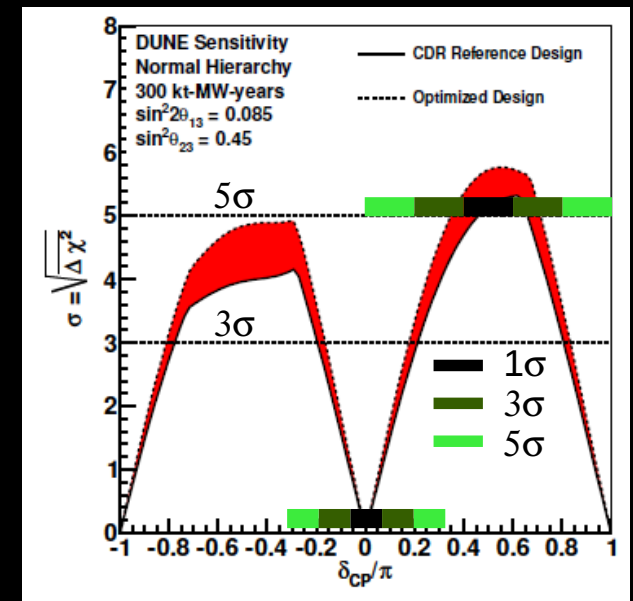
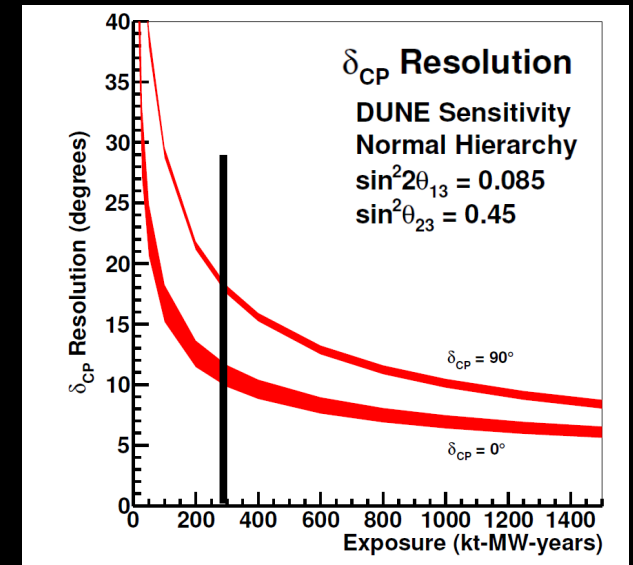
	CDR Reference Design	Optimized Design
$\nu$ mode (150 kt · MW · year)		
$\nu_e$ Signal NH (IH)	861 (495)	945 (521)
$\bar{\nu}_e$ Signal NH (IH)	13 (26)	10 (22)
Total Signal NH (IH)	874 (521)	955 (543)
Beam $\nu_e + \bar{\nu}_e$ CC Bkgd	159	204
NC Bkgd	22	17
$\nu_\tau + \bar{\nu}_\tau$ CC Bkgd	42	19
$\nu_\mu + \bar{\nu}_\mu$ CC Bkgd	3	3
Total Bkgd	226	243
$\bar{\nu}$ mode (150 kt · MW · year)		
$\nu_e$ Signal NH (IH)	61 (37)	47 (28)
$\bar{\nu}_e$ Signal NH (IH)	167 (378)	168 (436)
Total Signal NH (IH)	228 (415)	215 (464)
Beam $\nu_e + \bar{\nu}_e$ CC Bkgd	89	105
NC Bkgd	12	9
$\nu_\tau + \bar{\nu}_\tau$ CC Bkgd	23	11
$\nu_\mu + \bar{\nu}_\mu$ CC Bkgd	2	2
Total Bkgd	126	127

Number of events in the  $0.5 < E_\nu < 8.0 \text{ GeV}$  range, assuming 150 kt-MW-yr in each of the  $\nu$  and  $\bar{\nu}$  beam modes,  $\delta_{cp} = 0.0$ , and the NuFit 2014 oscillation parameters.

# The Physics of DUNE:

## Long-Baseline Physics: $\delta_{cp}$ and CPV

- DUNE measurement of  $\delta_{cp}$ 
  - Resolution on  $\delta_{cp}$  gets better as  $\sin(\delta_{cp}) \rightarrow 0$
  - Range on  $\delta_{cp}$  resolution from  $6^\circ$ - $10^\circ$  (~10 yr exposure)
- Sensitivity to CPV strongly depends on:
  - Statistics (thus the beam intensity, detector mass, run time)
  - The true value of  $\sin^2\theta_{23}$ ,  $\delta_{cp}$ , and the MH
  - Resolution on  $\delta_{cp}$  near  $\sin(\delta_{cp}) = 0$
  - Ability to constrain systematic uncertainties

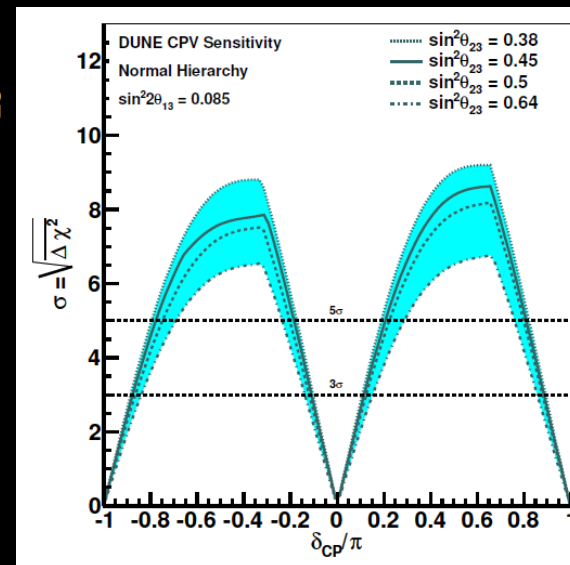


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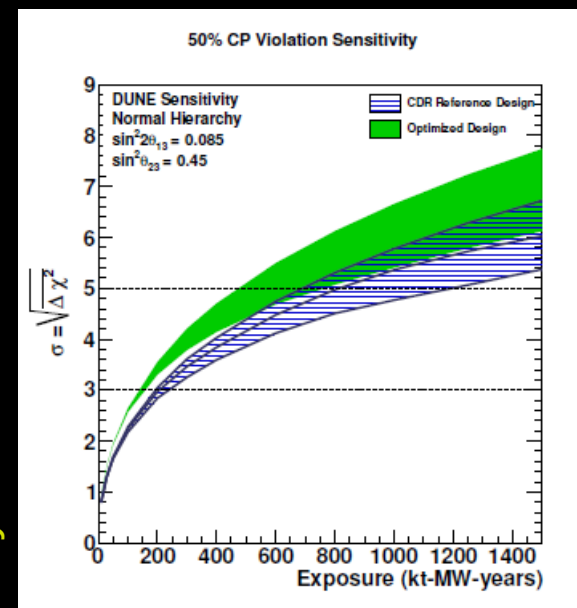
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  - **Ability to constrain systematic uncertainties**

Sensitivity variation bands for  $\sin^2\theta_{23}$



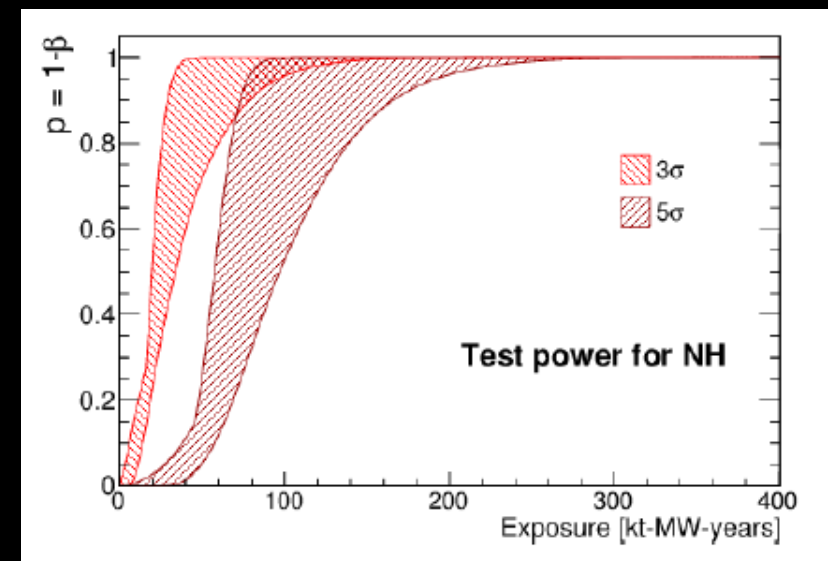
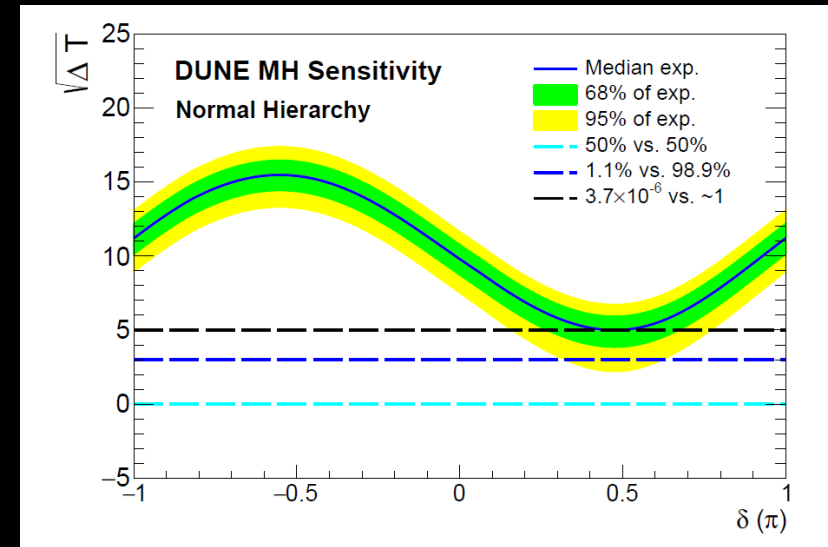
Sensitivity changes with systematic uncertainties



# The Physics of DUNE:

## Long-Baseline Physics: MH and the Rest

- DUNE will exclude the wrong MH at the 99% C.L. for all values of  $\delta_{cp}$
- The 99% C.L. result will come sooner for more favorable  $\delta_{cp}$  values
- DUNE will also constrain  $\sin^2(\theta_{13})$ ,  $\sin^2(\theta_{23})$ , and  $\Delta M^2_{31}$
- And has the potential to determine the  $\theta_{23}$  octant, and measure  $\nu_\tau$  appearance
- DUNE long-baseline physics goals also include:
  - Over-constrain the PMNS matrix
  - Search for exotic physics like NSI, LRI, CPT/Lorentz violation, compact extra dimensions, and sterile neutrinos



# The Physics of DUNE:

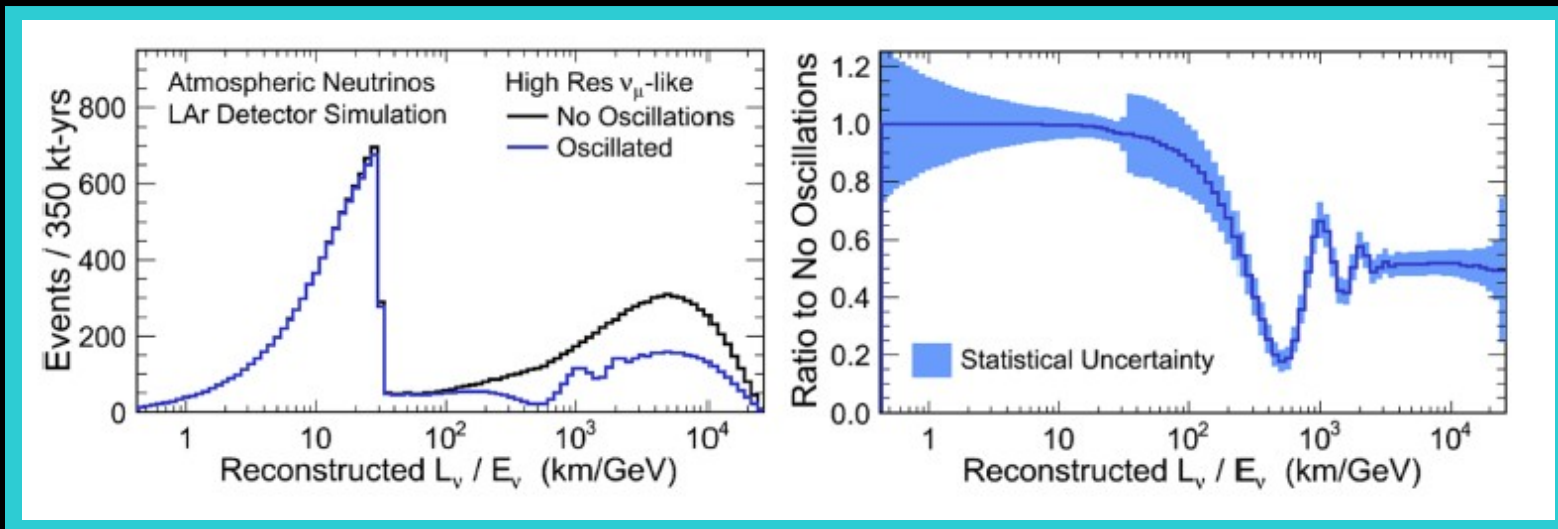
## Underground Physics: Proton Decay

- Signature of Baryon number violation
- Superior detection efficiency for K production modes
  - K PID through dE/dx
  - High spatial resolution and low energy thresholds → rejection atmospheric backgrounds
  - High Efficiency (>90%), high purity selections for  $p \rightarrow \nu + K^+$  and  $p \rightarrow \mu + K^0$
- Requires suitable triggering systems
- Efficiencies and background rates per Mt-yr:

Decay Mode	Water Cherenkov		Liquid Argon TPC	
	Efficiency	Background	Efficiency	Background
$p \rightarrow K^+ \bar{\nu}$	19%	4	97%	1
$p \rightarrow K^0 \mu^+$	10%	8	47%	< 2
$p \rightarrow K^+ \mu^- \pi^+$			97%	1
$n \rightarrow K^+ e^-$	10%	3	96%	< 2
$n \rightarrow e^+ \pi^-$	19%	2	44%	0.8

# The Physics of DUNE: Underground Physics: Atmospheric $\nu$

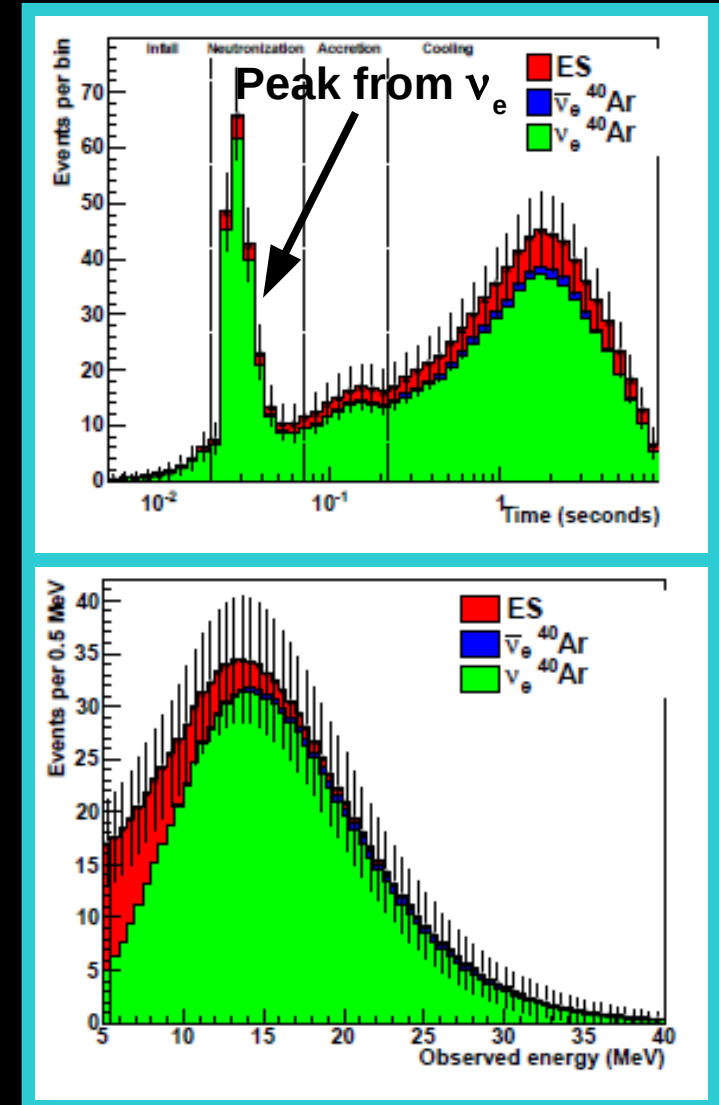
- Low energy thresholds gives superior L/E resolution
  - Fully reconstruct hadronic system
  - Low missing  $p_T$  improves angular resolution
- Good sensitivity to MH and  $\theta_{23}$  octant
- Combine with accelerator  $\nu$  data to improve oscillation physics measurements
- Sensitive to PMNS extensions / new physics
- Expect  $\sim 14k$  contained  $\nu_e$ -like events, and  $\sim 20k$  contained  $\nu_\mu$ -like events for a 350kt-yr exposure



# The Physics of DUNE:

## Underground Physics: Supernova Bursts

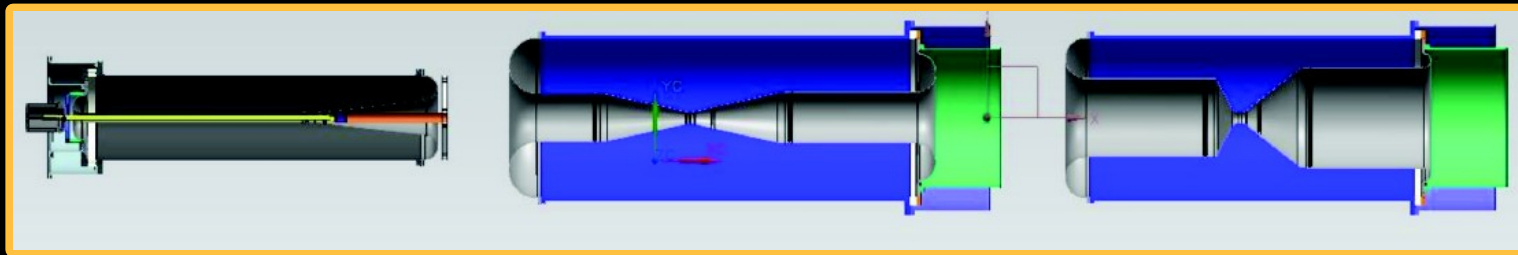
- Requires suitable triggering systems
- Other experiments rely on  $\bar{\nu}_e$  capture via inverse  $\beta$  decay
- DUNE will be able to observe the  $\nu_e$  flux through capture on Ar40
  - Unique sensitivity to the electron flavor component of the flux
  - Provides information on time, energy and flavor structure
  - Rates depend on core collapse model,  $\nu$  oscillation models, and distance.
  - Expect >3,000 events from a supernova at 10 kpc



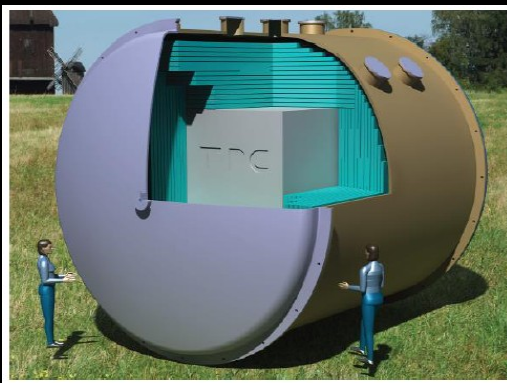


# DUNE Task Forces

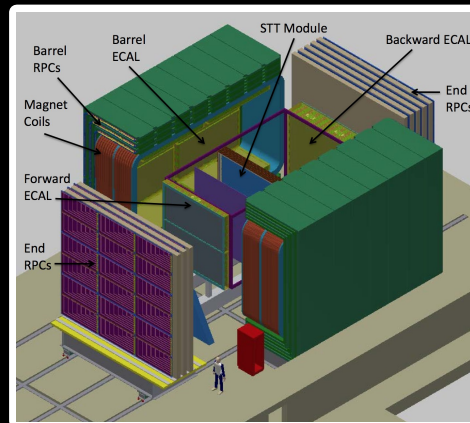
- Cross-working-group teams charged with simulating, evaluating, and optimizing the performance of the three main components of the experimental design
- **Beam Optimization**



- **Near Detector Optimization**

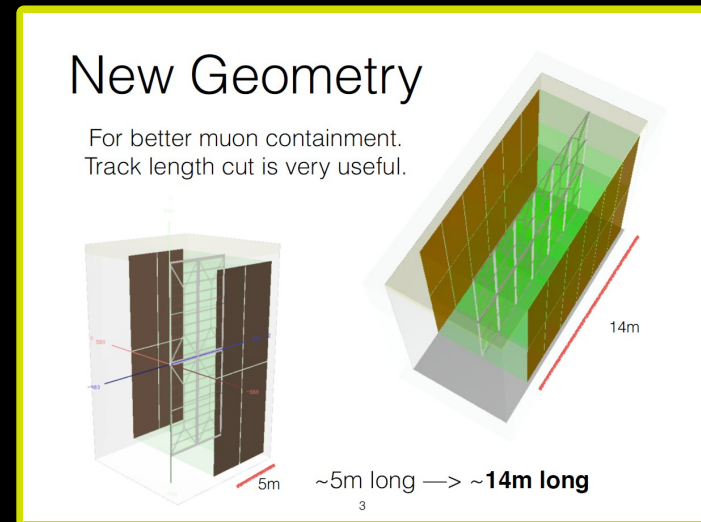


High-Pressure  
GAr TPC



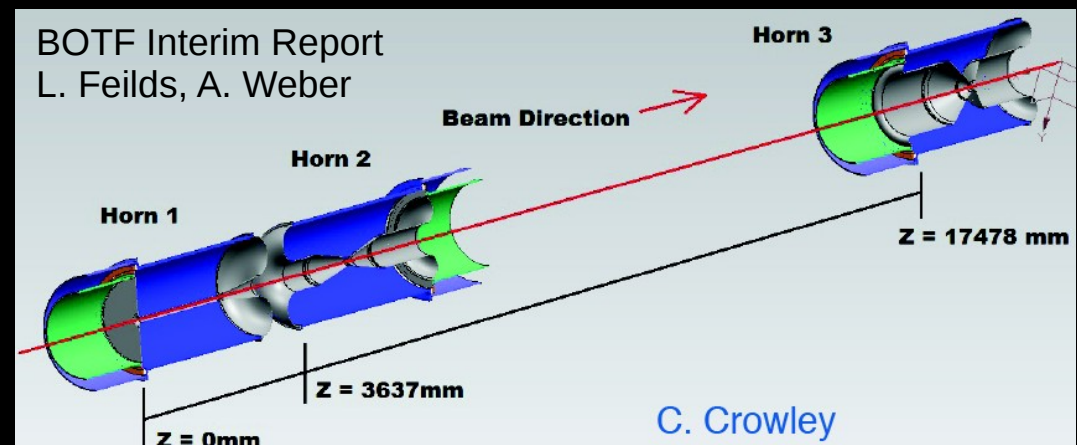
Fine-Grained  
Tracker

- **Far Detector Optimization**



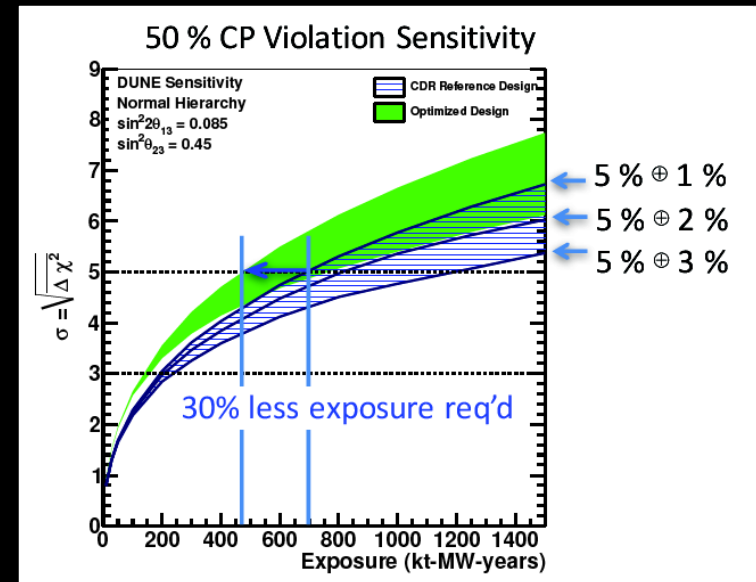
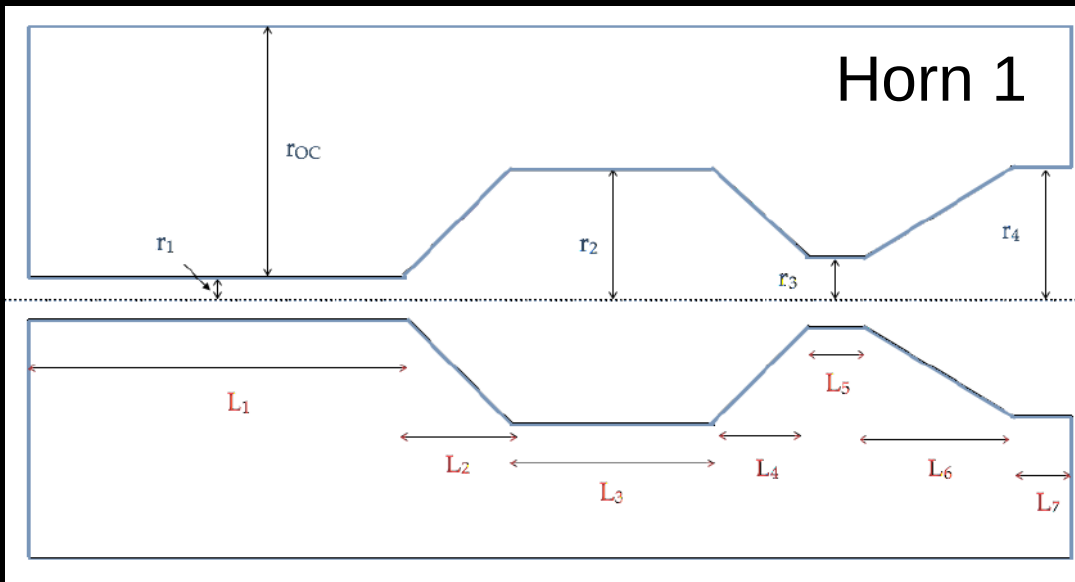
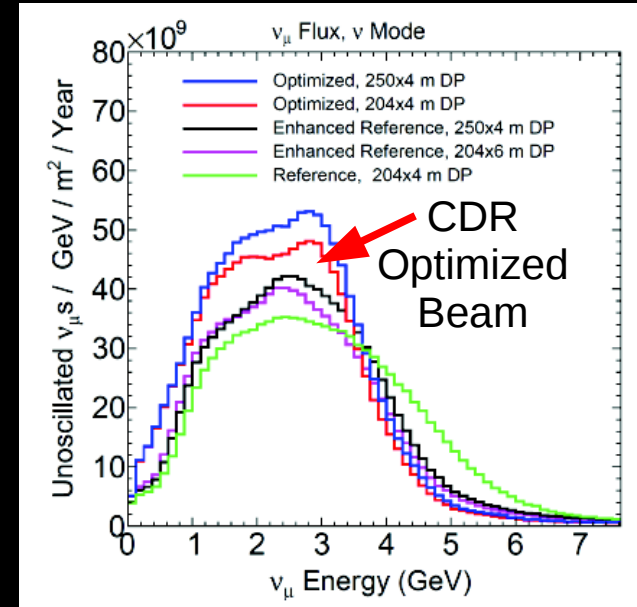
# Beam Optimization Task Force

- Charge:
  - Physics driven optimization of the beam line (target, horns, etc)
  - Study alternate designs and develop a cost benefit analysis
- Status:
  - Design has been optimized for multiple component sets (2 vs. 3 horns, multiple target designs, etc)
  - Realistic design based optimizations in advanced stages
  - Detailed studies of the design are in progress:
    - Physics sensitivities
    - Optimal run plan ( $v/\bar{v}$ )
    - Cost implications
    - Alternate metrics
    - Alternate optimization routines



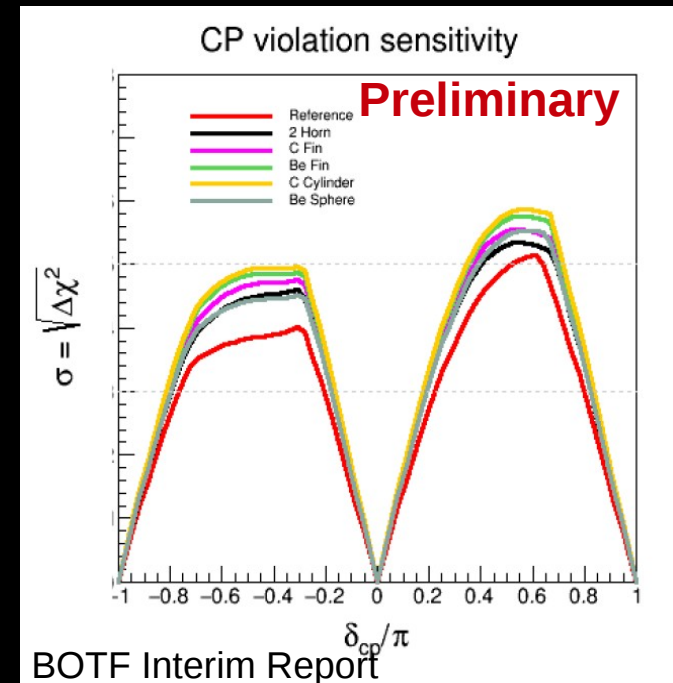
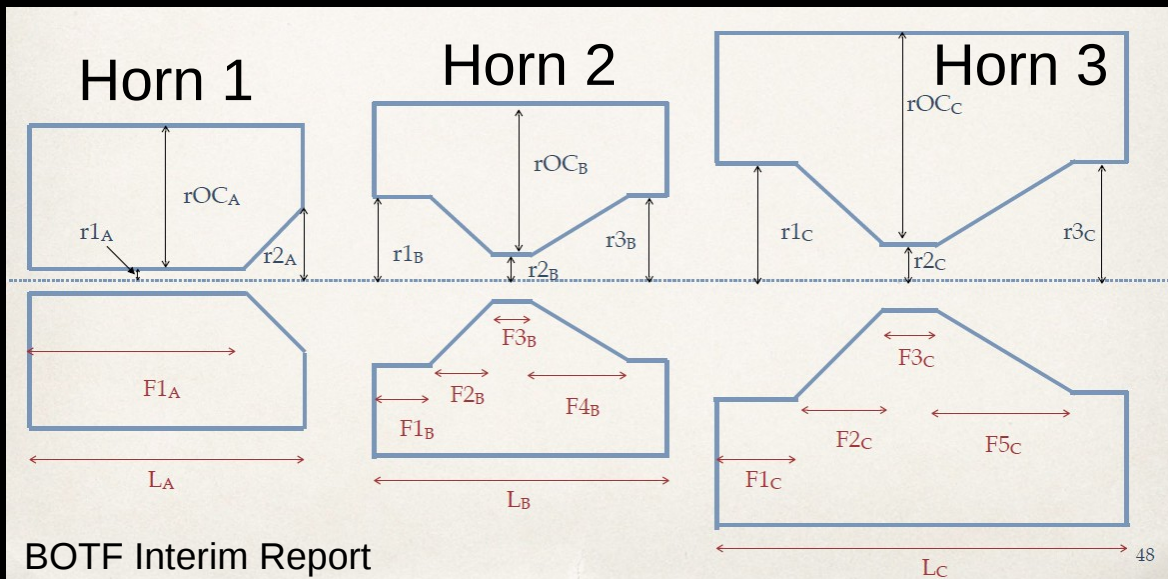
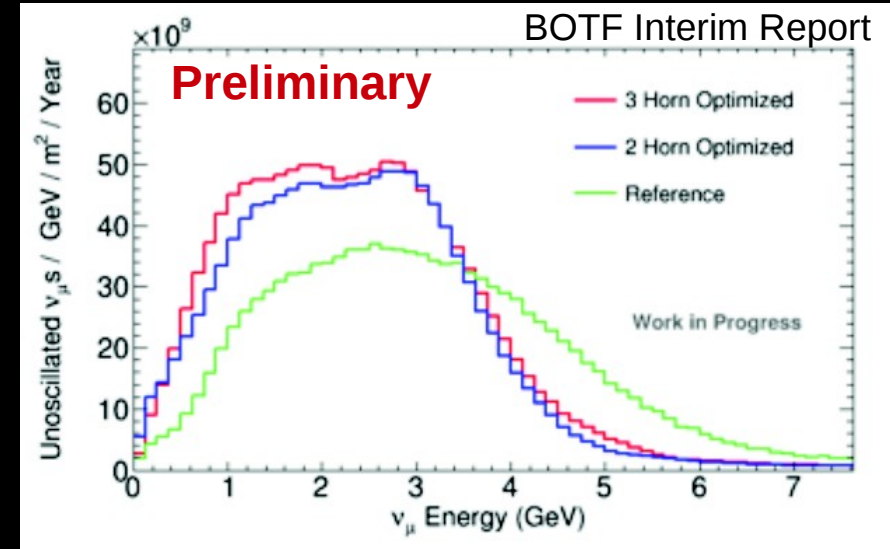
# Beam line Genetic Optimization

- Optimizations studies conducted for the DUNE CDR
- Genetic optimization of of:
  - Target and horn dimensions
  - Proton momentum
  - Decay pipe length
- Metric based on CPV sensitivity



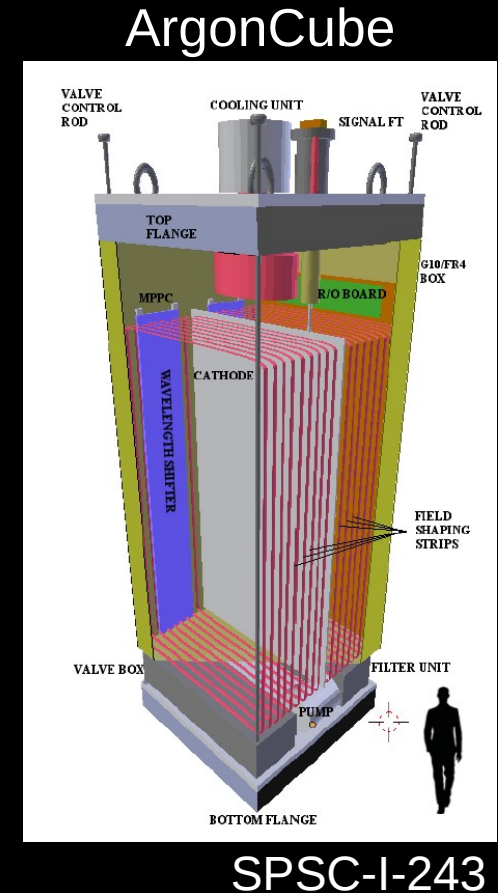
# Beam line Genetic Optimization

- Task force is building on the success of the CDR studies
- Optimization of 2 vs 3 horn design
- Studies of several target designs
- Shifted focus to engineering feasibility and design flexibility



# Near Detector Task Force

- Charge:
  - Develop full GEANT4 simulation of 3 technology options
    - Fine-Grained Tracker (FGT)
    - Modular Liquid Argon TPC (LAr TPC / ArgonCube)
    - High-Pressure Gaseous Argon TPC (HP GAR TPC)
  - Develop end-to-end simulation and analysis chain to evaluate the impact of each ND on CPV sensitivity
- Status:
  - Each step in the simulation and analysis chain, and interfaces between each step, have been developed
  - Full GEANT4 simulations have been completed
  - The VALOR framework is used for ND fits and a DUNE specific oscillation analysis has been developed
  - Progress on event reconstruction is hard fought
  - Detector uncertainties represent the next (and last) big challenge



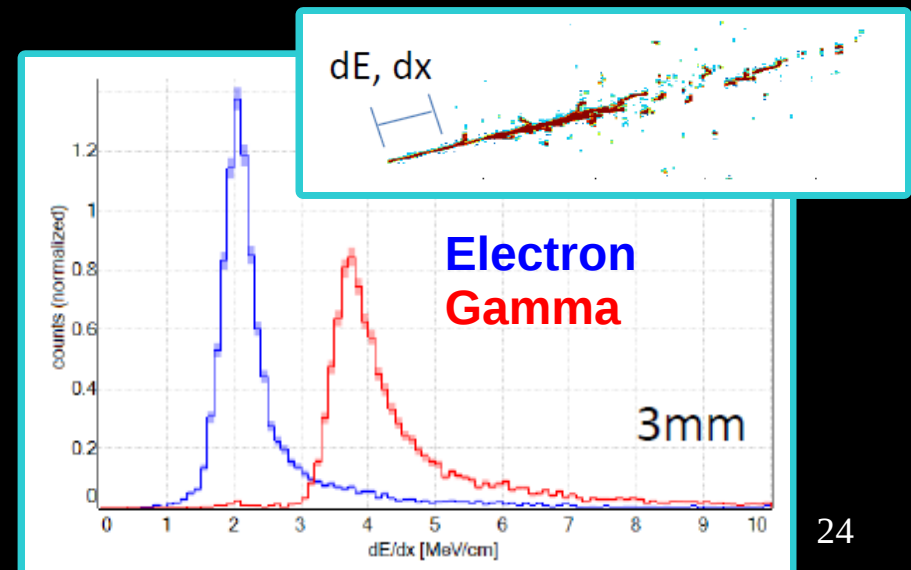
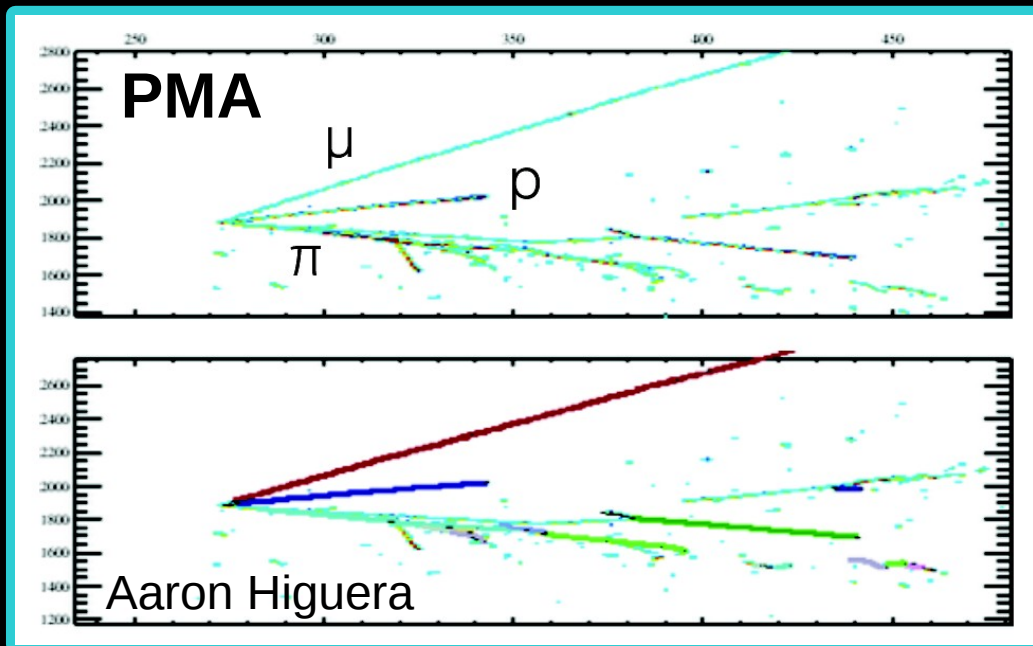


# Far Detector Optimization

- Charge:
  - Full GEANT4 simulation and reconstruction for reference and alternate designs
  - Optimization studies for FD components and configurations
  - Evaluate full range of FD physics topics
    - Oscillation: accelerator, atmospheric
    - Non-oscillation: proton decay, supernova bursts
- Status:
  - Detector simulation in advanced stages, including 2-phase
  - Recent non-accelerator event generation improvements
  - Reconstruction and PID algorithms in development
  - First round of optimization studies using full simulation tools underway
  - More progress on reconstruction required to draw conclusions

# LAr TPC Reconstruction

- Full simulation of beam  $\nu$ , atmospheric  $\nu$ , PDK, and Supernova events
- Huge progress has been made on reconstruction
  - Three reconstruction packages (PMA, Pandora, WireCell)
  - Exploring other options including machine learning techniques
  - Shower / track selection, particle ID, momentum and angle reconstruction
- Use of centralized software tools and infrastructure is crucial
  - LArSoft allows for easy collaboration with other LAr TPC experiments



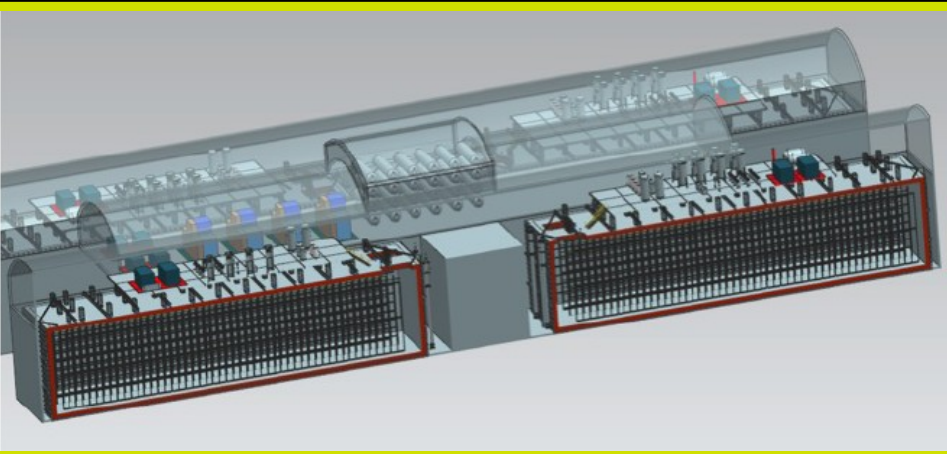
Dorota Stefan & Robert Sulej



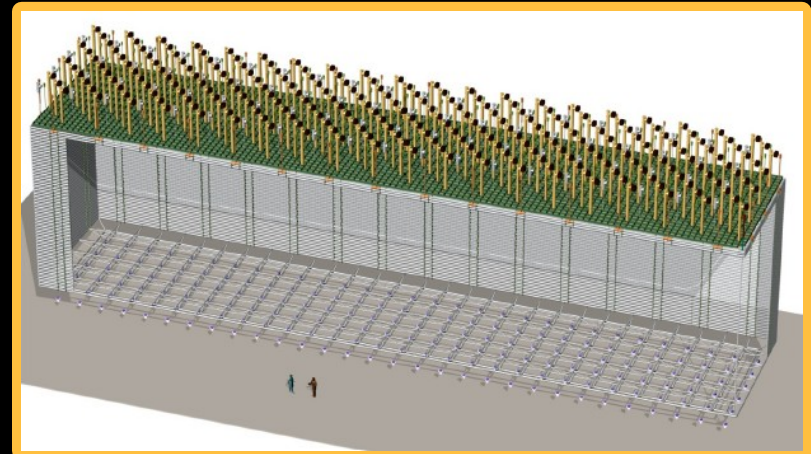
# Far Detector Options and R&D

- Two FD detector options:
- Single Phase
  - 35 ton (completed)
  - ProtoDune (2018)
  - Far Detector (1<sup>st</sup> module)
- Dual Phase
  - 311 (coming soon)
  - ProtoDune (2018)
  - Far Detector
- Important contribution from SBN Program detectors

Single phase, 2 modules



Dual phase, 1 module



# Conclusions

- LBNF will include:
  - A megawatt class  $\nu$  beam
  - Conventional facilities for near and far detectors
- The DUNE experiment will build 4 x 10 kt LAr TPCs and a highly capable ND at LBNF
- DUNE will determine the MH and measure  $\delta_{cp}$
- DUNE will provide a broad physics program including a wide variety of topics, including:
  - Conventional neutrino oscillations
  - Exotic neutrino oscillations
  - Neutrino interaction physics
  - Precision weak physics
  - Nucleon decay
  - Core collapse supernovae
  - Nuclear physics
  - Physics beyond the SM
- Optimization of the DUNE experimental design in progress<sup>26</sup>

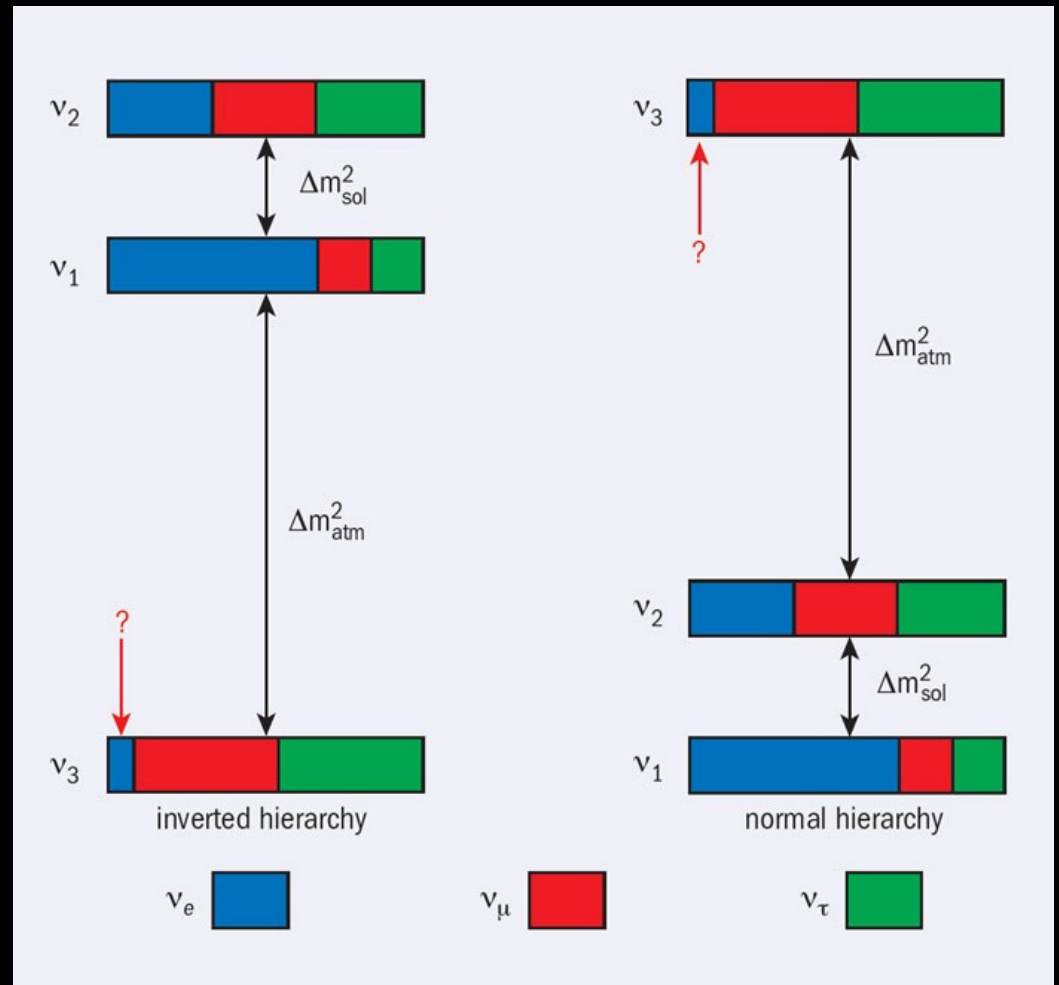
# Backup Slides

# Overview

- Physics potential of current  $\nu$  oscillation experiments
- The DUNE experimental setup
- The physics of DUNE
- The plan for DUNE infrastructure
- Inputs from the intermediate neutrino program
- Conclusions

# Unanswered Questions

- What are the  $\nu$  masses?
- Are  $\nu$  their own antiparticle?
- What is the  $\nu$  mass ordering?
- Is there CP violation (CPV) in the lepton sector, and what is the value of  $\delta_{cp}$ ?
- What is the  $\theta_{23}$  octant?
- Do protons decay?

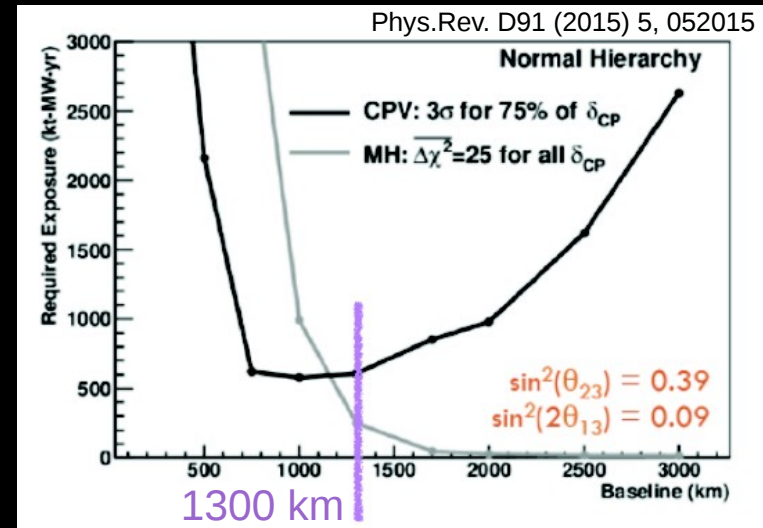


# DUNE and LBNF

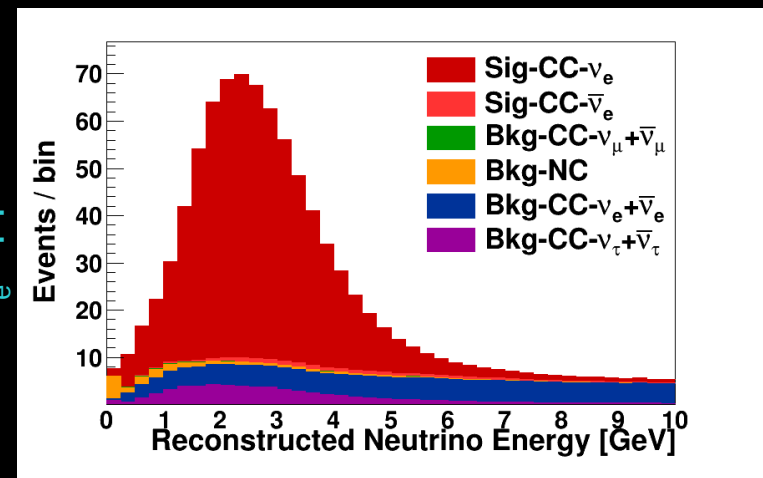
- Detectors and science collaboration will be managed separately from the neutrino facility and infrastructure
- Long-Baseline Neutrino Facility (LBNF)
  - Neutrino beam line
  - Near detector complex (but not the ND)
  - Far site (Sanford Lab) conventional facilities; detector hall, cryogenic systems
  - Operating costs for all of the above
- Deep-Underground Neutrino Experiment (DUNE)
  - Definition of **scientific goals** and **design requirements** for all facilities
  - **The Near and Far Detectors**
  - The scientific research program
- Close and continuous coordination between DUNE and LBNF will be required

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  - Sensitivity to  $\delta_{cp}$  and the MH in the same experiment

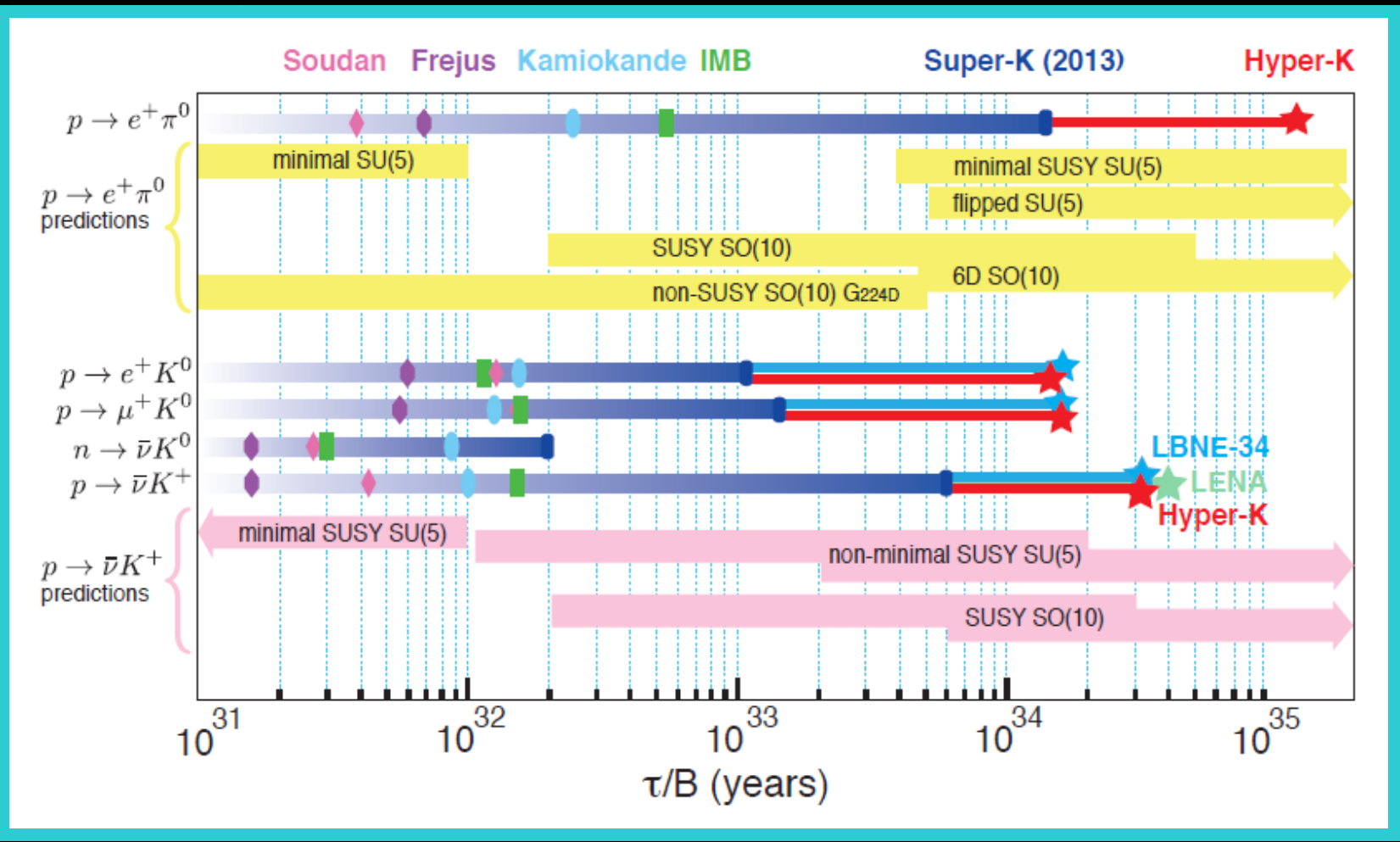


DUNE  $\nu_e$  appearance



# The Physics of DUNE: Underground Physics: Proton Decay

- S
- S
- F
- E



eric  
K<sup>0</sup>

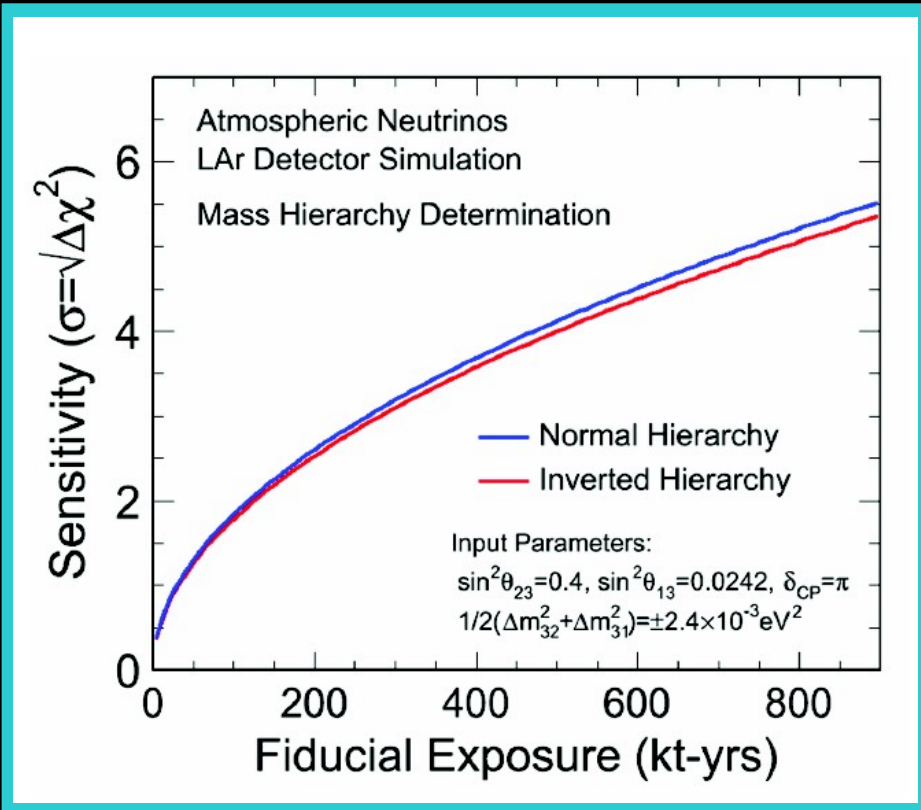
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# The Physics of DUNE: Underground Physics: Atmospheric $\nu$

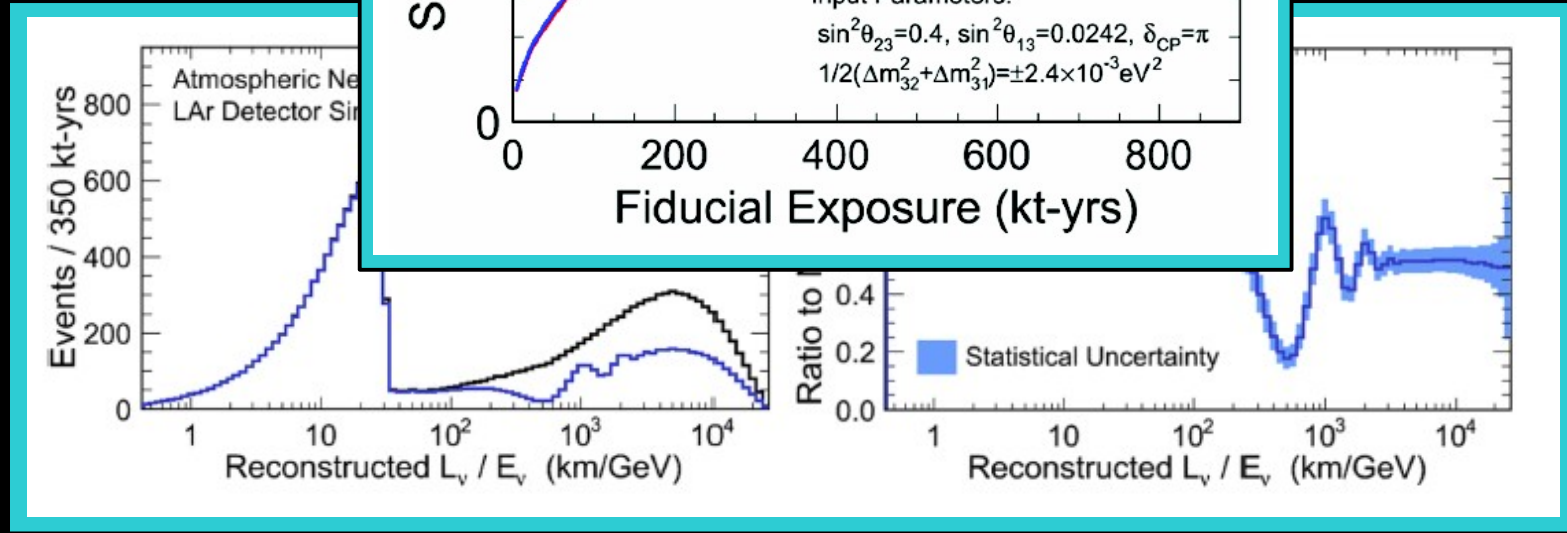
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  - Fully reconstruct
  - Low missing  $p_T$  in

- Good sensitivity to
- Combine with acc
- Sensitive to PMNS
- Expect ~14k cont
- 350kt-yr exposure



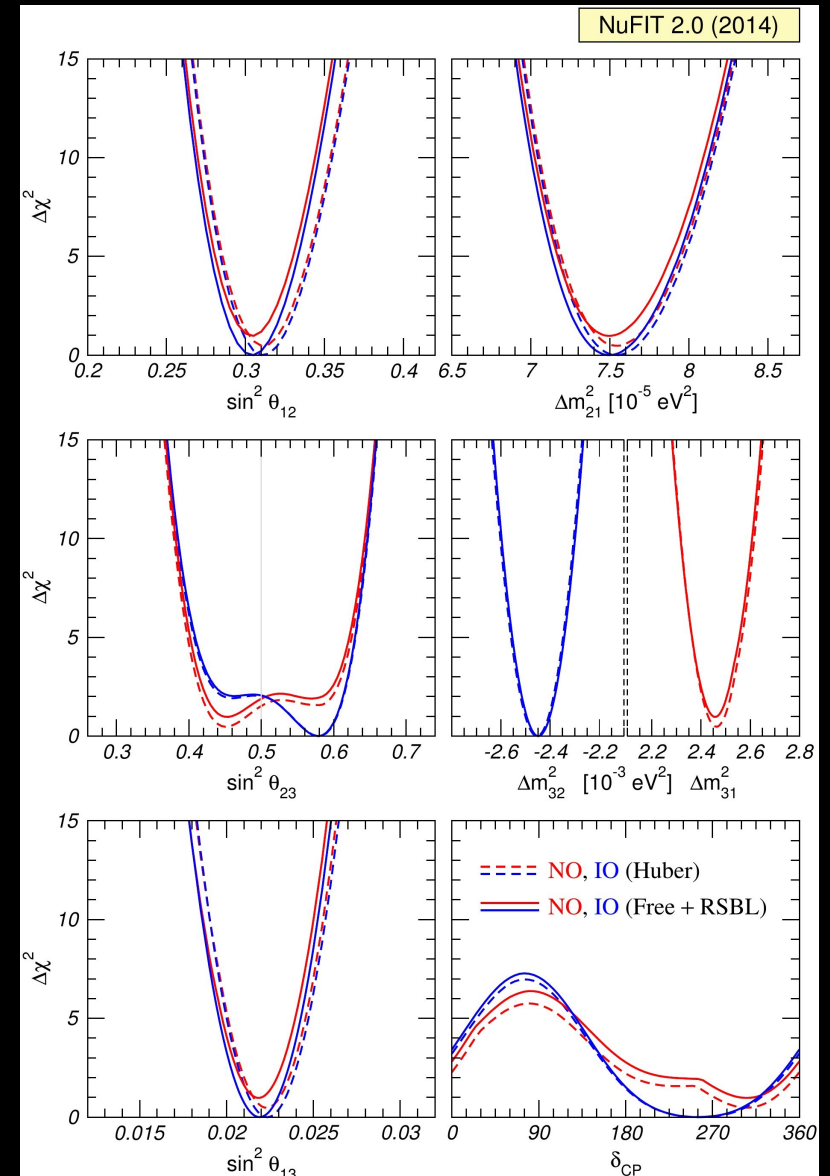
physics measurements

of  $\nu_\mu$ -like events for a



# The Current State of $\nu$ Oscillation Measurements

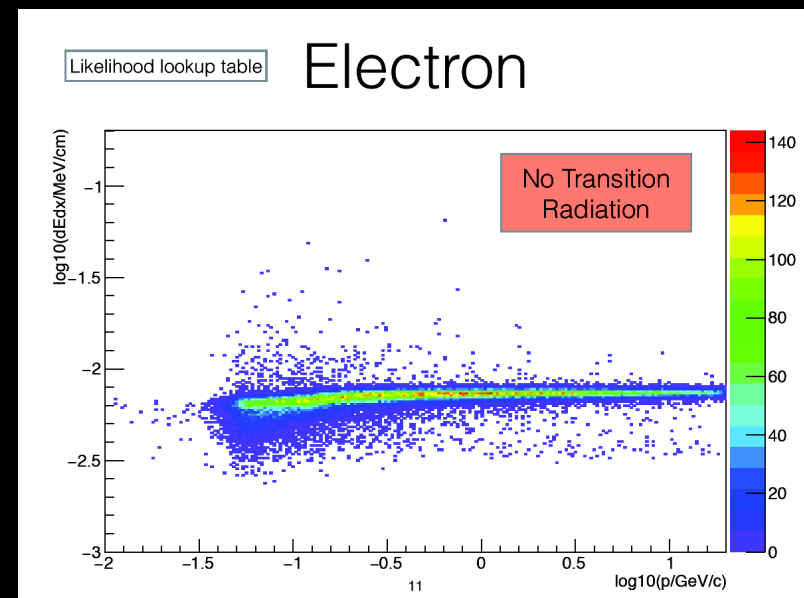
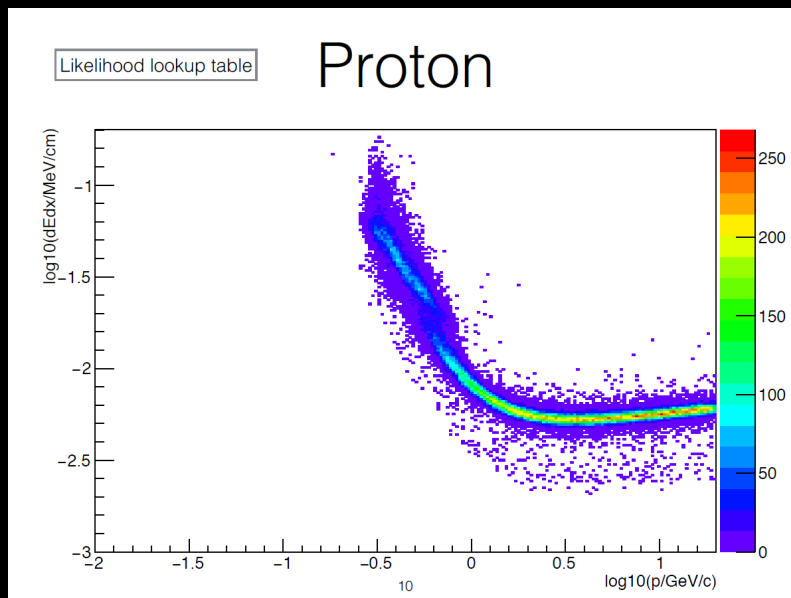
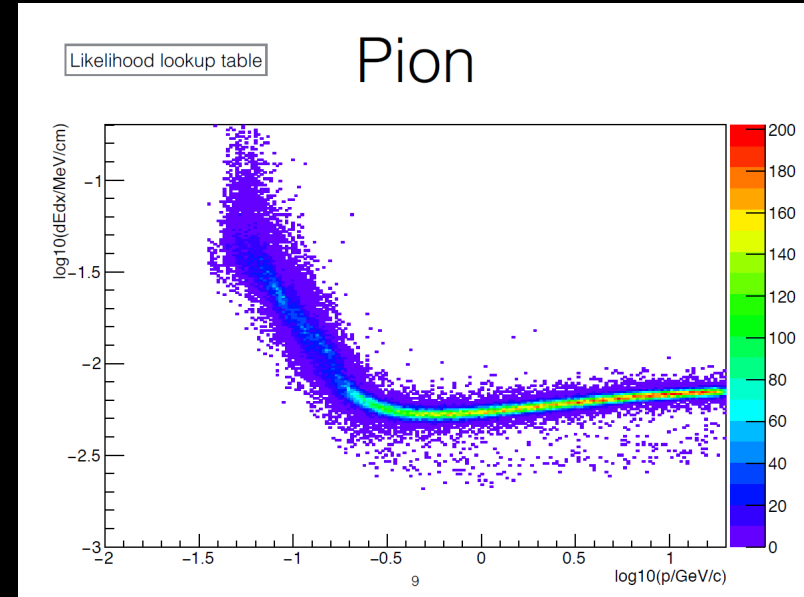
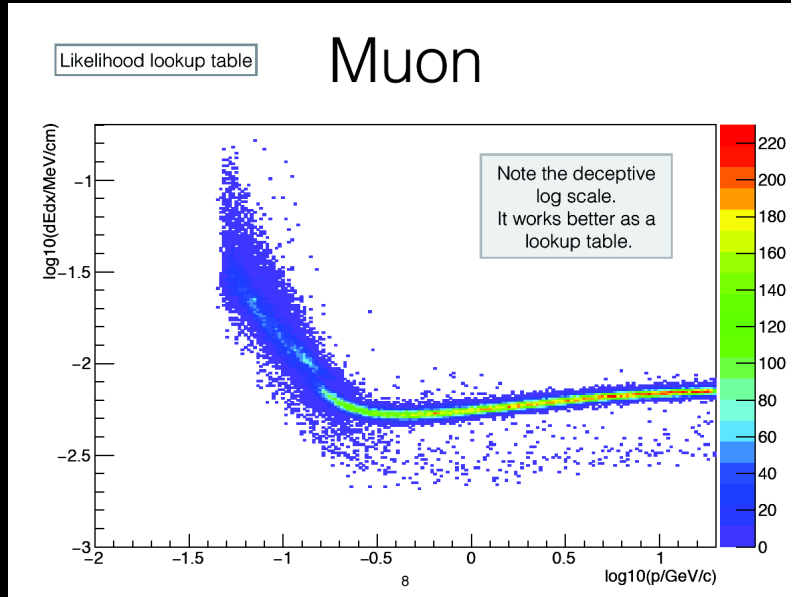
- PMNS matrix, factorized
- Numu  $\rightarrow$   $\nu_{\mu e}$  oscillation probability
- NuFit14 results



# The Physics of DUNE: Near Detector Physics

- The high resolution fine grained tracker (FGT) required for DUNE oscillation physics will allow for a multitude of  $\nu$  and other weak interaction physics measurements
- High statistics with excellent particle ID and reconstruction will allow for World leading measurements
- Full phase space differential measurements from  $4\pi$  coverage
- Precision cross section measurements of exclusive and inclusive channels, including many rare processes
- Variety of nuclear targets will help disentangle nuclear effects (both the nuclear initial state and final state interactions) from  $\nu$  interaction physics
- Precision electroweak and isospin measurements
- Exotic physics searches including heavy sterile neutrinos, light dark matter searches, and large  $\Delta m^2$  sterile  $\nu$  oscillations

# FGT dE/dx Profiles



# VALOR DUNE: Final state samples

## 2016a (2nd pass-through)

- $\nu_\mu$  CC
  1. 1-track QE enhanced ( $\mu^-$  only)
  2. 2-track QE enhanced ( $\mu^- + p$ )
  3.  $1\pi^\pm$  ( $\mu^- + 1\pi^\pm + X$ )
  4.  $1\pi^0$  ( $\mu^- + 1\pi^0 + X$ )
  5.  $1\pi^\pm + 1\pi^0$  ( $\mu^- + 1\pi^\pm + 1\pi^0 + X$ )
  6. Other
- Wrong-sign  $\nu_\mu$  CC
  7. Inclusive ( $\mu^+ + X$ )
- $\nu_e$  CC
  8. Inclusive ( $e^- + X$ )
- NC
  9. Inclusive

FHC  
+ RHC

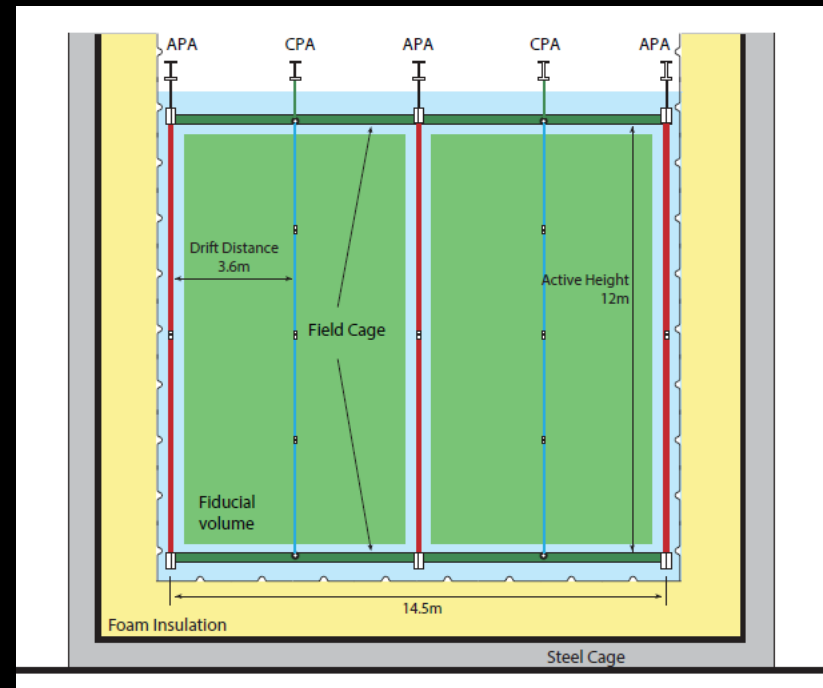
## 2016b (3rd pass-through)

- $\nu_\mu$  CC
  1. 1-track  $0\pi$  ( $\mu^-$  only)
  2. 2-track  $0\pi$  ( $\mu^- + \text{nucleon}$ )
  3. N-track  $0\pi$  ( $\mu^- + (>1) \text{ nucleons}$ )
  4. 3-track  $\Delta$ -enhanced ( $\mu^- + \pi^+ + p$ , with  $W_{reco} \approx 1.2 \text{ GeV}$ )
  5.  $1\pi^\pm$  ( $\mu^- + 1\pi^\pm + X$ )
  6.  $1\pi^0$  ( $\mu^- + 1\pi^0 + X$ )
  7.  $1\pi^\pm + 1\pi^0$  ( $\mu^- + 1\pi^\pm + 1\pi^0 + X$ )
  8. Other
- Wrong-sign  $\nu_\mu$  CC
  9.  $0\pi$  ( $\mu^+ + X$ )
  10.  $1\pi^\pm$  ( $\mu^+ + \pi^\pm + X$ )
  11.  $1\pi^0$  ( $\mu^+ + \pi^0 + X$ )
  12. Other
- $\nu_e$  CC
  13.  $0\pi$  ( $e^- + X$ )
  14.  $1\pi^\pm$  ( $e^- + \pi^\pm + X$ )
  15.  $1\pi^0$  ( $e^- + \pi^0 + X$ )
  16. Other
- NC
  17.  $0\pi$  (nucleon(s))
  18.  $1\pi^\pm$  ( $\pi^\pm + X$ )
  19.  $1\pi^0$  ( $\pi^0 + X$ )
  20. Other
- $\nu_e$ 
  21.  $\nu_e + e^-$  elastic
  22. Inverse muon decay  $\bar{\nu}_e + e^- \rightarrow \mu^- + \bar{\nu}_\mu$

FHC  
+ RHC

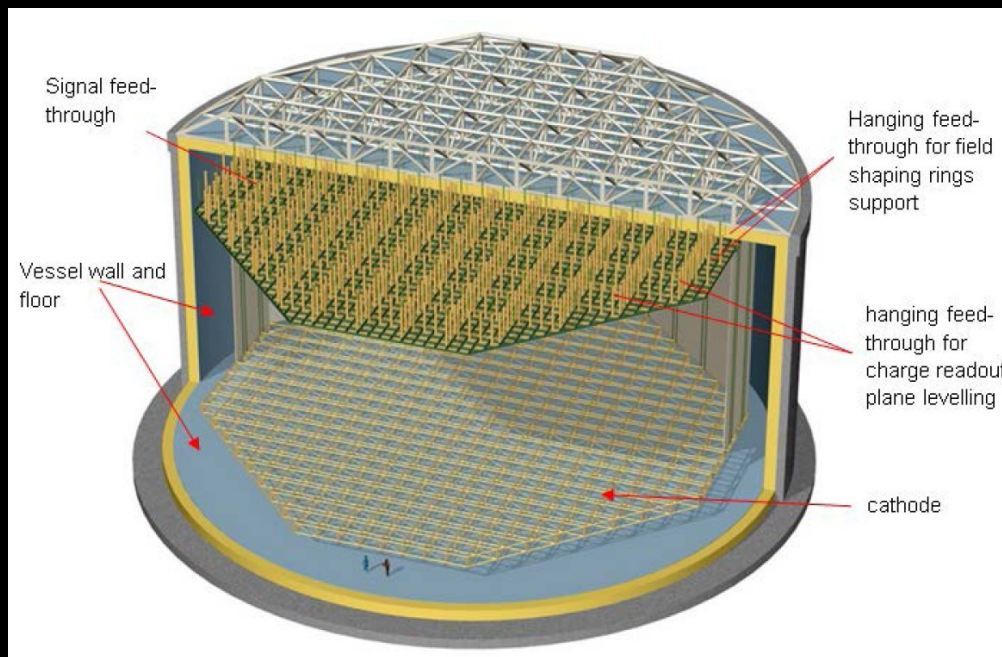
# Experimental Infrastructure: The DUNE Far Detector

- Heart of a deep underground neutrino and nucleon decay observatory
- Liquid Argon (LAr) Time Projection Chamber (TPC) with a 40 kt fiducial mass
- Staged construction with the goal of the first 10 kt by 2021/22
- Two potential designs:
  - Single phase
    - Current reference design
    - Based on ICARUS design
    - Horizontal drift ~3.6 m
    - Wire pitch of 5 mm
    - Detection and electronics in liquid
    - Modular approach
    - Well known cost and schedule



# Experimental Infrastructure: The DUNE Far Detector

- Heart of a deep underground neutrino and nucleon decay observatory
- Liquid Argon (LAr) Time Projection Chamber (TPC) with a 40 kt fiducial mass
- Staged construction with the goal of the first 10 kt by 2021/22
- Two potential designs:



## → Dual phase

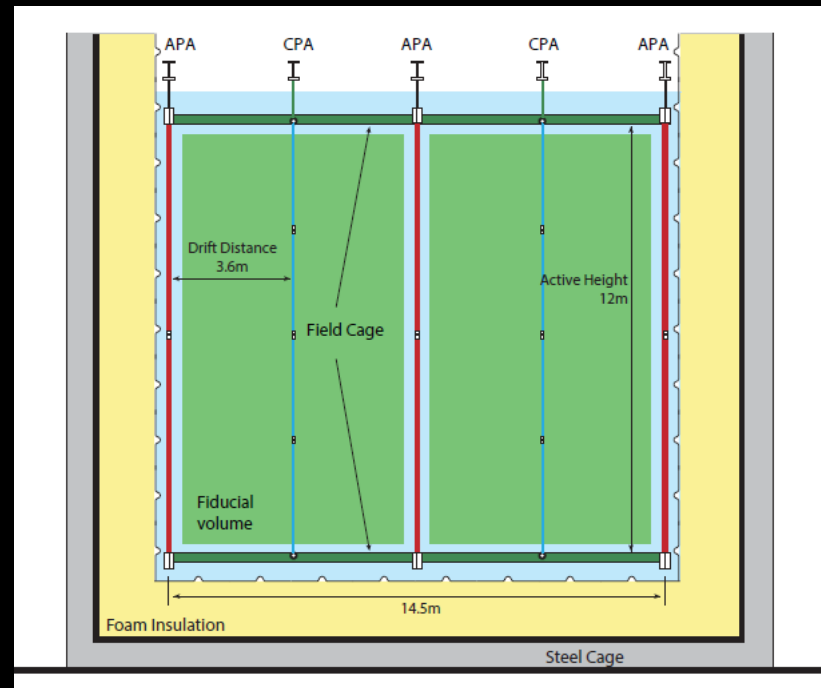
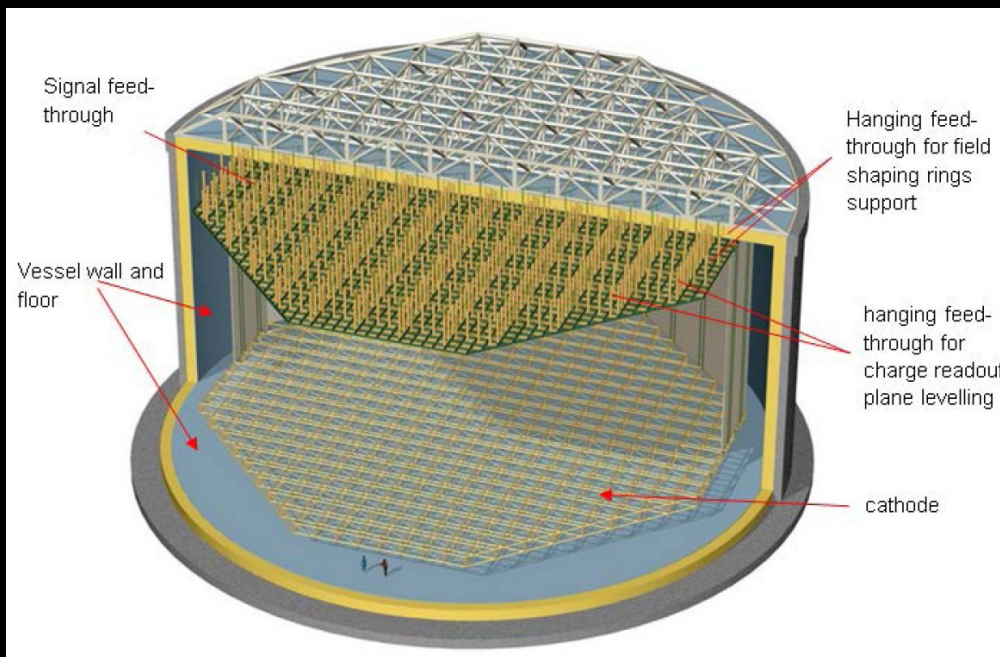
- Alternate design
- New technique; signal amplification
- Vertical drift ~10 - 20 m
- Detection and electronics in gas
- Adaptable to cryostat shape
- Low thresholds, high S/N ratio
- Pitch of 3 mm or less

# Experimental Infrastructure: The DUNE Far Detector

Heart of a deep underground neutrino and nuclear decay

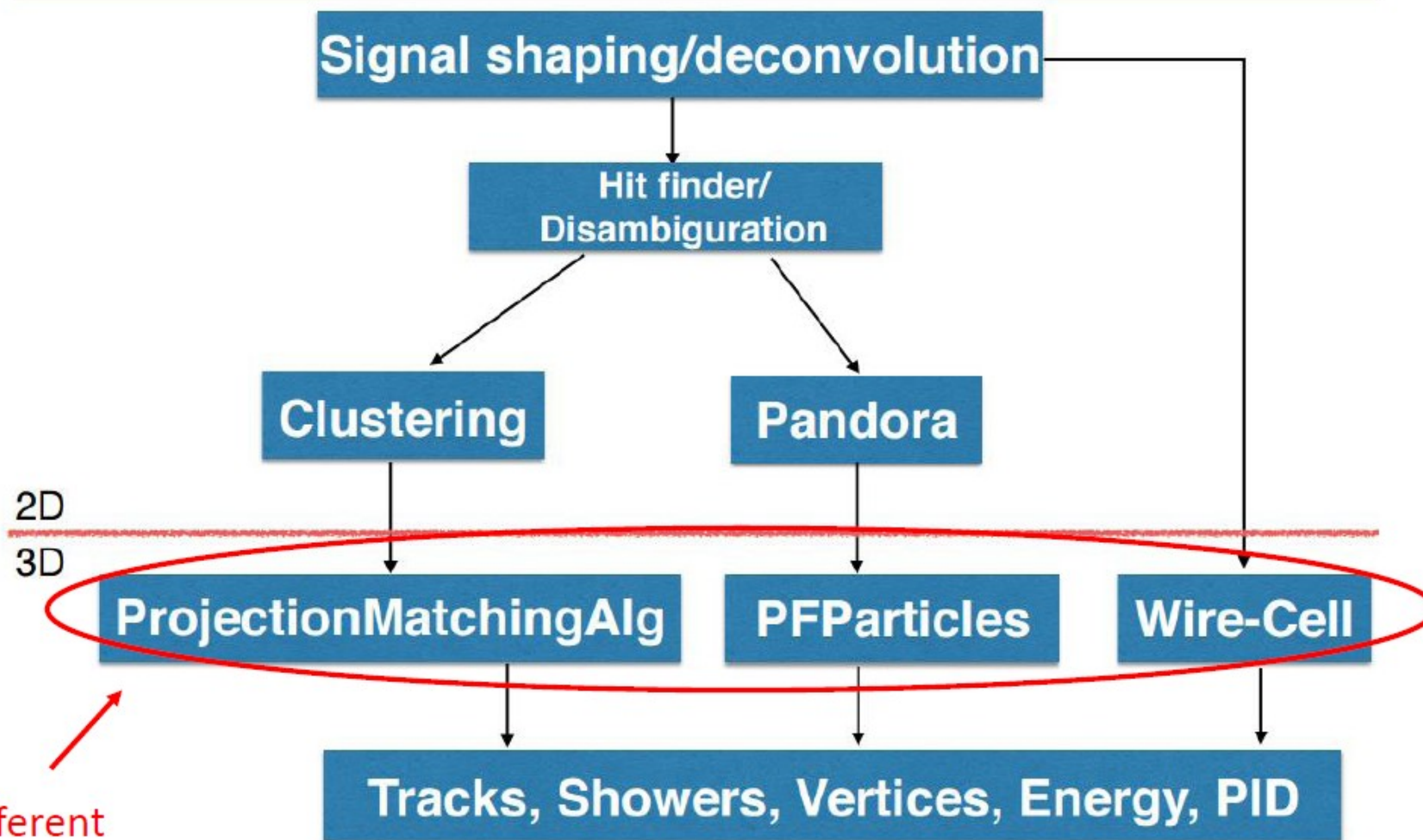
The CERN Neutrino Platform is working to build  $\sim 6 \text{ m}^3$  prototype detectors for both designs, and deploy them in CERN a charged particle test beam

Two potential designs:





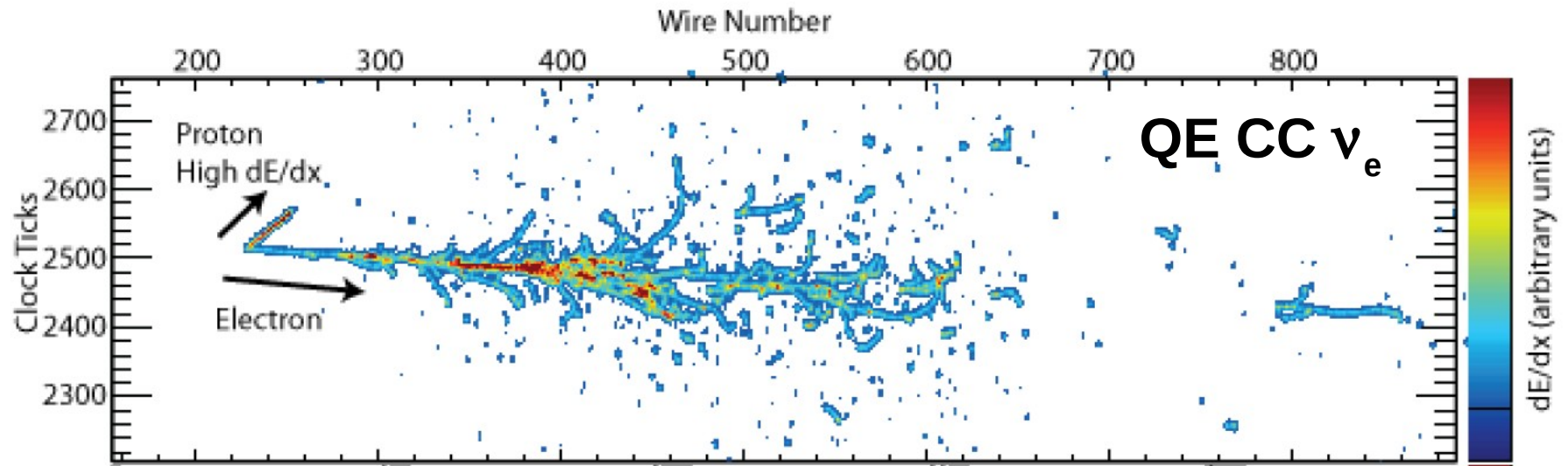
# Reminder: Reconstruction Chain



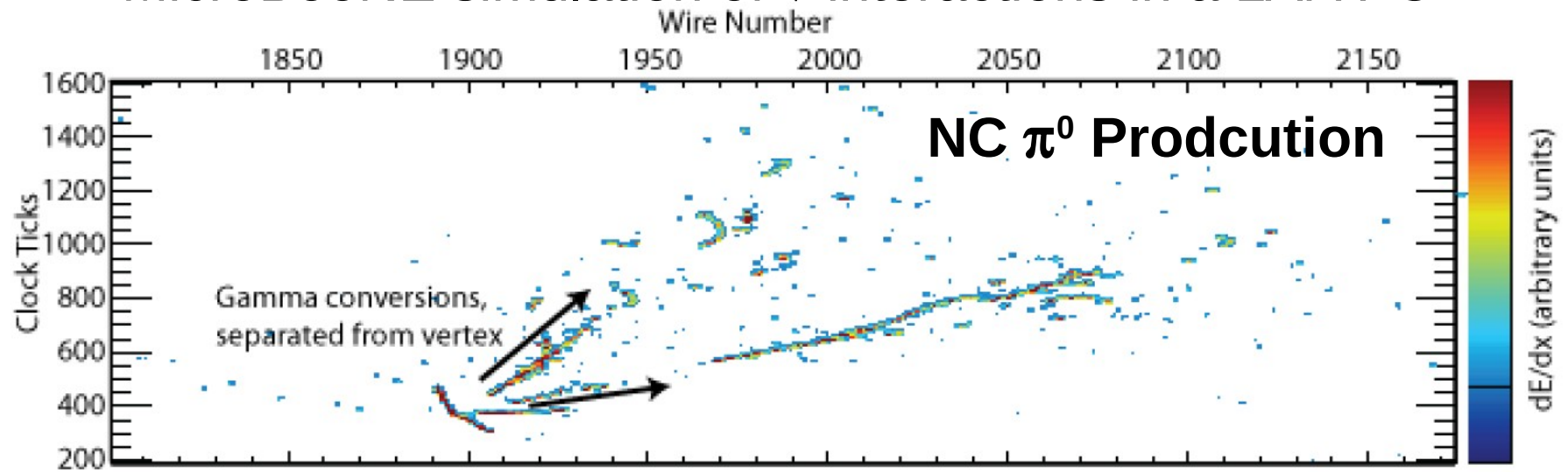
3 different approaches

Tingjun Yang

# Experimental Infrastructure:

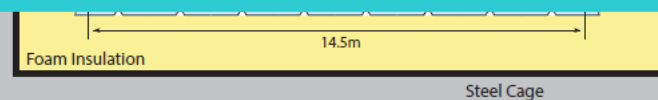


MicroBooNE simulation of  $\nu$  interactions in a LArTPC



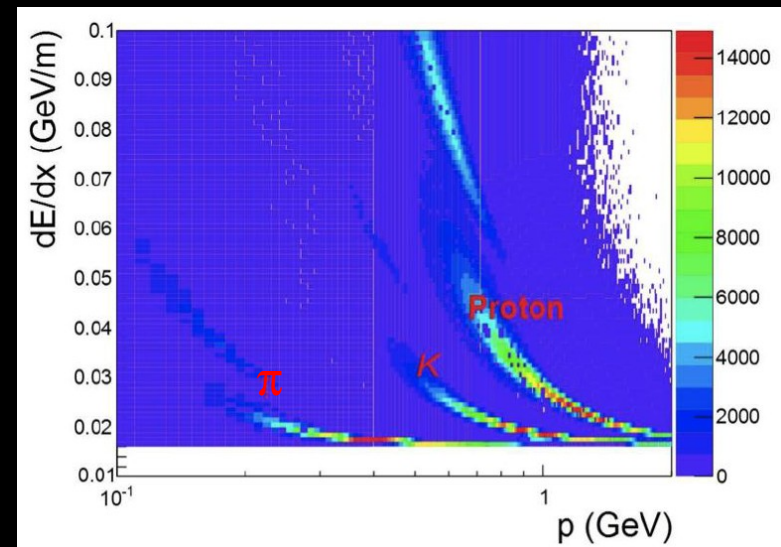
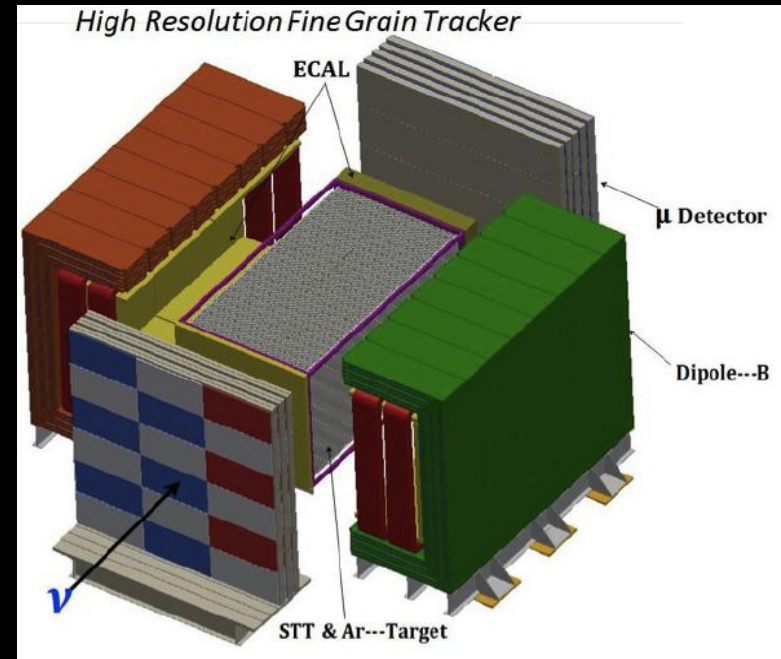
Sig  
thro

Vesse  
floor



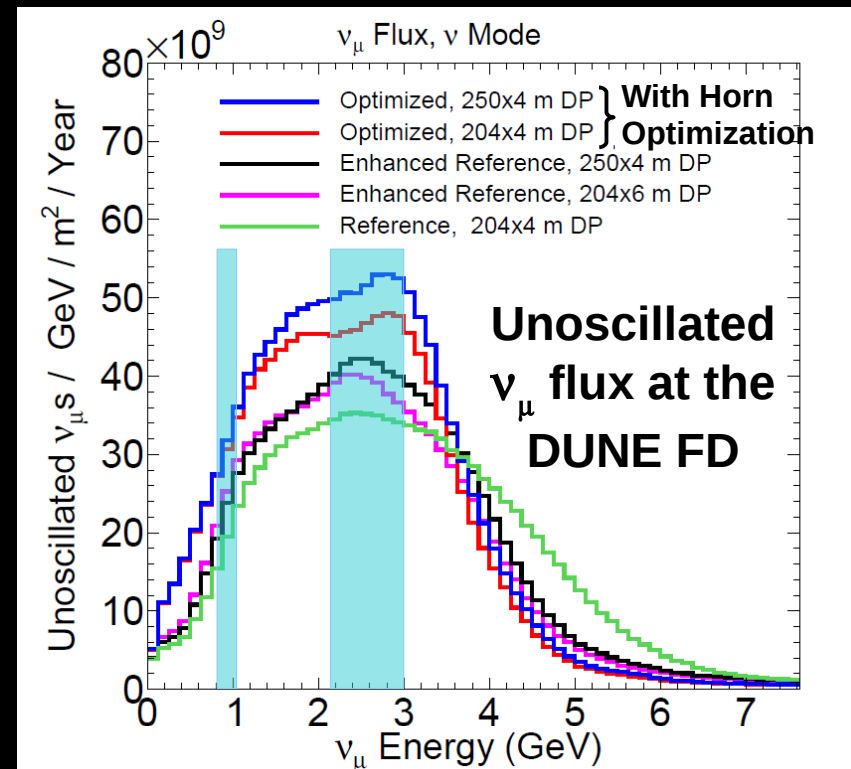
# Experimental Infrastructure: The DUNE Near Detector

- Detector requirements
  - Constrain **flux rate and shape** to the few % level
  - Charge ( $\nu/\bar{\nu}$ ) separation
  - Hadronic shower composition
    - Ar40 & Ca40 nuclei
    - $\nu/\bar{\nu}$  differences
  - Constrain relevant cross sections
  - Provide a wealth of physics measurements
- Detector Options
  - Fine Grained Tracker (reference)
  - LArTPC
  - High pressure GArTPC
  - Hybrid detector (ArTPC + FGT)



# Experimental Infrastructure: The FNAL → SURF Beam

- Beam requirements
  - 1.2 MW, upgradeable to 2.3 MW (120GeV protons):
    - POT/pulse:  $7.5 \times 10^{13}$  p
    - Cycle time: 1.2 sec
    - Uptime: 56%
  - Direction  $5.8^\circ$  downward
  - Wide-band spectrum covering the 1<sup>st</sup> and 2<sup>nd</sup> oscillation maxima
- Upgrades from reference design
  - PIP-II: increase p throughput
  - Horn current: 200 kA → 230 kA
  - Target design: C → Be, shape
  - Decay Pipe: 204 m → 250 m
  - Horn design optimization



- Can use 60 - 80 GeV protons
  - Increase flux at 2<sup>nd</sup> max
  - Reduces high energy tail
  - Need more POT to maintain power

# The Path to the Full Exposure

- A “Conceptual Design Review” is being held next month
- Goal: Install the first 10 kt underground on the 2021/22 timescale
  - Begin underground physics program, and engage collaboration
  - Test all aspects of the the underground installation and detector performance
  - Ready for beam physics program when beam turns on
- Remaining modules, up to 40 kt, installed in rapid succession
  - Initial 10 kt installation provides infrastructure for required conventional facilities
  - Opportunity for combination of multiple detector technologies
- Leverage intermediate neutrino program to inform design, and improve detector performance
- Construction of a fine grained near detector
- Collect beam data by 2024, and run for ~10 exposure-yr

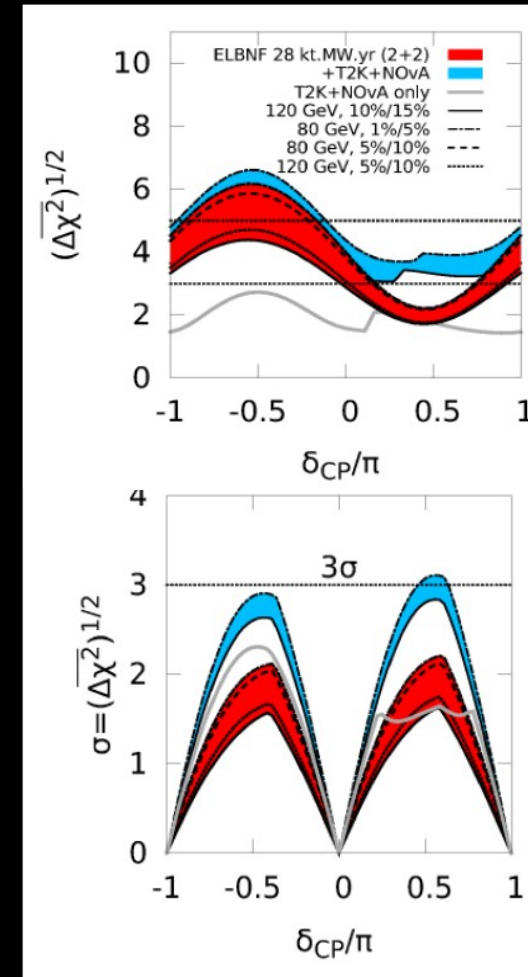
# Input From the Intermediate $\nu$ Program

- In addition to the in-situ measurements from the beamline monitoring, and the DUNE ND and FD, many external measurements are required
- NA61/SHINE and MIPP will provide data for hadron production model tuning used in beamline simulations
- Electron scattering at JLab will provide data on the nuclear structure of Ar
- Test beam LArTPCs: CAPTAIN, LArIAT, ProtoDUNE (single & double phase)
  - High statistics data on detector response required for calibrations
  - Allows for in-situ tests of detector components and comparison of detector technologies
- LArTPCs in neutrino beams: MicroBooNE, SBND, and ICARUS
  - Test and refine reconstruction algorithms and calibration methods
  - Measure cross sections and nuclear effects on Ar40
- Other cross section experiments like Minerva and ND280 (T2K) will map out cross sections over a wide energy range and nuclear targets
- Neutrino event generator development and tuning

# Physics with the First 10 kt\*

\*Assuming a 50 kt-yr exposure

- Baryon number violation
  - 50 kt-yr will competitive limits / signal events for  $p \rightarrow K + \bar{\nu}$
  - Early measurements of background rates for other decay channels
- Core-collapse supernova neutrinos
  - Largest detector sensitive to  $\nu_e$  via  $\nu_e + \text{Ar}^{40} \rightarrow e + \text{K}^{40}$
  - Prompt supernova alert due to early  $\nu_e$  production
  - 100's to ~1,000 events at ~10 kpc
- Atmospheric neutrinos
  - Provide ~2500  $\nu_e$  CC events
  - Test reconstruction and allow for leptonic and hadronic energy scale calibrations
- Accelerator neutrino (right)
  - Expected events:  $\nu_e$   $94 \pm 23$ ,  $\bar{\nu}_e$   $23 \pm 5$  (NH,  $\delta_{cp} = [-\pi/2, 0, \pi/2]$ )
  - Improved MH sensitivity over NOvA+T2K, even better combined
  - CPV sensitivity commensurate with NOvA+T2K, better combined



# Novel Features of the Experimental Design

- DUNE calls for unprecedented precision in a  $\nu$  experiment
- Achieving this precision will require hard work, innovation, and a start-of-the-art experimental design
- LArTPCs allows for high resolution of final state particle 4-momenta
  - The resolution  $\delta_{cp}$  largely limited by energy scale uncertainties which are limited by hadronic system reconstruction
  - Nearly background free to proton decay searches
  - Access to  $\nu_e$  flux from supernovas
- The DUNE FGT ND