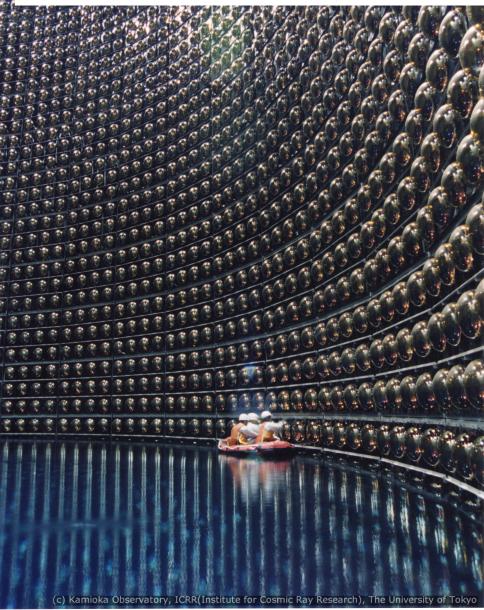
# ACCELERATOR NEUTRINO EXPERIMENTS

Louise Suter, Argonne National Laboratory

9<sup>th</sup> June 2016 Flavor Physics and CP Violation, Pasadena, CA

**FPCP** 

2015 noble prize in physics for *"the discovery of neutrino oscillations, which shows that neutrinos have mass"* 



- Takaaki Kajita for the first measurement of  $v_{\mu}$  disappearance looking at atmosphere v using Super-Kamiokande
- Arthur McDonald for leading the SNO collaboration who demonstrated that solar v were not disappearing on their way to Earth, instead they arrived at SNO with a different v flavor

#### Proved that neutrinos must have mass

New York Times, June 5th 1998 Mass Found in Elusive Particle; Universe May Never Be the Same

> Discovery on Neutrino Rattles Basic Theory About All Matter By MALCOLM W. BROWNE

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Mass mixing matrix factor $\begin{vmatrix} \nu_{e} \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix} = \begin{pmatrix} 1 \\ c_{23} & s_{23} \\ -s_{23} & c_{23} \end{pmatrix}$		$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$
$\Delta m_{32}^2 \simeq 2 \times 10^{-3}$ $L/E = 500 \text{ km/C}$	$eV^2$ $\Delta m_{31}^2 \approx \Delta m_{32}^2$ $\Delta m_{21}^2$ L/E	$\simeq 8 \times 10^{-5} \text{eV}^2$ = 15,000 km/GeV
$egin{array}{c}  u_\mu  ightarrow  u_\mu  ightarrow  u_ au \  u_\mu  ightarrow  u_ au \  atmospheric an  atmospher$	•	$ u_e  ightarrow  u_e  ightarrow  u_\mu +  u_ au$ solar and

long baseline

long baseline

reactor

 $c_{\alpha\beta} = \cos_{\alpha\beta} \ s_{\alpha\beta} = \sin_{\alpha\beta}$ 

L. Suter PCP 4 The ongoing and future accelerator neutrino program aims to answer some of the remaining big questions

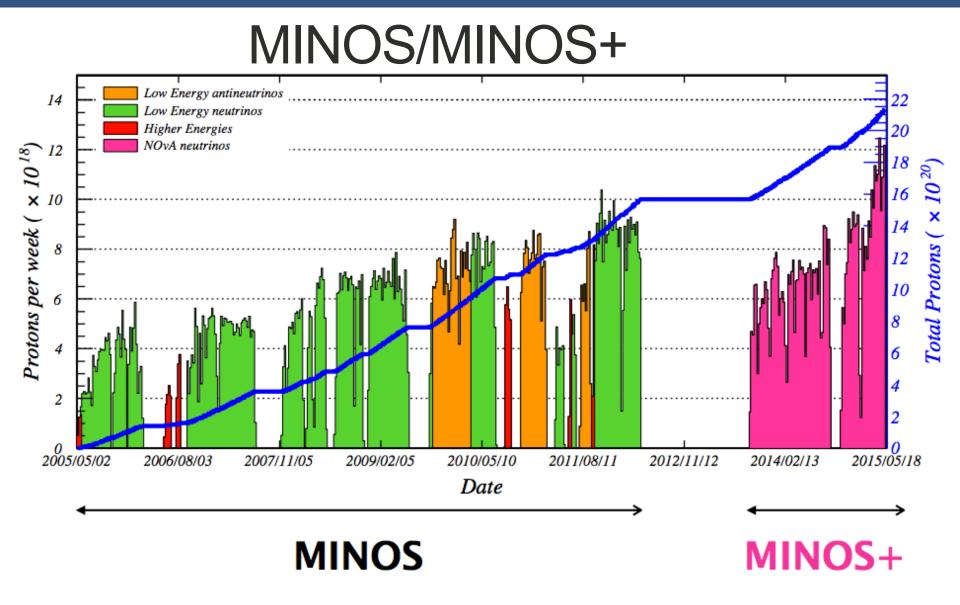
- Mass ordering?
- Nature of  $v_3$  i.e.  $\theta_{23}$  octant ?
- Is CP violated in the neutrino sector?
- Are the sterile neutrinos?

#### \_\_\_\_

#### MINOS/MINOS+

- Two detector long baseline neutrino oscillation experiment
- Near:
  - 23.7 ton fiducial mass
  - 1.04 km downstream from target
- Far:
  - 4.2 kiloton fiducial mass
  - Veto shield for cosmic suppression
  - 705m underground
- Both detectors are magnetized tracking/ sampling calorimeters, segmented into planes composed of 2.54 cm-thick steel planes and 1 cm-thick scintillator strips

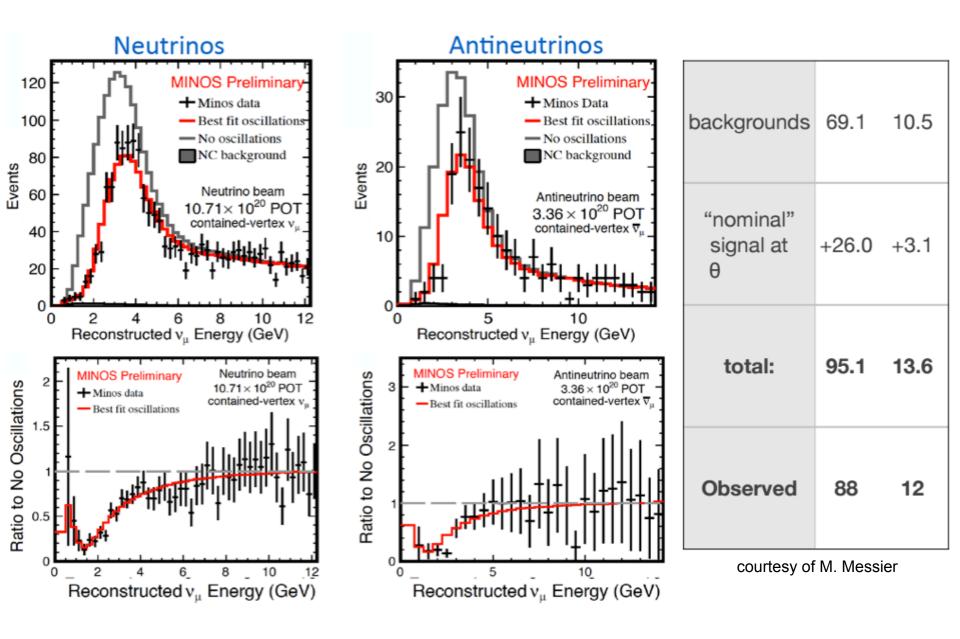




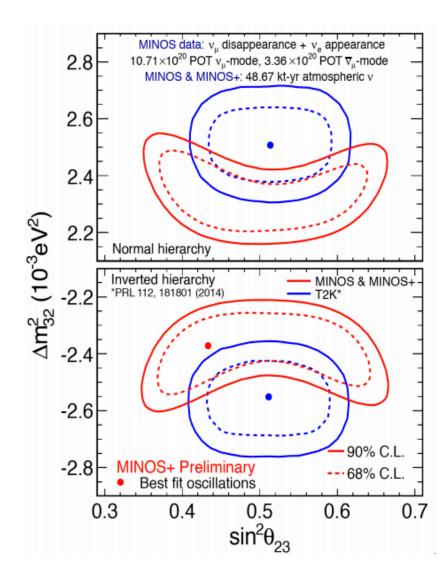
Low energy beam data 2003 to 2011

High energy beam data 2013 to 2016

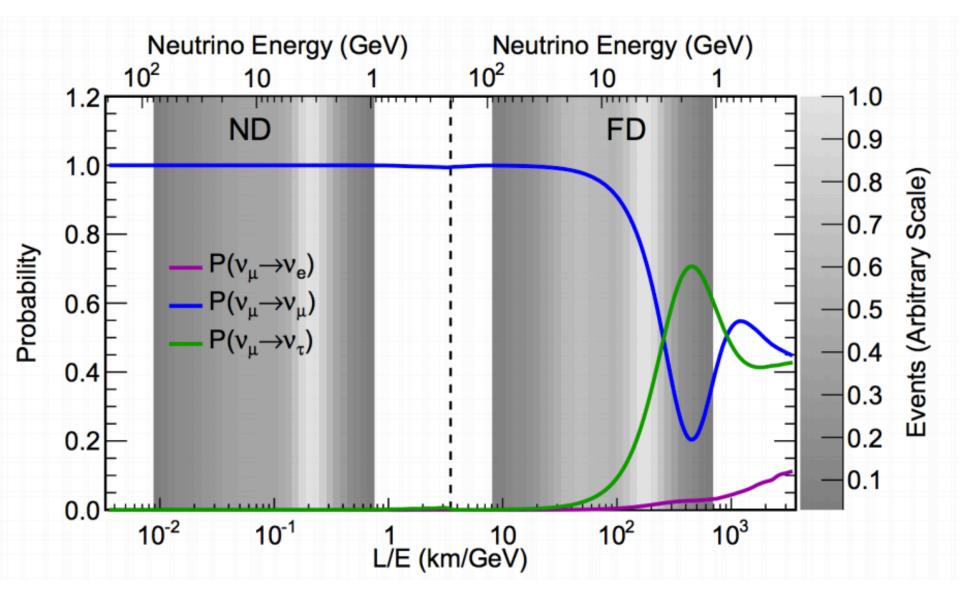
6

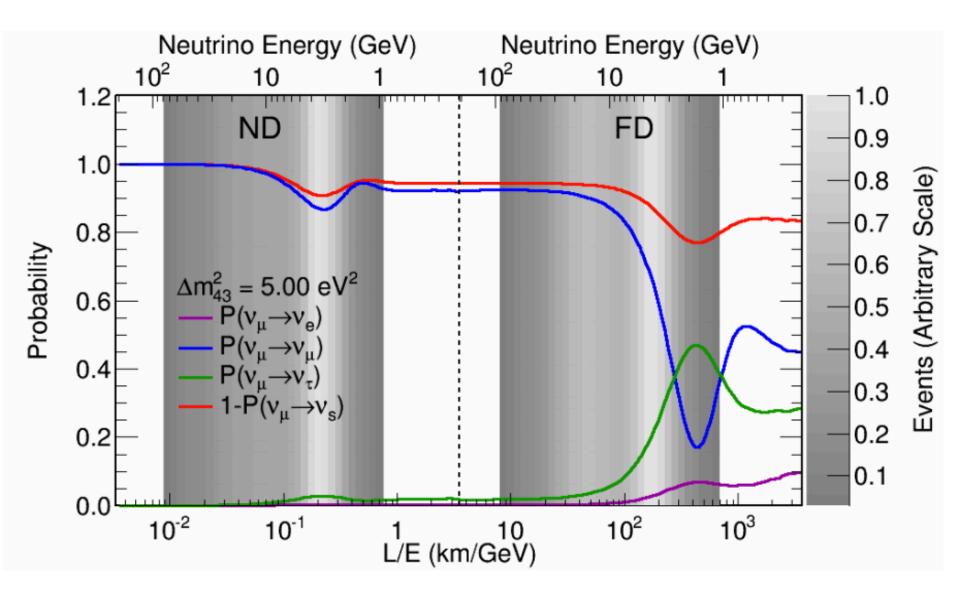


#### **Muon-Neutrino Disappearance**



Inverted Hierarchy  $\left|\Delta m^2_{32}\right| = 2.37^{+0.11}_{-0.07} \times 10^{-3} \mathrm{eV}^2$  $\sin^2 \theta_{23} = 0.43^{+0.19}_{-0.05}$  $0.36 < \sin^2 \theta_{23} < 0.65$  (90% C.L.) Normal Hierarchy  $\left|\Delta m_{32}^2\right| = 2.34^{+0.09}_{-0.09} \times 10^{-3} \text{eV}^2$  $\sin^2 \theta_{23} = 0.43^{+0.16}_{-0.04}$  $0.37 < \sin^2 \theta_{23} < 0.64$  (90% C.L.)

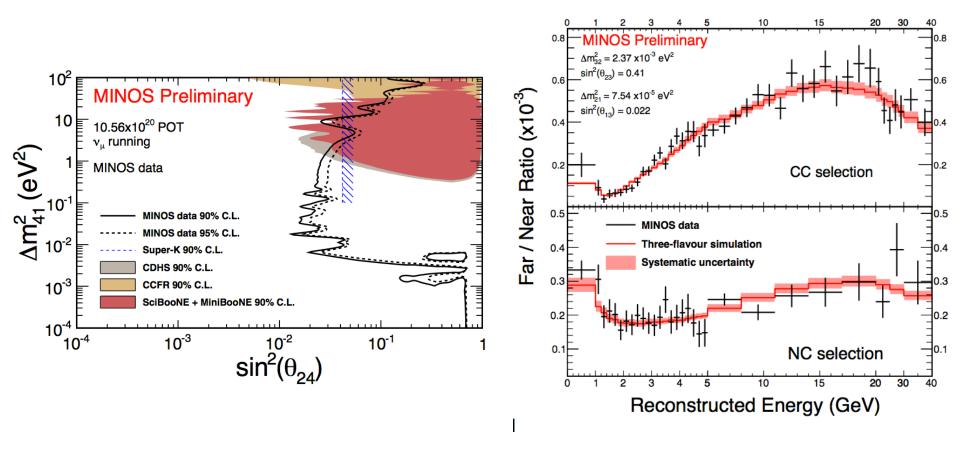




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#### Direct fit to F/N ratio for CC and NC events Set $\delta_{13}$ , $\delta_{14}$ , $\delta_{24}$ and $\theta_{14}$ to zero Parameters fit are: $\Delta m_{32}^2$ , $\Delta m_{41}^2$ , $\theta_{24}$ , $\theta_{23}$ , and $\theta_{34}$

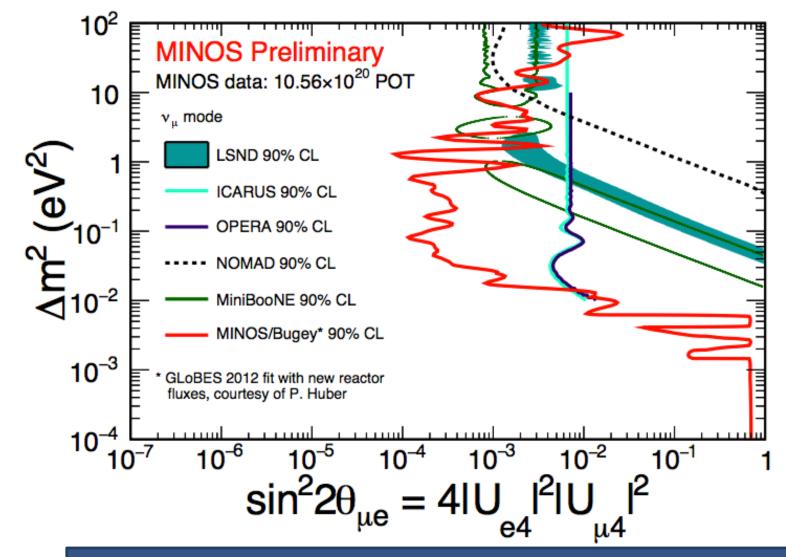


Exclude large previously unexplored region of  $\Delta m_{41}^2$ 

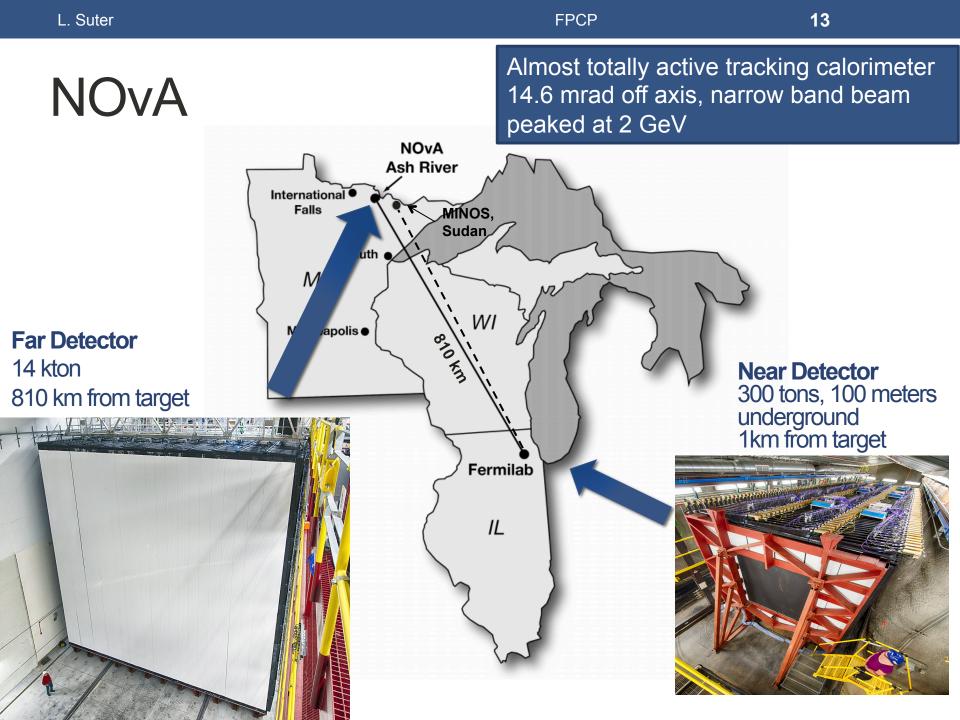
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#### Combine with Bugey to produce $sin^2 2\theta_{\mu e} \Delta m^2 (eV^2)$ limits



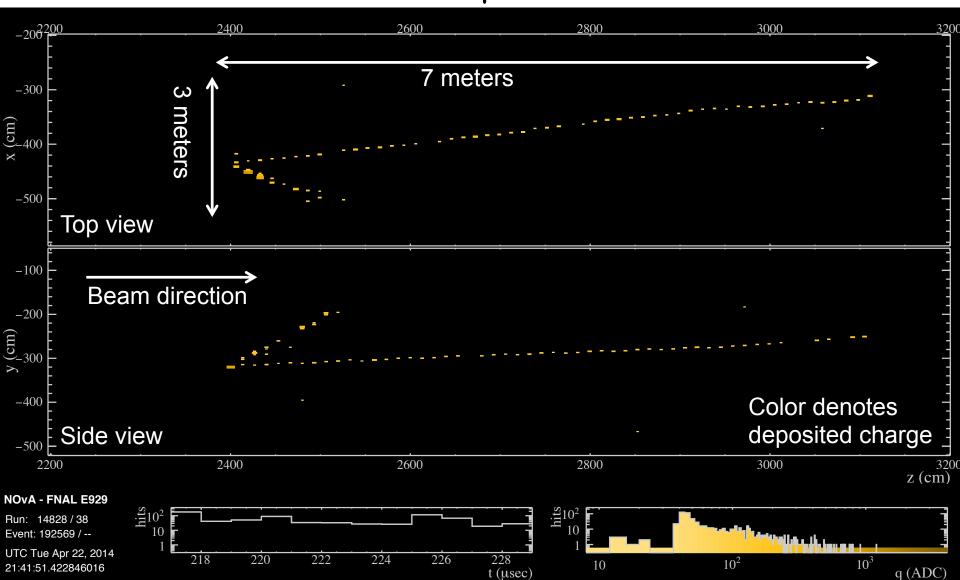
Combined MINOS & Bugey exclude most of the region allowed by LSND & MiniBooNE



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### Selected $v_{\mu}$ event

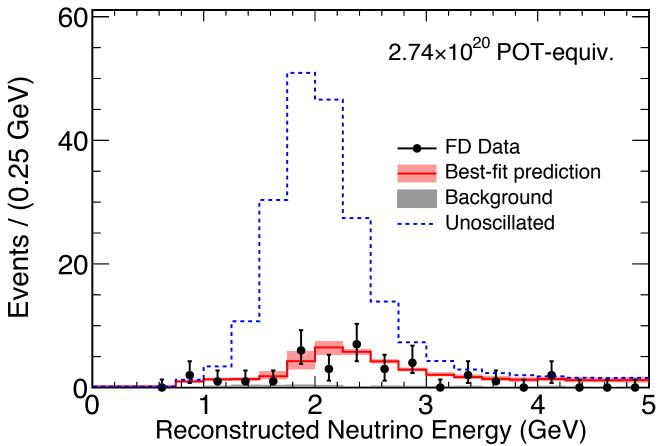


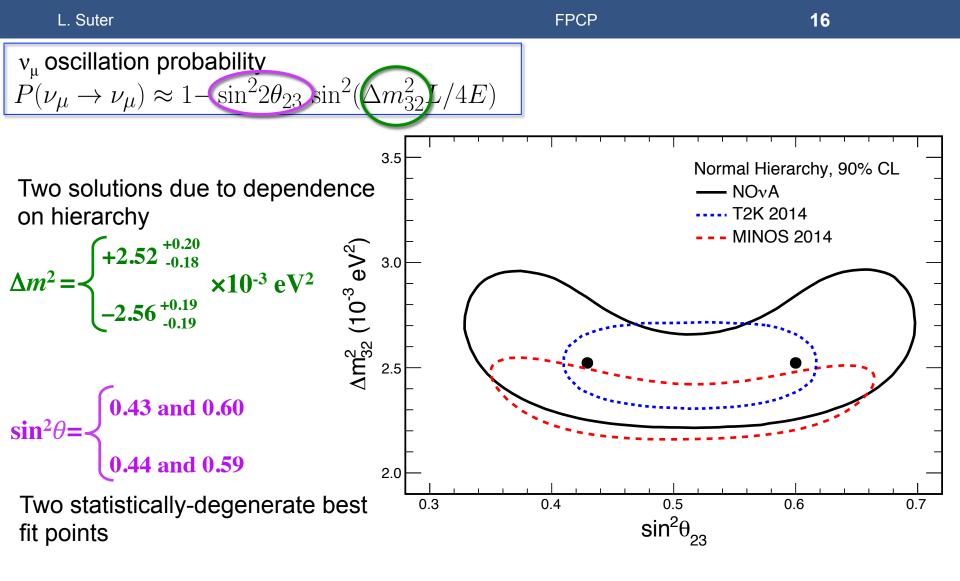
#### L. Suter

## $v_{\mu}$ Disappearance Results

- 211.8 ± 12.5 events predicted without oscillations
  - Including 2 ± 2 beam background and 1.4 ± 0.2 cosmic events
- 33 events observed!

**NOvA Preliminary** 





- Good compatibility with both MINOS and T2K
- With less than 10% of the nominal final statistics NOvA is already competitive with the world limits

## **Electron-neutrino Appearance**

Two separate electron-neutrino selection algorithms, LID and LEM Predicted 2 – 5 signal events on background of 1

	Total Bkg	Beam v <sub>e</sub>	NC	v <sub>µ</sub> CC	v <sub>r</sub> CC	Cosmic
LID	0.94 ± 0.09	0.47	0.36	0.05	0.02	0.06
LEM	1.00 ± 0.11	0.46	0.40	0.07	0.02	0.06

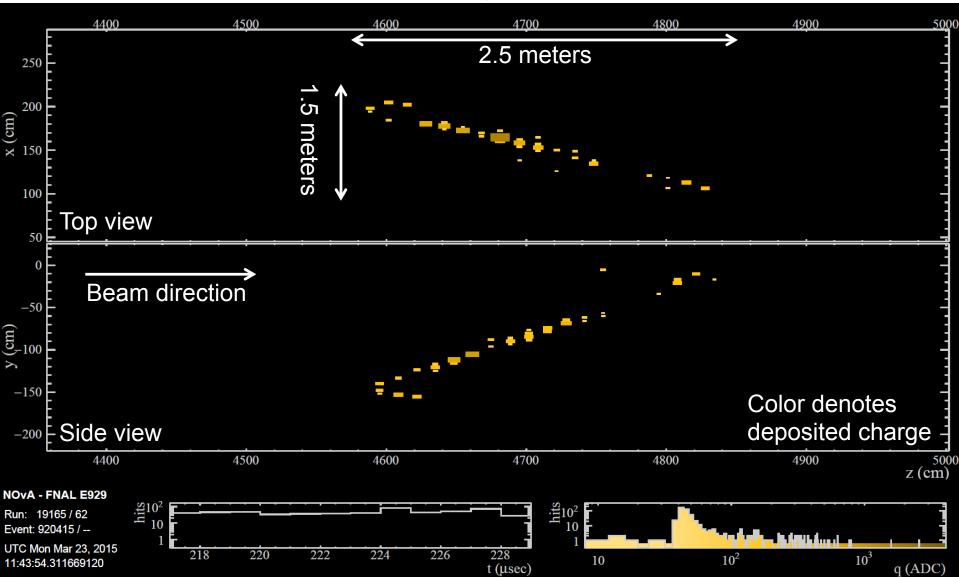
#### Signal prediction depends on oscillation parameters

Signal	Most	Least	0.09       NOvA         Contours 3 yr v and 3 yr v̄         0.08          Δm <sub>32</sub> 2  = 2.32 10 <sup>-3</sup> eV <sup>2</sup> sin <sup>2</sup> <sub>2</sub> (2θ <sub>13</sub> ) = 0.095		
	NH δ <sub>CP</sub> =3π/2	IH <b>δ</b> <sub>CP</sub> =π/2 _	Inverted hierarchy sin <sup>2</sup> (20 <sub>23</sub> ) = 1.00		
LID	5.62 ± 0.72	2.24 ± 0.29			
LEM	-5.91 ± 0.59	2.34 ± 0.23	I> <sup>©</sup> 0.04		
$0.03$ Normal hierarchy $0.01$ $0.02$ $0.04$ $0.06$ $0.08$ $v_e \text{ appearance } (v_{\mu} \rightarrow v_{\mu})$					

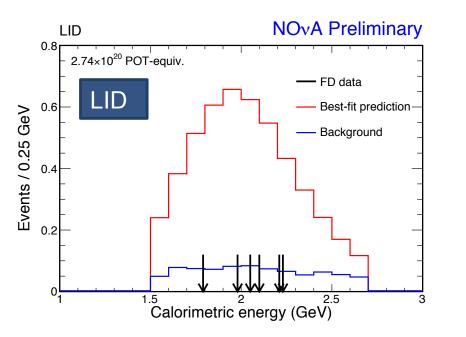
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### Selected $v_e$ event



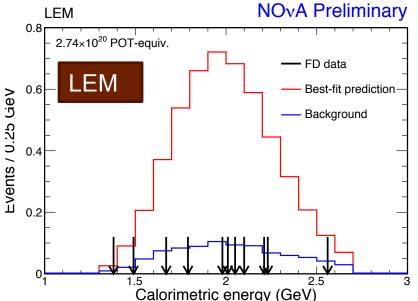
### $v_e$ Appearance Results



LID: Selected 6 events 3.3 $\sigma$  significance for v<sub>e</sub> appearance

LEM: Select 11 events 5.5 $\sigma$  significance for  $v_e$  appearance

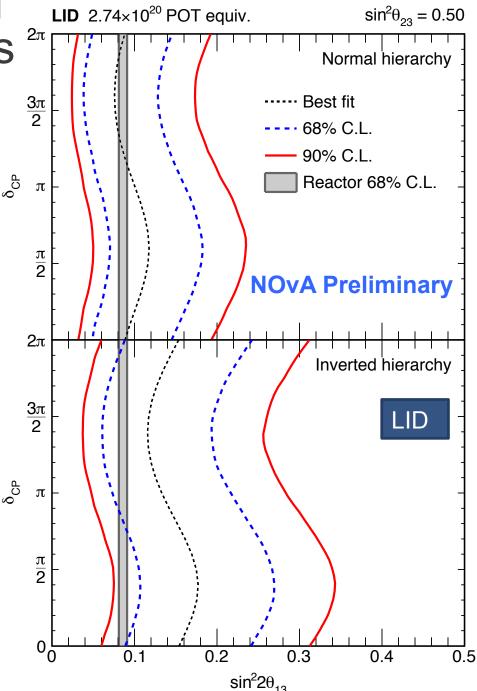
- All 6 LID events selected by LEM
- Trinomial probability of selecting this combination (11:6/5/0) is 9.2%



# $\nu_e$ Appearance Results ${\scriptstyle \text{LID}}$

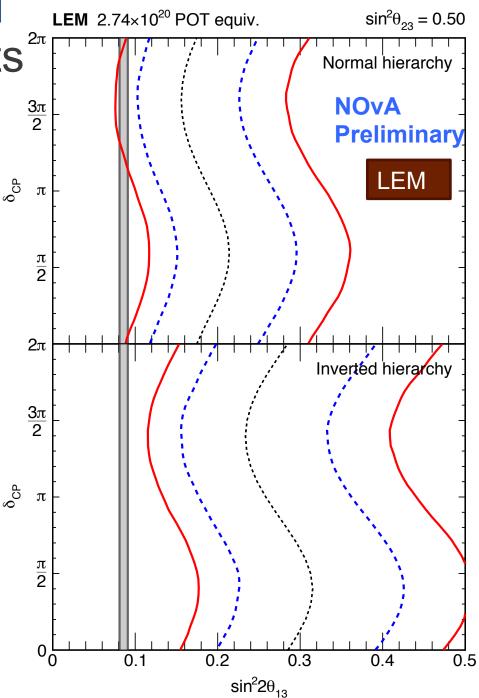
- Contours determined using Feldman-Cousins procedure <sup>b</sup>
  - Include uncertainties on solar parameters
  - Atmospheric ∆m<sup>2</sup> varied within new NOvA uncertainties
  - $sin^2\theta_{23}$  held fixed at 0.5
- LID results in good agreement with reactor measurements

LID: Selected 6 events  $3.3\sigma$  significance for  $v_e$  appearance



# $v_e$ Appearance Results

- LEM curves shift to the right
- Some tension with reactor results, particularly in IH

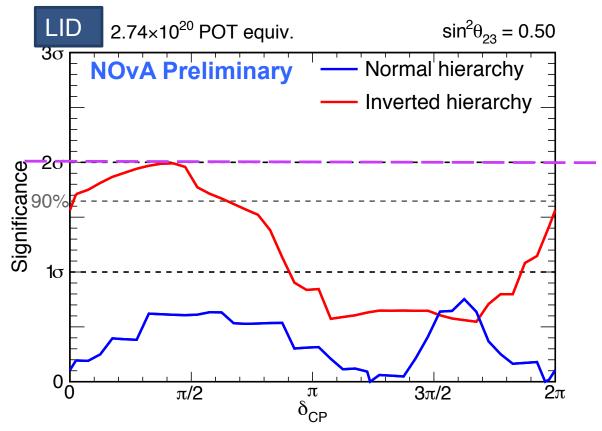


LEM: Select 11 events 5.5 $\sigma$  significance for  $v_e$  appearance

Statements holds for  $0.4 < \sin^2\theta_{23} < 0.6$ 

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- LID shows mild tension with IH, 0 <  $\delta_{\rm CP}$  < 0.8 $\pi$
- LEM disfavors IH at greater than  $2\sigma$  for all  $\delta_{\rm CP}$



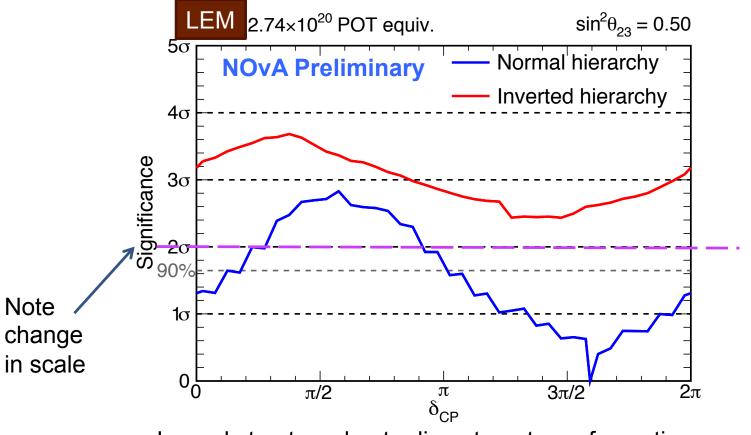
Jagged structure due to discrete nature of counting experiment

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Statements holds for  $0.4 < \sin^2\theta_{23} < 0.6$ 

23

- LID shows mild tension with IH, 0 <  $\delta_{\rm CP}$  < 0.8 $\pi$
- LEM disfavors IH at greater than  $2\sigma$  for all  $\delta_{\text{CP}}$



Jagged structure due to discrete nature of counting experiment

**FPCP** 

#### L. Suter

Far detector,

295 km from

source

Mt.lkenoyama

1.360m

Super-Kamiokande

#### FPC

### T<sub>2</sub>K

- T2K is a long-baseline neutrino experiment with a 600 MeV narrow band muon neutrino beam
- Detectors 2.5° off axis from neutrino beam

Mt.Noguchi-Goro Dake

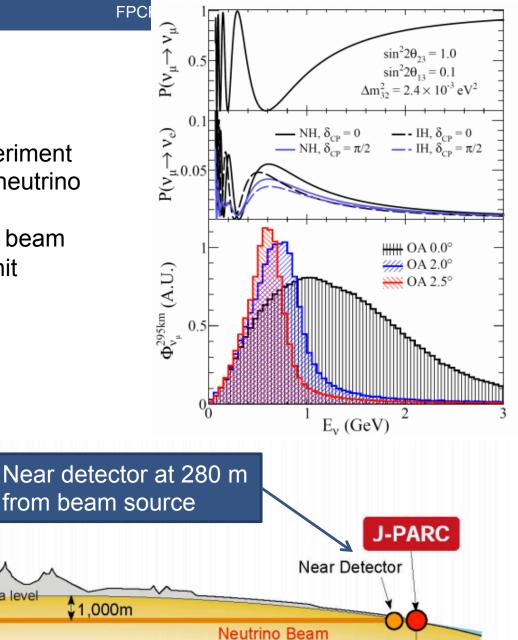
sea level

1,000m

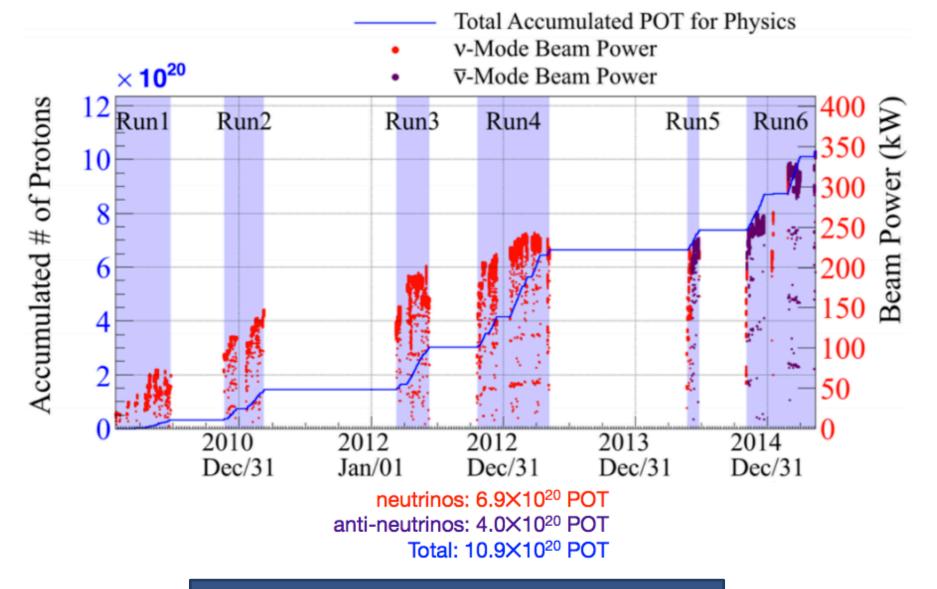
295km

2,924m

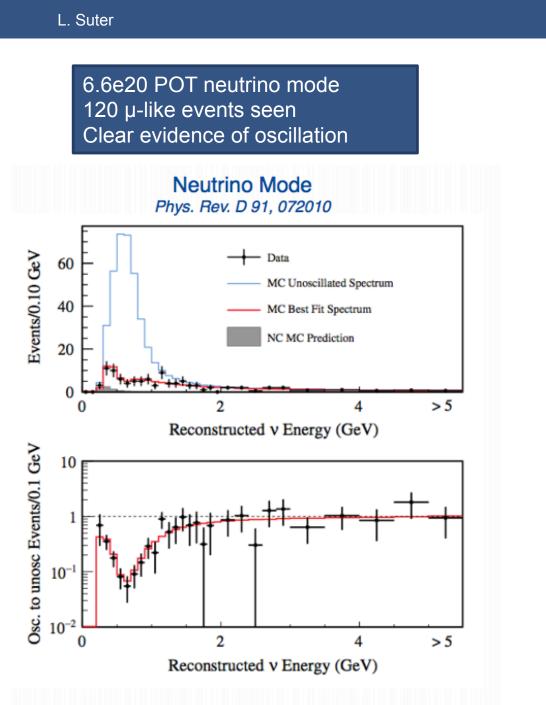
Neutrino energy spectrum tuned to hit oscillation maximum at far detector

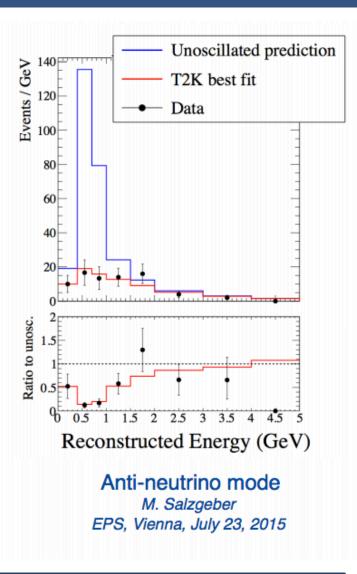






Been running in anti-neutrino mode since 2014

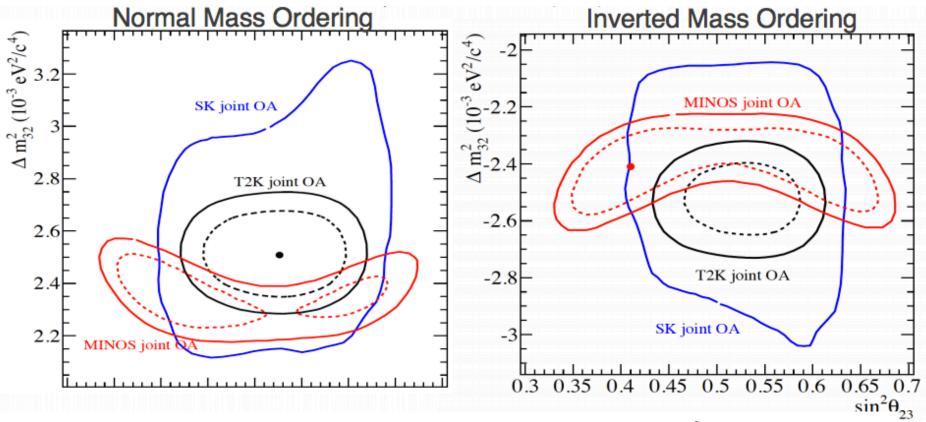




4e20 POT in antineutrino mode 34 μ-like events seen Clear evidence of oscillation

**FPCP** 

**FPCP** 



- T2K best fits: slight (<1σ) preference for Normal Ordering, sin<sup>2</sup>θ<sub>23</sub>=consistent with maximal
- MINOS best fits: slight preference for **Inverted Ordering**. Negligible preference for lower  $\theta_{23}$  octant

Without reactor constraints:

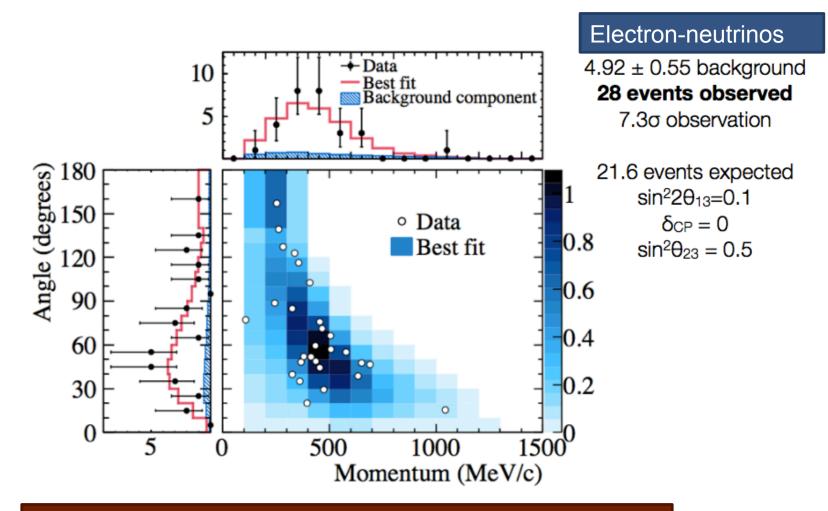
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Normal hierarchy:  $\sin^2 \theta_{23} = 0.514^{+0.055}_{-0.056}$ Inverted hierarchy:  $\sin^2 \theta_{23} = 0.511 \pm 0.055$ 

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#### **Electron-Neutrino and Anti-Neutrino Appearance**

**FPCP** 

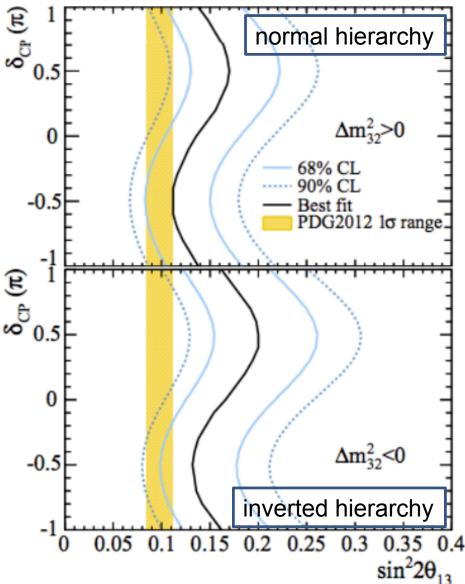


Anti-Neutrino: 3 events on a background of 1.8 seen Update expected this summer

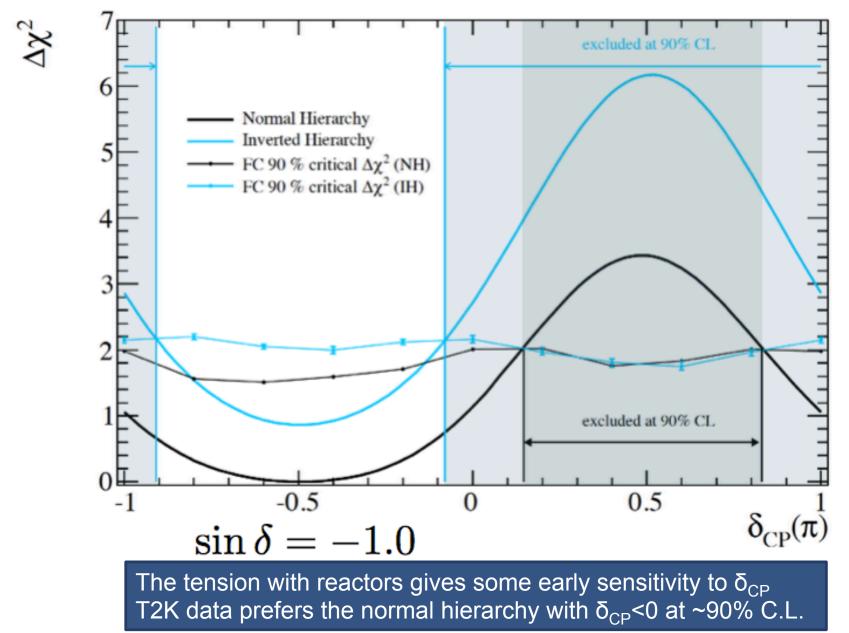
courtesy of M. Messier

# **Electron-Neutrino Appearance**

- T2K sin<sup>2</sup>2θ<sub>13</sub> result computed assuming sin<sup>2</sup>θ<sub>23</sub>=0.5
- Consistent at 90% CL (1.6σ)
- Excess seen by T2K nudges all remaining unknowns in direction to increase rates
  - normal hierarchy
  - θ<sub>23</sub> > 45°
  - δ<sub>CP</sub>= -π/2 (aka 3π/2)



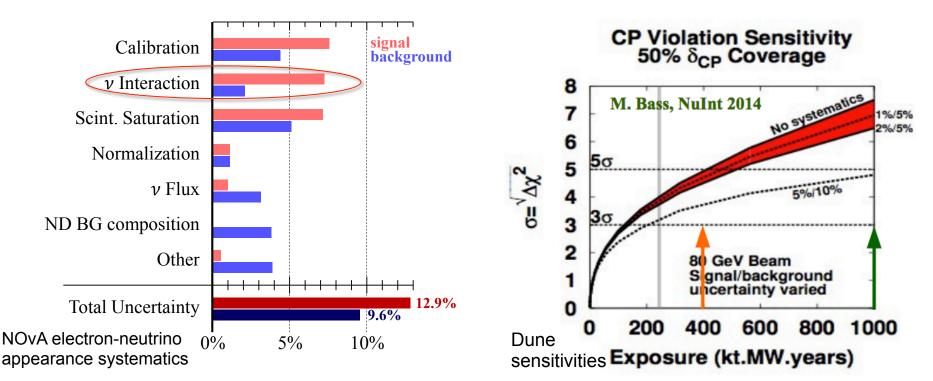
29



courtesy of M. Messier

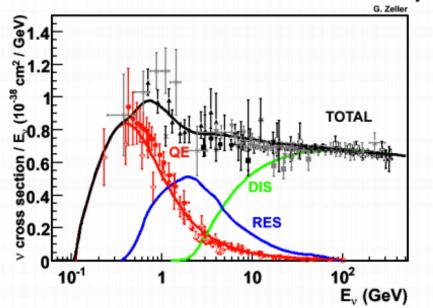
### **MINERvA**

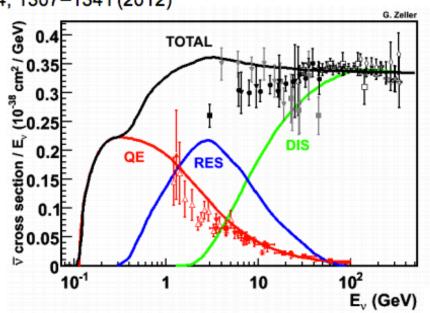
- Study neutrino-nucleus scattering at a few GeV, using Fermilabs NuMI 120 GeV beam
- Measure the effects of the nuclear environment on neutrino scattering
- Improve understanding of neutrino-nucleus cross section model by working with generators
- Benefits current and future neutrino oscillation experiments



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Quasielastic (QE)	Resonance	DIS
vµ	v	v
W	W	W
	$\Delta$ $\pi$	
$n \sim p$	$N \sim N$	<i>q</i> ~ ~

Rev. Mod. Phys. 84, 1307–1341 (2012)

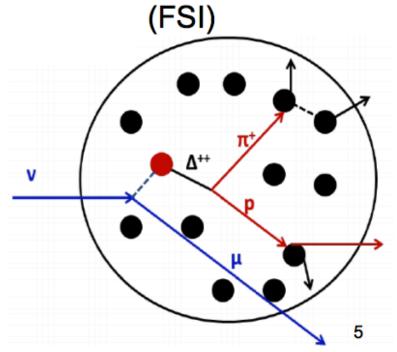




courtesy of D Ruterbories

Detector built with heavy nuclei for high interaction rate but introduces complications

Struck nucleon is not isolated short and long range correlations exist Particles created have to propagate out of the nucleus. So called final state interactions



### **MINERvA** ongoing analyses

- Not-yet-published MINERvA analyses in Low Energy Beam
- Nuclear effects at low-momentum transfer (arXiv: 1511.05944)
- Electron neutríno charged-current quasí-elastic scattering (arXív:1509.05729)
- DIS on nuclear target (Carbon, Iron and Lead)
- Neutral current diffractive pio production
- Neutrino-electron scattering
- Measurements of muon kinematics in neutrino induced charged pion production and antineutrino induced neutral pion production
- Double differential muon neutrino and antineutrino quasi-elastic scattering
- Neutrino and antineutrino flux and inclusive cross sections using the "low-nu" technique Erom Lawra Fields
- 3 Kaon production analyses (charged current, neutral current and coherent)
- neutríno charged current neutral píon production
- Antineutrino charged-current inclusive scattering on Carbon, Iron and Lead
- Neutríno charged-current inclusive scattering on Helium

Highest Priority Medium Energy Beam analyses

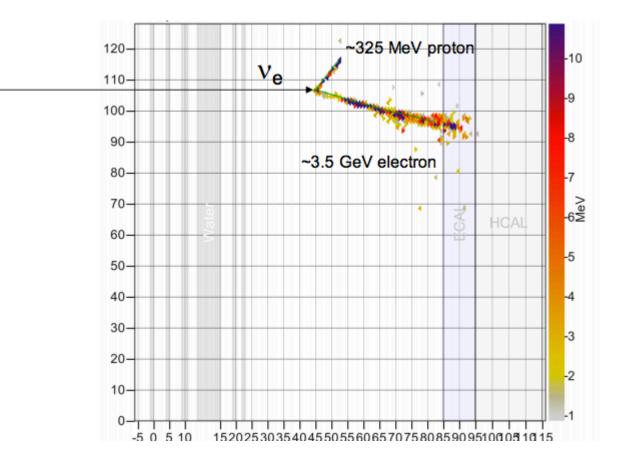
- Neutrino electron scattering on scintillator
- Neutrino DIS on scintillator, iron, carbon & lead
- CCQE on scintillator, iron, carbon & lead
- Pion production on scintillator, iron, carbon & lead
- Coherent pion production on scintillator, iron, carbon & lead
- Nuclear structure functions

Keep them coming.....

**FPCP** 

## CC QE via $v_e$ -nucleus scattering

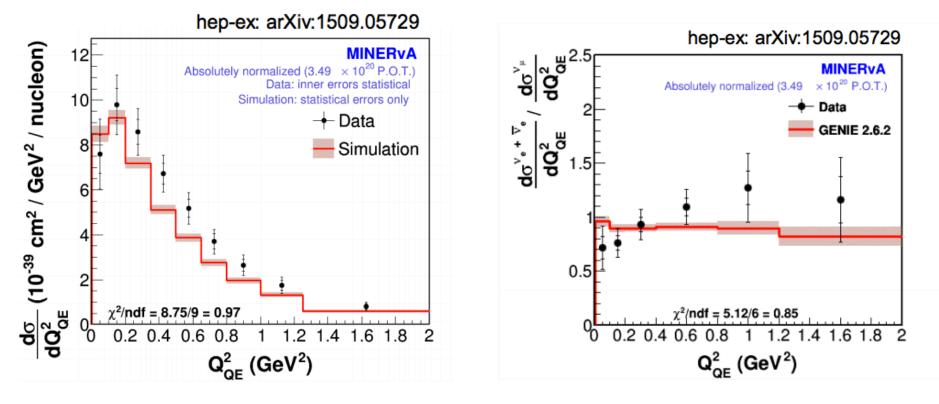
Dominant neutrino oscillation  $v_e$ -appearance signal process



courtesy of D Ruterbories

#### $\nu_{\rm e}$ CC QE differential cross section

## $\nu_{e}$ to $\nu_{\mu}$ differential cross section ratio



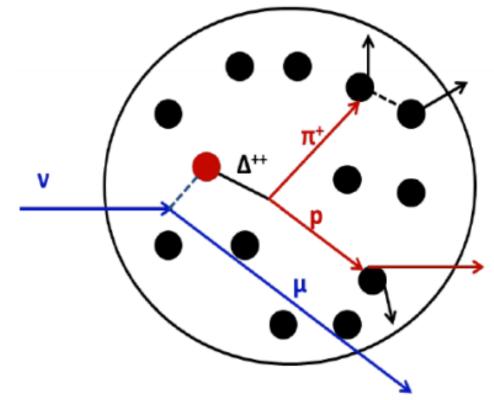
• Ratio is consistent with 1.0, but shape is not significant

Large correlated errors

**FPCP** 

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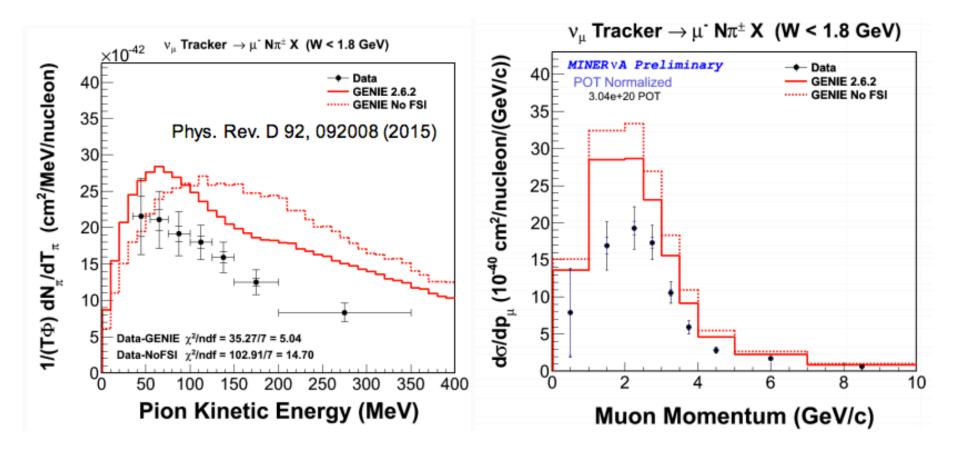
- Can investigate final state interactions by measuring the pion spectra
- Can investigate nuclear modifications by measuring Q<sup>2</sup> which is insensitive to pion FSI since you measure via the outgoing lepton



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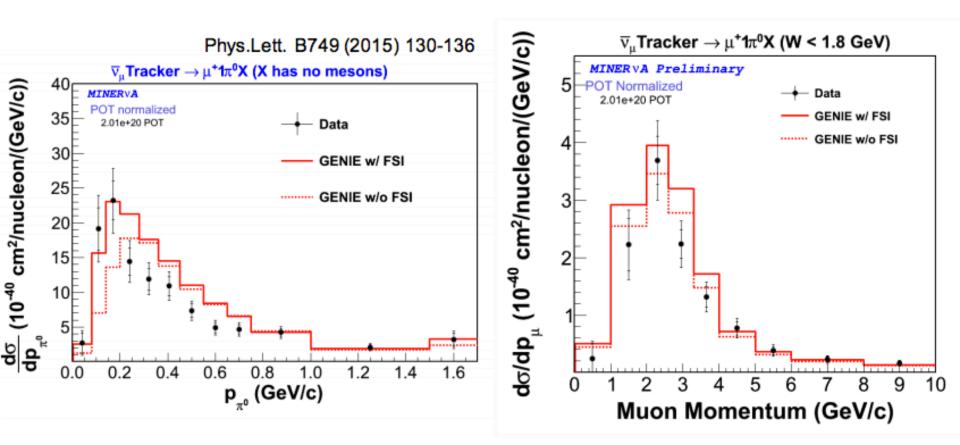
**FPCP** 

### **Charged Pion Production**

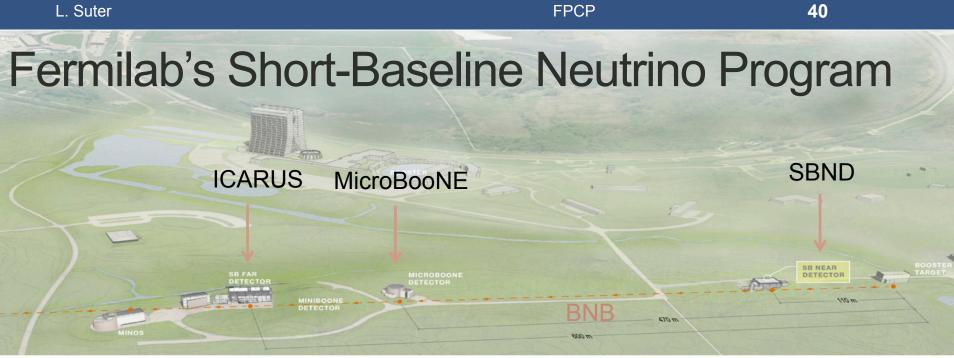


- Shape agrees, but cross section is over predicted
- FSI reduces the cross section due to pion absorption

### Neutral pion production



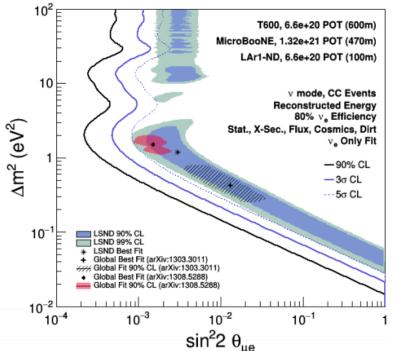
- Shape agrees
- FSI increases the cross section due to charge exchange



#### Three LArTPC detector program

- 1) MicroBooNE: currently taking data
- Near Detector SBND: Construction to start ~now, first data in 2018
- 3) Far Detector ICARUS@FNAL: Construction started 2015, first data in 2018

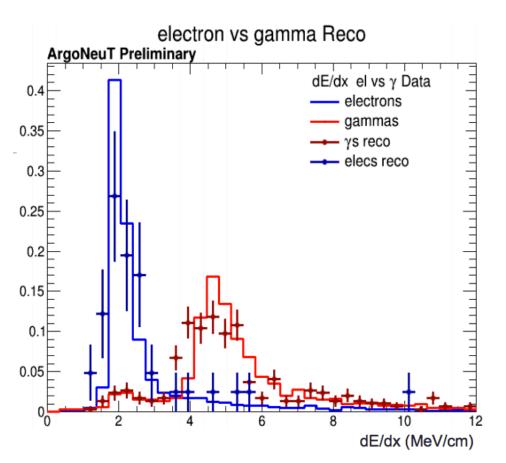
Designed to answer the question of sterile neutrinos and test bed for LArTPC technology

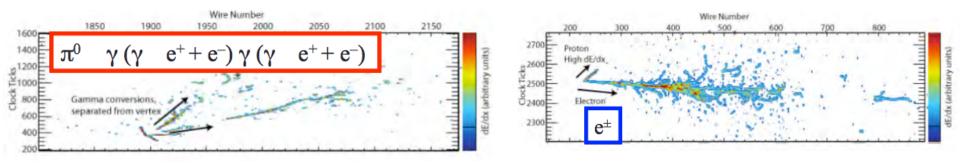


### **MircoBooNE**



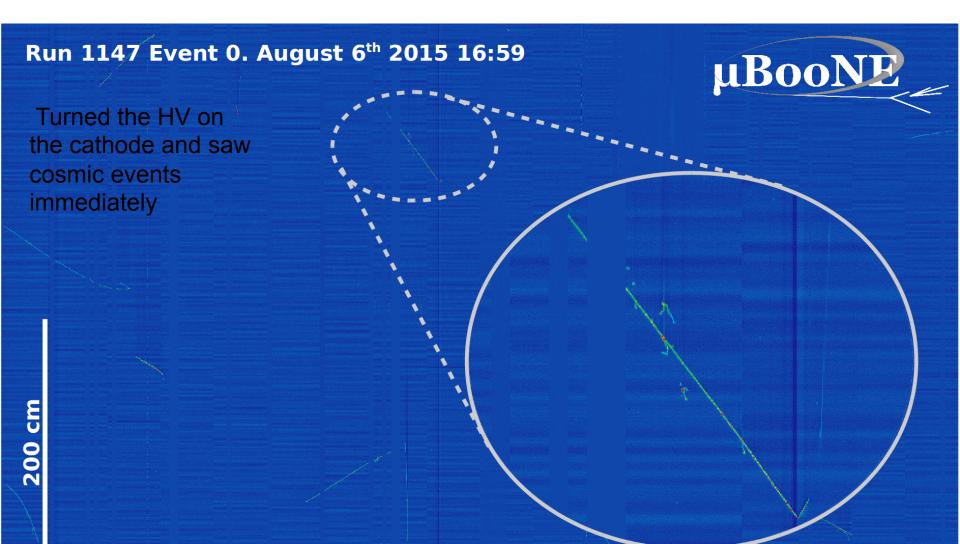
- LArTPC allow for precision electron-photon separation
- MicroBooNE aims to answer question of whether excess seen by MiniBooNE is electrons or photons
- Expect millions of events per year, active cross sections group

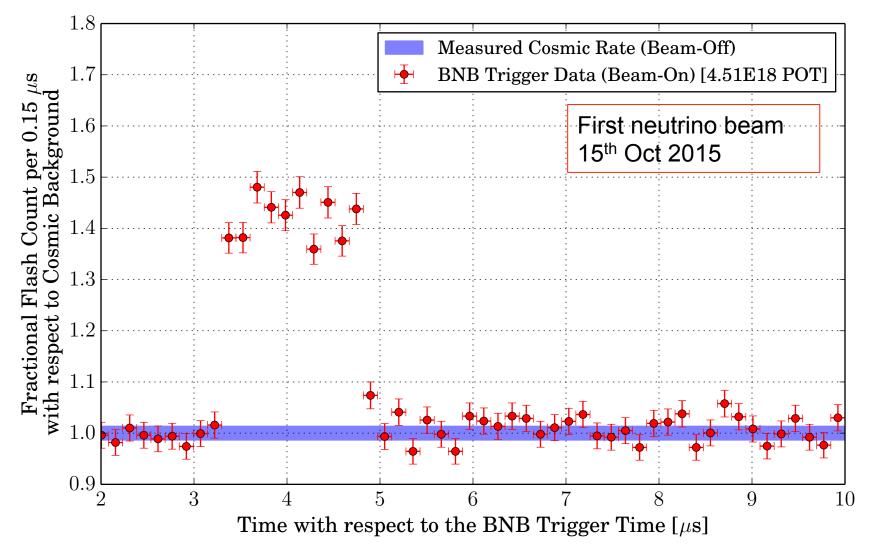




200 cm

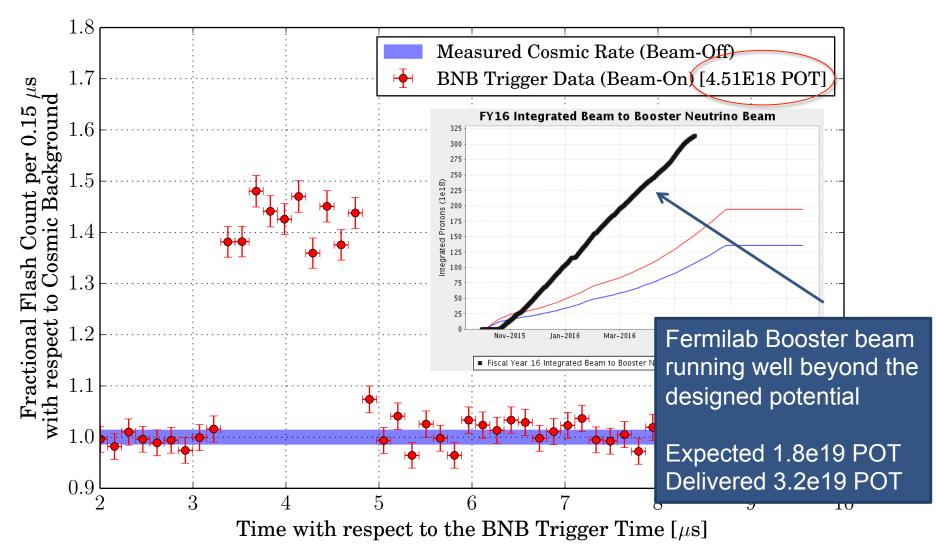
43





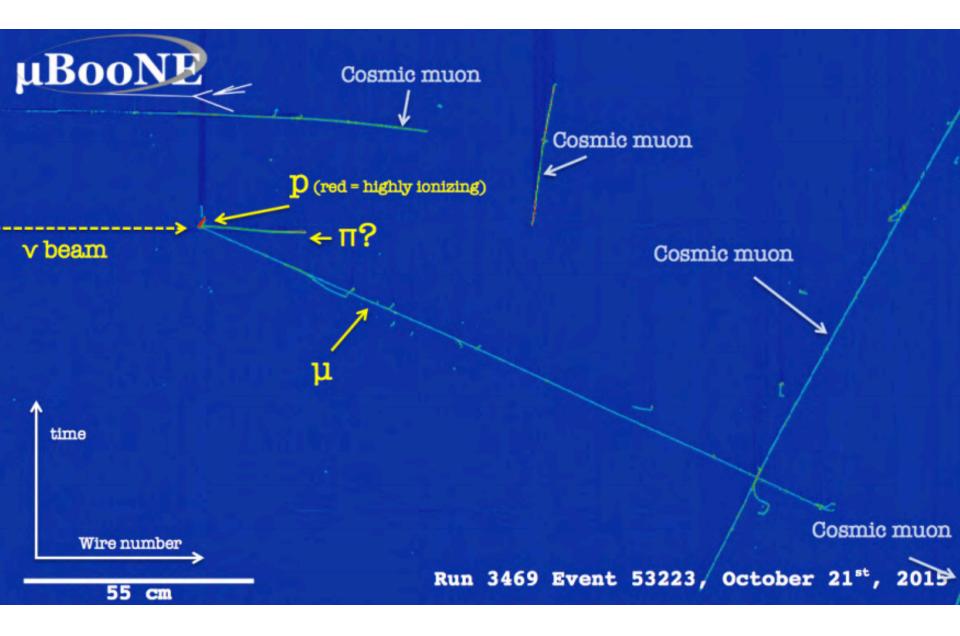
Detector on surface large cosmic ray background

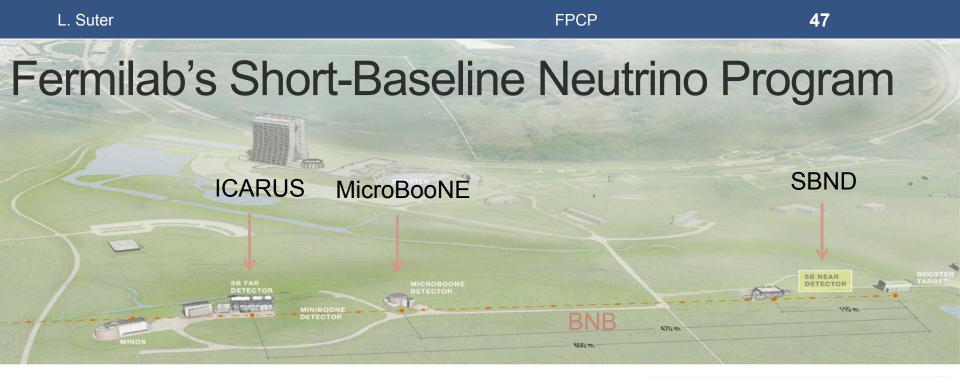
Events identified using reconstruction from both the wire readout and the light collection system



Detector on surface large cosmic ray background

Events identified using reconstruction from both the wire readout and the light collection system

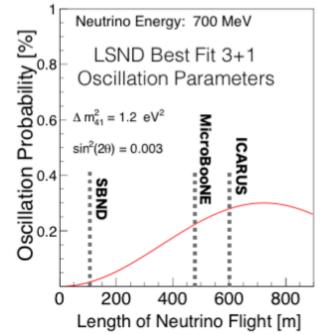




#### Three LArTPC detector program

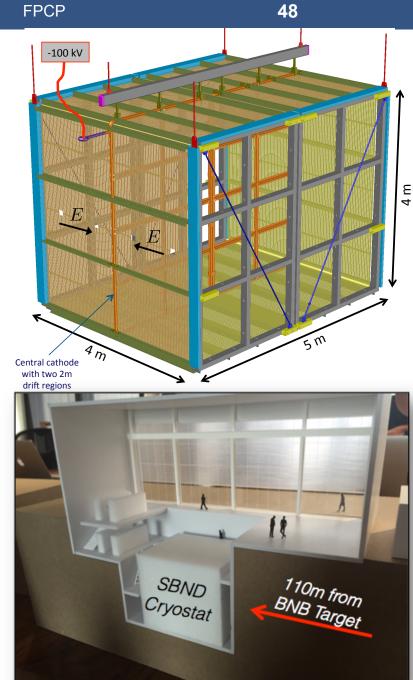
- 1) MicroBooNE: currently taking data
- Near Detector SBND: Construction to start ~now, first data in 2017
- 3) Far Detector ICARUS@FNAL: Construction started 2015, first data in 2017

Designed to answer the question of sterile neutrinos and test bed for LArTPC technology

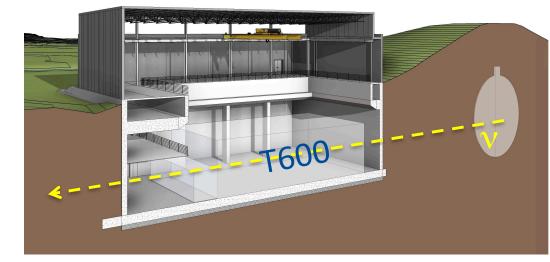


- Short baseline program with long baseline design
  - Near detector to provide a detailed characterization of the beam before oscillation
  - Allows for the cancelation of many of the dominant systematics
- Expects > 1.5 million neutrino events per year to able to perform high precision cross section measurements
- First data 2018

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## ICURAS @ FNAL



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- SBN Far Detector uses the largest LArTPC in the world
  - 6x larger than MicroBooNE
- Been running at Gran Sasso since 2010, was moved to CERN in 2014 to be refurbished
- Will be moved to FNAL 2017
- First data 2018



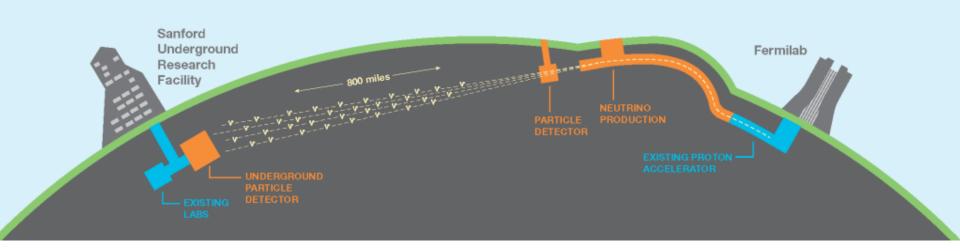
L. Suter

- Measurements using atmospheric neutrinos, reactor neutrinos and long-baseline neutrinos form a consistent picture
  - Large  $\theta_{23}$  (0.4 < sin<sup>2</sup> $\theta_{23}$  < 0.6)
  - Precisely known  $\theta_{13} = 8.4^{\circ}$
- Consistent hints favoring
  - π<δ<sub>CP</sub><2π</li>
  - Normal mass ordering
- First data from NOvA strengthens this picture
- New results from NOvA, T2k, and MINOS+ this summer, will be very exciting times
- New results from MINERvA and MircoBooNE will strengthen our understanding of neutrino interactions and our other measurements

# Backup

## DUNE

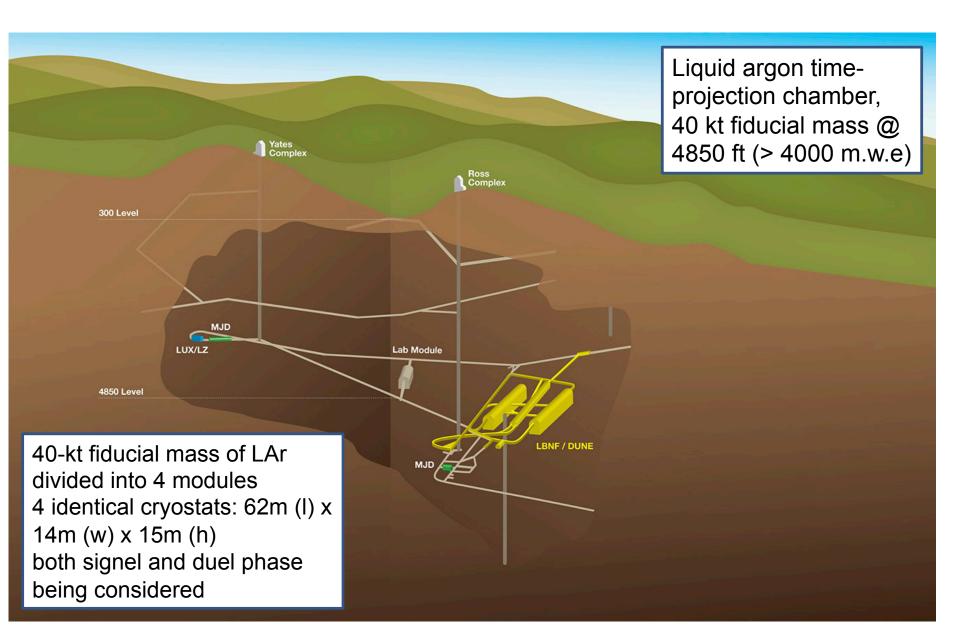
Wide-band on-axis neutrino beam produced by the primary 60-120 GeV proton beam. Initial power 1.2 MW, upgradable to 2.4 MW

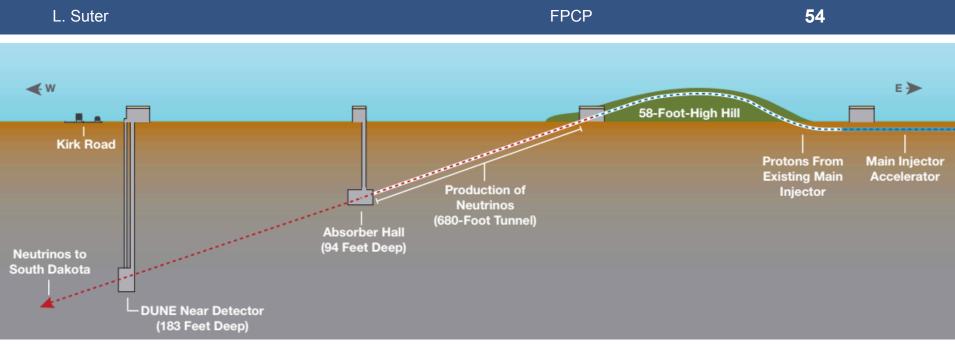


1300 km (800 mile) baseline, 1 st (2nd) oscillation max. @ ~2.4 (0.8) GeV

First data 2024, first beam 2026

Fine-grained near detector complex, enabling unprecedented studies of neutrino interactions



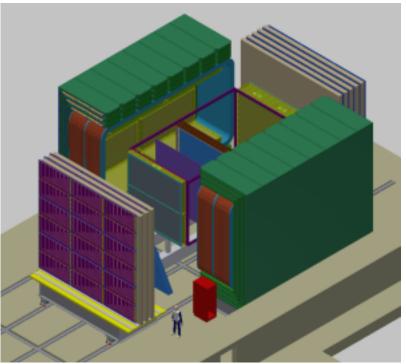


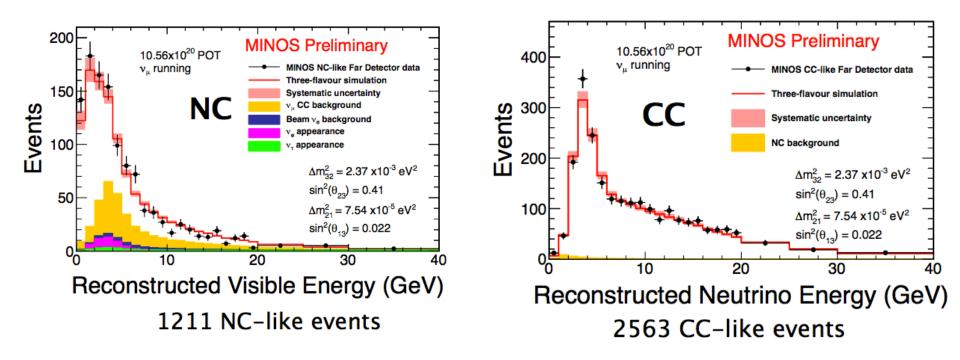
central straw-tube tracking (STT) system with embedded nuclear targets (CH, C, Ar, ...)

 $4\pi$  electromagnetic calorimeters (ECAL) dipole 0.4-T magnet surrounding STT and ECAL

 $4\pi$  muon identifiers

Additional detectors under study: magnetized LAr TPC and high-pressure gas TPC > 100 million events per year for a wide range of energies





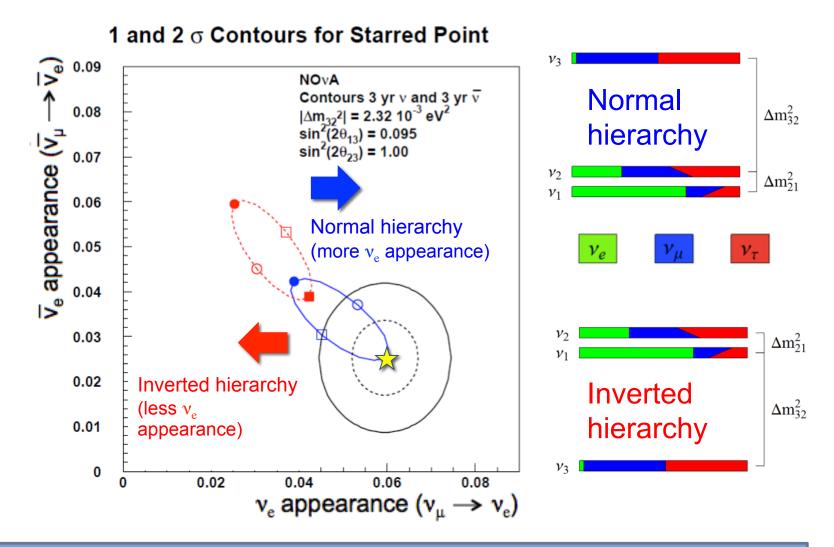
#### FD spectra three-flavor oscillated with MINOS 2012 CC-analysis fit values

$$R = \frac{N_{data} - \sum B_{CC}}{S_{NC}} \xrightarrow{\text{Predicted CC}}_{\text{from all flavors}} \\ S_{NC} \xrightarrow{\text{Predicted NC}}_{\text{interaction signal}} \\ \end{array}$$

R [0-3 GeV] = 1.10 + /- 0.06 + /- 0.07R [0-40 GeV] = 1.05 + /- 0.04 + /- 0.10

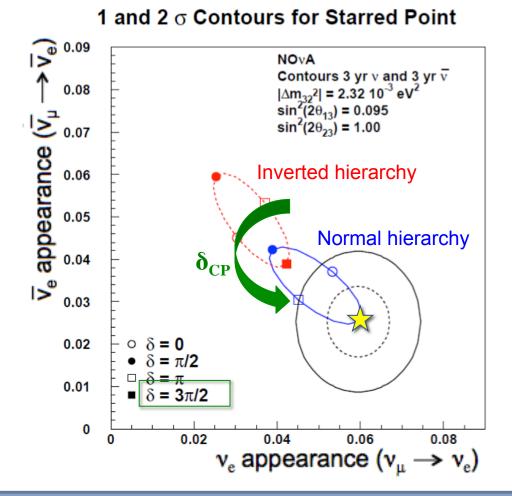
#### FP<u>CP</u>

### Mass hierarchy: Electron-neutrino appearance



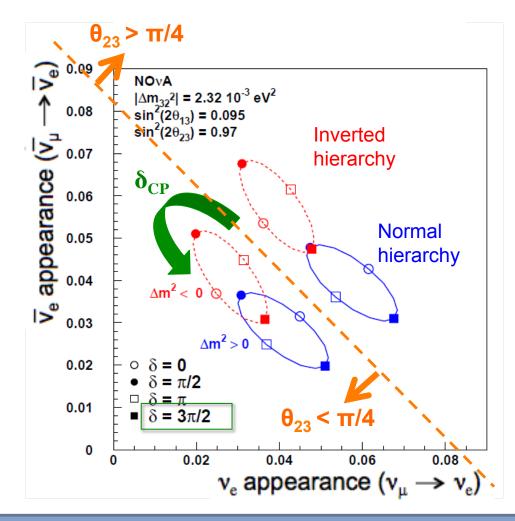
Number of electron neutrinos appearing depends on ordering of the mass states

### CP violation: Electron-neutrino appearance



Number of electron neutrinos appearing depends on **amount of CP violation in neutrino sector** 

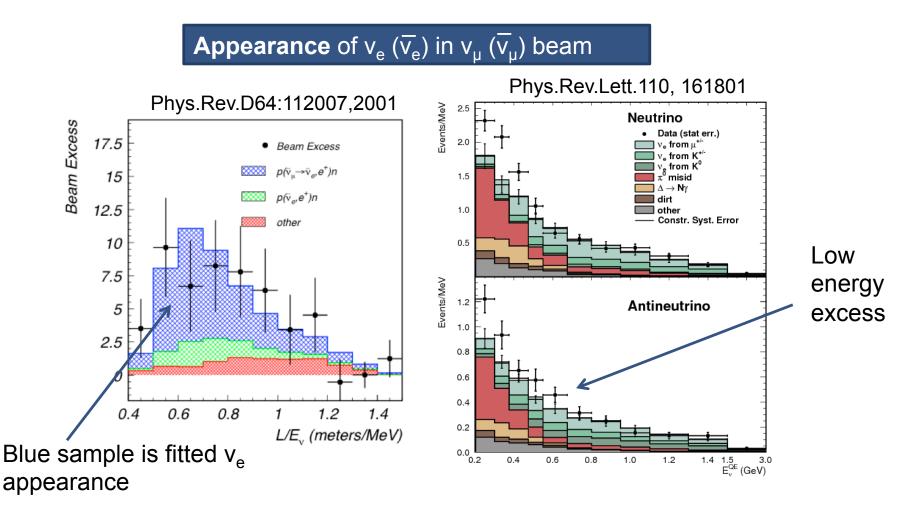
## $\theta_{23}$ octant : Electron-neutrino appearance



Number of electron neutrinos appearing depends on  $\theta_{23}$  octant

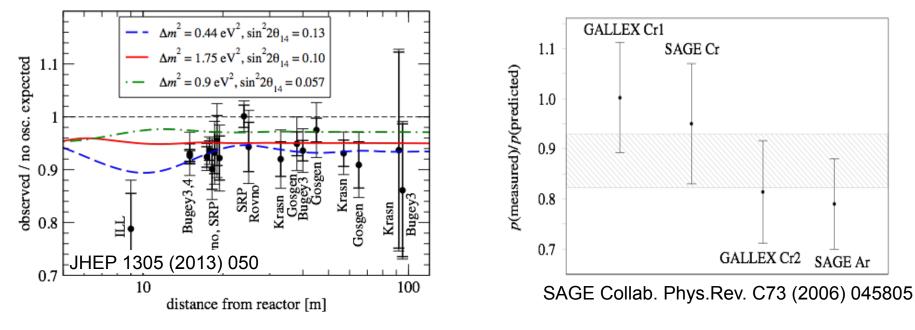
## **Sterile Neutrinos Hints**

LSND and MiniBooNE have experimental results which could be interpreted as due to a new neutrino with a mass ~1 eV



## **Sterile Neutrinos Hints**

Anomalies from rector neutrino experiments and rate measurements of radioactive sources on Gallium could also be interpreted in this context



Experiment	Type	Channel	Significance
LSND	DAR	$\bar{\nu}_{\mu} \to \bar{\nu}_e \ \mathrm{CC}$	$3.8\sigma$
MiniBooNE	SBL accelerator	$\nu_{\mu} \rightarrow \nu_{e} \ \mathrm{CC}$	$3.4\sigma$
MiniBooNE	SBL accelerator	$\bar{\nu}_{\mu} \to \bar{\nu}_e \ \mathrm{CC}$	$2.8\sigma$
GALLEX/SAGE	Source - e capture	$\nu_e$ disappearance	$2.8\sigma$
Reactors	Beta-decay	$\bar{\nu}_e$ disappearance	$3.0\sigma$

K. N. Abazajian et al. (arXiv:1204.5379 [hep-ph])