Supernova Data-Compression in MicroBooNE

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Outline

The MicroBooNE experiment

Continuous data-stream and the need for compression

The MicroBooNE data-compression scheme

Performance studies

Note: I will focus on the Monte-Carlo studies which I have been performing, but there is a lot of work on the development and firmware-implementation of this data-compression scheme performed by the rest of the Nevis electronics team.

The MicroBooNE Experiment

MicroBooNE is a 170 ton LAr TPC neutrino experiment which sits on the Booster beam at Fermilab.

MicroBooNE aims to:

- investigate the MiniBooNE Low Energy Excess.
- perform v-Ar cross-section measurements in the 1 GeV region.
- LAr neutrino detector technology R&D.
- + many more physics topics, including supernova neutrinos.

The MicroBooNE Experiment



Turned on in early August. First cosmic tracks!

40 cm

E

26

Run 1148 Event 778. August 6th 2015 17:16

— MicroBooNE vs. DUNE

Continuous Data Stream

MicroBooNE is sensitive to supernova neutrinos... ...but we don't have a supernova trigger.

We expect to see O(10) neutrino interactions (10s of MeV) in a O(10 sec) timespan. Hard to trigger on with a surface detector.

Instead, we rely on the SuperNova Early Warning System (SNEWS), a trigger from larger experiments mostly used by telescopes around the world.

The plan:



Continuous Data Stream

What does it mean to save data continuously?

The MicroBooNE TPC Data:

8256 wires on 3 planes 2 Bytes per sample @ 2 MHz sampling → 33 GB/sec.

Limited bandwidth \rightarrow we need a compression factor of x80

How do we plan on achieving this?

Huffman compression is not enough. \rightarrow use a lossy compression scheme

Continuous Data Stream



Continuous Data Stream - Physics Constraints

MicroBooNE's compression-scheme needs to efficiently save data associated with charge deposited by interactions from supernova neutrinos.

Supernova neutrino interactions in MicroBooNE:

 $v_{e} + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$

Produce 10s of MeV e- and 1-3 MeV de-exitation photons.

This means saving pulses from O(1-10 MeV) particles.

MicroBooNE Data-Compression Scheme

The goal is to limit the data-flow.

Do this by saving only a subset of the TPC data:

- remove quiet/noise regions of waveforms
- save pulses caused by drifting charge in TPC

General rules:

- compression rate limited by our S/N.
- we can't beat the cosmic muon background.



Saved

MicroBooNE Data-Compression Scheme

Possible envisioned compression scheme:



Step 1) search for baseline in stable region.

Step 2) save pulse if above/below a baseline-subtracted threshold.



Channel-by-channel settable parameters and thresholds.

...Scheme details evolving ¹⁰

Monte-Carlo Performance Studies

Regardless of how one chooses to compress the data, it is important to study the effects of any data-suppression scheme.

Questions to address:

- What compression factor is achieved?
- What are the inefficiencies?
- How much / what information is lost?
- Energy dependence?

Monte-Carlo Performance Studies



Compression Factor = outgoing/incoming data-size

Monte-Carlo Performance Studies



Pulse-finding efficiency for charge on single wire



Pulse-finding efficiency for charge from single particle



Simulation Framework

A simulation tool for a data-compression scheme needs to:

- be able to evaluate different compression schemes
- evaluate/validate the scheme's decision making logic
- evaluate the scheme's performance and efficiency



Simulation Framework

Framework is dependent on C++/ROOT. Designed so that it can be easily interfaced with other tools (i.e. LArSoft)



Conclusions

MicroBooNE is a LArTPC currently operating in the Booster Neutrino Beamline.

MicroBooNE has a dedicated data-stream for a supernova neutrino search.

This data-stream is subject to a lossy compression, to reduce data-rates.

The validation and evaluation of possible data-compression schemes is a necessary step in order to understand:

- compression scheme performance
- efficiencies and physics impact

Do this in a flexible and re-usable framework.

Backup

TPC Readout Electronics



Supernova Neutrino Flux



time-dependent luminosity and mean energy by flavor

Supernova de-excitation photons



MC efficiency estimation



Searching for deposited charge