

# Design of the LBNF Beamline

Jim Hylen, for the DUNE collaboration  
Fermilab Accelerator Division

DPF 2017

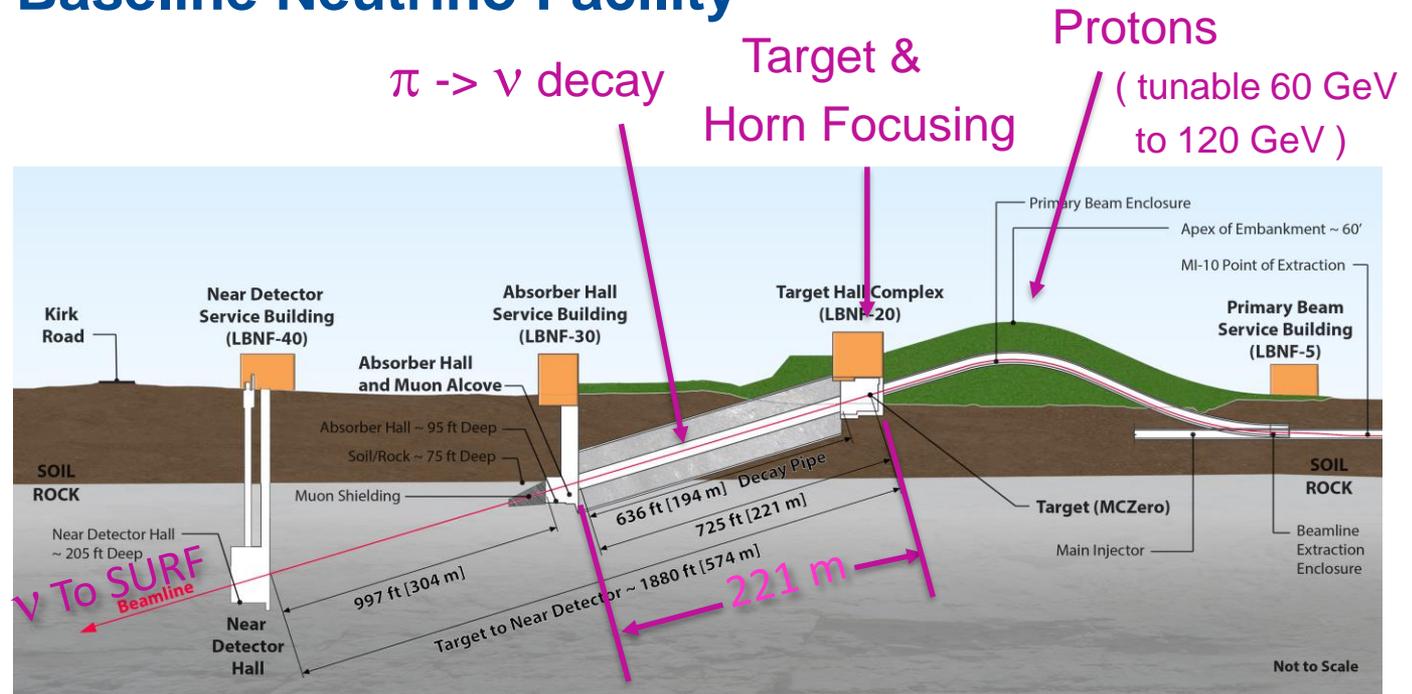
July 31, 2017



# LBNF = Long Baseline Neutrino Facility

Takes protons from Fermilab accelerator

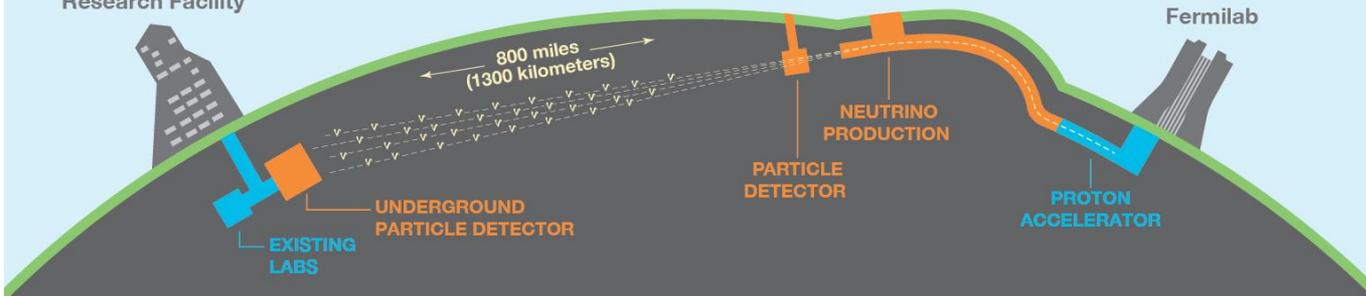
Produces beam of  $\nu_\mu$  or  $\bar{\nu}_\mu$



Aimed at DUNE detector in SURF 1300 km away

Sanford Underground Research Facility

Fermilab

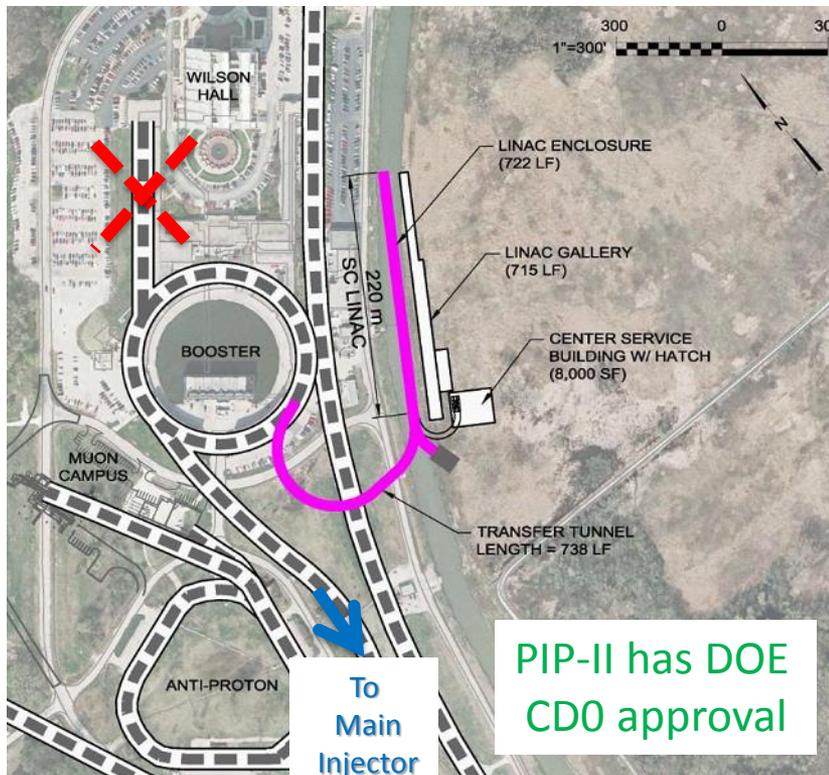


Study neutrino oscillation phenomena

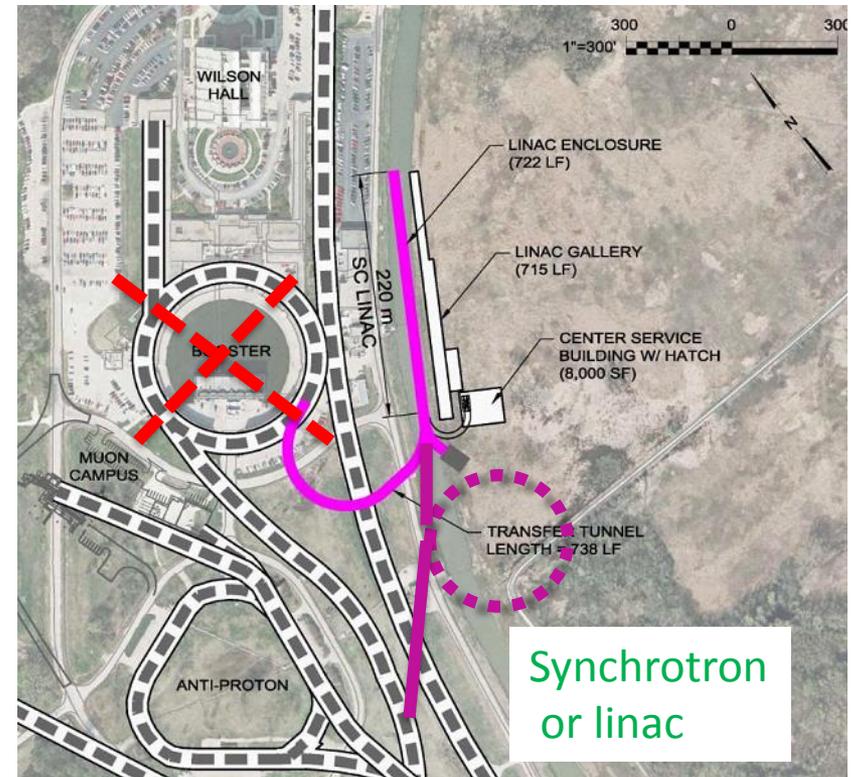
# Accelerator stages: Linac -> Booster -> Main Injector -> beamline

Fermilab NuMI neutrino beam recently upgraded; under Proton-Improvement-Plan I (PIP-I) went from 0.4 MW to 0.7 MW proton beam power (*achieved this year !*)

- LBNF to start operation at 1.2 MW with PIP-II new linac to Booster
- LBNF designed for upgrade to 2.4 MW with PIP-III replacement for Booster



LBNF initial target & horns for 1.2 MW



LBNF permanent parts 2.4 MW capable

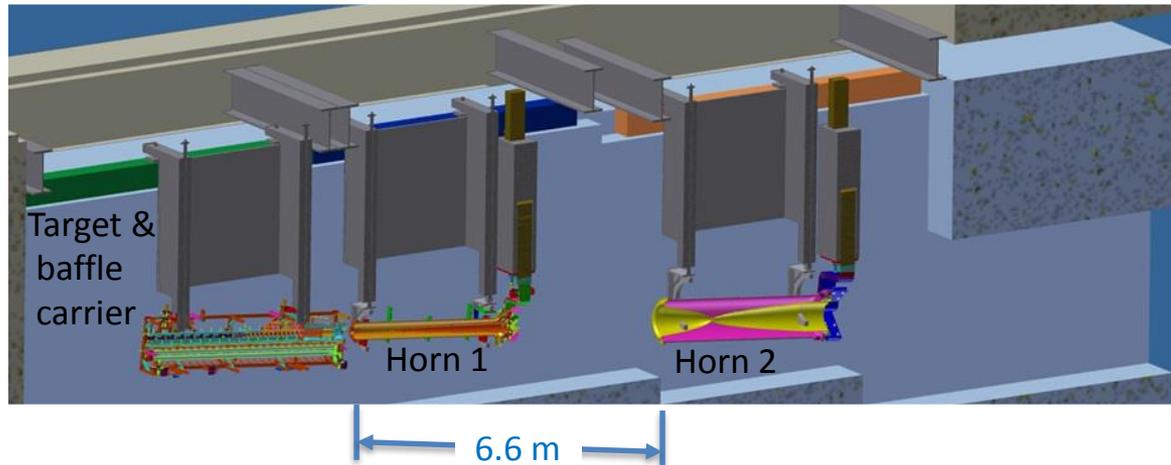
# Detector drivers for LBNF beam design

- Detector to measure CP violation;  $\nu_{\mu} \rightarrow \nu_e$  vs.  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$ 
  - Beam should produce as many neutrinos as possible around 1<sup>st</sup> and 2<sup>nd</sup> oscillation peaks ( $E_{\nu} \sim 2.4$  and  $0.8$  GeV for  $L=1300$  km)
- Does 3- $\nu$  mixing picture hold together ?
  - Broad energy spectrum to look for deviations as function of  $L/E$
  - Implies beam pointed at detector, rather than off-axis (T2K and NOVA)
- DUNE far detector is non-magnetized
  - $\nu$  vs  $\bar{\nu}$  selection done by beam focusing  $\pi^+$  or  $\pi^-$  (defocus other)
  - Toroidal horn magnetic field, (rather than e.g. solenoid-gradient focusing)
- Detector and beam will run for decades
  - Include flexibility to modify beam-line, e.g. for higher  $E_{\nu}$  spectrum
  - Implies possibly different target/horn shapes and locations

# Two different horn configurations on table, *plus staging options*

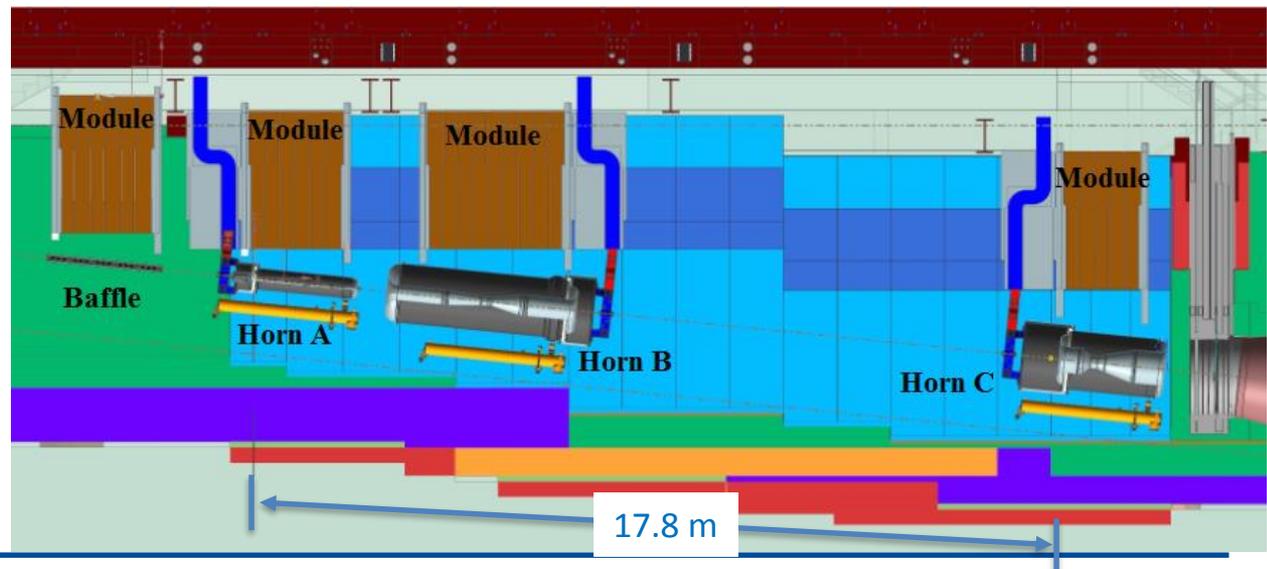
**CD-1 Reference** lower-cost starter, based on proven NUMI tech, ...upgraded/replaced later

- 2 horns
- 1 m long target
- Target inserted 2/3 way into horn 1



**Optimized design** recent optimized configuration for DUNE CP violation

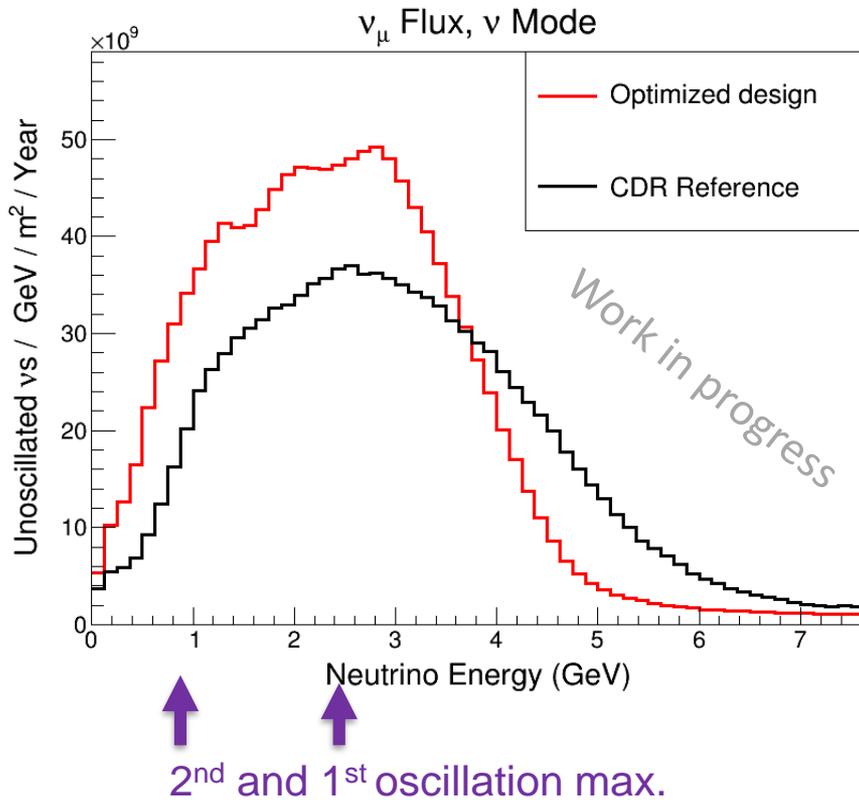
- 3 horns
- 2 m long target
- Target mounted entirely in horn A



**Optimized staged**

- start with just 2 of the 3 optimized horns

# Optimized versus reference design



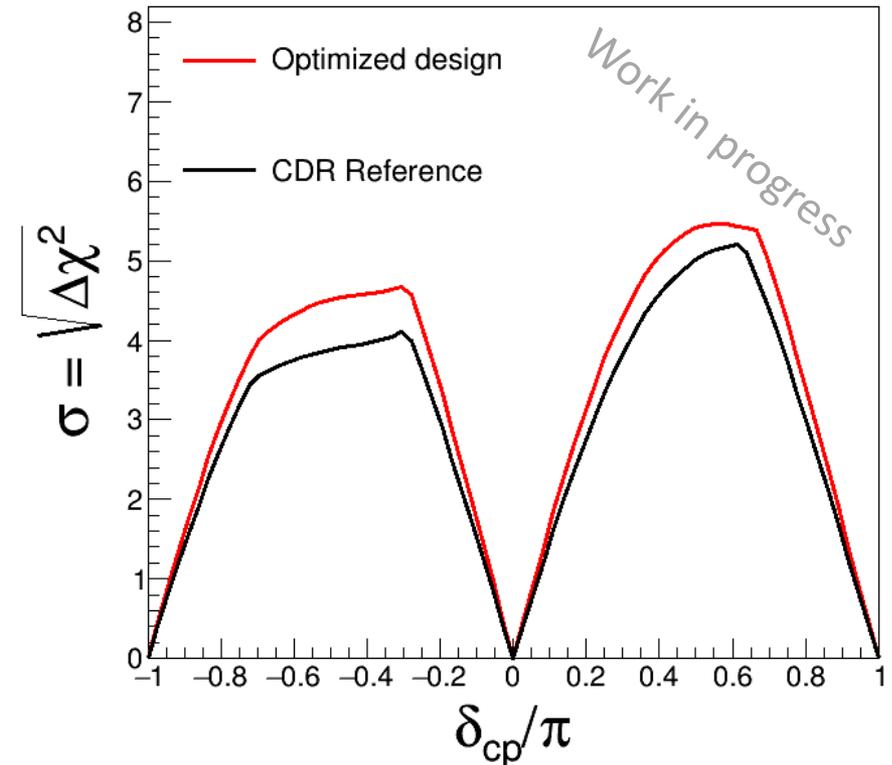
## Optimized system:

- Increases flux in oscillation region
- Decreases flux in high-energy tail
- Increases CP sensitivity

Sensitivity for (detector) x (beam) of 300 kT MW years exposure

- includes derating for beam down-time
- DUNE reference = 40 kT detector

## CP violation sensitivity

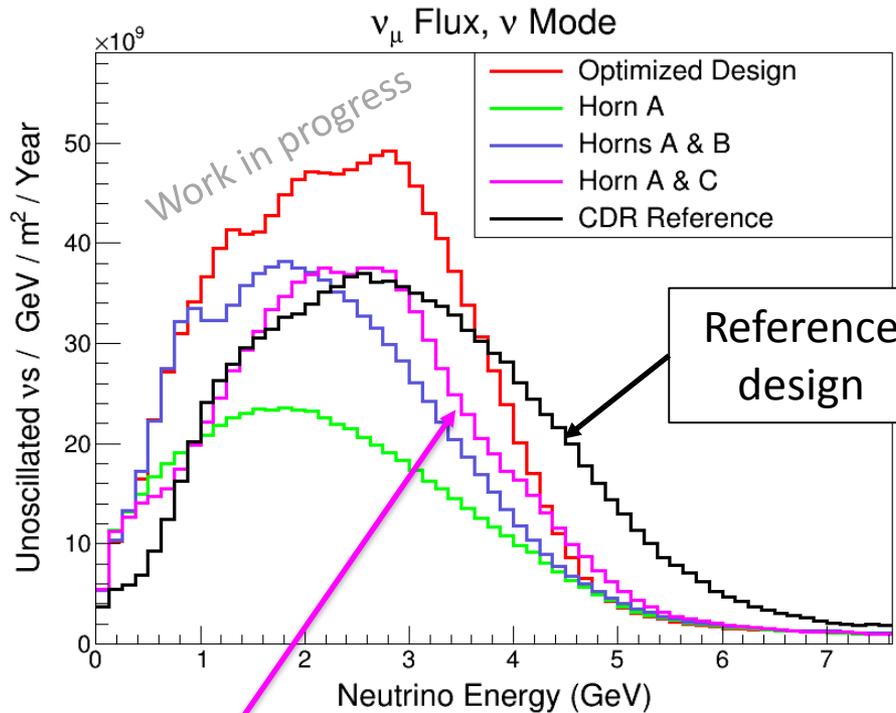


See Rowan Zaki's presentation

"Optimization of the LBNF Neutrino Beam"

# Beam-line Staging possibility

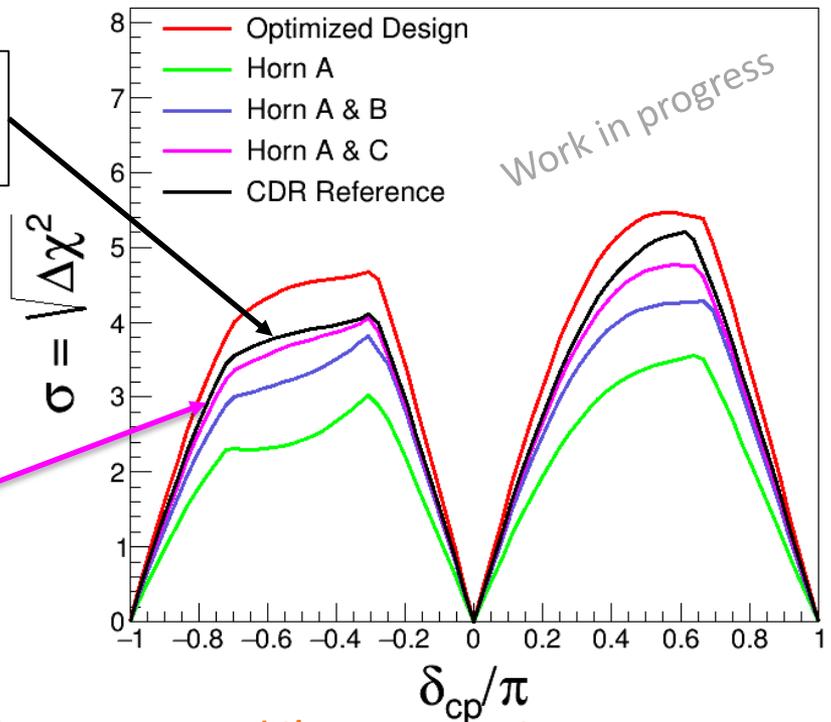
If do not have resources for fully-optimized beam at start



Sensitivity for detector x beam of 300 kT MW years exposure

- includes derating for beam down-time
- DUNE reference = 40 kT detector

## CP violation sensitivity



Using horns A&C from optimized design is nearly as good as the 2-horn reference design

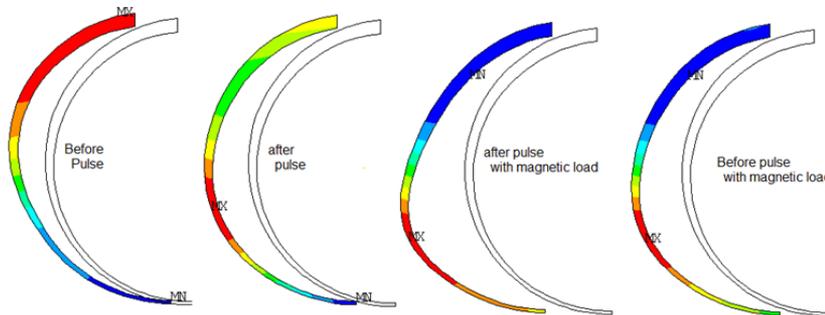
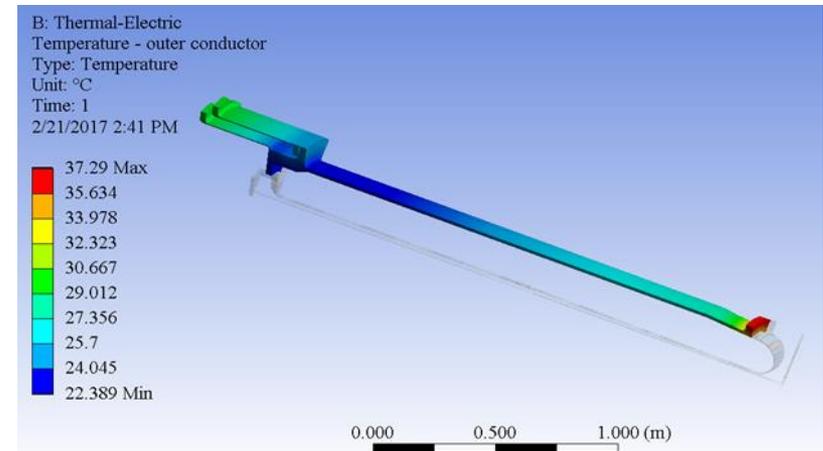
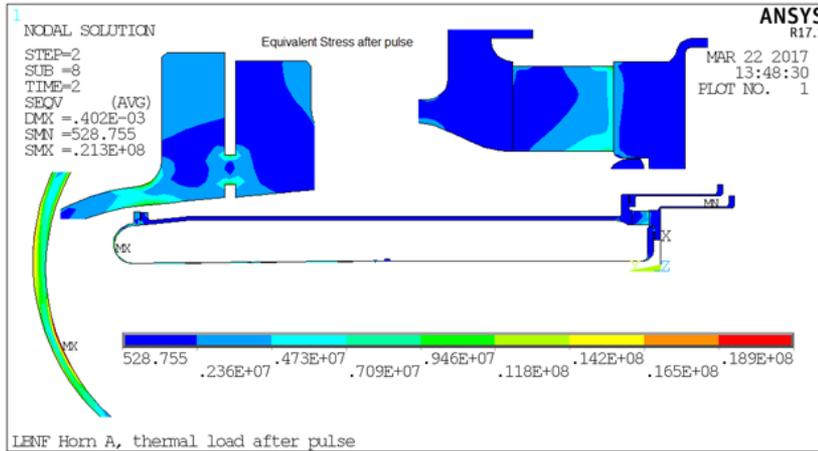
➤ would allow much easier later upgrade to fully optimized

See Rowan Zaki's presentation  
 "Optimization of the LBNF Neutrino Beam"

# Have completed conceptual design of optimized horns

## Temperature and Stress looks OK

FEA of Horn A, which has highest current density and beam heating



Point	Safety Factor for stress	
	No preload	With Preload
1	1.87	2.00
2	1.36	1.75
3	2.2	3.00
4	2.46	3.10
5	1.91	2.10
6	1.91	2.10



See further information in Cory Crowley's poster  
 "LBNF Optimized Horn Design & Target Integration"

# Energy & radiation deposition

- Much of work for design of high power neutrino beam is radiation and rad safety
  - Prompt, air-borne, ground-water, residual, remote handling, radiation damage, ...

Let's start by looking at where the beam power ends up.

System	RD (kW)	OD (kW)	OD/RD
Target Pile	952	1238	1.30
Decay Pipe Region	452	542	1.20
Hadron Absorber	786	400	0.51
Misc: infrastructure, binding energy, sub-thrshld	144	151	1.05
<i>Neutrino power</i>	66	69	1.05
<b>Total</b>	<b>2400</b>	<b>2400</b>	

For 2.4 MW proton beam power

← kW deposited in region

For Ref. Design (RD) and Opt. Design (OD)

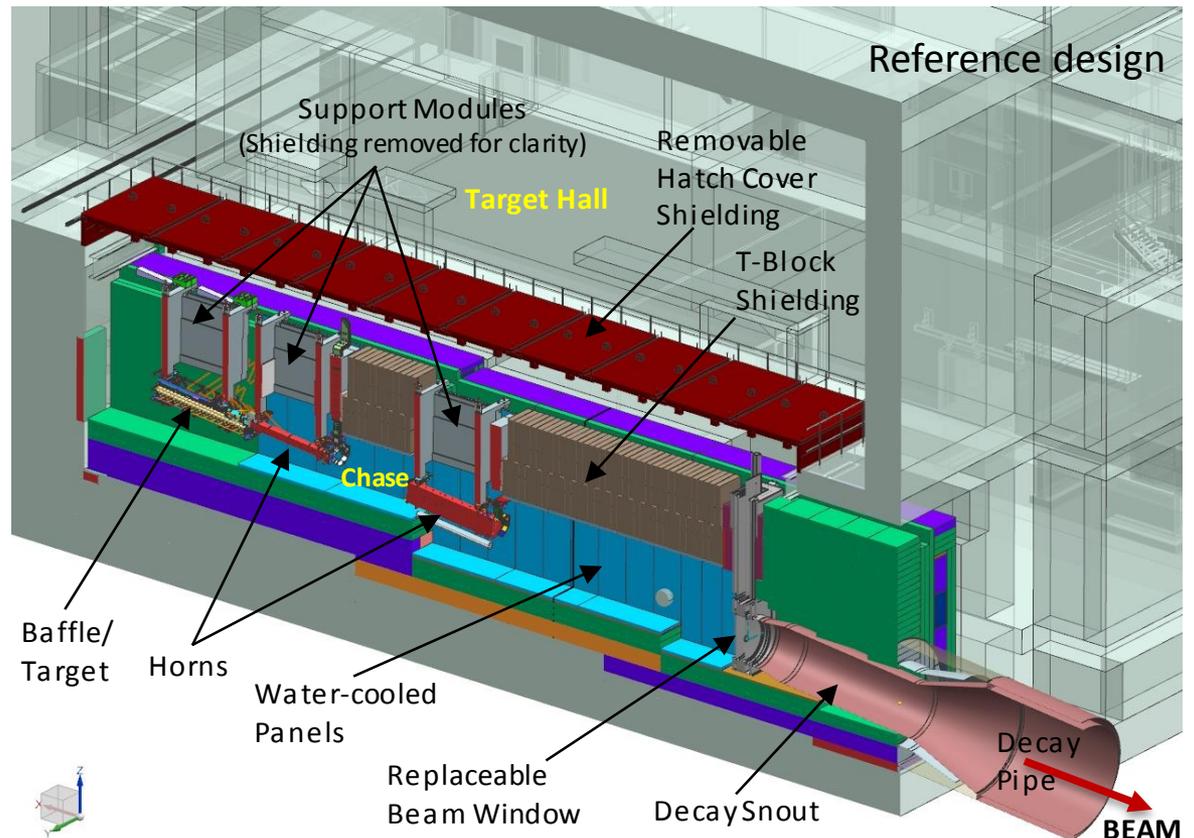
~  $10^{-13}$  watt deposited in far detector !

*MARS Monte Carlo*

# Target pile *inside* target hall

- Radiation shielding around target:
  - 1.8 m steel + 1 m concrete thick on sides
  - 3 m steel + 15 cm Borated Poly on top
- Component alignment:  
< 1 mm
- (nearly) sealed gas volume
- Target & horn handling  
all remotely done due to high residual radiation

*Recent work:*  
Changed gas in pile from air to N<sub>2</sub>;  
eliminates <sup>41</sup>Ar production,  
also ozone + nitric acid corrosion



# Cooling design choices

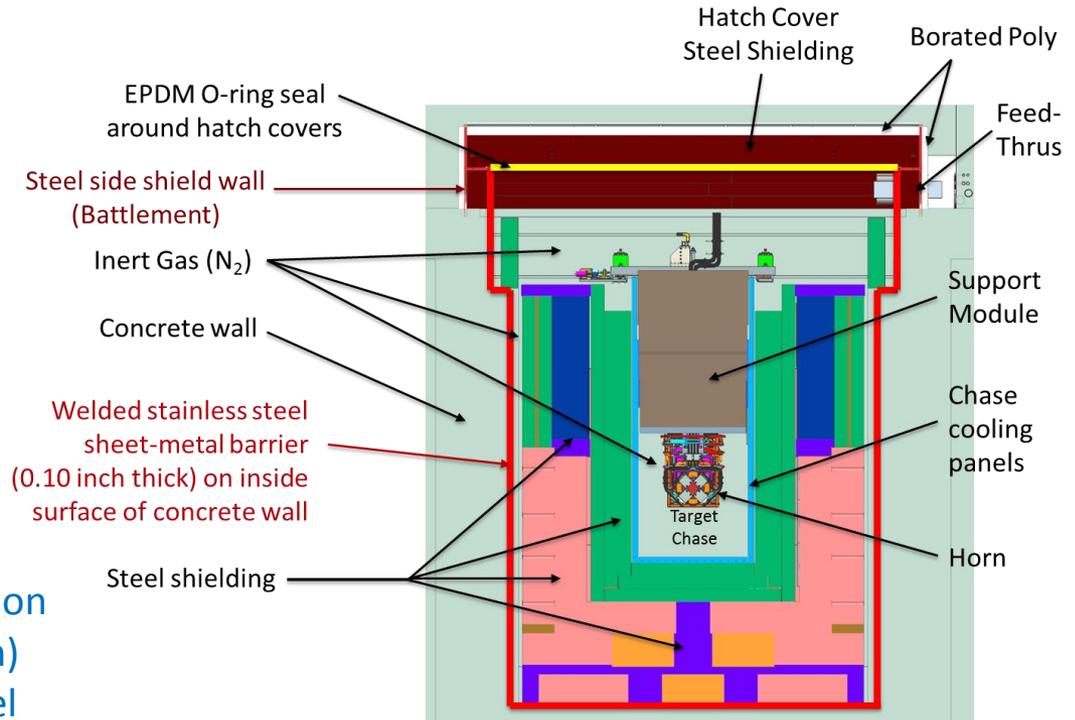
LBNF designed for 30 year lifetime. All water piping required to be replaceable/repairable.  
→ Use gas cooling for permanent/unreachable structures.

Replaceable water cooling panels are used for innermost steel layer

Bulk shielding is all cooled by 35,000 ft<sup>3</sup>/minute flow of N<sub>2</sub> gas

## Lessons learned being applied:

- Concrete is all outside the N<sub>2</sub> vessel
- Vessel includes all gas in high radiation (containing short-lived air-activation)
- Steel (emitting tritium) is all in vessel
- Continuously purge tritium by slow N<sub>2</sub> release (1 to 7 cfm)



See Joseph Angelo's poster

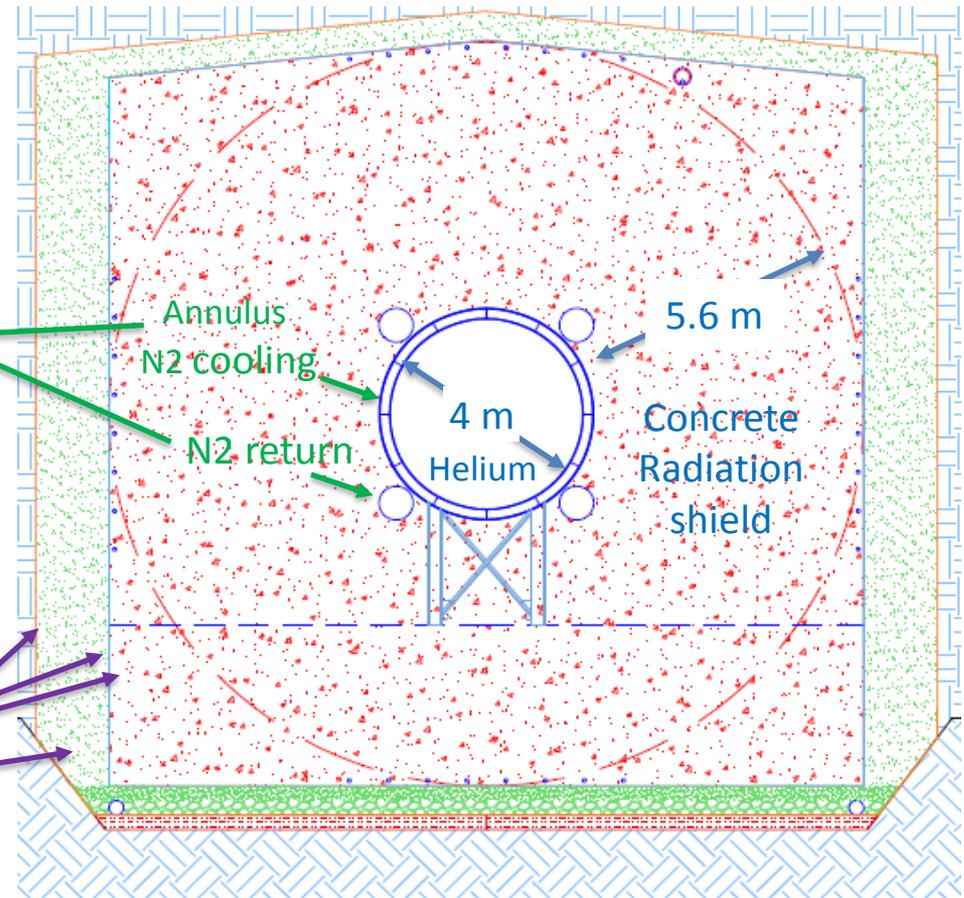
"Design of a Nitrogen Cooled Target Shield Pile for the LBNF Beamline"

# Decay pipe region

Recent work:

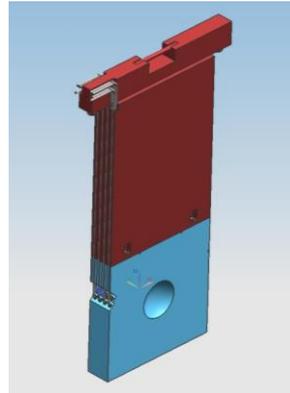
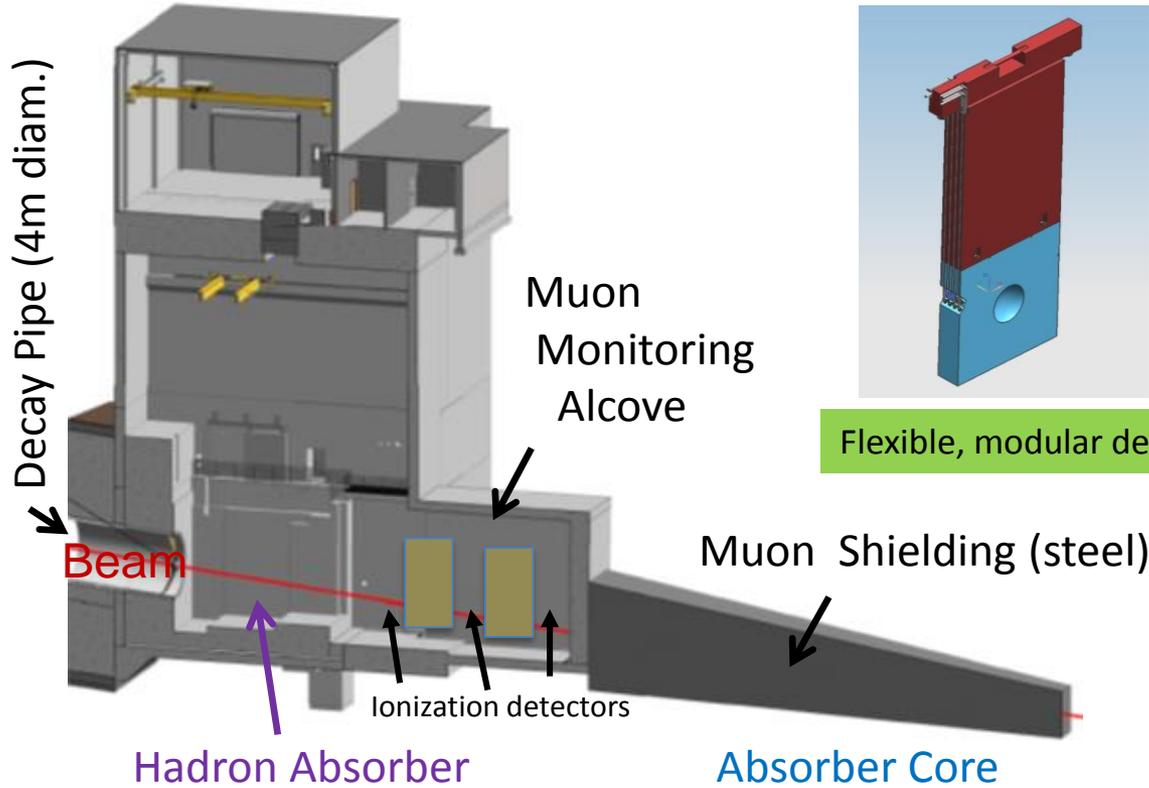
Changed gas cooling from air to N<sub>2</sub>,  
reduces <sup>41</sup>Ar production, also corrosion

- Pipe is 4 m diam., 194 m long
- Static helium fill  
( 10% more v compared to air fill )
- Cooled by flowing 35,000 ft<sup>3</sup>/minute of nitrogen gas
- Structure dominated by concrete radiation shield
- Multiple features to keep water out  
Water barriers  
Fall-back water drainage



Core blocks are replaceable  
via Remote Handling  
(each 1 ft thick)

## Absorber region

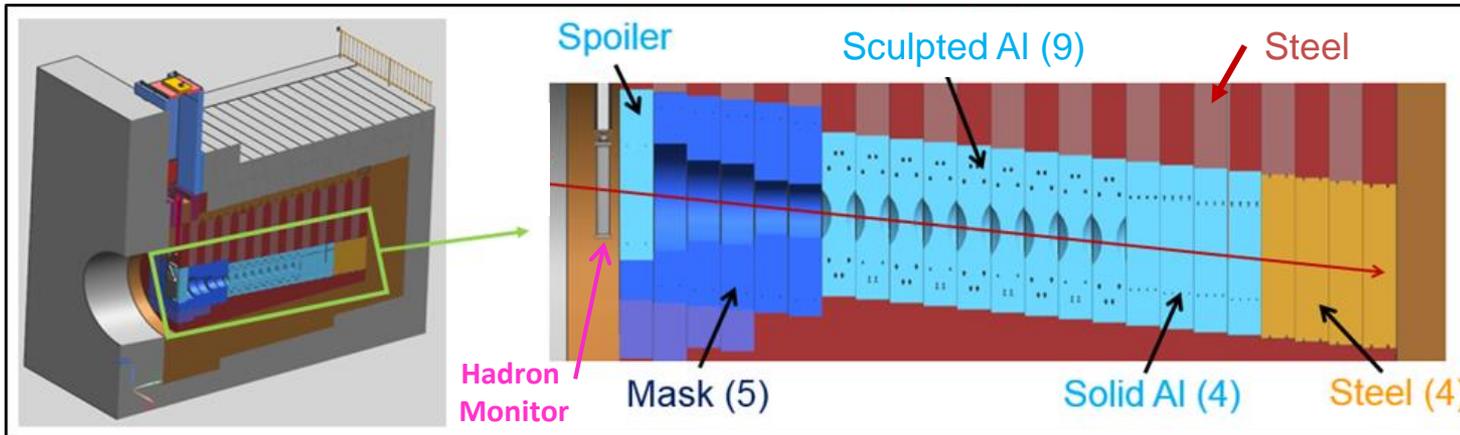


Flexible, modular design

Recent work:

Energy deposition is *LESS* for opt. beam than for ref. beam

Opt. beam may allow widening mask holes & elimination of sculpting of core blocks, thus improving muon monitoring capability

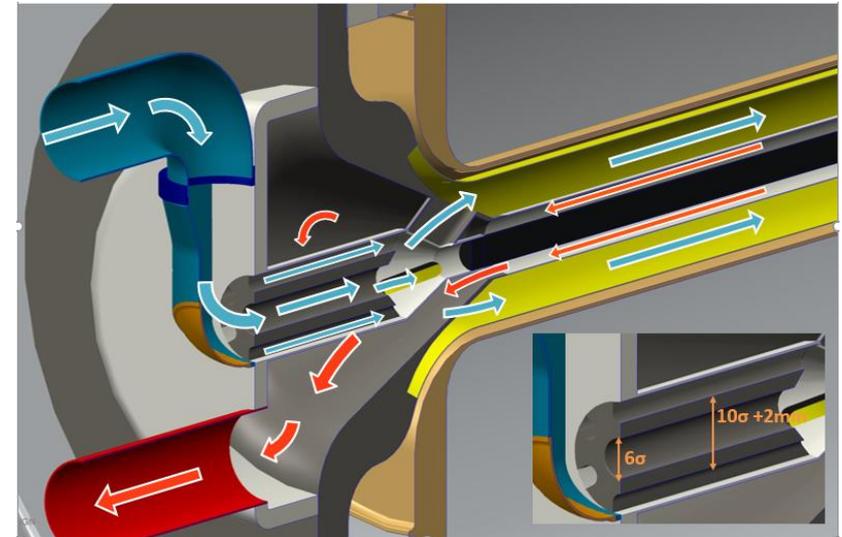
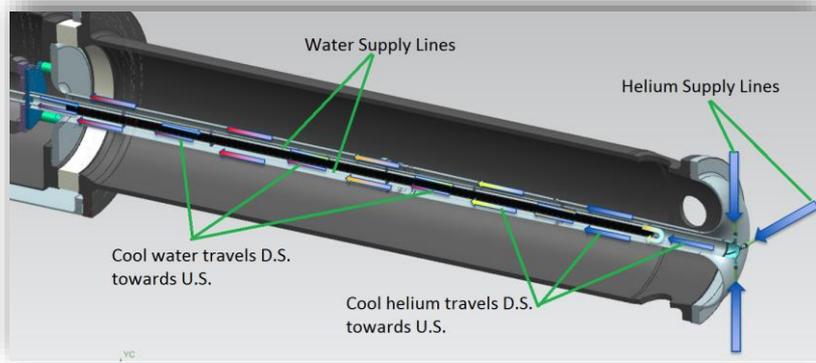


Core:  
water-cooled

Rest of shielding:  
forced air-cooled

# Target alternatives for opt. beam: both 2 m long graphite

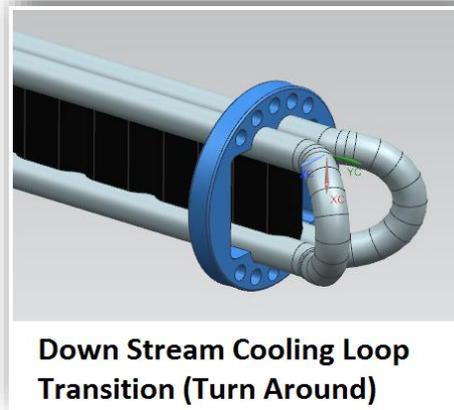
- Existing design (NuMI-like)
  - Water-cooled
- Being developed by RAL-UK
  - Helium-cooled



Graphite cylinders centered in coaxial Titanium tubes carrying helium

Graphite at significantly higher temperature;  
Radiation damage partially anneals at high T  
*Target may last longer !*

Graphite fins  
brazed to  
Titanium tubes  
carrying water



# Summary: Continue to make progress on LBNF beam design

## Beam focusing:

- Have complete reference conceptual design for a NuMI-like 2-horn system
- Nearing completion of conceptual design for a 3-horn system with longer target, **optimized for DUNE detection of  $\nu$  CP violation**. This system also allows an attractive 2-horn staging scenario.

Plan to make decision in the fall of 2017 on which course to pursue through preliminary design.

## Target pile atmosphere and decay pipe cooling:

- Completed conceptual design for replacing air fill with nitrogen
  - Eliminates production of radioactive  $^{41}\text{Ar}$
  - Eliminates Ozone and Nitric Acid corrosion

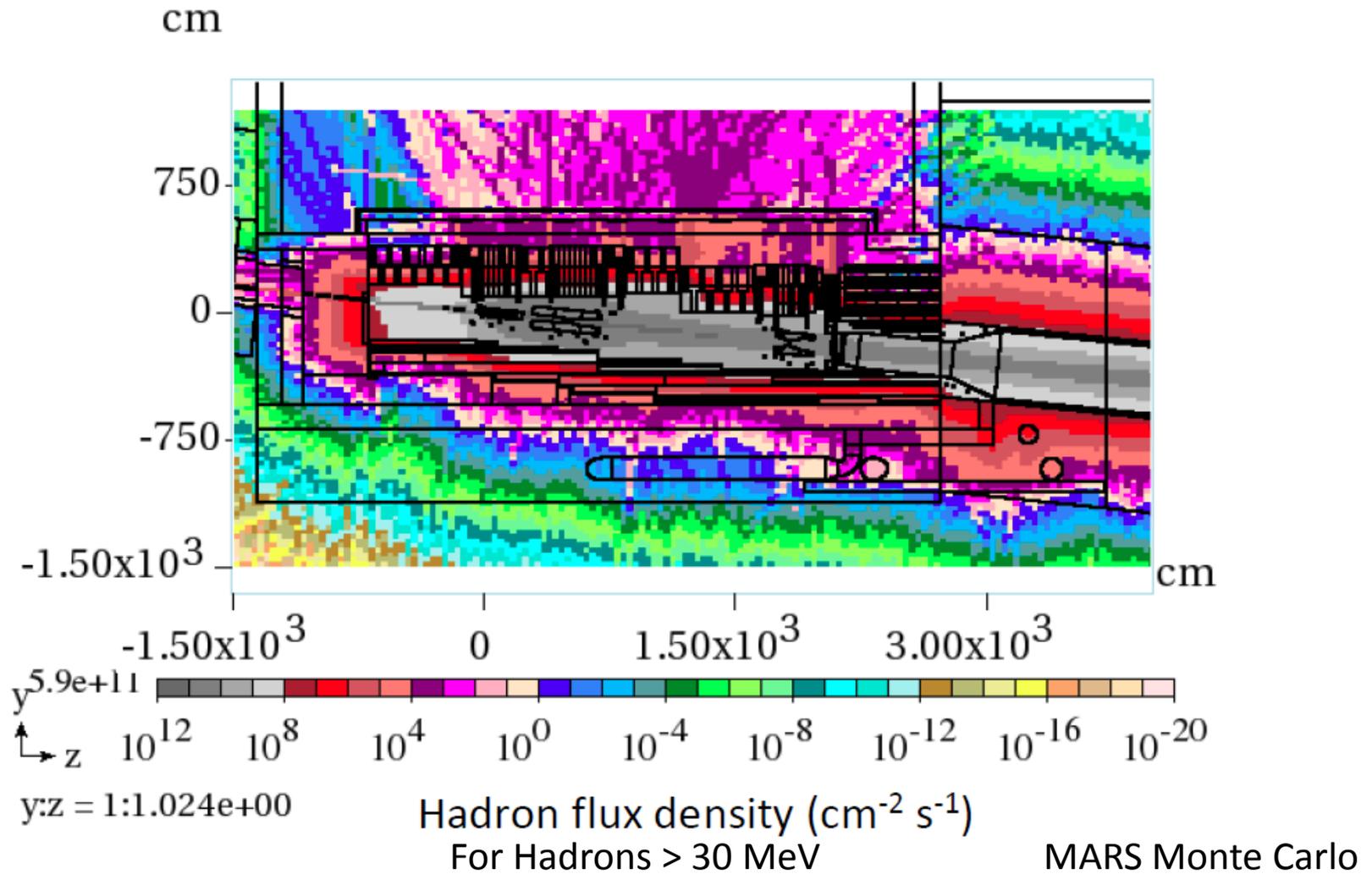
## Target

- Developing Helium cooled target design, alternate to reference water cooled design

*Beam to DUNE about a decade from now*

# BACK-UP

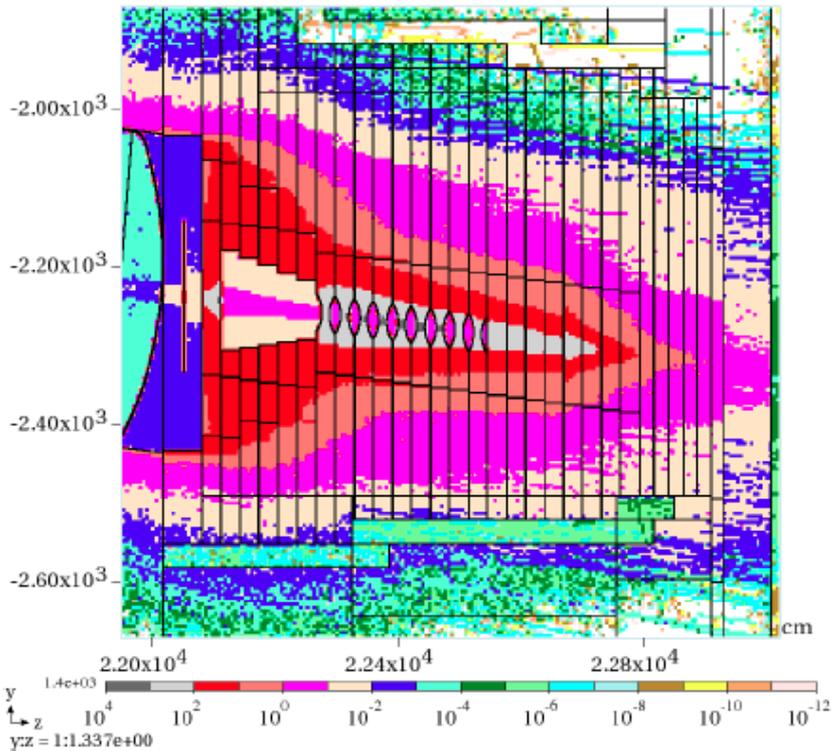
# Hadron flux for calculating air activation in target pile



# Power Density (mW/cm<sup>3</sup>) in Hadron Absorber

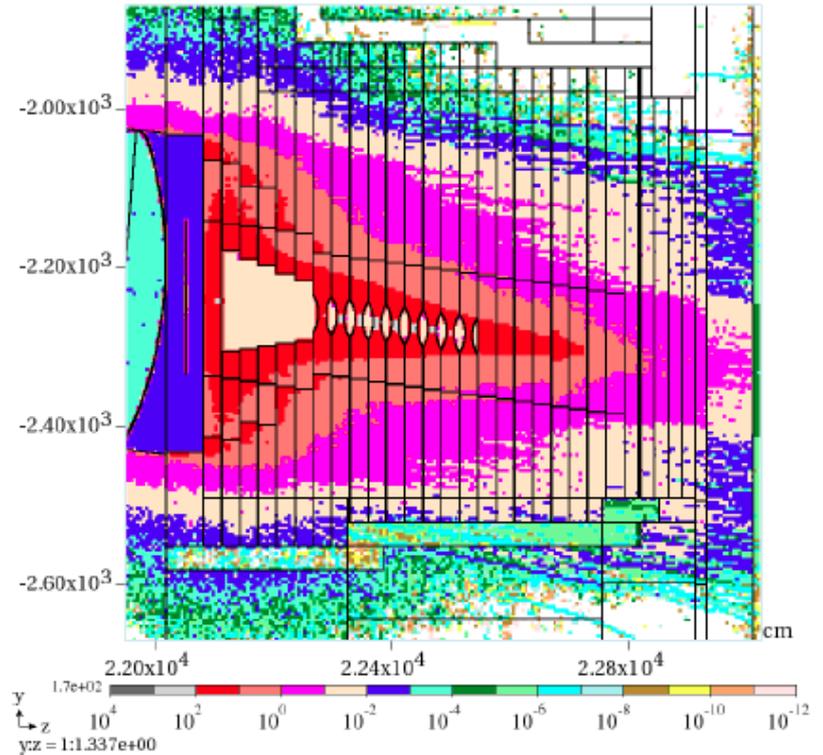
RD

Peak: 1.6 W/cm<sup>3</sup>



OD

Peak: 0.24 W/cm<sup>3</sup>



MARS Monte Carlo