



# Space Charge Effect Calibration: Planning

#### Michael Mooney BNL

ProtoDUNE Measurements Meeting August 9<sup>th</sup>, 2016



### Introduction



- We have heard recently that it is very likely that there will be no UV laser system at protoDUNE with which to calibrate out space charge effects (SCE), among other things
  - This will impact our calibration strategy significantly!
- Placement of CRT panels important consideration for properly calibrating out SCE
  - It has not been shown yet at e.g. MicroBooNE that we can obtain a clean sample of t<sub>o</sub>-tagged tracks with the light-collection system
- Highlight considerations for cosmic ray tagger (CRT) in this talk, including placement and how to do calibration
  - <u>Jacob's talk</u>: preliminary answers to partial set of relevant questions
  - <u>This talk</u>: more questions to be answered; also calibration strategy, impact on CRT needs, and required inputs





# Calibrating w/ Muon Tracks



• Two samples of  $t_0$ -tagged tracks can provide SCE corrections:

- <u>Single tracks</u> enable corrections at TPC faces by utilizing endpoints of tracks (correction vector approximately orthonormal to TPC face)
- <u>Pairs of tracks</u> enables corrections in TPC bulk by utilizing • unambiguous point-to-point correction looking at track crossing points
- Require high-momentum tracks (plenty from cosmics, beam halo) 5

### **Corrections at TPC Faces**





- Claim on previous slide is that the correction at TPC faces using single tracks is the correction vector obtained by projecting the track end point onto the closest TPC face
- True at most boundaries as only one SCE component is large
- TPC edges (boundaries in Y <u>and</u> Z) will still need pairs of tracks

### Why Crossing Points?



- As Igor pointed out at protoDUNE Science Workshop, a single laser track is not enough to obtain the SCE correction vector
- Principle applies to calibration with muon tracks as well!

### Front/Back CRT Panels





- Discussed with
   Flavio possible
   arrangement of CRT
   panels on front and
   back of detector
- 8+8 panels on front,
  8+8 panels on back
- Would be useful to tag t<sub>o</sub> for both muon halo tracks and cosmic muon tracks
- 32 panels in total, but possibly more to use elsewhere?

### **Track Samples**



- With anode planes and front/back CRT panels, you get three samples of t<sub>o</sub>-tagged tracks:
  - Cosmics crossing both anode planes (left)
  - Cosmics crossing a CRT panel (middle)
  - Muon halo tracks crossing a CRT panel (right)



- Combining these t<sub>o</sub>-tagged track samples, we get complete coverage for single tracks!
- However, if you want to calibrate in the bulk, you need track pairs, and they should be at relatively large angle w.r.t. each other
- Near top of TPCs would have much lower statistics CRT coverage on top helps (muon halo, tag from top CRT)
  - Front/back CRT cosmics will help fill in these areas as well (not shown) 10



- Combining these t<sub>o</sub>-tagged track samples, we get complete coverage for single tracks!
- However, if you want to calibrate in the bulk, you need track pairs, and they should be at relatively large angle w.r.t. each other
- Near top of TPCs would have much lower statistics CRT coverage on top helps (muon halo, tag from top CRT)
  - Front/back CRT cosmics will help fill in these areas as well (not shown) 11





- We can perform a calibration of SCE w/o a laser system using cosmic tracks,muon halo tracks IF we can tag t<sub>o</sub> with high reliability
  - Use both single tracks and track pairs for calibration of TPC faces and TPC bulk, respectively
- Best way to do this is extensive CRT system
  - Light-collection system likely not able to reliably (high degree of certainty as required in calibration) tag  $t_o$
- Installing CRT panels on front/back of detector in discussion
  - Need to know number of tracks we can utilize for the measurement per unit time including all possible calibration samples
  - Jacob has looked at **cosmic tracks** passing through front/back CRT
  - Need to combine this with look at e.g. anode-anode crossing tracks, but preliminary conclusion is that top CRT panels probably not necessary, but helpful (more statistics in crucial regions)
  - Also need input about **beam halo** rate and spatial distribution!





# BACKUP SLIDES





- Space charge: excess electric charge (slow-moving ions) distributed over region of space due to cosmic muons passing through the liquid argon
  - Modifies E field in TPC, thus track/shower reconstruction
  - Effect scales with L<sup>3</sup>, E<sup>-1.7</sup>





### **SpaCE:** Overview

- Code written in C++ with ROOT libraries
- ◆ Also makes use of external libraries (ALGLIB)
- Primary features:
  - Obtain E fields analytically (on 3D grid) via Fourier series
  - Use interpolation scheme (RBF radial basis functions) to obtain E fields in between solution points on grid
  - Generate tracks in volume line of uniformly-spaced points
  - Employ **ray-tracing** to "read out" reconstructed {x,y,z} point for each track point – RKF45 method
- First implemented effects of uniform space charge deposition without liquid argon flow (only linear space charge density)
  - Also can use **arbitrary space charge configuration**
    - Can model effects of liquid argon flow (however, interpretation is difficult)

### **Impact on Track Reco.**



- Two separate effects on reconstructed **tracks**:
  - Reconstructed track shortens laterally (looks rotated)
  - Reconstructed track bows toward cathode (greater effect near center of detector)
- Can obtain straight track (or multiple-scattering track) by applying corrections derived from data-driven calibration



### Compare to FE Results: E

- Looking at central z slice (z = 5 m) in x-y plane (MicroBooNE)
- Very good shape agreement compared to Bo Yu's 2D FE (Finite Element) studies
- Normalization differences understood (using different rate)



### Compare to FE Results: E

- Looking at central z slice (z = 5 m) in x-y plane (**MicroBooNE**)
- Very good shape agreement here as well
  - Parity flip due to difference in definition of coordinate system





- Compare 30 x 30 x 120 field calculation (left) to 15 x 15 x 60 field calculation with interpolation (right) for MicroBooNE
- Include analytical continuation of solution points beyond boundaries in model – improves performance near edges





#### Sample "Cosmic Event" **K**VEN BROOK NATIONAL LABORATORY MicroBooNE [<u></u>10 z E<sup>10</sup> **Nominal Drift** Half Drift 6-Field Field 5-5 500 V/cm 250 V/cm 4-3. 2-Ж 0

0.5

2.5

X [m]

1.5

1

0.5

1

1.5

2

1 Im

**0**.5

2.5

X [m]

1.5

1

0.5

0

0.5

1

1.5

2

1 (m)

0.5



### Complications



- Not accounting for non-uniform charge deposition rate in detector → significant modification?
- ◆ Flow of liquid argon → likely significant effect!
  - Previous flow studies in 2D... differences in 3D?
  - Time dependencies?



### Liquid Argon Flow



BROOKHAVEN

NATIONAL LABORATORY

## **Smoking-gun** Test for SCE



- ◆ Can use cosmic muon tracks for calibration
  - Possibly sample smaller time scales more relevant for a particular neutrino-crossing time slice
  - Minimally: data-driven cross-check against laser system calibration
- ◆ Smoking-gun test: see lateral charge displacement at track ends of non-contained cosmic muons → space charge effect!
  - No timing offset at transverse detector faces (no E<sub>v</sub> distortions)
  - Most obvious feature of space charge effect









- oCE to produce digride coment mana
- Can use SpaCE to produce displacement maps
  - Forward transportation:  $\{x, y, z\}_{true} \rightarrow \{x, y, z\}_{sim}$ 
    - Use to **simulate** effect in MC
    - Uncertainties describe accuracy of simulation
  - **Backward transportation**:  $\{x, y, z\}_{reco} \rightarrow \{x, y, z\}_{true}$ 
    - Derive from calibration and use in data or MC to correct reconstruction bias
    - Uncertainties describe remainder systematic after bias-correction
- Two principal methods to encode displacement maps:
  - **Matrix representation** more generic/flexible
  - Parametric representation (for now, 5<sup>th</sup>/7<sup>th</sup> order polynomials) fewer parameters
    - Uses matrix representation as input → <u>use for LArSoft</u> <u>implementation</u>

### Nominal SP Geometry





- Nominal SP protoDUNE geometry:
  - Drift (X): 3.6 m
  - Height (Y): 5.9 m
  - Length (Z): 7.0 m
- Dimensions used for simulations slightly different (to simplify calculations):
  - Drift (X): 3.6 m
  - Height (Y): 6.0 m
  - Length (Z): 7.2 m











#### BROOKHAVEN SP/DP Comp. – E Field Dist. BROOKHAVEN



### BROOKHAVEN SP/DP Comp. – E Field Dist. BROOKHAVEN



#### BROOKHAVEN SP/DP Comp. – Spatial Dist. BROOKHAVEN



# SP/DP Comp. – Spatial Dist. BROOKHAVEN





### **Modified** Geometry



- Modified ProtoDUNE geometry:
  - Drift (X): 2.2 m
  - Height (Y): 5.9 m
  - Length (Z): 7.0 m
- Dimensions used for simulations slightly different (to simplify calculations):
  - Drift (X): 2.4 m
  - Height (Y): 6.0 m
  - Length (Z): 7.2 m











### Bulk Calibration w/ Cosmics BROOKHAVE

- Fill in displacement correction map gaps using **cosmic muons**
- <u>One idea</u>: correction from center of line connecting points of closest approach (separation **d**) between two tracks (<u>before</u> and <u>after</u> SCE)
  - Get "true" muon track from PCA fit to already-calibrated points
  - Weight each contribution by e<sup>-d/D</sup> (where D is tunable parameter)
  - Use only high-momentum cosmics to minimize MCS effects
- Relies on first correcting points at boundaries, high stats to average out MCS, and knowing track t<sub>o</sub>



**M. Mooney** 

arxiv:1511.01563

**Update Correction <u>to</u> Point P**