

# Searching for Milky Way Satellite Galaxies with DECam

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David N. Schramm Fellow  
Fermilab

DPF 2017  
July 31, 2017

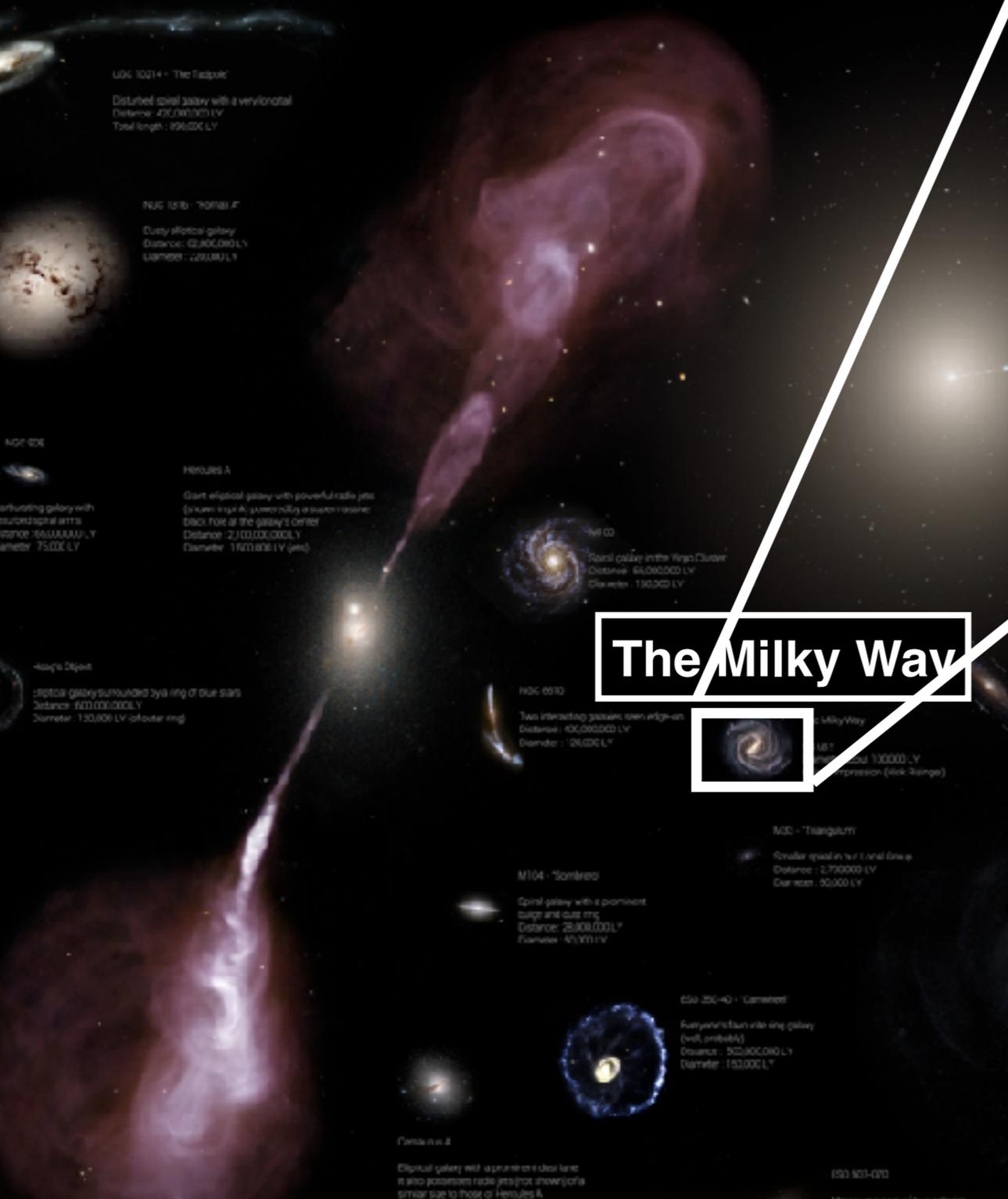


# What are Dwarf Galaxies?

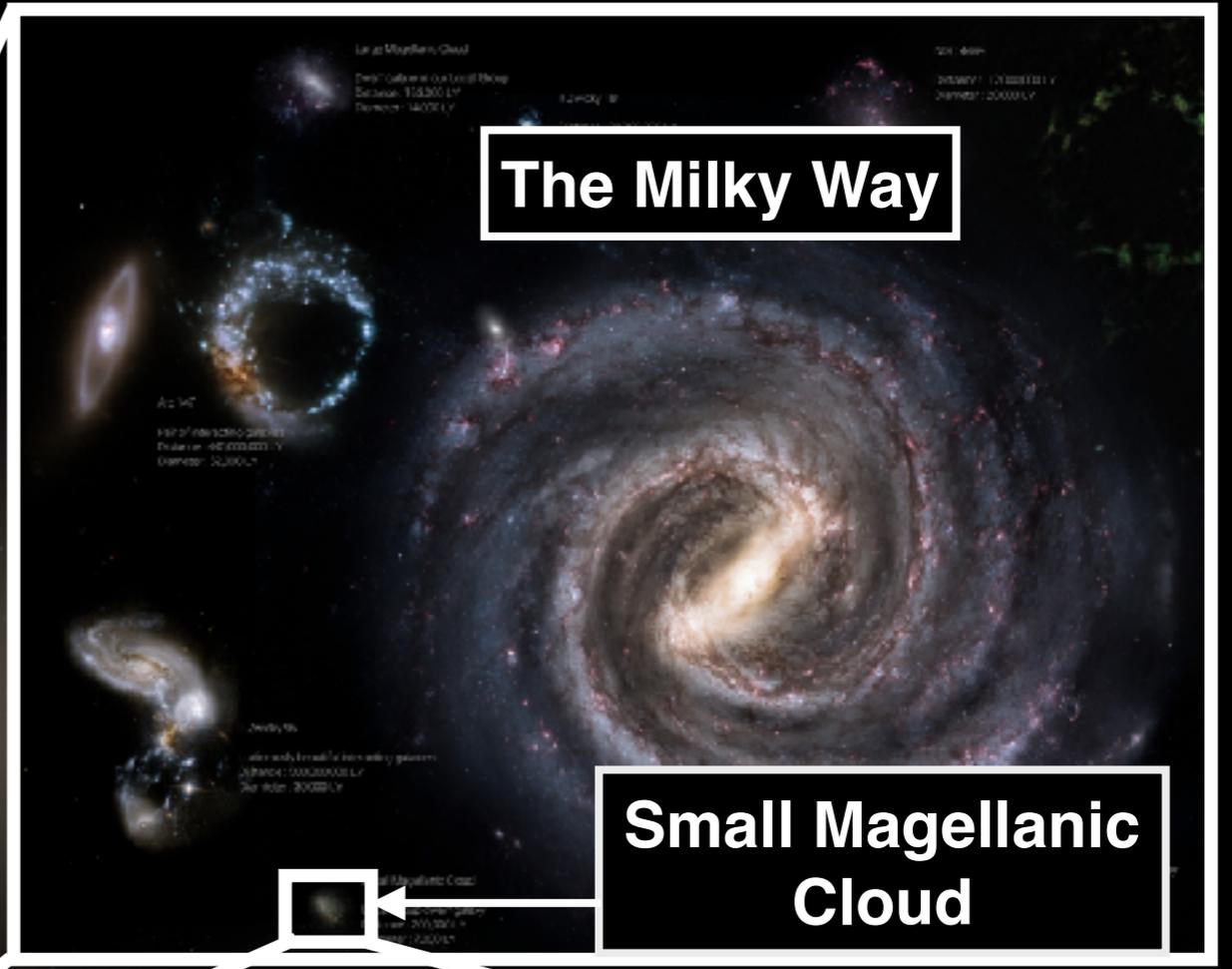




# What are Dwarf Galaxies?

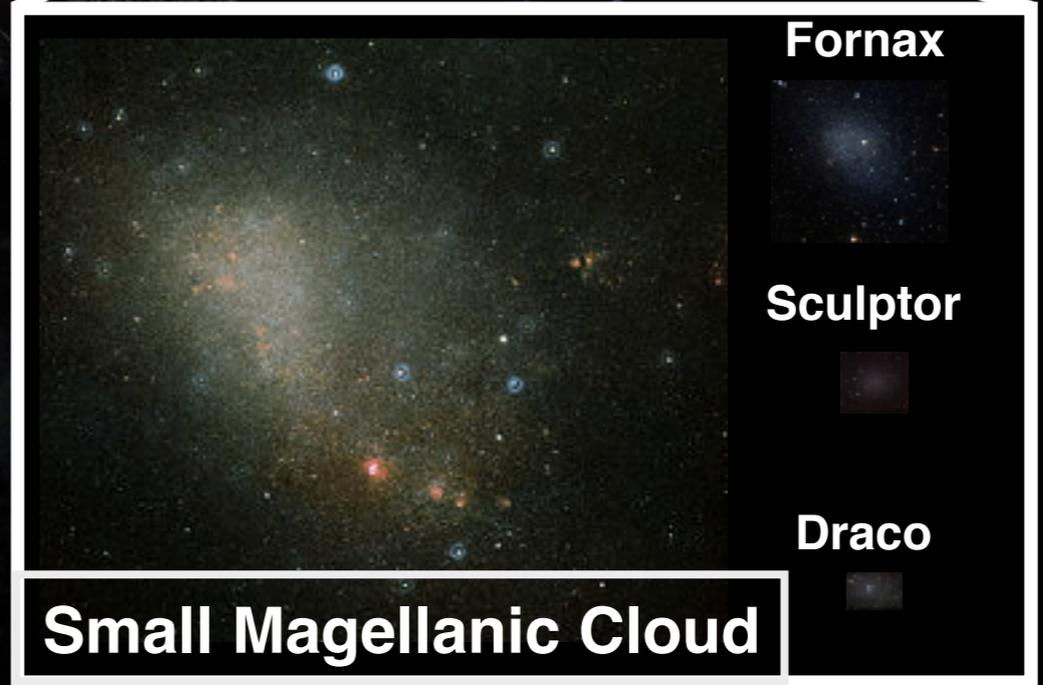


The Milky Way



The Milky Way

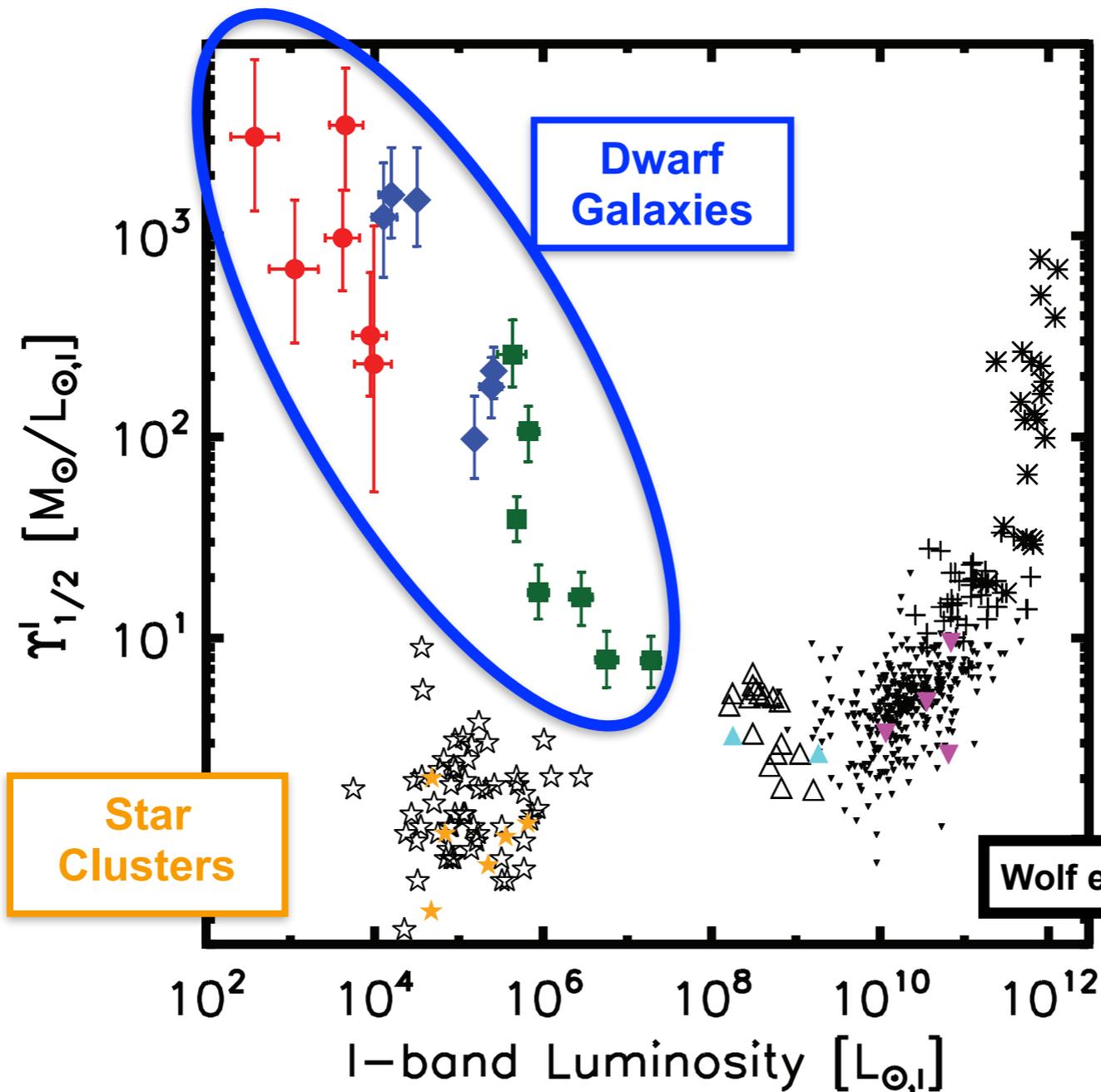
Small Magellanic Cloud



Small Magellanic Cloud

# Why are dwarf galaxies important?

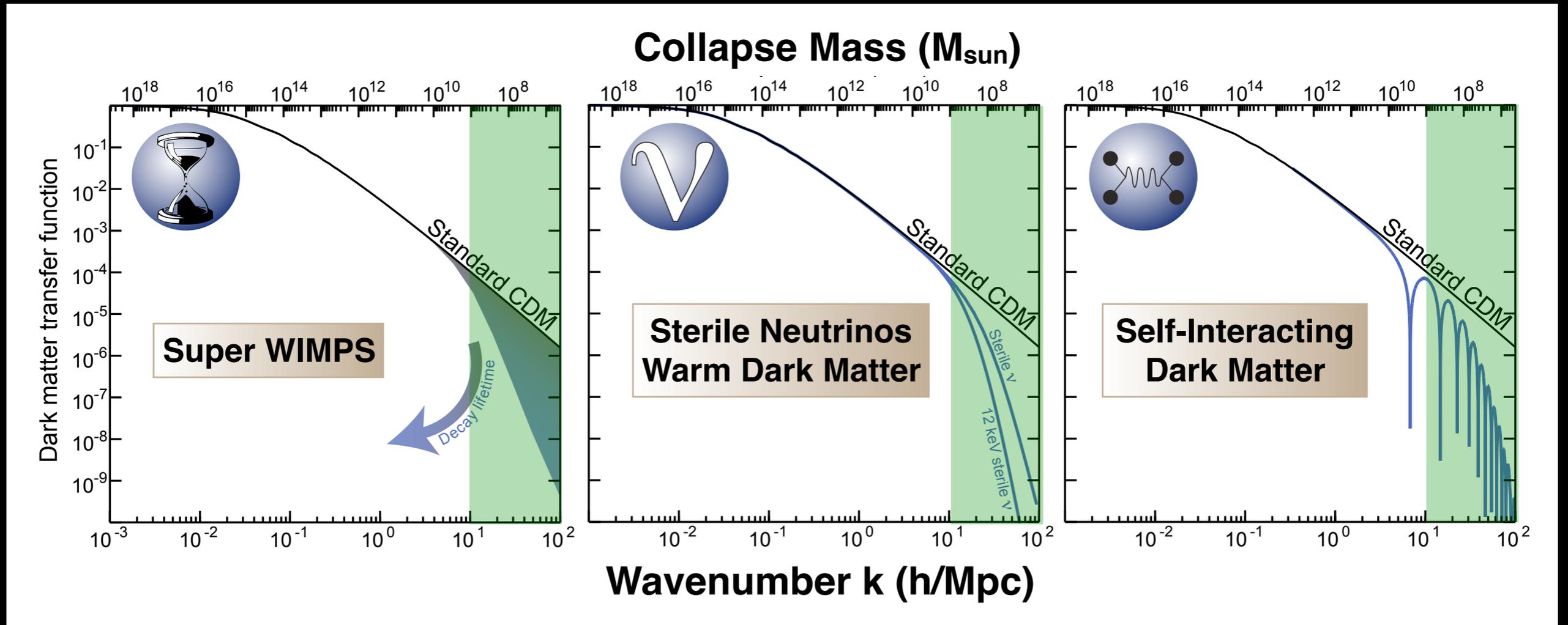
(Gravitational Mass)  
Visible Mass



Dwarf galaxies are the most dark-matter-dominated objects known

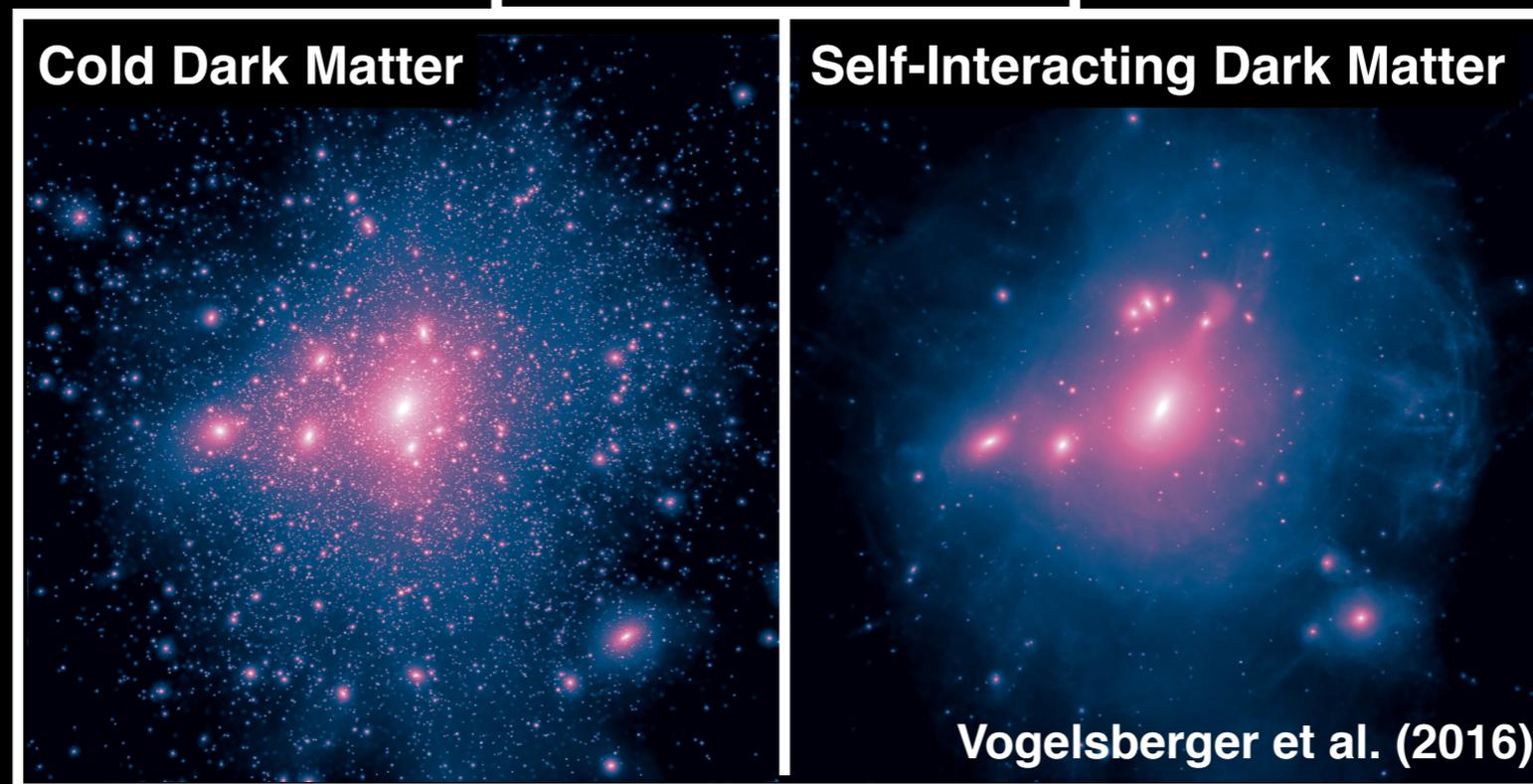
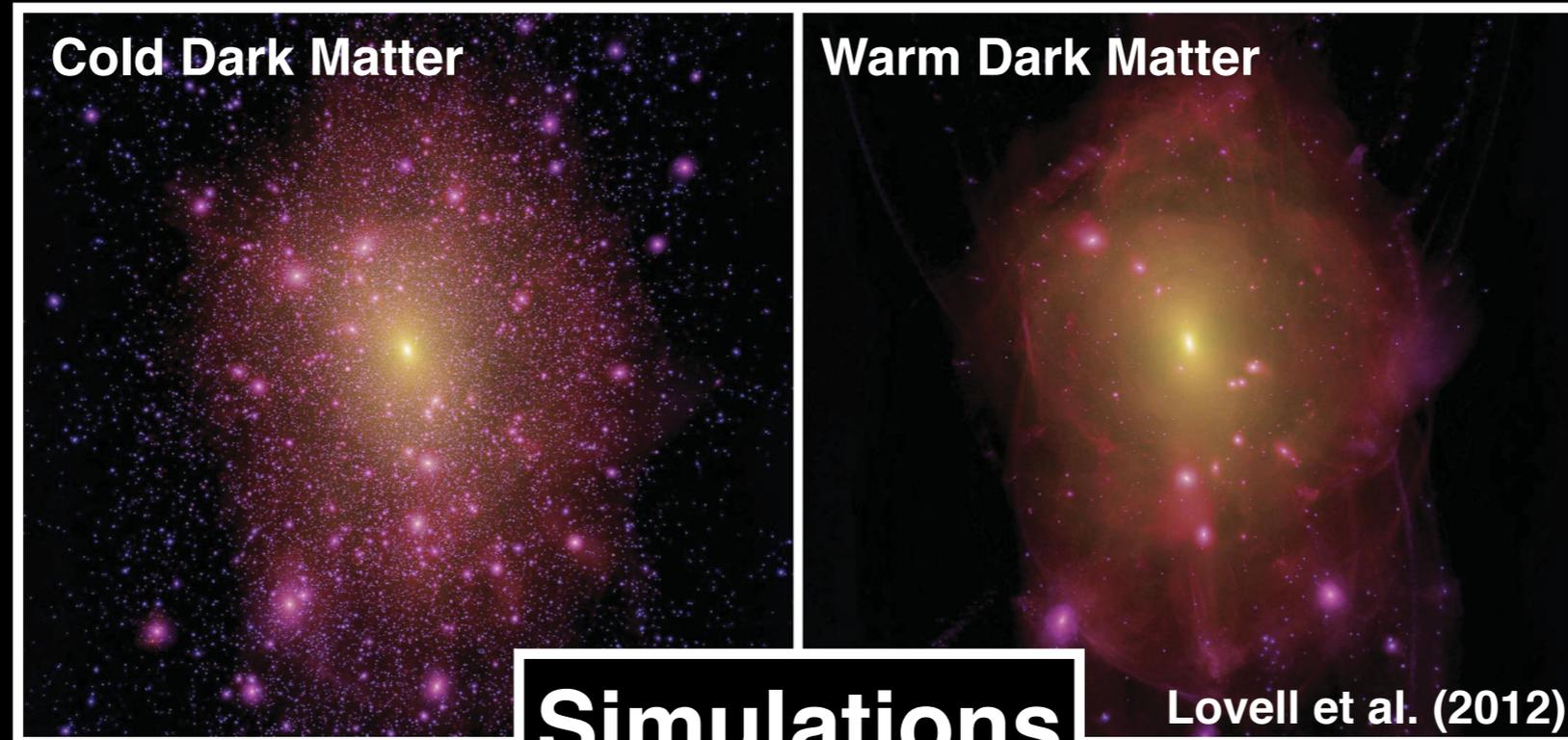
“Brightness” →

# Smallest Structures Probe Fundamental Characteristics of Dark Matter



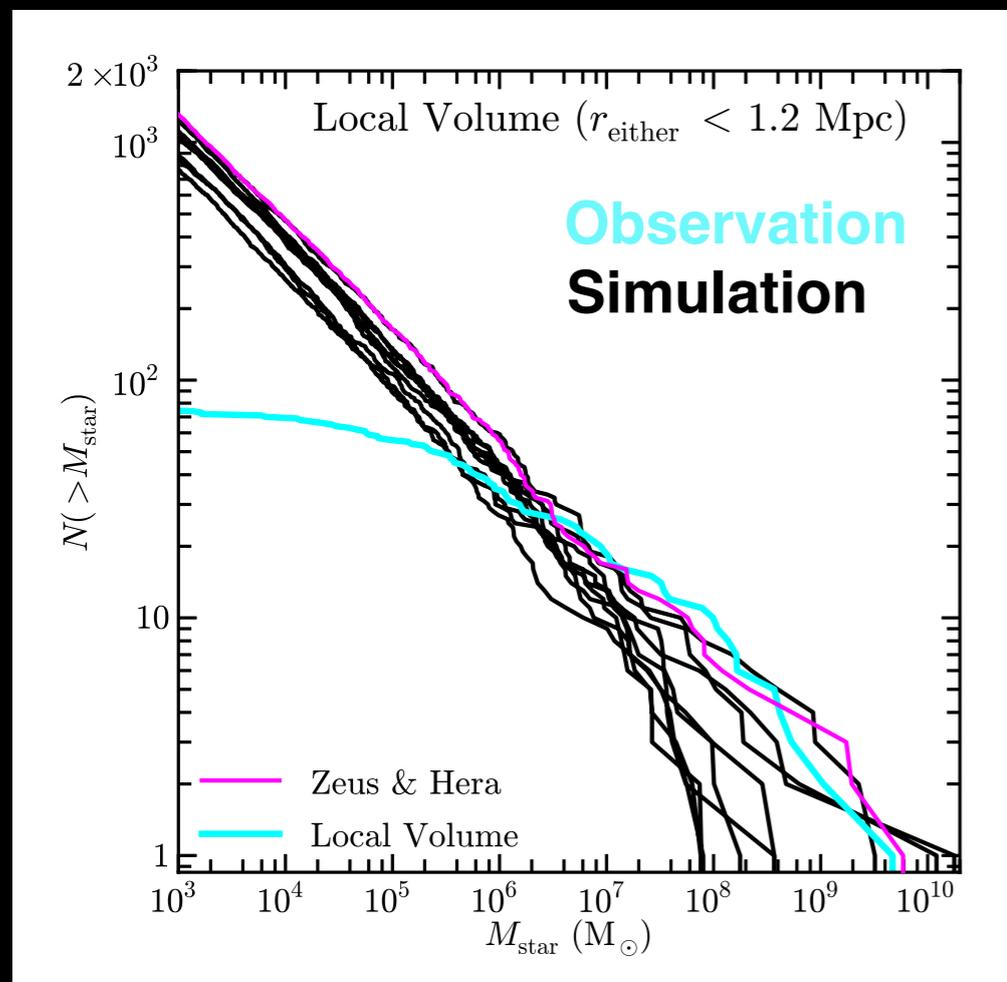
**Deviations from Cold Dark Matter could be detected in the abundance and densities of the smallest structures.**

# Smallest Structures Probe Fundamental Characteristics of Dark Matter



# Observational Challenges to Cold Dark Matter

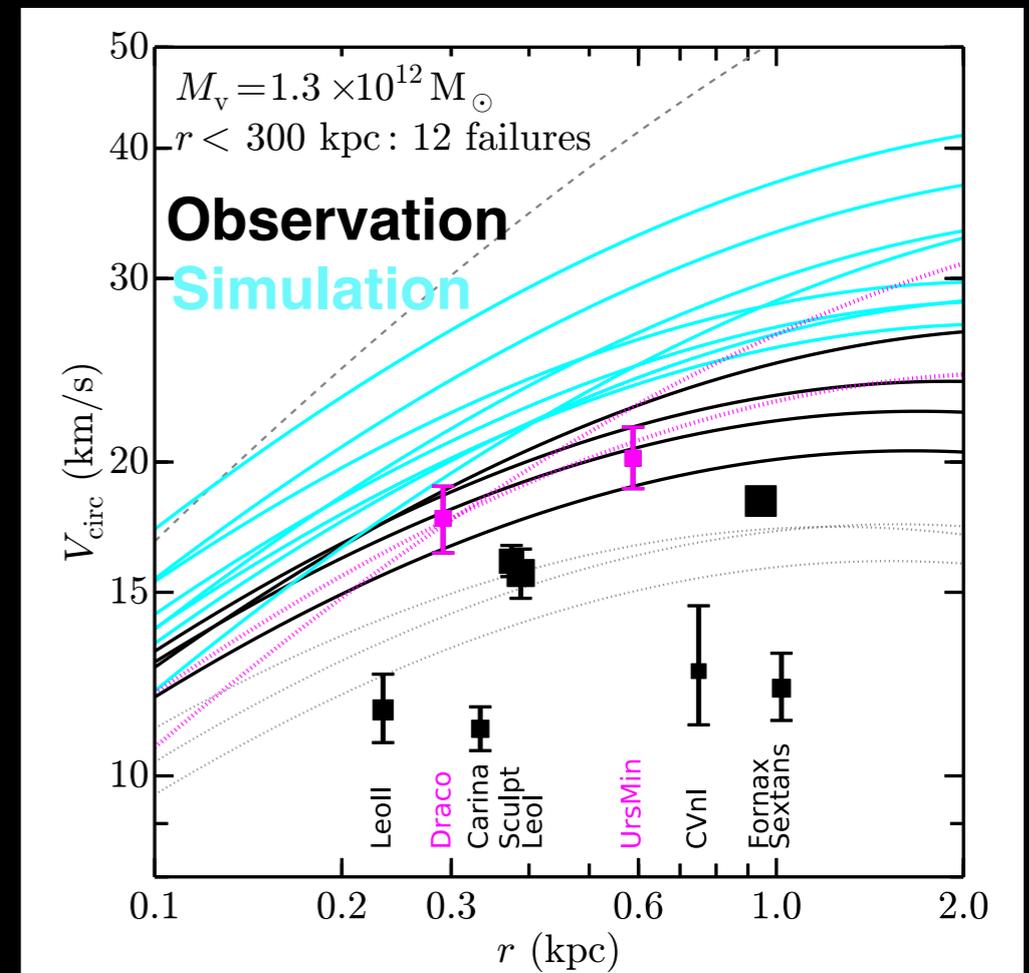
## The “Missing Satellites” Problem



Garrison-Kimmel et al. (2014)

Fewer satellites are observed compared to simulations

## The “Too Big to Fail” Problem

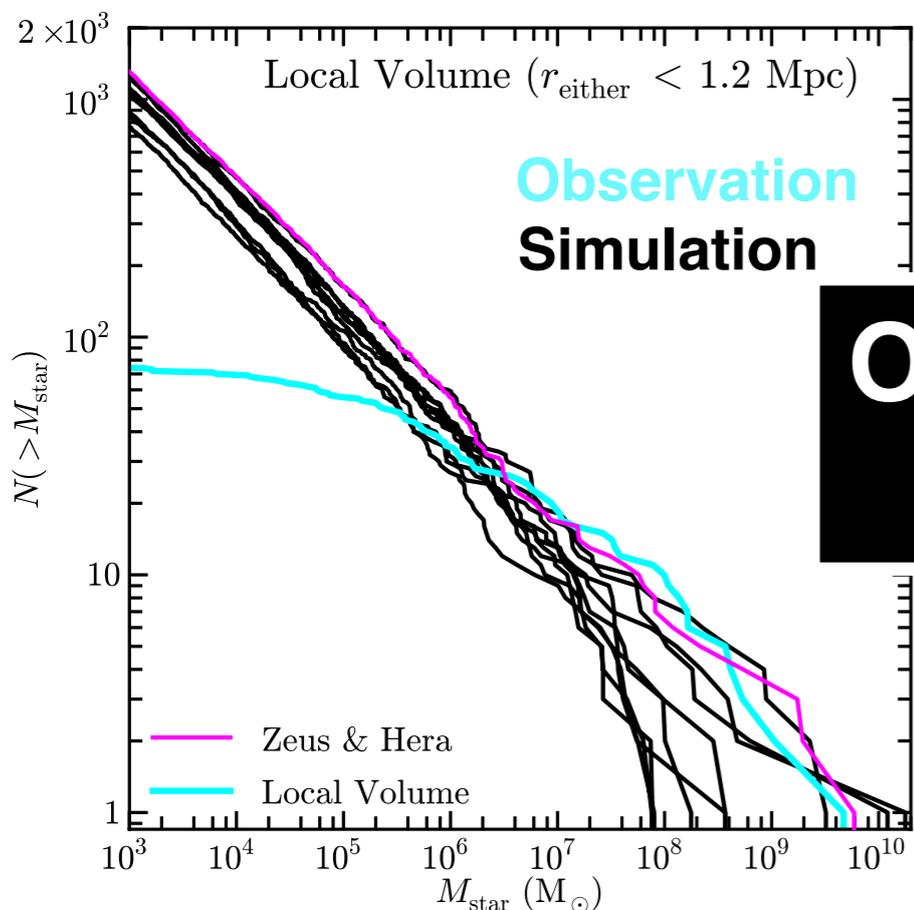


Garrison-Kimmel et al. (2014)

Observed satellites are under-dense compared to simulations

# Observational Challenges to Cold Dark Matter

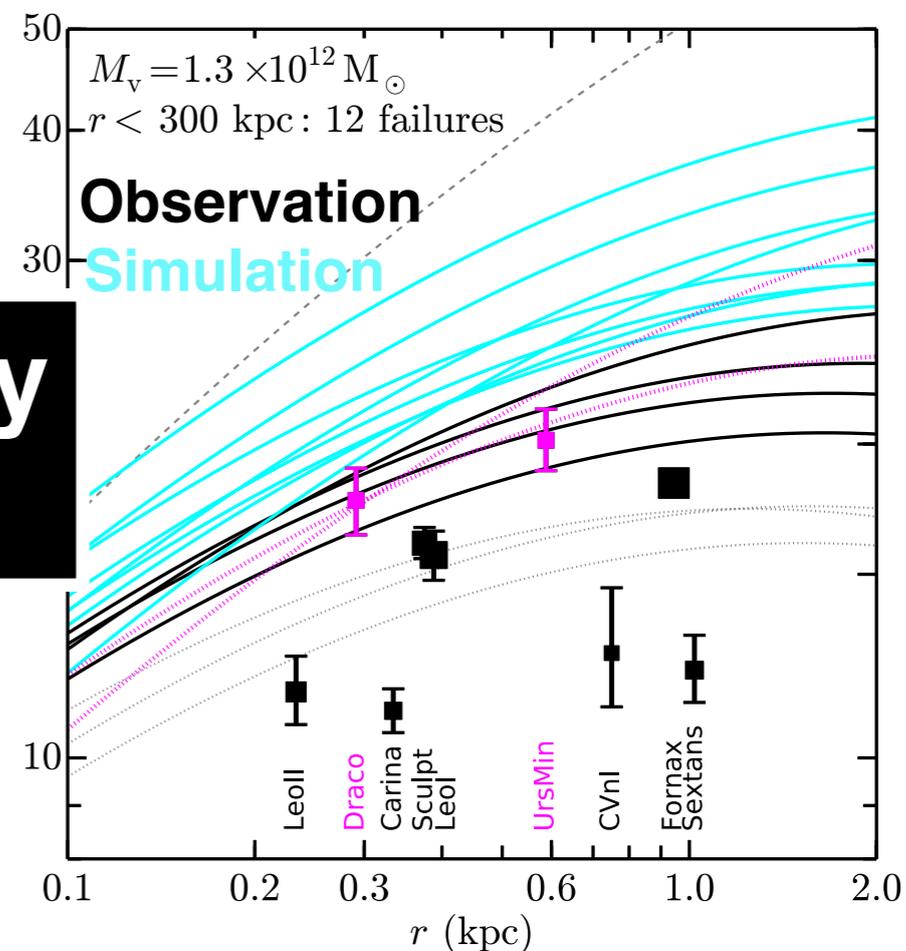
## The “Missing Satellites” Problem



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## The “Too Big to Fail” Problem



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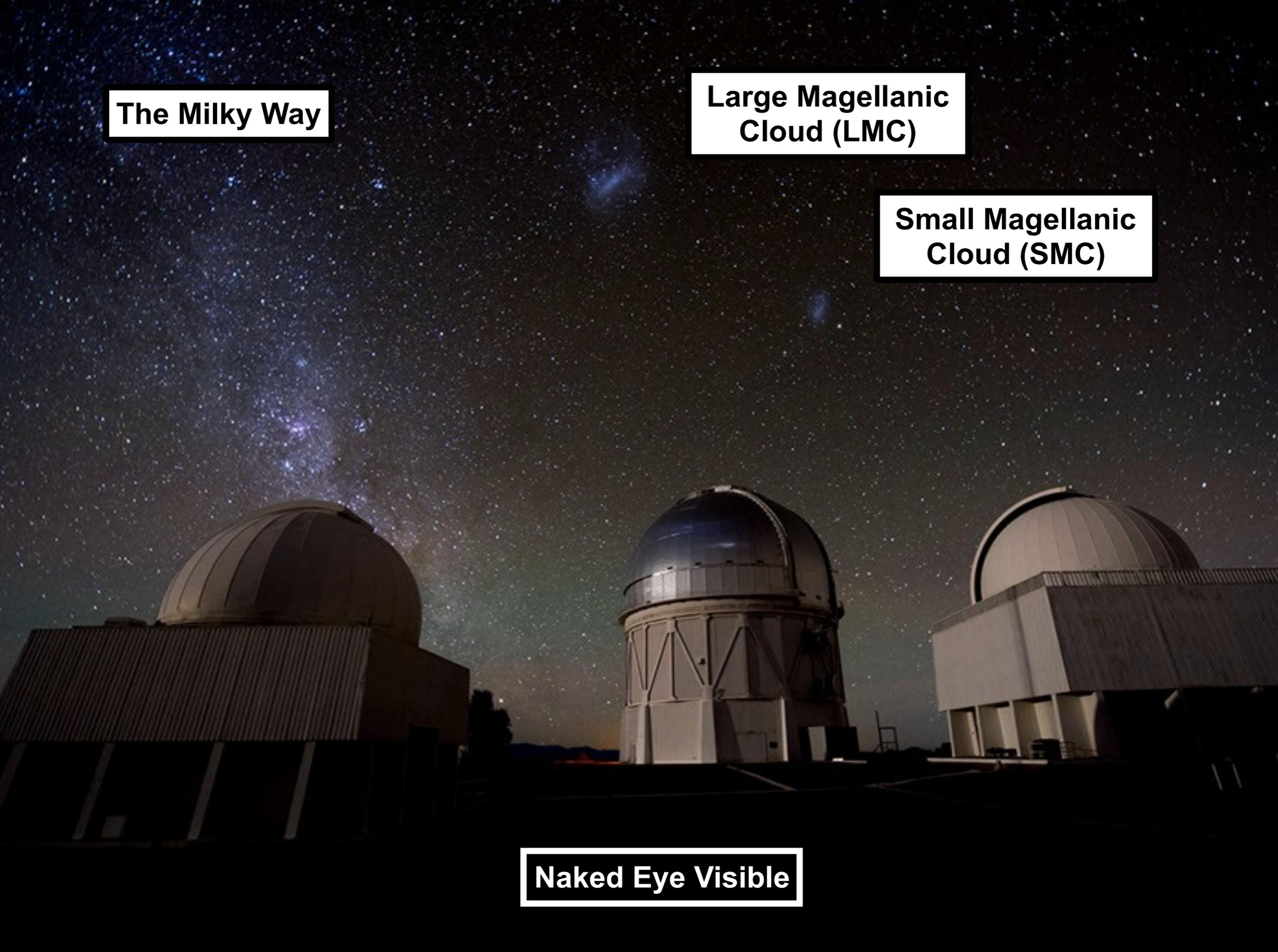
**Observationally Limited!**

**The Milky Way**

**Large Magellanic  
Cloud (LMC)**

**Small Magellanic  
Cloud (SMC)**

**Naked Eye Visible**



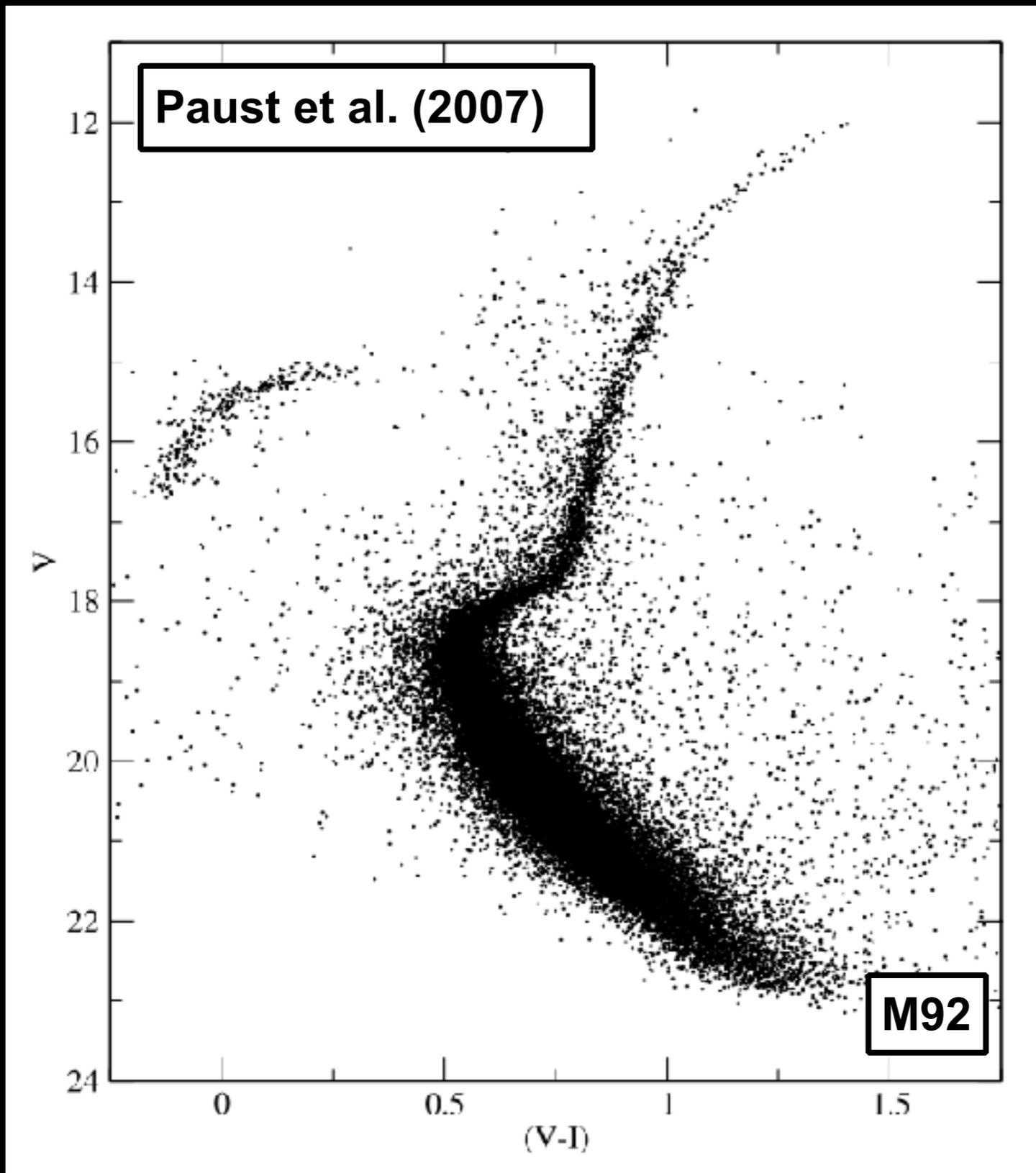
# Sculptor Dwarf Galaxy



ESO/DSS2

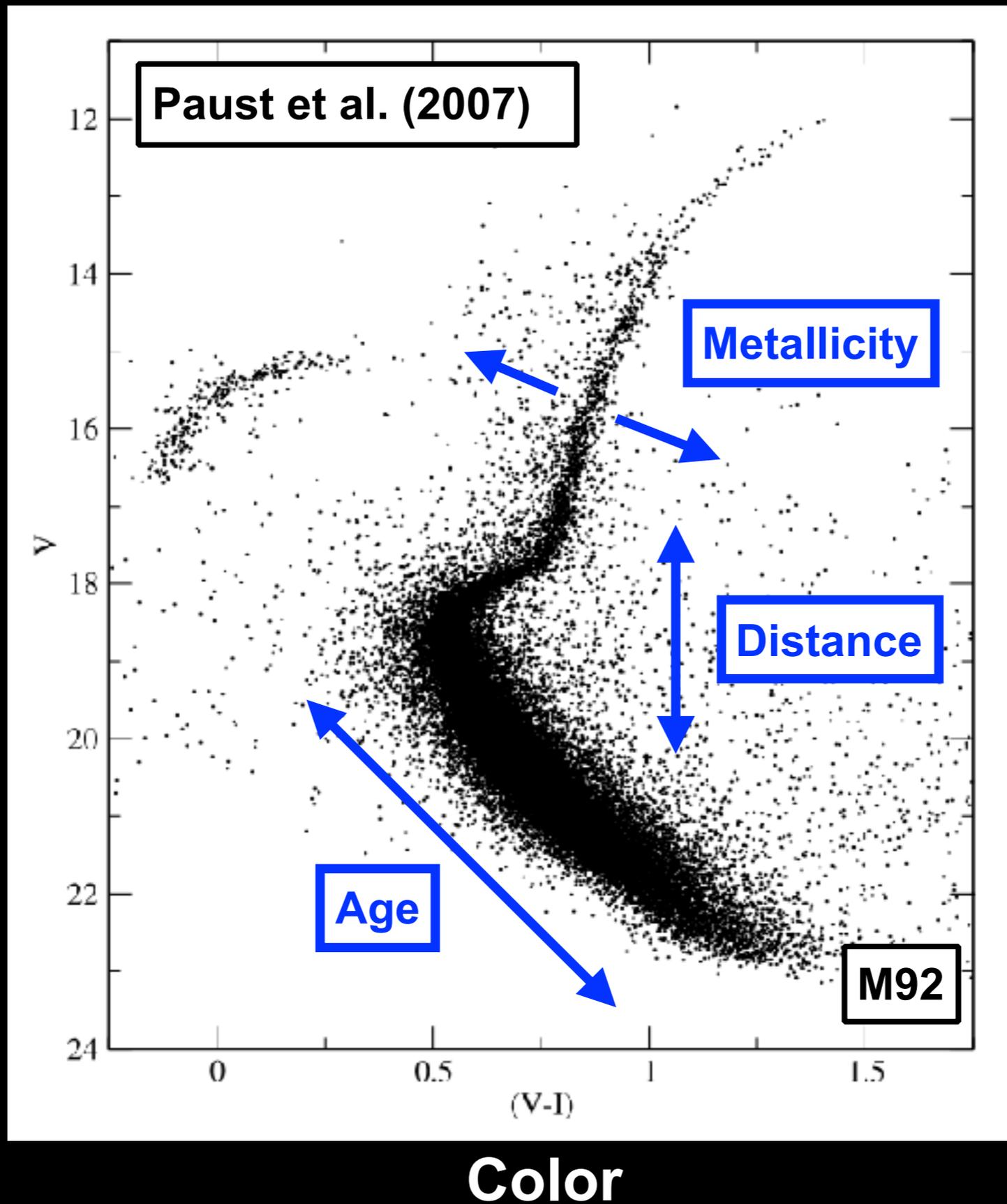
1.2m Telescope  
Photographic Plates

Brighter  
↑  
Magnitude  
↓  
Fainter



Color  
Bluer ← → Redder

Brighter  
↑  
Magnitude  
↓  
Fainter



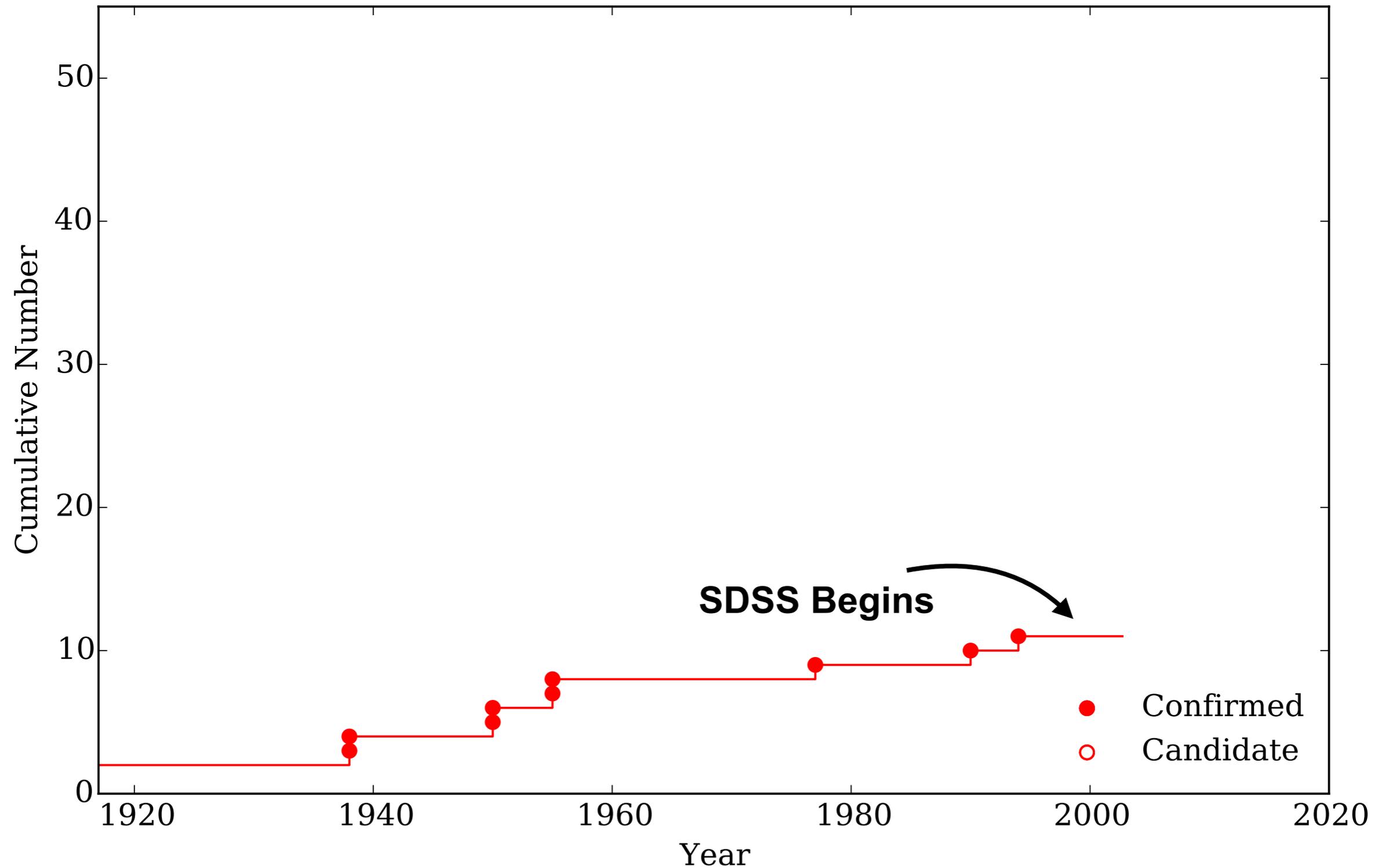
**Measure:**

- Age
- Metallicity
- Distance

**NOTE: We can't measure dark matter content from photometry alone...**

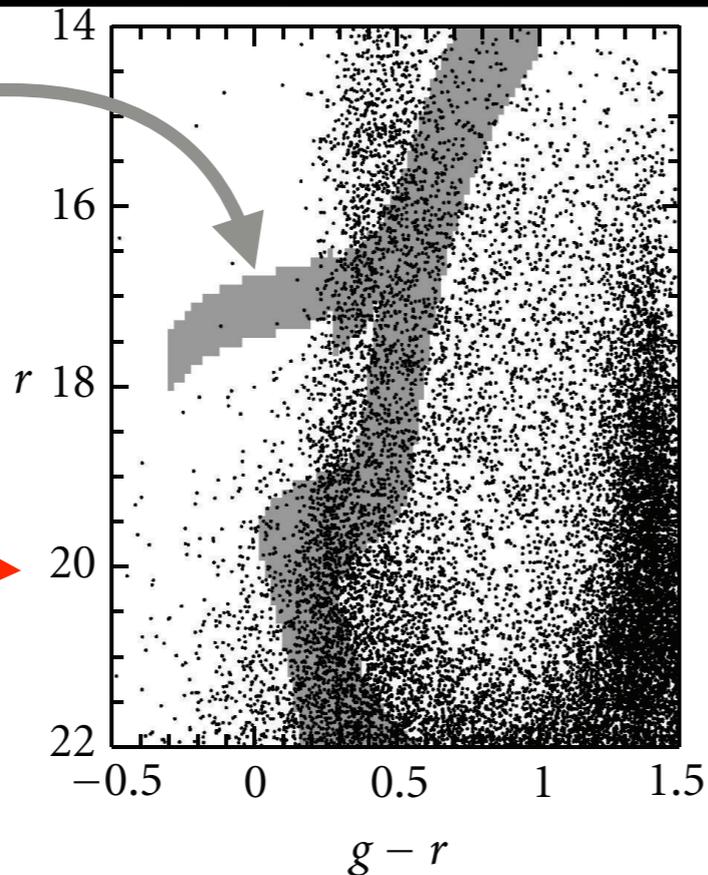
**Spectroscopy talk by Ting Li**

Bluer ← Color → Redder



Koposov et al. (2008)  
Walsh et al. (2009)  
Willman et al. (2010)

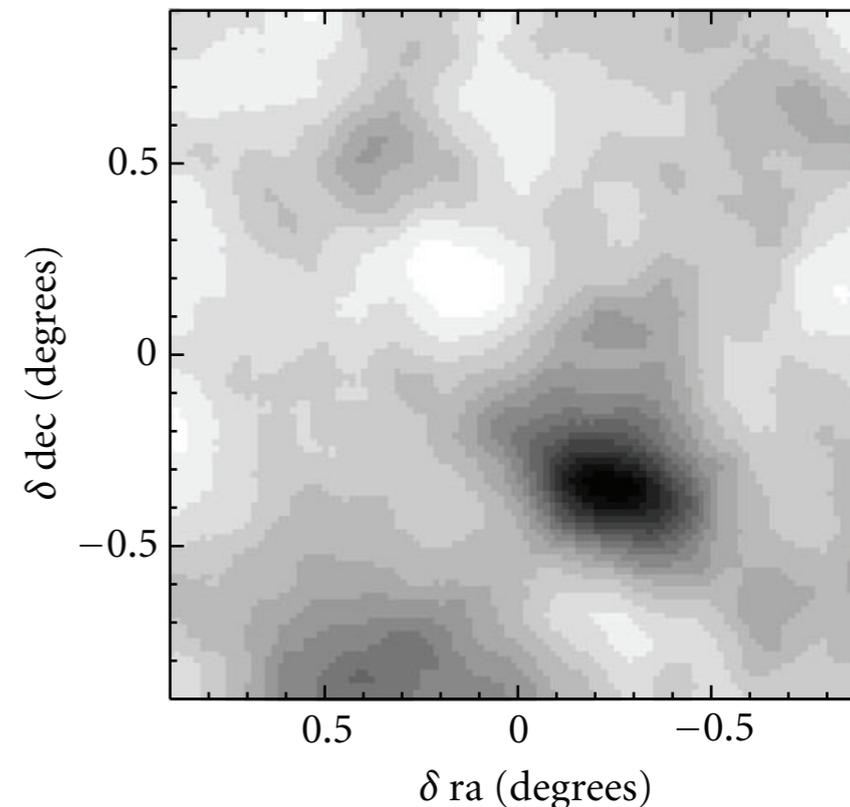
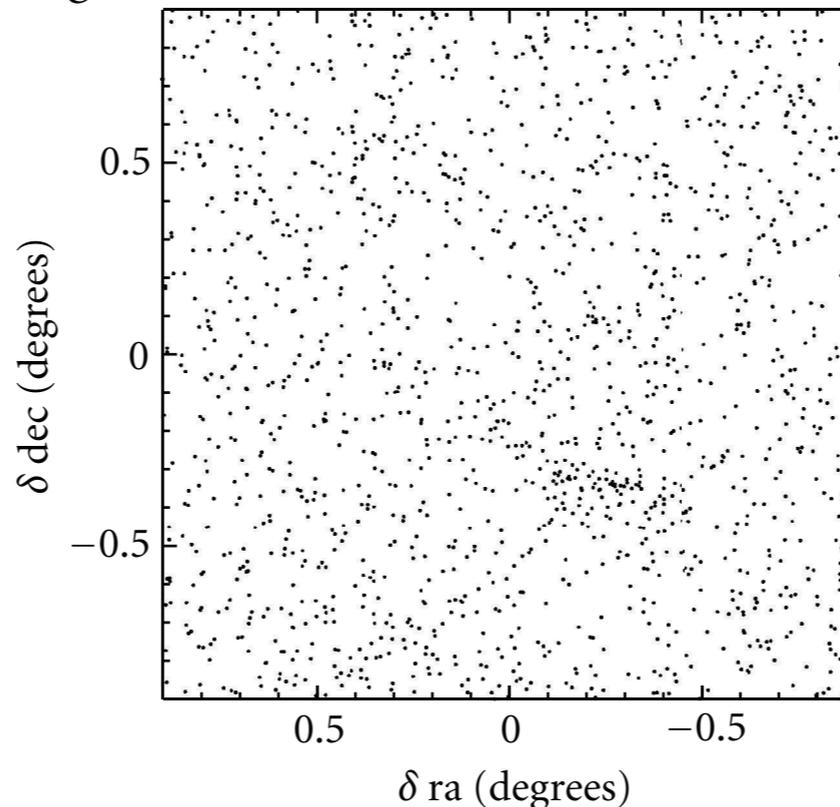
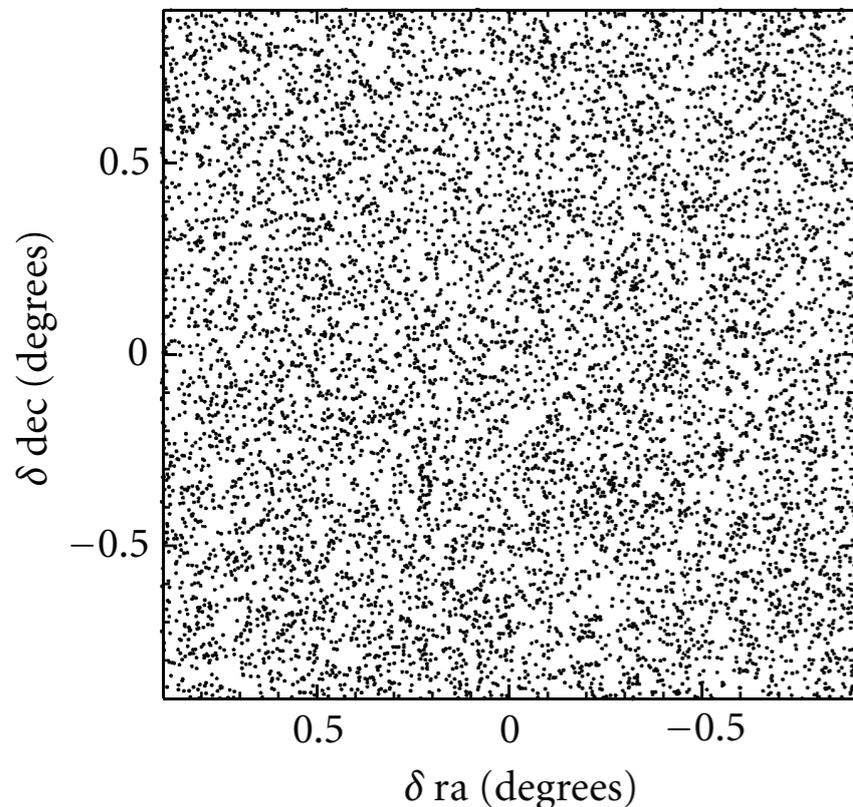
Stellar Isochrone



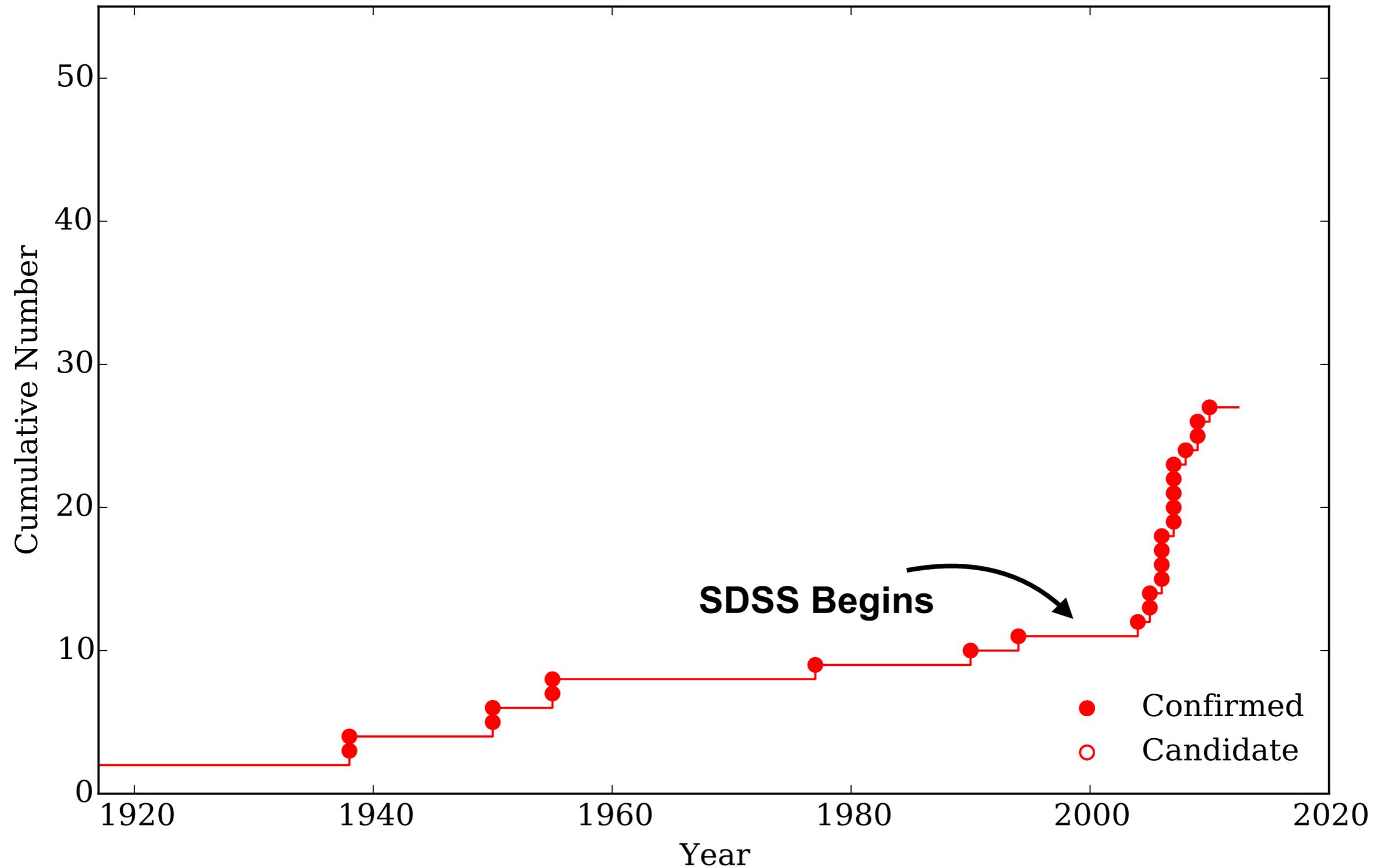
1) Start with a large catalog of stars

2) Apply a selection in color-magnitude space based on a stellar isochrone

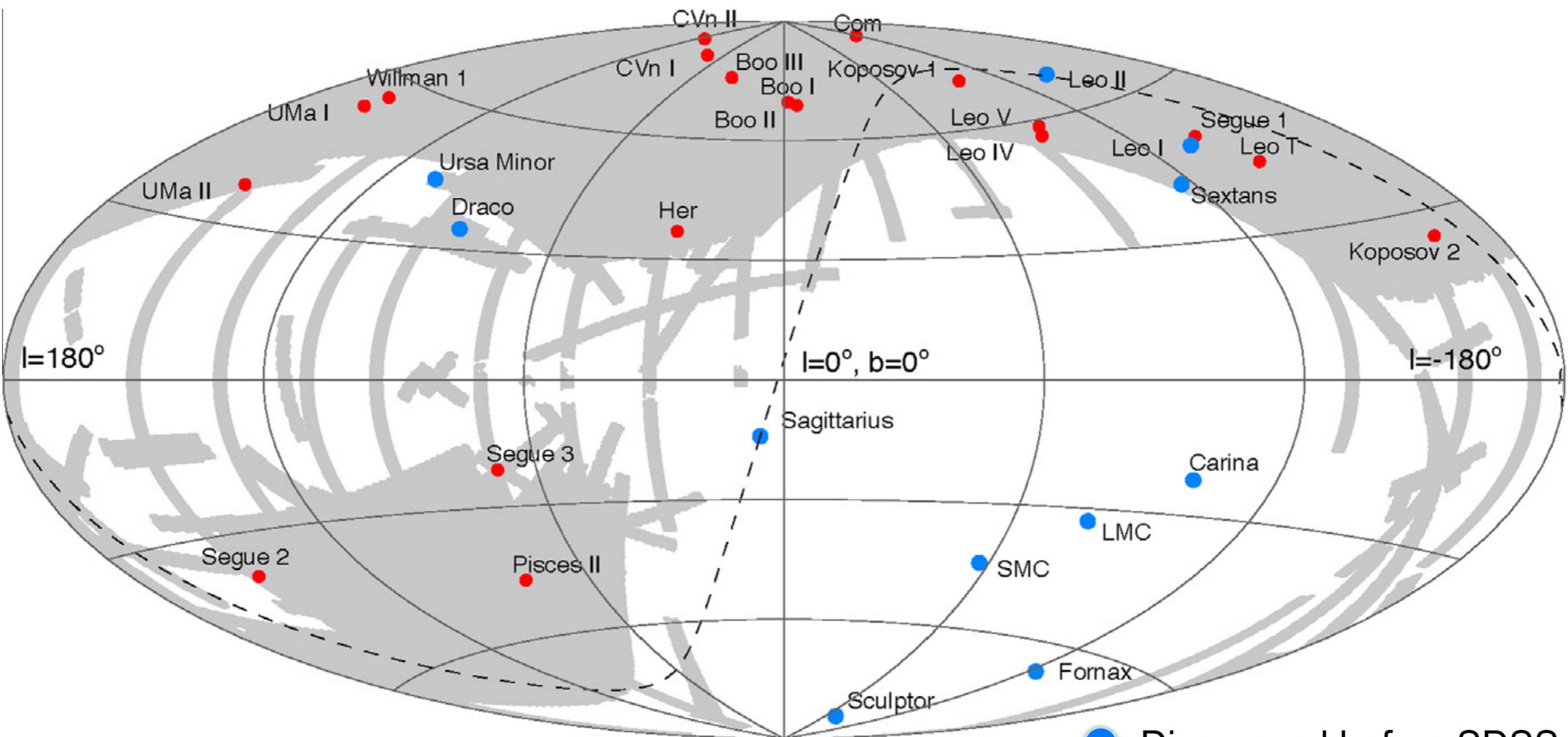
3) Convolve with a spatial kernel



# Dwarf Galaxy Discovery Timeline



# SDSS Sky Coverage

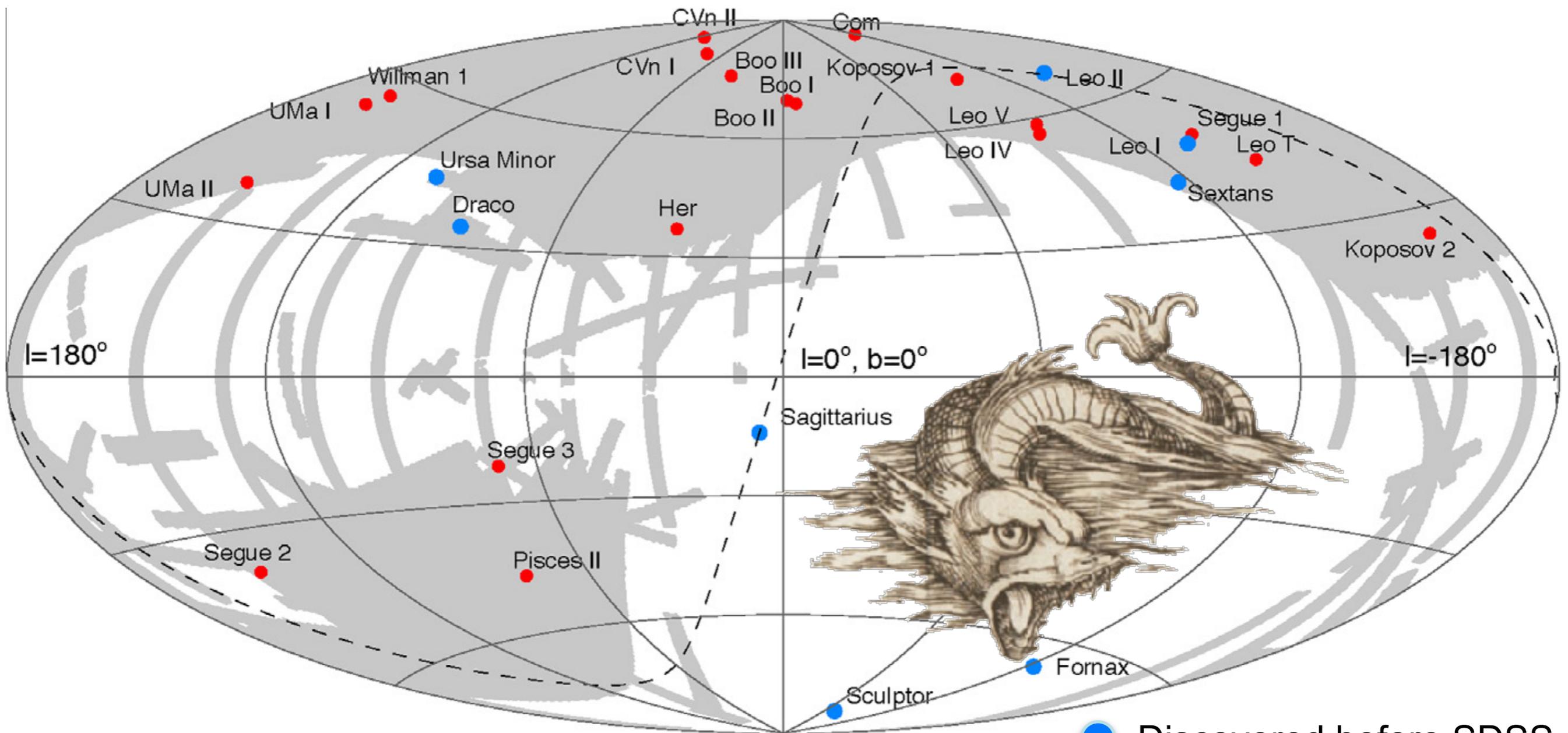


☐ Sky Covered by SDSS

- Discovered before SDSS (classical dwarfs)
- Discovered with SDSS (ultra-faint dwarfs)

(Belokurov 2013)

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# The Dark Energy Survey

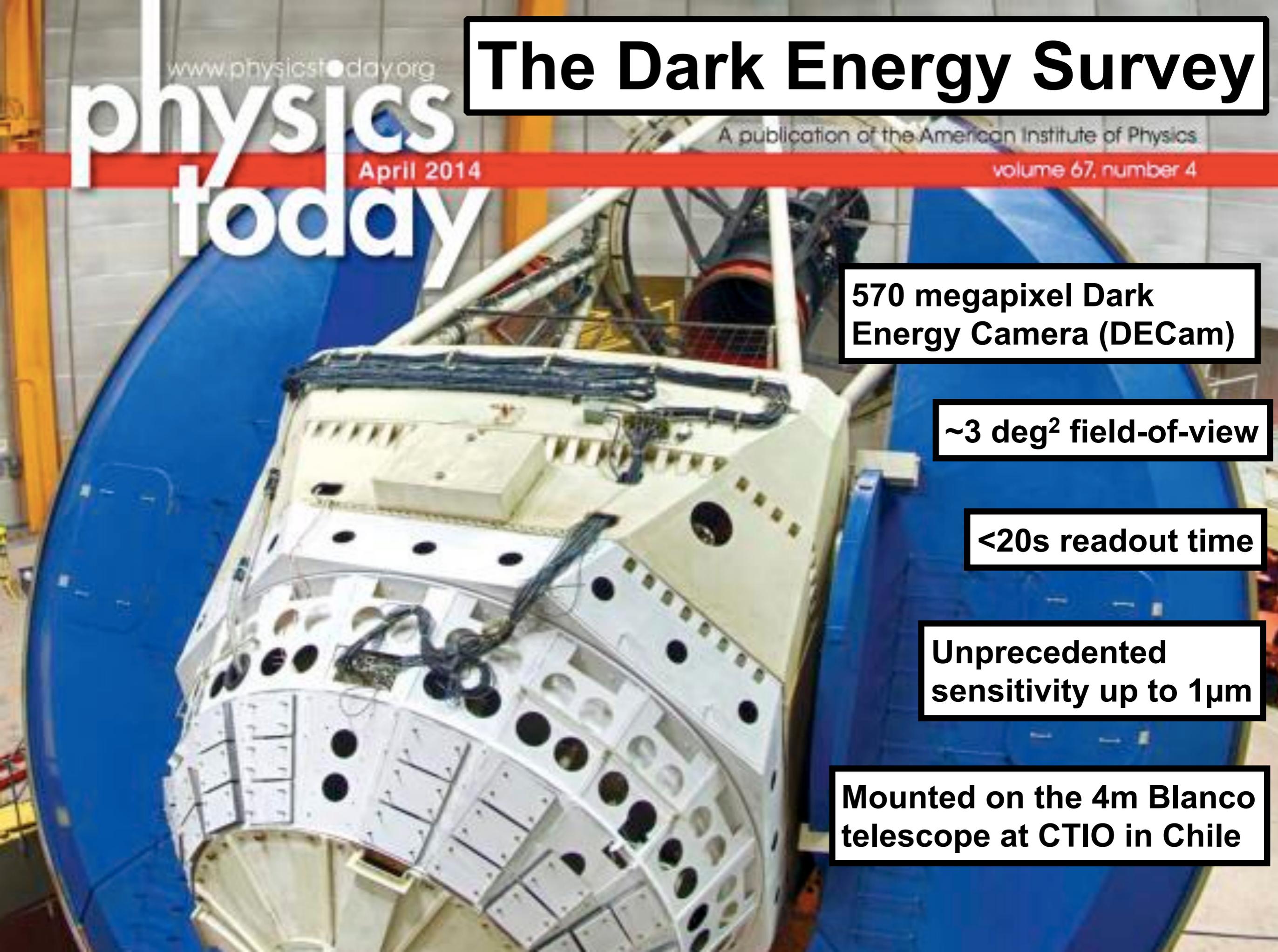
**570 megapixel Dark Energy Camera (DECam)**

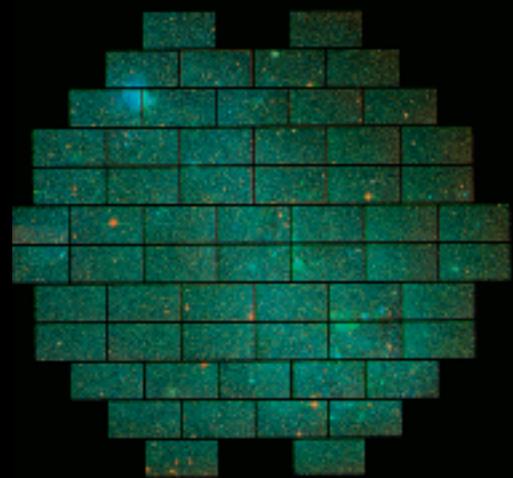
**~3 deg<sup>2</sup> field-of-view**

**<20s readout time**

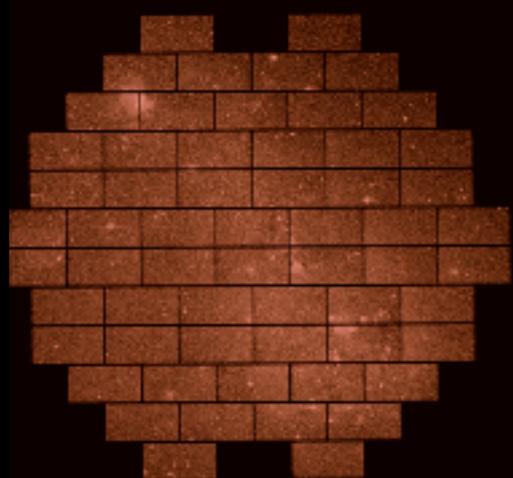
**Unprecedented sensitivity up to 1 $\mu$ m**

**Mounted on the 4m Blanco telescope at CTIO in Chile**

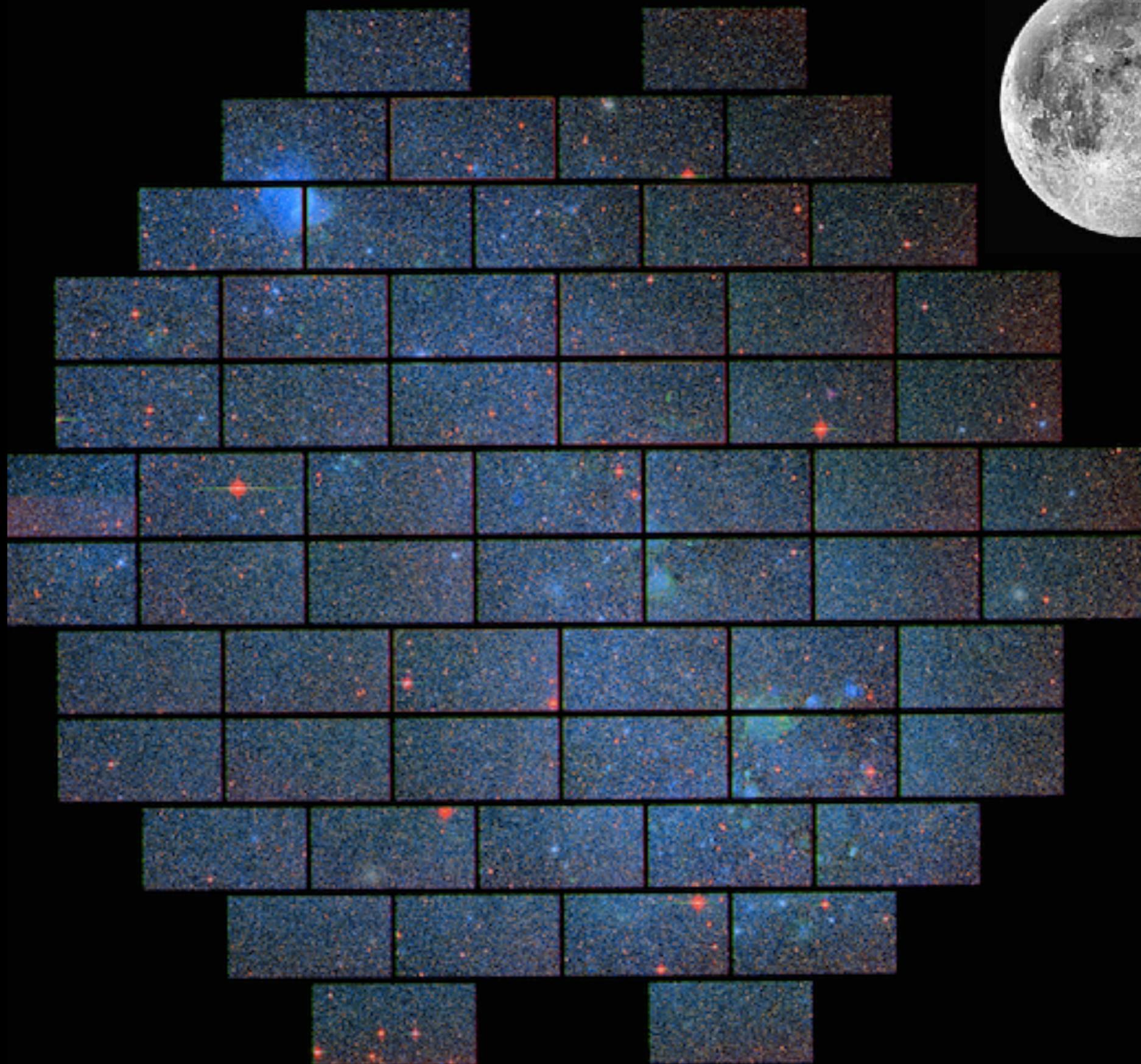
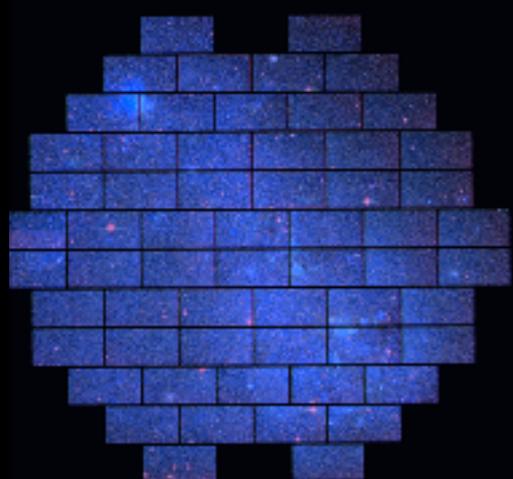


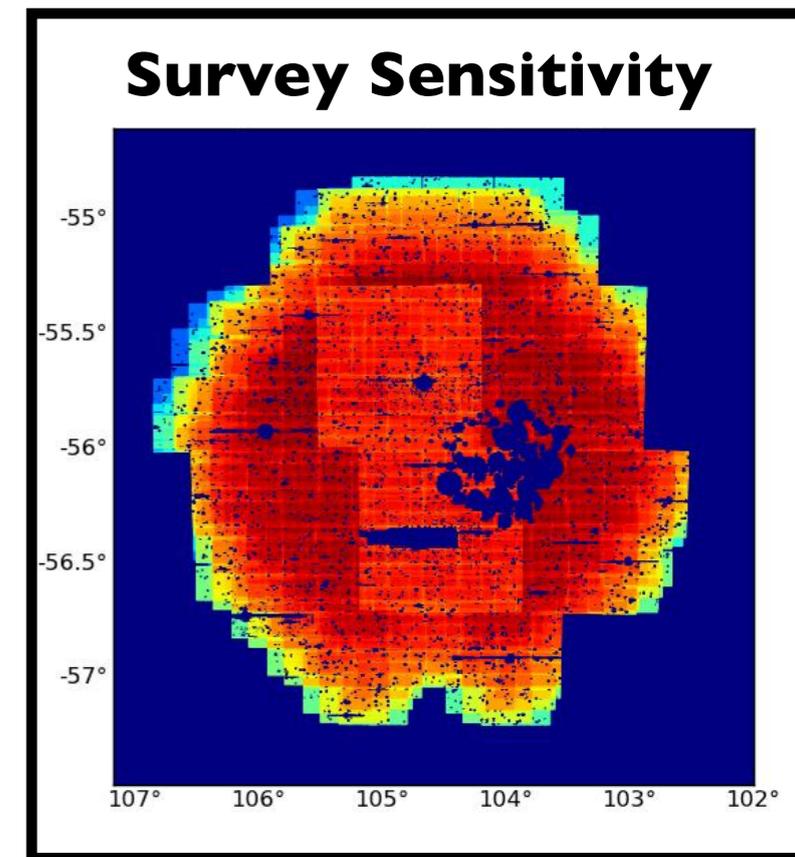
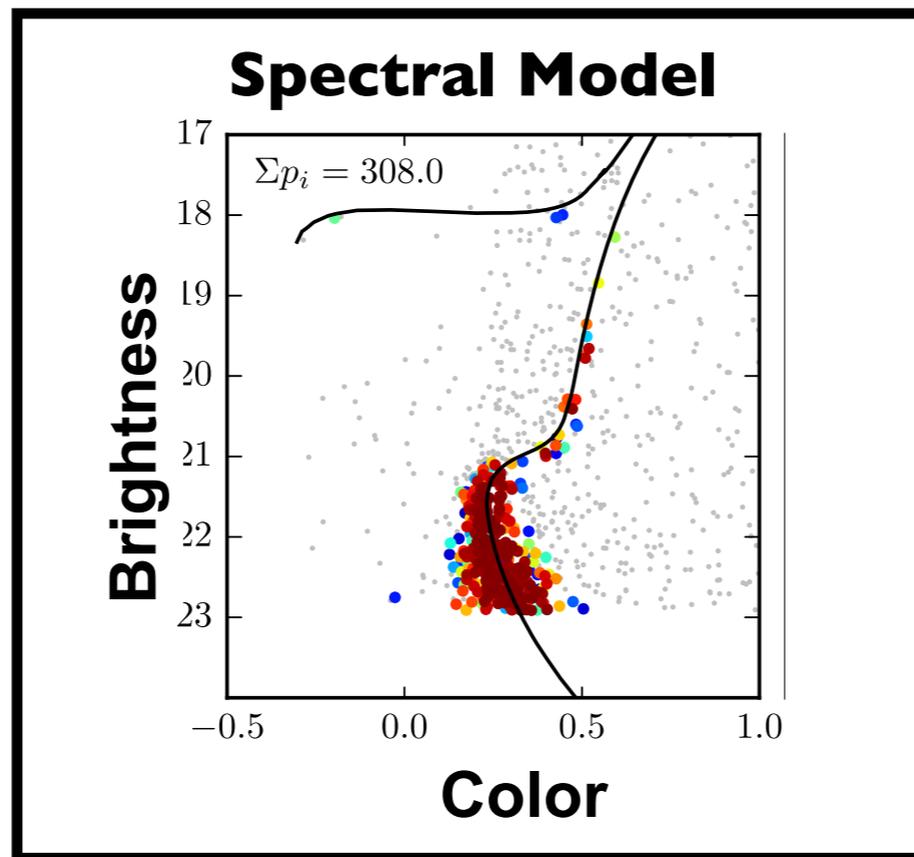
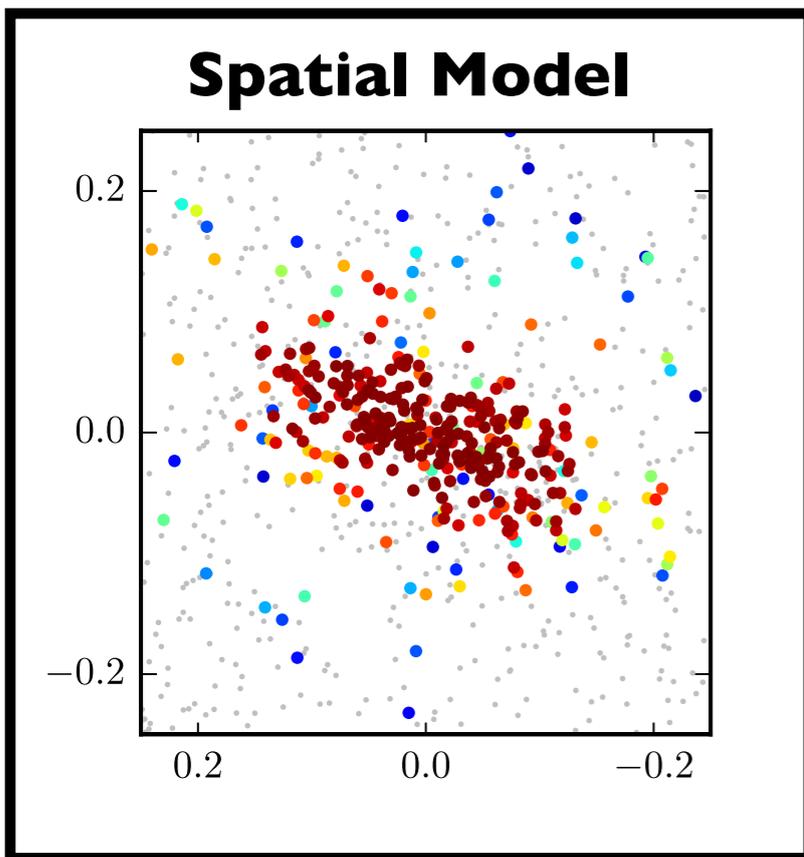


+



+





$$p_i = \frac{\lambda u_i}{\lambda u_i + b_i}$$

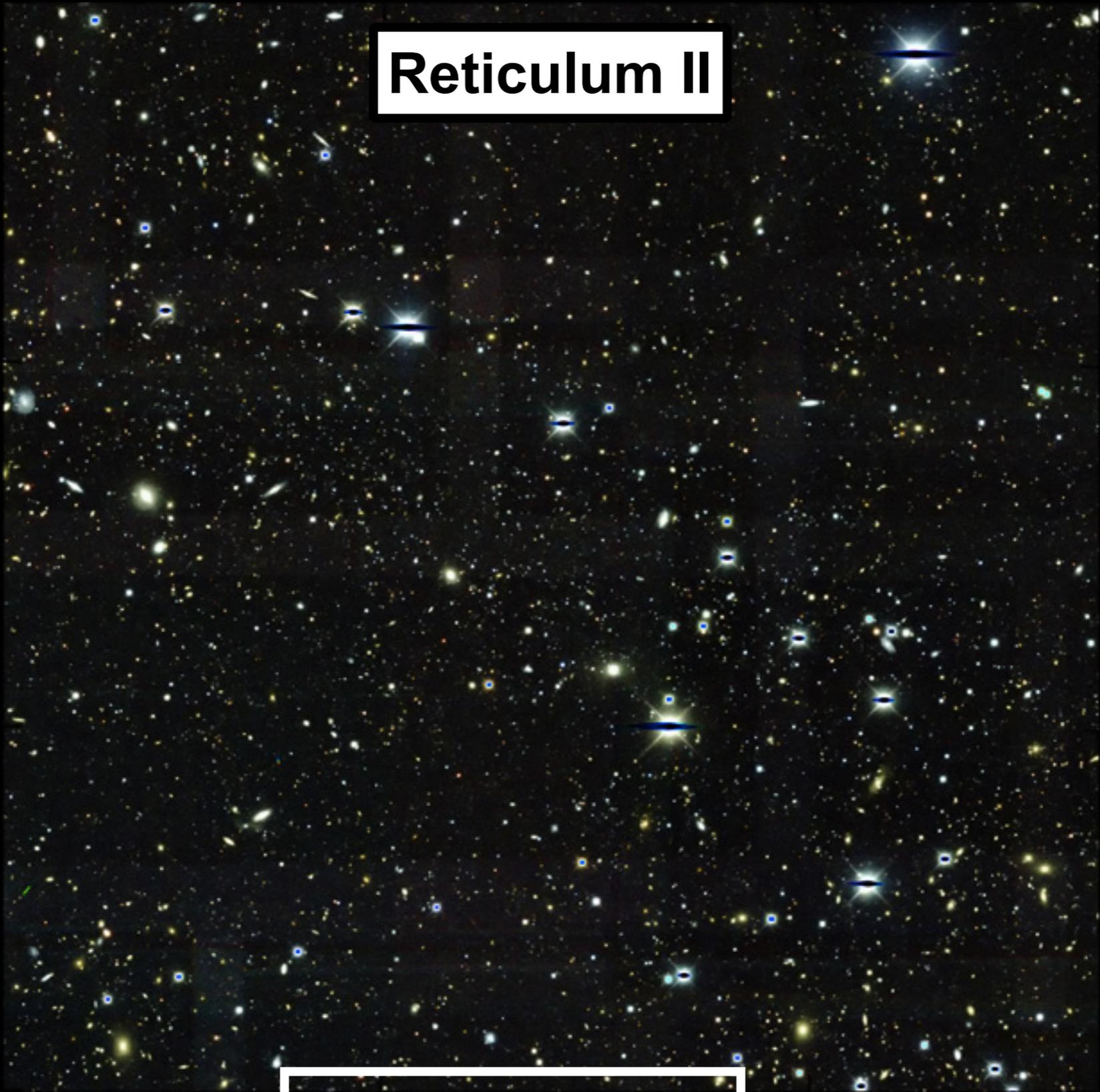
$$\lambda = \frac{1}{f} \sum_{i \in \text{Stars}} p_i$$

$$\log L = - \sum_{i \in \text{Stars}} \log(1 - p_i) - f\lambda$$

**A likelihood analysis to simultaneously combine spatial and spectral information**

- $u_i$  = signal probability
- $b_i$  = background probability
- $\lambda$  = number of stars in the dwarf
- $f$  = observable fraction of stars

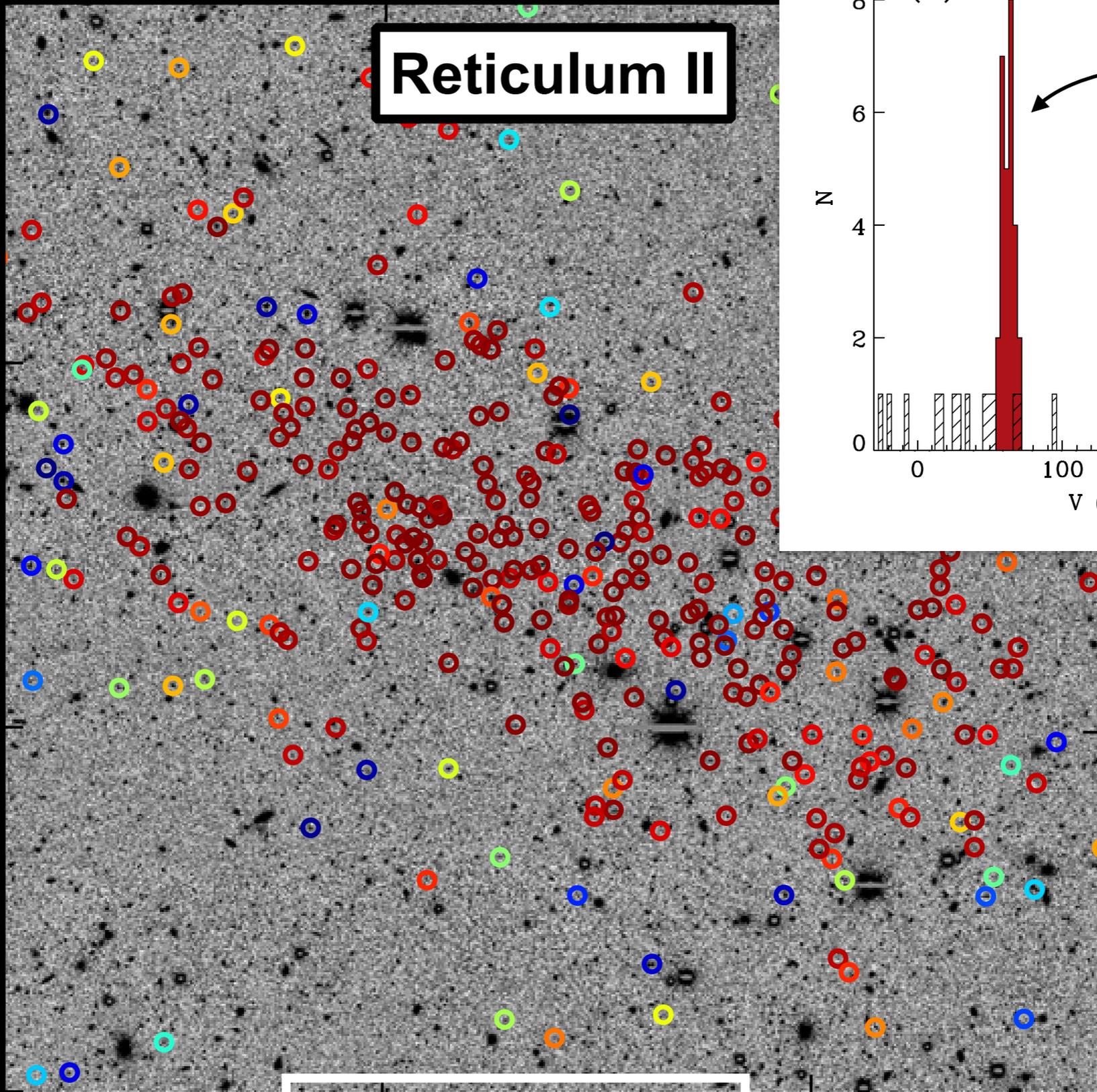
**This technique naturally yields a membership probability for each star; important for spectroscopic targeting**



**Reticulum II**

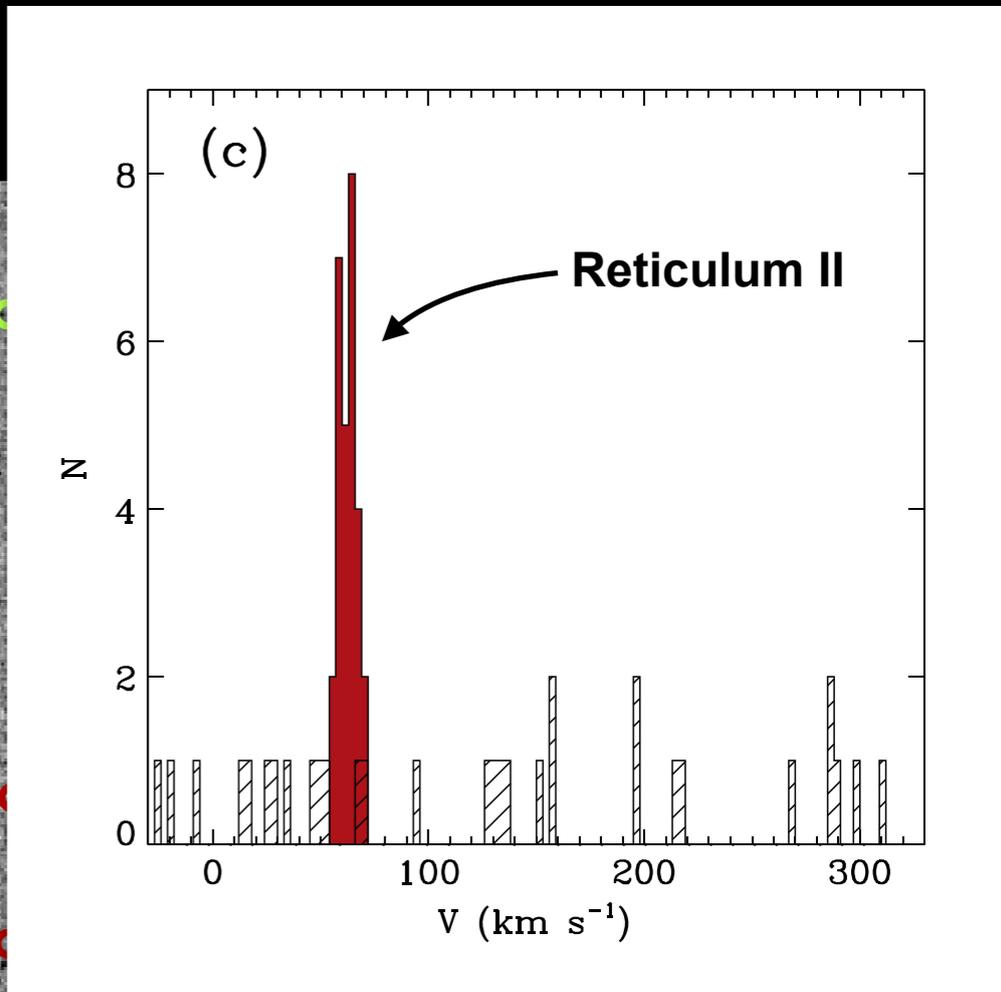
**4m Telescope  
DECam CCD Camera**

**Bechtol, ADW et al. (2015)**



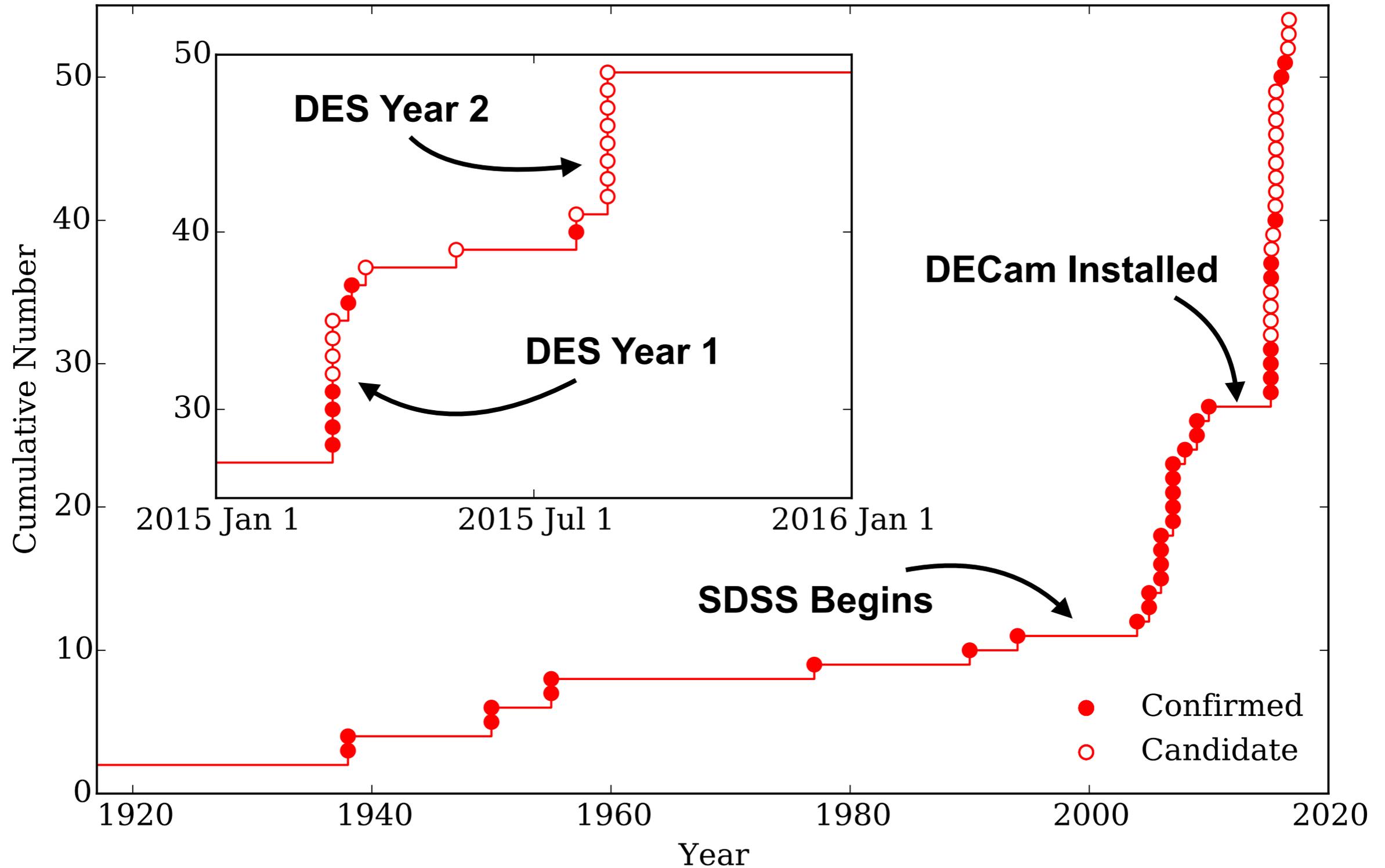
Colors correspond to the membership probability assigned to each star by the likelihood analysis

4m Telescope  
DECam CCD Camera

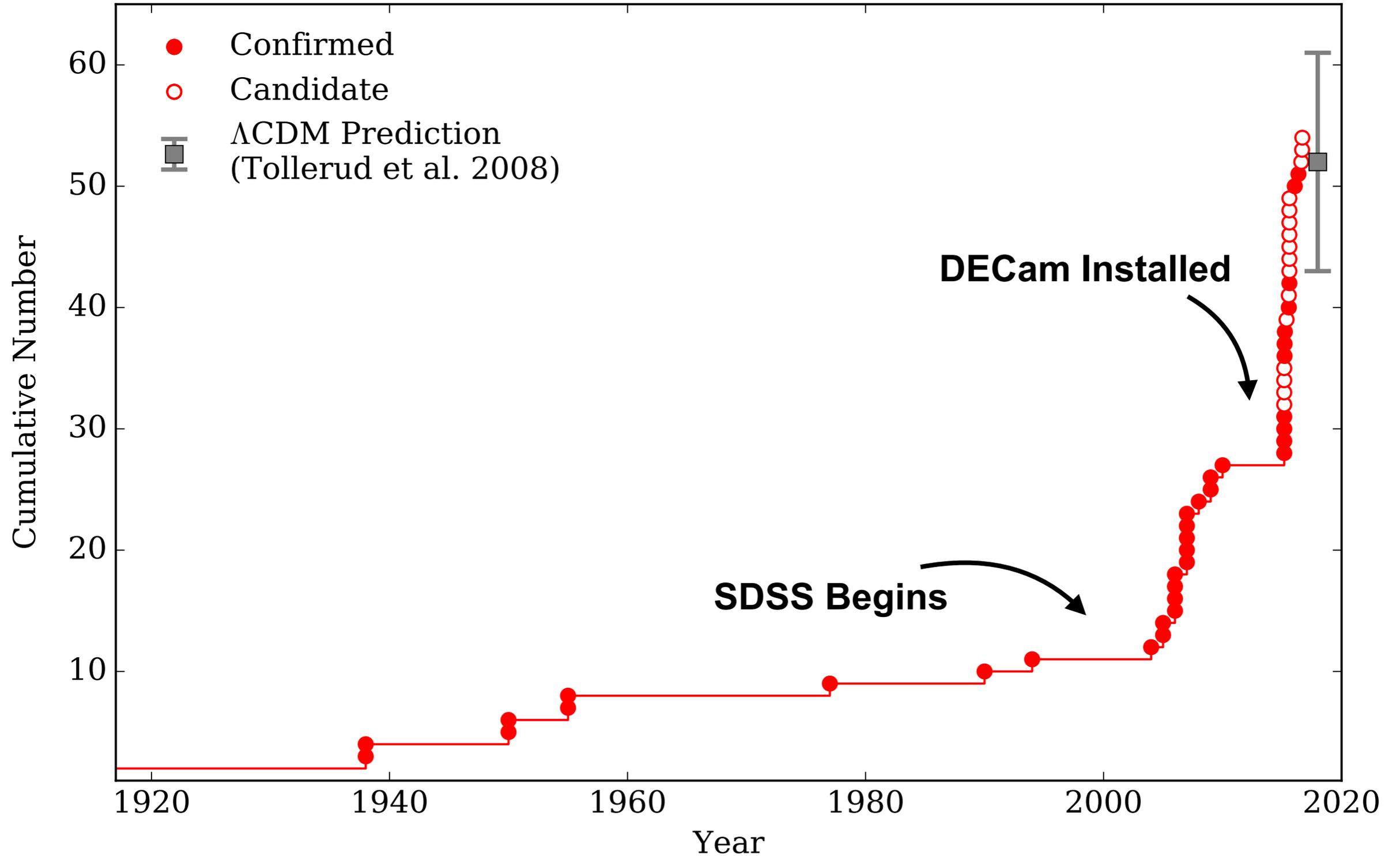


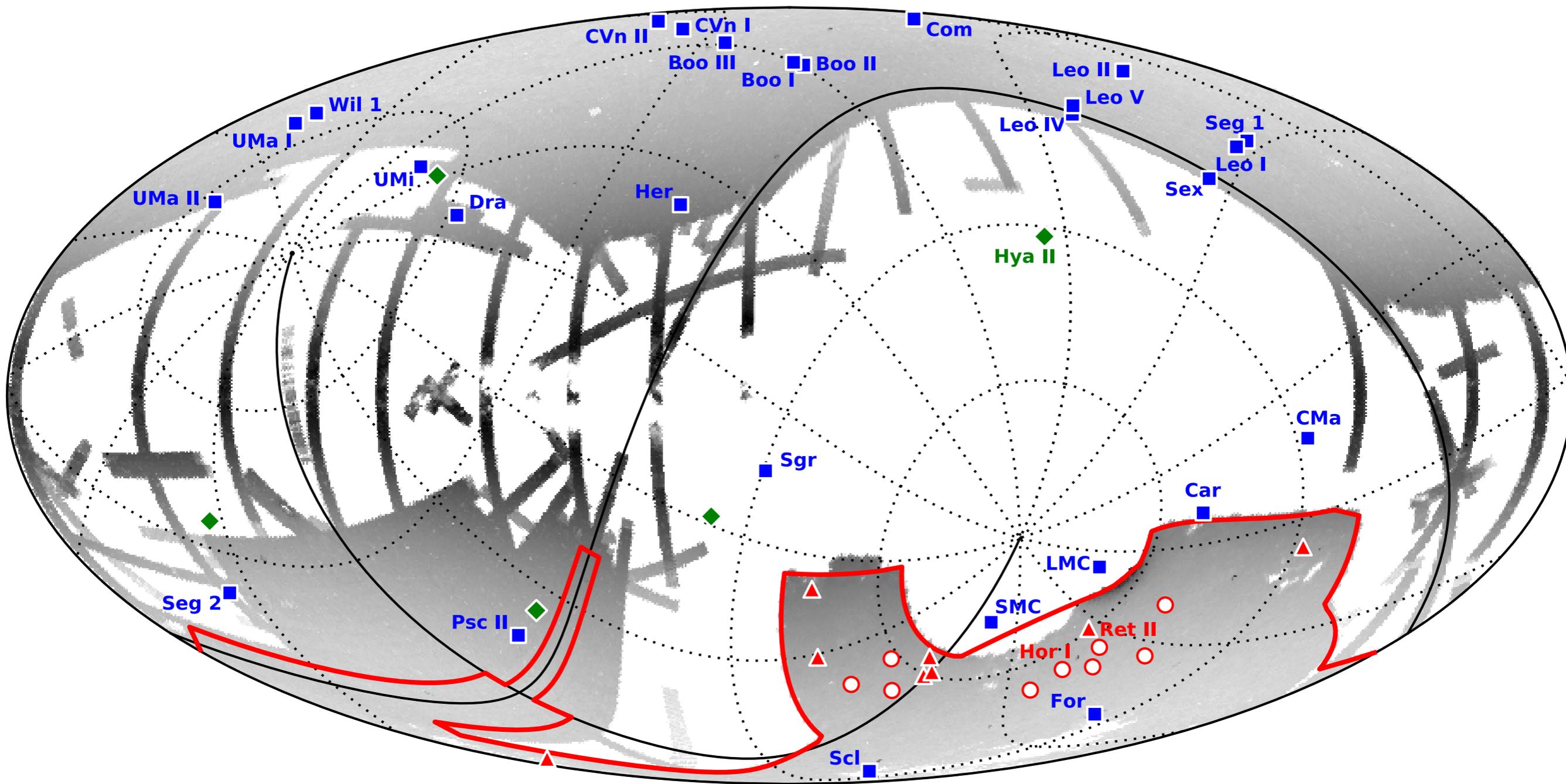
Bechtol, ADW et al. (2015)

# Dwarf Galaxy Discovery Timeline



# Dwarf Galaxy Discovery Timeline





**Blue** - Previously discovered satellites

**Green** - Discovered in 2015 with PanSTARRS, SDSS, etc.

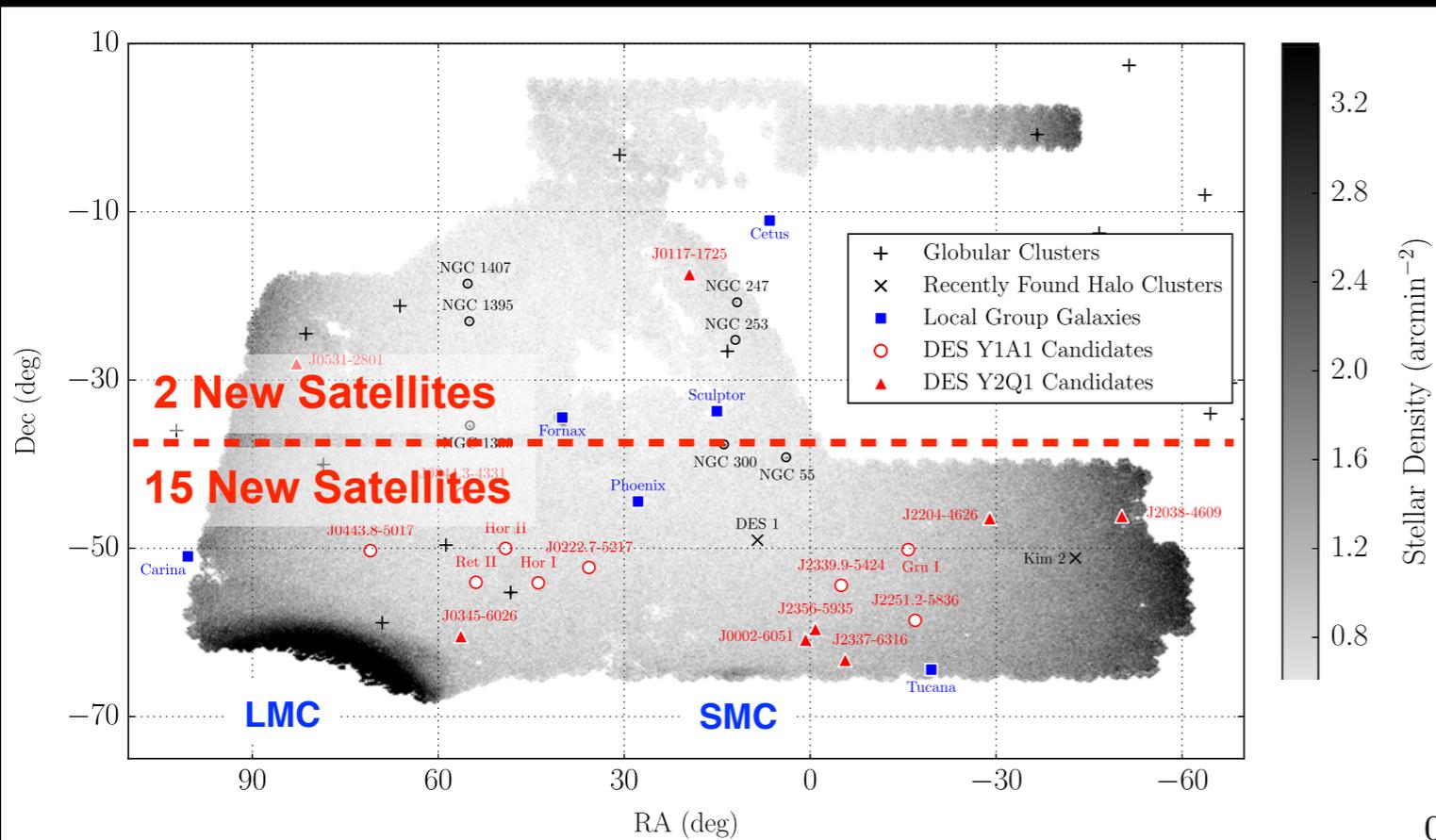
**Red outline** - DES footprint

**Red circles** - DES Y1 satellites

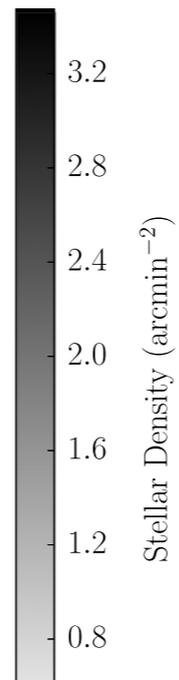
**Red triangles** - DES Y2 satellites



# Satellites of the Magellanic Clouds?



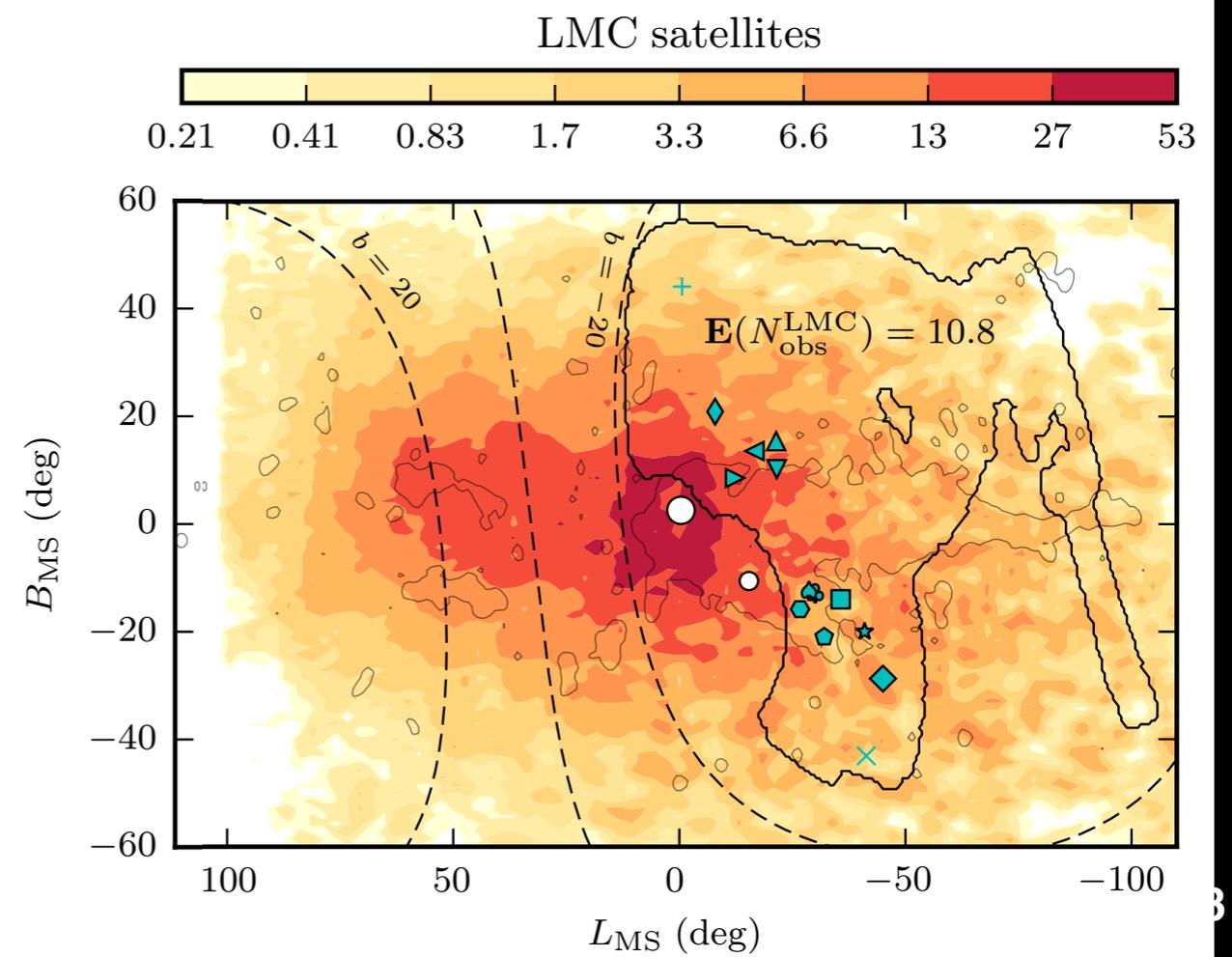
ADW et al. ApJ 813, 109 (2015)



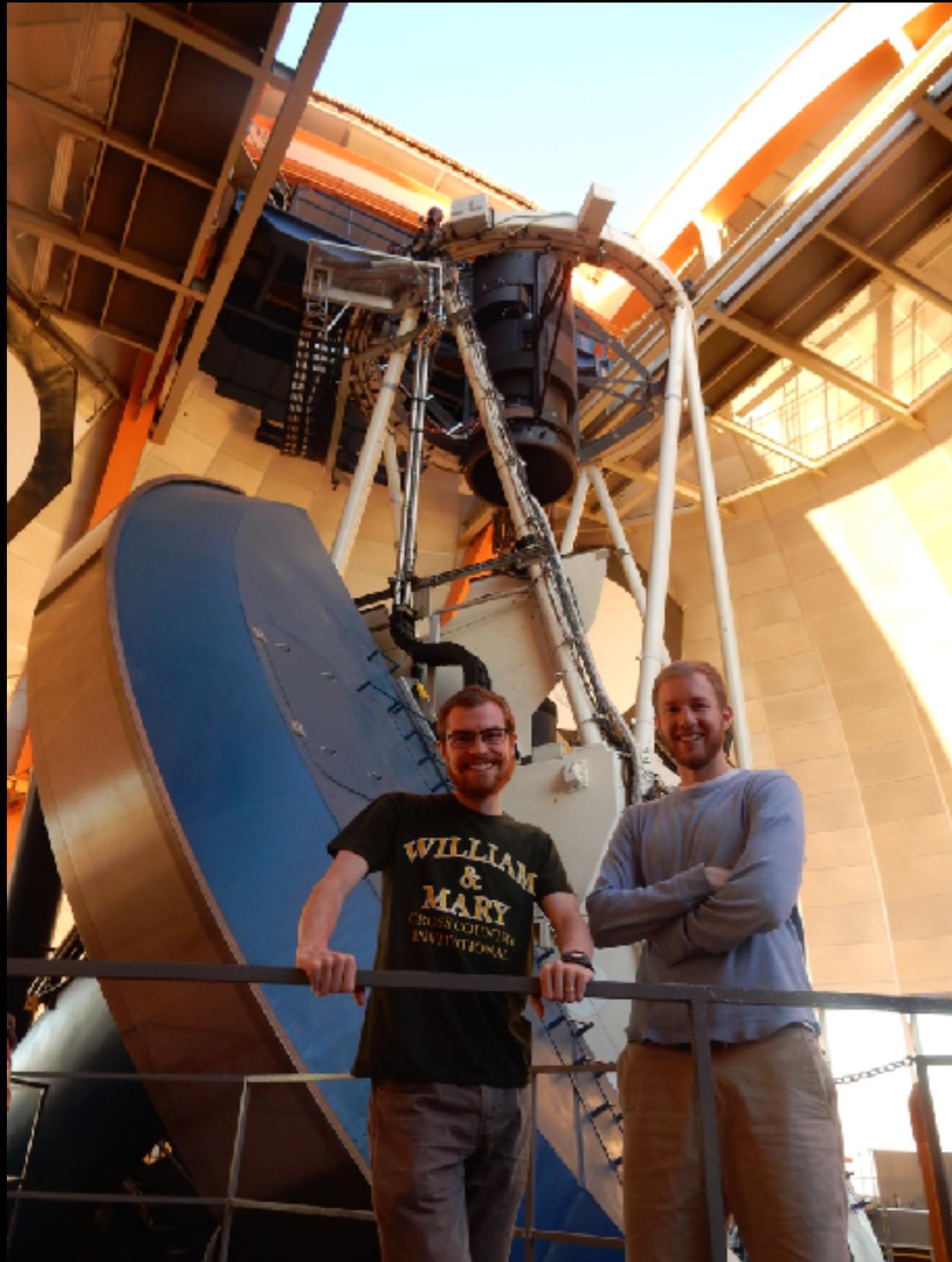
There is  $\sim 3\sigma$  evidence that DES satellites are not isotropically distributed.

Jethwa et al. MNRAS 461, 2 (2016)

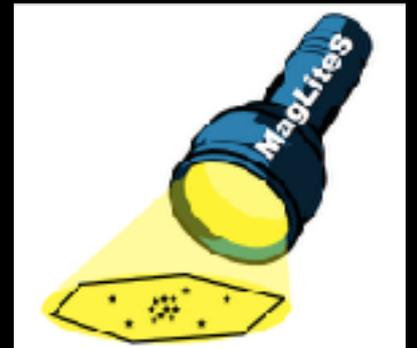
This anisotropy could be explained by an association with the Magellanic Clouds



# Magellanic Satellites Survey



**(MagLiteS)**



**DECam Program for 12 nights  
in 2016-2017**

**PI: Keith Bechtol**

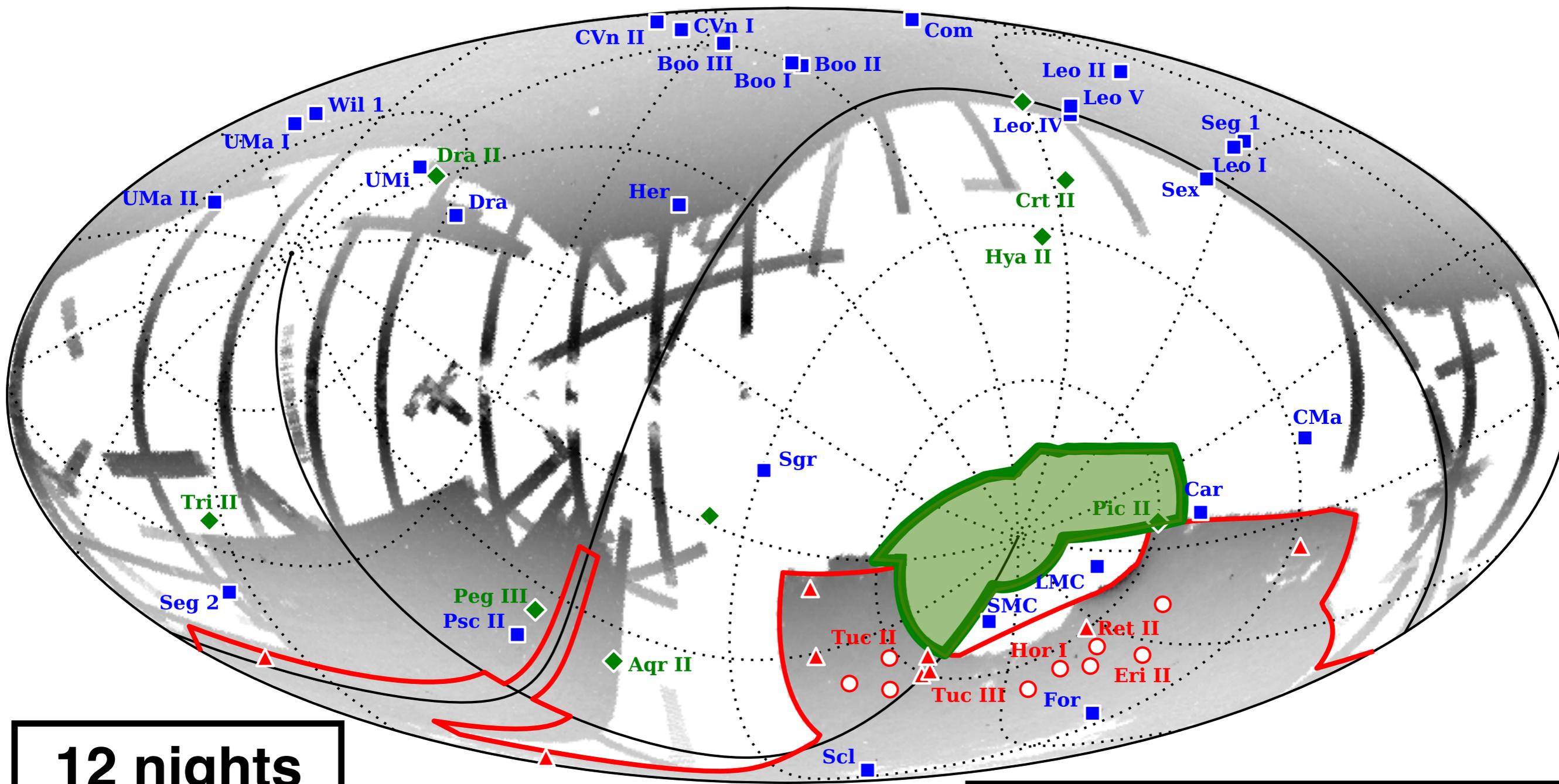
**Deputy PI: ADW**

**Funding through the NASA  
Guest Investigator Program**

**PI: ADW**

**Collaboration of ~45 members  
across ~20 institutions**

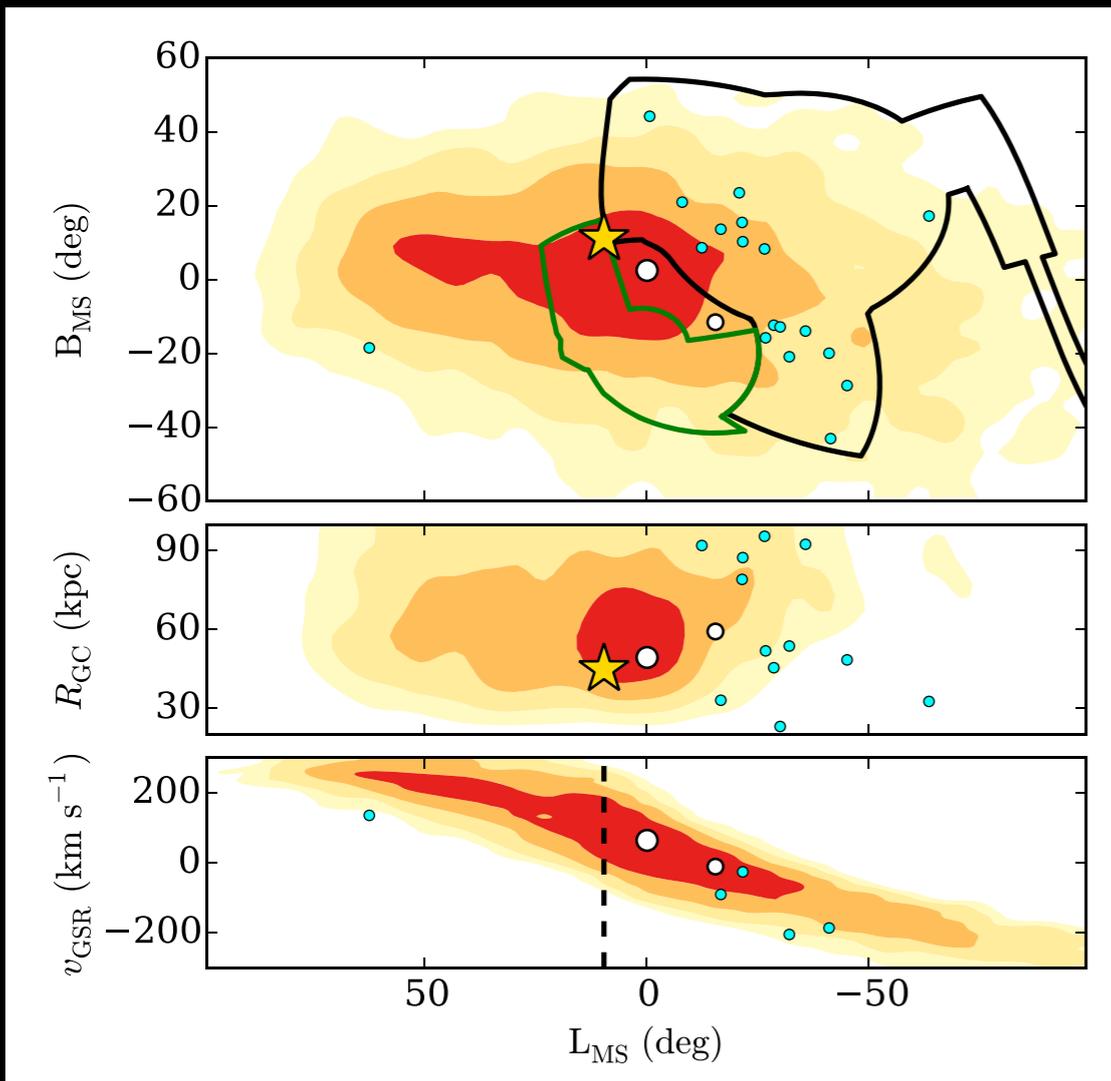
# Magellanic Satellites Survey (MagLiteS)



**12 nights**  
**~1300 deg<sup>2</sup>**  
**3 tilings**

**Roughly comparable**  
**in depth to DES Y2**  
**in g and r-bands**

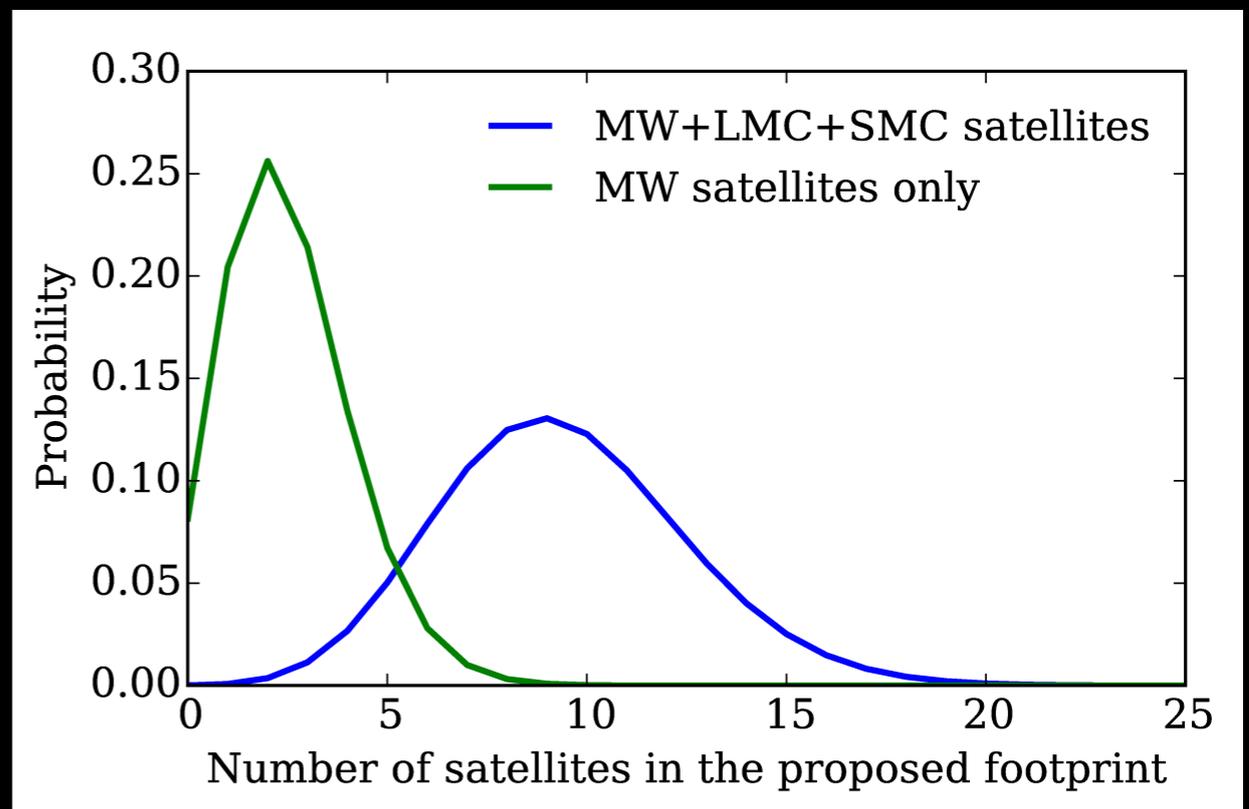
# Satellites of the Magellanic Clouds?



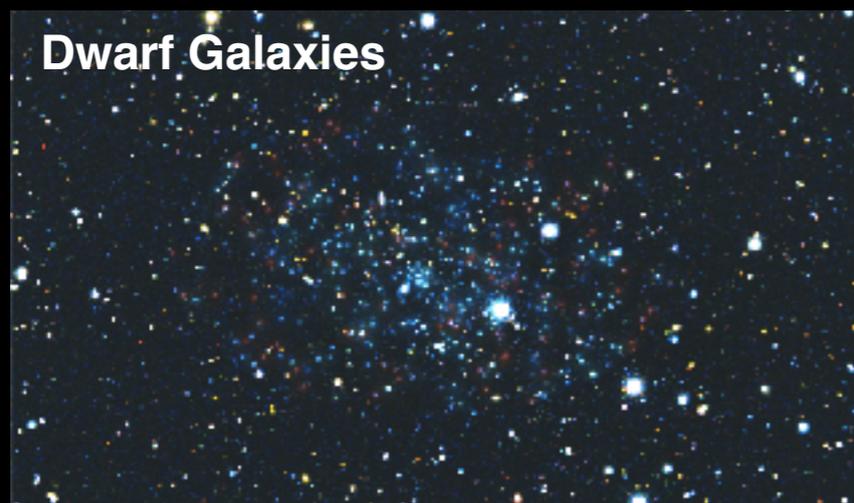
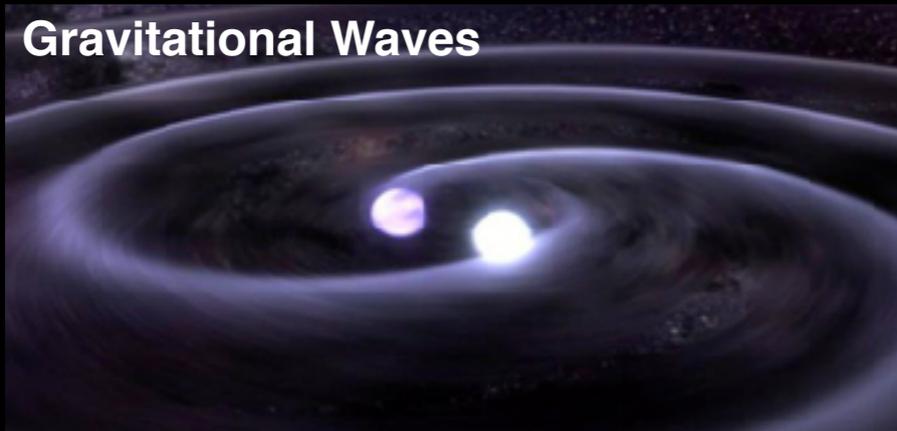
ADW et al. ApJL 833, 5, 2016

**First satellite found in a 1/4<sup>th</sup> of the MagLiteS data; other candidates being investigated.**

**Simulations predict ~3 dwarf galaxies for an isotropic distribution and ~10 galaxies for a Magellanic Cloud association.**



# Blanco Imaging of the Southern Sky



**NOAO DECam Program for 12 nights in 2017A**

**Co-PIs: Soares-Santos & ADW**

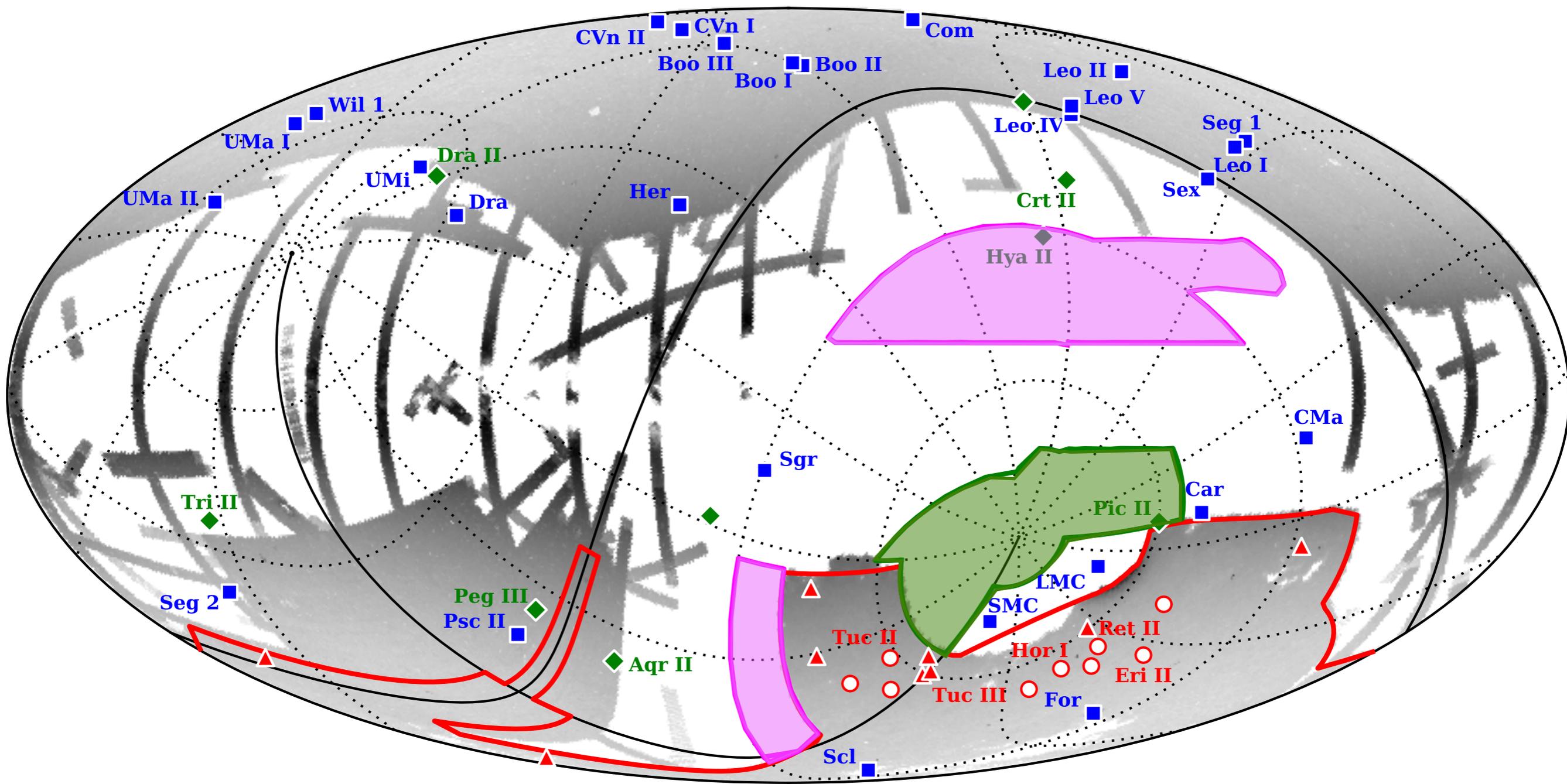
**3 Science Drivers:**

- **Dwarf Galaxy Searches**
- **Gravitational Wave Follow-up**
- **Search for Planet 9**

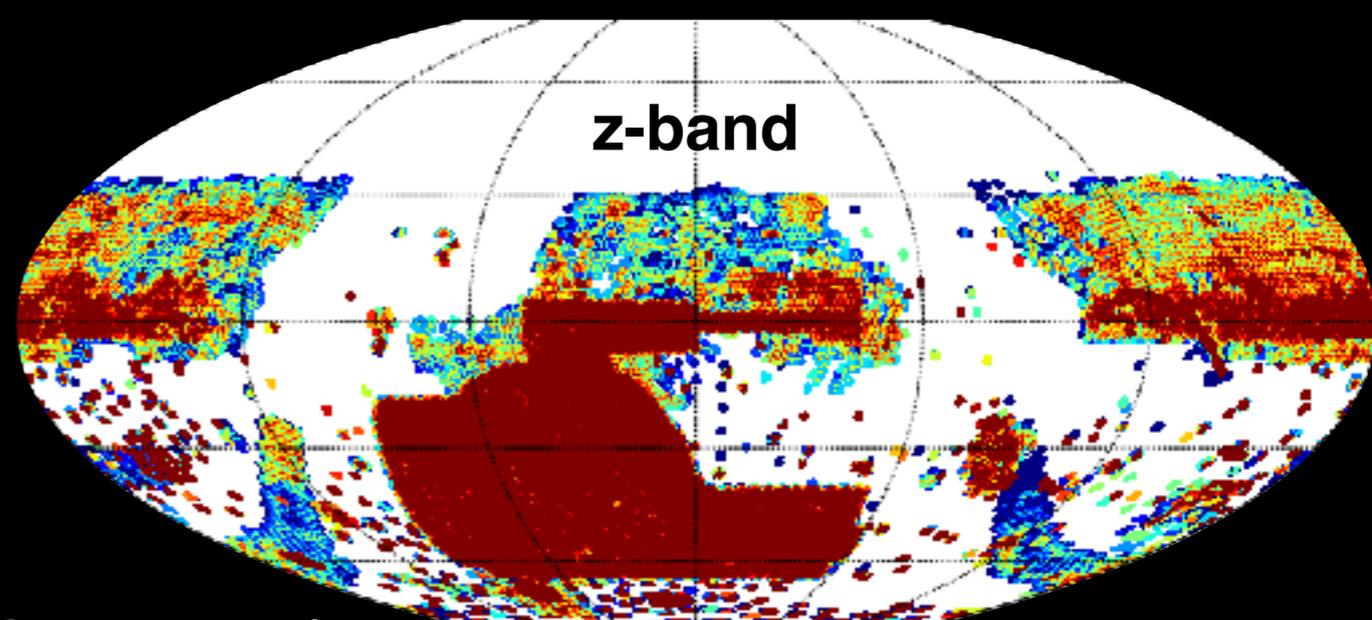
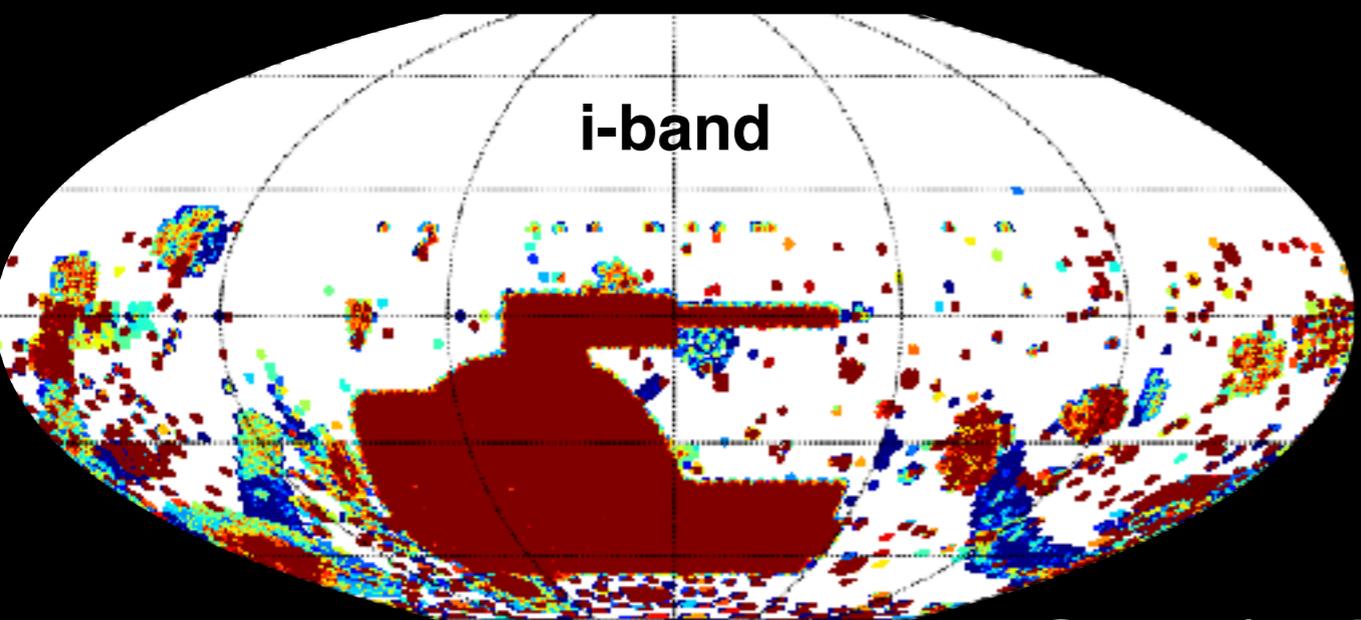
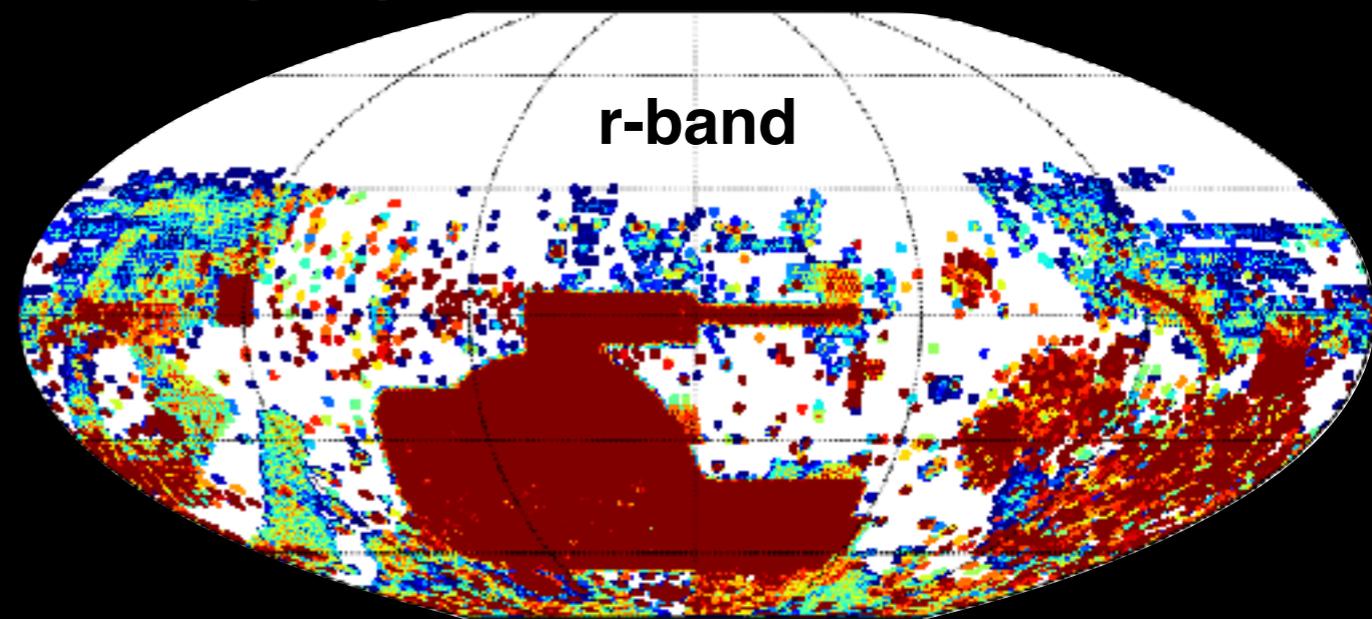
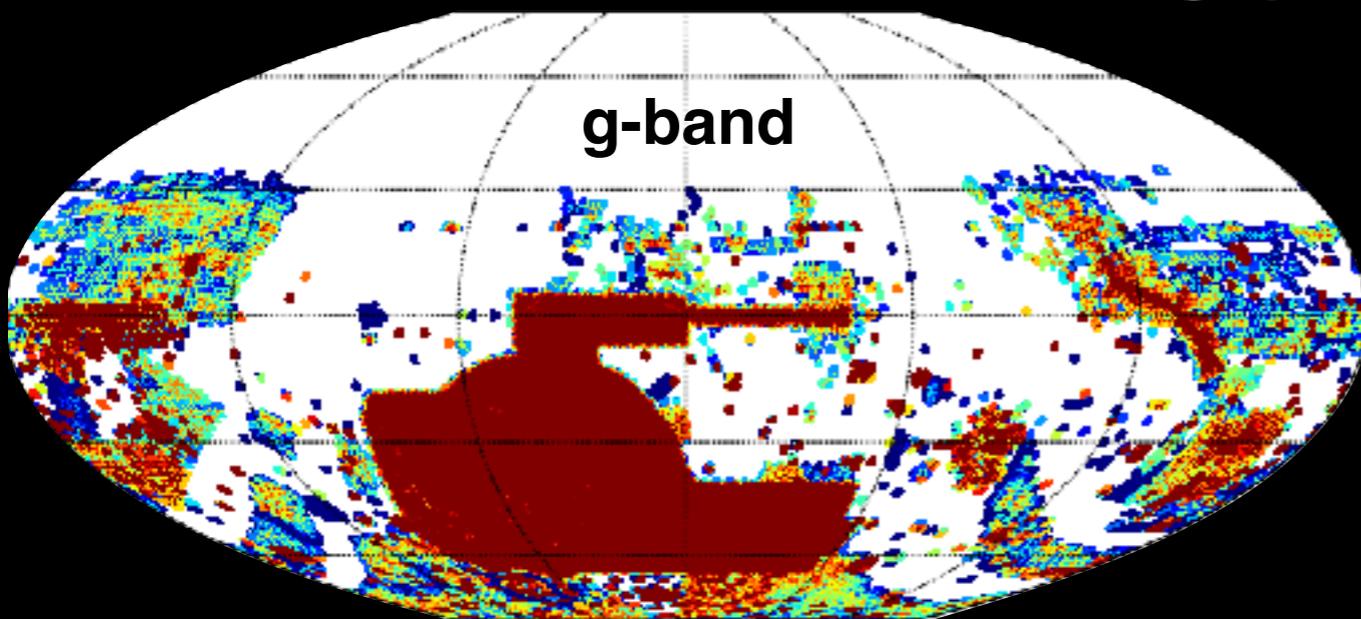
**Cover  $\sim 2000 \text{ deg}^2$  in 2017;  
eventually cover the entire  
sky in g,r,i,z bands**

**Collaboration of  $\sim 35$  members  
across  $\sim 10$  institutions**

# Blanco Imaging of the Southern Sky (BLISS)

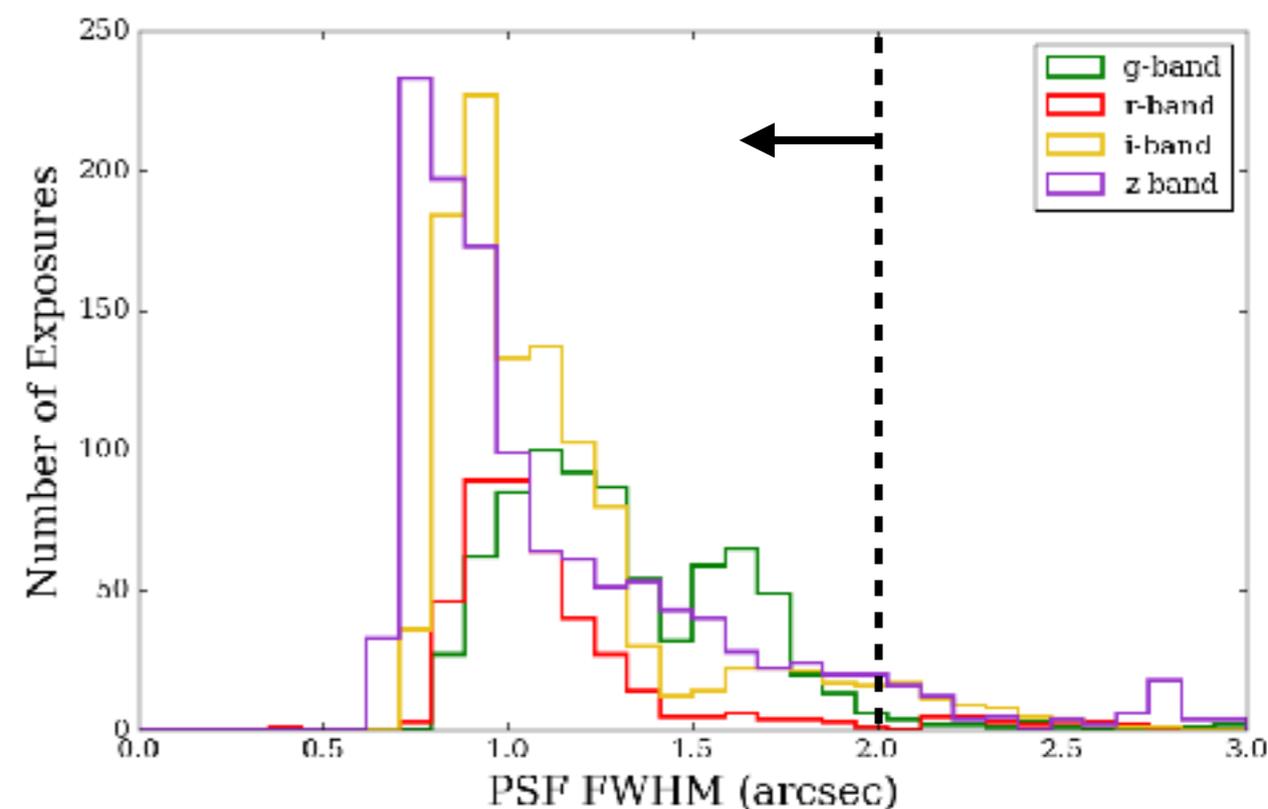
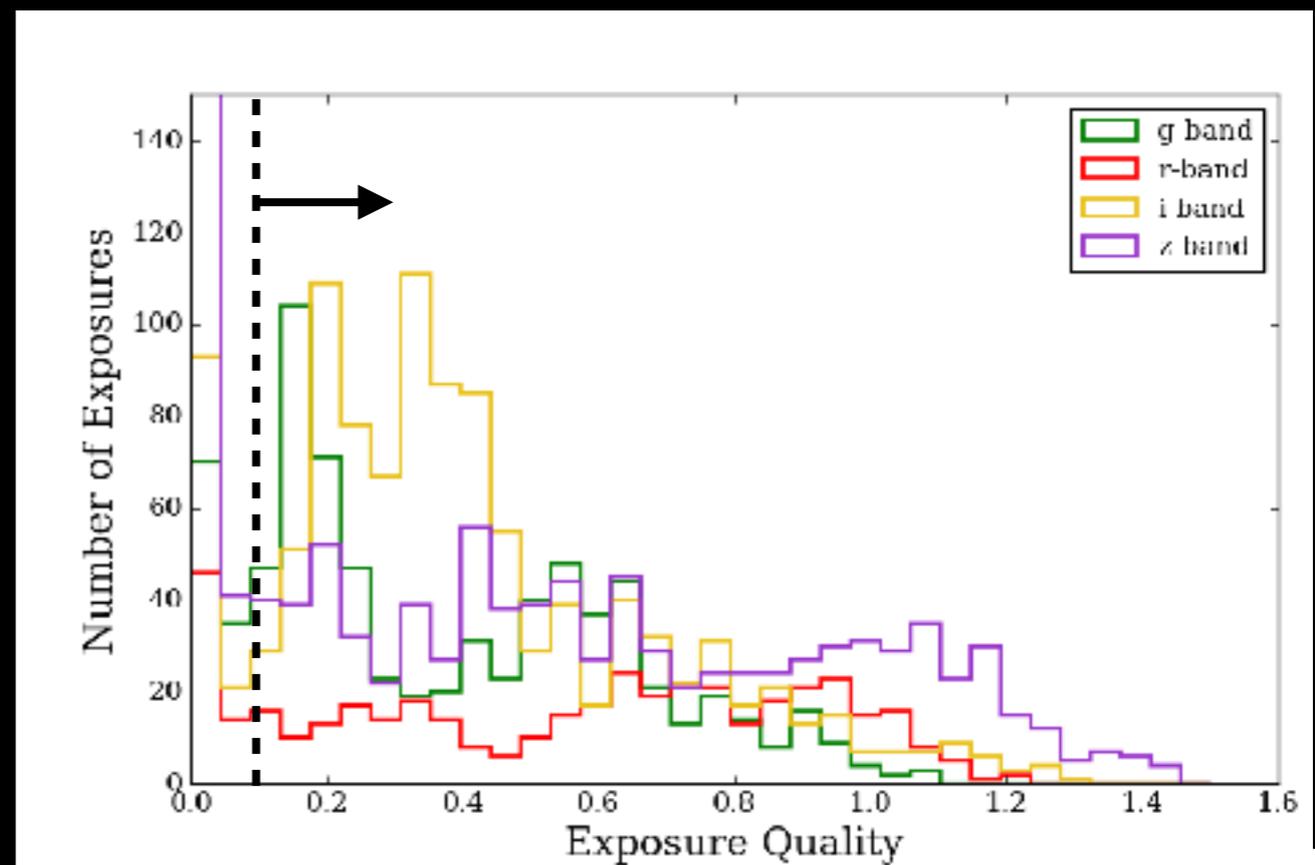
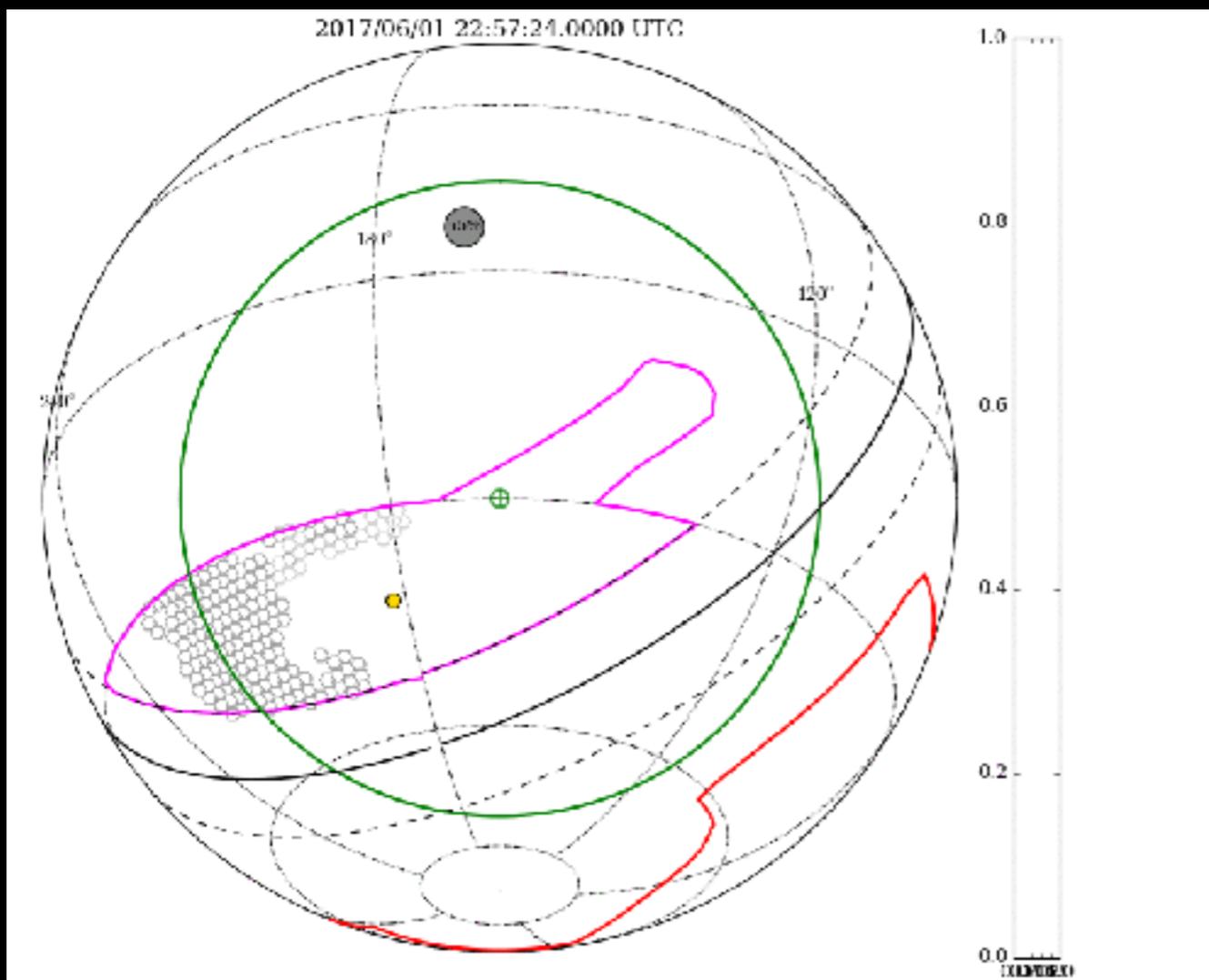


# BLISS also uses all pre-existing DECam Data



Sum( $t_{\text{eff}} \times t_{\text{exp}}$ )  
log-scale from 30s (blue) to 300s (red)

# Blanco Imaging of the Southern Sky



- Most exposures pass cuts on exposure quality
- 2017A data covers  $\sim 2200$  deg<sup>2</sup> in any single band



# CDM Predictions for Future Dwarf Discoveries

