



Long-Baseline Sterile Neutrino Searches: The MINOS(+) and NOvA Experiences

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Foreword

Δm² (eV²)

10

10-1

Antineutrino

10-2

10-1

sin²20

10-3

- Strongest indication (3.8 σ) of large Δm^2 oscillations from LSND's observation of excess $\bar{\nu}_e$ appearance in $\bar{\nu}_\mu$ beam over 30 m short baseline
- No evidence seen by similar muon decay-at-rest experiment KARMEN
- MiniBooNE saw ~ 3σ indication for anomalous $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$, but at somewhat different L/E from LSND



Neutrino

 10^{-3}

Phys. Rev. Lett.110, 161801

10-2

- However, no evidence for sterile mixingdriven disappearance seen
- Long-Baseline experiments like MINOS(+), NOvA, and DUNE can provide powerful probes of v_{μ} , \bar{v}_{μ} disappearance driven by sterile mixing

 $sin^2 2\theta$

Long-Baseline $v_{\mu} \rightarrow v_{s}$ Mixing Probe

- Look for energy-dependent depletion of CC events at the FD modulated by three-flavor survival probability
 - plot to the right assumes 3+1 model

- NC interactions initiated by v_e , v_μ , v_τ are topologically indistinguishable
 - NCs are insensitive to 3-flavor oscillations
- However, v_µ→v_s mixing would reduce NC interaction rate as v_s do not interact in the detectors
- Look for energy-dependent depletion of NC events at the FD



What LBL Expts. and Others Measure

$$P_{ee} \approx 1 - 2\sin^2 2\theta_{14} \times \sin^2 \frac{\Delta m_{41}^2 L}{E} \qquad |U_{e4}|^2$$

• MiniBooNE, LSND, and KARMEN, T2K ND (v_e, \bar{v}_e appearance): θ_{14} , θ_{24} $P_{\mu e} \approx 2 \sin^2 2\theta_{14} \sin^2 \theta_{24} \times \sin^2 \frac{\Delta m_{41}^2 L}{E} \qquad 4|U_{e4}|^2|U_{\mu 4}|^2$

• MiniBooNE, CDHS, CCFR (v_{μ} , \bar{v}_{μ} disappearance): θ_{24}

$$P_{\mu\mu} \approx 1 - 2\sin^2 2\theta_{24} \times \sin^2 \frac{\Delta m_{41}^2 L}{E} \qquad |U_{\mu4}|^2$$

• MINOS/MINOS+, SuperK, NOvA (v_{μ} , \bar{v}_{μ} disappearance and NC): θ_{24} , θ_{34}

$$1 - P_{\mu s} \approx \cos^2 \theta_{24} \sin^2 \theta_{34} \times \sin^2 \frac{\Delta m_{41}^2 L}{E} \sin^2 \frac{\Delta m_{31}^2 L}{E} \qquad |U_{\mu 4}|^2, \ |U_{\tau 4}|^2$$

MINOS/MINOS+ in a Nutshell

Main Injector Neutrino Oscillation Search





Long-baseline neutrino oscillation experiment



- Measure NuMI Neutrino beam energy and flavor composition with two detectors over 735 km
 - L/E ~ 500 km/GeV



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Event Topologies

MINOS and 4-Flavor Oscillations

- ▶ $v_{\mu} \rightarrow v_{s}$ mixing causes energy-dependent depletion of NC and v_{μ} -CC energy spectra w.r.t 3-flavor mixing
- Small Δm_{43}^2 (> Δm_{32}^2):
 - FD spectral distortions at energies above 3-flavor oscillation maximum
 - No ND effects
- Medium Δm^2_{43} :
 - Rapid oscillations at FD average out
 - No ND effects
 - Counting experiment
- Large Δm^2_{43} :
 - Rapid oscillations at FD average out
 - ND spectral distortions affect extrapolation to FD

MINOS FD CC and NC Selections

- MINOS was designed to separate v_μ CC interactions from NCs, but isolating a pure NC sample is more difficult
 - $\bullet\,$ Main background originates from inelastic (high-y) $v_{\mu}\,$ CC events
 - NC events selected with 89% efficiency and 61% purity in FD
 - ${\ensuremath{\, \circ }}$ 97% of v_e CC selected as NC
 - In ND, need to worry about reconstruction failures due to pile-up => one of the largest systematic uncertainties in NC analysis

MINOS FD CC and NC Energy Spectra

- Comparison with 3-flavor prediction for full MINOS low-energy beam neutrino mode sample: 10.56×10²⁰ POT
- Selected v_µ-CC and CC candidates in both detectors
 - $\odot~2563~\nu_{\mu}\text{-}CC\text{-like}$ events in FD
 - 1211 NC-like events in FD
- Looked at model-independent rate measurement for NC Selection

 $R = \frac{N_{data} - \sum B_{CC}}{S_{NC}} \qquad \begin{array}{c} \text{Predicted CC}\\ \text{background}\\ \text{from all flavors} \end{array}$ $R = 1.08 \pm 0.11 \text{ (0 - 40 GeV)}$ $R = 1.11 \pm 0.10 \text{ (0 - 3 GeV)}$

No evidence for oscillations into sterile neutrinos at Δm²₄₃≈0.5 eV²

MINOS 4-Flavor Analysis Strategy

- Assume 3+1 sterile neutrino mixing scenario
 Apply oscillations to both ND and FD
 - Use distance to meson decay point
 - Fit for $|\Delta m^2_{32}|$, θ_{23} , $|\Delta m^2_{43}|$, θ_{24} , θ_{34}
- To account for ND distortions, fit oscillated
 F/N ratio directly to F/N data ratio
 Include constraint on ND rate
- Carefully assessed systematic uncertainties affecting high-energy tail of spectrum with respect to previous CC and NC analyses
 Re-evaluated beam flux uncertainties
- Log-likelihood surfaces are Feldman-Cousins corrected

Systematics

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MINOS Disappearance Limit

MINOS 90% C.L. exclusion limit ranges over 4 orders of magnitude in Δm^2_{43} ! Strongest constraint on v_{μ} disappearance into v_s for $\Delta m^2_{43} < 1 \text{ eV}^2$

Comparison to Appearance Results

- With MiniBooNE Neutrino Mode
- Assuming 3+1 model, combine MINOS disappearance 90% C.L. limit in θ₂₄ to Bugey reactor experiment 90% C.L. disappearance limit in θ₁₄
- Working with Daya Bay to produce MINOS & Daya Bay combined limit

Bugey limit computed from GLoBES 2012 fit using new reactor fluxes, provided by Patrick Huber

MINOS data increases tension between null and signal results for $\Delta m^2_{43} < 1 \text{ eV}^2$

Comparison to Appearance Results

- With MiniBooNE Antineutrino Mode
- Assuming a 3+1 model and CPT conservation so SBL neutrino and antineutrino oscillations are identical
- Working on sterile neutrino search in 3.4×10²⁰ POT of MINOS antineutrino running

MINOS data increases tension between null and signal results for $\Delta m^2_{43} < 1 \text{ eV}^2$

Improvements with MINOS+

- CC events at the FD highly suppressed by 3-flavor oscillations, and absence of high energy tail limits range of Δm²₄₃ that can be probed.
- However, NC events will be the largest sample measured by NOvA

NOvA NC Disappearance due to $v_{\mu} \rightarrow v_{s}$ Mixing

• For first analysis, will assume $\Delta m^2_{41} < 0.5 \text{ eV}^2$, so that we have no oscillations at the ND, but rapid oscillations average out at the FD

NOvA NC Selection

- Main challenge for NC analysis is cosmic neutron interactions in the upper part of the FD. Have developed preliminary cut-based NC selection:
- ND Purity = 63.4%;

- Efficiency [All cuts/(DQ+Fid.)] = 32.2%
- FD Purity = 69.3% or 83.6% without cosmics; Efficiency [All cuts/(DQ+Fid.)] = 13.5%

NOvA Simulation

NOvA Sensitivities vs MINOS

• NOvA plot shows constraints on θ_{24} and θ_{34} for the cut-based NC selection compared with prediction using a perfect NC selection

Brief Thoughts on DUNE Sterile Searches

- On-axis broadband beam will confer DUNE sensitivity over a wide range of sterile neutrino masses
- NC channel provides a 3-flavor-independent handle on sterile-induced v_{μ} disappearance
- Underground detector, exquisite NC/CC separation, and 1.2 MW beam provide much more powerful sensitivity than MINOS
- Can also do SBL searches for v_e appearance using the ND, FD atmospherics, Xtra Dimens.

DUNE V., disappearance

Reconstructed Energy (GeV)

(⊽,+v,) CC

CDR Reference Desig

150 kt-MW-vr ⊽ mode

sin²(θ₂₃)=0.45

کی ²⁵⁰

Events/0.25 120 100

R. Gandhi, B. Kayser, M. Masud, S. Pakrash, arXiv:1508.06275

Summary

- Long-baseline sterile neutrino searches complement current and future SBL searches, and provide unique sensitivity to large portions of parameter space and sterile mixing matrix elements such as $|U_{\tau 4}|^2$
- DUNE should have powerful sensitivity to sterile mixing over long-baselines. Sensitivity to v_e appearance at the ND should be investigated. FD atmospheric searches à la SuperK also possible.
- NC disappearance channel should offer a handle in disentangling CP violation from light sterile neutrino admixture, assuming Fermilab's SBL program would have not fully excluded that possibility by then.

Backup

3+1 Oscillations

$$P(\nu_a \to \nu_b) = \delta_{ab} - 4\sum_{j>i} \mathcal{R}(U_{aj}^* U_{bj} U_{ai} U_{bi}^*) \sin^2 \Delta_{ji} + 2\sum_{j>i} \mathcal{I}(U_{aj}^* U_{bj} U_{ai} U_{bi}^*) \sin 2\Delta_{ji}$$

$$\begin{split} P_{\nu_{\mu} \to \nu_{\mu}} &= 1 - 4 \bigg\{ |U_{\mu3}|^2 (1 - |U_{\mu3}|^2 - |U_{\mu4}|^2) \sin^2 \Delta_{31} + |U_{\mu4}|^2 |U_{\mu3}|^2 \sin^2 \Delta_{43} \\ &+ |U_{\mu4}|^2 (1 - |U_{\mu3}|^2 - |U_{\mu4}|^2) \sin^2 \Delta_{41} \bigg\}, \\ P_{\nu_{\mu} \to \nu_{\alpha}} &= 4 \mathcal{R} \bigg\{ |U_{\mu3}|^2 |U_{\alpha3}|^2 \sin^2 \Delta_{31} + |U_{\mu4}|^2 |U_{\alpha4}|^2 \sin^2 \Delta_{41} \\ &+ U_{\mu4}^* U_{\alpha4} U_{\mu3} U_{\alpha3}^* (\sin^2 \Delta_{31} - \sin^2 \Delta_{43} + \sin^2 \Delta_{41}) \bigg\} \\ &+ 2 \mathcal{I} \bigg\{ U_{\mu4}^* U_{\alpha4} U_{\mu3} U_{\alpha3}^* (\sin 2\Delta_{31} - \sin 2\Delta_{41} + \sin 2\Delta_{43}) \bigg\}, \end{split}$$

3+1 Oscillations

$$\begin{split} |U_{e3}|^2 &= c_{14}^2 s_{13}^2, \\ |U_{e4}|^2 &= s_{14}^2, \\ |U_{\mu3}|^2 &= s_{13}^2 s_{14}^2 s_{24}^2 + c_{13}^2 s_{23}^2 c_{24}^2 - \frac{1}{2} s_{23} s_{14} \sin(2\theta_{13}) \sin(2\theta_{24}) \cos(\delta_1 + \delta_2), \\ |U_{\mu4}|^2 &= c_{14}^2 s_{24}^2, \\ |U_{\tau3}|^2 &= s_{13}^2 s_{14}^2 c_{24}^2 s_{34}^2 + c_{13}^2 s_{23}^2 s_{24}^2 s_{34}^2 + c_{13}^2 c_{23}^2 c_{34}^2 + \frac{1}{2} s_{23} s_{14} s_{34}^2 \sin(2\theta_{13}) \sin(2\theta_{24}) \cos(\delta_1 + \delta_2) \\ &- \frac{1}{2} c_{23} s_{14} c_{24} \sin(2\theta_{13}) \sin(2\theta_{34}) \cos\delta_1 - \frac{1}{2} c_{13}^2 s_{24} \sin(2\theta_{23}) \sin(2\theta_{34}) \cos\delta_2, \\ |U_{\tau4}|^2 &= c_{14}^2 c_{24}^2 s_{34}^2, \\ |U_{\tau4}|^2 &= s_{13}^2 s_{14}^2 c_{24}^2 c_{34}^2 + c_{13}^2 s_{23}^2 s_{24}^2 c_{34}^2 + c_{13}^2 c_{23}^2 s_{34}^2 + \frac{1}{2} s_{23} s_{14} c_{34}^2 \sin(2\theta_{13}) \sin(2\theta_{24}) \cos(\delta_1 + \delta_2) \\ &+ \frac{1}{2} c_{23} s_{14} c_{24} \sin(2\theta_{13}) \sin(2\theta_{34}) \cos\delta_1 + \frac{1}{2} c_{13}^2 s_{24} \sin(2\theta_{13}) \sin(2\theta_{34}) \cos\delta_2, \\ |U_{\tau4}|^2 &= s_{13}^2 s_{14}^2 c_{24}^2 c_{34}^2 + c_{13}^2 s_{23}^2 s_{24}^2 c_{34}^2 + c_{13}^2 c_{23}^2 s_{34}^2 + \frac{1}{2} s_{23} s_{14} c_{34}^2 \sin(2\theta_{13}) \sin(2\theta_{24}) \cos(\delta_1 + \delta_2) \\ &+ \frac{1}{2} c_{23} s_{14} c_{24} \sin(2\theta_{13}) \sin(2\theta_{34}) \cos\delta_1 + \frac{1}{2} c_{13}^2 s_{24} \sin(2\theta_{23}) \sin(2\theta_{34}) \cos\delta_2, \\ |U_{\tau4}|^2 &= c_{14}^2 c_{24}^2 c_{34}^2. \end{split}$$

Sterile Neutrinos

- Oscillations into light sterile neutrinos are a proposed explanation for anomalies seen in short-baseline accelerator experiments, reactor experiments (following re-evaluation of $\bar{\nu}_e$ reactor flux to be 3.5% higher), and from gallium anomaly
- Severe tension between appearance, and disappearance measurements or appearance measurements that see no signal

Kopp, Machado, Maltoni, Schwetz, arXiv:1303.3011

• Short and long-baseline accelerator, reactor, radioactive, and atmospherics experiments planned to resolve these anomalies over next 10-15 years

MINOS NC Event Selection

• NC/CC event separation achieved via cuts on topological variables

- Discard events with length > 47 planes
- Discard events with a track > 6 planes longer than the shower
- Same selection applied to data and MC in Far Detector

MINOS NC Systematics

- Normalization: 2.2%
 - Livetime, Near/Far reconstruction efficiency, fiducial mass
- Relative Hadronic Calibration: 2.1%
 - Inter-Detector calibration uncertainty
- Absolute Hadronic Calibration: [±10%, ±6.5%]
 - Hadronic Shower Energy Scale(±5.6%), Intranuclear rescattering([±8%, ±4%])
- Muon energy scale: 2%
 - Uncertainty in dE/dX in MC
- CC Contamination of NC-like sample: ±15%
- NC contamination of CC-like sample: ±25%
- Cross-section uncertainties:
 - m_A (QE) and m_A (Res): ±15%
 - KNO scaling: ±33%
- Near Detector NC Selection: ±10% in 0-1 GeV bin
- Far Detector NC Selection: ±5% if E < I GeV, <2.5% if E > I GeV
- Beam uncertainty: I σ error band around beam fit results

Uncertainty	ΔR (0-120 GeV)
Absolute E_{Hadronic}	0.4%
Relative E_{Hadronic}	0.0%
Normalization	3.2%
CC Background	2.1%
ND Selection	2.7%
FD Selection	2.5%
Total	5.3%

MINOS+ Sensitivity to Extra-Dimensions

MINOS+ Decay, Decoherence

