

# LONG-BASELINE OSCILLATION PHYSICS IN DUNE

---

Elizabeth Worcester (BNL)

Neutrino – Latin America Workshop

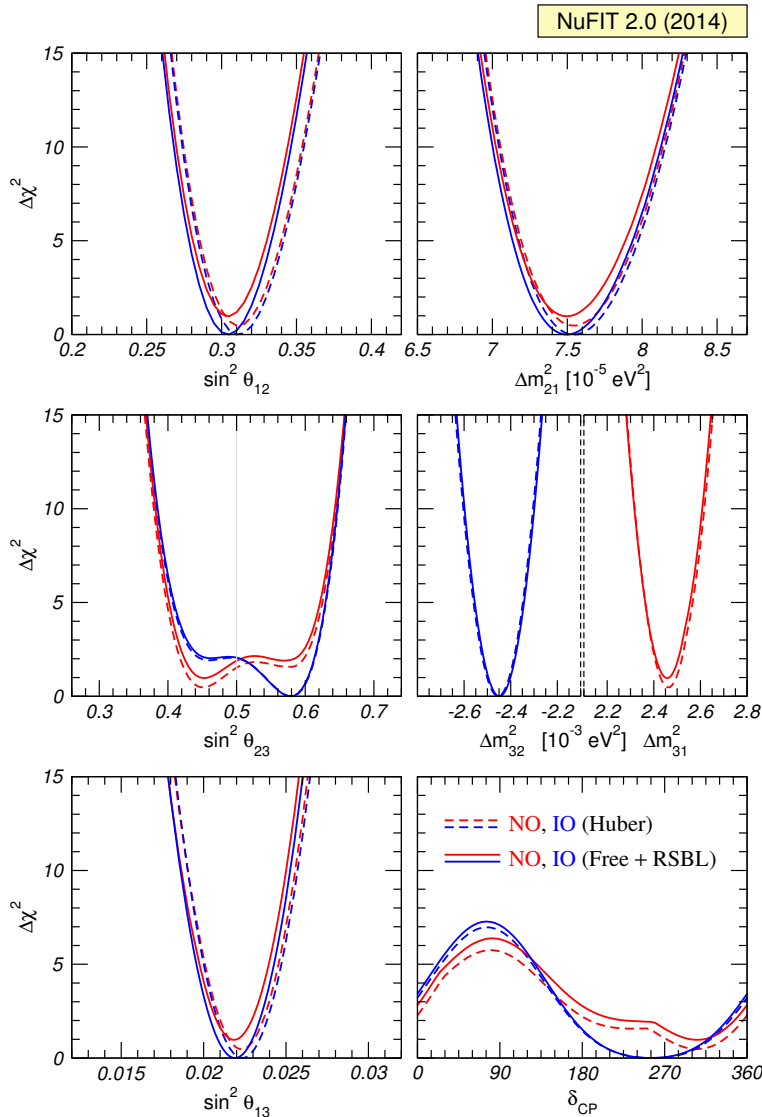
April 28, 2016

# Overview



- Introduction to long-baseline neutrino oscillation
- Oscillation sensitivity in DUNE
- Systematic uncertainties
- Additional physics topics

# Oscillation Parameters



- NuFit 2014
  - <http://www.nu-fit.org/>
  - Includes results through NOW 2014
  - $\theta_{13}$ ,  $\theta_{12}$ ,  $\Delta m^2_{21}$ ,  $|\Delta m^2_{32}|$  each known to a few percent
  - $\theta_{23}$  known to  $\sim 6\%$  (octant unknown)
  - Some preference for  $\delta_{CP} < 0$
- Further constraints expected from existing and planned experiments:
  - Hints from T2K and NOvA suggest  $\delta_{CP} < 0$
  - External constraints on mixing angles improve early sensitivity
  - Measurements or hints of MH or  $\delta_{CP}$  value could influence run plans
- Ultimate DUNE goals include precise measurements of  $\theta_{13}$ ,  $\theta_{23}$ ,  $\Delta m^2_{32}$ , and  $\delta_{CP}$  for unitarity and sum rule tests

# $\nu_e$ Appearance

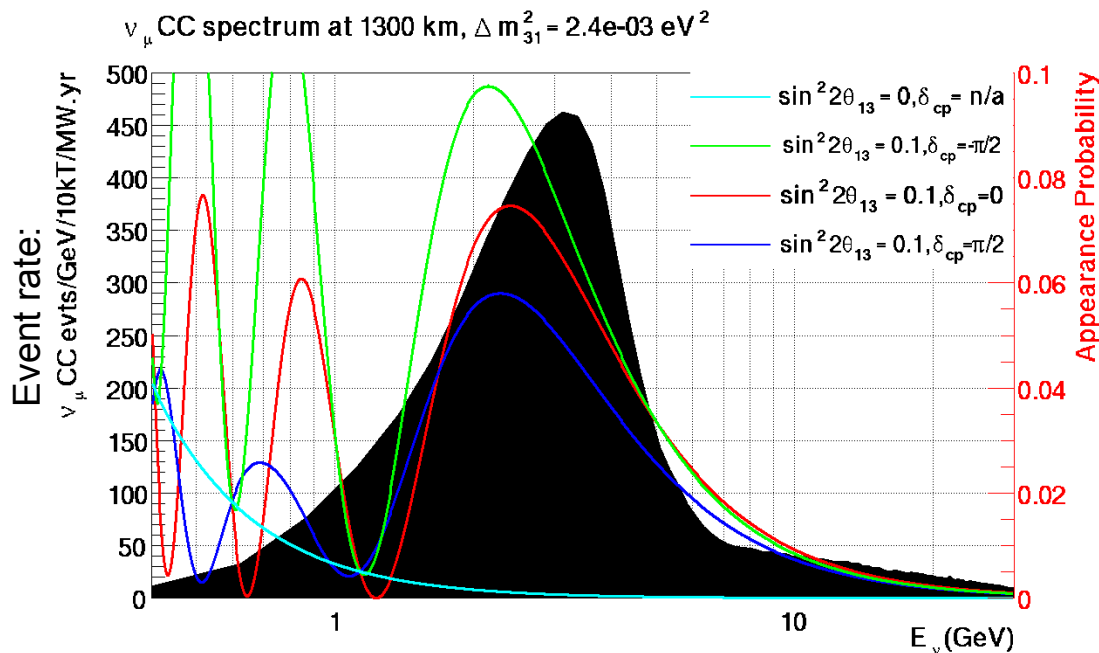
$$P(\nu_\mu \rightarrow \nu_e) \approx \boxed{\sin^2 2\theta_{13}} \boxed{\sin^2 \theta_{23}} \frac{\sin^2(\Delta_{31} - aL)}{(\Delta_{31} - aL)^2} \Delta_{31}^2 +$$

$$\alpha \boxed{\sin 2\theta_{13}} \boxed{\cos \delta} \frac{\sin(aL)}{(aL)} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \cos \Delta_{32} -$$

$$\alpha \boxed{\sin 2\theta_{13}} \boxed{\sin \delta} \frac{\sin(aL)}{(aL)} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \sin \Delta_{32}$$

$$a = G_F N_e \sqrt{2}$$

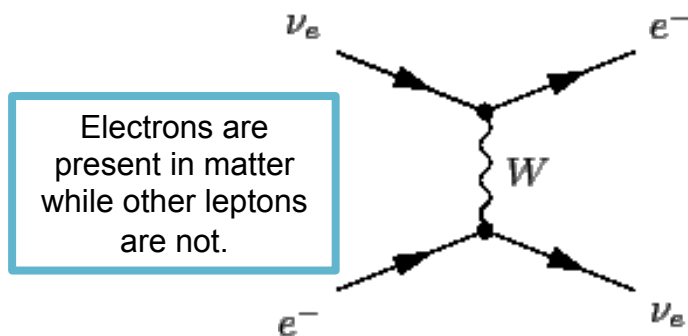
$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{4E}$$



- $\nu_e$  appearance amplitude depends on  $\theta_{13}$ ,  $\theta_{23}$ ,  $\delta_{CP}$ , and matter effects – measurements of all four possible in a single experiment
- Large value of  $\sin^2(2\theta_{13})$  allows significant  $\nu_e$  appearance sample

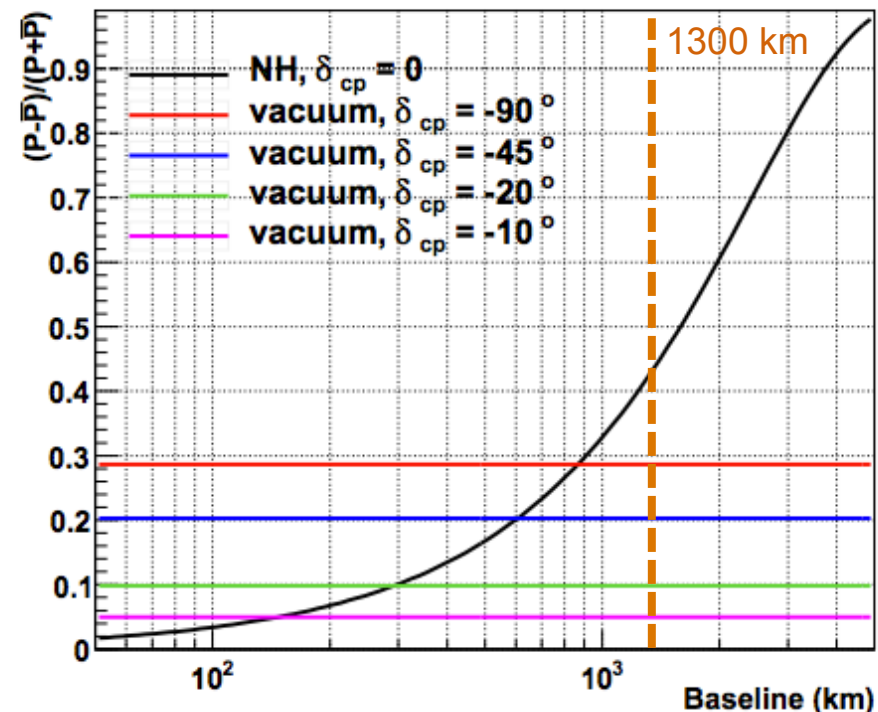
# Matter and CP Asymmetry

Charged-current coherent forward scattering on electrons:



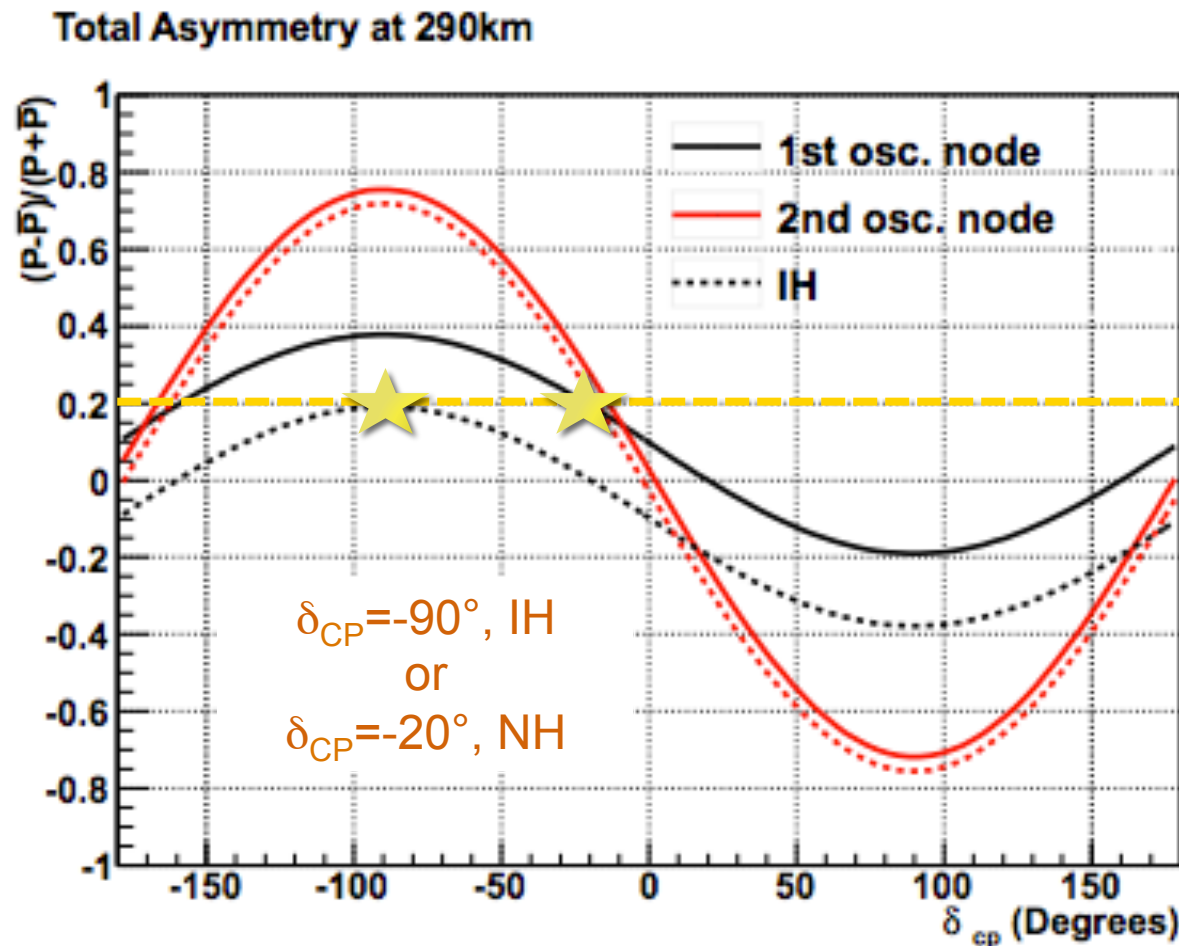
- CC process occurs for electron neutrinos only; muon and tau have only NC interactions with electrons
- Normal hierarchy: matter effect enhances appearance probability for neutrinos and suppresses it for antineutrinos (opposite for IH)

CP asymmetries in  $\nu_\mu \rightarrow \nu_e$  at 1<sup>st</sup> osc. node



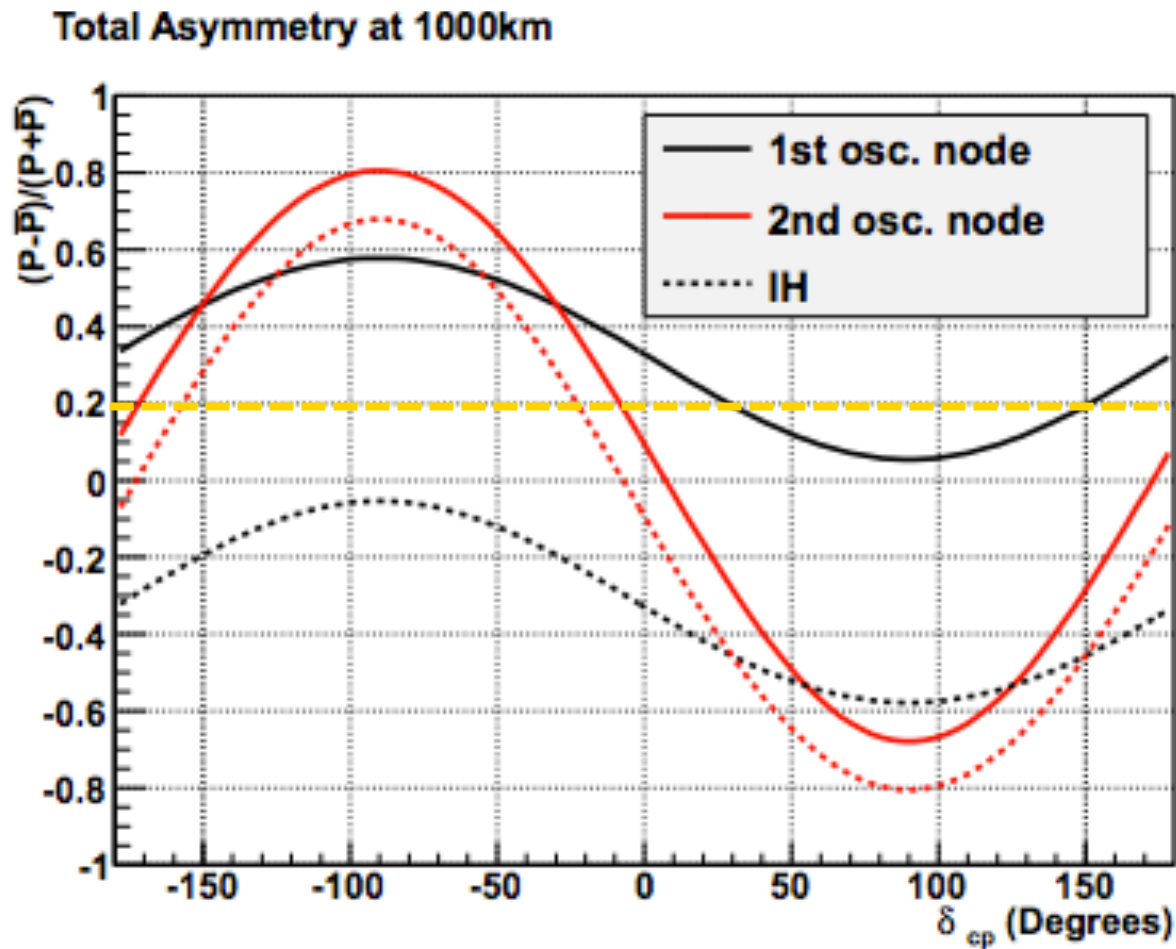
Matter asymmetry very important for long-baseline experiments!

# Matter and CP Asymmetry



Degeneracy between CP and matter asymmetry  
for 1<sup>st</sup> oscillation node at short baseline

# Matter and CP Asymmetry



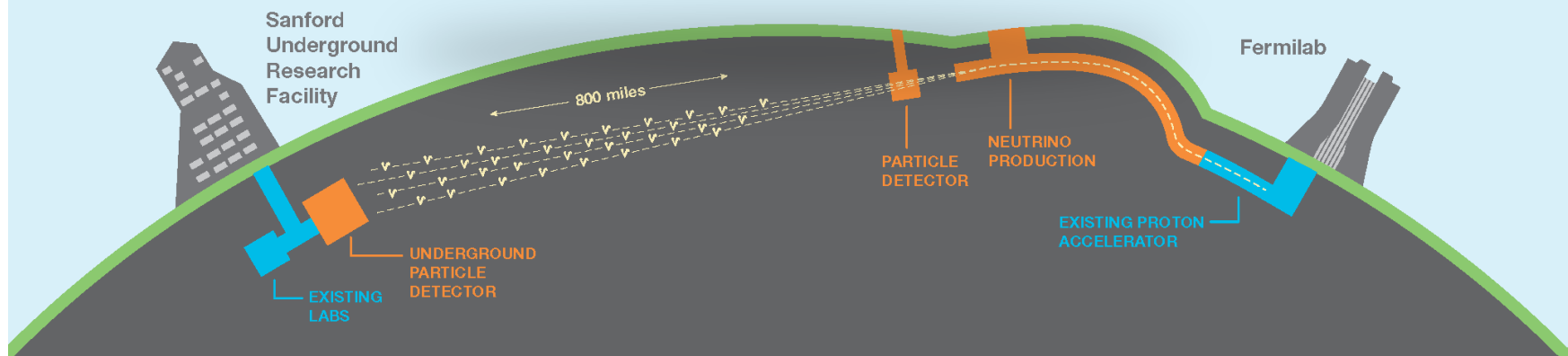
Longer baseline breaks degeneracy  
between CP and matter asymmetry



# DUNE



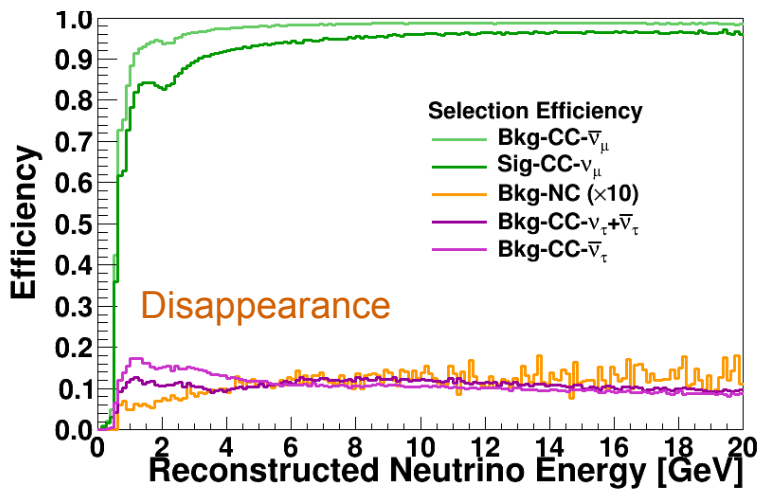
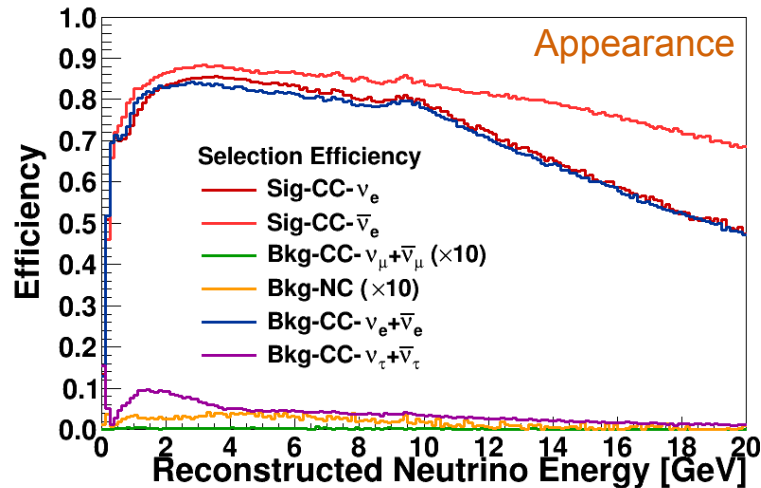
Measure  $\nu_e$  appearance and  $\nu_\mu$  disappearance in a wideband neutrino beam at 1300 km to measure MH, CPV, and neutrino mixing parameters in a single experiment.





# Signal Efficiency

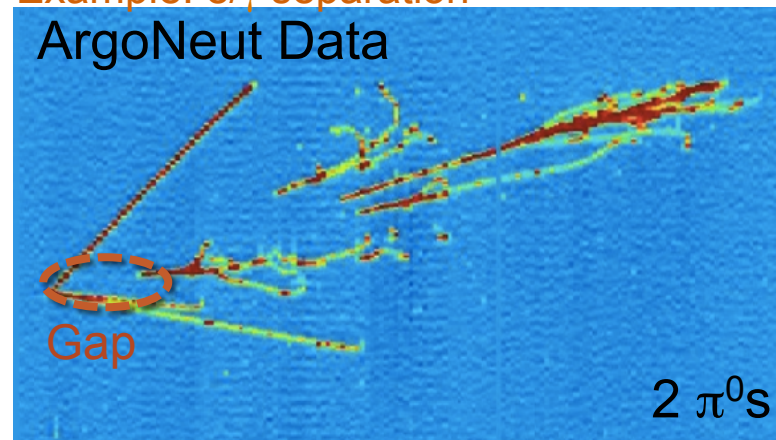
DUNE CDR:



- Expected efficiency from analysis of parameterized Fast MC
  - Detector performance parameters based on previous experiments and simulations
  - Reconstruction/analysis efforts underway to demonstrate performance with DUNE simulation
- Current generation of experiments have much to contribute to understanding of reconstruction/analysis in LArTPCs

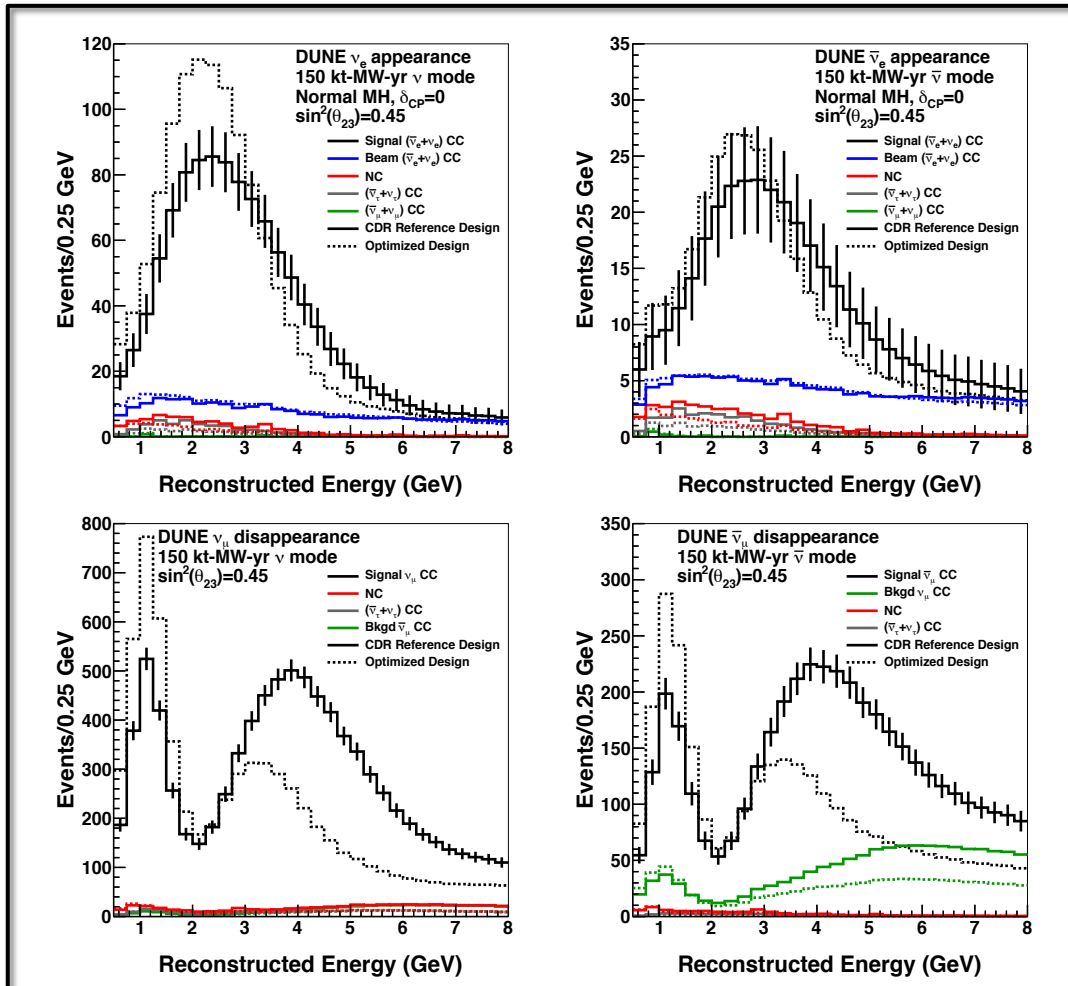
Example: e/ $\gamma$  separation

ArgoNeut Data



# Sensitivity Calculations

DUNE CDR:

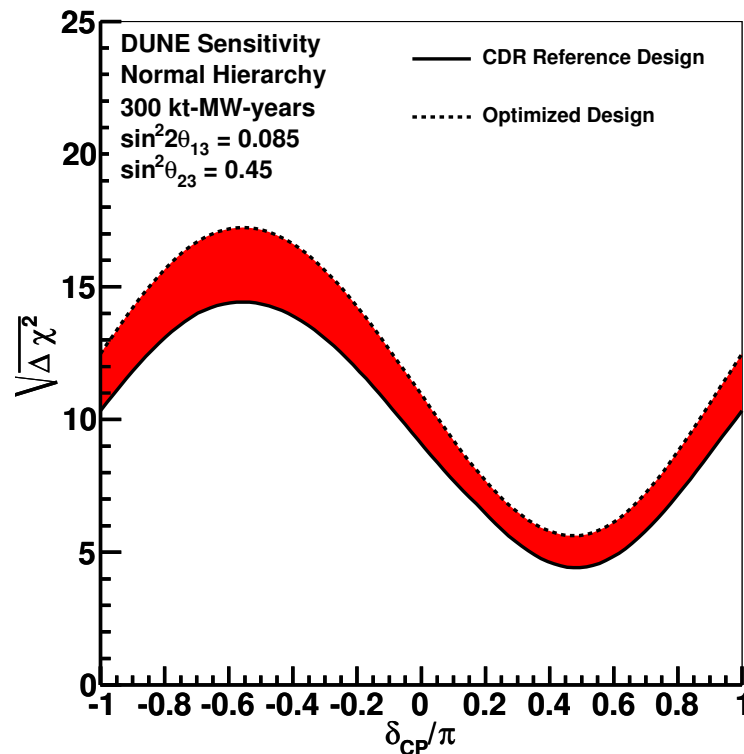


- GLoBES-based fit to four FD samples
- Two neutrino beam line designs considered
- GENIE event generator
- Reconstructed spectra predicted using detector response parameterized at the single particle level
- Simple systematics treatment
- GLoBES configurations to be made public soon

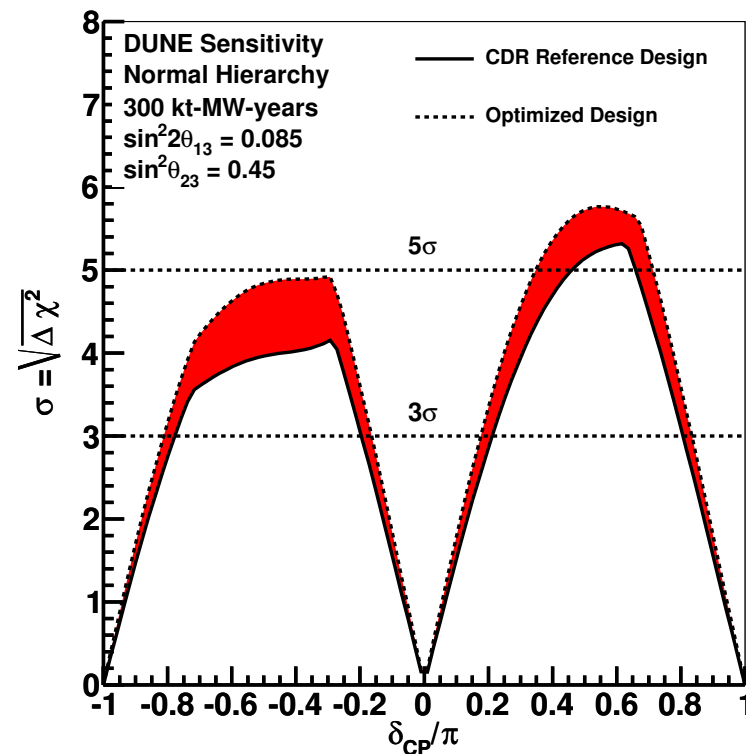
# DUNE Sensitivity

DUNE CDR:

Mass Hierarchy



CP Violation



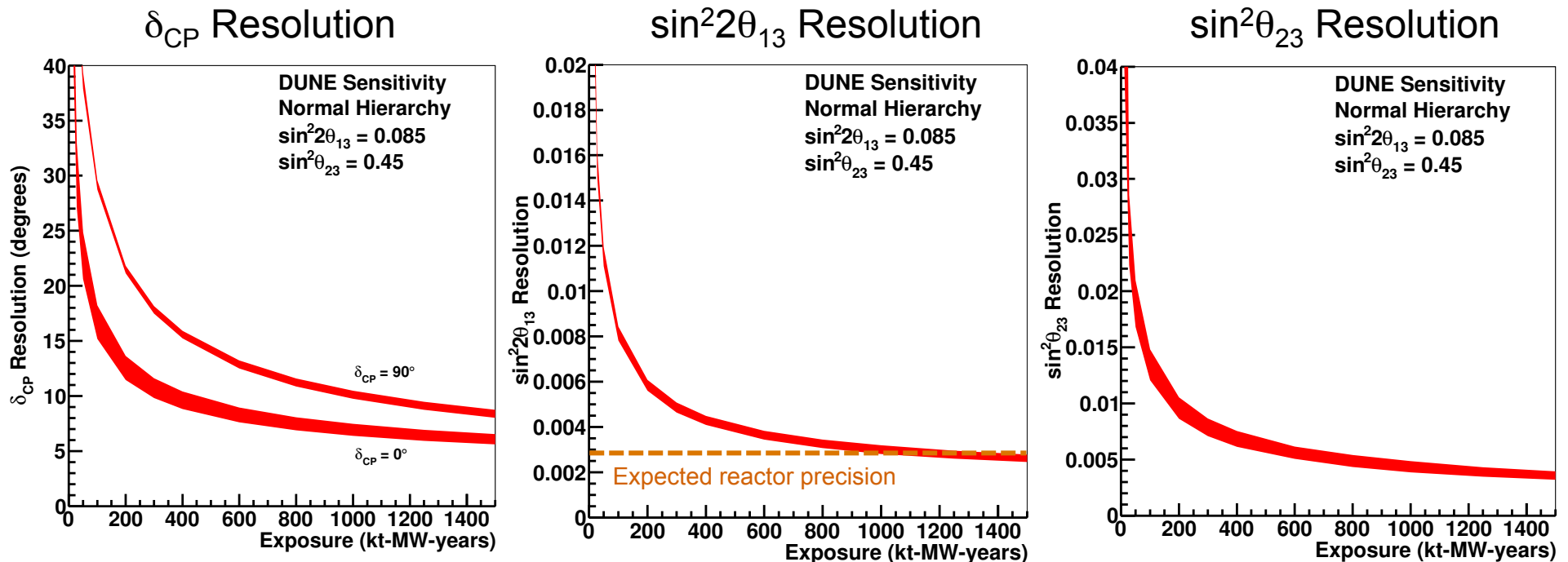
Width of band indicates variation among differing neutrino beam designs.

Exposure is 300 kt.MW.yr = 40 kt x 1.07 MW x (3.5 $\nu$ +3.5 $\bar{\nu}$ ) years.

Includes simple normalization systematics and oscillation parameter variations.

# Oscillation Parameter Sensitivity

DUNE CDR:



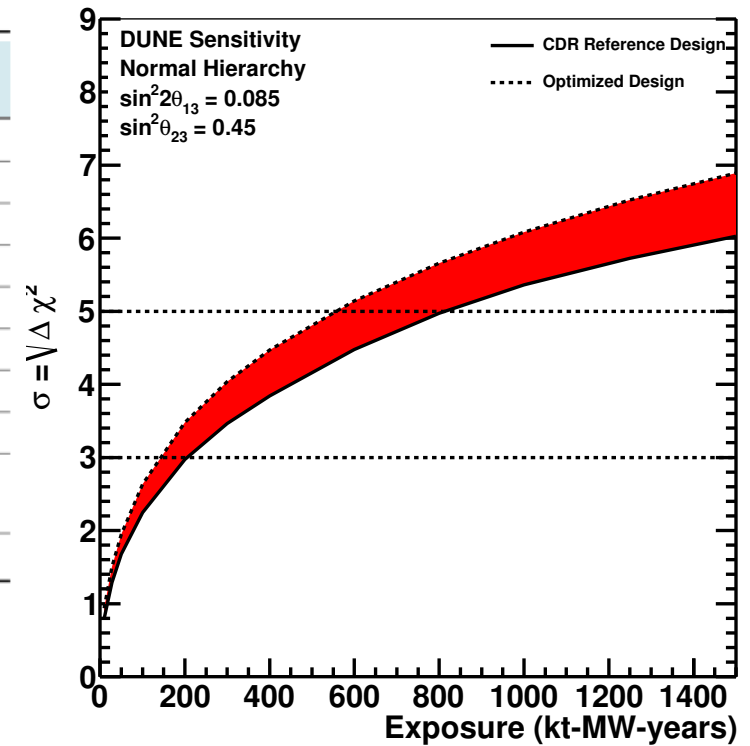
# DUNE Sensitivity Over Time

## DUNE CDR:

Physics milestone	Exposure kt · MW · year (reference beam)	Exposure kt · MW · year (optimized beam)
$1^\circ \theta_{23}$ resolution ( $\theta_{23} = 42^\circ$ )	70	45
CPV at $3\sigma$ ( $\delta_{CP} = +\pi/2$ )	70	60
CPV at $3\sigma$ ( $\delta_{CP} = -\pi/2$ )	160	100
CPV at $5\sigma$ ( $\delta_{CP} = +\pi/2$ )	280	210
MH at $5\sigma$ (worst point)	400	230
$10^\circ$ resolution ( $\delta_{CP} = 0$ )	450	290
CPV at $5\sigma$ ( $\delta_{CP} = -\pi/2$ )	525	320
CPV at $5\sigma$ 50% of $\delta_{CP}$	810	550
Reactor $\theta_{13}$ resolution ( $\sin^2 2\theta_{13} = 0.084 \pm 0.003$ )	1200	850
CPV at $3\sigma$ 75% of $\delta_{CP}$	1320	850

Interesting measurements will be made throughout the DUNE physics program!

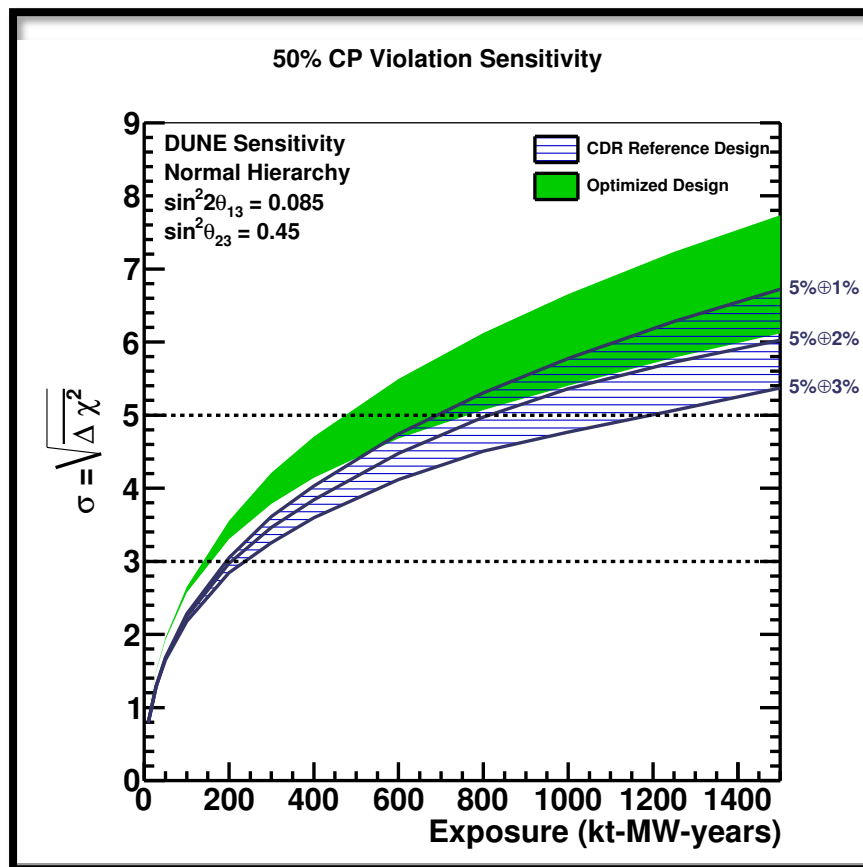
## CP Violation



Initial beam power: 1.07 MW at 80 GeV  
Planned upgrade to > 2 MW

# Systematic Uncertainty

## DUNE CDR:



- CPV measurement statistically limited for  $\sim 100$  kt-MW-years
- Sensitivities in DUNE CDR are based on GLoBES calculations in which the effect of systematic uncertainty is approximated using uncorrelated signal normalization uncertainties.
  - $\nu_\mu = \bar{\nu}_\mu = 5\%$
  - $\nu_e = \bar{\nu}_e = 2\%$
- Uncertainty in  $\nu_e$  appearance sample normalization must be  $\sim 5\% \oplus 2\%$  to discover CPV in a timely manner.

# Sources of Uncertainty

Source of Uncertainty	MINOS $\nu_e$	T2K $\nu_e$	Goal for DUNE $\nu_e$
Beam Flux	0.3%	3.2%	2%
Interaction Model	2.7%	5.3%	~2%
Energy Scale ( $\nu_\mu$ )	3.5%	Included above	Included in 5% $\nu_\mu$ uncertainty
Energy Scale ( $\nu_e$ )	2.7%	2.5% includes all FD effects	2%
Fiducial Volume	2.4%	1%	1%
Total Uncertainty	5.7%	6.8%	3.6%
Used in DUNE sensitivity calculations:			5% $\oplus$ 2%

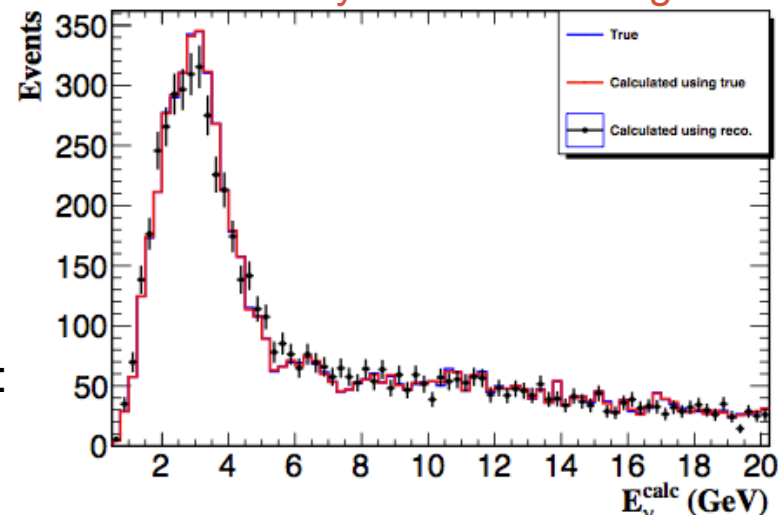
DUNE goals are for the *total* normalization uncertainty on the  $\nu_e$  appearance sample. The DUNE analysis will be a 3-flavor oscillation fit such that uncertainties correlated among the four FD samples will largely cancel.



# Strategy for Flux

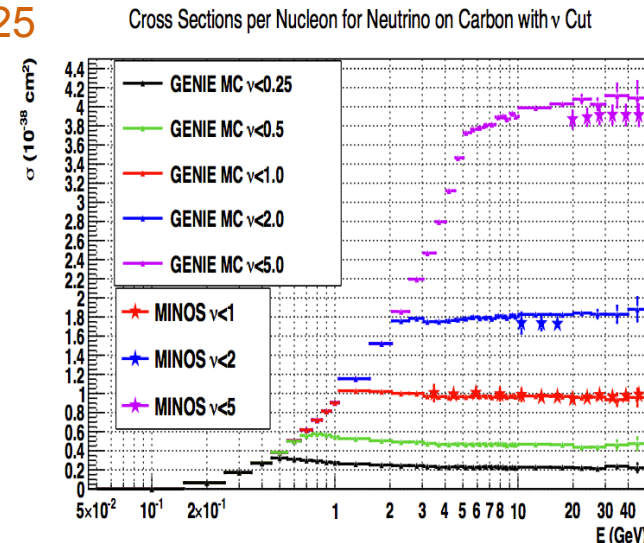
- Constrain absolute flux with near detector measurements of fully-leptonic neutrino interactions
  - Cross-sections known to high precision
  - Neutrino-electron scattering:  $\sim 3\%$  stat. ( $E_\nu < 5$  GeV)
  - Inverse muon decay:  $\sim 3\%$  stat. ( $E_\nu > 11$  GeV)
- Constrain flux shape using low- $\nu_0$  method: 1-2%
- Low- $\nu_0$  measurement for both  $\nu_e$  and  $\nu_\mu$  flux, in combination with hadron production data (NA61/SHINE), constrains ND/FD flux ratio at the 1% level

Fast MC study of  $\nu$ -e scattering:



120 events in MINERvA  $\rightarrow$  13% constraint on NuMI flux, statistically limited (J. Park thesis)

arXiv:1201.3025  
(Bodek et al.):



# Strategy for Interaction Model



- Prospects for improved interaction models:
  - Improved models becoming available
  - Intermediate neutrino program measurements in LAr TPCs
- ND constraint:
  - High precision near detector designed to constrain cross-section and hadronization uncertainties, resolving many individual particles produced by resonance and DIS interactions
  - Argon nuclear targets in ND allows significant cancellation of cross-section uncertainties common to near and far detectors
- FD constraint:
  - Four FD samples allow cancellation of uncertainties that are correlated between  $\nu_e/\nu_\mu$  or  $\nu/\bar{\nu}$

# Improving Interaction Models

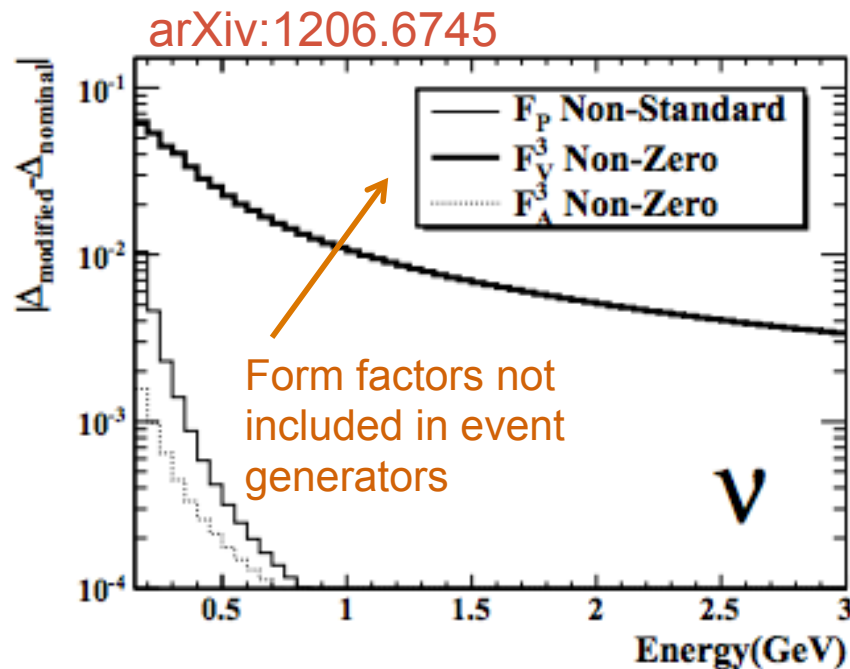


- Worldwide effort that will benefit DUNE!
- Alternative models being implemented in GENIE include:
  - Long- and short-range correlations among nucleons
  - Effect of random phase approximations
  - Meson exchange currents
  - 2p-2h effects in CCQE
  - Effective spectral functions
  - Coherent pion production
  - Alternative model of DIS interactions
  - Variation of tunable parameters within existing models
- Comparisons among generators
- Neutrino interaction data available or coming soon from:
  - ArgoNeuT, MINERvA, CAPTAIN-MINERvA, NOvA-ND, T2K-ND280,  $\mu$ BooNE, SBND, ICARUS, ...
- Electron-argon scattering data coming soon from JLab

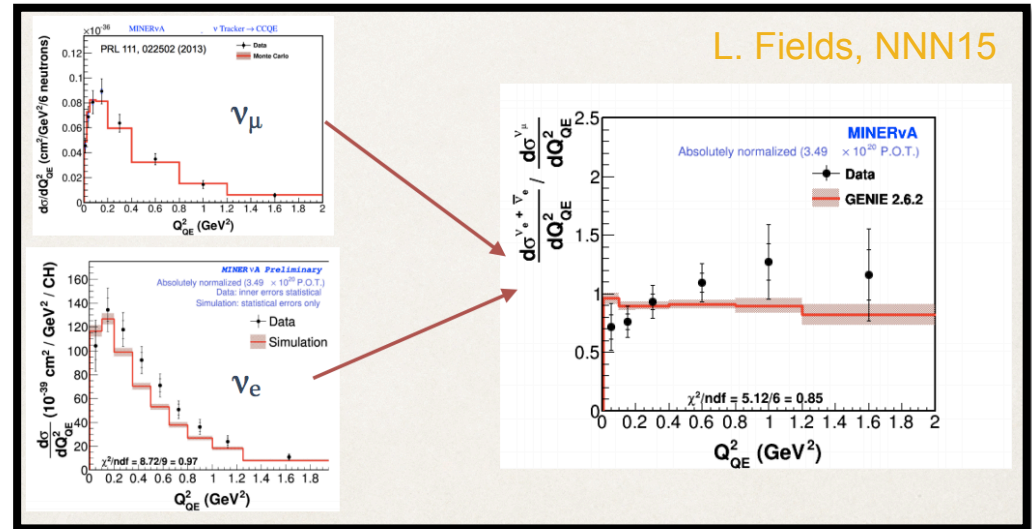
DUNE  
collaborators  
active in all of  
these efforts!

# Constraints on Cross-Section Ratios

- Theoretical and experimental constraints on variation in  $\nu/\bar{\nu}$  and  $\nu_e/\nu_\mu$  cross-section ratios determines how much four far-detector samples can constrain uncertainty from cross-section models
  - Current nominal variation for DUNE studies is 10% for  $\nu/\bar{\nu}$  and 2.5% for  $\nu_e/\nu_\mu$
  - Even a consensus on the right order of magnitude for these uncertainties would be valuable



MINERvA CCQE  
arXiv:1509.05729



# Strategy for Detector Effects

- DUNE LArTPC expected to perform better than existing appearance experiments in reconstruction of  $\nu_e$  interactions
  - Purity of quasielastic-like sample improved by detection of low-energy hadronic showers
  - Low threshold and good resolution improves calorimetric reconstruction
  - Experience from Intermediate Neutrino Program LArTPCs expected to inform simulation, reconstruction, and calibration of DUNE's far detector
- Calibration program
  - LArIAT, CAPTAIN, DUNE 35-ton prototype, protoDUNE
- Improved neutrino interaction model will reduce impact of imperfect reconstruction of neutrons and low-energy protons on analysis

DUNE 35-ton APAs:





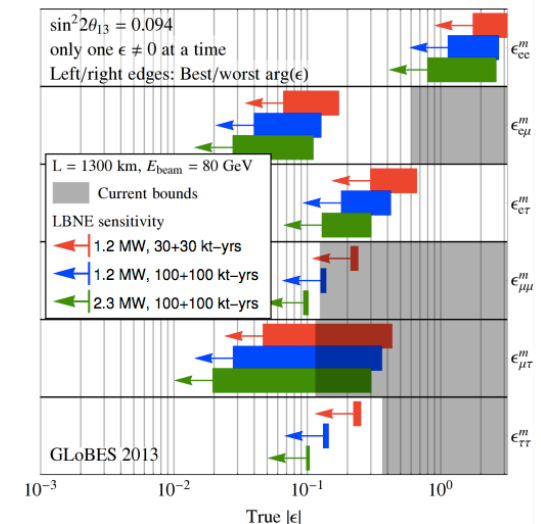
# Additional Physics Topics

- $\nu_\tau$  appearance
  - Expect >100 events per year at far detector (dependent on flux)
  - Higher energy tune would produce larger samples
- New physics affecting long-baseline oscillation
  - NSI
  - Long-range interactions
  - Large extra dimensions
  - Lorentz/CPT violation
  - Mixing with sterile  $\nu$
- Phenomenology
  - Sum rules; impact of CP violation; quantification of goals for precision parameter measurements

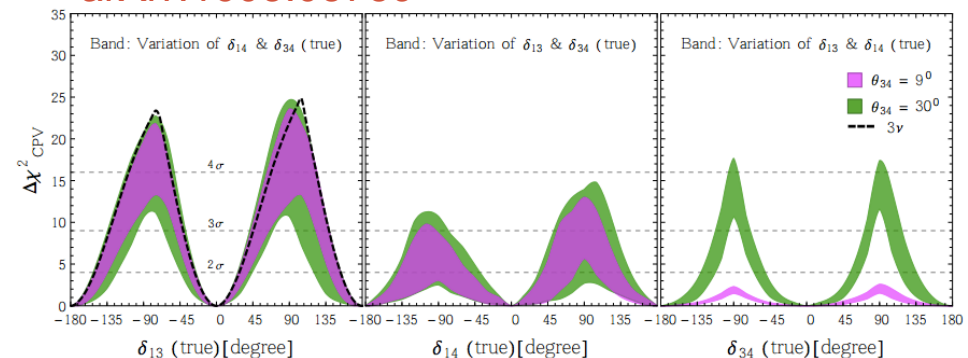
Individual effort could make huge impact on any of these topics!

arXiv:1307.7335

NC NSI discovery reach ( $3\sigma$  C.L.)



arXiv:1603.03759



# Summary



- DUNE will be sensitive to neutrino mass hierarchy, CP violation, and precision measurements of oscillation parameters  $\theta_{13}$ ,  $\theta_{23}$ ,  $\Delta m^2_{32}$ , and  $\delta_{CP}$  in a single experiment.
- DUNE will be sensitive to new physics.
- DUNE will make interesting physics measurements at every phase of experimental operations.
- Reconstruction and event selection are critical.
  - Still work in progress
- Constraint of systematic uncertainty is critical.
  - Still work in progress
  - Existing and near-future experiments are making many measurements needed to constrain DUNE systematics
- Theory/Phenomenology input is critical.
  - Improving/understanding neutrino interaction models
  - Understanding impact of precision measurements