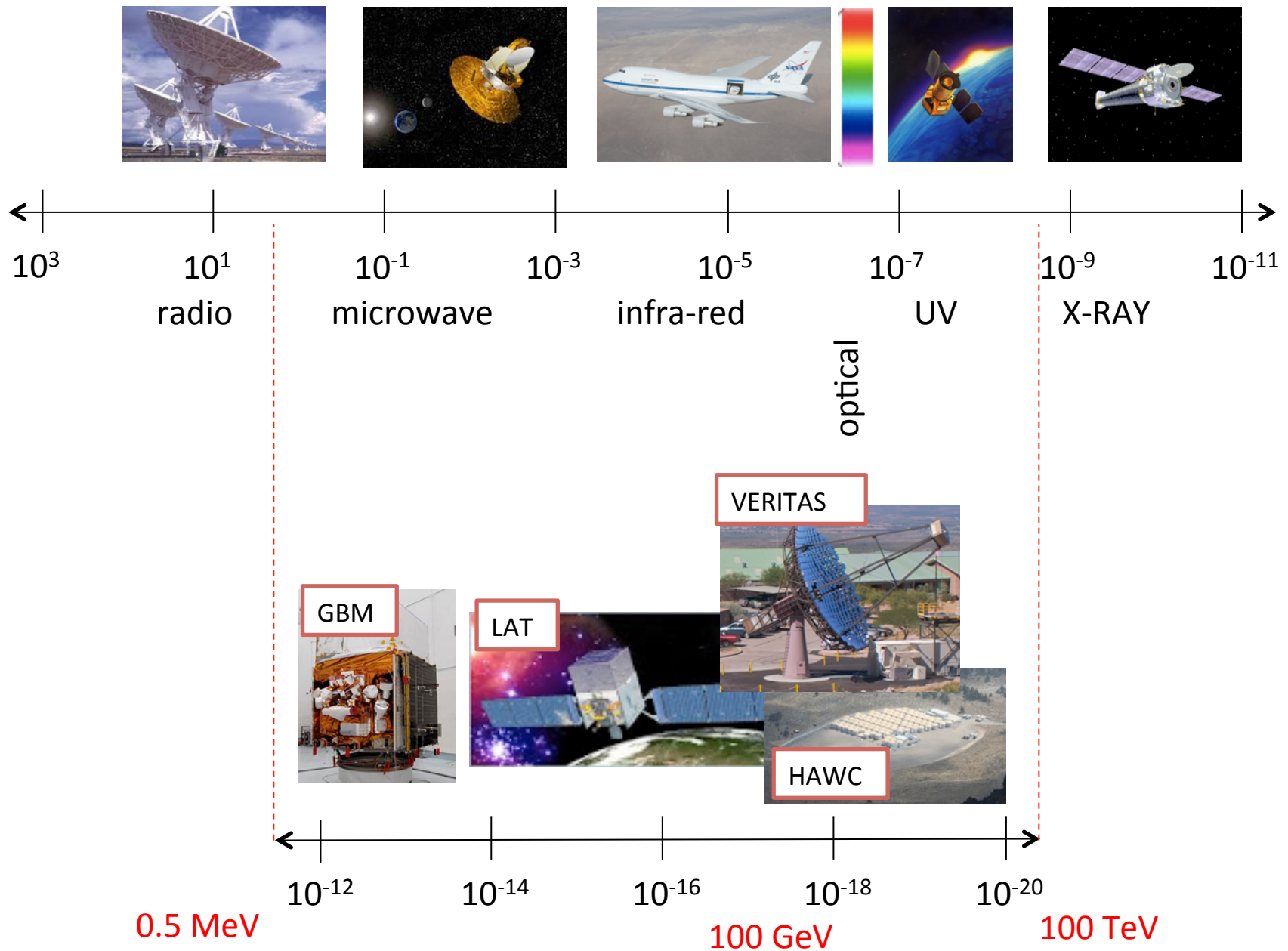


# TeV Particle Astrophysics with the High Altitude Water Cherenkov (HAWC) Gamma-Ray Observatory

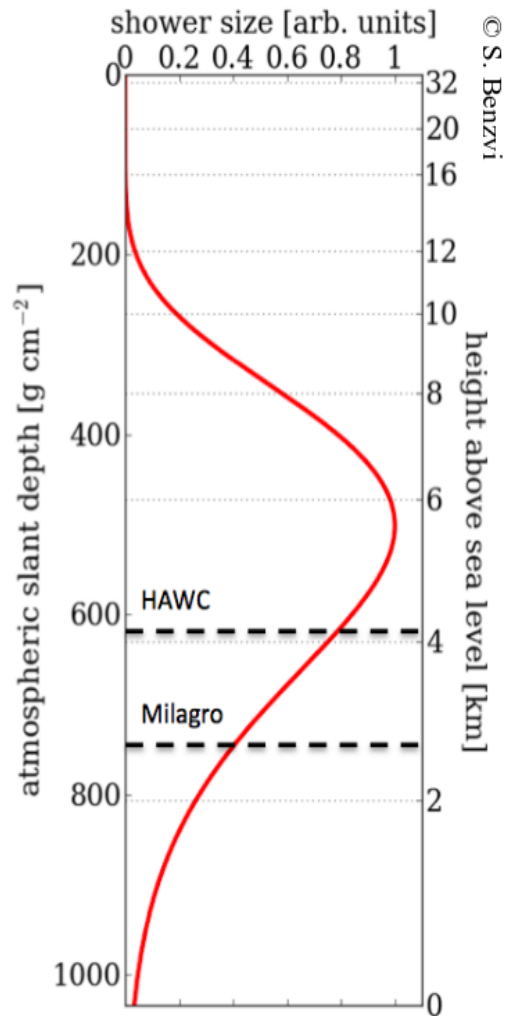
Detecting the Highest Energy Light  
with a Telescope Made of Water

Kirsten Tollefson  
for the HAWC Collaboration  
Michigan State University

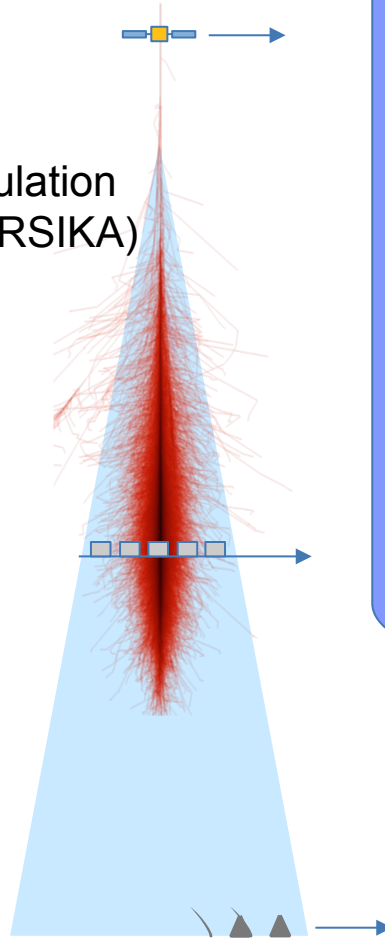




# Gamma-Ray Detectors



Simulation  
(CORSIKA)



**Satellite Detector**  
30 MeV – 300 GeV

**Wide Field of View,  
Continuous  
Operations**



**Extensive Air Shower (EAS) Detector**  
100 GeV – 100 TeV



**Imaging Atmospheric Cherenkov  
Telescope (IACT)**  
50 GeV – 20 TeV

**TeV  
Sensitivity**

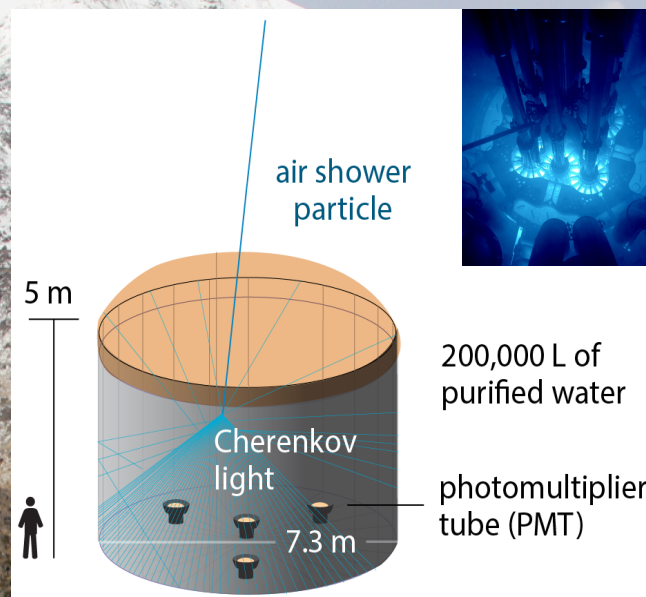


Pico de Orizaba  
5610 meters

Latitude 19°N, Longitude = 97°W.  
In the Mexican state of Puebla,  
4hr drive East of Mexico City.

HAWC  
4100 meters

300 water tanks covering 22,500 m<sup>2</sup>

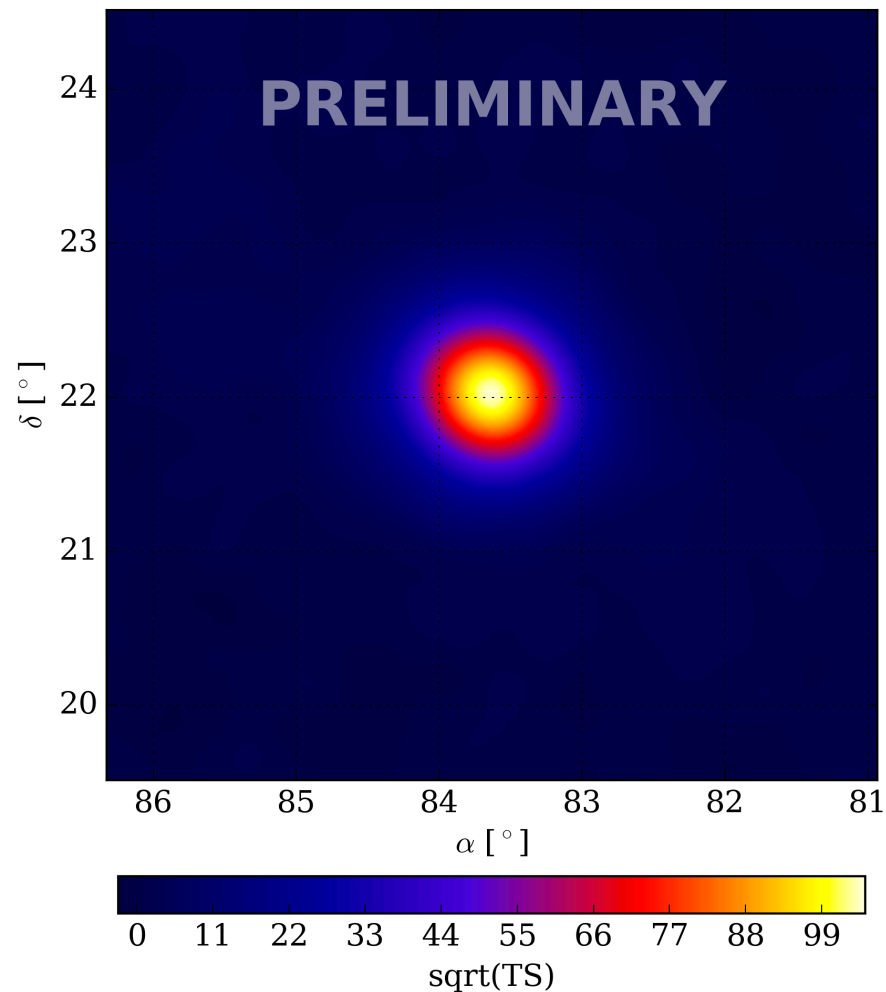


Taken March 18, 2015





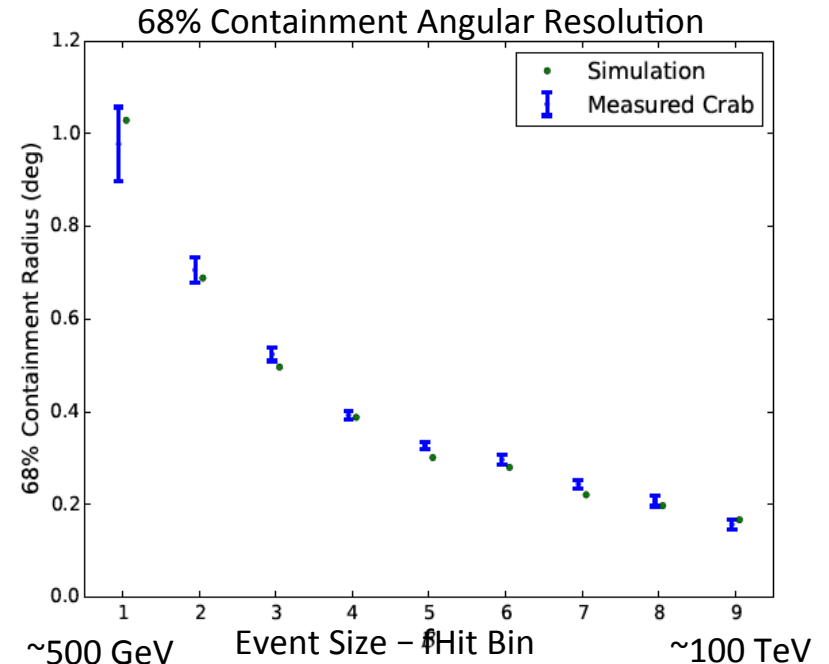
# Crab Nebula as a Calibration Source



K. Tollefson

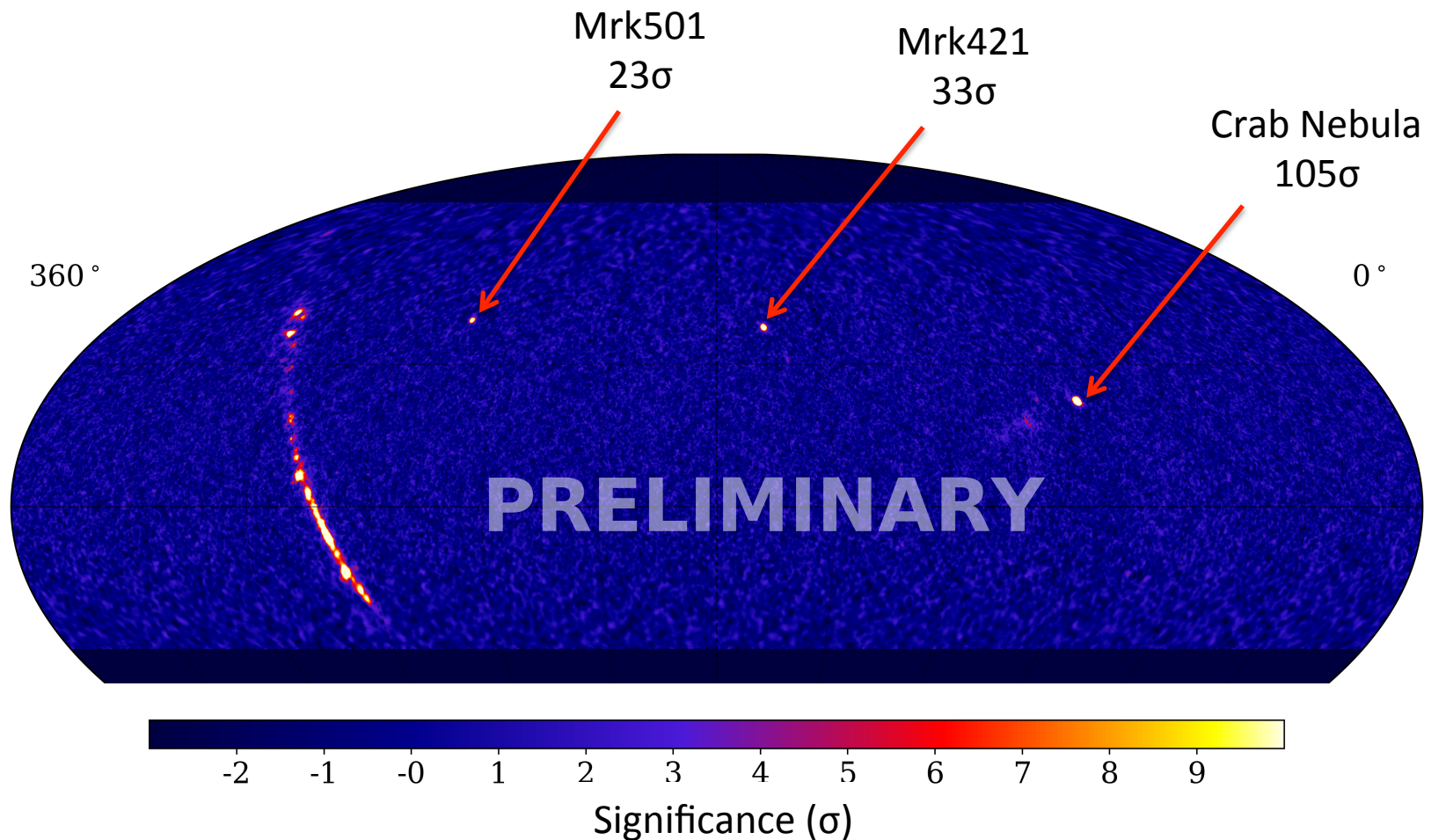
[ApJ, 843:39](#) (July 1, 2017)

- $>100\sigma$  in 1 year of data
- Observe at  $>5\sigma$  with each transit
- Angular resolution  $0.2^\circ$  at high energies





# 25 Month Sky Map with 100 GeV to 100 TeV Photons

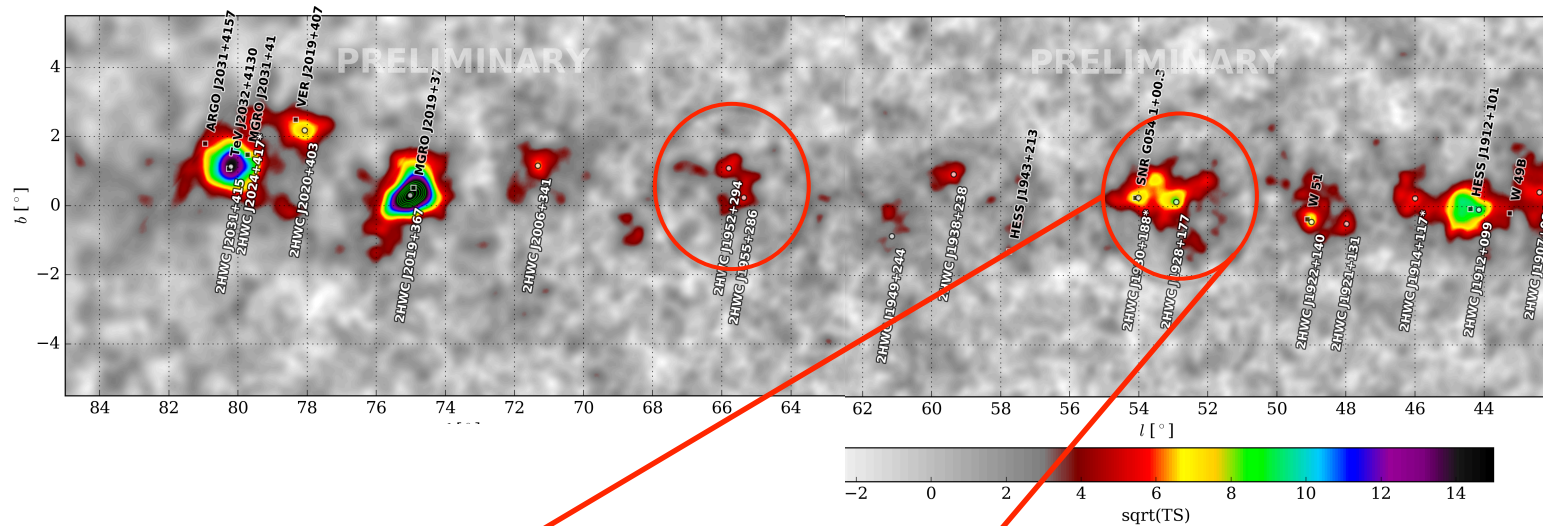


2HWC Catalog (17 months of data) published: [ApJ, 843:88](#) (July 10, 2017)

**39 sources found, 10 were new**

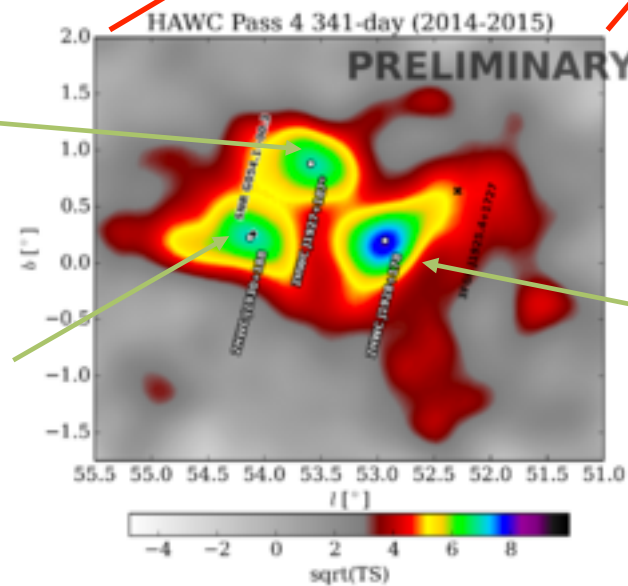


# Known **AND** New TeV Sources



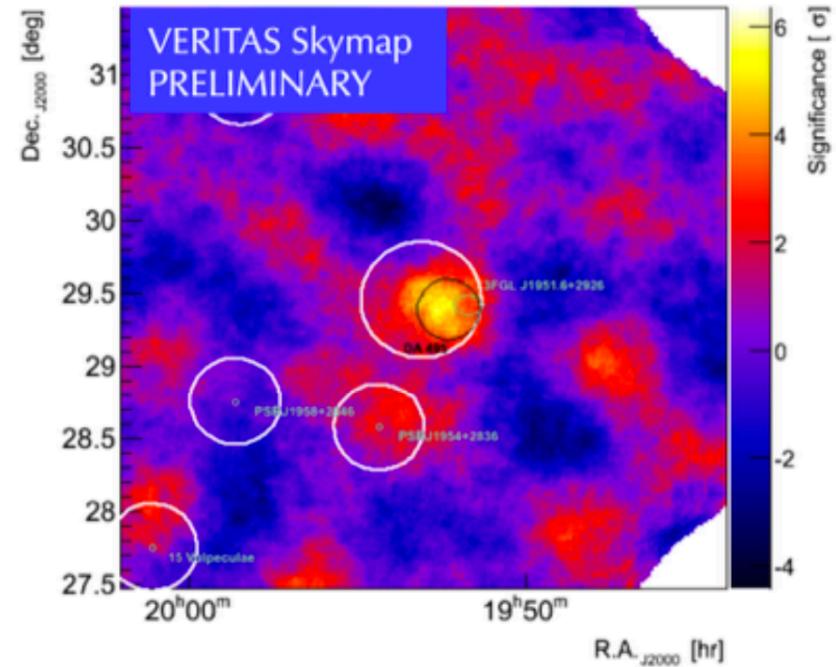
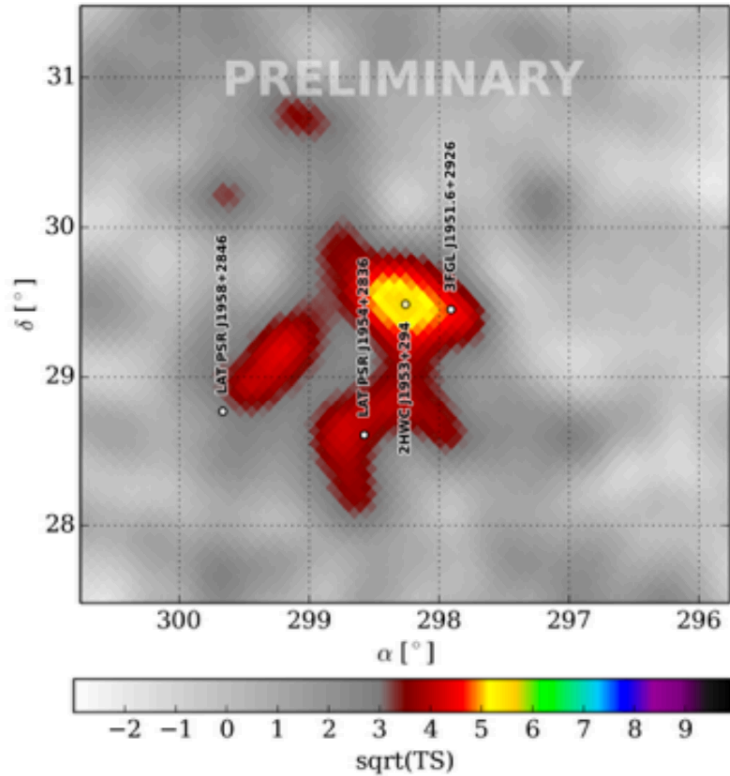
Association unclear

Supernova remnant  
with very energetic  
pulsar



Pulsar ~8kpc (26,000 ly) away

# VERITAS Confirms HAWC detection

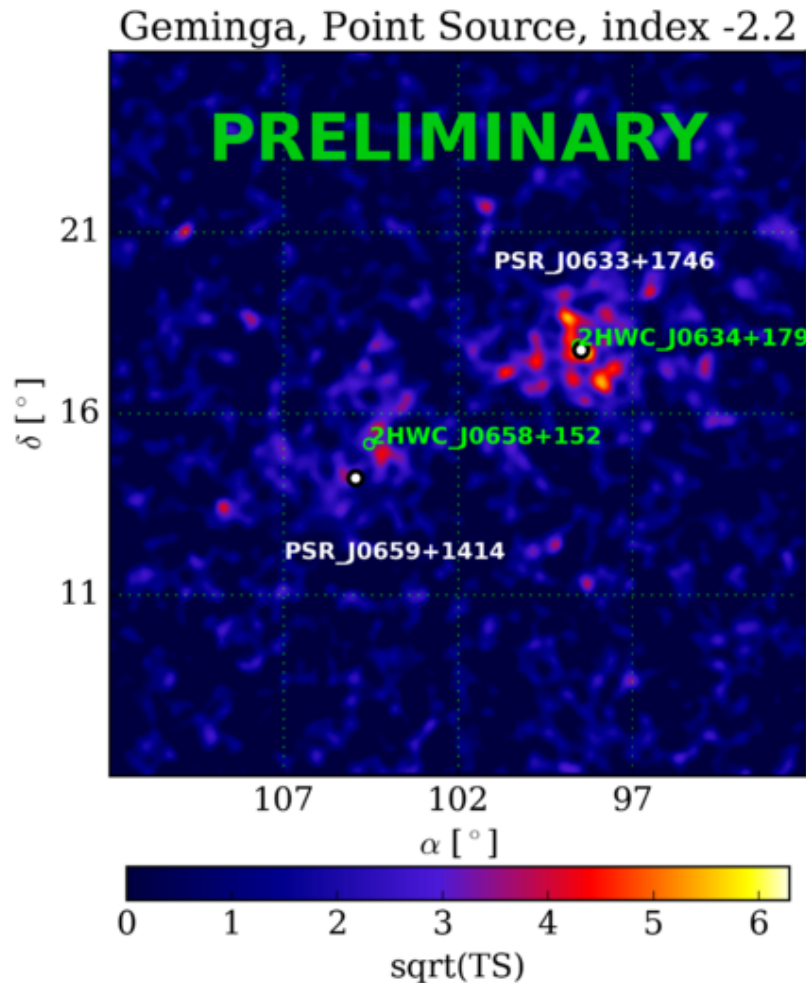


| Name           | $\sqrt{TS}$ | Index            | Flux for index at<br>7 TeV [ $\text{TeV}^{-1}\text{cm}^{-2}\text{s}^{-1}$ ] |
|----------------|-------------|------------------|---|
| 2HWC J1953+294 | 5.58        | $-2.76 \pm 0.15$ | $1.1\text{e-}14 \pm 4.2\text{e-}15$   |

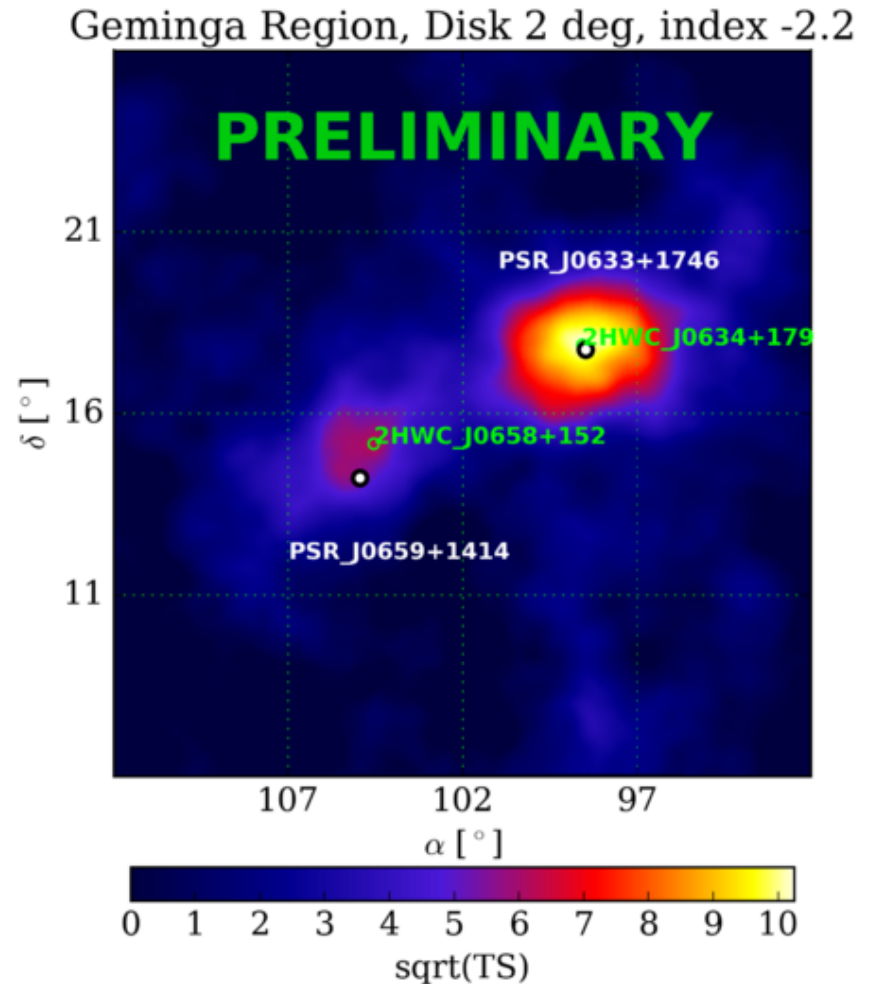


# HAWC Sees Extended Sources

**Geminga:** Closest known (250 parsecs) middle-aged pulsar



Treated as Point Source

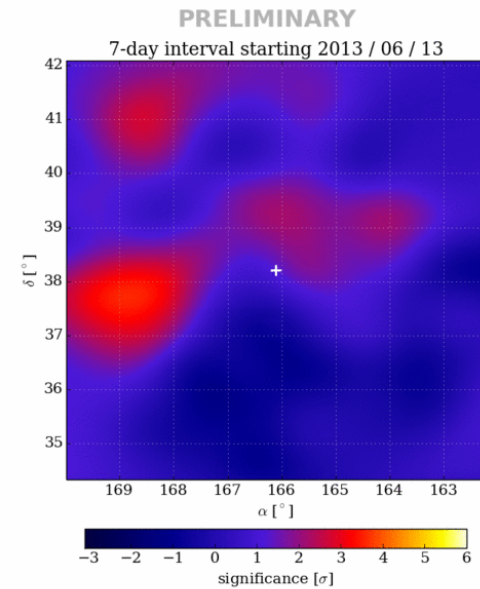
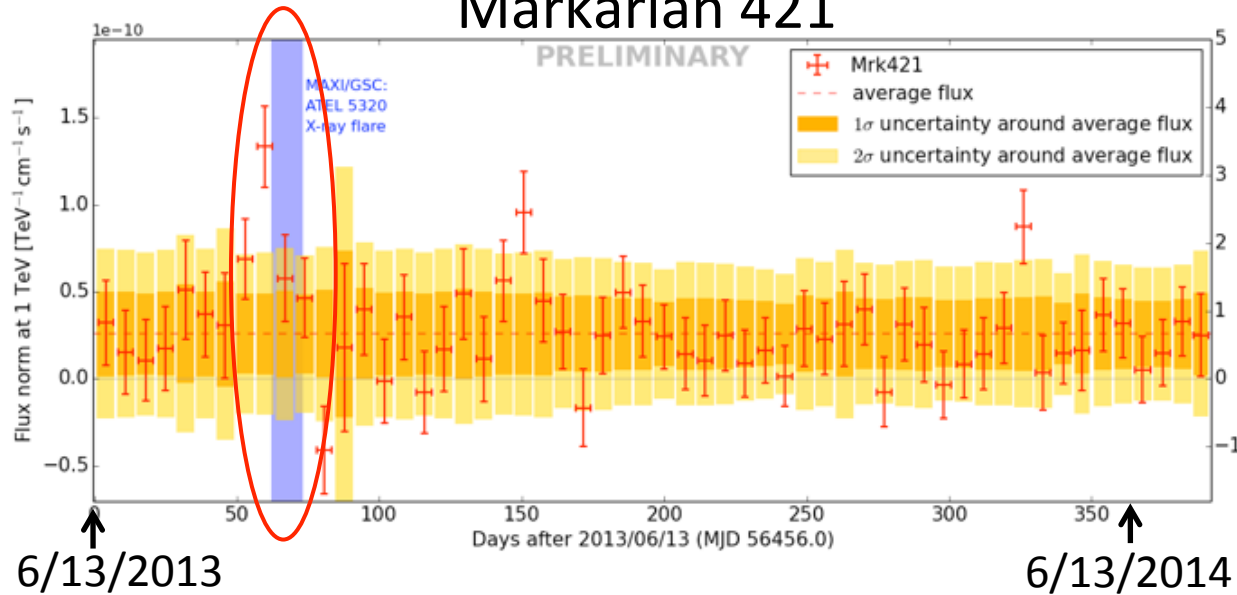


Treated as Extended Source

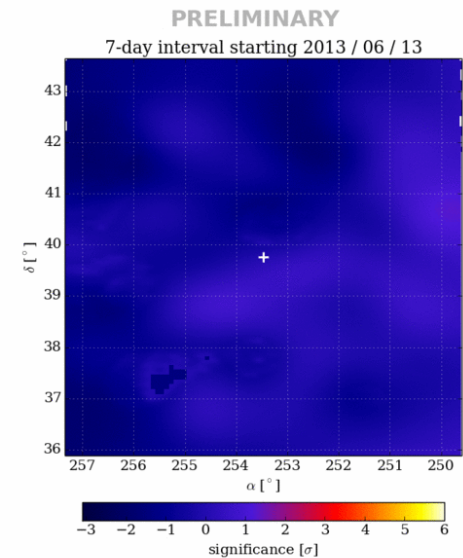
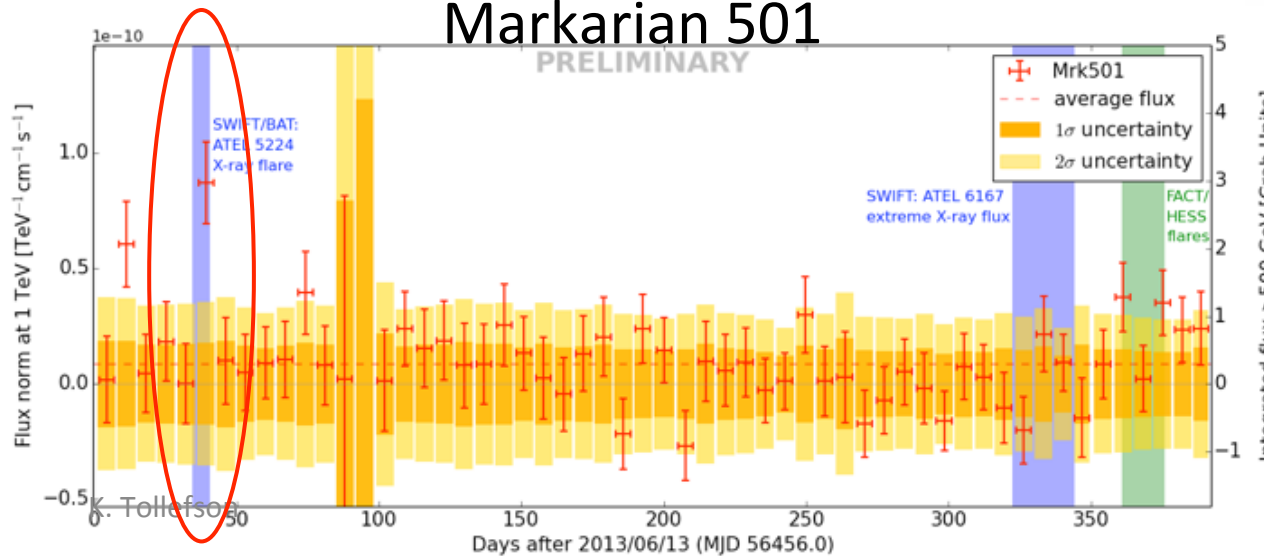
# HAWC Sees Flaring Sources

[ApJ, 841:100 \(June 1, 2017\)](#)

## Markarian 421



## Markarian 501

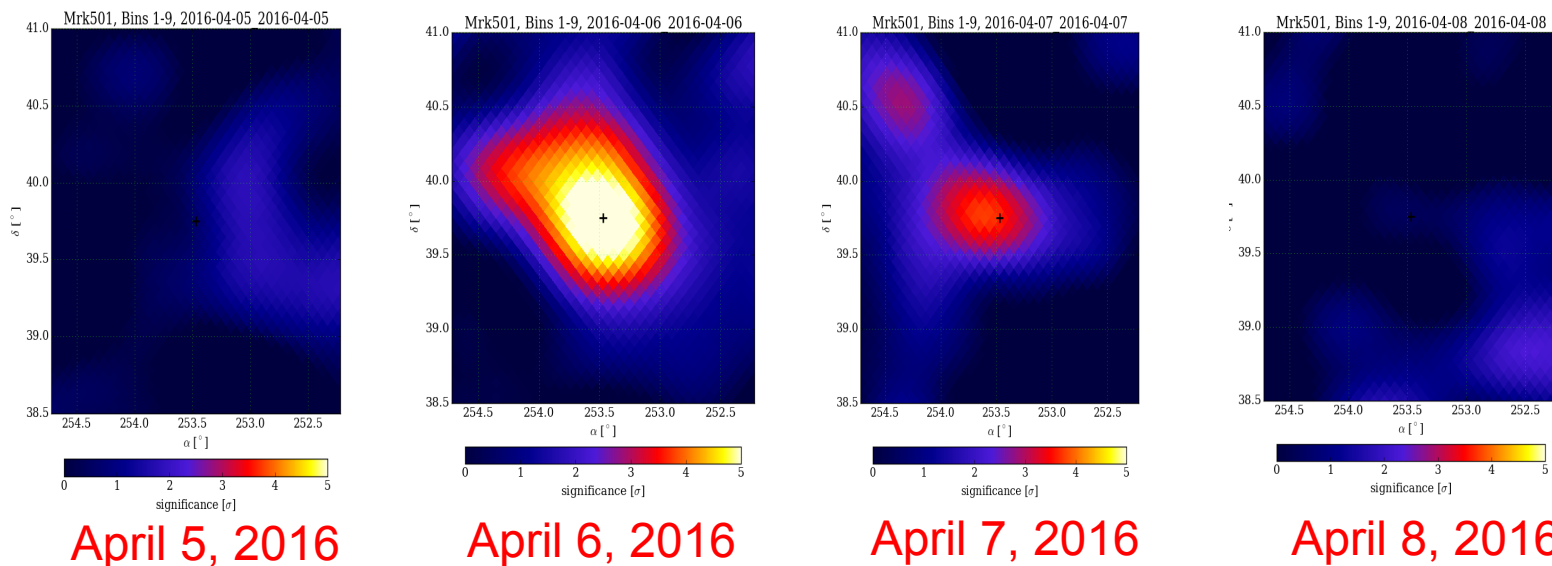




# Real-Time Alerts

HAWC monitors all gamma-ray sources visible to it every day.

Sends alerts such as Astronomer's Telegram (ATel)  
to immediately alert community of activity.



**HAWC detection of increased TeV flux state for  
Markarian 501**

ATel #8922; *Andrés Sandoval (IF-UNAM), Robert Lauer (UNM), Joshua Wood (UMD) on  
behalf of the HAWC collaboration  
on 7 Apr 2016; 23:38 UT*

*Credential Certification: C. Michelle Hui (c.m.hui@nasa.gov)*

Real-time Flare  
Monitoring published:  
[ApJ 843:116](#)  
(July 10, 2017)



# LIGO Gravitational Wave Events

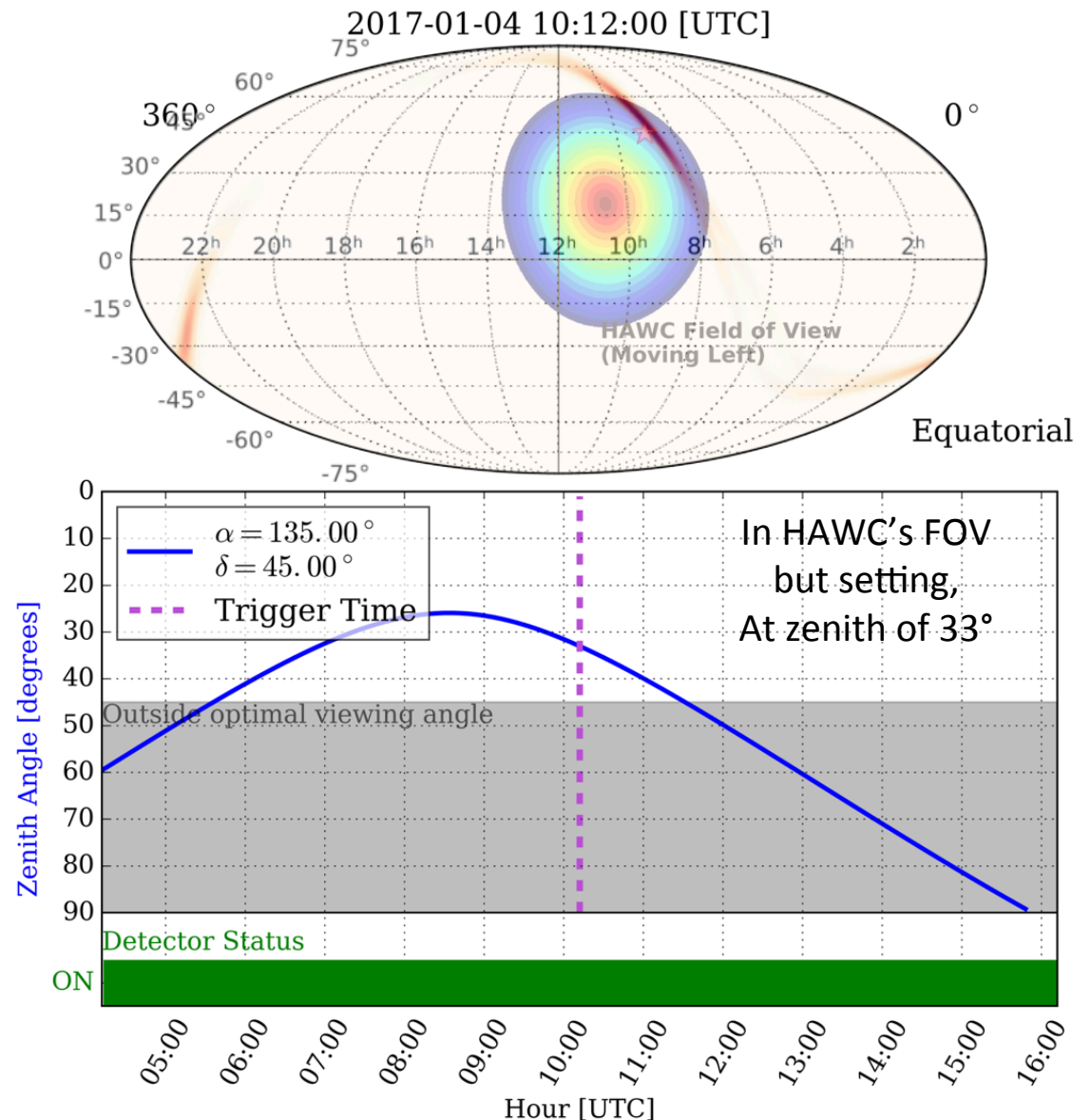
## GW170104

- Jan. 4, 2017 at 10:11:58.6 UTC
- BH-BH of  $31M_{\odot} + 19M_{\odot}$
- $z = 0.18 \pm 0.08$
- [PRL 118, 221101 \(2017\)](#)

**No transients found in the HAWC data on short (0.1 to 100s) and long (24 hours) time scales in window around the LIGO event.**

Results reported promptly to MOU partners.

K. Tollefson



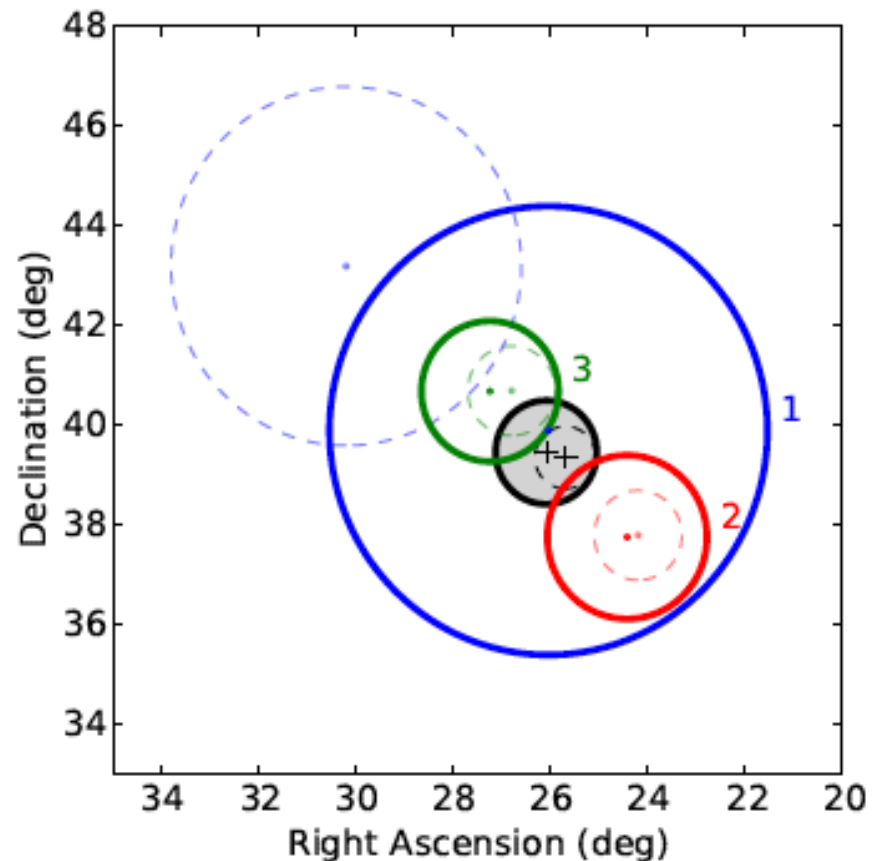


# Triplet Event

- 3  $\nu_\mu$  candidates within 100s consistent with point source origin on Feb. 17, 2016
- Probability to detect at least 1 triplet from atmospheric backgrounds is 32%

$$\pi^0 \rightarrow \gamma\gamma$$

$$\pi^\pm \rightarrow \mu^\pm \nu_\mu \rightarrow \nu_\mu \nu_e$$



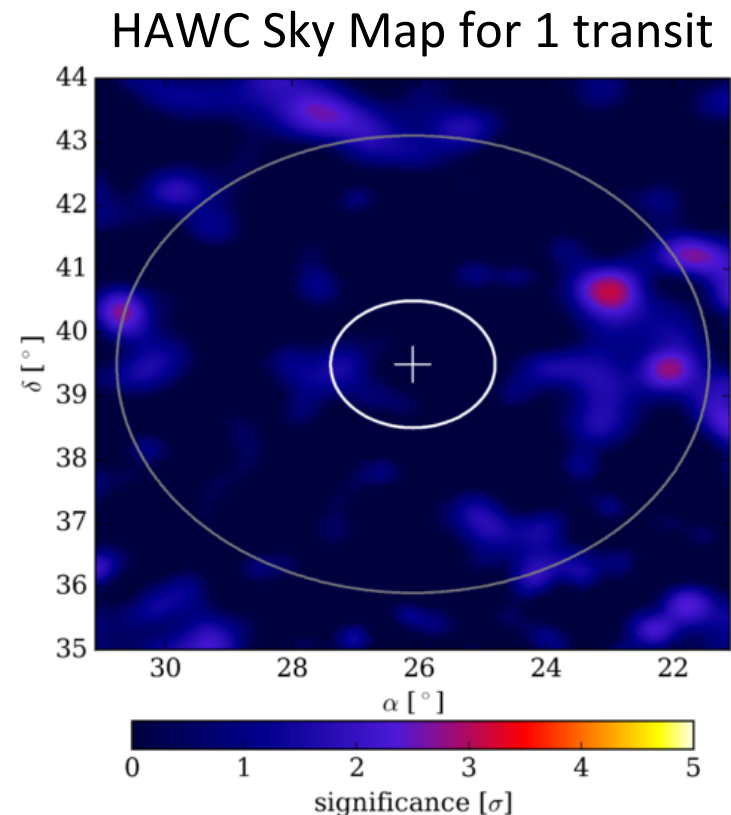
Notified other experiments to search for EM counterpart

# IceCube Triplet Event

[arXiv:1702.06131](https://arxiv.org/abs/1702.06131)



- 8 observatories did follow-ups from visible frequencies up to gamma rays
  - 3 optical: ASAS-SN, LCO, MASTER
  - 2 x-ray: XRT and BAT
  - 3 gamma-ray: FermiLAT, VERITAS, HAWC
- HAWC: Event position had just entered FOV and observed a full transit ( $\sim 6$  hours for zenith  $< 45^\circ$ )
- **No EM counterparts were observed**





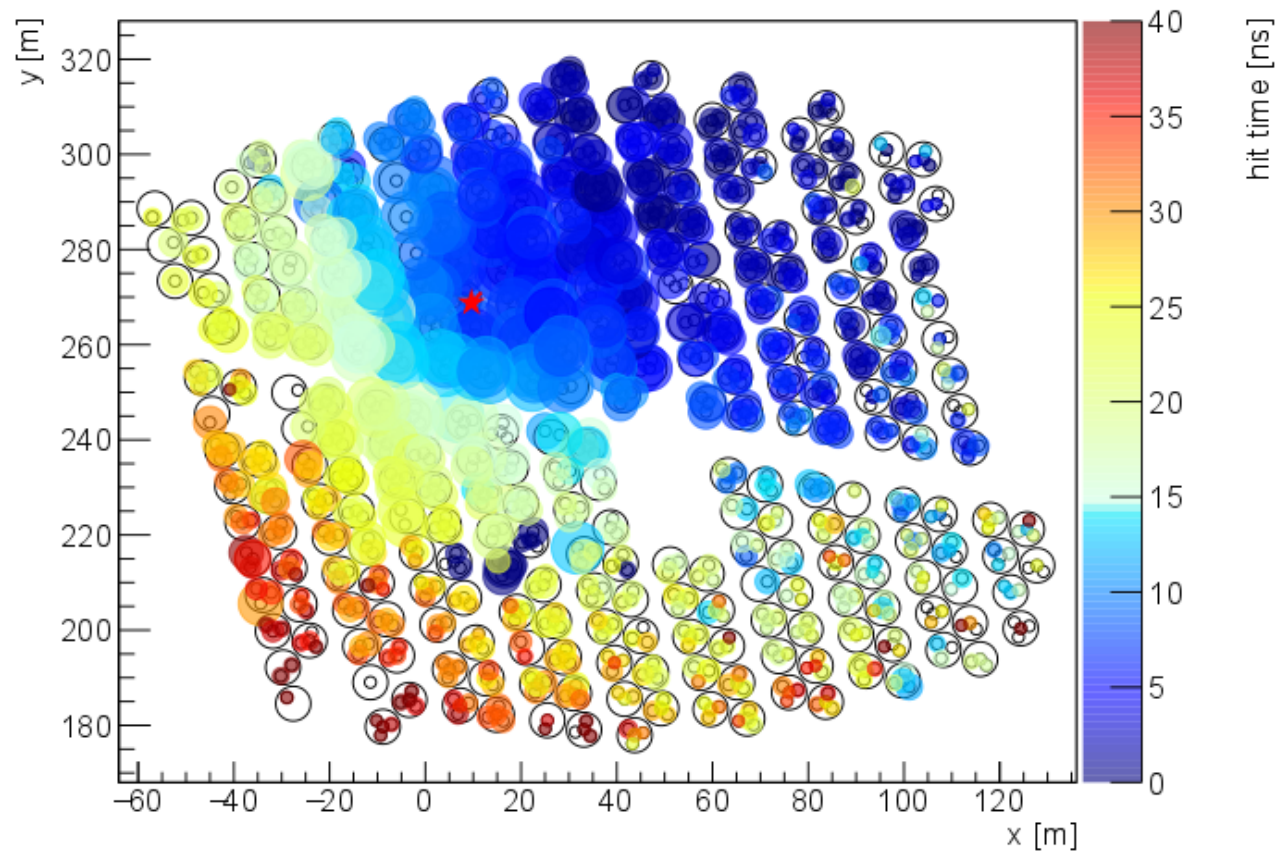


# Additional Topics

- **Dark Matter Searches**
  - See **Pat Harding's talk** in Dark Matter session tomorrow at 11:30am in Hornets Nest
  - Dwarf Spheroidal Galaxy limits submitted to ApJ [arXiv:1706:01277](https://arxiv.org/abs/1706.01277)
- Gamma-ray Burst (GRB) Searches
  - [ApJ 843:88](https://doi.org/10.1086/8118888) (July 10, 2017) with first 18 months of data
- Many more astrophysics results
  - See ICRC2017 conference proceedings for the latest <http://www.icrc2017.org/>



# Highest Energy Event (So Far)



**Neural Net Energy = 60 TeV**



# Ongoing Upgrade

350 **Outrigger** tanks will cover an area 4x HAWC  
increasing sensitivity 3-4x above 10 TeV





# Summary

- HAWC started full operations in March 2015 and is performing great!
  - Most sensitive gamma-ray experiment above 10 TeV
  - Several publications using first 1.5 years of data, more expected soon
  - Sending real-time alerts, following-up on alerts, doing multi-wavelength and multi-messenger analyses with other experiments
- Upcoming Attractions:
  - Analysis improvements, including better energy estimators
  - Outriggers will improve resolution at highest energies
  - Joint Fermi/VERITAS/HAWC working groups are standardizing and combining results