

# Status Report on the Light Maps Production

Anne CHAPPUIS – Isabelle DE BONIS – Dominique DUCHESNEAU – Laura ZAMBELLI

*WA105 SB Meeting*

*22 June 2016*

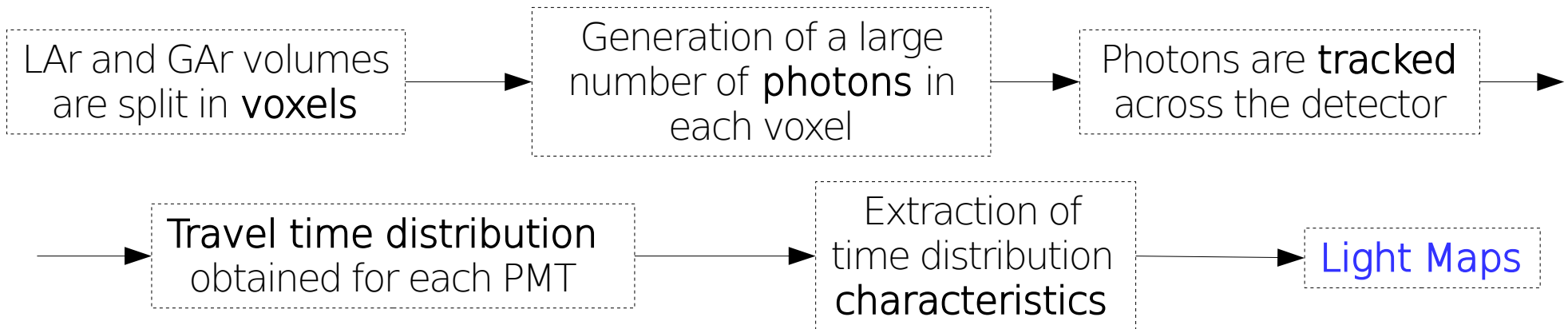


**WA105** 

The text 'WA105' is rendered in a large, bold, black serif font. To the right of the text is a detailed, 3D-style illustration of the WA105 detector structure, showing its complex geometry and components.

# Introduction

- Today we will present our progress on the whole **map production procedure**.
- **Working chain :**



- **Remark**

We are currently implementing the **TPB + PMMA layers** on the cathode, and studying comparisons between the two designs (TPB coat on PMTs)

→ We plan to **present** this work in the next SB meeting.

# Procedure for Light Maps Production

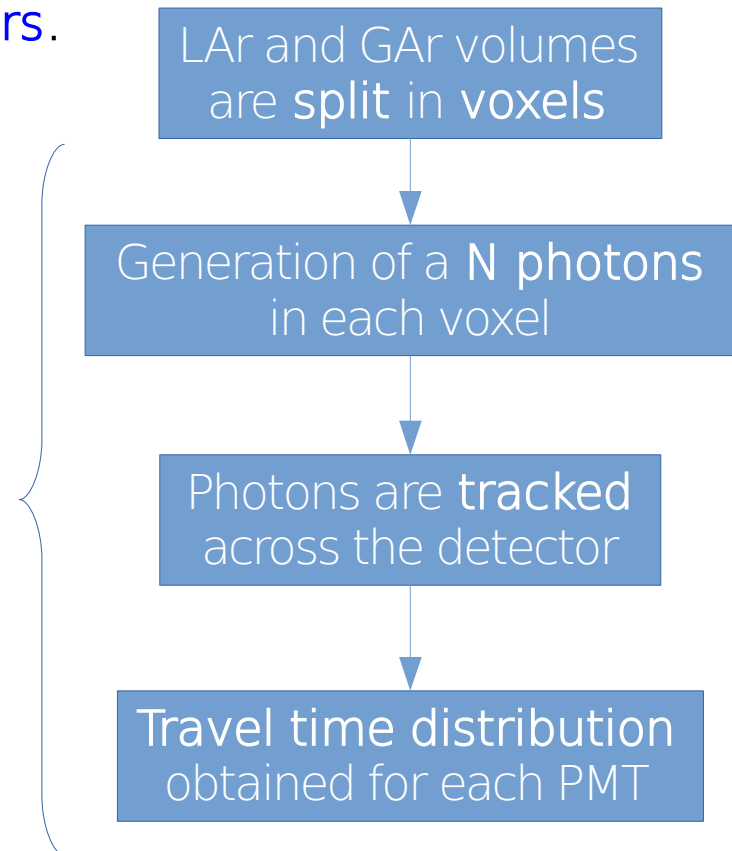
- Script which creates a 3D-histogram with **voxel centers**.

LAr and GAr volumes  
are split in voxels

# Procedure for Light Maps Production

- Script which creates a 3D-histogram with **voxel centers**.
- Scripts which loop on the voxel histogram.
  - Build a **macro file** to initialize the run :
    - Production point : voxel center
    - Number of photons (N)
    - $E = 9,69 \pm 0,22 \text{ eV}$
    - Isotropic emission
  - Launch **LightSim** with the macro file

One file **.root** per voxel.  
Each file contains **36 histograms**.



# Procedure for Light Maps Production

- Script which creates a 3D-histogram with **voxel centers**.

- Scripts which loop on the voxel histogram.

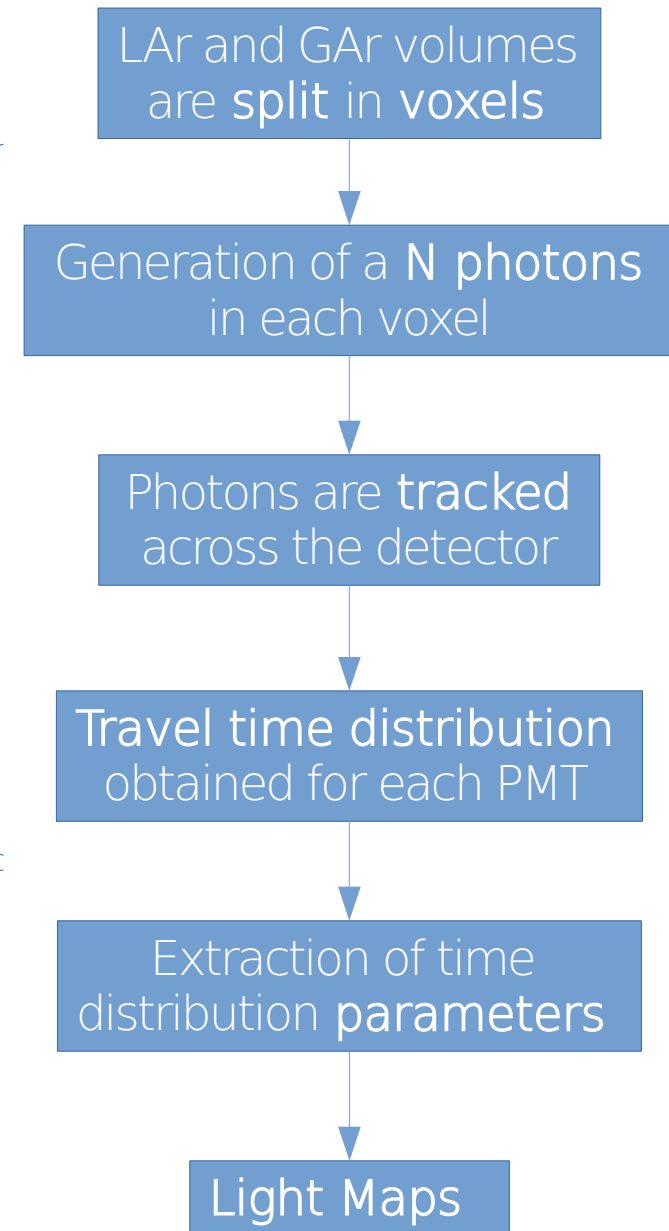
- Build a **macro file** to initialize the run :
  - Production point : voxel center
  - Number of photons (N)
  - $E = 9,69 \pm 0,22$  eV
  - Isotropic emission
- Launch **LightSim** with the macro file

One file **.root** per voxel.  
Each file contains **36 histograms**.

- Script which **analyzes** all the **.root** files.

- Loop on the voxels and 36 PMTs
- Extract **time distribution parameters** ( $n_{\text{param}}$ )
- Build a **3D-histogram** for each parameter and PMT.

Map file (.root) with  $(n_{\text{param}} \times 36)$  **3D-histograms**.



# Procedure for Light Maps Production

- Script which creates a 3D-histogram with voxel centers.

- ✓ **Scripts which loop on the voxel histogram.**

- ✓ Build a **macro file** to initialize the run :

Production point : voxel center  
Number of photons (N)  
 $E = 9,69 \pm 0,22$  eV  
Isotropic emission

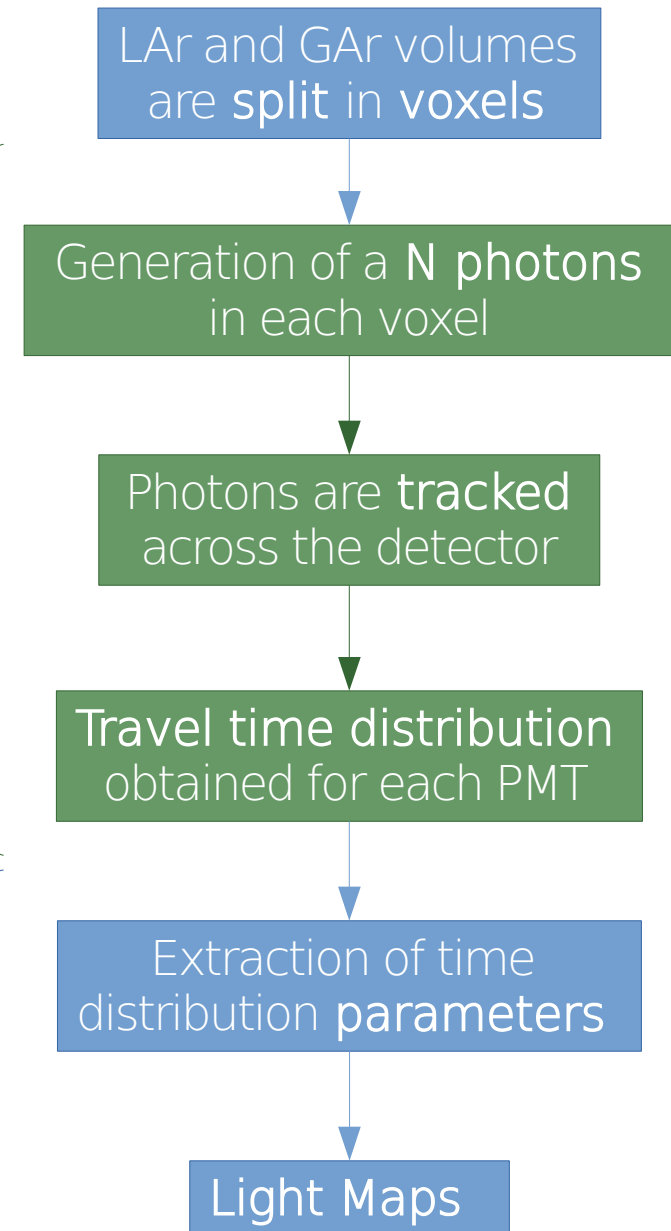
- ✓ Launch **LightSim** with the macro file

One file **.root** per voxel.  
Each file contains **36 histograms**.

- Script which analyzes all the **.root** files.

- Loop on the voxels and 36 PMTs
- Extract **time distribution parameters** ( $n_{\text{param}}$ )
- Build a **3D-histogram** for each parameter and PMT.

Map file (**.root**) with ( $n_{\text{param}} \times 36$ ) 3D-histograms.

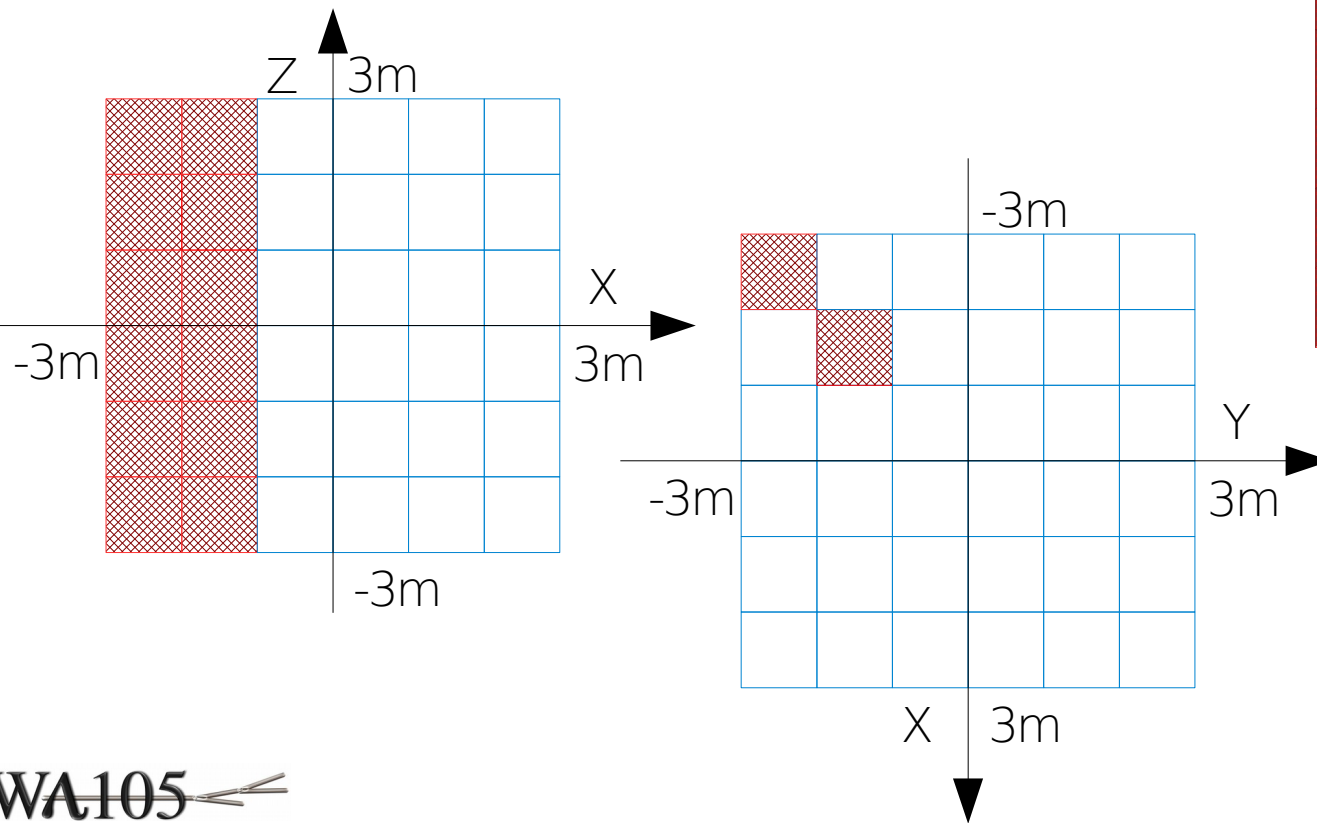
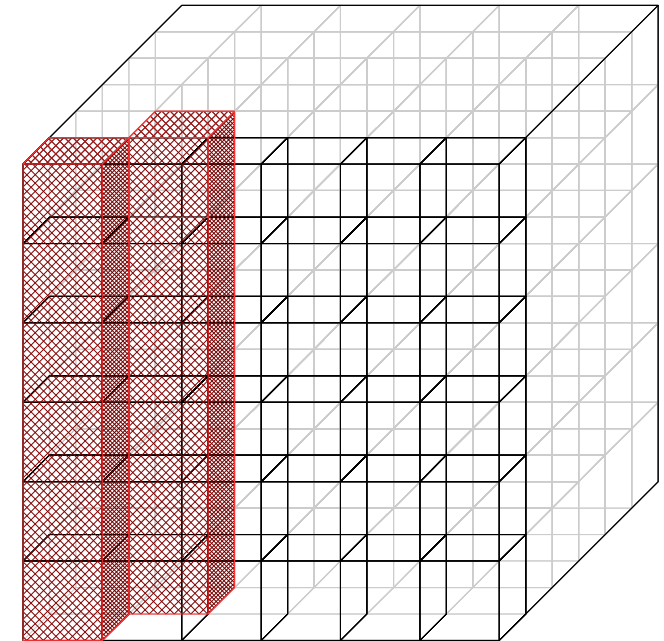


# Photon Tracking and Time Distribution Generation

- **Test** of the procedure :  
→ Use of a **simplified** voxel map description.

(1x1x1)m<sup>3</sup> voxels  
→ (6x6x6) = 216 voxels

- We look at only **two** columns (12 voxels) :
  - X and Y **fixed** at -2.5m and -1.5m.
  - $Z = \{-2.5 ; -1.5 ; -0.5 ; 0.5 ; 1.5 ; 2.5\}$  m



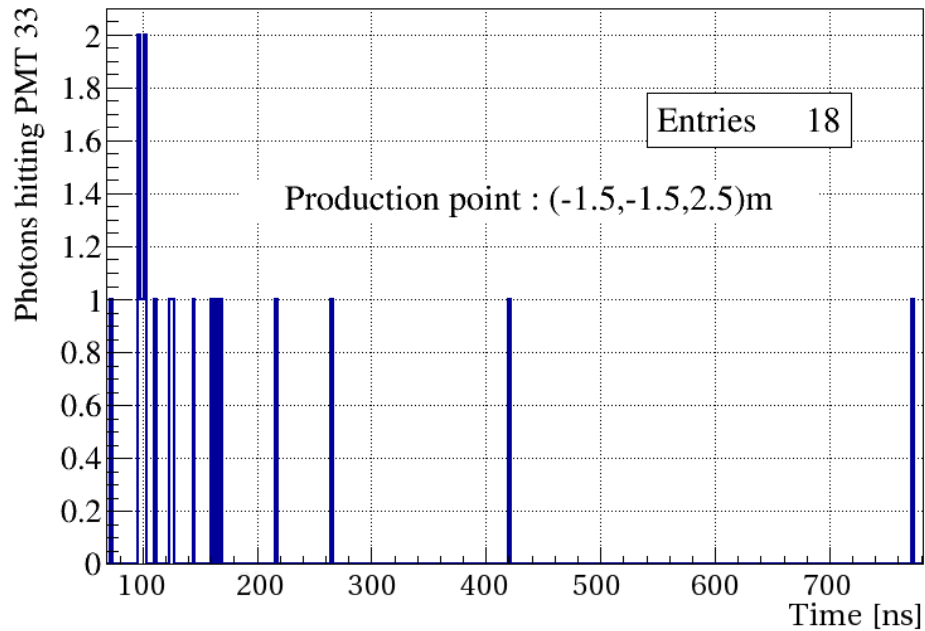
## Remark :

Exploiting the **X-Y symmetry** of the detector  
→ Generate photons on  $\frac{1}{4}$  of the detector to save time

# Photon Tracking and Time Distribution Generation

Number of photons  $N$  : First try with  $10^7$  photons

→ Most of the PMTs don't see enough photons to extract the time distribution characteristics !



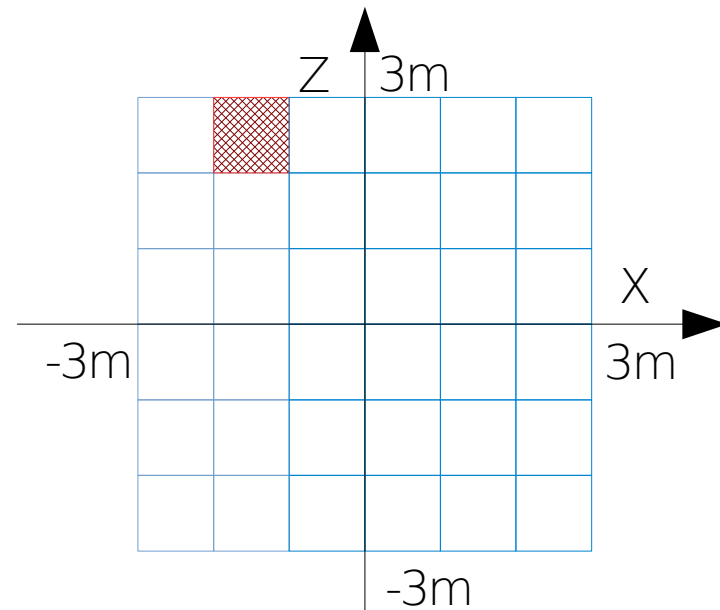
PMT Grid

1	2	3	4	5	6
7	8	9	10	11	12
13	14	15	16	17	18
19	20	21	22	23	24
25	26	27	28	29	30
31	32	33	34	35	36

For voxel at the **top** of the tank :

A lot of photons are lost during the travel in LAr !

→ Only **a few** of them finally reach the PMTs.

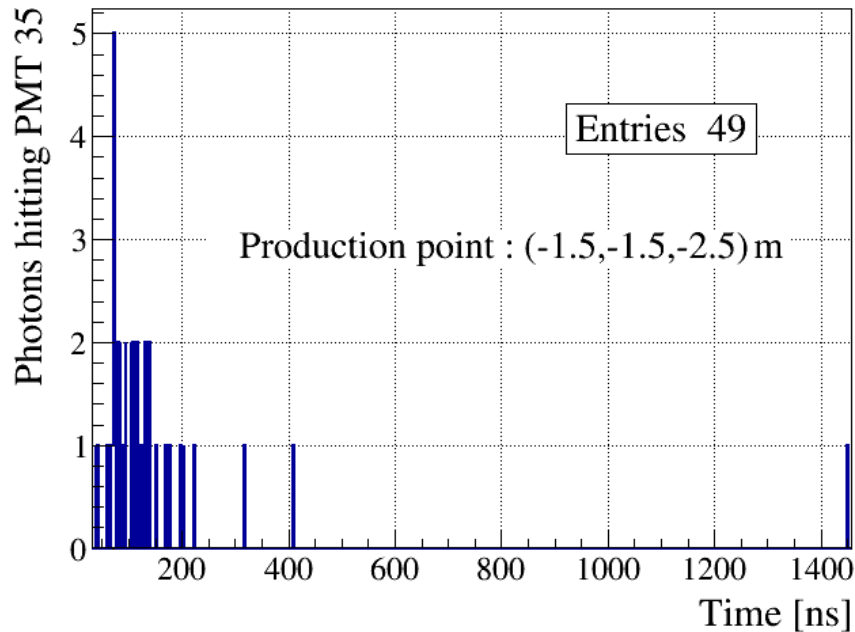




# Photon Tracking and Time Distribution Generation

Number of photons  $N$  : First try with  $10^7$  photons

→ Most of the PMTs don't see enough photons to extract the time distribution characteristics !

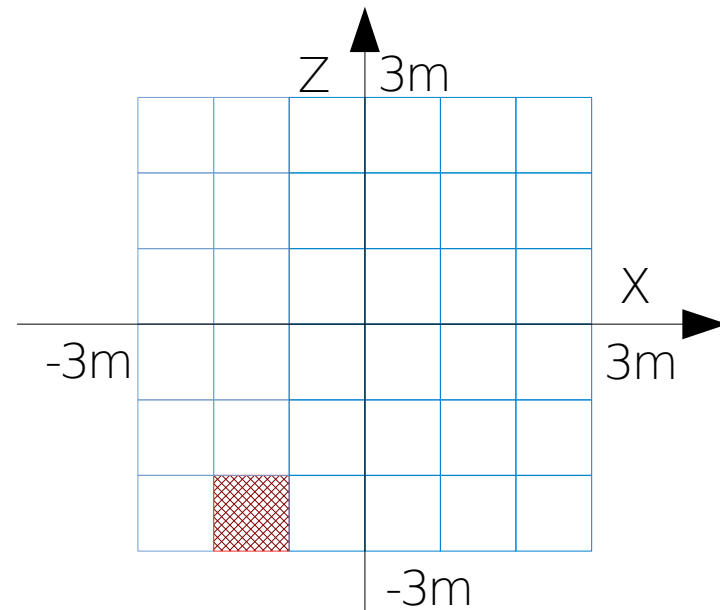


PMT Grid

1	2	3	4	5	6
7	8	9	10	11	12
13	14	15	16	17	18
19	20	21	22	23	24
25	26	27	28	29	30
31	32	33	34	35	36

Same problem for voxels at the **bottom** of the tank :

→ Most of the photons will not reach **distant PMTs.**



# Photon Tracking and Time Distribution Generation

→ It's better with  $N = 10^8$  but it is time consuming ! ( $\sim 5h$  per voxel)

Number of voxels	Queue	N	CPU	% of distributions with less than 100 photons
6	long (CC IN2P3)	$10^7$	$\sim 2h30$	2.5 %
		$10^8$	$\sim 24h$	27.1%

→ We choose  $10^8$  photons for this study, but this number could be optimized with respect to the voxel position.

# Procedure for Light Maps Production

- Script which creates a 3D-histogram with voxel centers.

- ✓ **Scripts which loop on the voxel histogram.**

- ✓ Build a **macro file** to initialize the run :

Production point : voxel center  
Number of photons (N)  
 $E = 9,69 \pm 0,22$  eV  
Isotropic emission

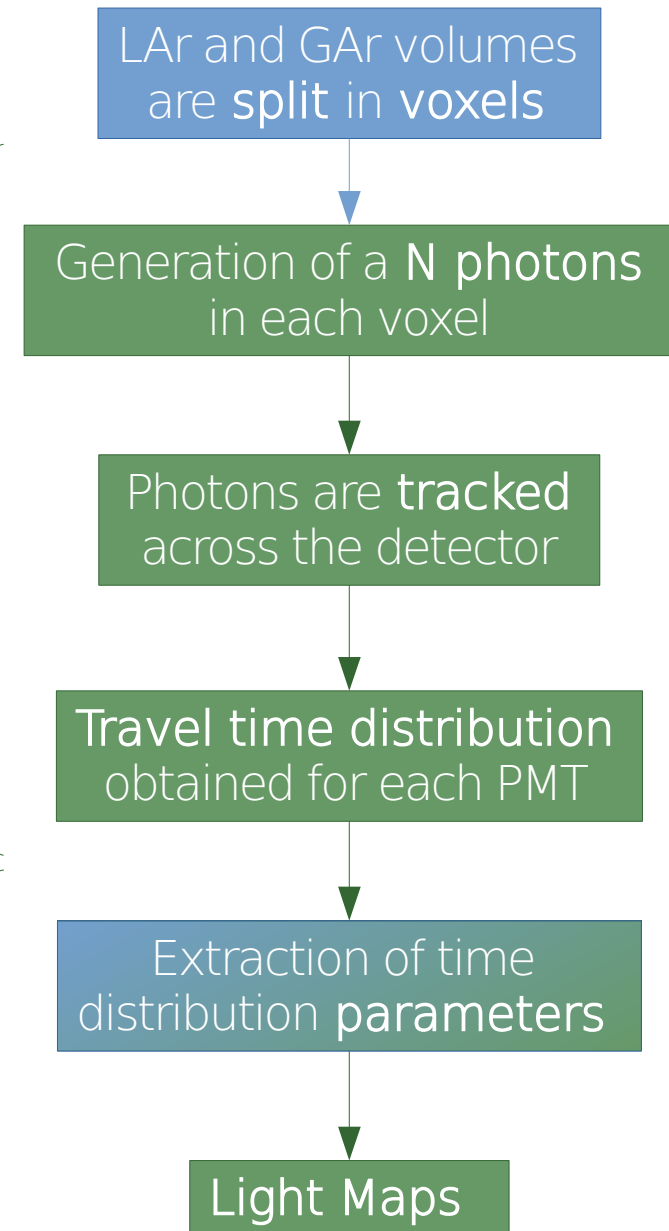
- ✓ Launch **LightSim** with the macro file

One file **.root** per voxel.  
Each file contains **36 histograms**.

- ✓ **Script which analyzes all the .root files.**

- ✓ Loop on the voxels and 36 PMTs
- ✓ Extract **time distribution parameters** ( $n_{\text{param}}$ )
- ✓ Build a **3D-histogram** for each parameter and PMT.

Map file (.root) with ( $n_{\text{param}} \times 36$ ) **3D-histograms**.



# Extraction of time distribution characteristics

→ Which **characteristics** are needed to **reconstruct** all time distributions ?

→ WA105 SB Meeting, 20 April 2016

Currently for the available maps in QSCAN and LightSim, time distributions are :

- Reconstructed with 2 parameters : tpeak and trms
- Considered as exponential function starting at tpeak, with  $\tau = \frac{trms}{\sqrt{2}}$

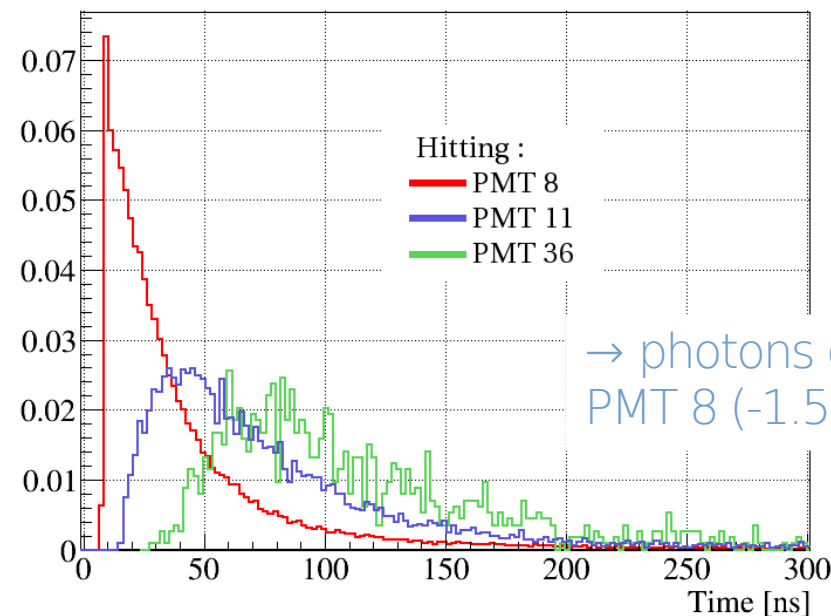
**BUT :**

PMT Grid

1	2	3	4	5	6
7	8	9	10	11	12
13	14	15	16	17	18
19	20	21	22	23	24
25	26	27	28	29	30
31	32	33	34	35	36

Y  
↓ X

Time distribution normalized to the total detected photons numbers

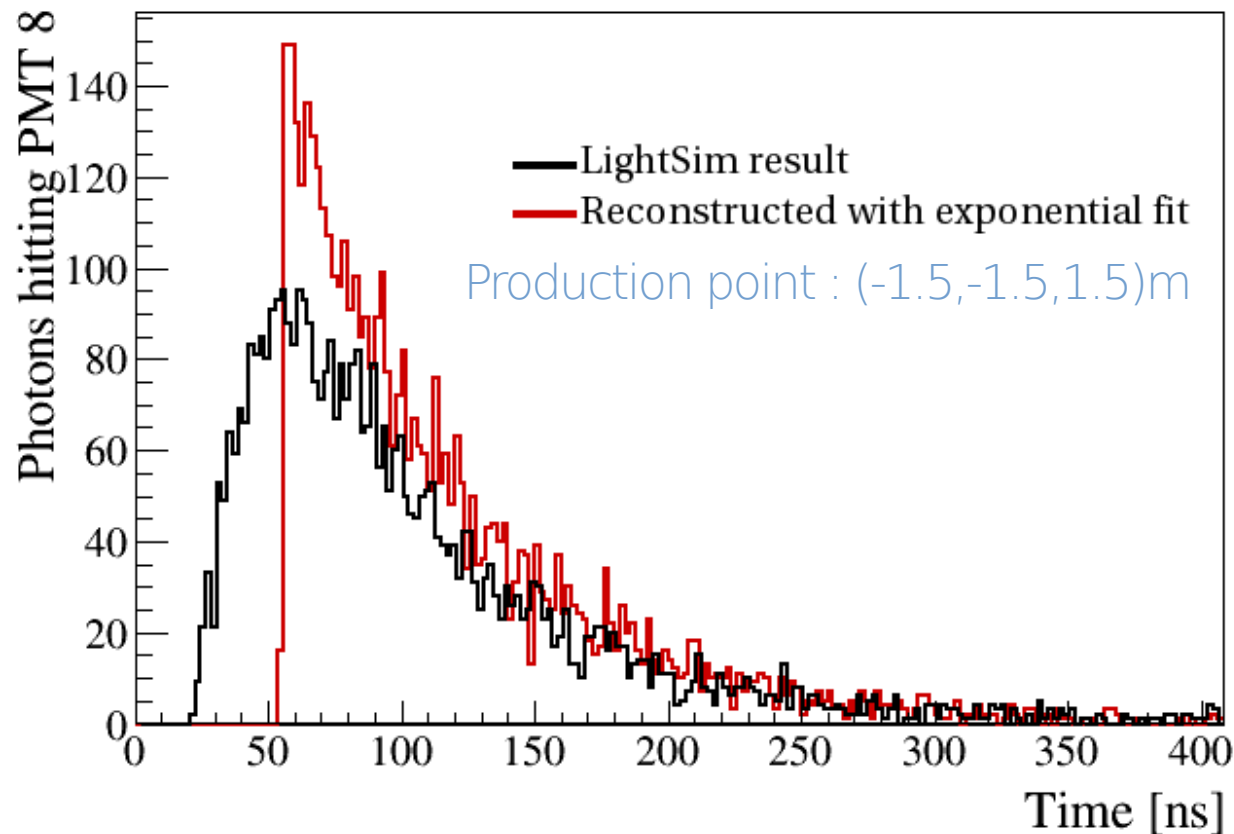


→ Exponential approximation seems reasonable for the PMT8

→ It's not the case for the more distant PMTs and upper voxels !

# Extraction of time distribution characteristics

Comparison between the distribution obtained with LightSim and the "reconstructed" one based on an exponential fit



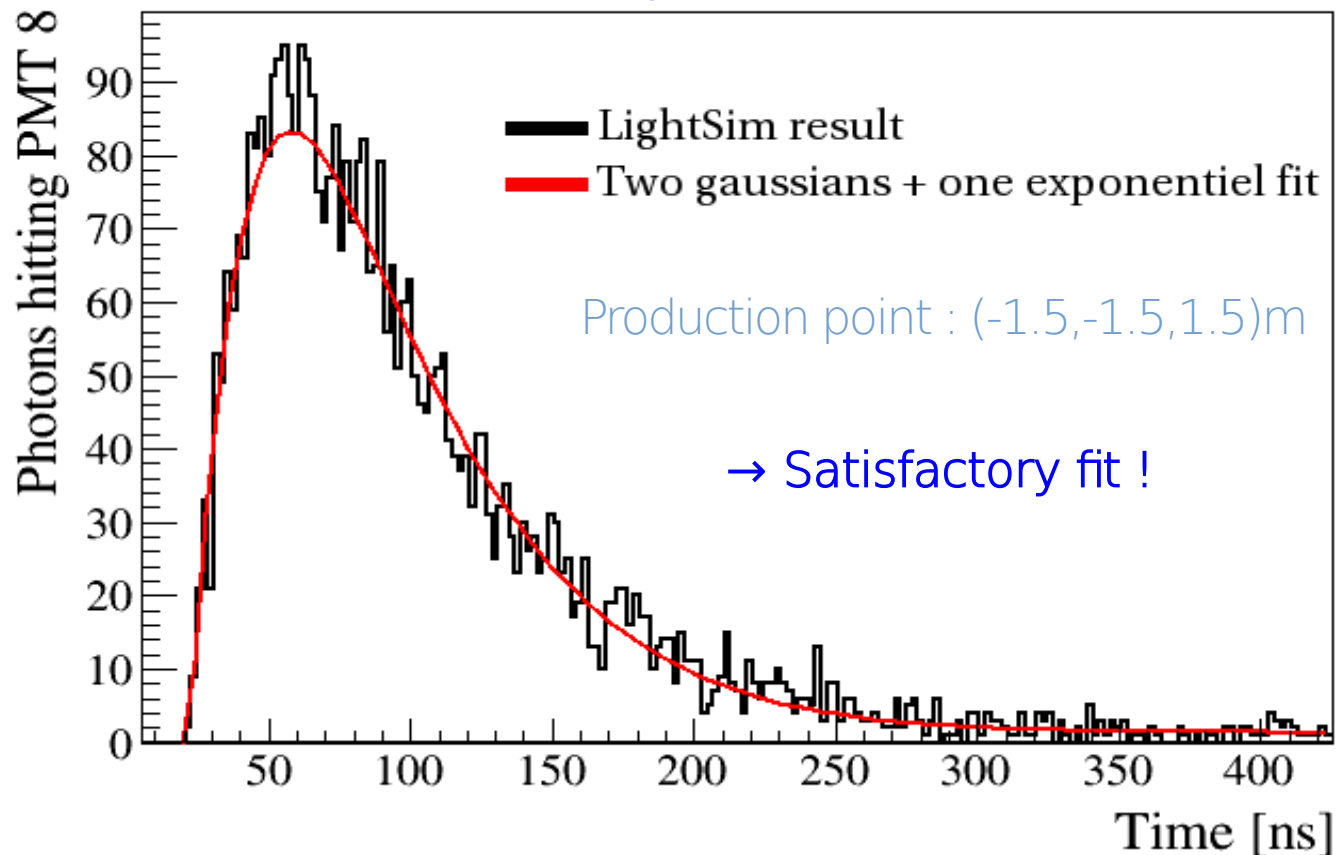
→ 2 parameters used

→ But the contribution **before tpeak** is not reconstructed correctly.

# Extraction of time distribution characteristics

New fit : using two gaussian functions + one exponential function.

→ 4 parameters



Maps construction :

- Use an **exponential** fit for closer PMTs
- Use **two gaussian + one exponential** fit for other distributions (distant PMTs and upper voxels)

# Conclusion and Perspectives

The map production procedure is completely implemented and tested :

- **Scripts** are written
- Production of a final .root with time distribution characteristics for **12 voxels**.

**Next steps :**

- Implementation of the **voxel structure** produced by Silvestro (available on svn)
- Optimization of the **number of photons** generated per voxel.
- Implementation of the **improved** time distribution **parametrisation** needed in the map.
- Implementation of the **TPB and PMMA layers** on the cathode. **(already well-advanced)**
- Modify the **implemented geometry** to match the latest detector design.  
(ex: PMT-cathode distance, add CRP plane, etc)  
→ According to CERN-SPSC-2016-017