# Extracting neutrino oscillation parameters using a simultaneous fit of the $\nu_{\rm e}$ appearance and $\nu_{\mu}$ disappearance data in the NOvA experiment



Prabhjot Singh, University of Delhi, India

On Behalf of the NOvA Collaboration

DPF 2017

Fermilab

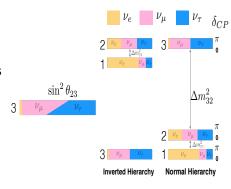
August 03, 2017





#### Current questions in neutrino physics

- Ordering of the neutrino masses: Normal or Inverted hierarchies?
- CP violation: do the neutrinos and anti-neutrinos oscillate differently?
- $\theta_{23}$  angle: is the mixing between  $\nu_{\mu}$  and  $\nu_{\tau}$  maximal or not?



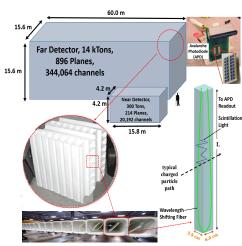
#### Introduction to the NOvA Experiment

- NOvA (NuMI Off-axis  $\nu_e$  Appearance) is a two detector, long baseline, neutrino oscillation experiment
  - NuMI (Neutrinos at Main Injector) beam is 97.5%  $\nu_{\mu}$
  - NOvA is 14.6 mrad off-axis from the NuMI beam, observed energy spectrum peaks at 2 GeV peak
  - Near detector, ND, on-site at Fermilab.
  - Far detector, FD, at Ash River, MN, 810 km from the target.
- Oscillation channels
  - ullet  $u_{\mu} 
    ightarrow 
    u_{\mu}$  disappearance
  - ullet  $u_{\mu} 
    ightarrow 
    u_{e}$  appearance
  - Anti-neutrino modes

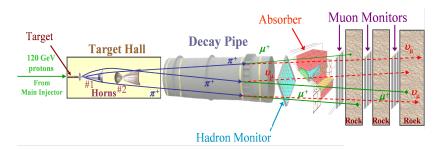


#### **NOvA Detectors**

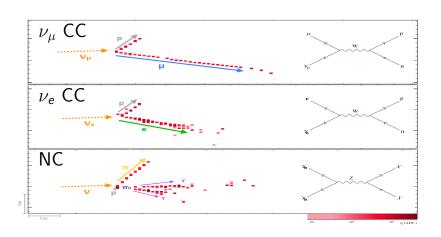
- NOvA detectors are functionally identical
  - Polyvinyl chloride (PVC) and low Z scintillator detectors.
  - NOvA cells are arranged in alternating horizontal and vertical planes for 3D tracking.
- The ND is used to measure the beam before oscillations and the FD measures the oscillated spectrum.



#### NuMI (Neutrinos at Main Injector) beam



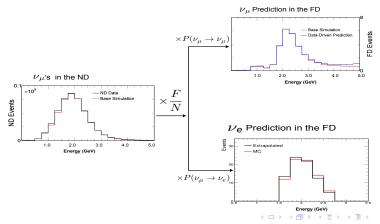
- A beam of high energy protons from the Main Injector are impinge on a fixed graphite target producing pions and kaons.
- Magnetic horns selects particles of the desired charge & momentum and focus them into a narrow beam.
- Charged pions and kaons spontaneously decay into muons and neutrinos.
- Rock after absorber filters out muons neutrinos produced in decay pipe.
- We are left with the neutrino beam.



- $\nu_{\mu}$  CC is the signal for the  $\nu_{\mu}$   $(\bar{\nu}_{\mu})$  disappearance analysis.
- $\bullet$   $\nu_e$  CC is the signal for the  $\nu_e$   $(\bar{\nu}_e)$  appearance analysis.
- NC is a background in above analyses.

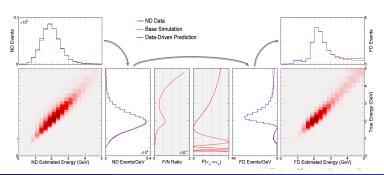
#### Extrapolation: ND Role

- ND spectra are used to predict the signal and backgrounds in the FD.
- Both  $\nu_{\mu} \to \nu_{\mu}$  disappearance and  $\nu_{\mu} \to \nu_{e}$  appearance analyses use  $\nu_{\mu}$ 's in the ND.
- The P( $\nu_{\alpha} \to \nu_{\beta}$ ) and the simulated spectra from ND and FD are used in the form of Far over Near, F/N, ratios to make the FD predictions.



#### Extrapolation: Far detector signal prediction

- $\bullet$  The reconstructed ND  $\nu_{\mu}$  CC energy spectrum is used to correct the FD simulated prediction.
- True energy distribution is corrected so that the reconstructed data & MC agree at the ND.
- The FD and ND MC spectra along with the oscillation probability for the  $\nu_{\mu} \to \nu_{\mu}$  disappearance and  $\nu_{\mu} \to \nu_{e}$  appearance are used to extrapolate corrected ND true spectra to the FD.
- The extrapolated true energy distribution is converted to the reconstructed energy distribution by using  $E_{true} o E_{reco}$  distribution.



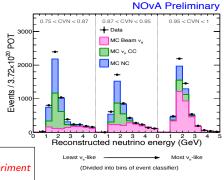
#### Decomposition and Extrapolation of Backgrounds in $\nu_e$ appearance

- We use deep learning to seperate the signal and backgrounds into 3 bins of ID confidence from least  $\nu_e$  like to the most  $\nu_e$  like.
- The  $\nu_e$  selected beam  $\nu_e$  CC,  $\nu_\mu$  CC and NC interactions in the ND are backgrounds in the FD.

- The ND spectrum is decomposed into NC, beam  $\nu_e$  CC and  $\nu_\mu$  CC.
- Each component is extrapolated from the ND to the FD for background predictions in the FD.

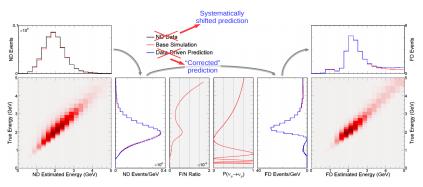
See talk:

Deep Learning Applications in the NOvA Experiment
- Fernanda Psihas



#### Extrapolation: Systematic uncertainties

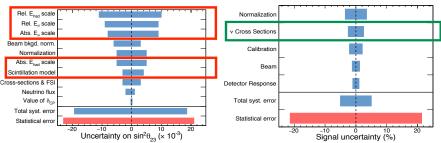
- Systematic uncertainties in the analysis are extrapolated from the ND to the FD using the same extrapolation technique.
- The ND Data is replaced by the ND prediction under a systematic shift.
- The corrected ND true energy spectra is extrapolated to the FD.



10 / 20

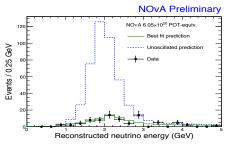
## Systematic uncertainties in the $\nu_{\mu}$ disappearance and $\nu_{\rm e}$ appearance analyses

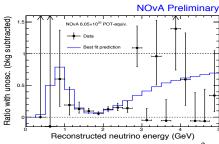
- Systematic uncertainties are included as nuisance parameters in the fit for all analyses.
- The functionally identical near and far detectors allow most uncertainties to cancel when predicting the FD spectrum.



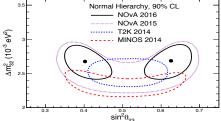
- In the disappearance analysis, uncertainties due to detector response are dominant but smaller than the statistical uncertainties.
- In the appearance analysis, cross section uncertainties are important for signal, but small compared to statistical uncertainties.

#### $u_{\mu}$ fit to the disappearance data



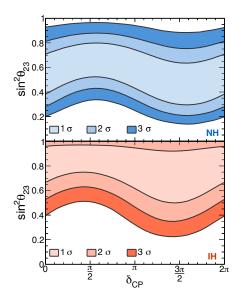


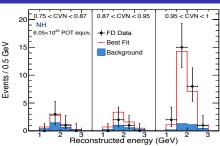




- Disappearance data is fit for the  $\Delta m^2_{32}$  and  $\sin^2 \theta_{23}$  parameters.
- Expected 473 events, observe 78 events, clear evidence of neutrino oscillations.
- Best fit in the Normal Hierarchy:  $\Delta m_{32}^2 = (2.67 \pm 0.11) \times 10^3 \text{ eV}^2$   $\sin^2 \theta_{23} = 0.404^{+0.030}_{-0.020} \text{ or } 0.624^{+0.022}_{-0.030}$
- -0.030
- Maximal-mixing disfavoured at 2.6  $\sigma$ .

#### $\nu_e$ fit to the appearance data





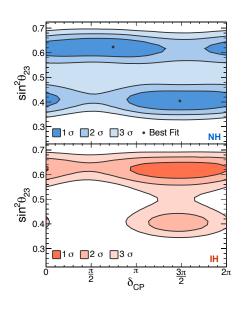
- Appearance data is fit for the  $\sin^2 \theta_{23}$  and  $\delta_{CP}$  parameters.
- Constrain parameters

$$\bullet \ \sin^2 2\theta_{13} = 0.085 \, \pm \, 0.005,$$

• 
$$\Delta m_{32}^2 = +2.44 \pm 0.06$$
 (NH)  
= -2.49 ± 0.06 (IH)

• Observe 33 events in the FD,  $> 8\sigma$  significance of  $\nu_e$  appearance.

#### Simultaneous fit of the disappearance and appearance data



- The constraints on the oscillation parameters can be improved by combining NOvA's  $\nu_e$  appearance data with its  $\nu_\mu$  disappearance data.
- Appearance and disappearance data are simultaneous fit for the  $\sin^2\theta_{23}$ ,  $\delta_{CP}$  and  $\Delta m_{32}^2$  parameters.
- $\sin^2 2\theta_{13} = 0.085 \pm 0.005$  is a constraint from the reactor experiments.
- Two statistically degenerate best fit points are in Normal Hierarchy:

$$\sin^2 \theta_{23} = 0.404, \ \delta_{CP} = 1.48\pi \ \text{and} \ \sin^2 \theta_{23} = 0.623, \ \delta_{CP} = 0.74\pi$$

• Inverted Hierarchy with lower  $\theta_{23}$  octant for all values of  $\delta_{CP}$  is rejected at >93% CE.

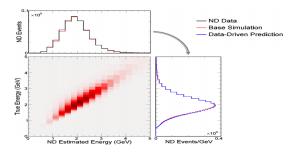
#### Summary

- NOvA is a two detector, long baseline, accelerator based neutrino oscillation experiment.
- The Near Detector measures NuMI beam before oscillations and helps in signal and background predictions in the Far Detector.
- Having two functionally identical detectors is advantageous for cancelling uncertainties that are common between the two detectors.
- Best fit to  $\nu_{\mu}$  disappearance data is a non-maximal value of  $\theta_{23}$ , maximal mixing disfavoured at 2.6  $\sigma$ .
- A simultaneous fit of the disappearance and appearance data rejects Inverted Hierarchy with lower  $\theta_{23}$  octant for all values of  $\delta_{CP}$  at >93% CL.
- Updated oscillation results soon with 50% more data.

### Backup

#### Extrapolation: ND Spectra

- The P( $\nu_{\alpha} \rightarrow \nu_{\beta}$ ) depends on  $E_{true}$ , but detectors measure  $E_{reco}$ .
- Detectors/reconstruction have different sensitivities to different processes, which have different  $E_{true} \leftrightarrow E_{reco}$ .

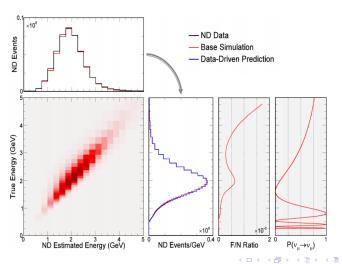


- The reconstructed ND  $\nu_{\mu}$  CC energy spectrum is used to correct the FD simulated prediction.
- True energy distribution is corrected so that reconstructed data & MC agree at the ND.

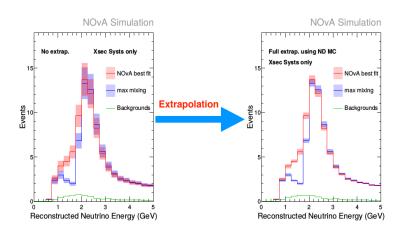
NOvA Analysis

#### Extrapolation: Far over Near ratios

• The FD and ND MC spectra along with the oscillation probability for the  $\nu_{\mu}$  ( $\bar{\nu}_{\mu}$ )  $\rightarrow$   $\nu_{\mu}$  ( $\bar{\nu}_{\mu}$ ) disappearance and  $\nu_{\mu}$  ( $\bar{\nu}_{\mu}$ )  $\rightarrow$   $\nu_{e}$  ( $\bar{\nu}_{e}$ ) appearance are used to extrapolate corrected ND true spectra to the FD.

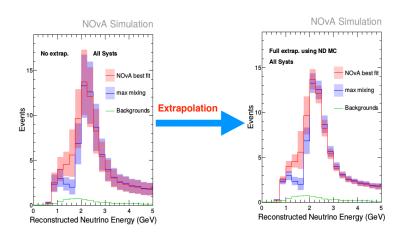


#### Extrapolation: all XS uncertainties



- These plots show the effect of cross-section uncertainty with and without extrapolation.
- The cross-section uncertainty is reduced when we use extrapolation method.

#### Extrapolation: all $\nu_{\mu}$ uncertainties



- ullet These plots show the effect of all  $u_{\mu}$  uncertainties with and without extrapolation.
- The uncertainties are reduced when we use extrapolation method.