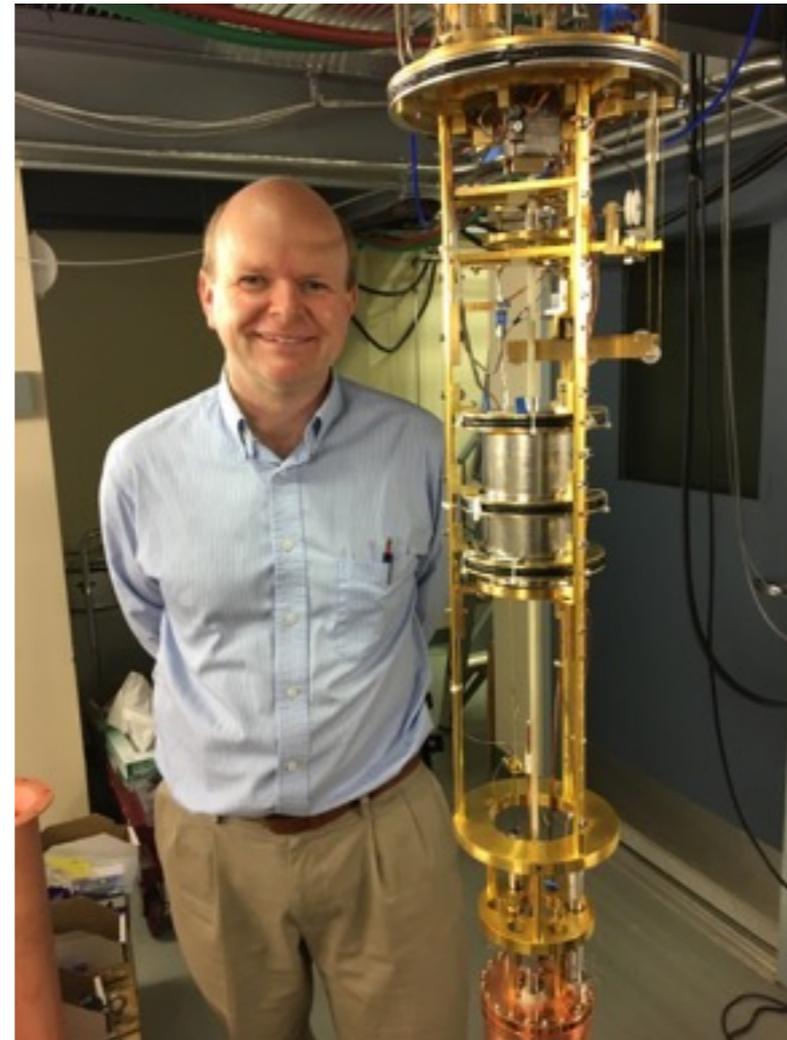


# Overview of axions, axion-like particles

Ed Daw

The University of Sheffield



# The Strong CP problem

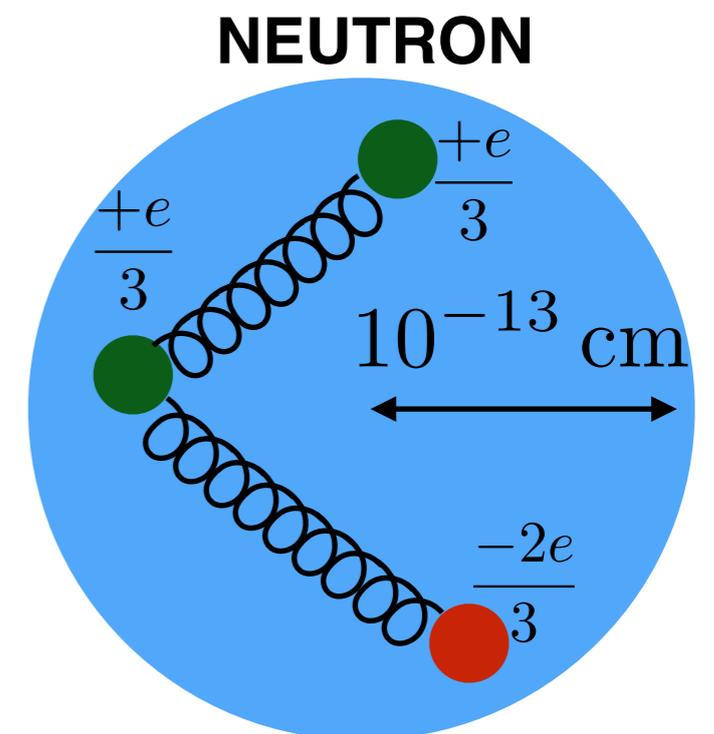
Standard model symmetry group is  $\underbrace{SU(3)}_{\text{NON-ABELIAN}} \times \underbrace{SU(2)}_{\text{NON-ABELIAN}} \times \underbrace{U(1)}_{\text{ABELIAN}}$

$$\mathcal{L}_{\text{CPV}} = \frac{(\Theta + \arg \det M)}{32\pi^2} \vec{E}_{\text{QCD}} \cdot \vec{B}_{\text{QCD}}$$

**CP CONSERVING!**
         
 **CP VIOLATING**
         
 **CP CONSERVING**

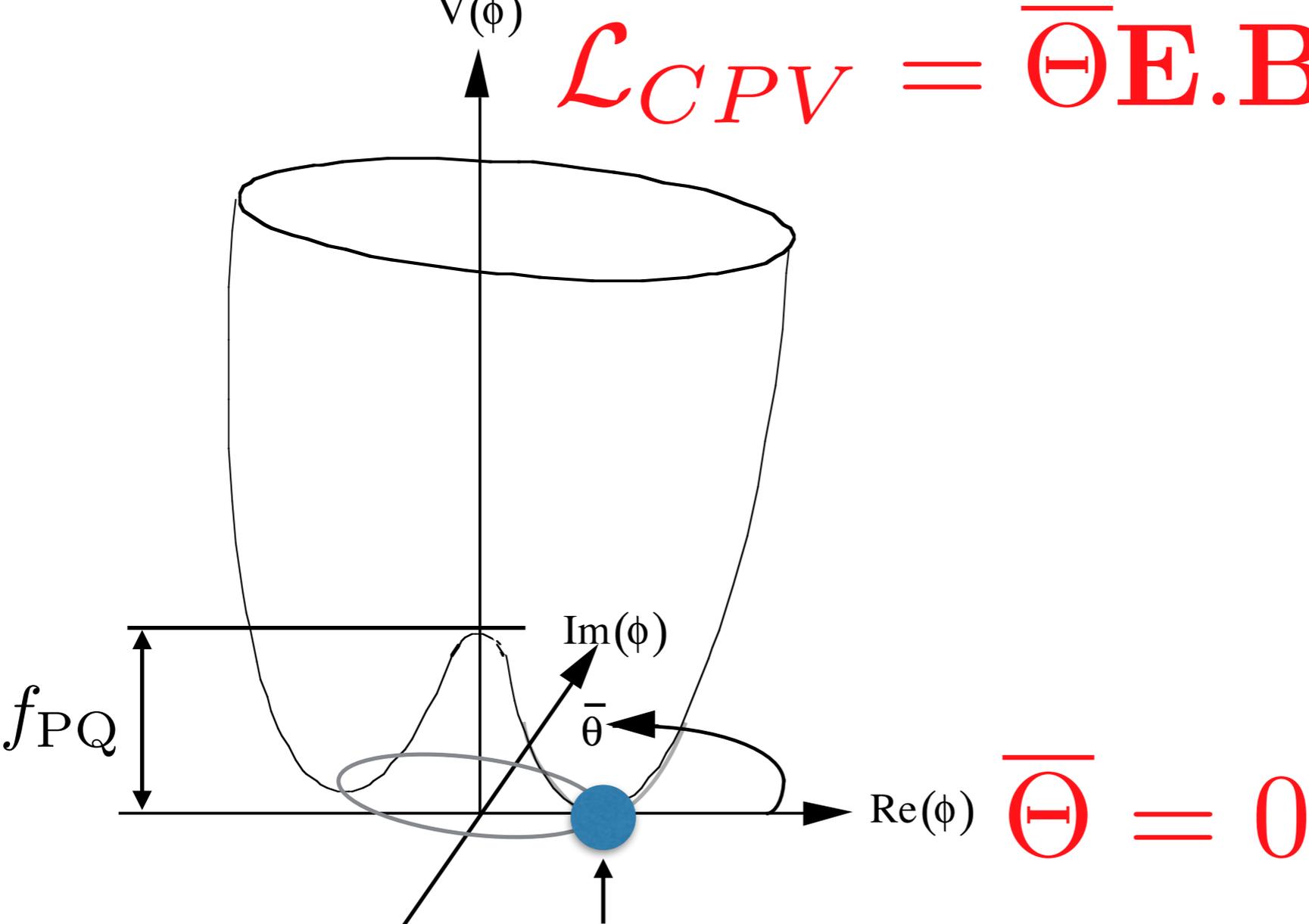
Evidence for CP conservation in the SU(3) strong interactions from multiple measurements of neutron and nuclear electric dipole moments. For example, neutron EDM  $< 10^{-26}$  e-cm.

Even simple dimensional arguments show that this is unexpected. Why do the intricate SU(3) QCD interactions conserve CP when the less intricate SU(2) QED interactions do not? This is the strong CP problem.

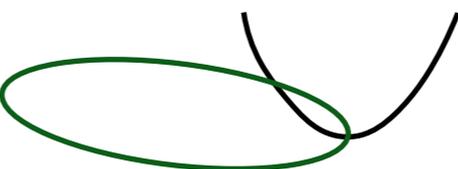


# The Peccei Quinn Mechanism Axions and ALPs

$$\mathcal{L}_{CPV} = \overline{\Theta} \mathbf{E} \cdot \mathbf{B}$$



**About Minimum: small curvature (hence small mass) with respect to  $\bar{\theta} = \arg(\phi)$  large curvature (hence large mass) with respect to  $\text{Re}(\phi)$**

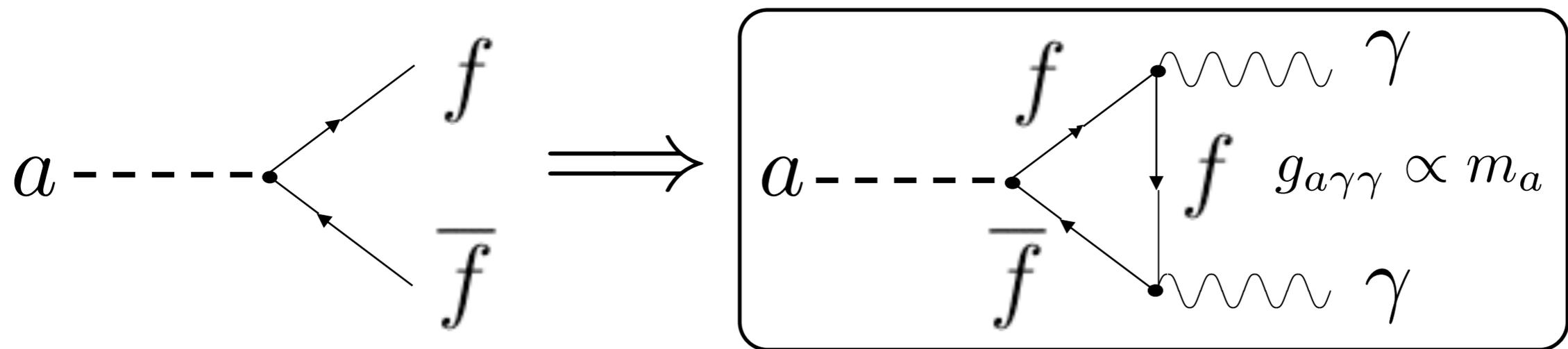
Axion DOF  ALP DOF



The  
University  
Of  
Sheffield.

# Axion Phenomenology

The axion is a pseudoscalar; has the same quantum numbers as the  $\pi^0$ , and the same interactions, but with strengths scaled to the axion mass

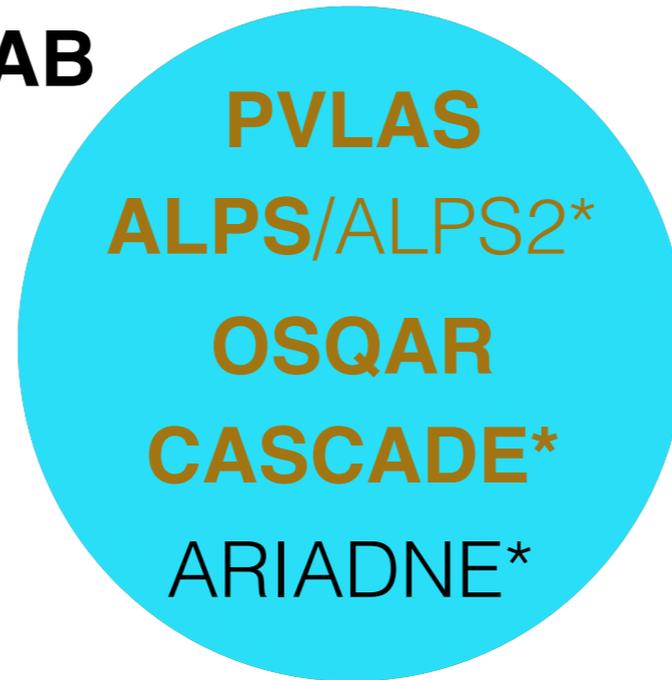


$$f_{PQ} \propto \frac{1}{m_a}$$

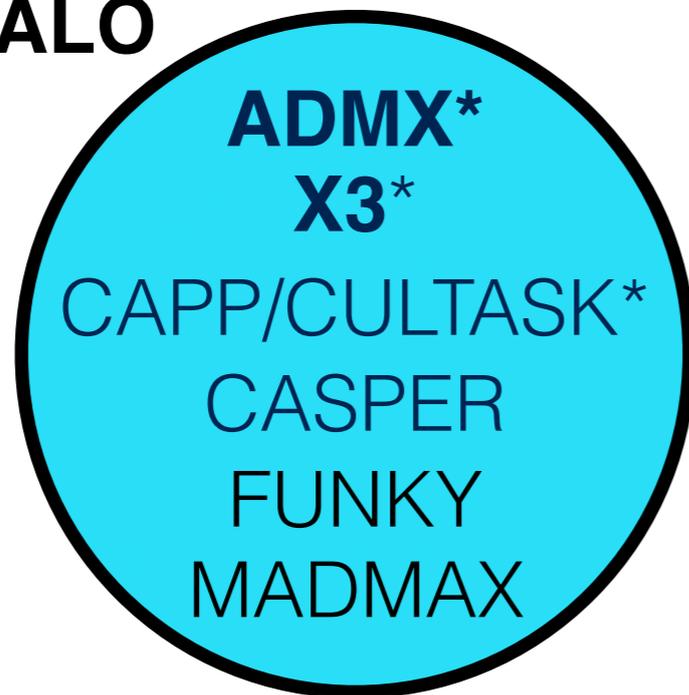
$$\Omega_{PQ} \propto \frac{1}{m_a^{\frac{7}{6}}}$$

# Axion Sources for Lab Searches

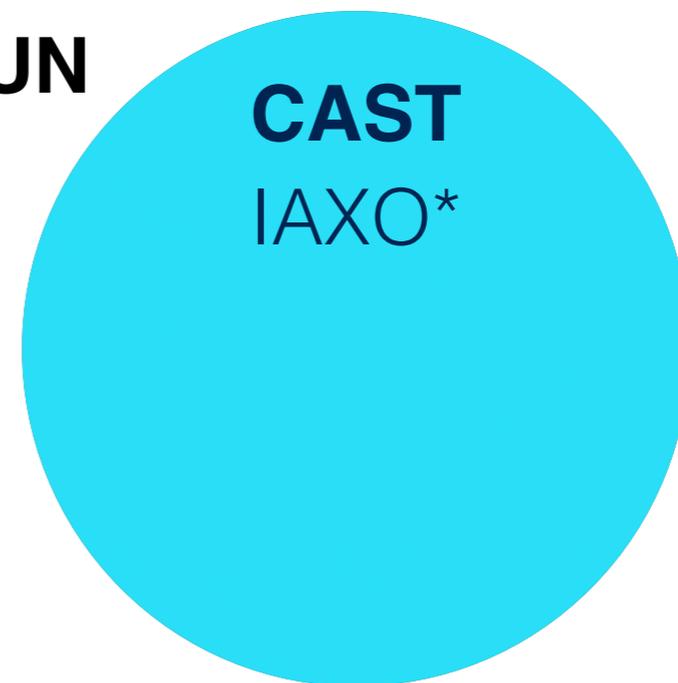
## LAB



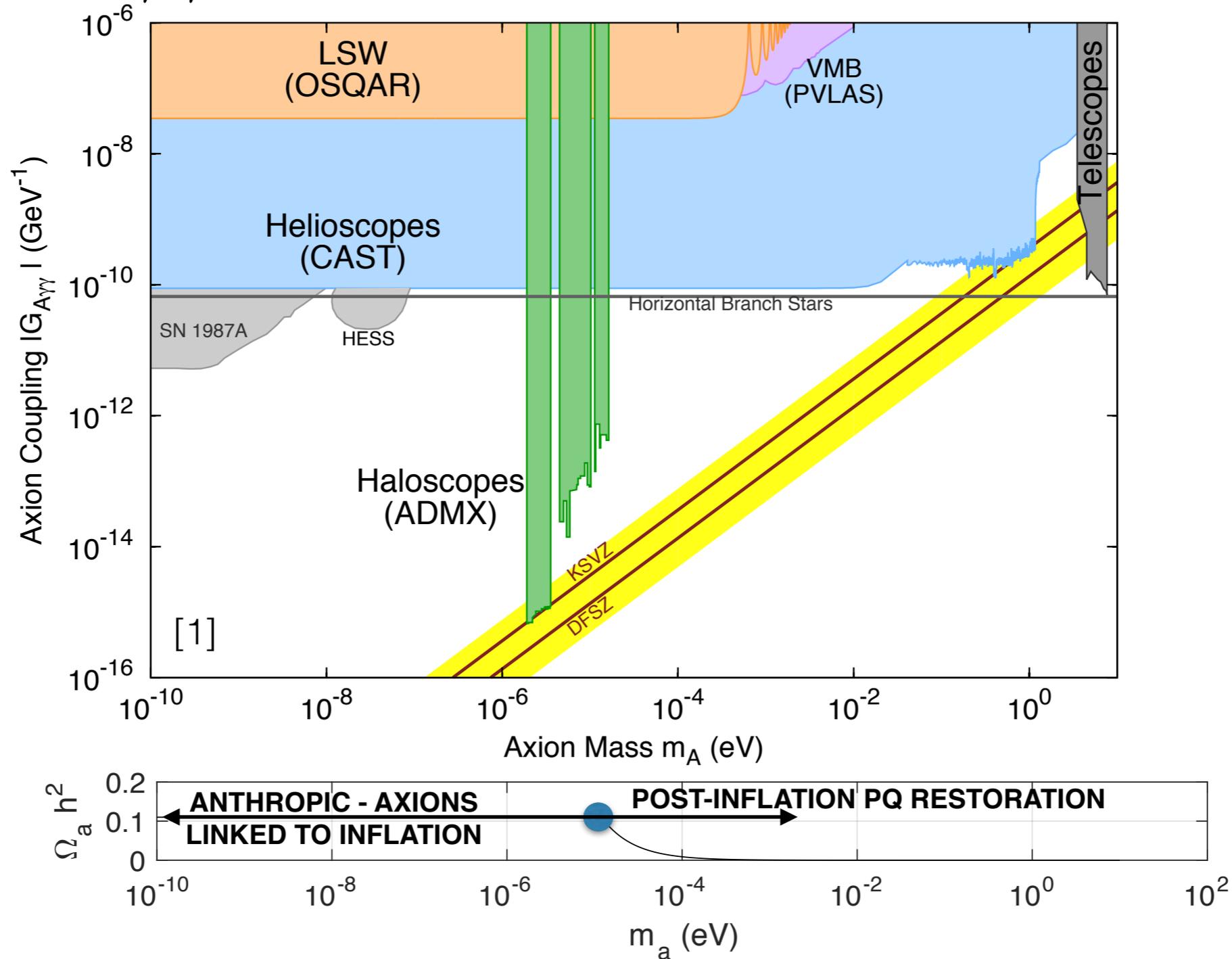
## HALO



## SUN

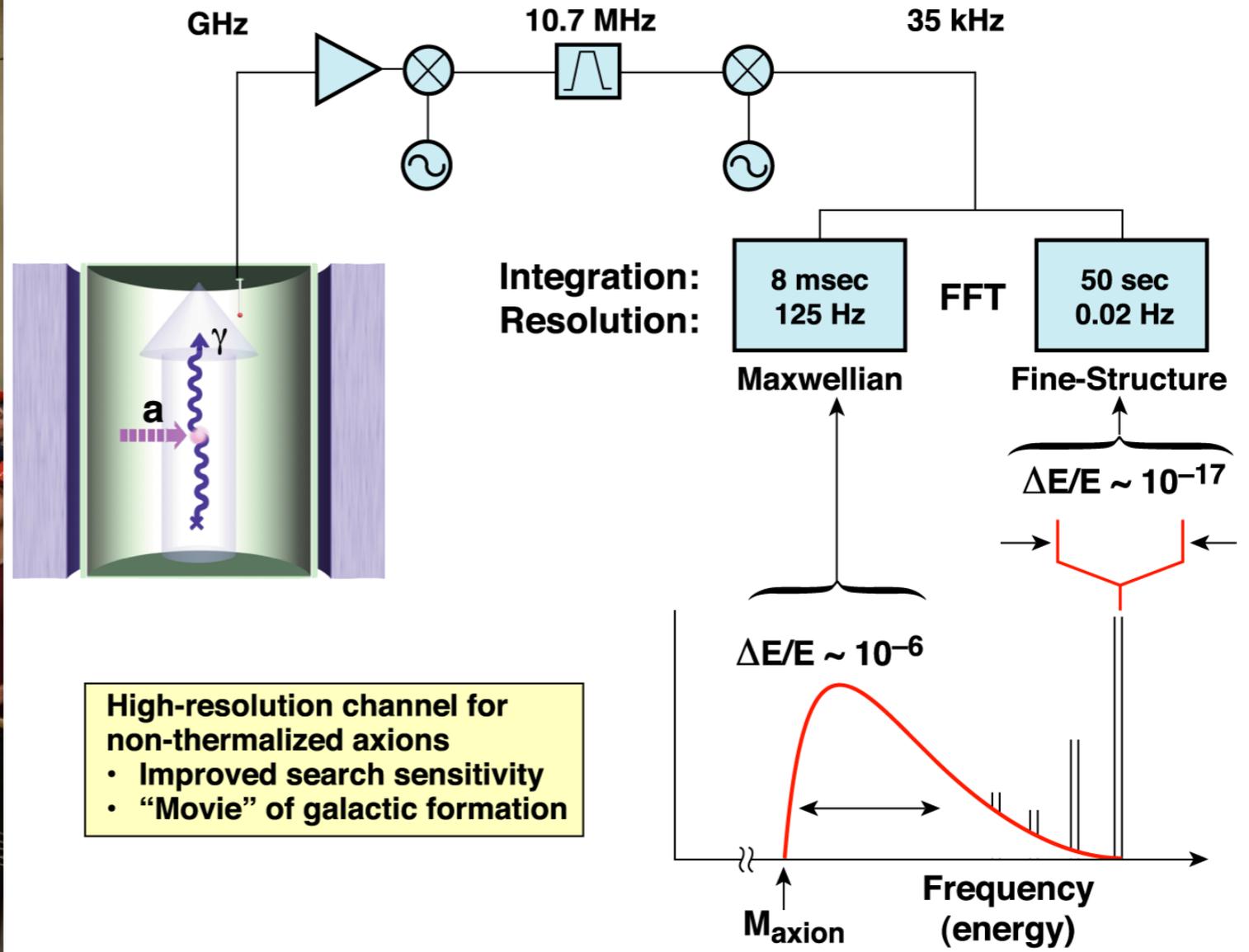


# $g_{a\gamma\gamma}$ vs. $m_a$ parameter space



[1] K.A. Olive et al. (Particle Data Group), Chin. Phys. C, 38, 090001 (2014) and 2015 update 2016 revision by A. Ringwald, L. Rosenberg, G. Rybka,

# Resonant Cavity Detectors - **ADMX**



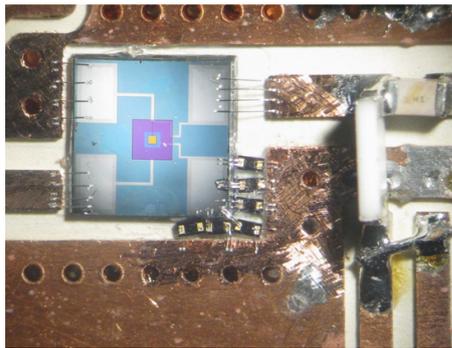
High-resolution channel for non-thermalized axions

- Improved search sensitivity
- "Movie" of galactic formation

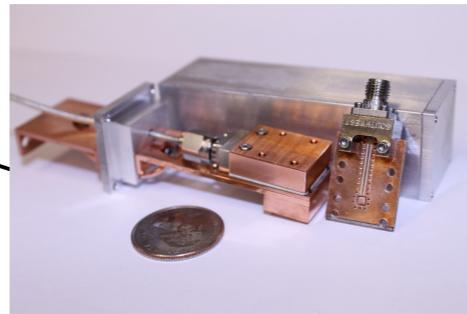
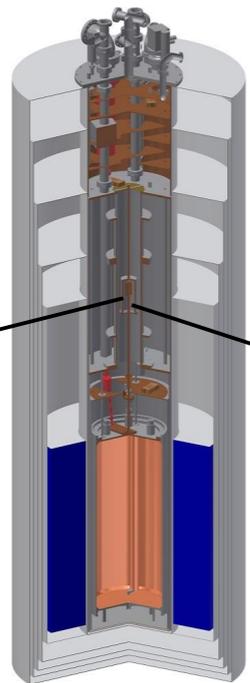
QUANTITY IN BRACKETS IS

$$P_\gamma = \left( \frac{g_{a\gamma\gamma}^2 \rho_a \hbar^2}{m_a^2 c} \right) 2\pi c^2 \epsilon_0 B_0^2 V f_{010} \nu_a Q$$

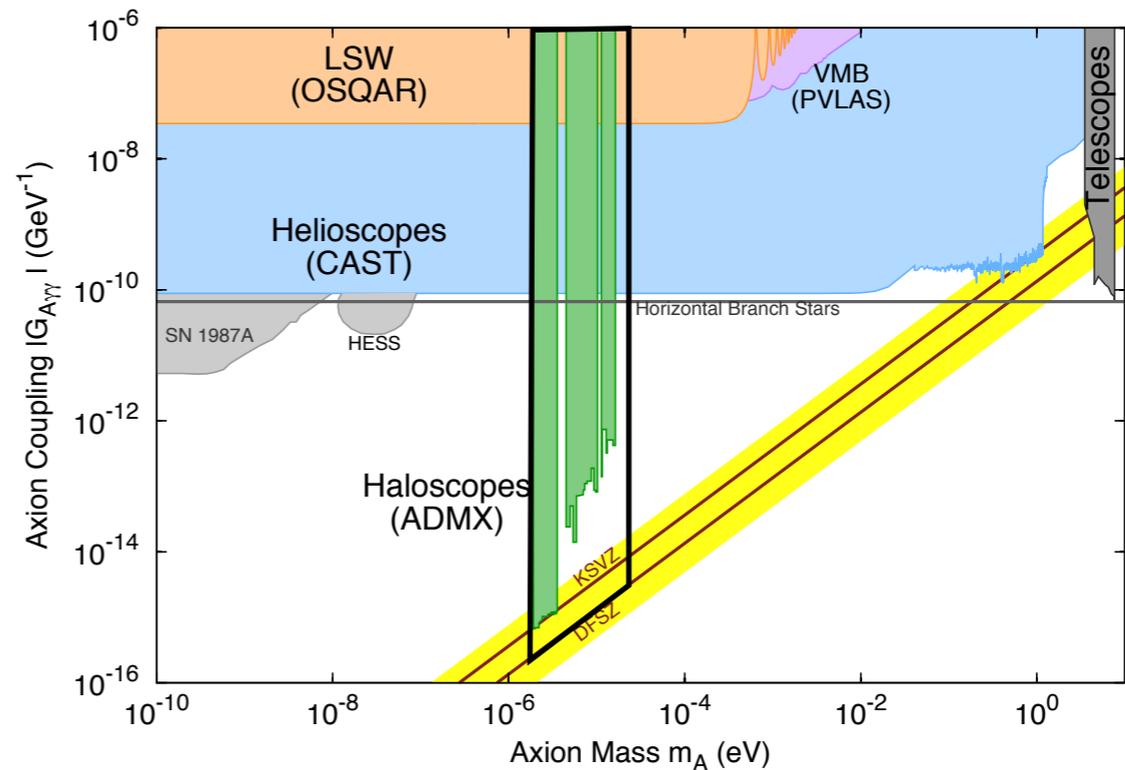
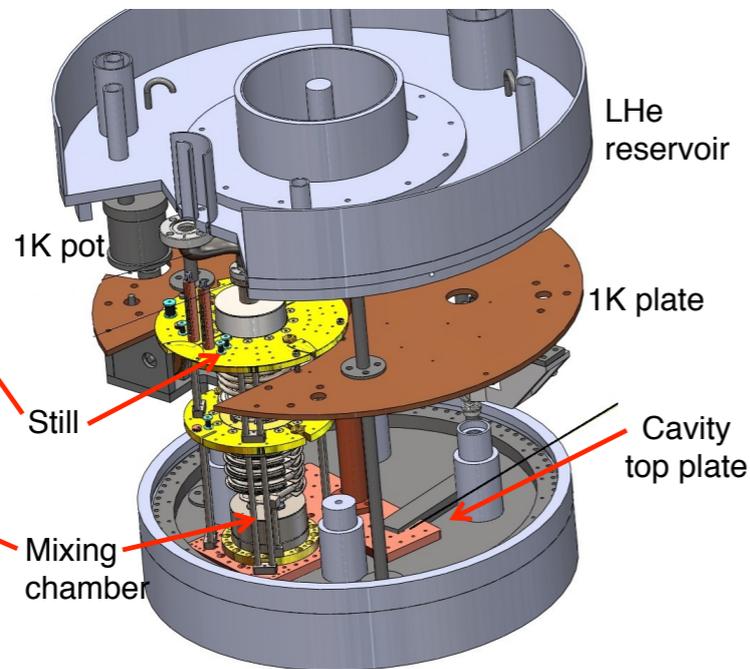
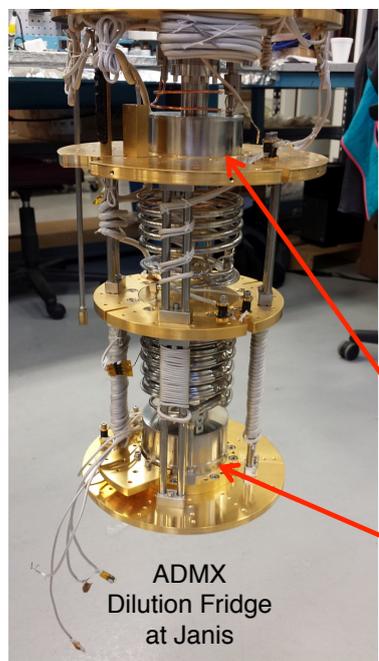
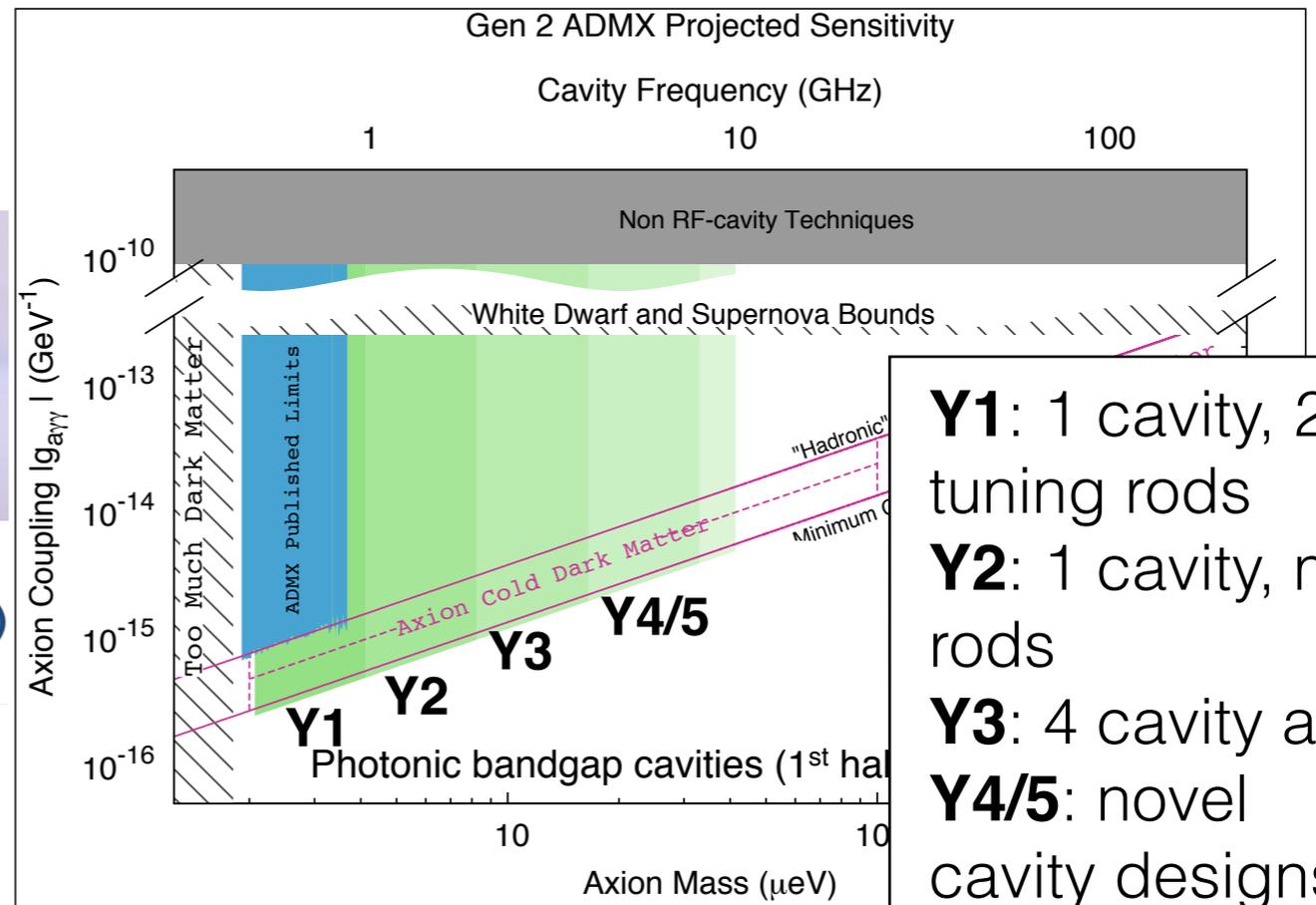
# ADMX II Reach



**SQUIDs**  
(at lower frequencies)

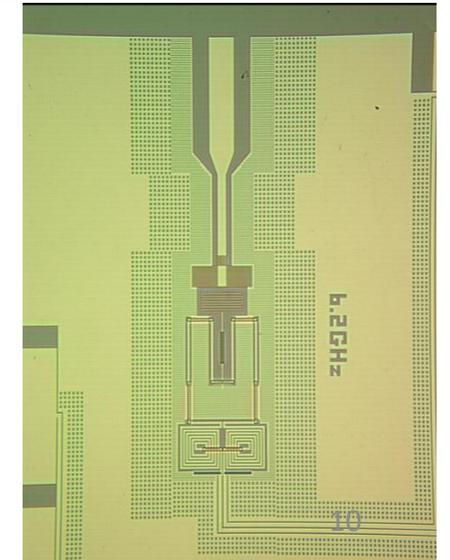


**"JPAs"**  
(at higher frequencies)

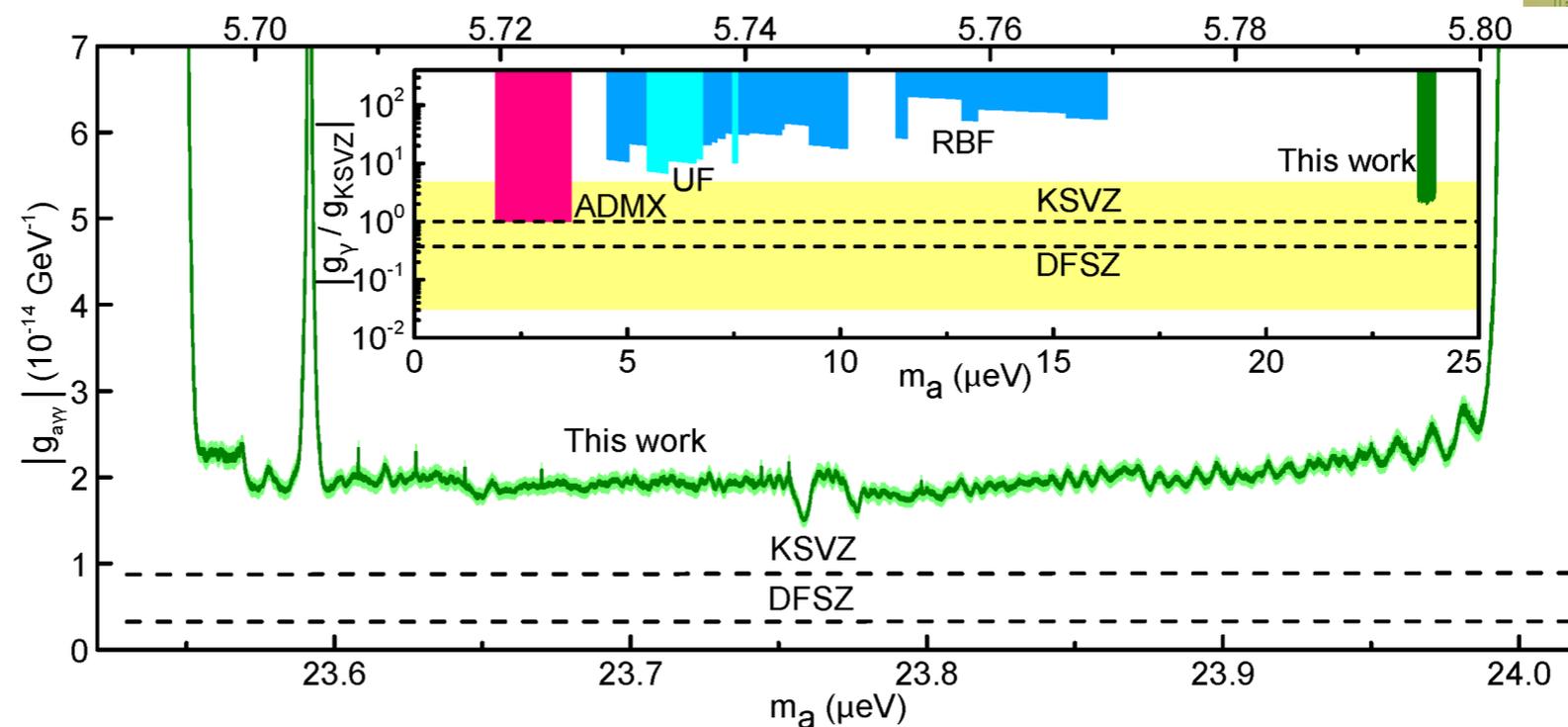


# X3 - Yale (formerly ADMX-HF)

- Probes the 5-25 GHz (20-100  $\mu\text{eV}$ ) axion mass region.
- Uses a 25 mK dilution refrigerator.
- Uses a highly uniform 9 T magnet.
- Uses Josephson Parametric Amplifiers (JPAs)

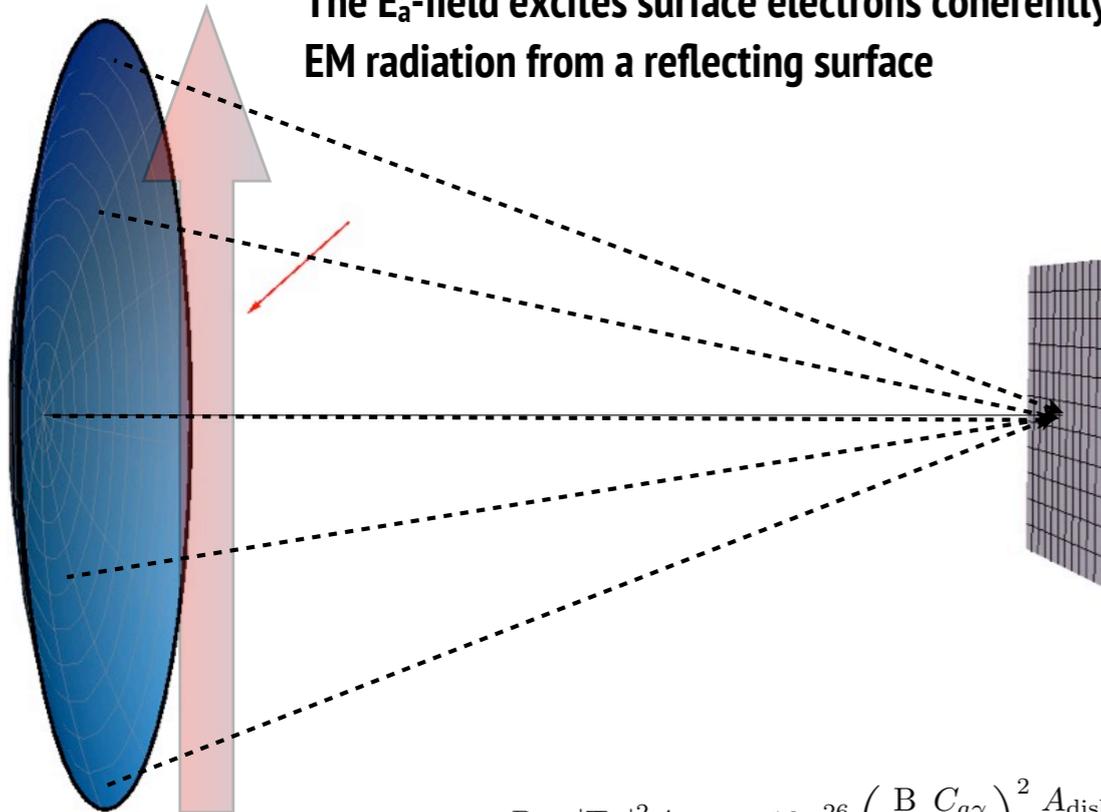


[From a talk by Tim Shokair for X3 collaboration, 2015.]

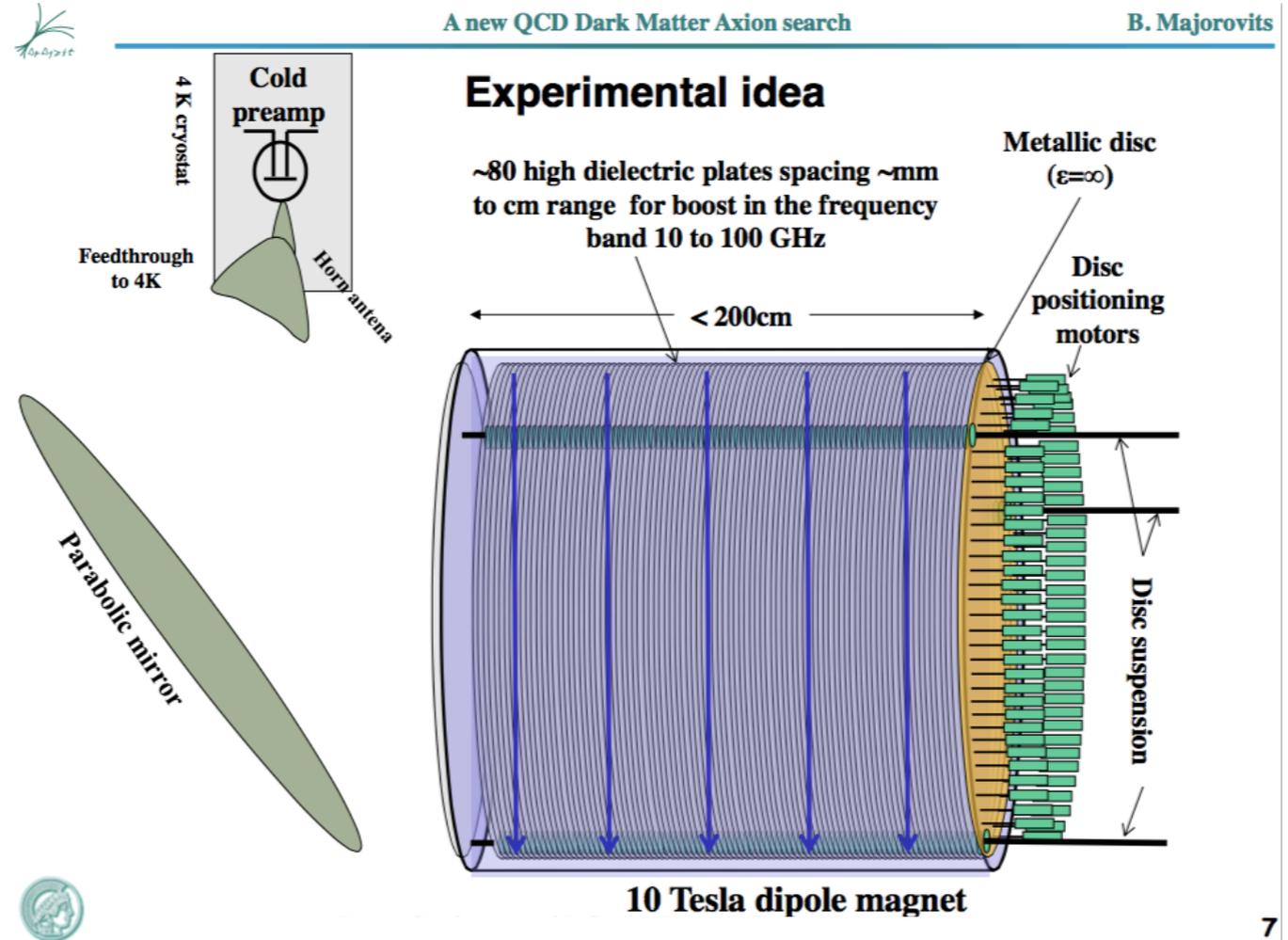


# Dish reflector searches

The  $E_a$ -field excites surface electrons coherently  
EM radiation from a reflecting surface



$$P \sim |\mathbf{E}_a|^2 A_{\text{dish}} \sim 10^{-26} \left( \frac{B}{5T} \frac{C_{a\gamma}}{2} \right)^2 \frac{A_{\text{dish}}}{1 \text{ m}^2} \text{ Watt}$$



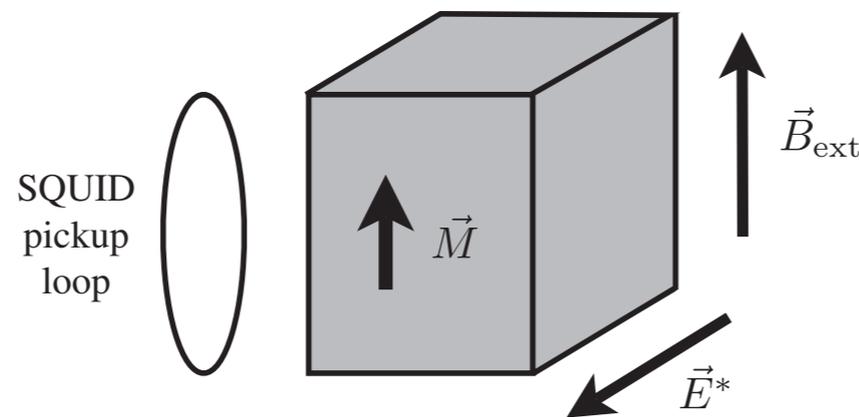
From talk by Javier Redondo, Bela Majorovits,  
Warsaw Workshop on Non-Standard Dark Matter, June 2016

Disc reflector idea employed by proposed MADMAX, FUNKY

# Nuclear EDM Searches

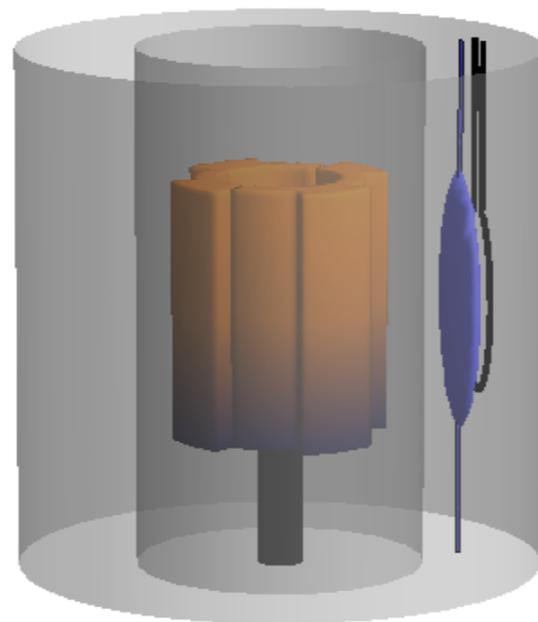
CASPER is a search for axion to nucleon coupling generating an oscillating EDM in a sample. Sensitive to KSVZ axions,

$$m_a \leq 10^{-9} \text{ eV}$$



PHYS. REV. X **4**, 021030 (2014)

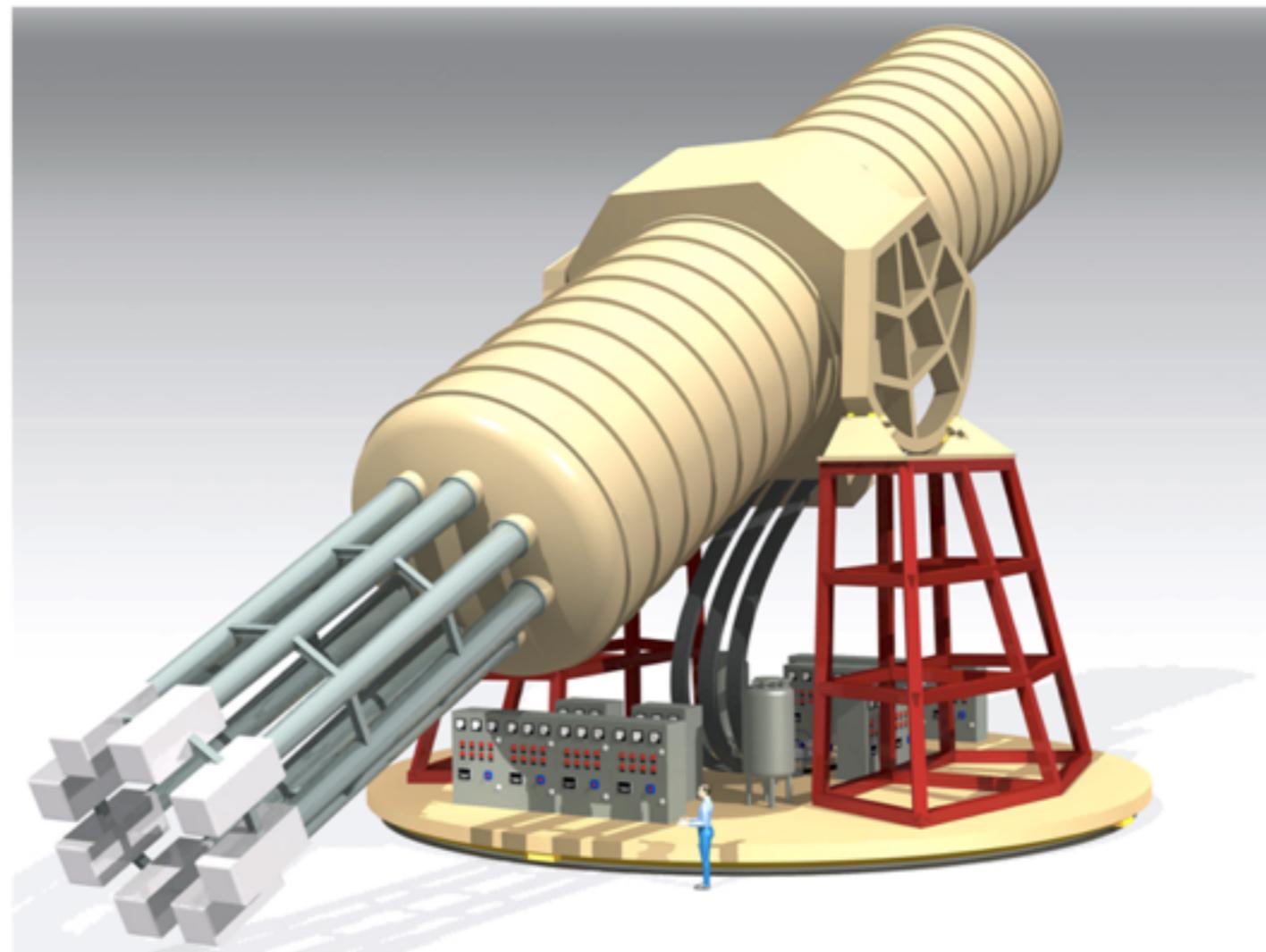
ARIADNE is a search for short-range axion mediated forces between a source mass and an NMR sample.

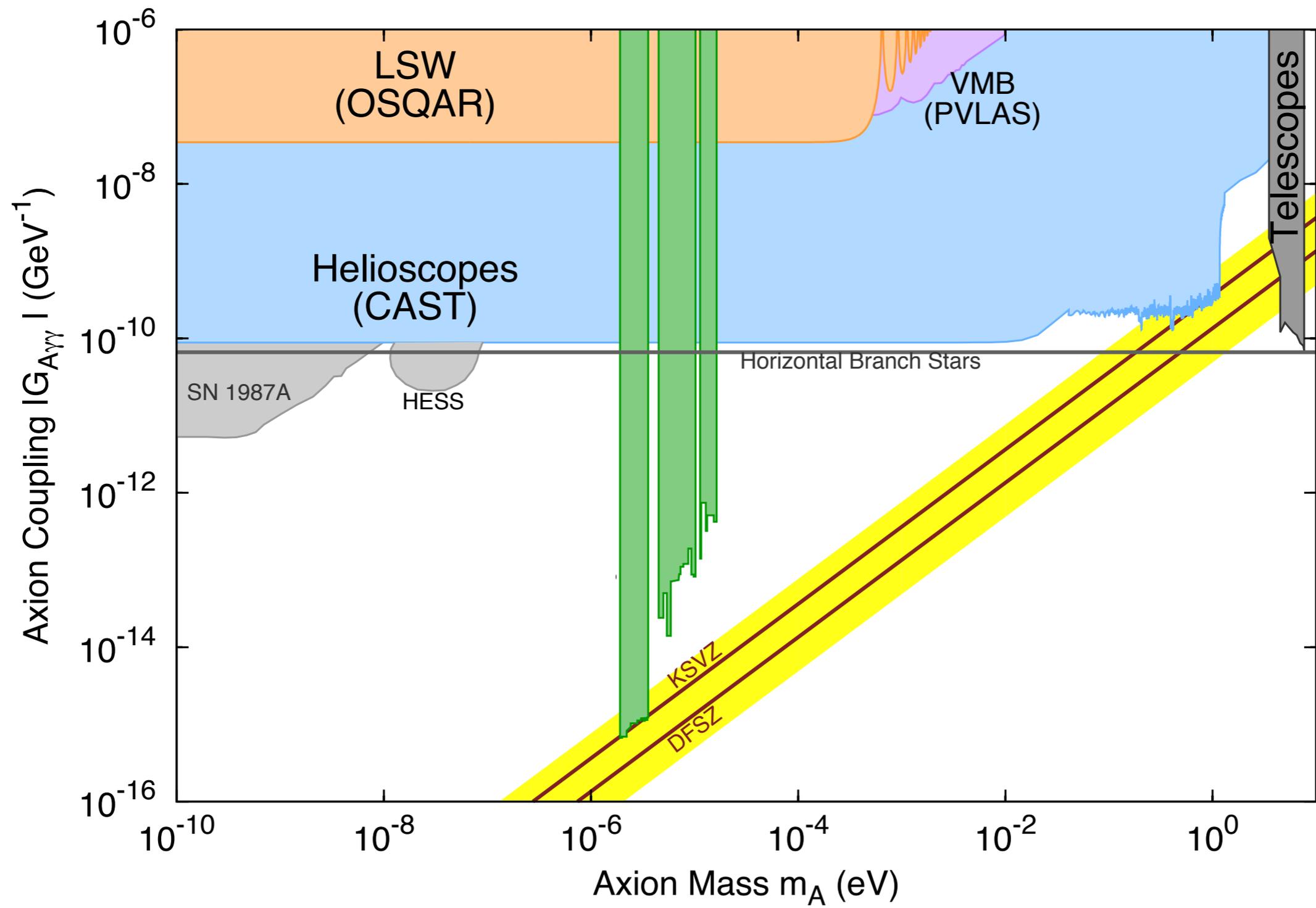


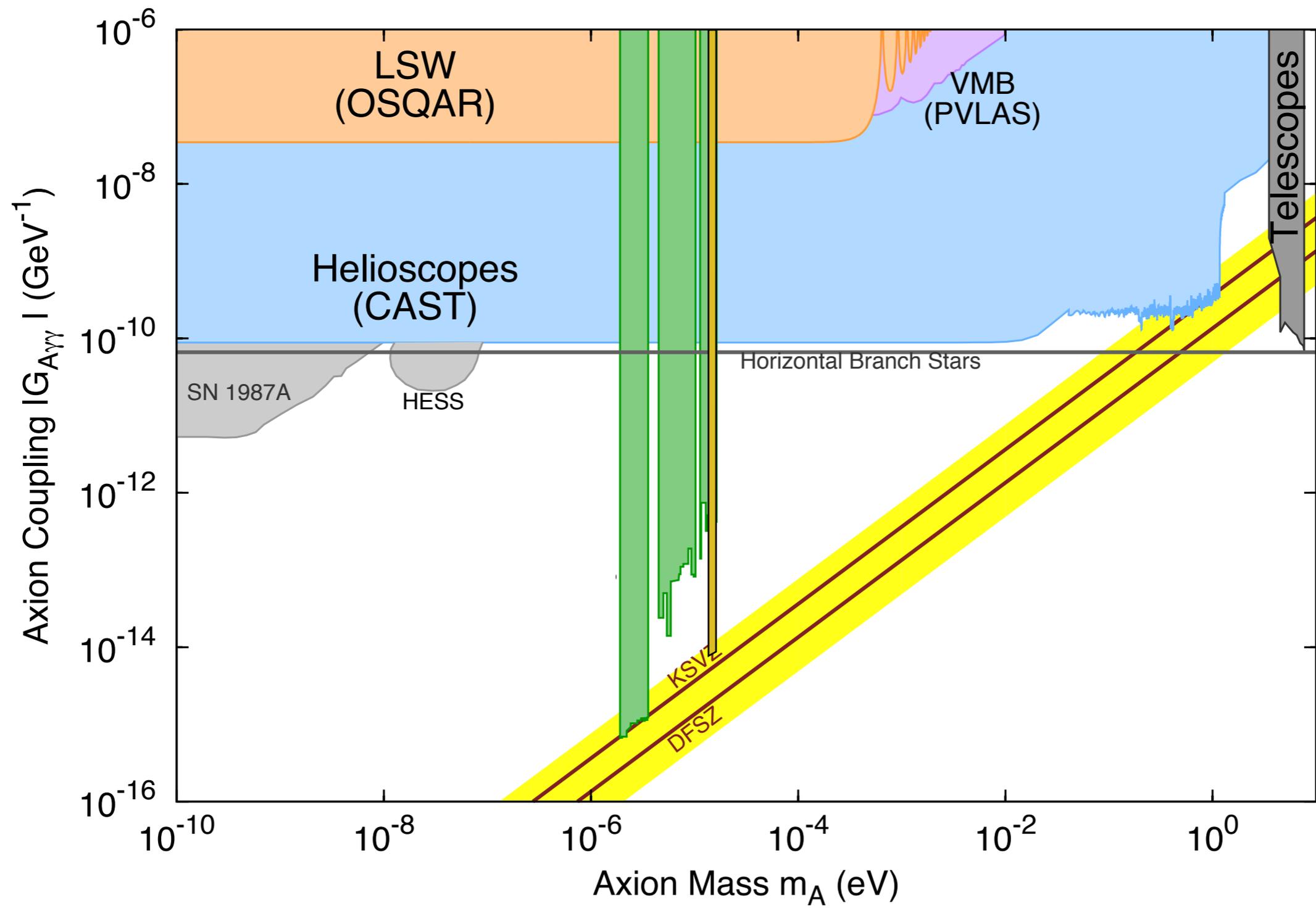
# IAXO

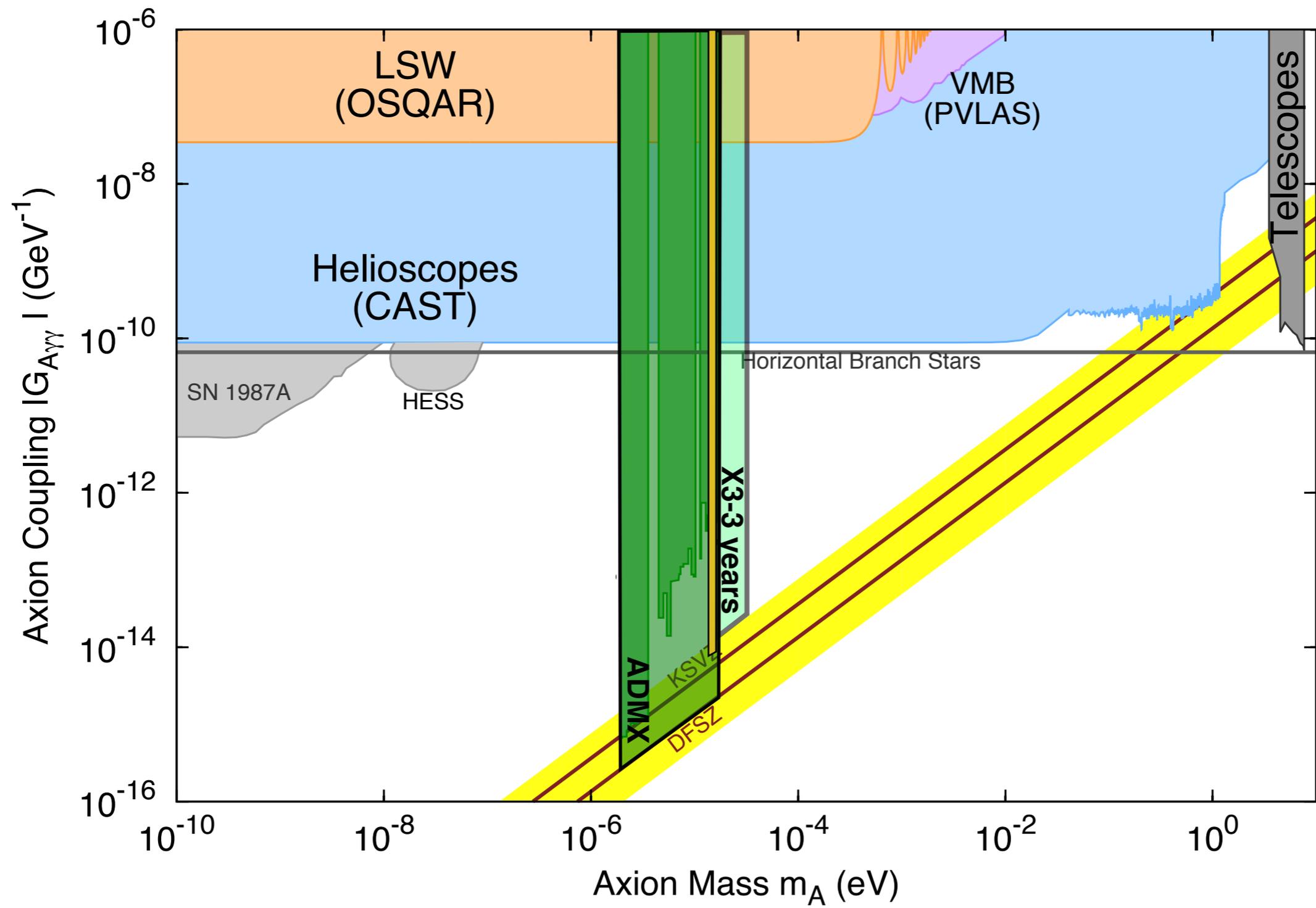
See Talk by Igor Irastorza TODAY (Tuesday), Elisa Ruiz-Choliz on Thursday

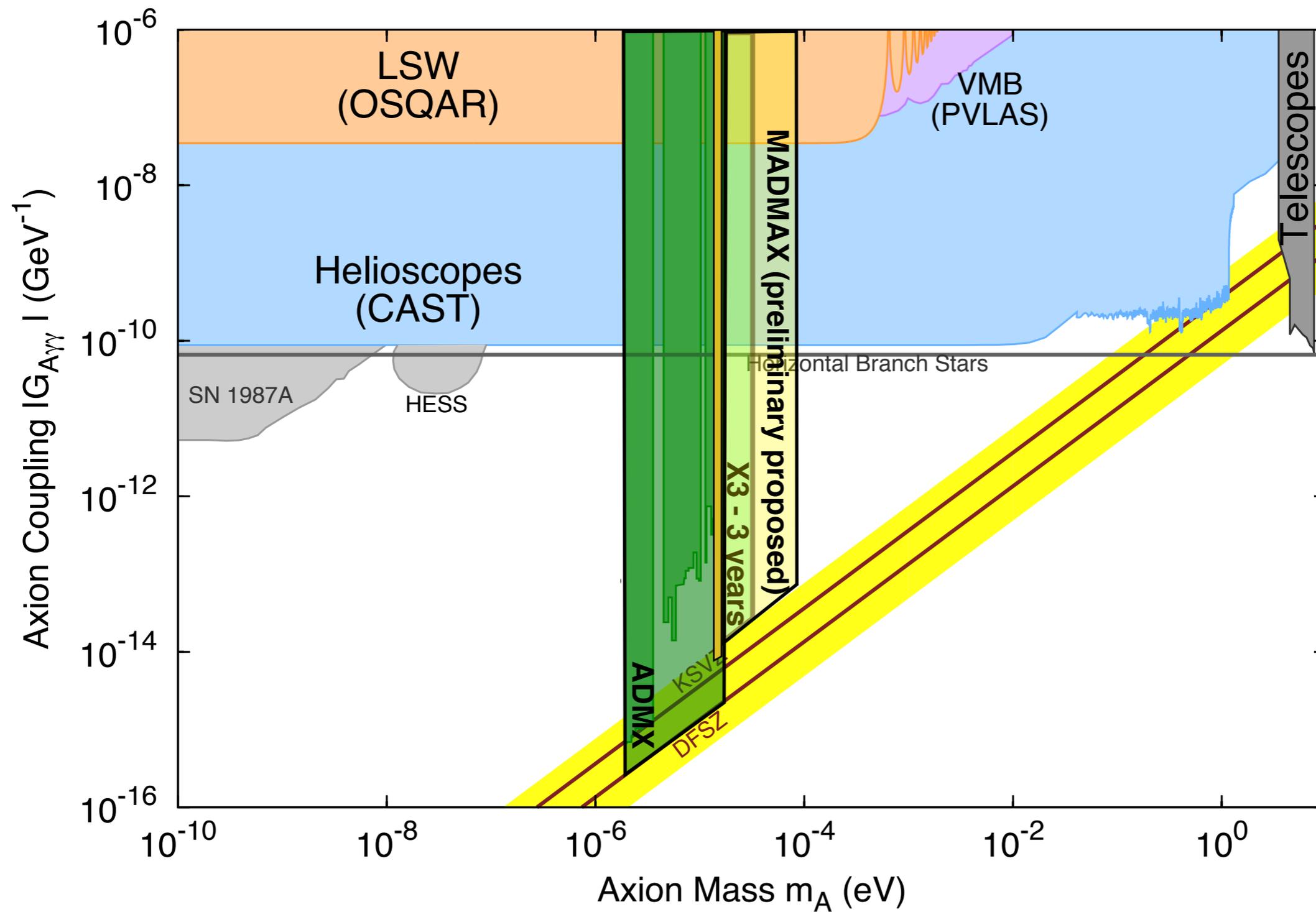
A proposed large scale axion haloscope, with a greatly increased axion conversion volume, new electronics, and a very large high-field magnet. Projected sensitivity to DFSZ/KSVZ axions above  $0.01\text{eV}$ .

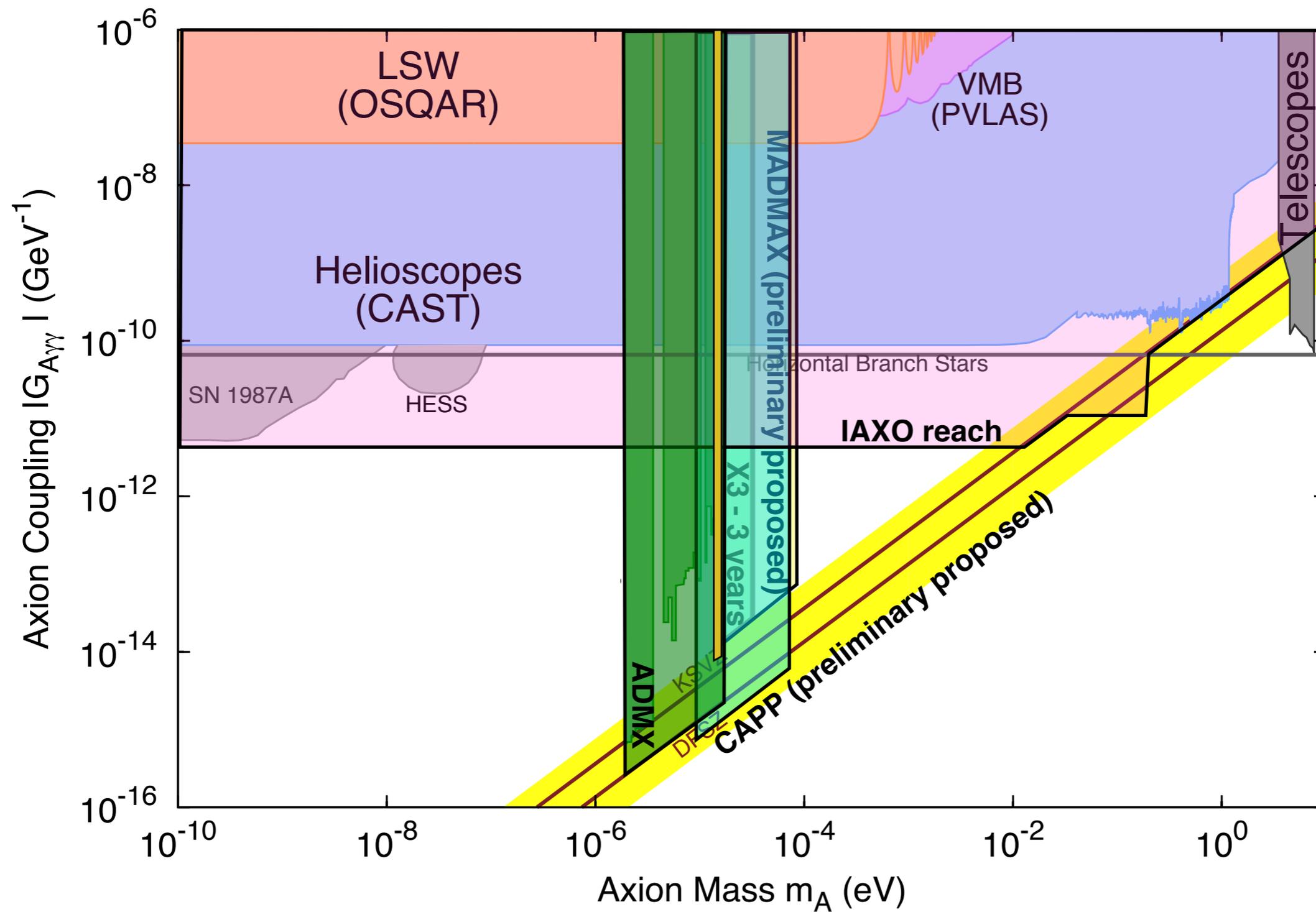












# Conclusions

ADMX2 running imminent.

X3 first results available (preliminary).

Other cavity searches ramping up.

‘Mirror’ searches promising.

IAXO greatly increase reach of haloscopes into axion territory.

NMR / spin precession methods developing

**Real prospects for axion discovery as rate of coverage of mass range ramps up.**