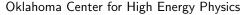
DIMENSION 7 NEUTRINO MASS GENERATION AND ITS IMPLICATIONS FOR THE LHC AND THE DARK MATTER

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K.S. Babu, S. Nandi, Z. Tavartkiladze, Phys. Rev. D 80, 071702 (2009)
 S. Bhattacharya, S. Jana and S. Nandi, Phys. Rev. D 95, no. 5, 055003 (2017)
 K. Ghosh, S. Jana and S. Nandi, arXiv:1705.01121
 T. Ghosh, S. Jana, S. Nandi, (to appear)

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DIMENSION 7 NEUTRINO MASS GENERA



- To provide a new mechanism for light neutrino mass generation with new mass scale at the TeV.
- To connect the neutrino physics with the physics that can be explored at the LHC.
- To connect neutrino physics to dark matter.

Outline of Talk



- Introduction
- Model and the Formalism
- Phenomenological Implications
 - ► for LHC
 - for Dark Matter
- Conclusions and Outlook



- The existence of neutrino masses are now firmly established. $m_{\nu} \sim 10^{-2} \text{ eV} \Rightarrow 1 \text{st}$ and only indication for physics beyond the SM
- m_{ν} is about a billion times smaller the quark and charged lepton masses
- What is the mechanism for such a tiny neutrino mass generation ?

Introduction



- Most popular mechanism : see-saw, $m_{\nu} \sim \frac{m_D^2}{M}$ \Rightarrow dimension 5 operator : $L_{eff} = \frac{f}{M}LLHH$ The observed neutrino mass, $m_{\nu} \sim 10^{-2}$ eV.
- If $M = M_{PL}$, then m_{ν} is too small
- If $M = M_{GUT}$, then m_{ν} is still too small
- $M \sim 10^{14}$ GeV is needed \rightarrow A new symmetry breaking scale (N_R)
- This scale is too high \rightarrow No connection can be made to the physics to be explored at the LHC or Tevatron \Rightarrow need $M \sim$ TeV.

Introduction



- It is possible that the dimension 5 operator does not contribute to neutrino masses in a significant way.
 ⇒ next operator (dimension 7) : L_{eff} = ^f/_{M³}LLHH(H[†]H)
- If $M=M_{PL}$, This by itself is not enough to make $M\sim$ TeV, need $f\sim 10^{-9}$
- We propose a model in which $f \sim y_1 y_2 \lambda_4$ with each $\sim 10^{-3}$ (domain of natural values)
- This gives $M \sim \text{TeV}$ scale to obtain neutrino masses in the range $10^{-2} 10^{-1}$ eV. \Rightarrow Connect to physics at the LHC, as well as dark matter

Model & Formaism



- Gauge Symmetry : $SU(3)_C \times SU(2)_L \times U(1)_Y$
- Usual SM particles +a pair of vector-like SU(2) triplet leptons, $\Sigma \equiv \begin{pmatrix} \Sigma^{++} \\ \Sigma^{+} \\ \Sigma^{0} \end{pmatrix} \sim (1,3,2), \ \bar{\Sigma} \equiv \begin{pmatrix} \bar{\Sigma}^{0} \\ \bar{\Sigma}^{-} \\ \bar{\Sigma}^{--} \end{pmatrix} \sim (1,3,-2)$ +a new isospin $\frac{3}{2}$ scalar quadruplet, $\Delta \equiv \begin{pmatrix} \Delta^{+++} \\ \Delta^{+} \\ \Delta^{0} \\ +a \text{ new scalar singlet, } S \sim (1,1,0).$
- Δ has positive mass square, but acquires a tiny induced VEV through Higgs potential via interaction with H.
- Σ has interactions with SM lepton doublets, H as well as Δ .

Model & Formalism



Higgs Potential

$$V(H,\Delta) = \mu_H^2 H^{\dagger} H + \mu_{\Delta}^2 \Delta^{\dagger} \Delta + \frac{\lambda_1}{2} (H^{\dagger} H)^2 + \frac{\lambda_2}{2} (\Delta^{\dagger} \Delta)^2$$
(1)
+ $\lambda_3 (H^{\dagger} H) (\Delta^{\dagger} \Delta) + \lambda_4 (H^{\dagger} \tau_a H) (\Delta^{\dagger} T_a \Delta) + \{\lambda_5 H^3 \Delta^* + h.c.\},$

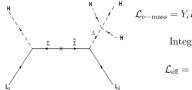
$$V(H,\Delta,S) = V(H,\Delta) + \mu_S^2 S^2 + \frac{\lambda_7}{2} S^4 + \lambda_8 (H^{\dagger} H) S^2 + \lambda_9 (\Delta^{\dagger} \Delta) S^2,$$
(2)

• Minimization of V \Rightarrow Neutral component of Δ acquires an induced VEV at the tree level, $v_{\Delta} = -\lambda_5 v^3 / M_{\Delta}^2$

Model & Formalism



Neutrino Mass Generation in the Model



$$_{-\mathrm{mass}} = Y_i L_i H^* \Sigma + \overline{Y}_i L_i \Delta \overline{\Sigma} + M_{\Sigma} \Sigma \overline{\Sigma} + h.c.,$$

Integrating out the $\Sigma, \overline{\Sigma}$ fermions

$$\mathcal{L}_{\text{eff}} = -\frac{(Y_i \overline{Y}_j + Y_j \overline{Y}_i) L_i L_j H^* \Delta}{M_{\Sigma}} + h.c.$$

 EWSB induces a VEV on the CP-even neutral component of the quadruplet

$$v_{\Delta} = -\lambda_5 v^3 / M_{\Delta}^2$$

• This leads to dimension 7 neutrino mass at tree level with Δ replaced by HHH/M_{Δ}^2 .

$$(m_{\nu})_{ij} = \frac{(Y_i Y_j' + Y_i' Y_j) v_{\Delta} v}{M_{\Sigma}} = -\frac{\lambda_5 (Y_i Y_j' + Y_i' Y_j) v^4}{(M_{\Sigma} M_{\Delta^0}^2)}.$$

• $m_{\nu} \sim 10^{-2} - 10^{-1}$ eV range with M_{Σ} and M_{Δ} at the TeV scale with $(Y_1, Y_2, \lambda_5) \sim 10^{-3}$ S. Nandi (OK State) DIMENSION 7 NEUTRINO MASS GENERAL July 26, 2017 9 / 27

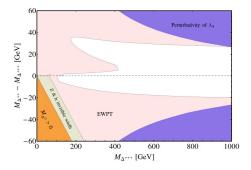


Signals at LHC :

- Productions of Δ 's in the model
- Decay modes
- Mass Bounds
- Signals
- Other Implications

Phenomenological Implications Constraints





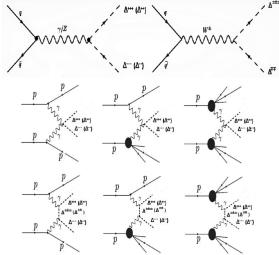
- $M_{\Delta_i}^2 = M_{\Delta^0}^2 q_i \frac{\lambda_4}{2} v^2 (q_i = 1, 2, 3)$
- $\Delta M > 0 \implies M_{\Delta^{+++}} < M_{\Delta^{++}} < M_{\Delta^+} < M_{\Delta^0}$
- $\Delta M < 0 \implies M_{\Delta^{+++}} > M_{\Delta^{++}} > M_{\Delta^+} > M_{\Delta^0}$
- EWPT \implies *S*, *T* parameters constrains the parameter space in $M_{\Delta^{\pm\pm\pm}} \Delta M$ plane

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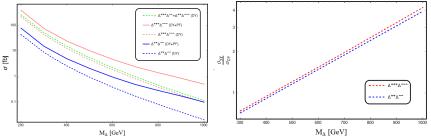
A. Productions



- Drell Yan Production
- Photon Photon Fusion



A. Productions

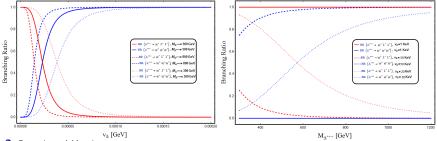


- $\Delta^{\pm\pm\pm}(\Delta^{\pm\pm})$ pair and associated production at the LHC happens via Drell-Yan (DY)
- Photon fusion (PF) is another process \Rightarrow Photon PDF is available from NNPDF, CTEQ, MRST
- $\bullet\,$ For larger $\Delta^{\pm\pm\pm}(\Delta^{\pm\pm})$ PF contribution is significant
 - \rightarrow However uncertainty in available photon PDFs are significant (${\sim}25$
 - 30%)

Babu, Jana, (2016) [arXiv:1612.09224], K.Ghosh, Jana, Nandi, (2017) [arXiv:1705.01121]



B. Decay Modes of Triply Charged Scalar

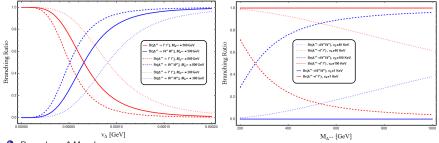


Depends on ΔM and v_Δ

- For $\Delta M \ge 0 \implies$
 - $\overline{\Delta}^{\pm\pm\pm} \rightarrow I^{\pm}I^{\pm}W^{\pm}(W^{\pm}W^{\pm}W^{\pm})$ dominates for small (large) v_{Δ}
 - $\Delta^{\pm\pm} \rightarrow I^{\pm}I^{\pm}(W^{\pm}W^{\pm})$ dominates for small (large) v_{Δ}
 - $\blacktriangleright\,$ Crossover happens at $\sim 10^{-4}~{\rm GeV}$
 - ▶ For $\Delta M \gtrsim 2-20$ GeV Cascade Decay $\Delta^{\pm\pm} \rightarrow H^{\pm\pm\pm}W^{*\mp}$ dominates
- For $\Delta M < 0 \implies$
 - $\Delta^{\pm\pm\pm}
 ightarrow \Delta^{\pm\pm} W^{\pm}$ always happens
 - $\Delta^{\pm\pm} \rightarrow I^{\pm}I^{\pm}(W^{\pm}W^{\pm})$ dominates for small (large) v_{Δ}
 - ► For $\Delta M \gtrsim 2 20$ GeV Cascade Decay $\Delta^{\pm\pm} \rightarrow H^{\pm}_{P} W^{*\pm}_{*}$ dominates $_{\sim\sim\sim}$

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Depends on ΔM and v_{Δ}

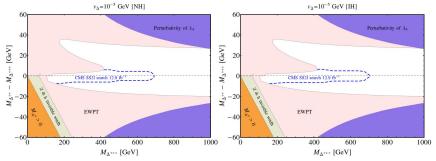
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 - ► For $\Delta M \gtrsim 2-20$ GeV Cascade Decay $\Delta^{\pm\pm}_{-} \rightarrow H^{\pm}_{-} W^{*\pm}_{-}$ dominates $_{\supset Q,Q}$





C. Mass Bounds



 $\bullet~$ CMS and ATLAS searches for $\Delta^{\pm\pm}$ in SS2I final states

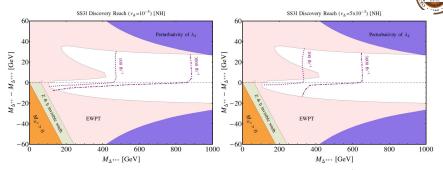
- However they assume 100% BR in various leptonic channel
- CMS analysis at 12.9 fb⁻¹ provides strongest limit CMS PAS HIG-16-036
- We obtain the strongest limits from $\mu^\pm\mu^\pm(e^\pm e^\pm)$ for NH (IH)
- No bound for $W^{\pm}W^{\pm}$ dominated channel

D. Searching for $\Delta^{\pm\pm\pm}$ at the LHC



- Although $\Delta^{\pm\pm}$ has a better prospect to be found at the LHC, this particle is not exclusive to this model
- \bullet To verify/falsify this model we also need to search for $\Delta^{\pm\pm\pm}$
- $\Delta^{\pm\pm}$ searches looses sensitivity for $\Delta M\gtrsim 5$ GeV
- For $\Delta^{\pm\pm\pm}$ one needs to look at SS3I channel \implies sensitivity remains the same for all $\Delta M > 0$
- Search Strategy \implies 3 isolated SS leptons (e, μ) , $p_T(l_1, l_2, l_3) > (30, 30, 20)$ GeV, $\not E_T > 30$ GeV, Z-veto
- Major BGs $\rightarrow t\bar{t}(Z/\gamma^*), t\bar{t}W^{\pm}, t\bar{t}t\bar{t}, I^+I^-VV(V=Z,W^{\pm})$
- After cut $t \bar{t} W^{\pm}$ dominates $ightarrow \sigma_{BG}^{\textit{total}} pprox 5 imes 10^{-3}$ fb

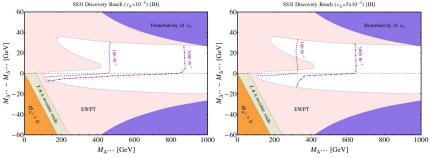
E. Future Prospects of SS3I Search



- Discovery potential upto 450 (950) GeV at 100 (3000) fb⁻¹ for *IIW* dominated region Discovery potential upto 500 (950) GeV at 100 (3000) fb⁻¹ for *IIW* dominated region
- Discovery potential upto 350 (700) GeV at 100 (3000) fb $^{-1}$ for WWW dominated region
- Covers the whole area available for $\Delta M > 0$ scenarios
- Similar results for NH and IH

E. Future Prospects of SS3I Search





- Discovery potential upto 450 (850) GeV at 100 (3000) fb⁻¹ for *IIW* dominated region
- Discovery potential upto 350 (650) GeV at 100 (3000) fb $^{-1}$ for $W\!W\!W$ dominated region
- Covers the whole area available for $\Delta M > 0$ scenarios
- Similar results for NH and IH

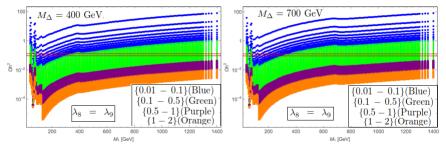
Dark Matter



- To bring both the issues of neutrino masses and DM together under one umbrella with a minimal possible extension of the SM
- The DM is an electroweak (EW) singlet scalar *S*, odd under an imposed exact *Z*₂ symmetry, interacting to SM through 'Higgs-portal' coupling, while all other particles are even under *Z*₂.
- Additional interactions with Δ, the scalar singlet DM S survives a large region of parameter space by relic density constraints from WMAP/PLANCK and direct search bounds from updated LUX data.
- The relevant parameter space of this model is spanned by : $\Rightarrow \{M_s, M_{\Delta}, \lambda_8, \lambda_9\}$

Dark Matter Relic Density





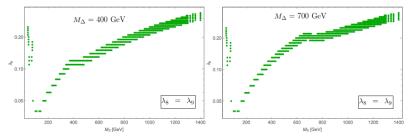
- The scalar singlet S introduced here interacts with the scalar quadruplet Δ and can annihilate through $SS \rightarrow \Delta^0 \Delta^0, \Delta^+ \Delta^-, \Delta^{++} \Delta^{--}, \Delta^{+++} \Delta^{---}$ on top of annihilations to SM particles through Higgs portal interactions.
- Relic density of the DM in the present universe is obtained by the annihilation cross-section of the DM as

$$\Omega h^2 = \frac{0.1 \mathrm{pb}}{\langle \sigma v \rangle}.$$
 (3)

 $\Rightarrow \text{ Annihilation cross-section}: \langle \sigma v \rangle = \langle \sigma v \rangle_{SS \to SM} + \langle \sigma v \rangle_{SS \equiv \Delta\Delta \equiv S} \quad \text{ and } \quad v \in SS = SM$

Dark Matter Relic Density





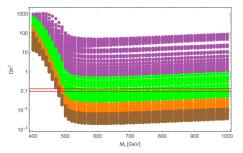
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Dark Matter Relic Density





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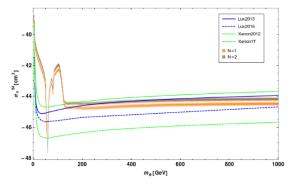
$$\Omega h^2 = \frac{0.1 \text{pb}}{\langle \sigma v \rangle}.$$
(5)

 $\Rightarrow \text{Annihilation cross-section} : \langle \sigma v \rangle = \langle \sigma v \rangle_{SS \to SM} + \langle \sigma v \rangle_{SS \to \Delta\Delta}$

Dark Matter



Direct Detection Constraints

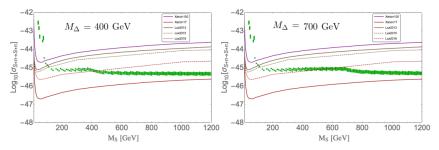


 Scalar singlet extension of SM to accommodate DM through Higgs portal interaction is under tension as the allowed region of relic density space has been ruled out to a very large DM mass excepting for the Higgs resonance by non-observation in direct search experiment, especially the LUX and XENON data

Dark Matter



Direct Detection Constraints



 Thanks to the additional interactions with Δ, the scalar singlet DM S survives a large region of parameter space by relic density constraints from WMAP/PLANCK and direct search bounds from updated LUX data.

Conclusion



- Presented a model to generate tiny neutrino masses via tree-level d=7 operator
- The scale of new physics \simeq TeV
- The model has doubly charged, as well as tripply charged scalar
- The decay of the triply charged scalar can produce diplaced vertex at the LHC
- The doubly charged scalar can be probed upto 1 TeV at the LHC in the SS2I channel
- The triply charged scalar can be probed upto 500 GeV at the LHC in the SS3I model
- A singlet scalar is a viable dark matter candidate in this model.

The End

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