

# The Neutrino “Livingston” Plot

INSS 2017. Fermilab. 2017. Aug. 17.

## **Supergroup 22+23:**

James Ellison

David Friant

Adam Lister

Shivesh Mandalia

Ben Messerly

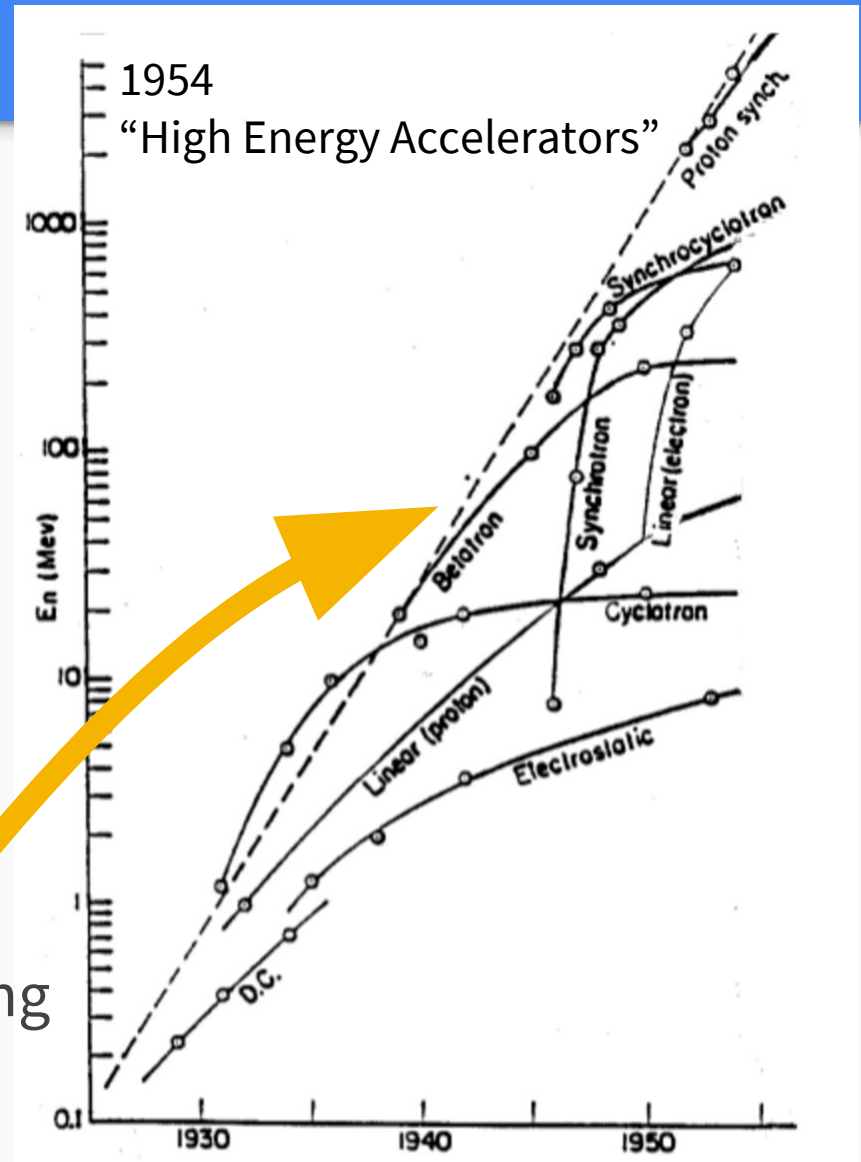
# The Original

M. Stanley Livingston

- FNAL director 1967-1970

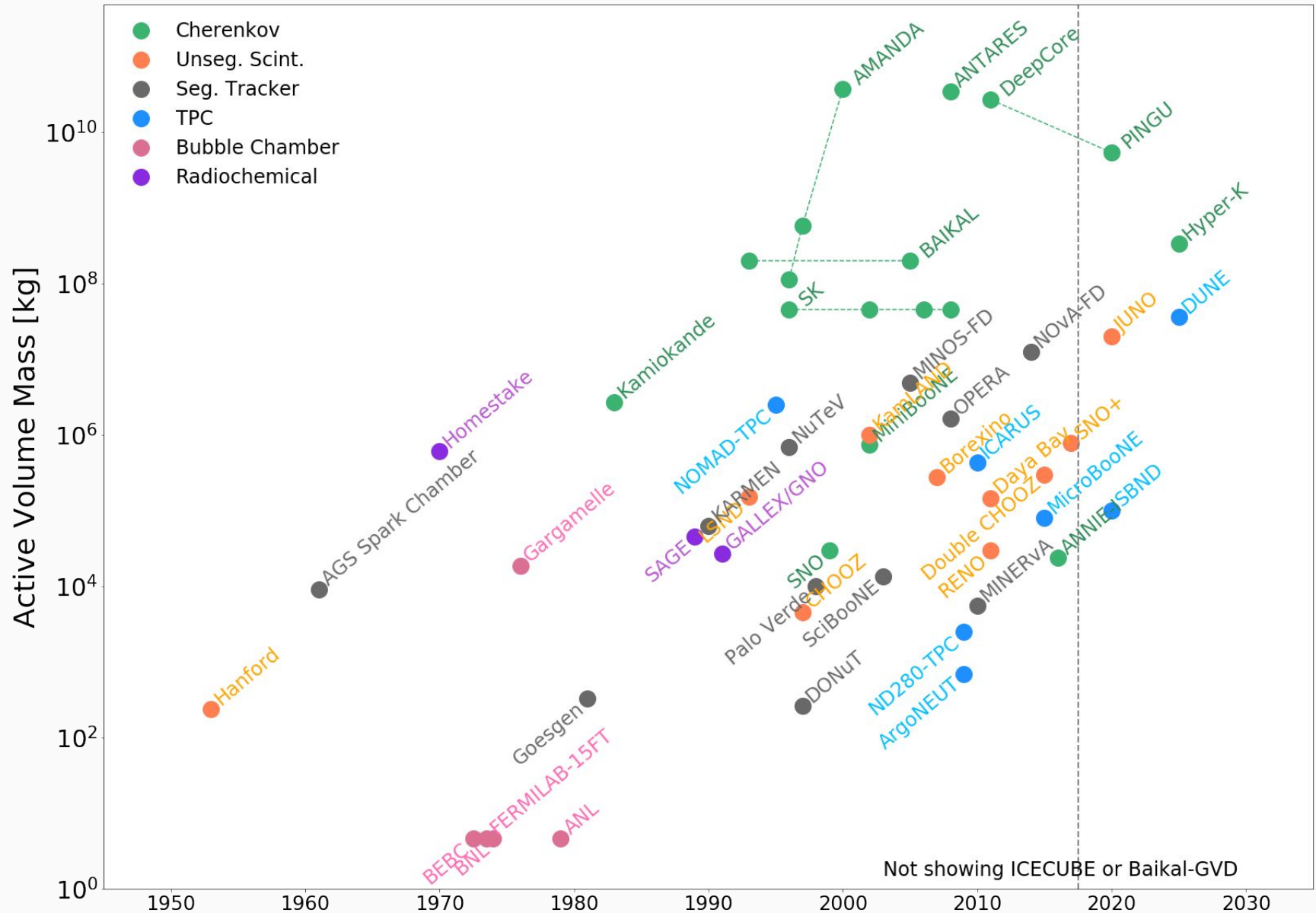


Accelerator CoM energy was doubling every six years!

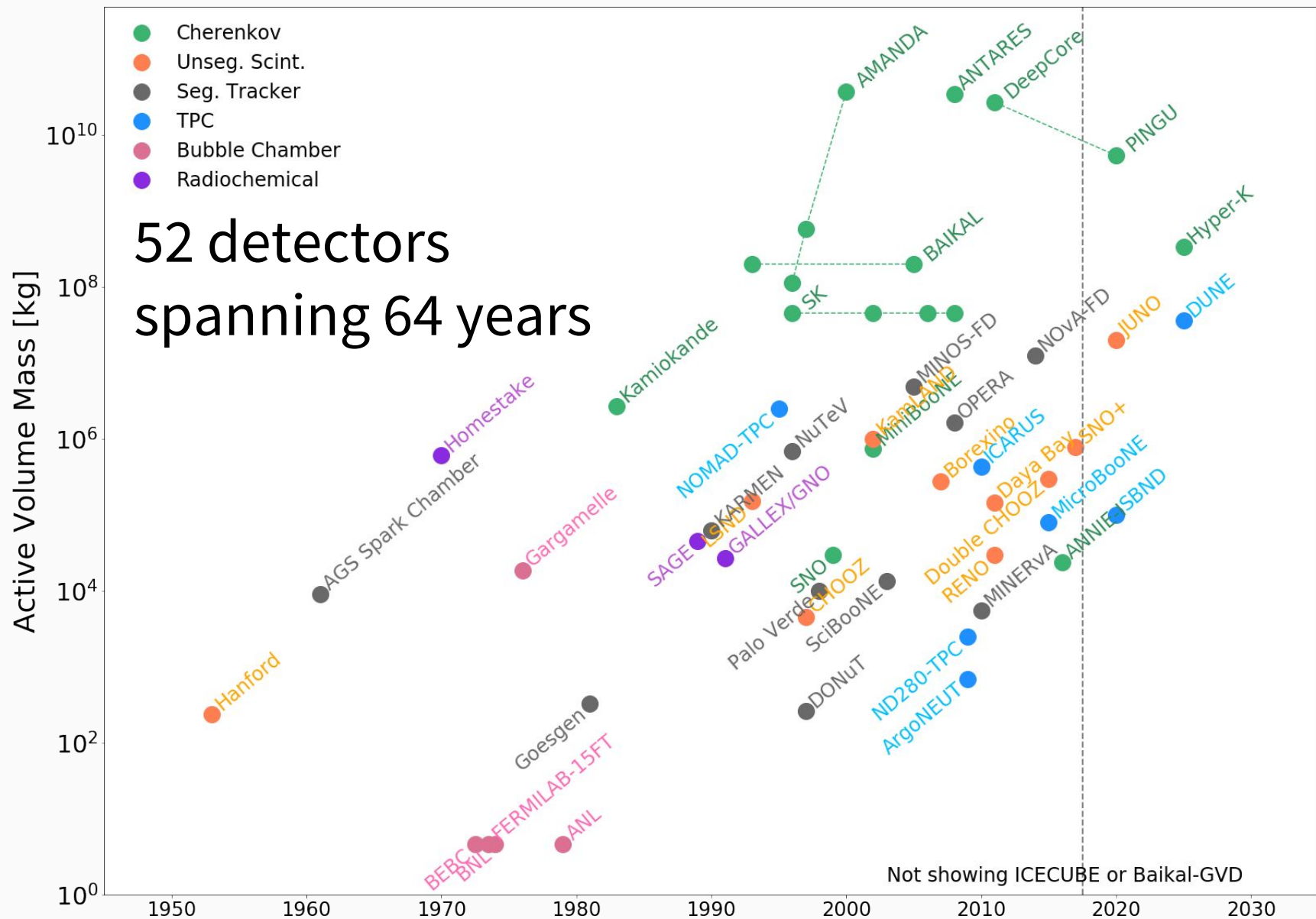


# Neutrino Detector Mass

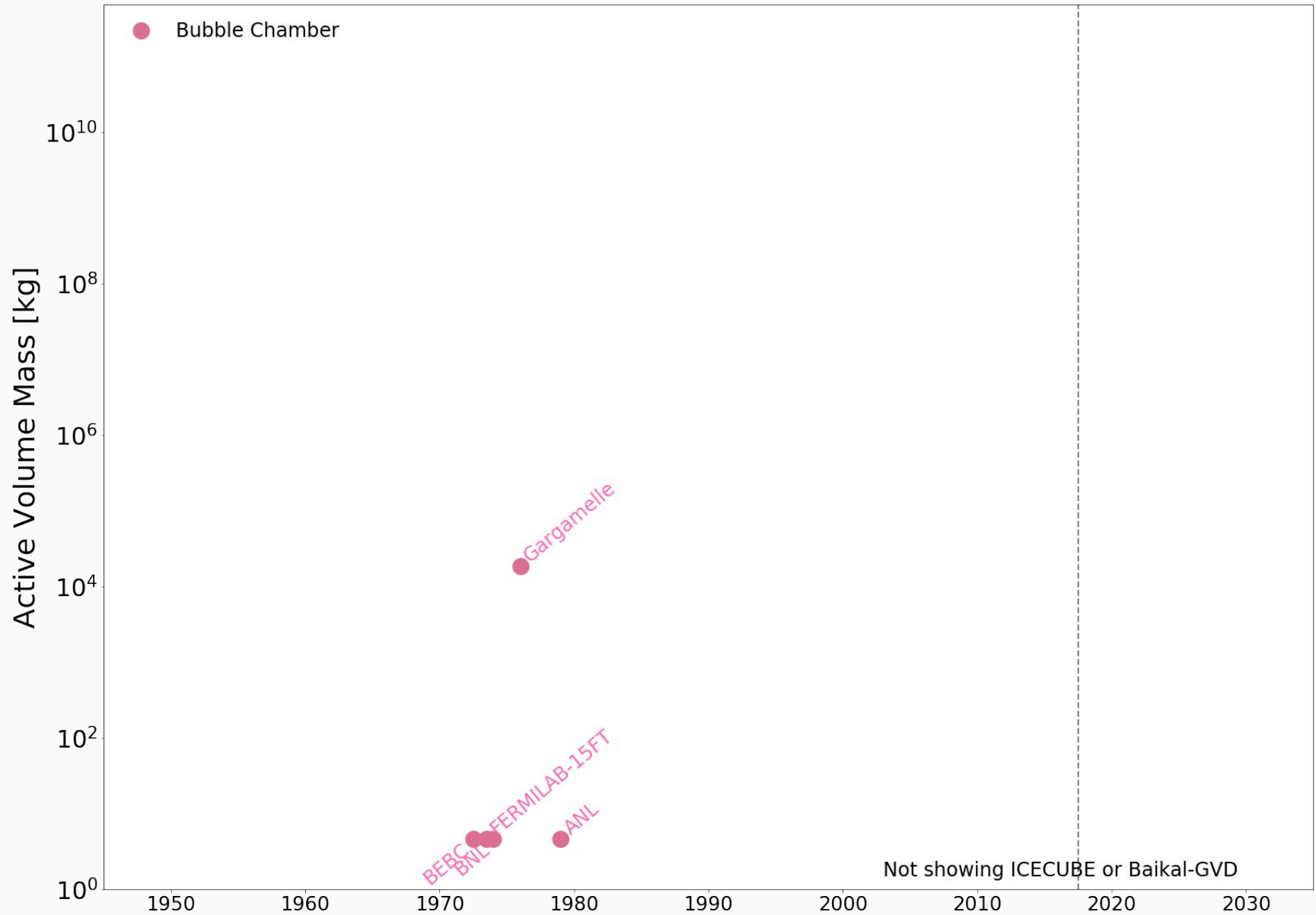
# Neutrino Detector Mass



# Neutrino Detector Mass



# Neutrino Detector Mass – Bubble Chamber



# Neutrino Detector Mass – Radiochemical

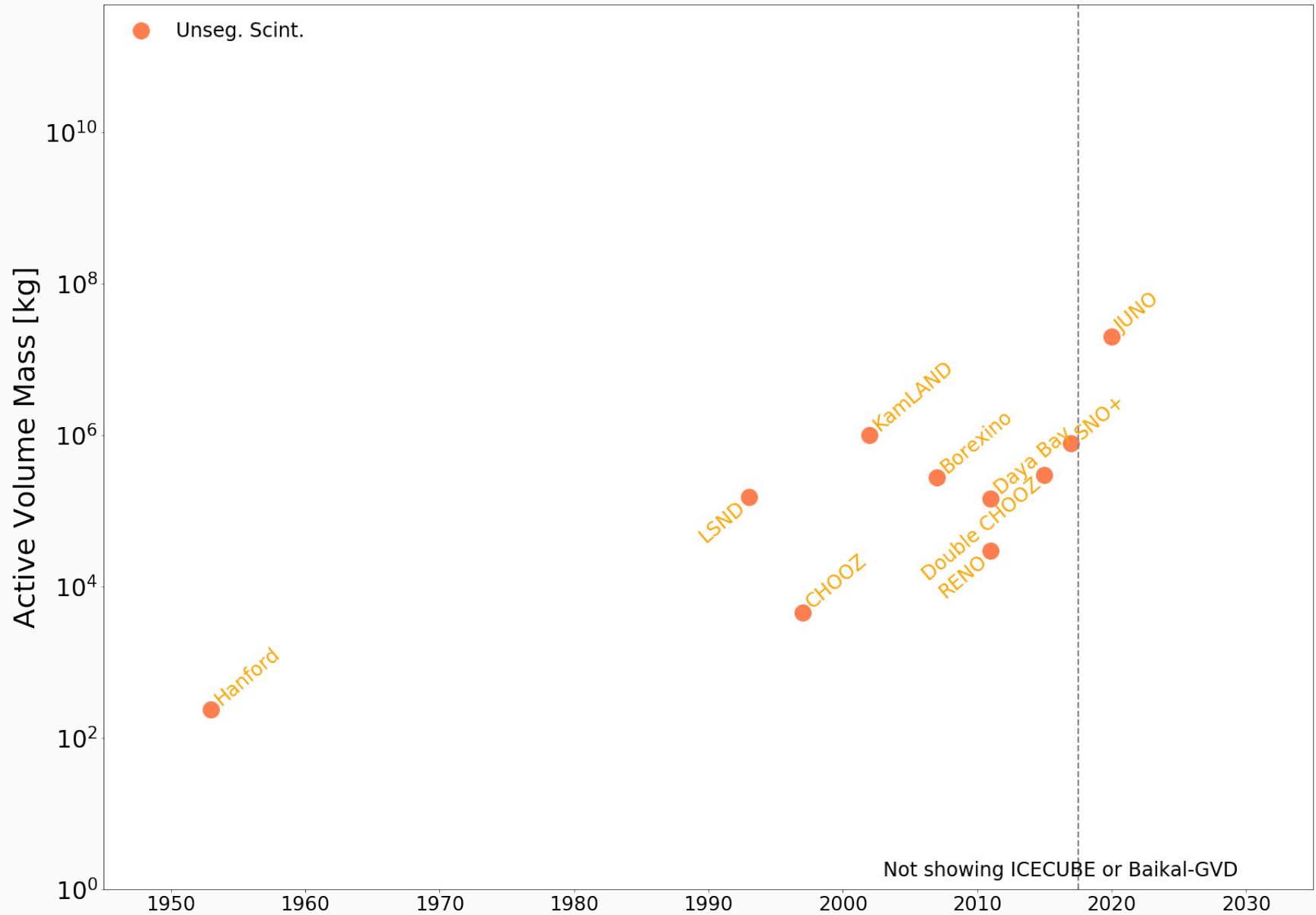


# Neutrino Detector Mass – Segmented Tracker

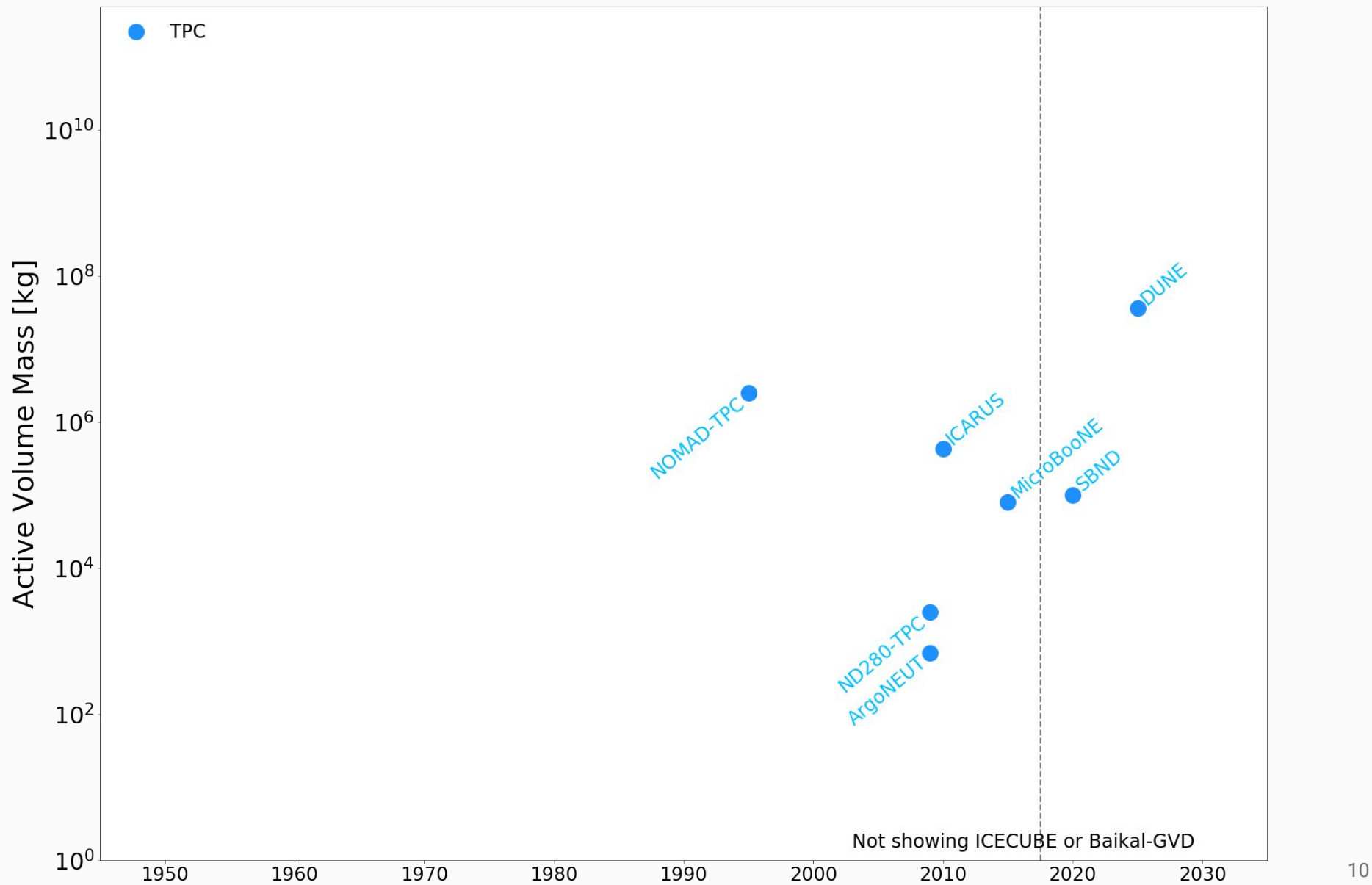




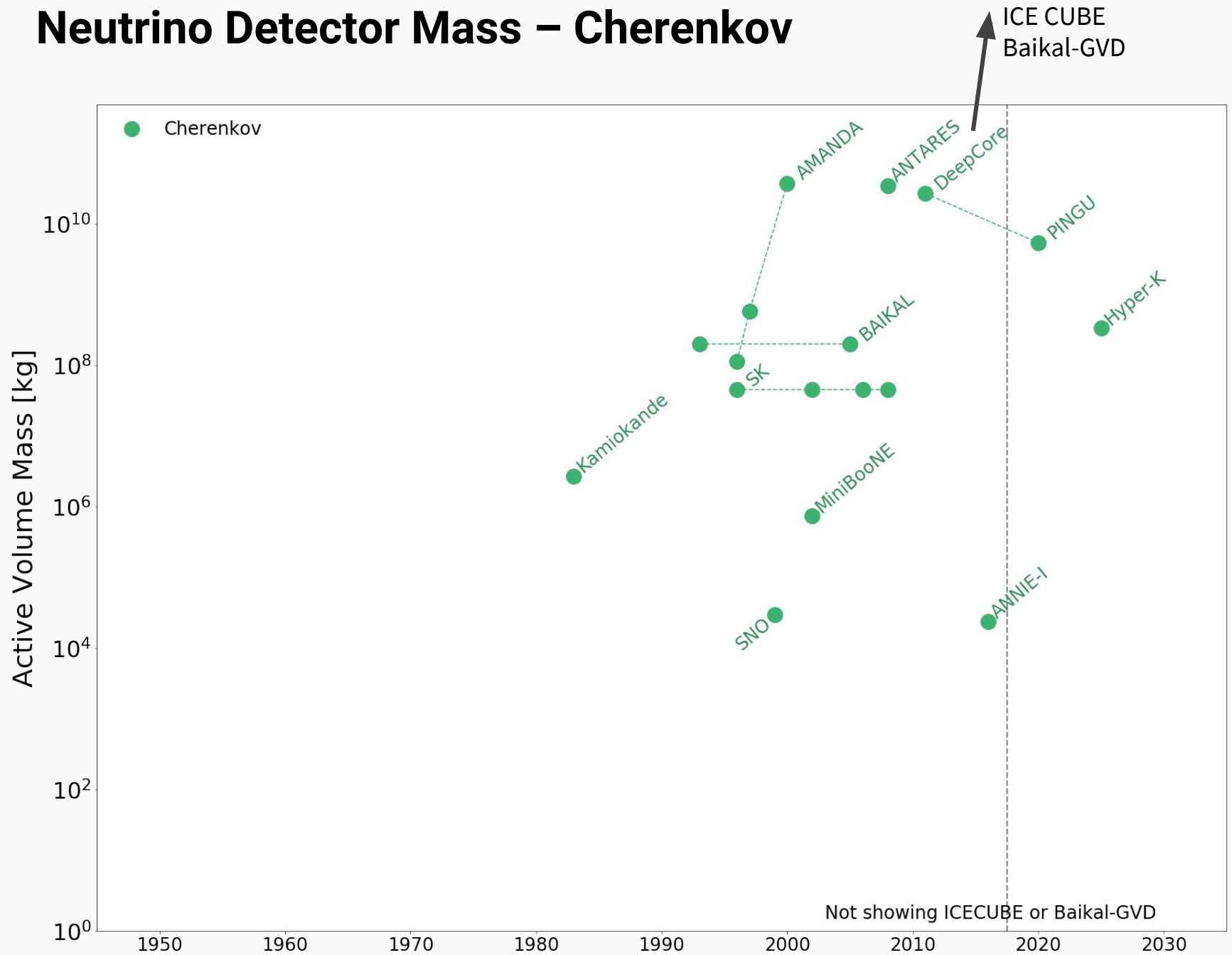
# Neutrino Detector Mass – Unsegmented Scintillator



# Neutrino Detector Mass – TPC

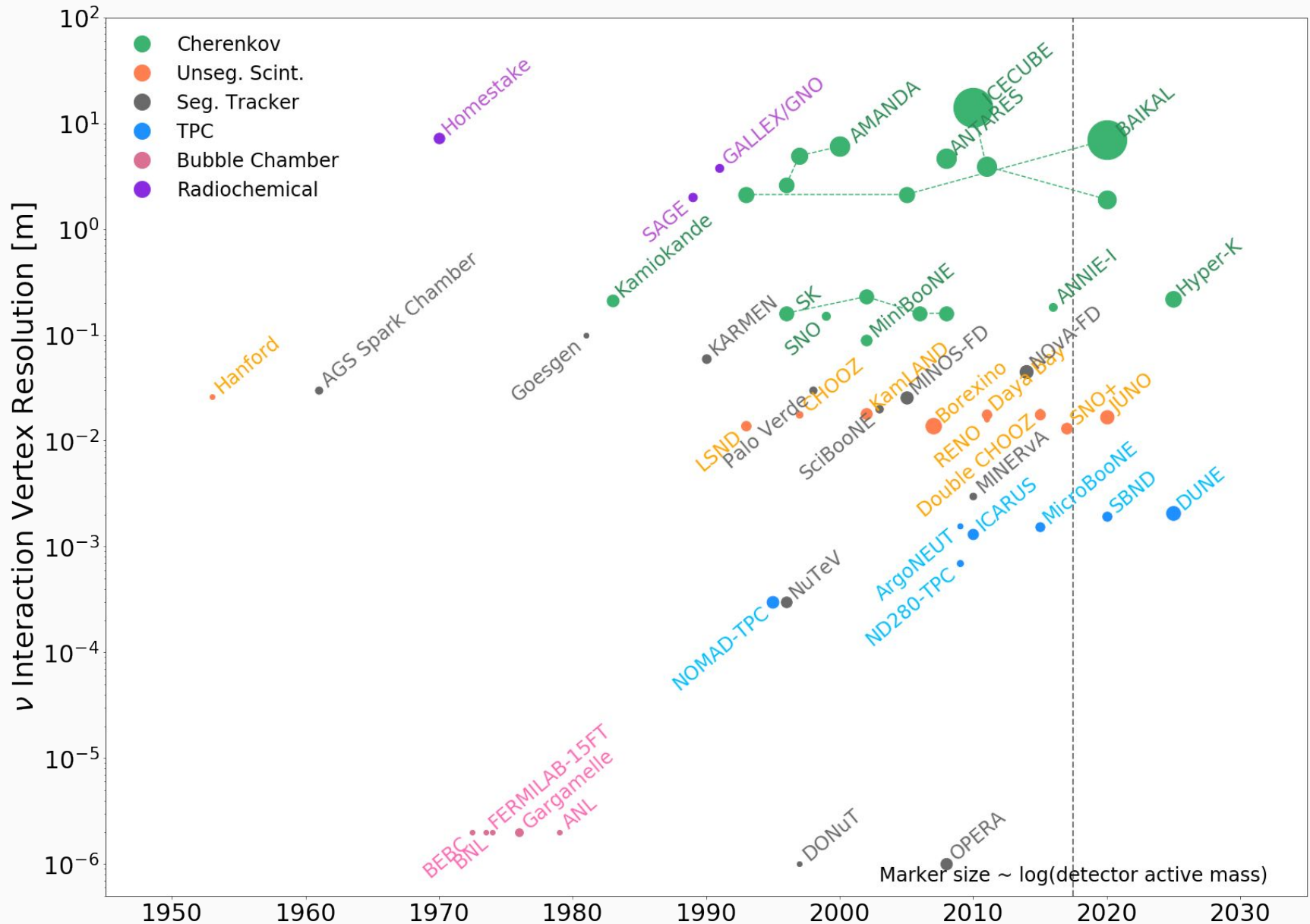


# Neutrino Detector Mass – Cherenkov

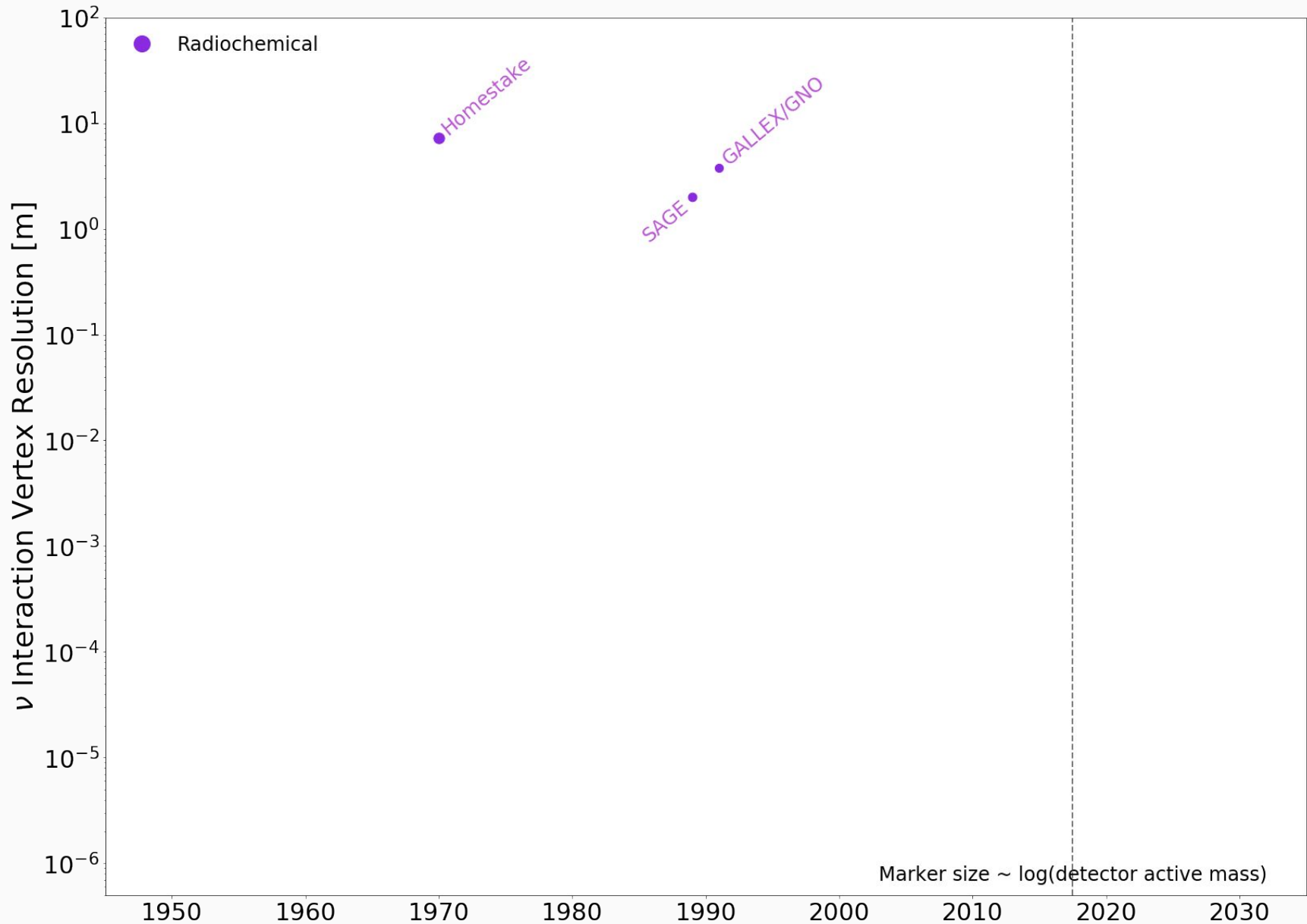


# Neutrino Interaction Vertex Resolution

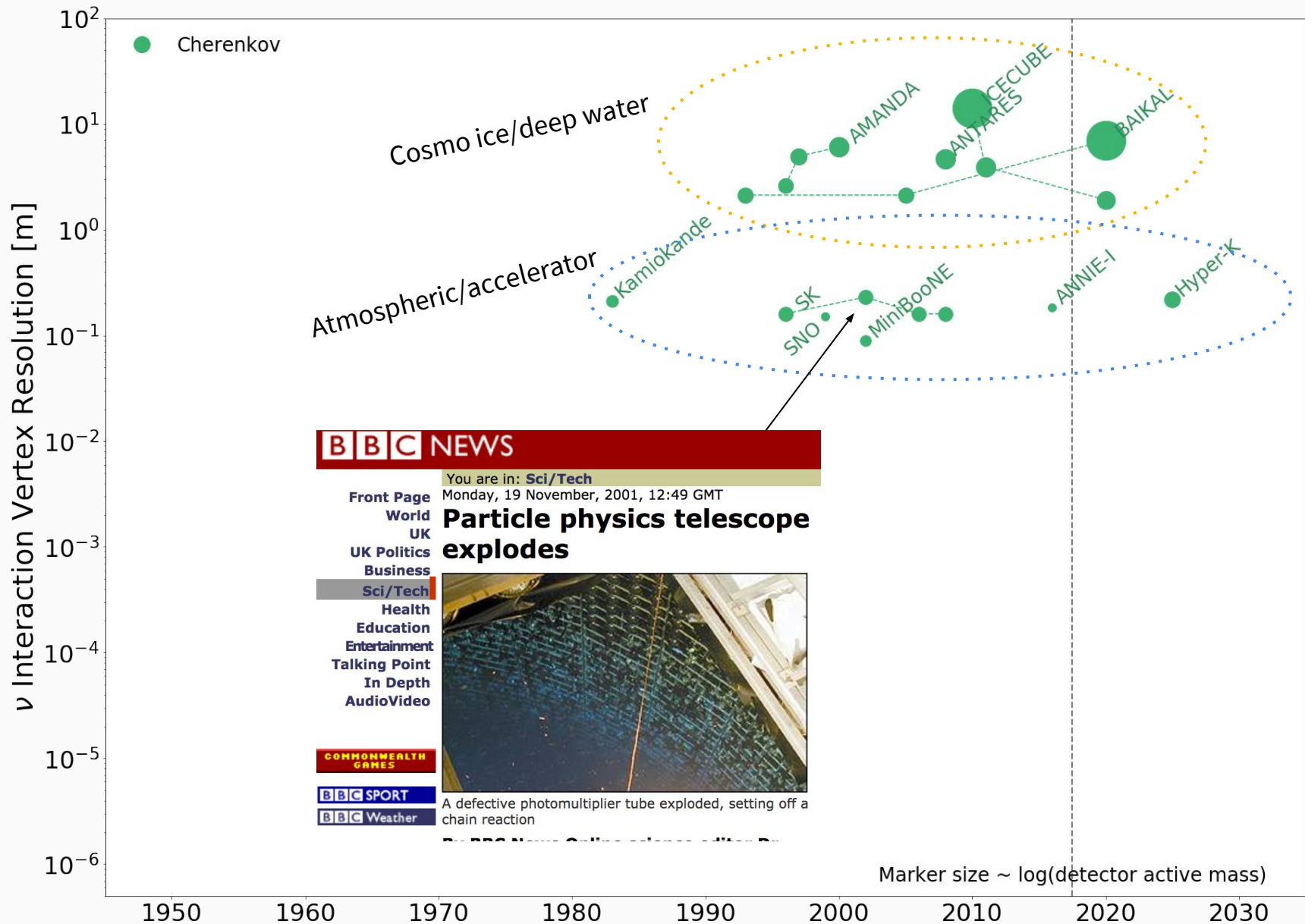
# Neutrino Interaction Vertex Resolution



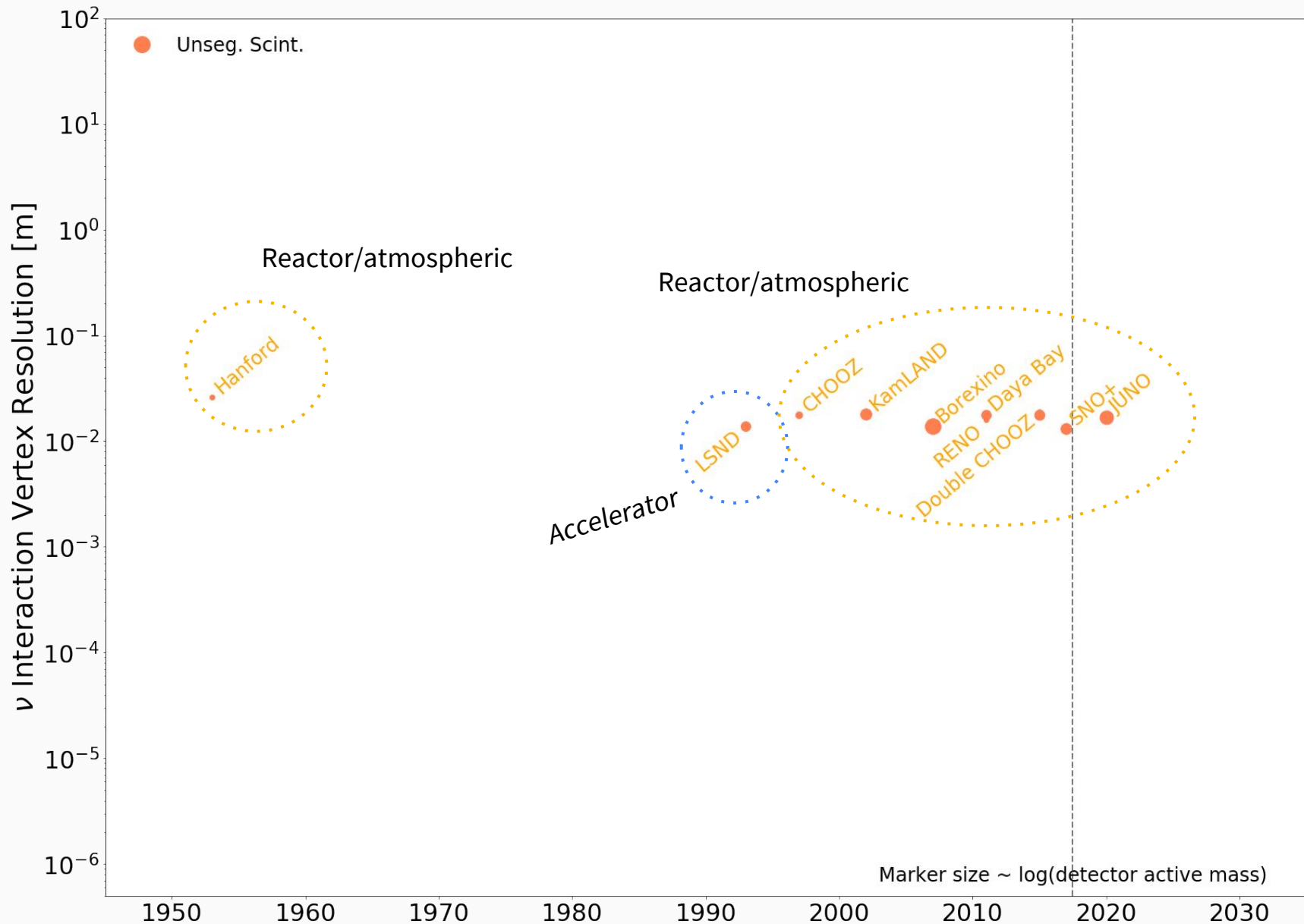
# Neutrino Interaction Vertex Resolution – Radiochemical



# Neutrino Interaction Vertex Resolution – Cherenkov

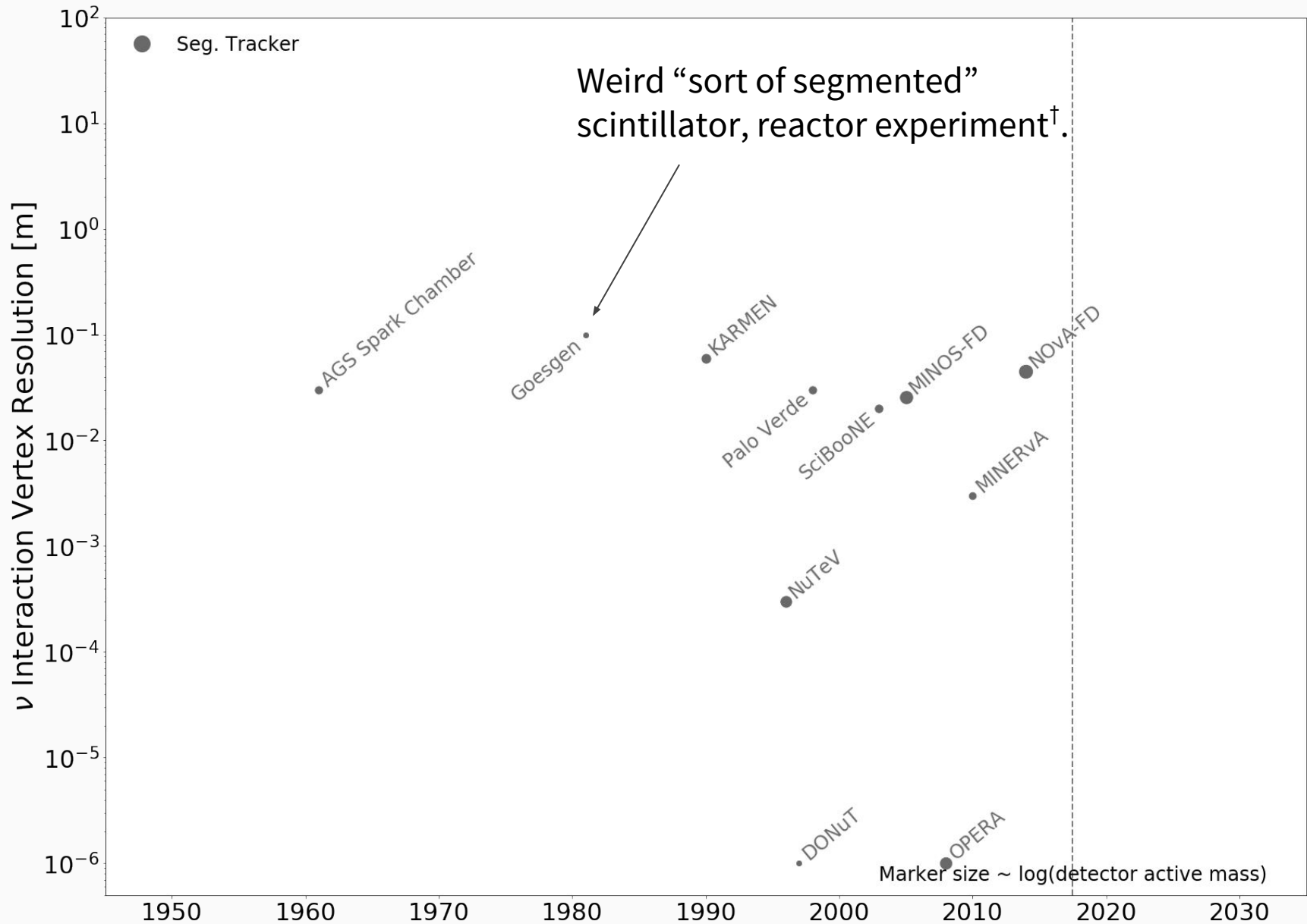


# Neutrino Interaction Vertex Resolution – Unseg Liq Scin



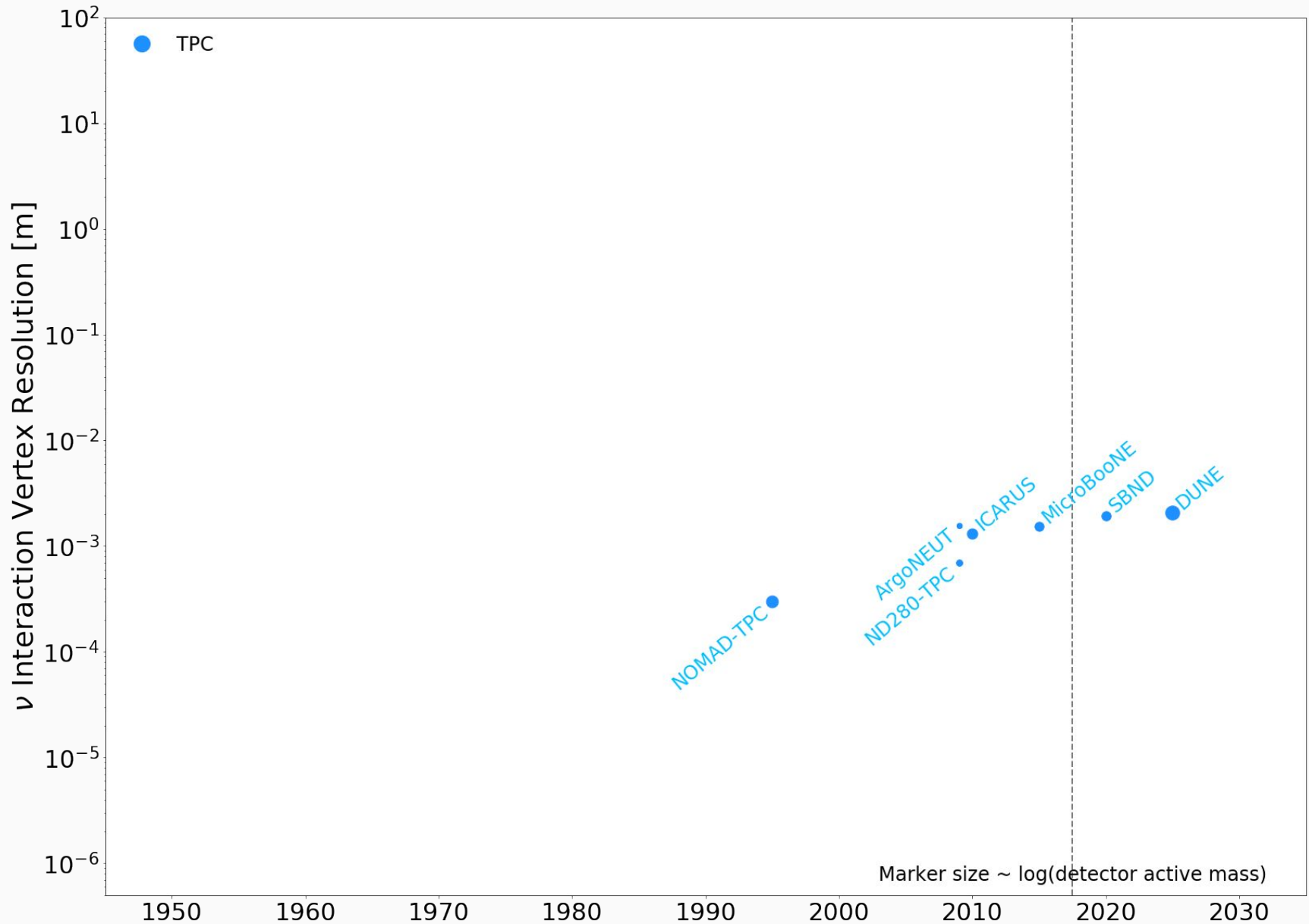


# Neutrino Interaction Vertex Resolution – Seg Tracker

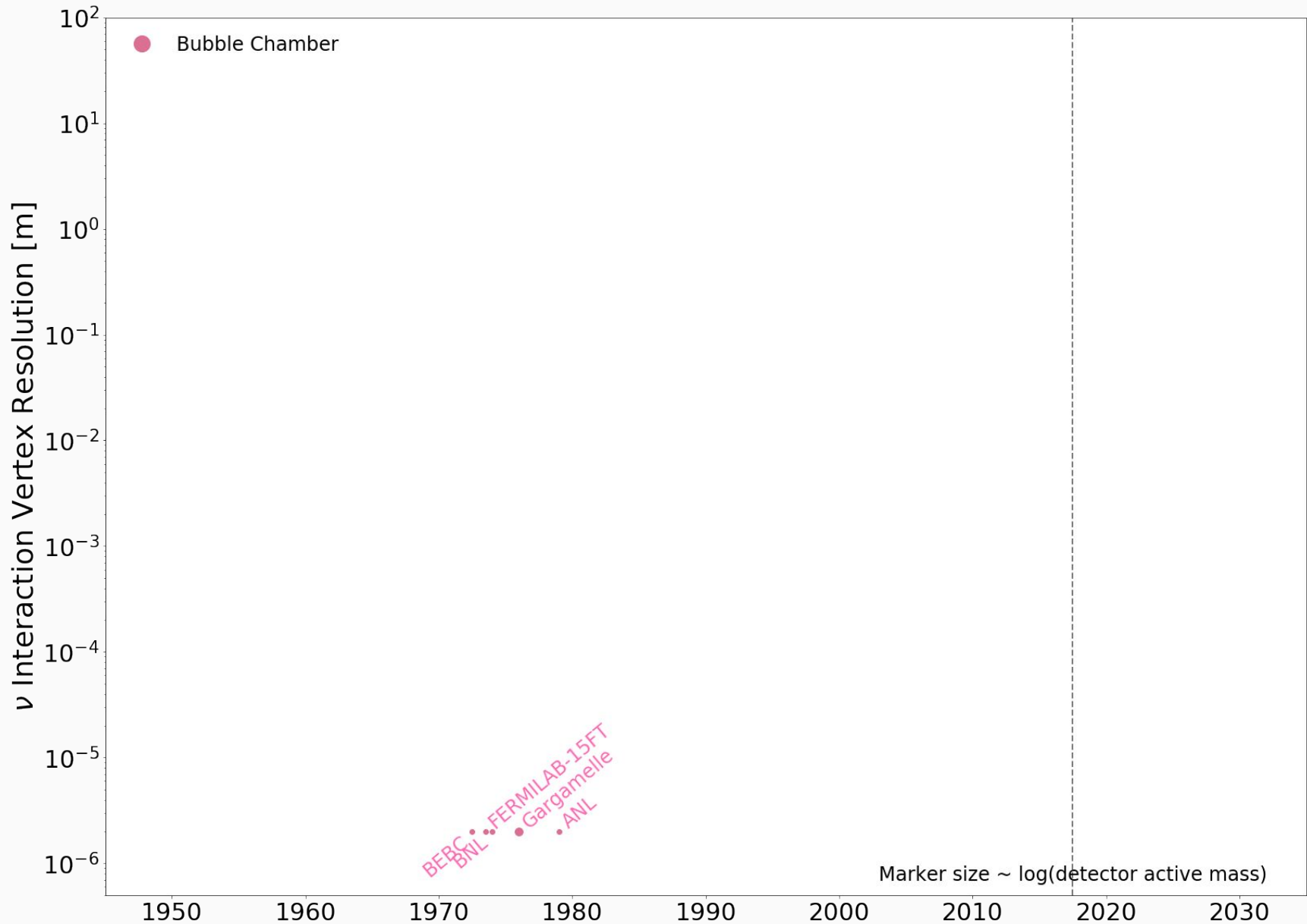


<sup>†</sup> <https://journals.aps.org/prd/pdf/10.1103/PhysRevD.34.2621>

# Neutrino Interaction Vertex Resolution – TPC



# Neutrino Interaction Vertex Resolution – Bubble Chamber



# Conclusion

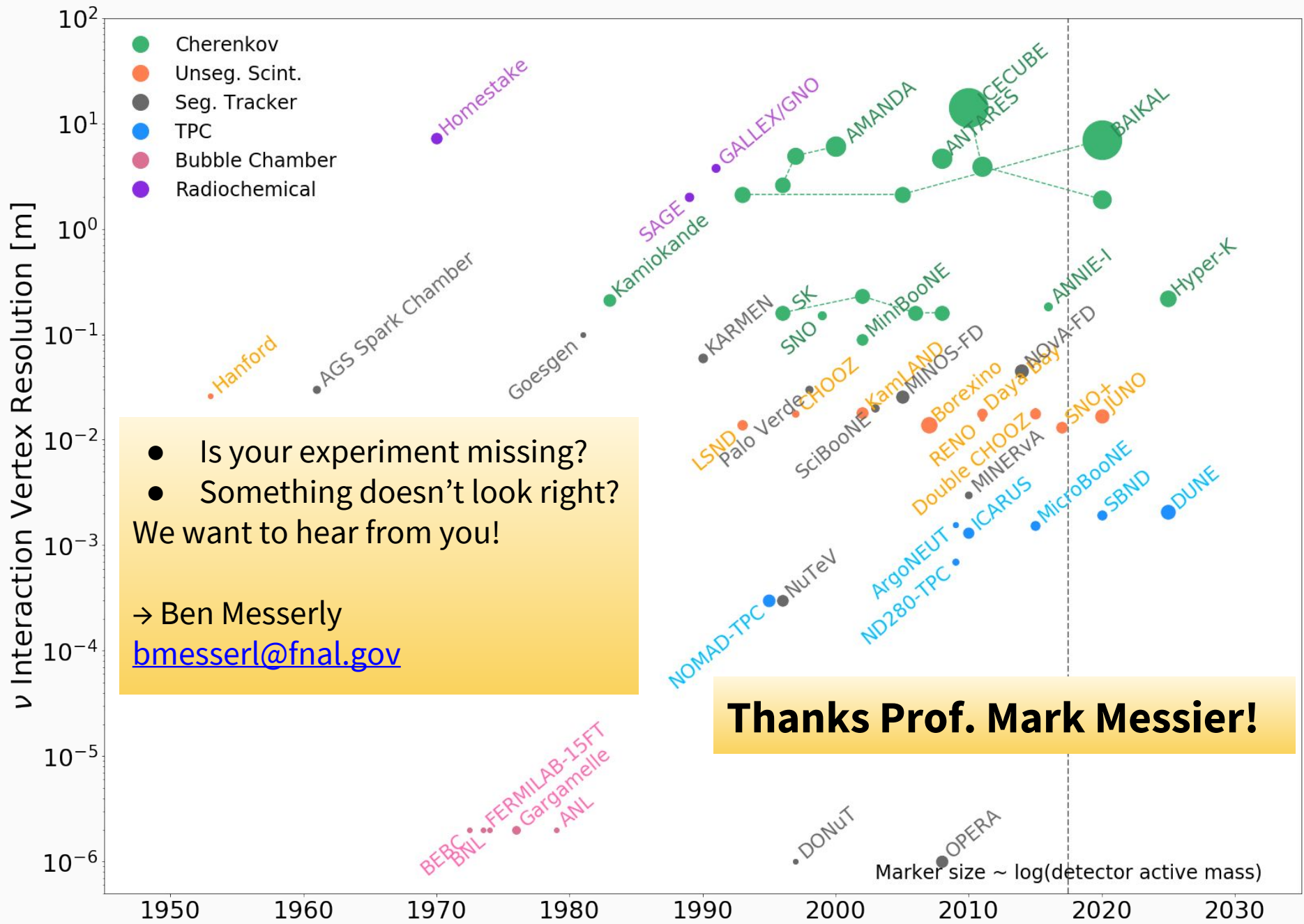
# Our Data, Calculations, and Citations

B	C	D	E	F	G
Start	End	Experiment	Detector Type	Location	mass [kg]
1961	1962	<b>AGS Spark Chamber</b>	Tracking Calorimeter	BNL	9.07E+03
1996	2005	<b>AMANDA</b>	Cherenkov	South Pole	5.00E+09
1969	1979	<b>ANL</b>	Bubble Chamber	Argonne National Laboratory	2.3kg (H), 4.6kg (D)
2016	-	<b>ANNIE</b>	Cherenkov	Fermilab	
2008	-	<b>ANTARES</b>	Cherenkov	Mediterranean Sea	
2009	2010	<b>ArgoNEUT</b>	TPC	Fermilab	6.90E+02
1998	-	<b>Baikal</b>	Cherenkov	Baikal Lake, Russia	2.00E+08
1973	1984	<b>BEBC</b>	Bubble Chamber	CERN	
1974		<b>BNL 7ft</b>	Bubble Chamber	Brookhaven National Laboratory	

<https://docs.google.com/spreadsheets/d/1HKEvm6tsoRX3EGpANM3n6nKIHn5NldHmfSU16YjVcoM/edit?usp=sharing>

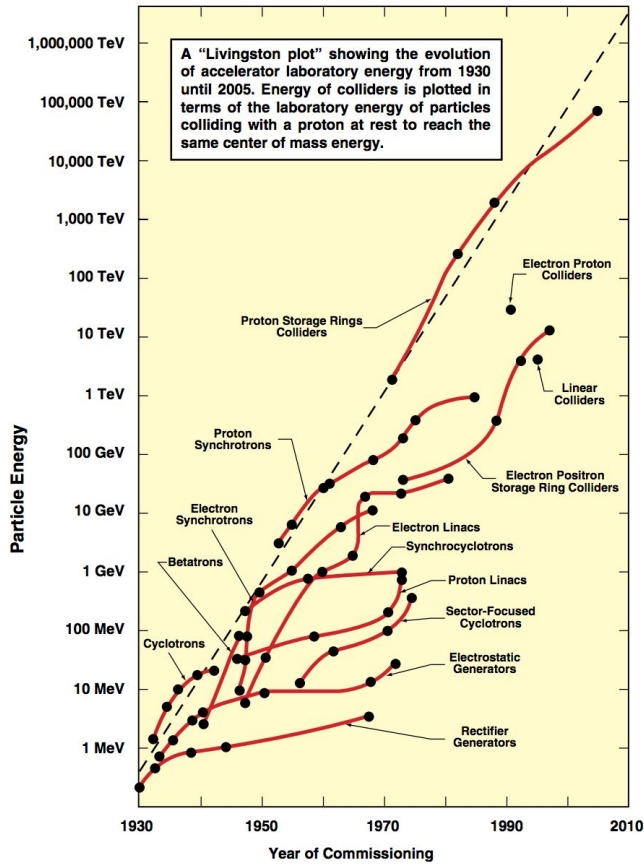
## Notes on the method

- Many caveats + exceptions!
- Paper skimming, ancient pre-internet experiment archeology
- Extensive approximations, back-of-the-envelope calculations
- Where available use quoted resolutions instead of calculated resolutions
  - Calculations agree with quoted values within an order of magnitude; few exceptions.
  - See water cherenkov “*n.b.*” in backup slides.
- *Most* experiments don’t care about and don’t provide enough info to estimate resolution.

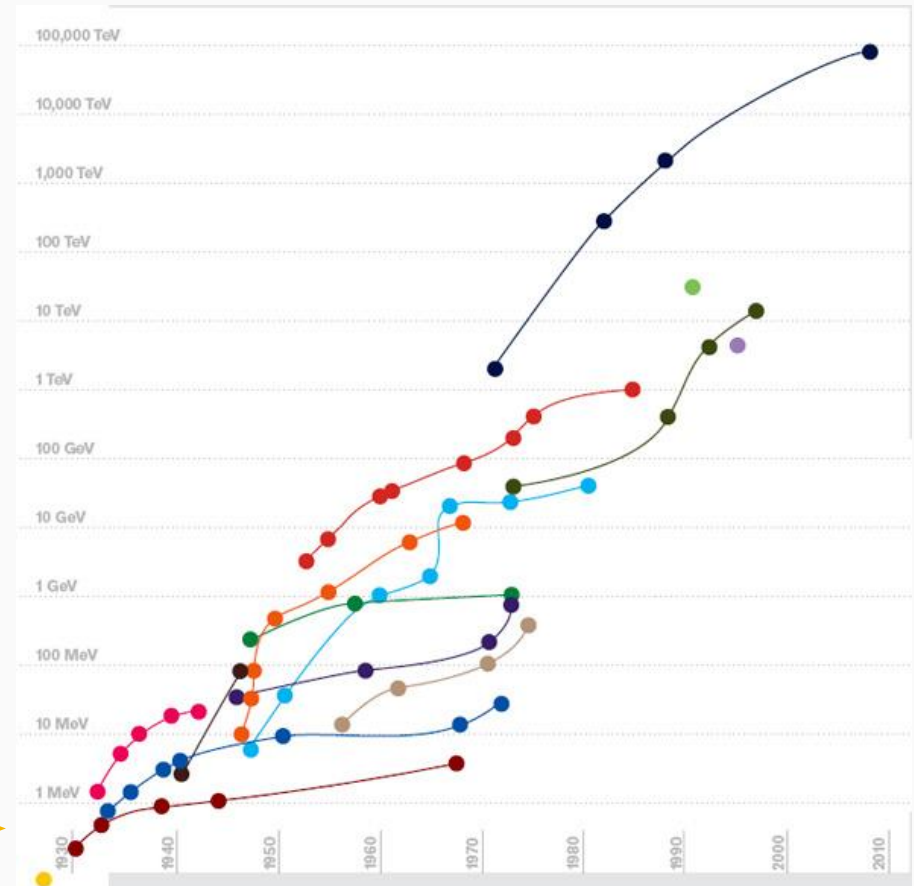


# Backup

# Many Iterations of the Livingston Plot



2001  
Snowmass  
Accelerator  
R&D Report



Symmetry  
Magazine.  
Oct 1, 2009.





# Calculation: Cherenkov and Unsegmented Scintillator

$$\text{Position Resolution} = \frac{\text{Resolution of a single PMT}}{\sqrt{\text{Photons detected}}}$$

$$\text{Resolution of a single PMT} = \text{timing resolution} \times c \times n$$

$$\text{Photons detected} = N_{\text{photons}} \times e^{-\frac{x_0}{\lambda_{Abs}}} \times \text{Quantum Eff} \times \text{PMT coverage}$$

$$N_{\text{photons}} = \frac{\text{photons produced}}{\text{track length}} \times \text{track length} \times \sin^2(\theta_c) \quad (\text{Cherenkov})$$

# Calculation: Cherenkov and Unsegmented Scintillator

$$\text{Position Resolution} = \frac{\text{Resolution of a single PMT}}{\sqrt{\text{Photons detected}}}$$

$$\text{Resolution of a single PMT} = \text{timing resolution} \times c \times n$$

$$\text{Photons detected} = N_{\text{photons}} \times e^{-\frac{x_0}{\lambda_{\text{Abs}}}} \times \text{Quantum Eff} \times \text{PMT coverage}$$

$$N_{\text{photons}} = \frac{\text{photons produced}}{\text{track length}} \times \text{track length} \times \sin^2(\theta_c) \quad (\text{Cherenkov})$$

- $x_0 \sim$  detector scale
- $\lambda_{\text{Abs}} \sim 10 \text{ m}$
- *Quantum Efficiency* = 20%
- PMT coverage  $\sim 10\text{-}40\%$
- Coverage =  $\Sigma (\text{SA}_{\text{PMT}}) / \text{SA}_{\text{det}}$ . For Ice Cube-style detectors  $\sim \Sigma_{\text{plane}} (\text{SA}_{\text{PMT}}) / \text{SA}_{\text{det}}$

# Calculation: Cherenkov and Unsegmented Scintillator

$$\text{Position Resolution} = \frac{\text{Resolution of a single PMT}}{\sqrt{\text{Photons detected}}}$$

$$\text{Resolution of a single PMT} = \text{timing resolution} \times c \times n$$

$$\text{Photons detected} = N_{\text{photons}} \times e^{-\frac{x_0}{\lambda_{Abs}}} \times \text{Quantum Eff} \times \text{PMT coverage}$$

$$N_{\text{photons}} = \frac{\text{photons produced}}{\text{track length}} \times \text{track length} \times \sin^2(\theta_c) \quad (\text{Cherenkov})$$

*N.B.* For all detectors in these two categories, we used a **10 MeV electron** for photon source.

Not what reactor or solar experiments ( $E_\nu \sim 1\text{MeV}$ ) had in mind when they were determining their resolutions!

Generally: higher energy track  $\rightarrow$  more photons  $\rightarrow$  better resolution.

# Calculation: Cherenkov and Unsegmented Scintillator

$$\text{Position Resolution} = \frac{\text{Resolution of a single PMT}}{\sqrt{\text{Photons detected}}}$$

$$\text{Resolution of a single PMT} = \text{timing resolution} \times c \times n$$

$$\text{Photons detected} = N_{\text{photons}} \times e^{-\frac{x_0}{\lambda_{\text{Abs}}}} \times \text{Quantum Eff} \times \text{PMT coverage}$$

$$N_{\text{photons}} = \frac{\text{photons produced}}{\text{track length}} \times \text{track length} \times \sin^2(\theta_c) \quad (\text{Cherenkov})$$

E.g. we found that using a 10 MeV electron for a reactor experiment ( $E_v \sim 5$  MeV) lead to a calculated resolution ~ an order of magnitude better than the quoted resolution.

Similarly, 10MeV too small for cosmo experiments.

For consistency, used 10 MeV electron, though next iteration might adjust.

# Calculation: Cherenkov and Unsegmented Scintillator

$$\text{Position Resolution} = \frac{\text{Resolution of a single PMT}}{\sqrt{\text{Photons detected}}}$$

$$\text{Resolution of a single PMT} = \text{timing resolution} \times c \times n$$

$$\text{Photons detected} = N_{\text{photons}} \times e^{-\frac{x_0}{\lambda_{Abs}}} \times \text{Quantum Eff} \times \text{PMT coverage}$$

$$N_{\text{photons}} = \frac{\text{photons produced}}{\text{particle energy}} \times \text{particle energy} \quad (\text{Scintillator})$$

## New Experiments

Resolution ~

wire spacing (x&y) \* drift distance / (drift time \* sampling rate)

# Segmented Tracking

## Category Includes:

- Spark Chamber (1962  $\nu_{\mu}$  discovery)
- $\nu_{\tau}$  Emulsion (OPERA, DONuT)
- Standard Scintillator Trackers (NuTeV, MINERvA, MINOS, NOvA)

## Resolution:

- Largely determined by dimensions/spacings of segments
- Though experiments had their own fancy methods which could improve resolution.

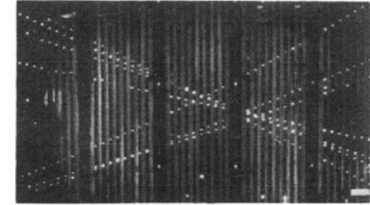


FIG. 4. Land print of Cosmic-ray muons integrated over many incoming tracks.

