

Theory of Chemical Tagging

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The Carnegie Observatories

Galactic Archaeology

- Understanding the chemical/dynamical evolution of galaxies
 - When, where and how stars formed [initial point]
 - What are their elemental abundances [chemical evolution]
 - What are their orbits now [dynamical evolution]

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Gaia : 10^9 stars

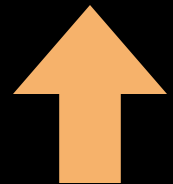
Galactic Archaeology

- Understanding the chemical/dynamical evolution of galaxies

● When, where and how stars formed [initial point]

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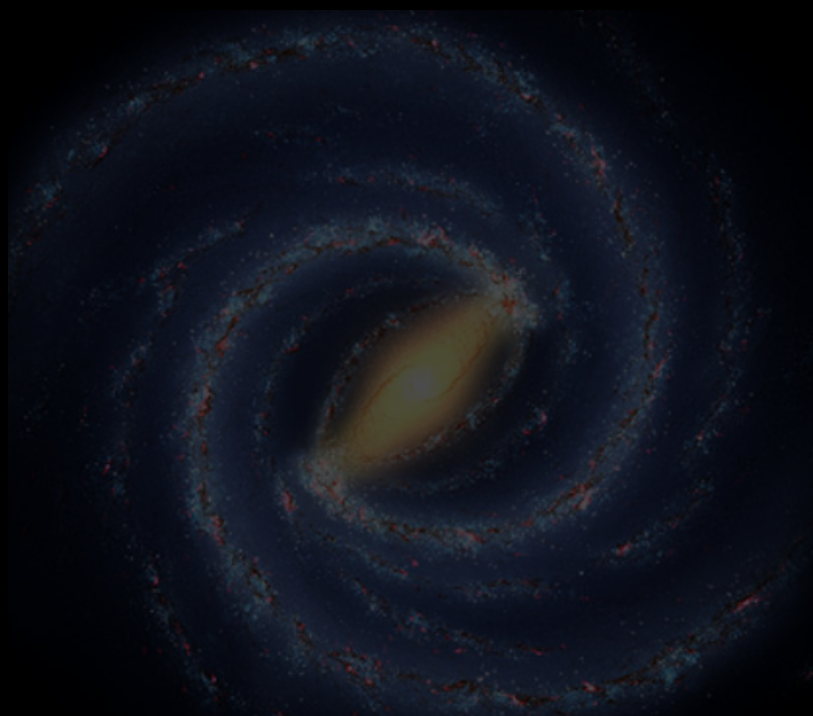
GALAH, Gaia-ESO,
APOGEE: 10^6 stars

Galactic Archaeology

- Understanding the chemical/dynamical evolution of galaxies

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Star forming associations quickly disperse after their formation (~ 100 Myrs)



Assuming $\sigma_{\text{stars}} = \sigma_{\text{ISM}} = 10$ km/s

Kinematic information is not sufficient to reconstruct the birth origins of stars

After 100 Myrs

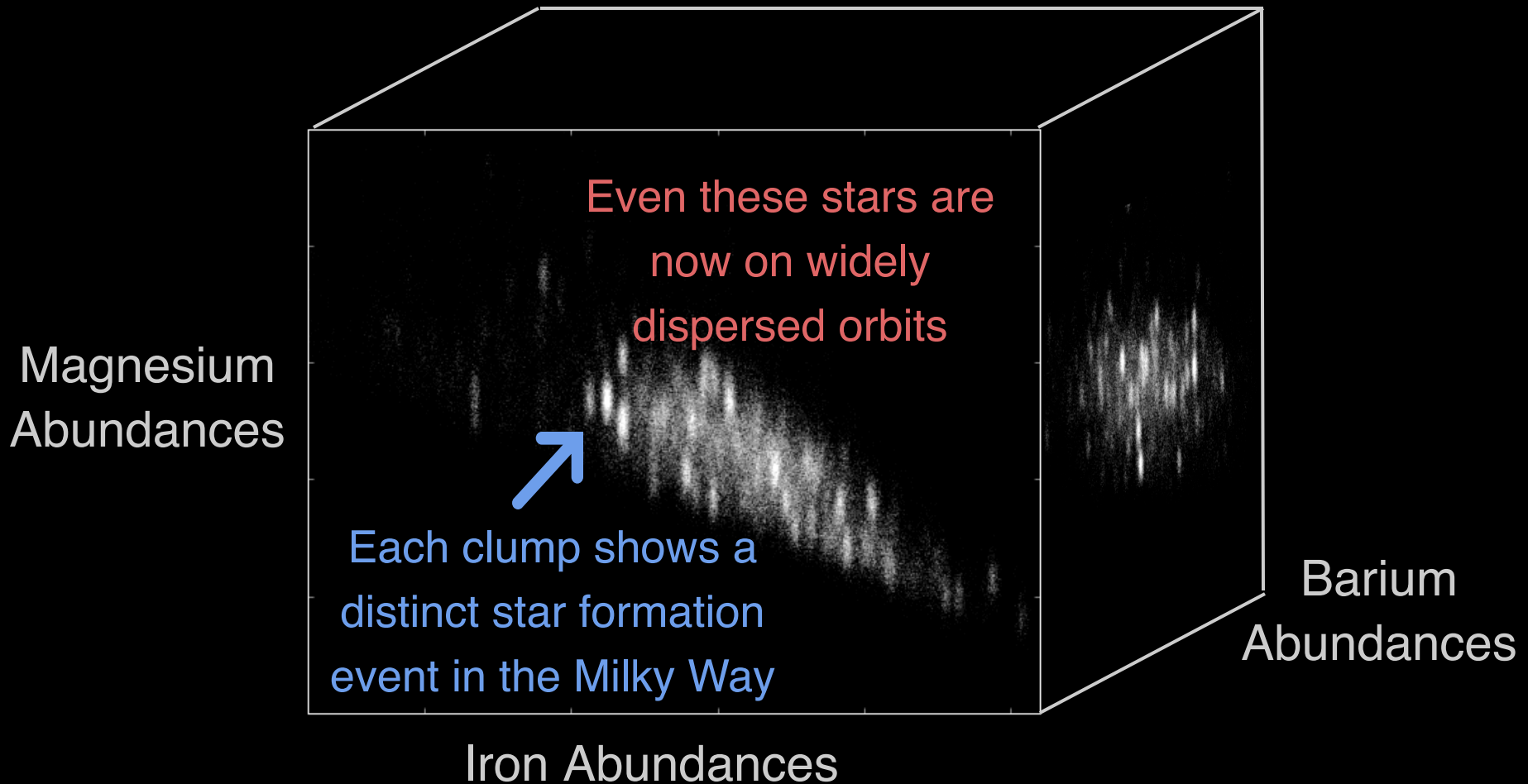
Dispersing ink in water

Lada & Lada 2003

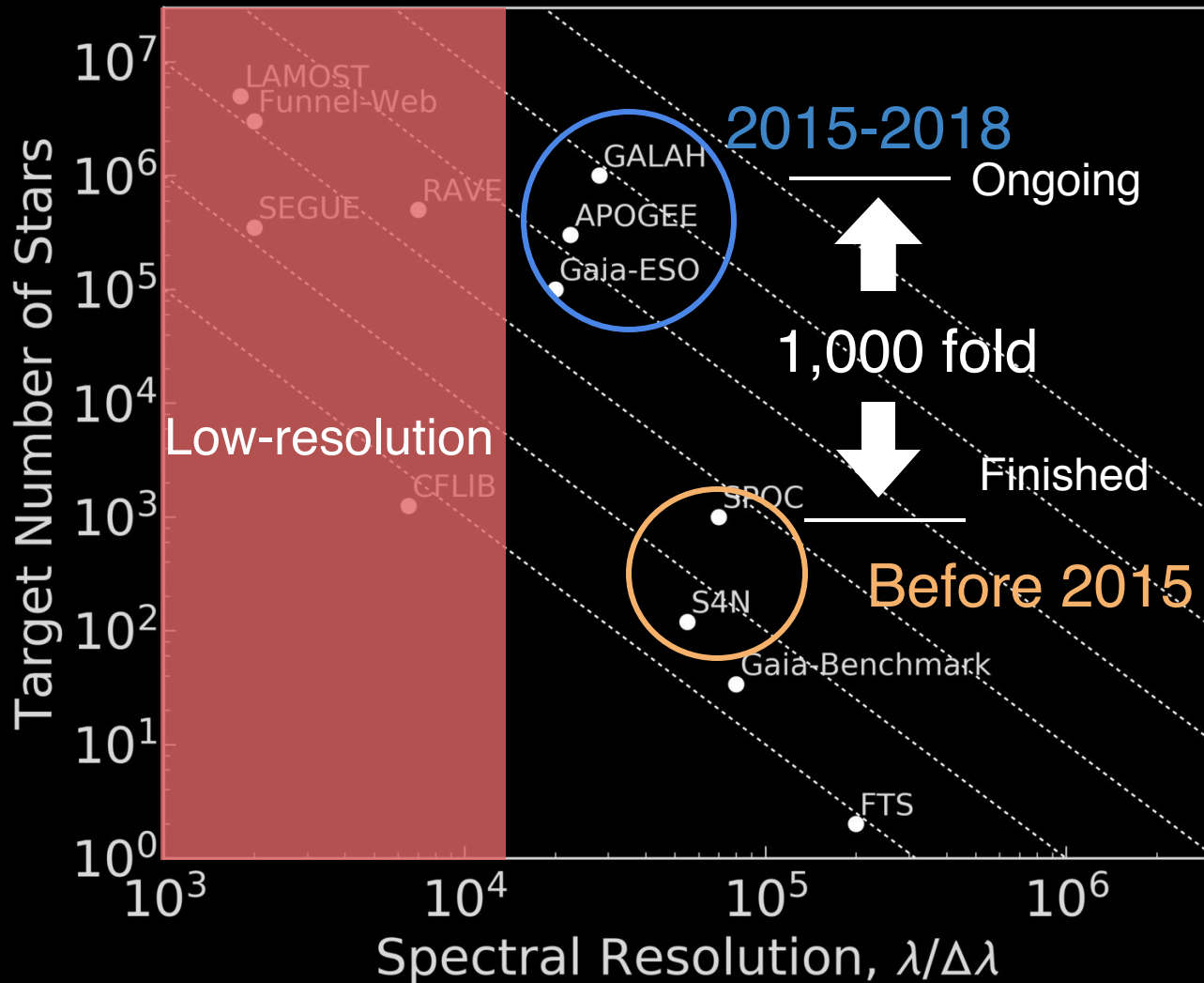
What is chemical tagging ?

Freeman & Bland-Hawthorn, 2002

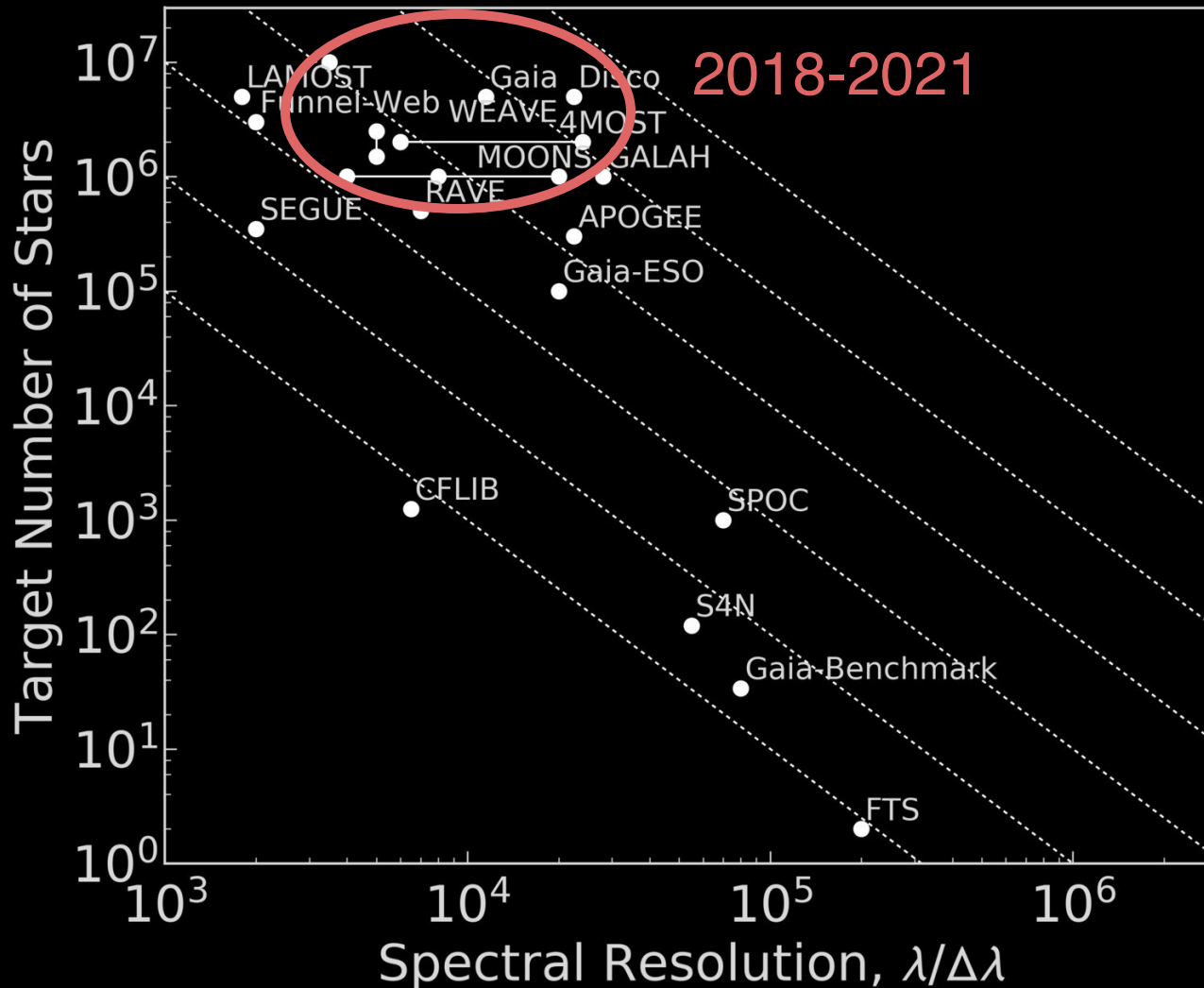
Chemically tagging the Milky Way in a multidimensional chemical space



Large spectroscopic surveys is rapidly changing the landscape of chemical tagging

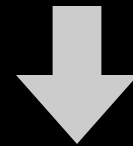


Large spectroscopic surveys is rapidly changing the landscape of chemical tagging



Why sample size matters for chemical tagging

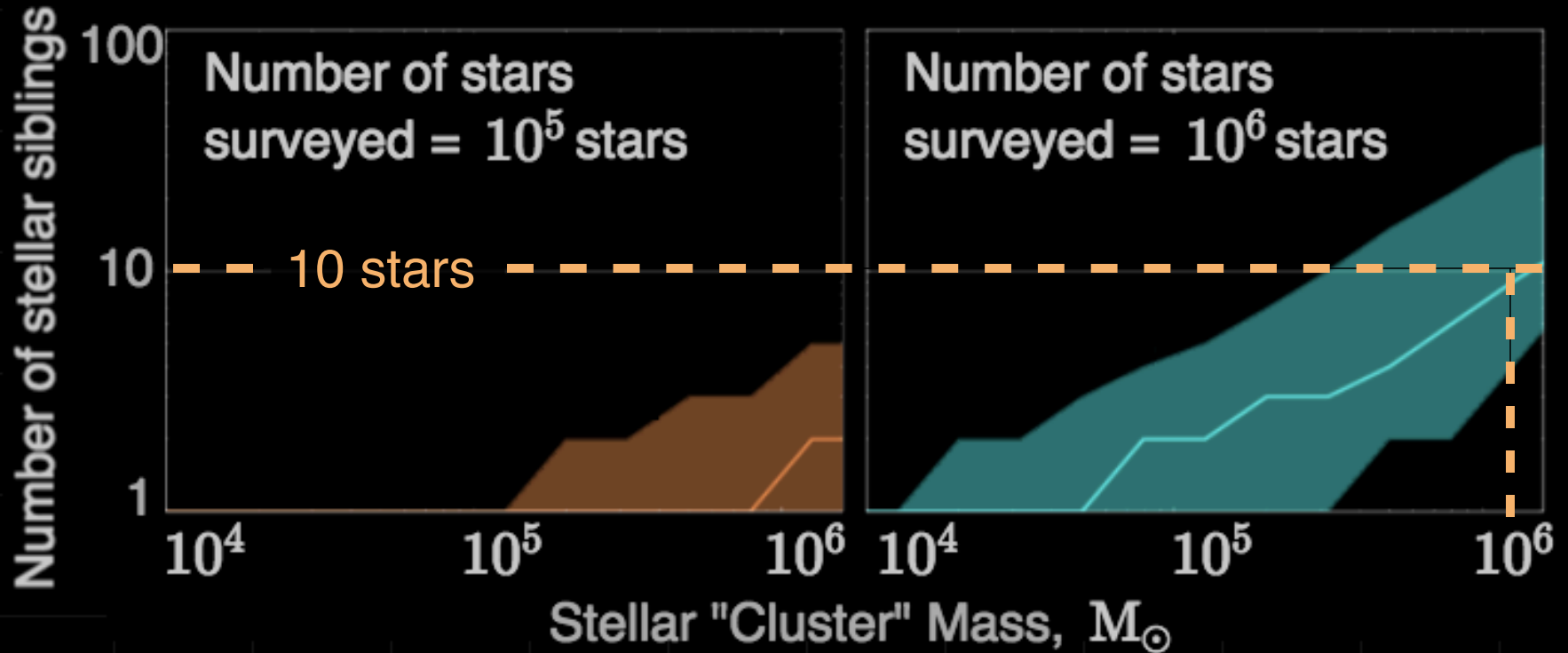
Most stars are now dispersed and are not observable



Low sampling rate

Mock Milky Way

Chemical tagging needs $> 10^6$ stars



Also see Bland-Hawthorn's work

Milky Way models

YST et al. 2015

Visualizing 30D chemical space with the tSNE-projection



Non-linear transdimensional mapping
while preserving substructures

$$f : 30\text{D} \rightarrow 2\text{D or } 3\text{D}$$

Due to the small sampling rate, currently,
it is hard to recover disrupted clusters

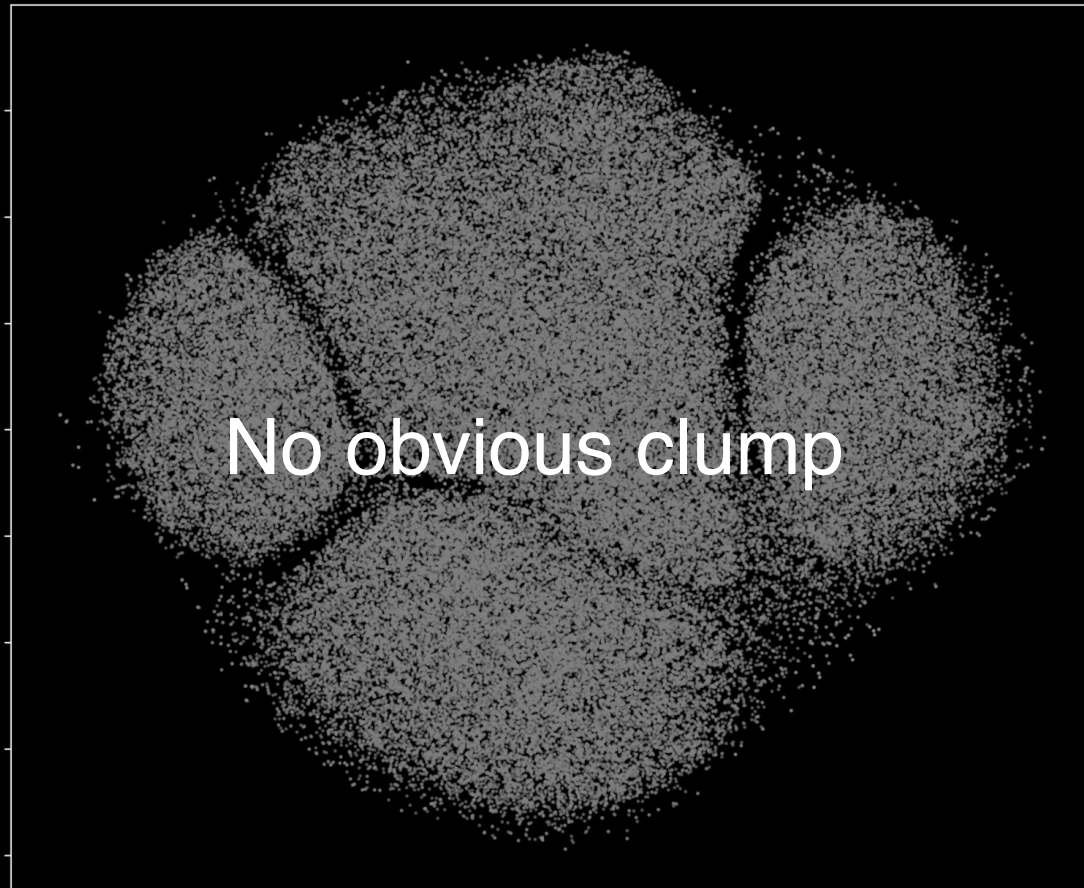
$$N_{\text{star}} = 3 \times 10^5$$

tSNE project of a
simulated Milky Way
with 30 elemental
abundances

Model empirically
the chemical space
using Galah iDR2

see YST et al. 2015

YST et al. 2016a

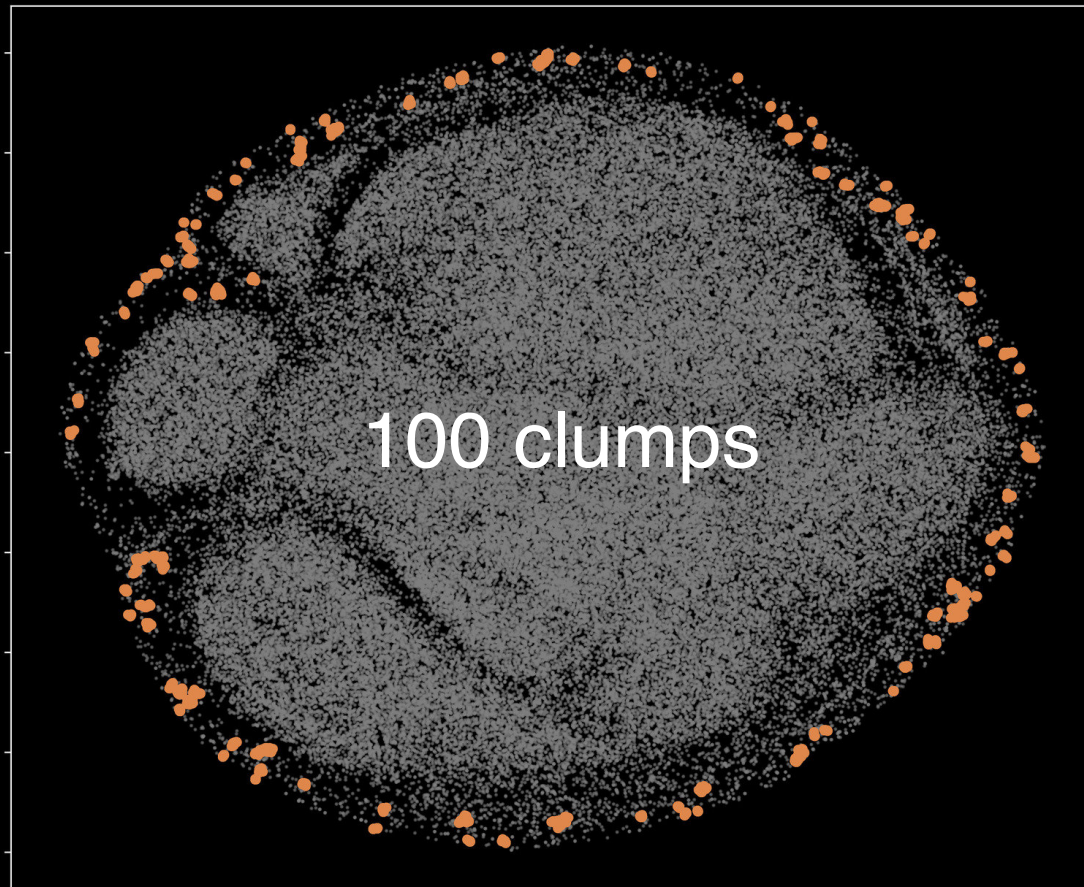


No obvious clump

tSNE projected chemical space

Galah, as a path finder, should find ~ 100 disrupted clusters when it is completed

$$N_{\text{star}} = 10^6$$



● Recovered clusters

Model empirically
the chemical space
using Galah iDR2

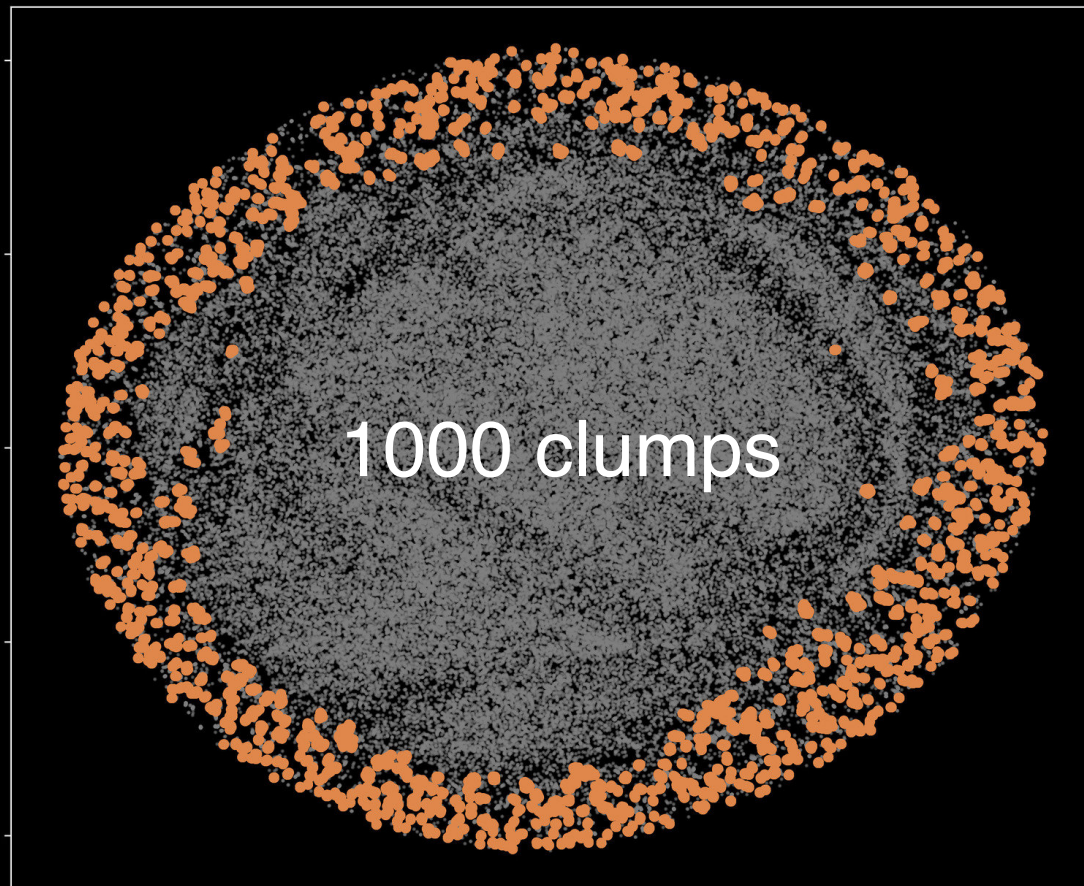
see YST et al. 2015

YST et al. 2016a

tSNE projected chemical space

Going beyond 10^6 stars will unravel
the full potential of chemical tagging

$$N_{\text{star}} = 10^7$$



● Recovered clusters

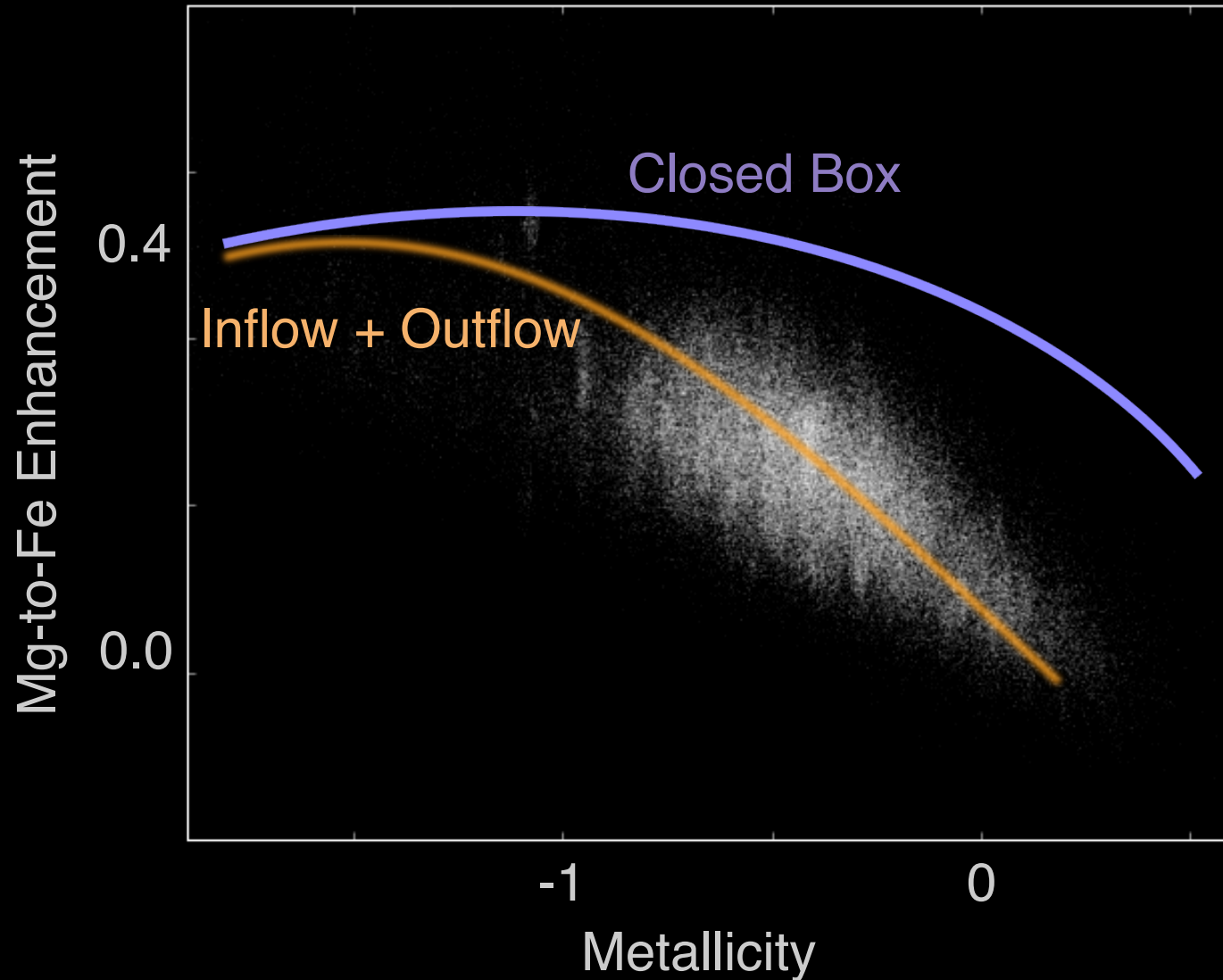
Model empirically
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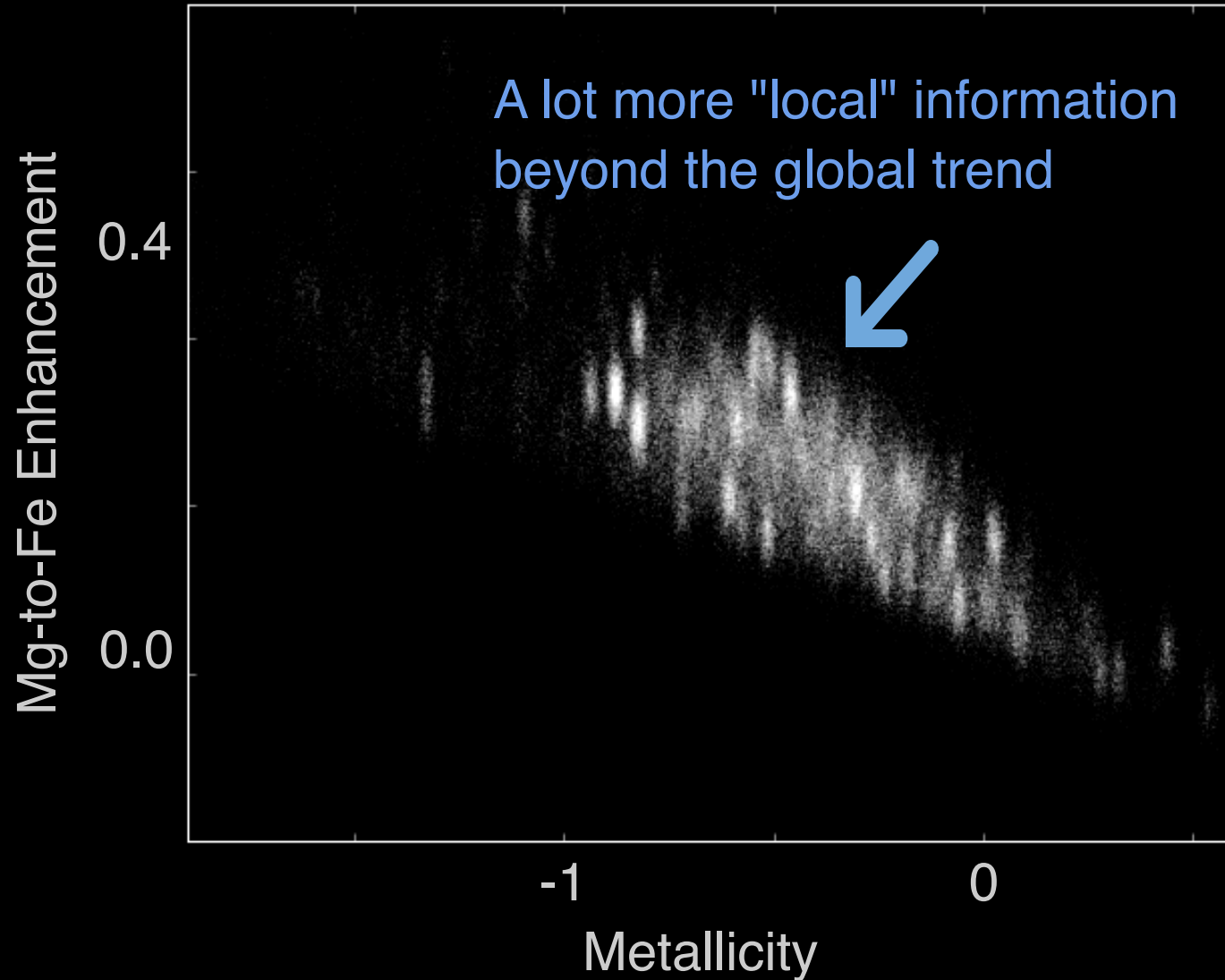
YST et al. 2016a

tSNE projected chemical space

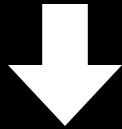
Classical Galactic archaeology



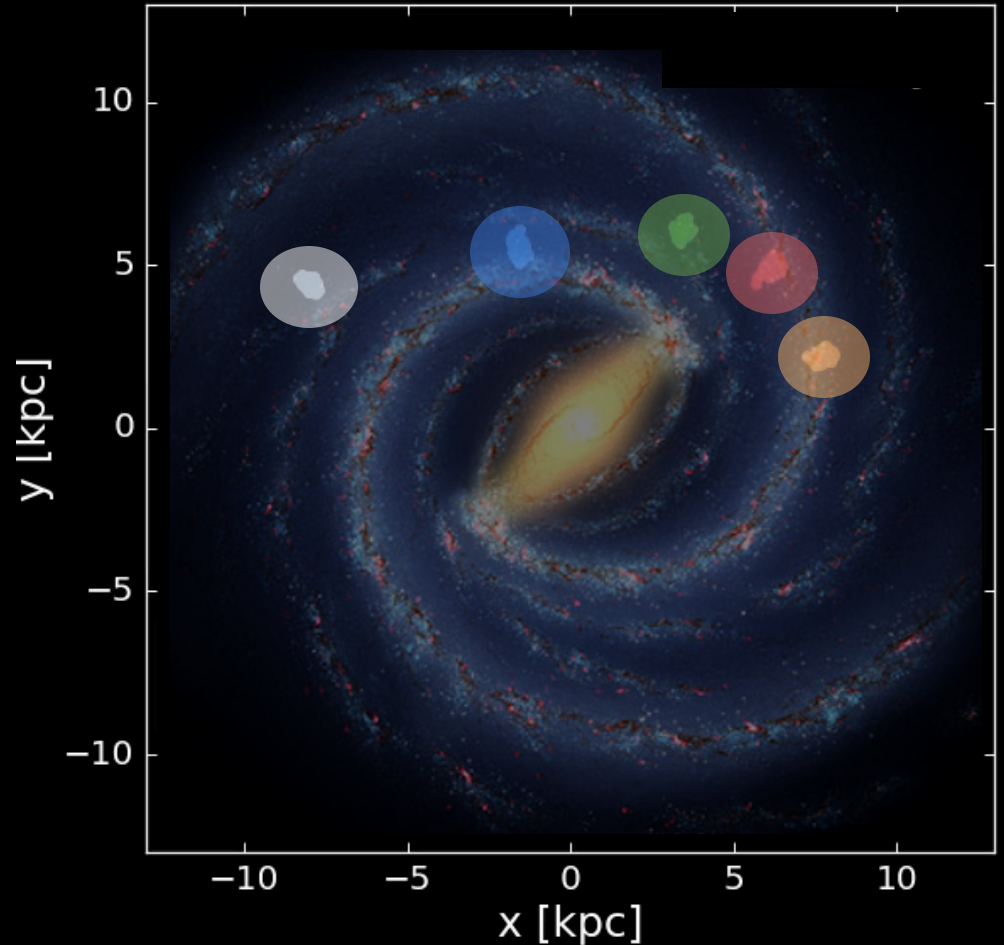
The new era of Galactic archaeology



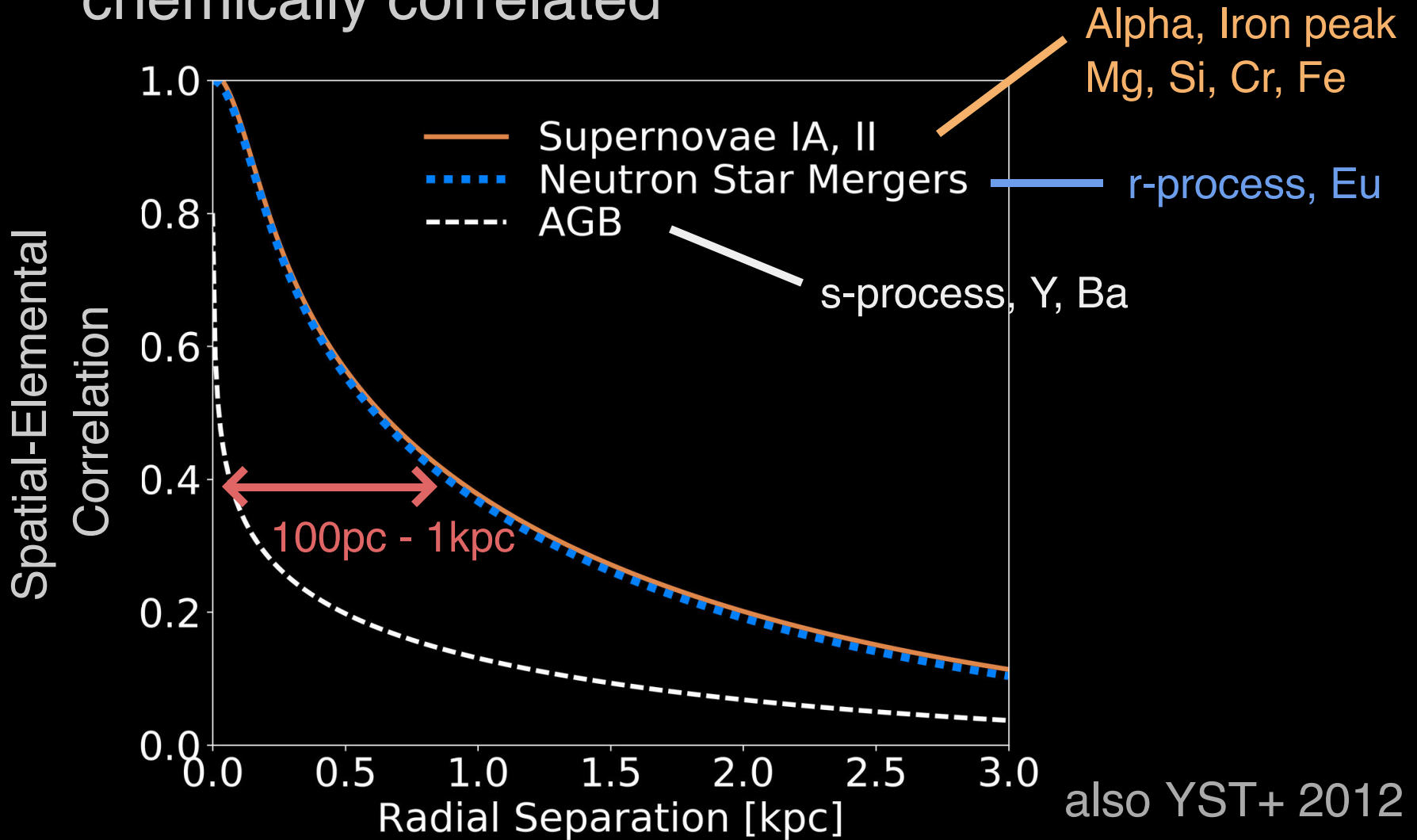
What is a "star cluster" in chemical tagging



The characteristic radius in which chemical abundances of stars are unique



The characteristic radius in which stars are chemically correlated



also YST+ 2012
Krumholz & YST 2017

Reconciling different results

- Chemical tagging can work

Hogg et al. 2016, Kos et al. 2017, Chen et al. 2017

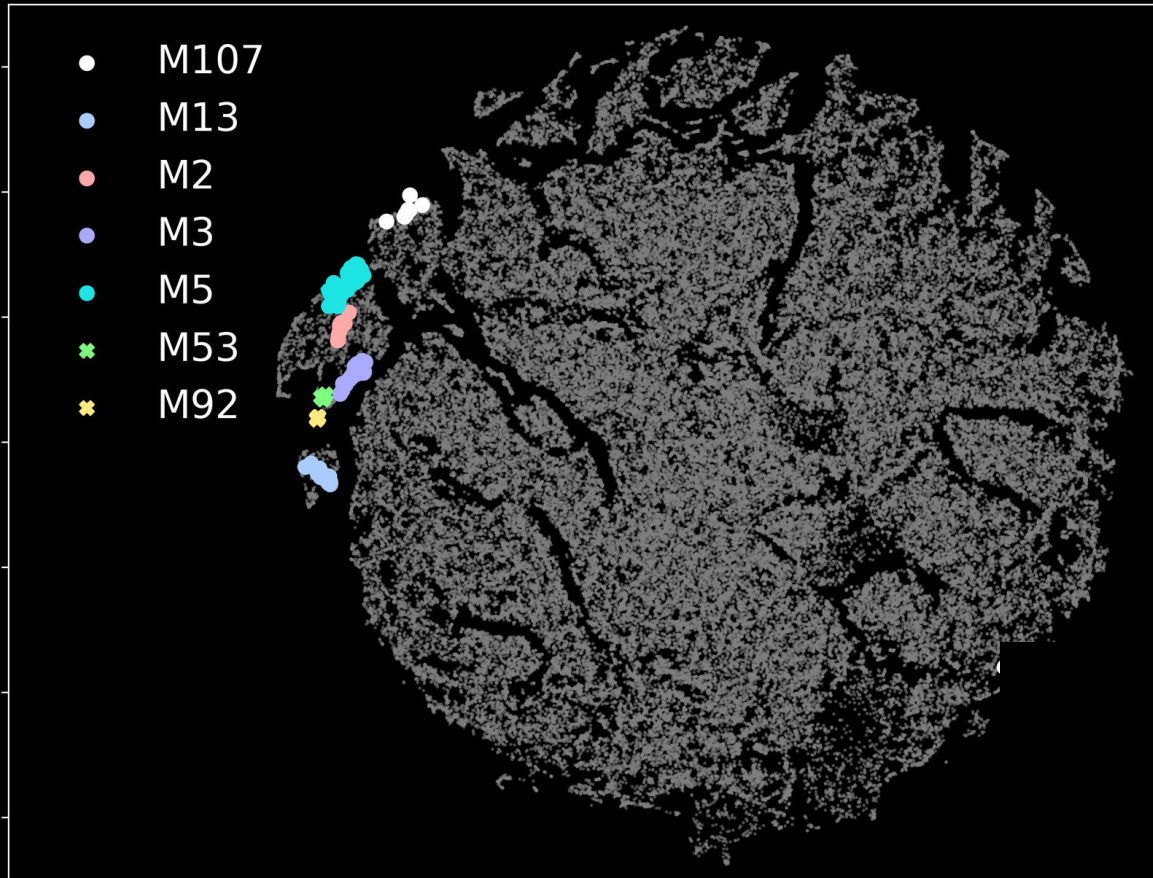
Globular clusters can be chemically tagged

- Chemical tagging cannot work

Blanco Cuaresma et al. 2015, Smiljanic et al. 2017

Open clusters are too chemically similar

Globular clusters can be easily tagged because

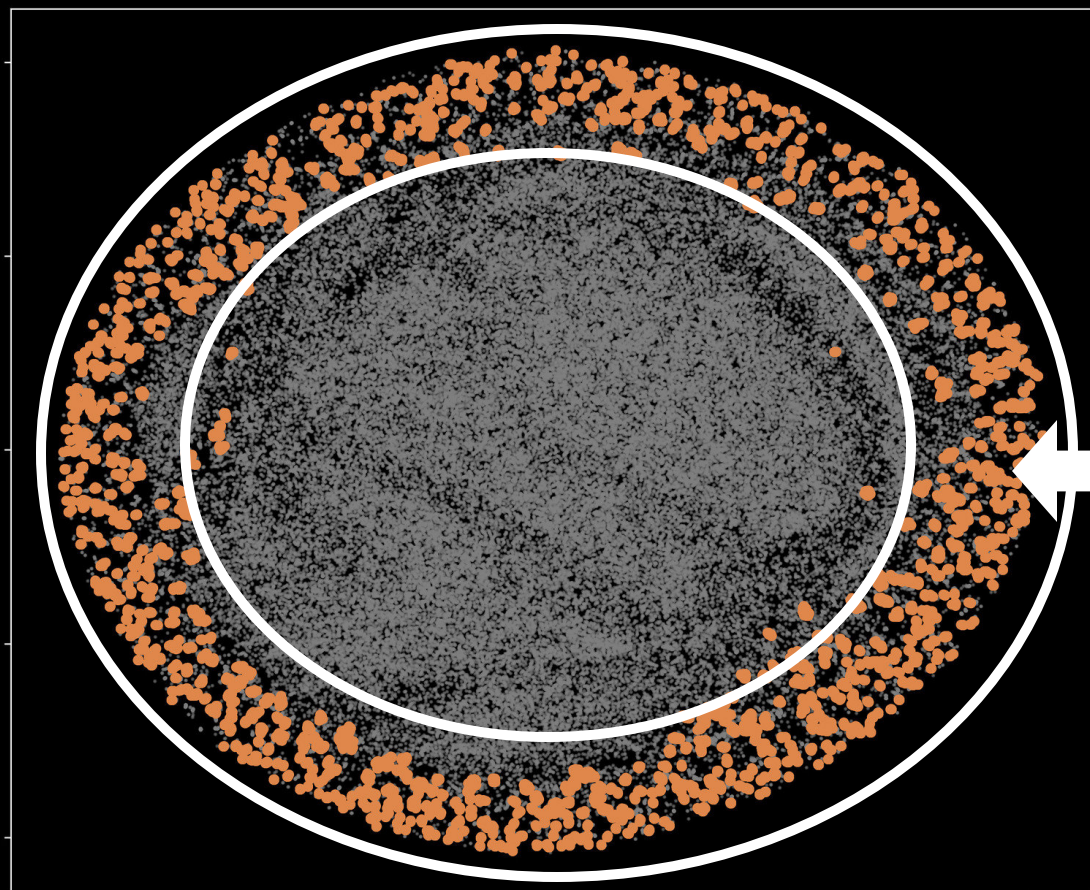


~~(1) Special targets~~
→ high sampling rate

(2) Located in
chemical regions
with a low
density/background

tSNE projection of APOGEE chemical and radial velocity space

Some parts of the chemical space will remain too crowded for "strong" chemical tagging



$$N_{\text{star}} = 10^7$$

● Recovered clusters

Chemically outlying/unique enough for "strong" chemical tagging

tSNE projected chemical space

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"Statistical" Chemical Tagging

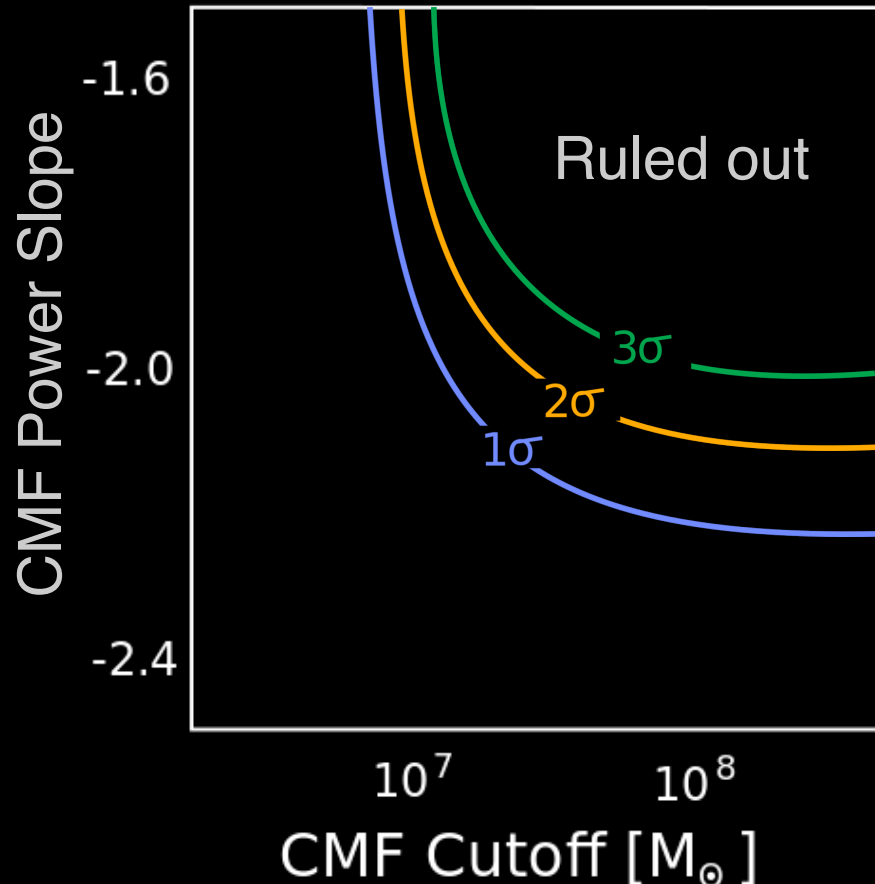
Study the elemental-elemental correlation

Only small clusters

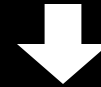
With massive clusters

More "clumpy" in
chemical space

The Milky Way cluster mass function for the old disk population (first 5 billion years)



Analyzing APOGEE DR12



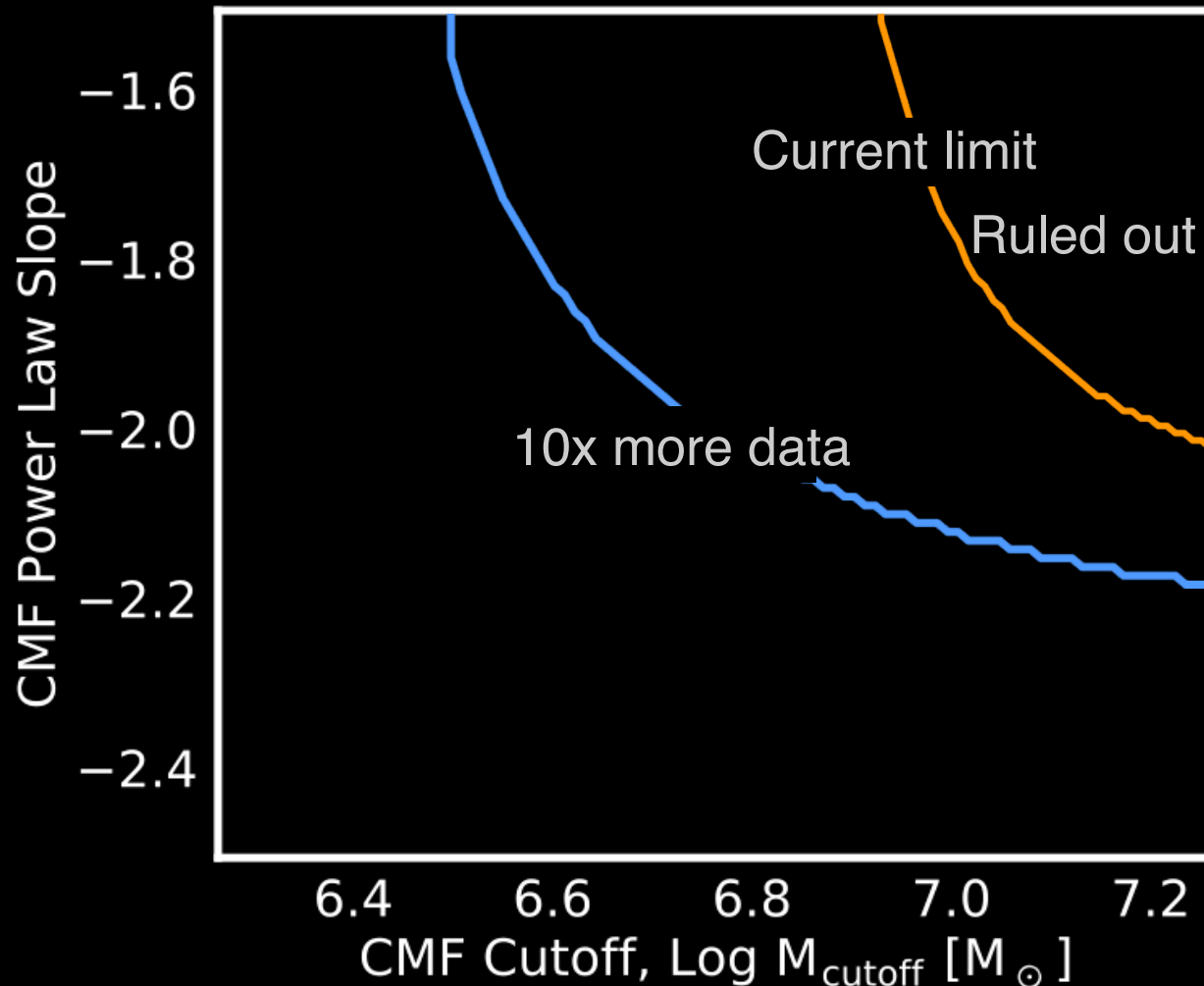
nil detection



The Milky Way did not form any chemically homogeneous clusters $> 10^7 M_{\odot}$

YST et al. 2016a

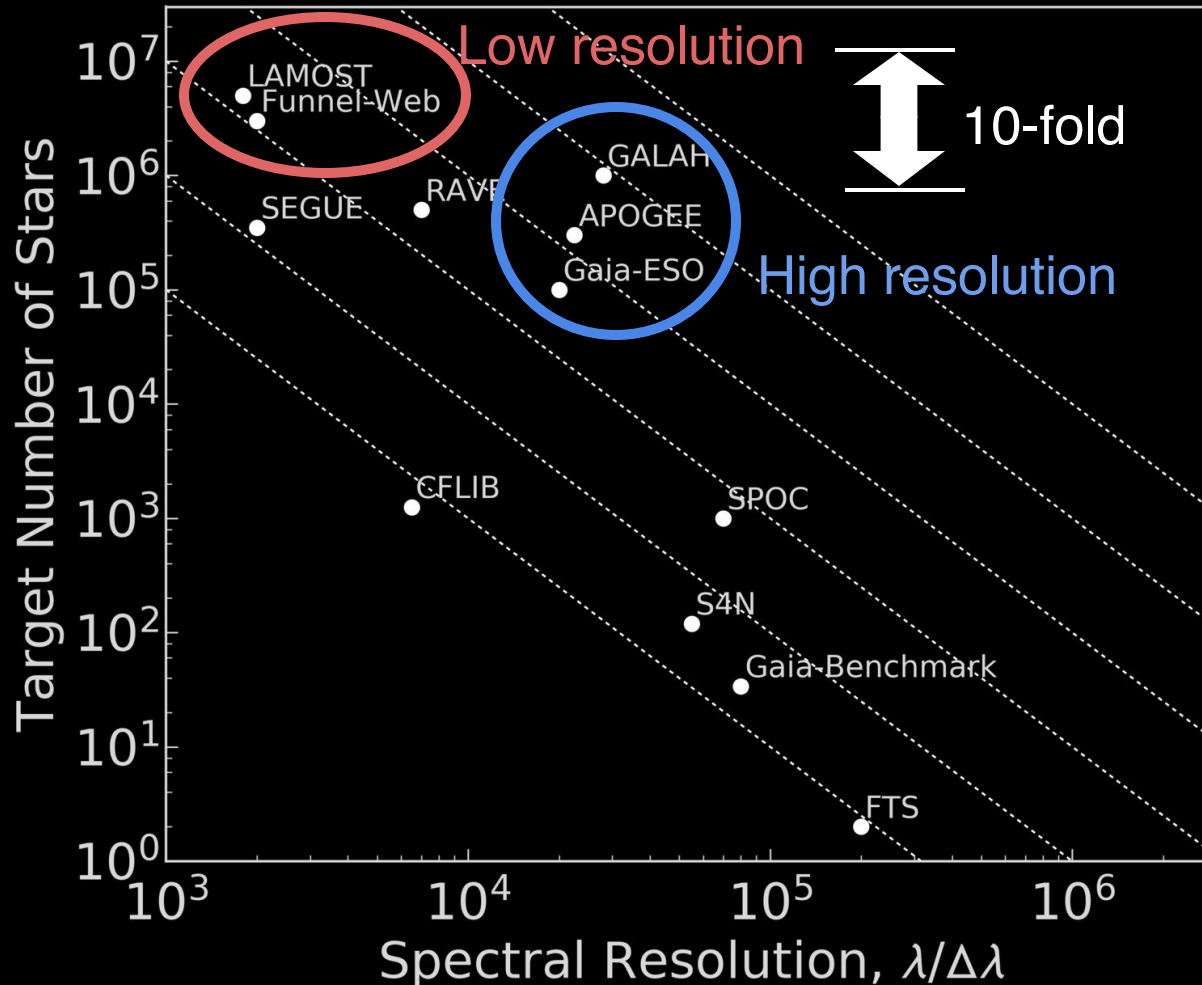
GALAH will answer whether $10^6 M_{\odot}$ chemically homogeneous star forming associations existed



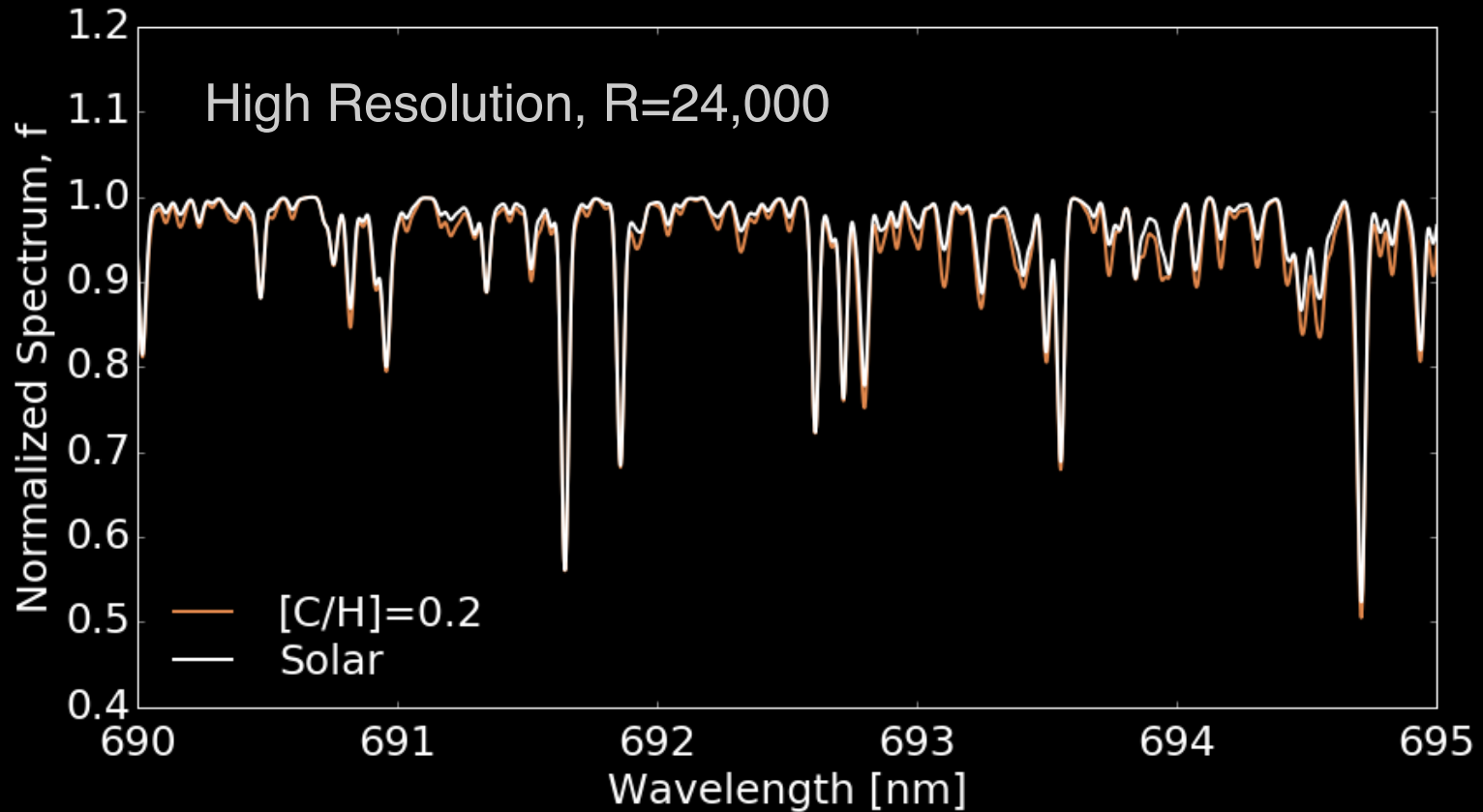
Moving forward

Spectra from more stars ($> 10^6$ stars)

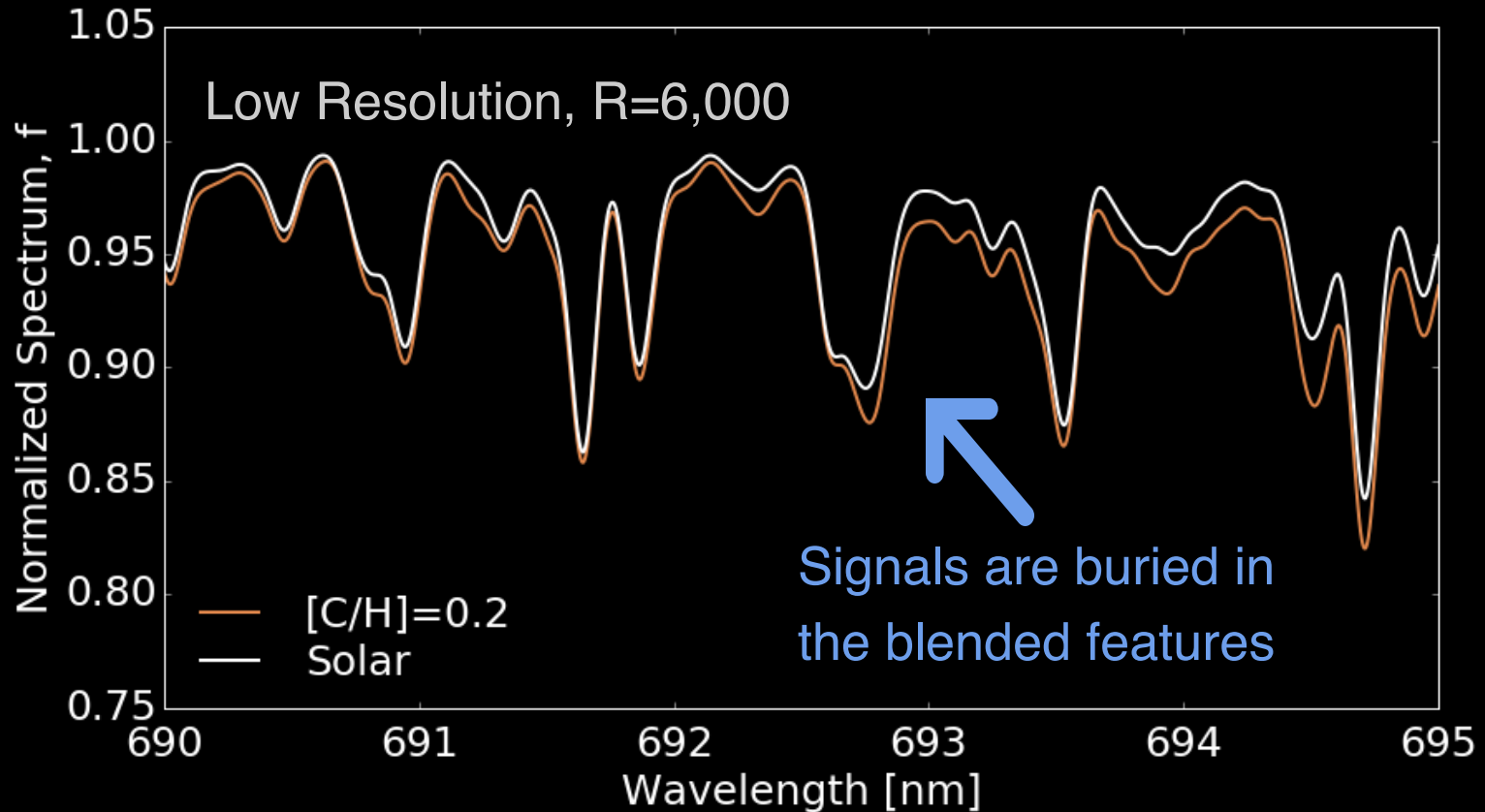
Extract elemental abundances from low-resolution spectra to improve significantly the sample size



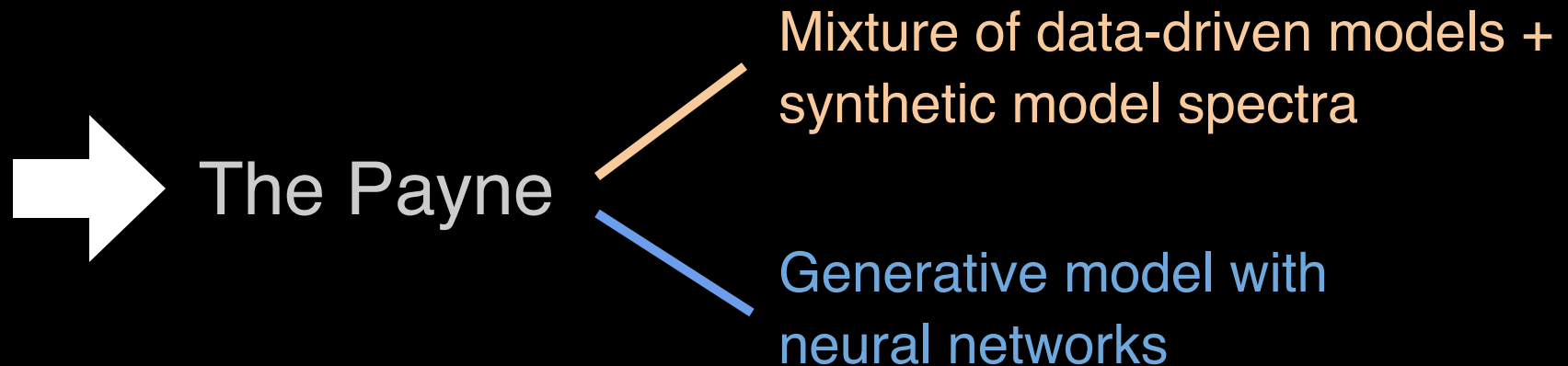
The previous bottleneck for low-resolution spectra :
modeling blended features with > 20 parameters



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modeling blended features with > 20 parameters



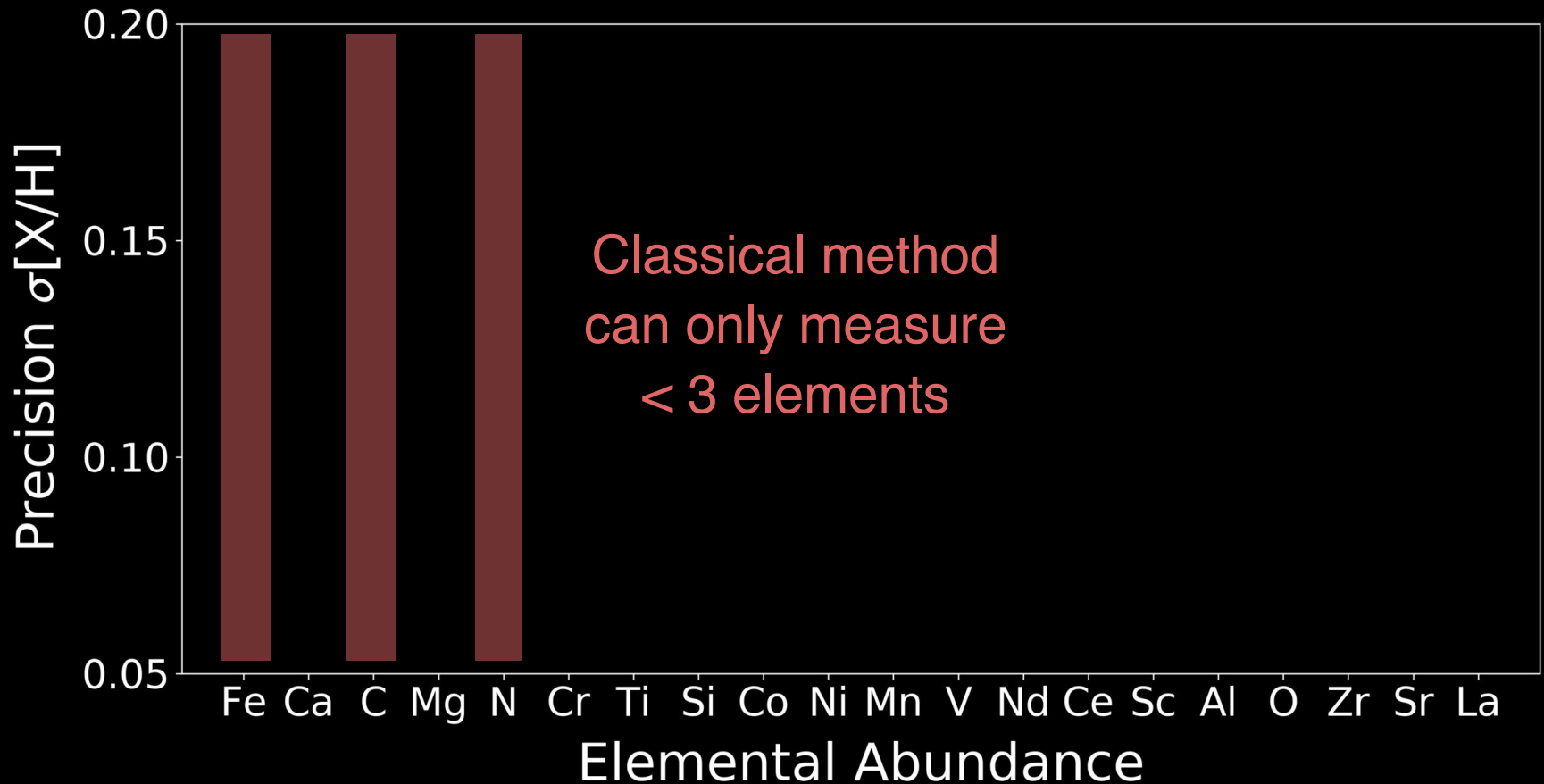
To extract information from blended features, we have to fit > 20 elemental abundances simultaneously



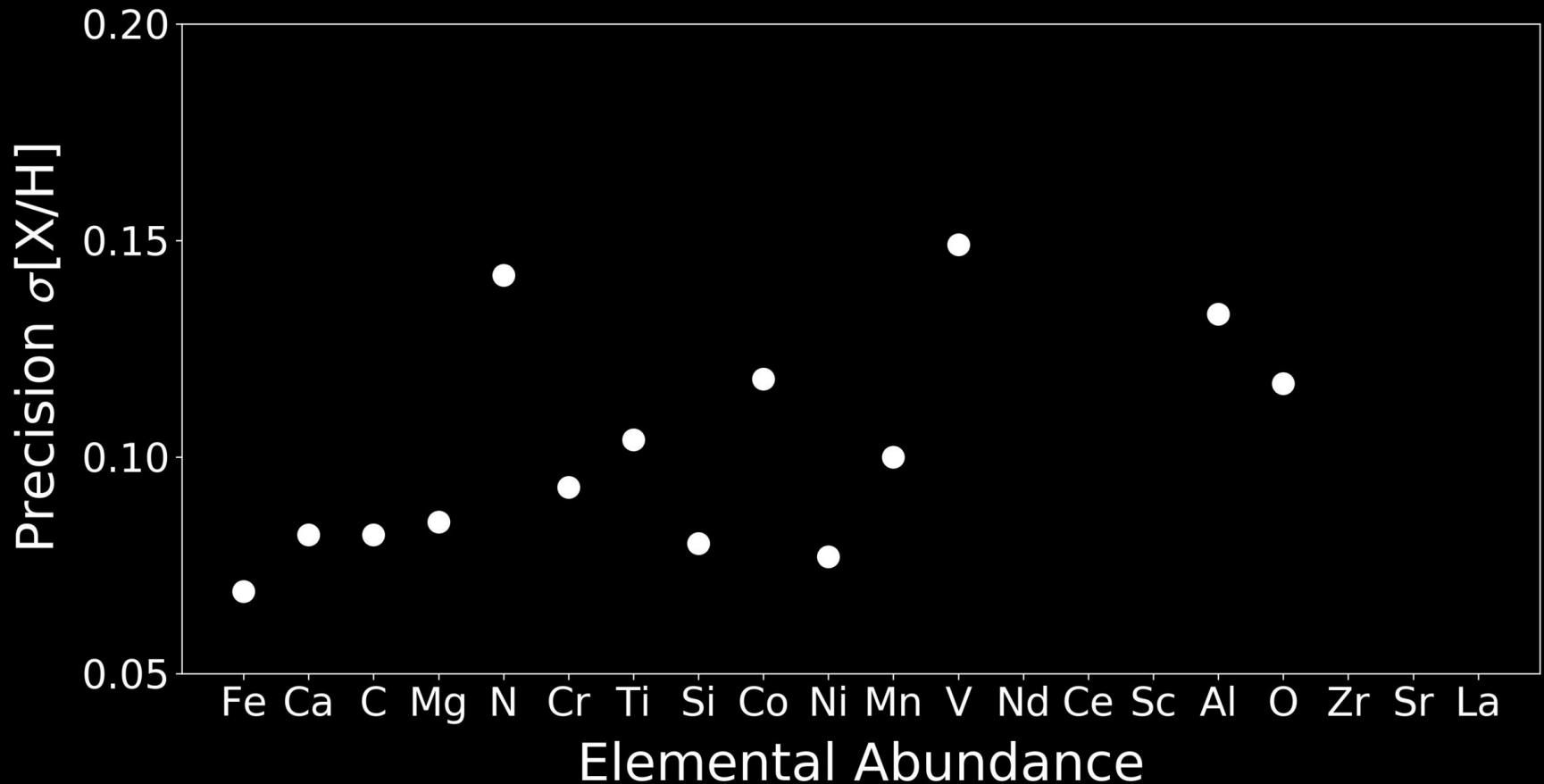
See also "The Cannon"
Work from A. Casey, M. Ness

YST et al. 2016b, 2017a, 2017b
Rix, YST et al. 2016

We measured 14 elemental abundances from
LAMOST $R = 1,800, S/N_{\text{pixel}} > 30$ spectra

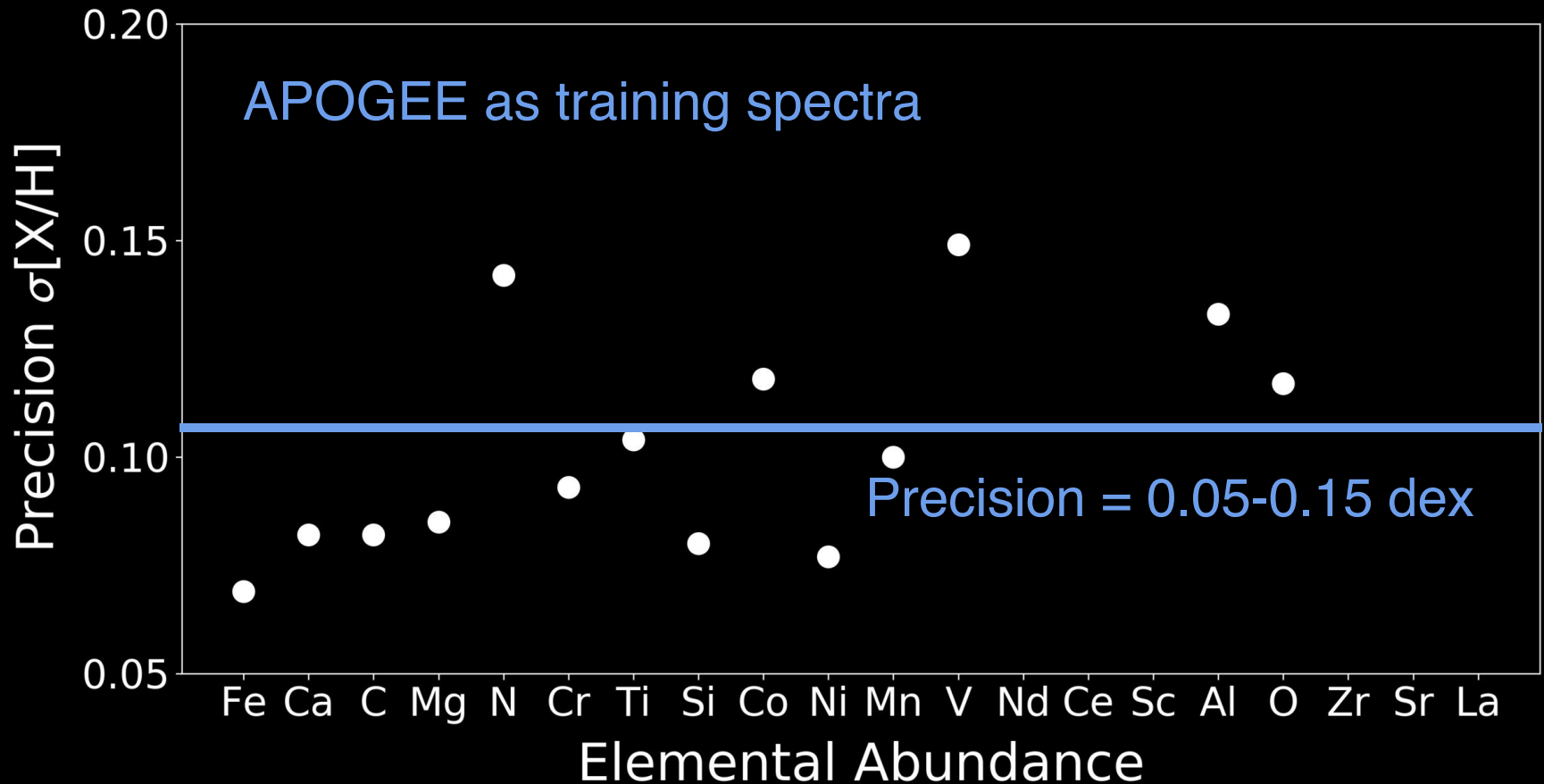


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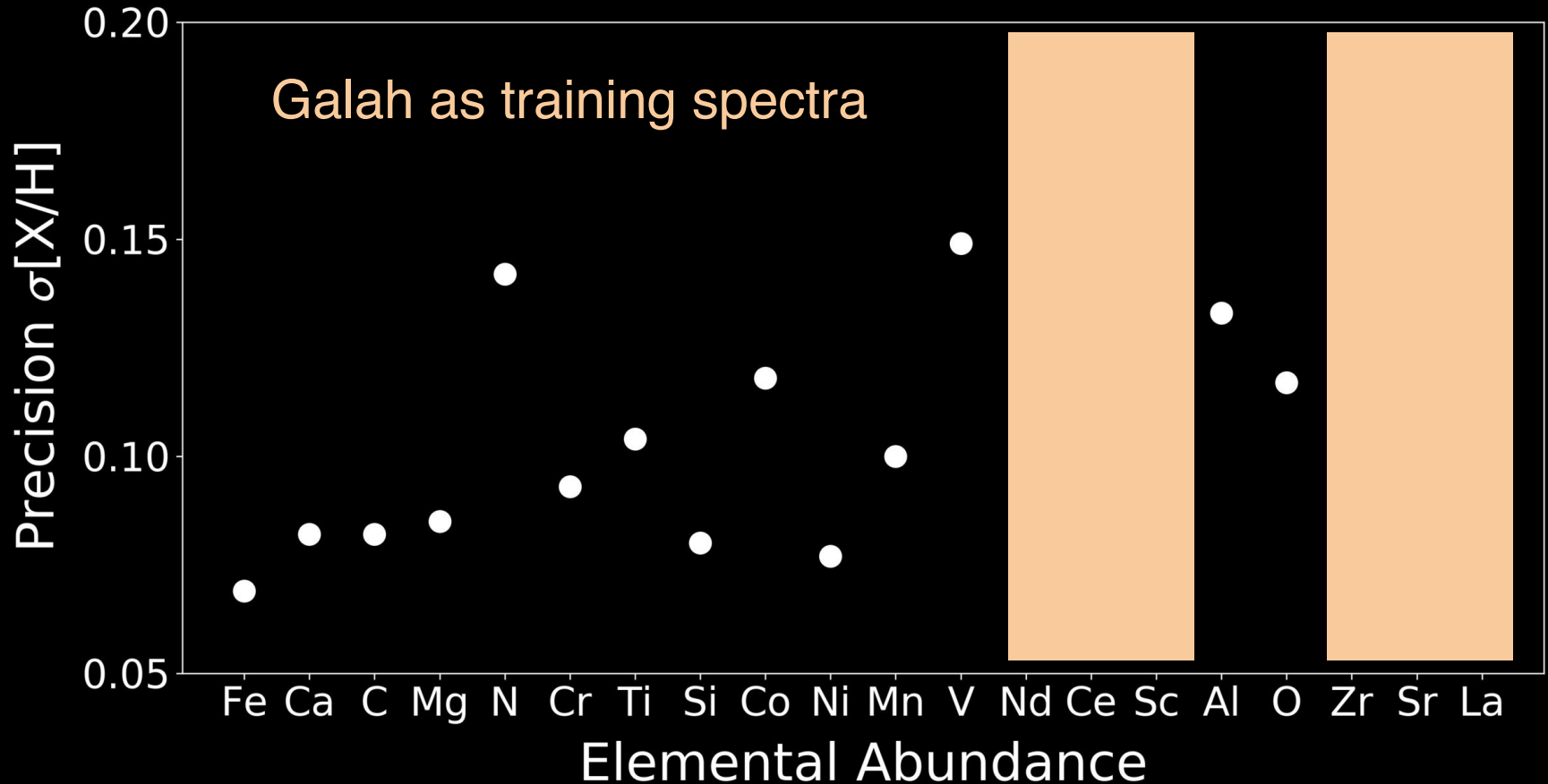


YST et al. 2017b

We measured 14 elemental abundances from
LAMOST $R = 1,800, S/N_{\text{pixel}} > 30$ spectra



Using Galah as training spectra will enable > 20 elemental abundances from $R=2000$ spectra



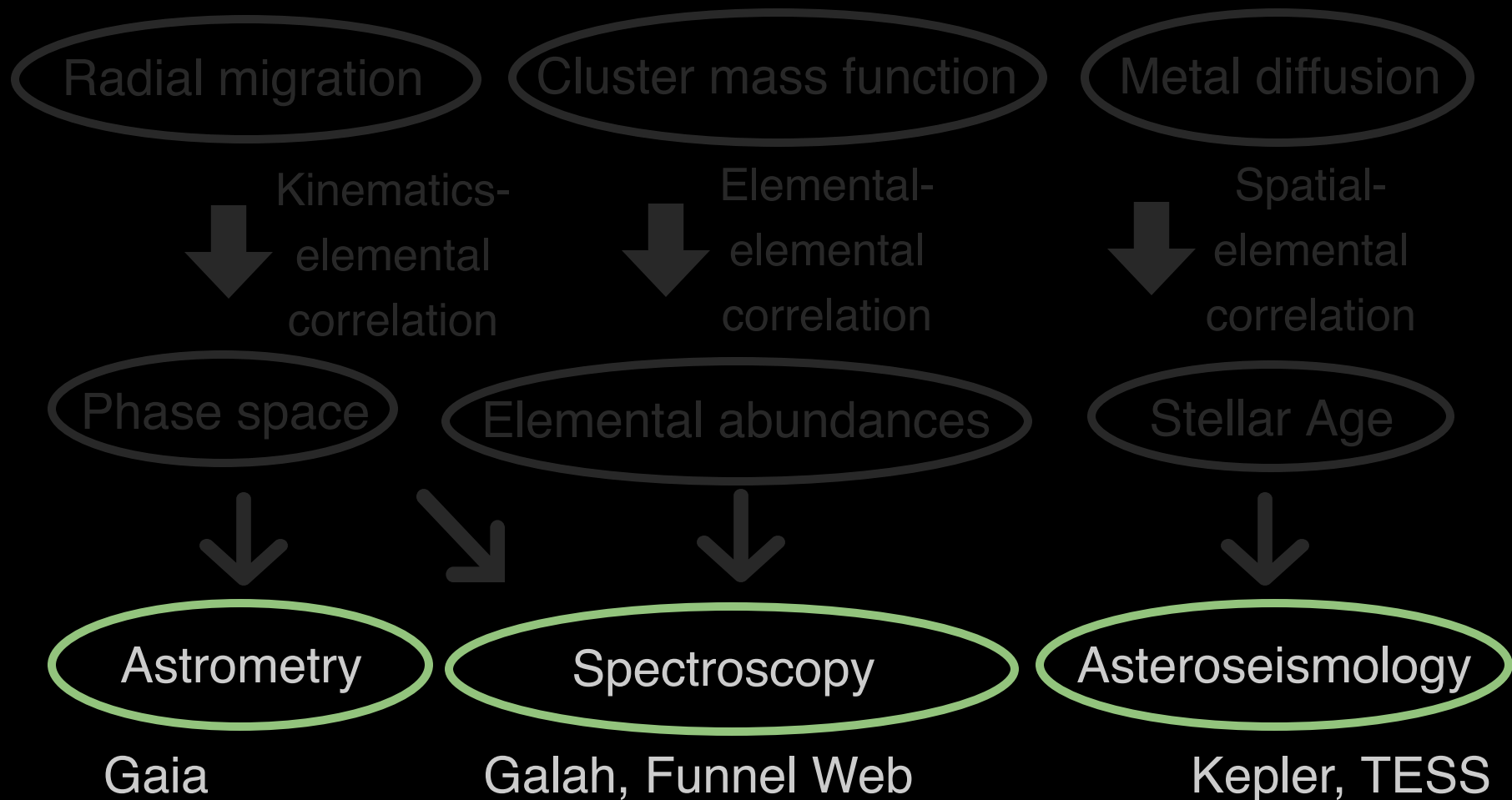
YST et al. 2017a

Summary: "strong" chemical tagging

- Reconstruct individual star clusters, even the members are now on widely dispersed orbits
- A "weak" form has been realized - e.g., tagging globular clusters
- The strong form is still largely limited by the sampling rate and requires $> 10^6$ stars
 - ➔ A combination of high- and low-resolution surveys

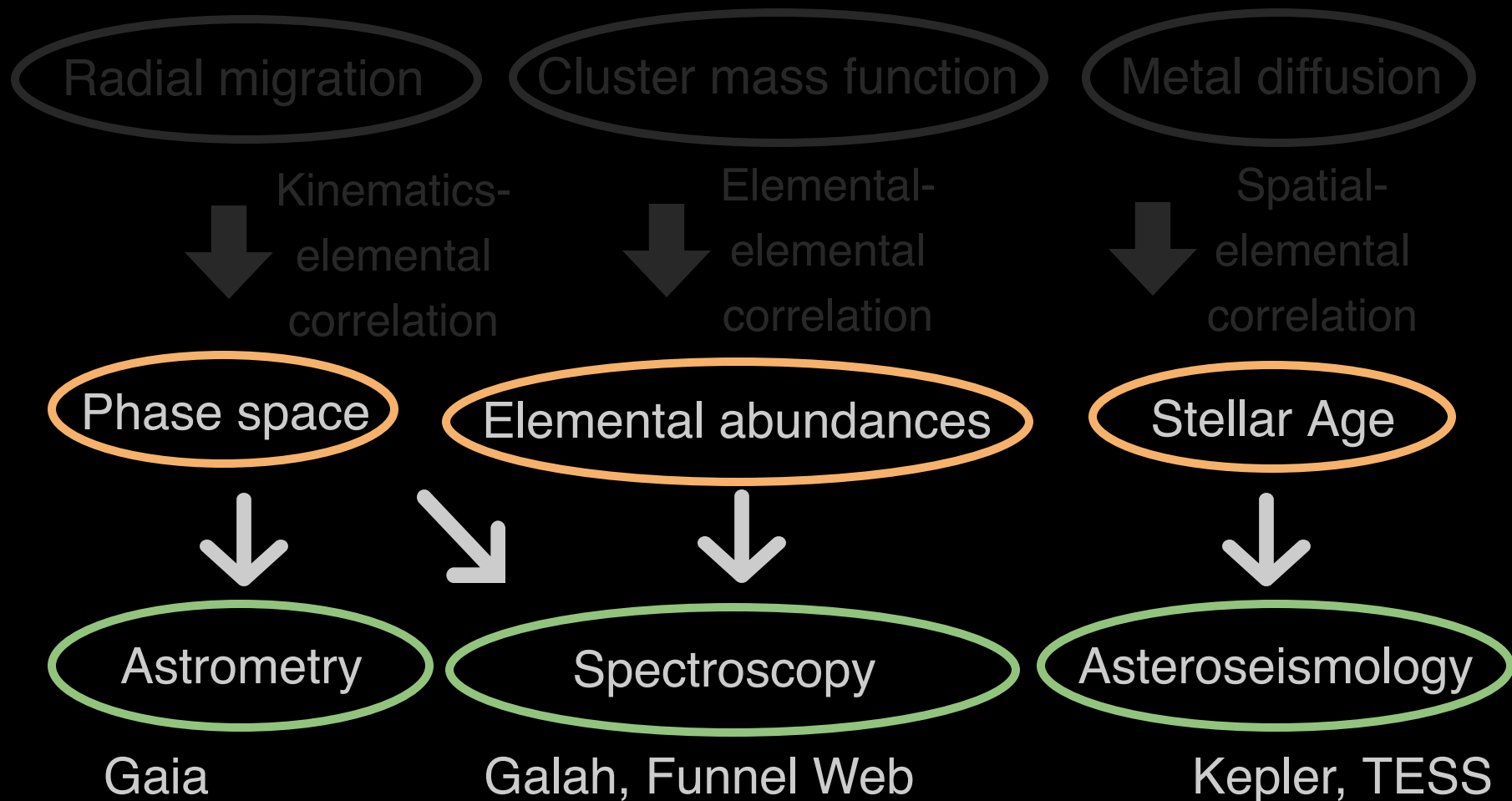
Summary : ○ Theory ○ Parameters ○ Observables

"statistical" chemical tagging



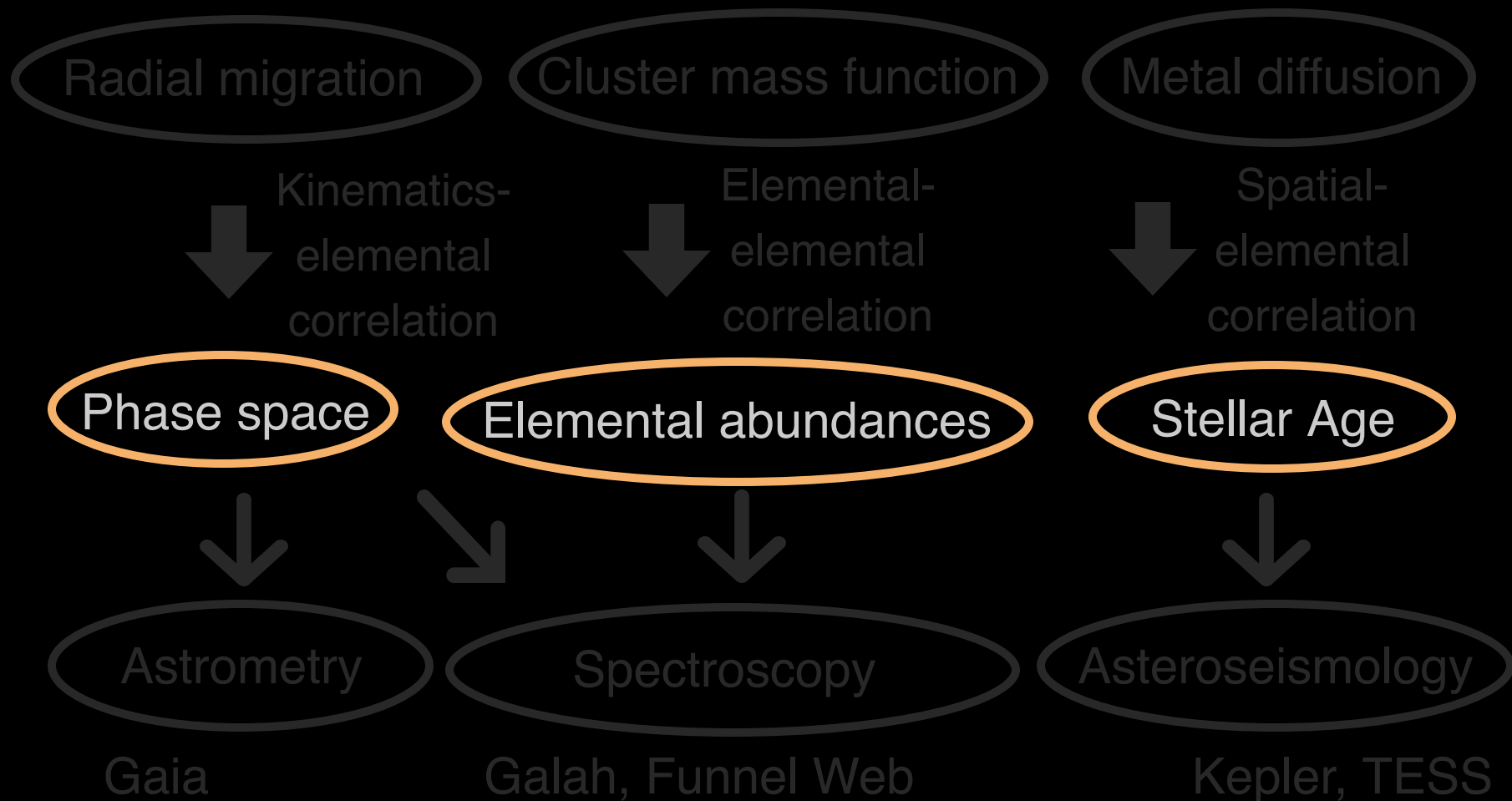
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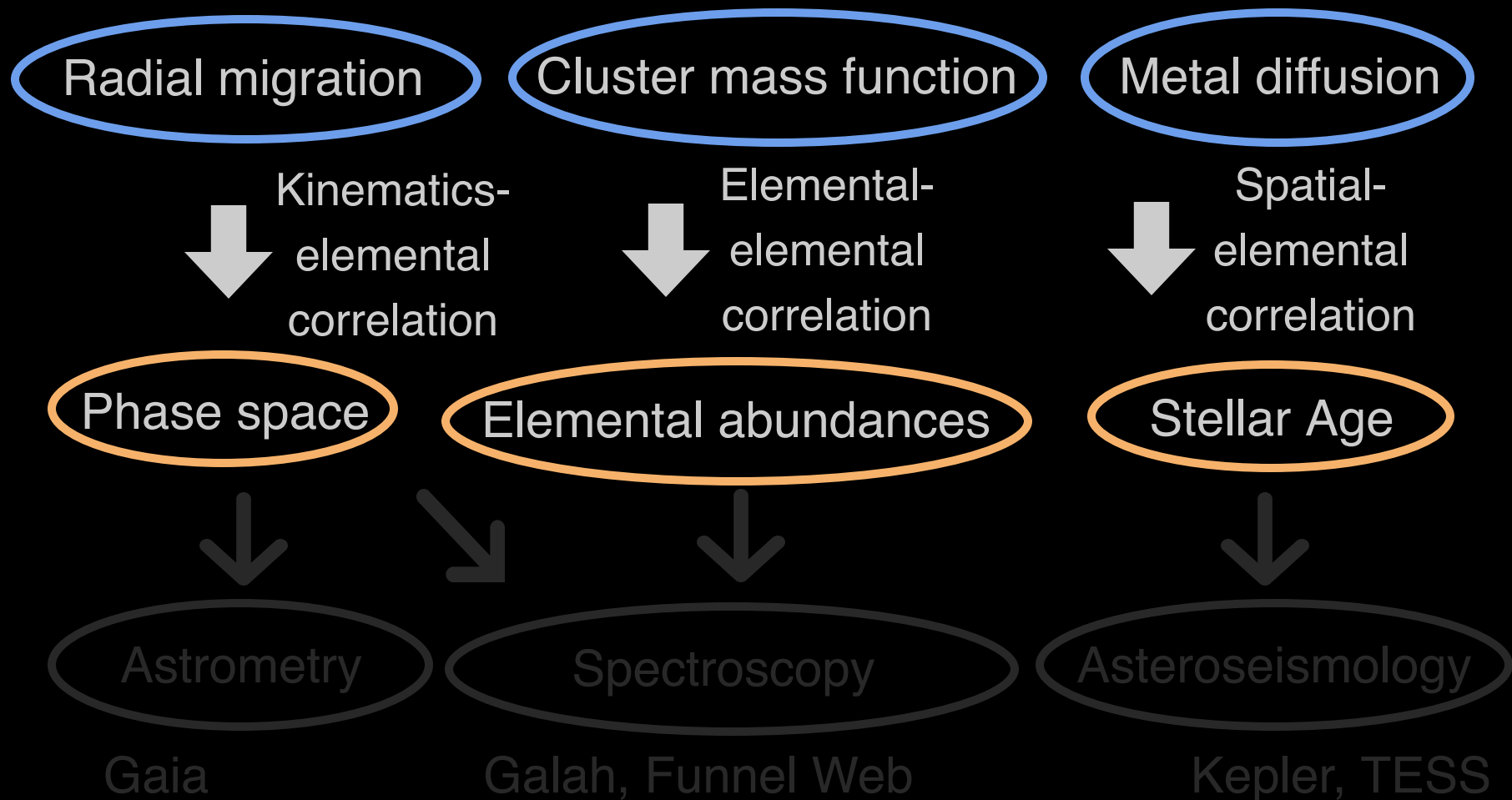
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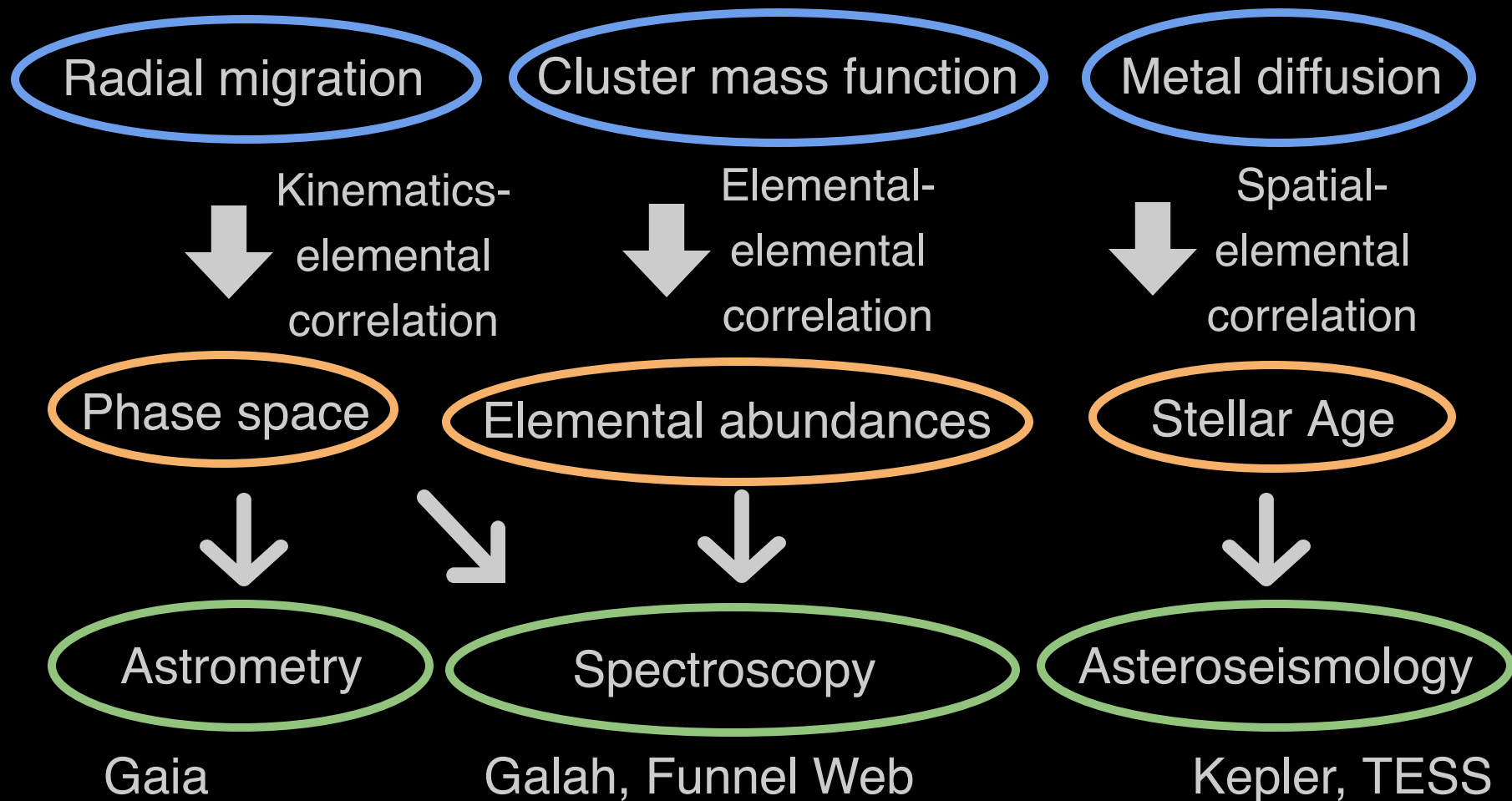
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Summary : ○ Theory ○ Parameters ○ Observables

"statistical" chemical tagging



Understanding the dynamical history of the Milky Way

Cold water

Hot water

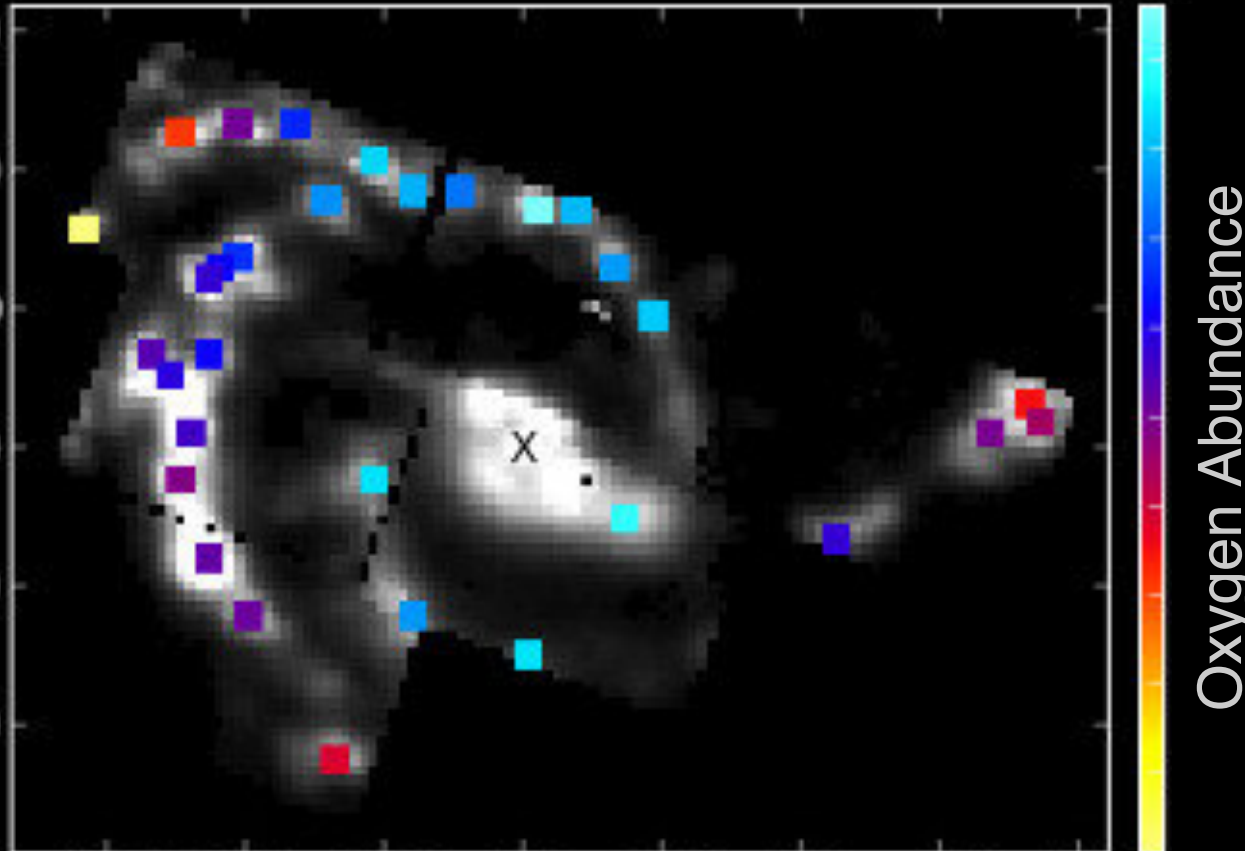
"Cold" dispersion

"Hot" dispersion

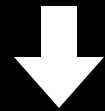
Connection to other IFU extragalactic studies

e.g., MaNGA, SAMI, Hector

spatial-elemental correlation
of H II regions



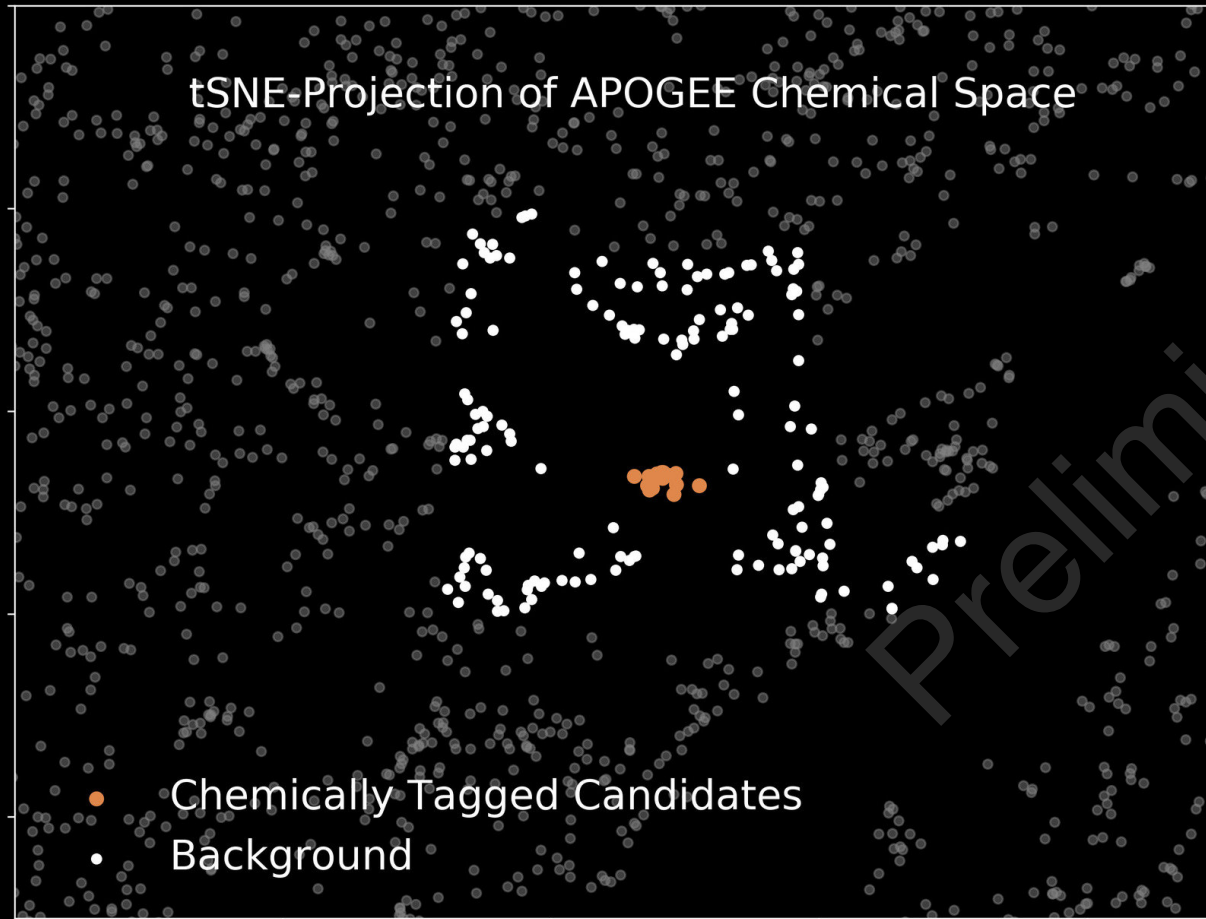
Oxygen Abundance



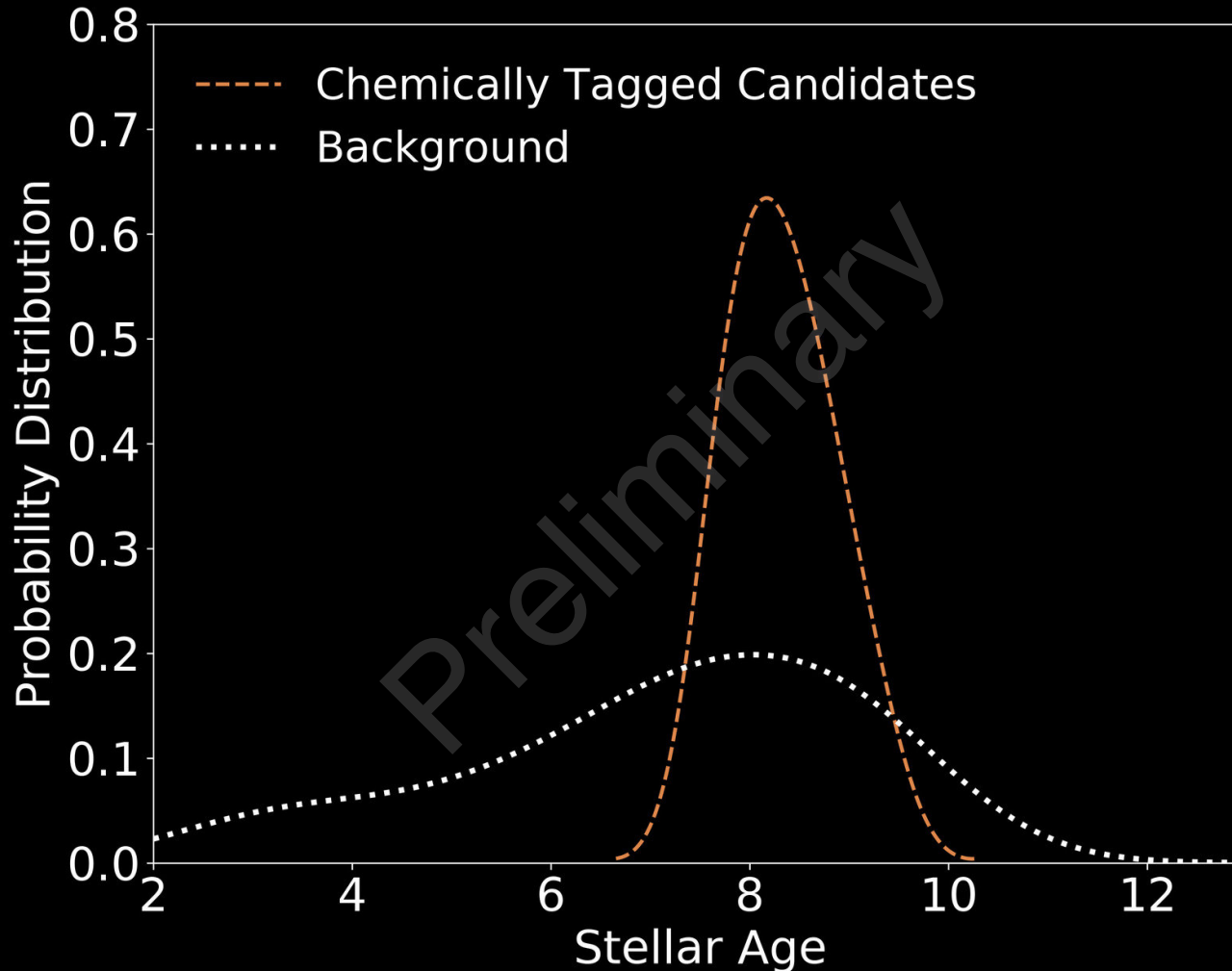
Characteristic
radius for
chemical tagging

Credit: Michael Dopita

We recover a disrupted cluster (candidate) through chemical tagging in APOGEE



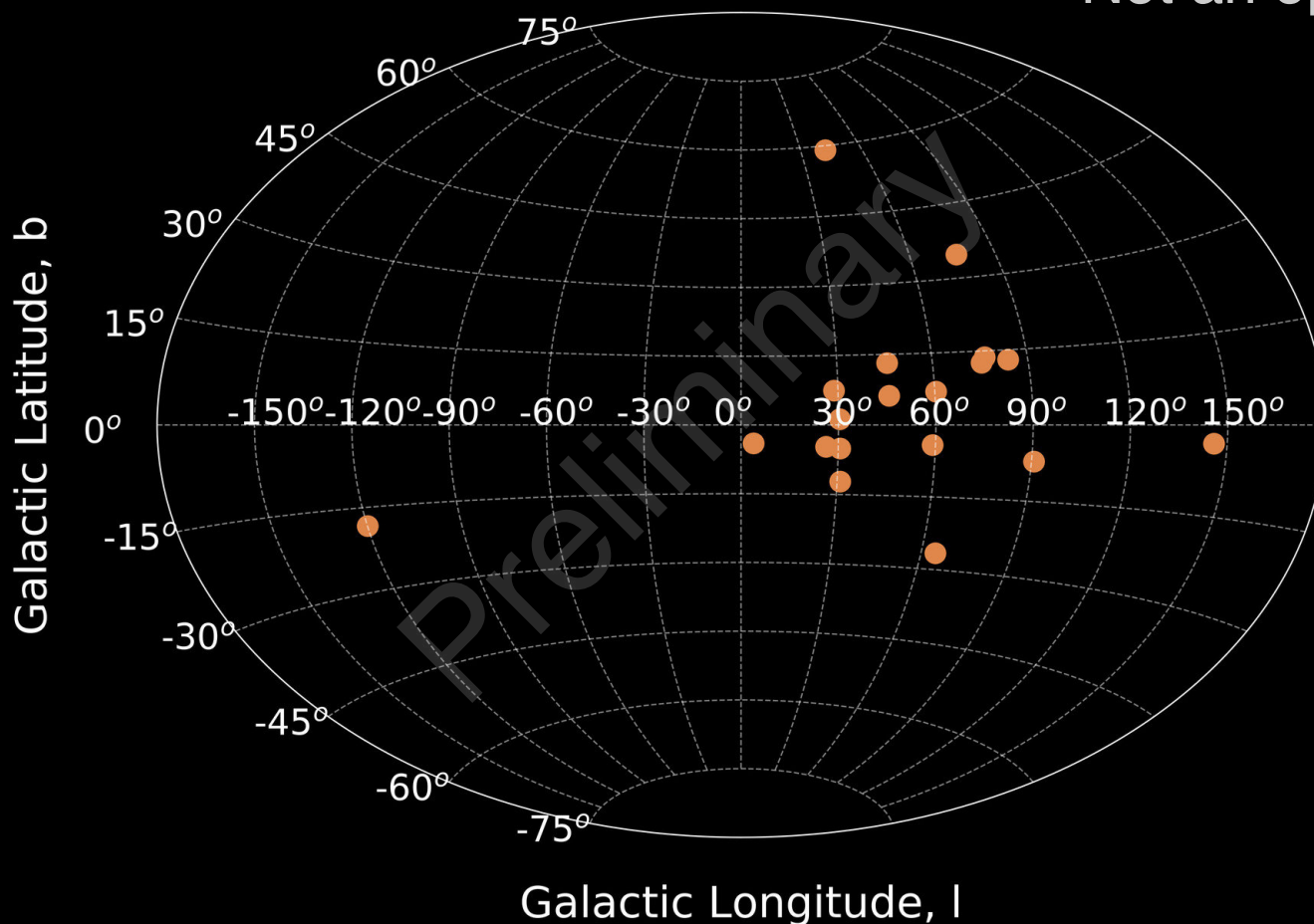
The candidates have similar stellar ages



Stellar ages are derived using C/N ratio

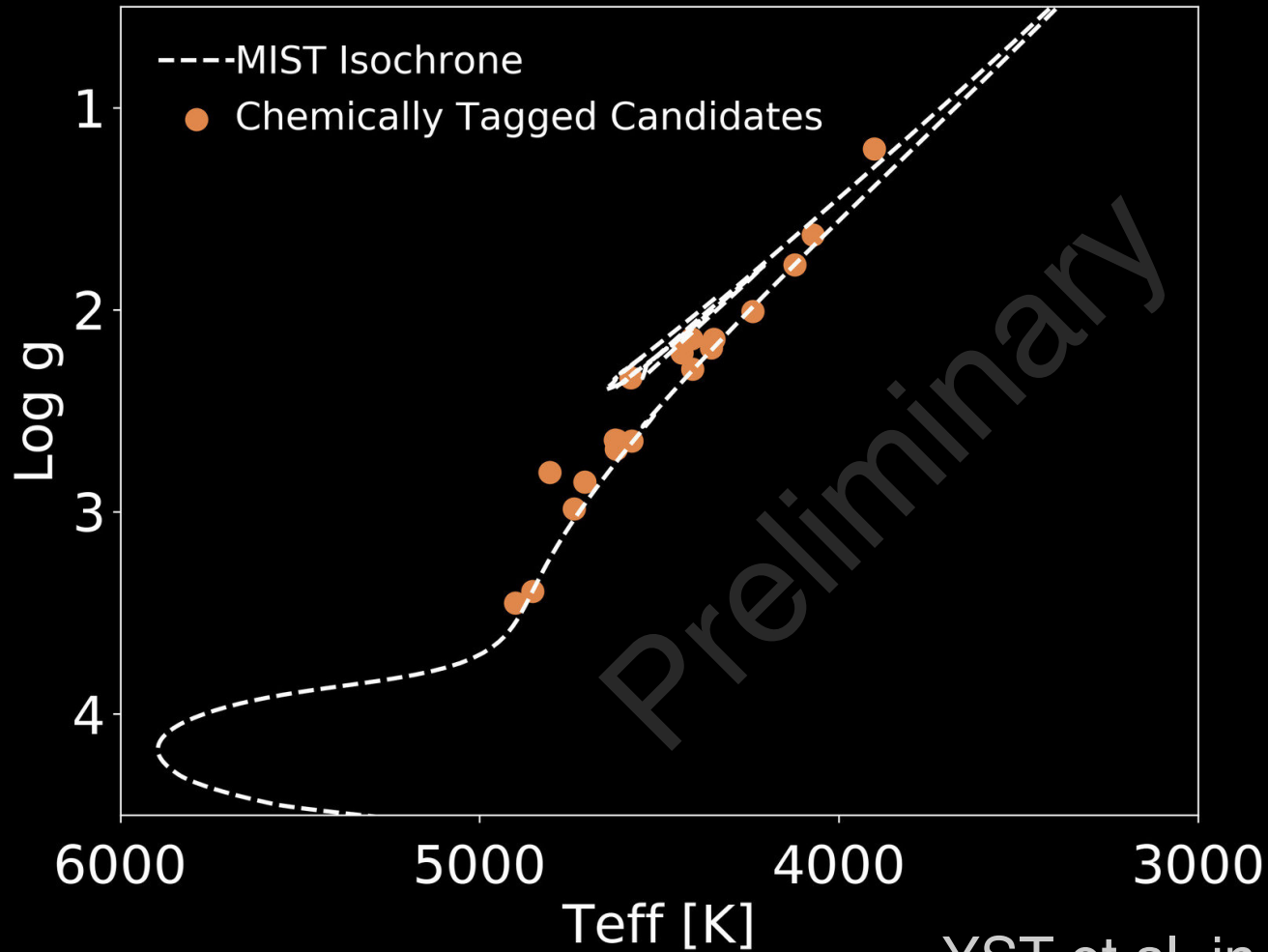
These candidates are spatially dispersed

Not an open cluster



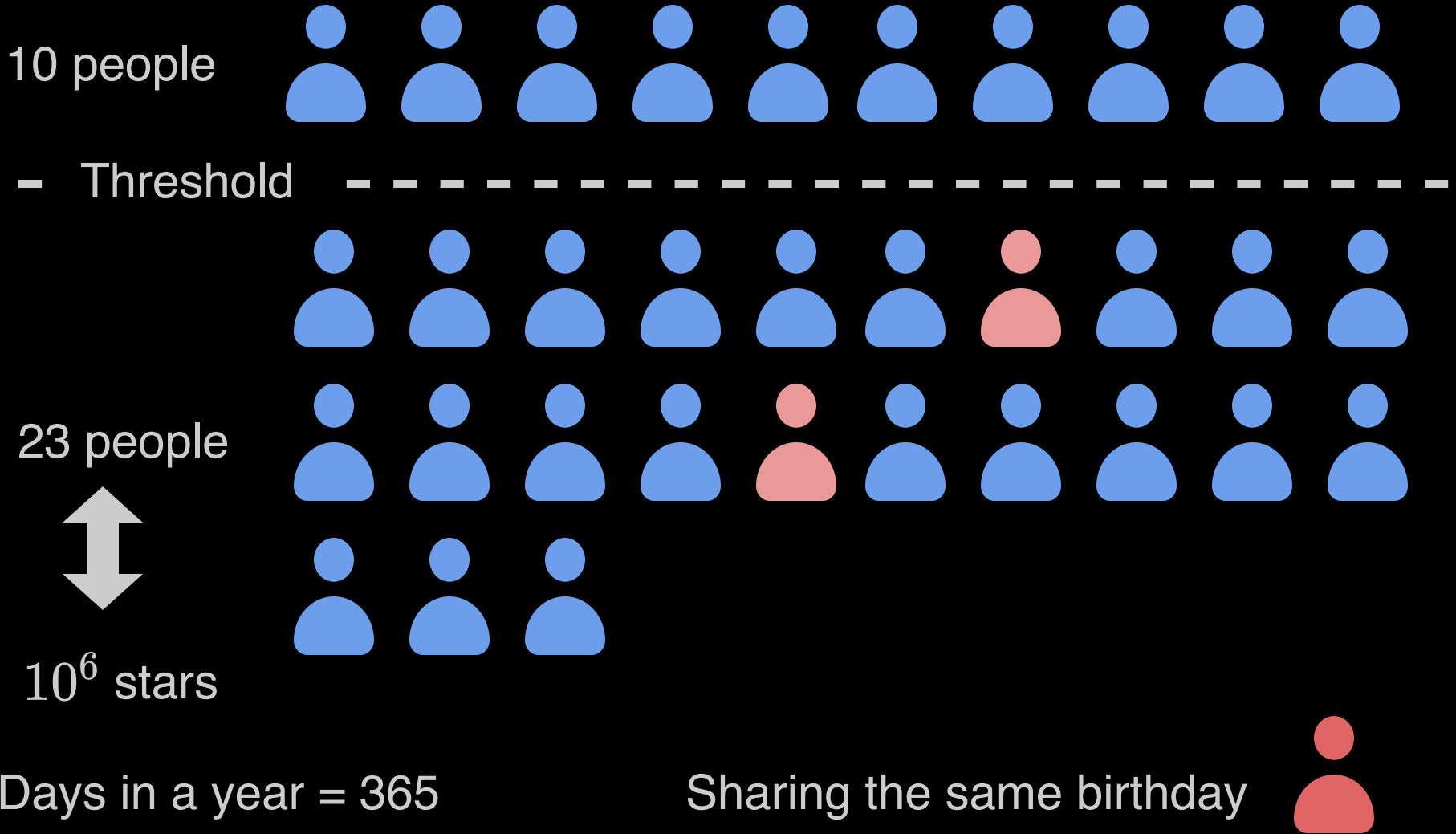
YST et al. in prep

The candidates agree with the isochrone perfectly

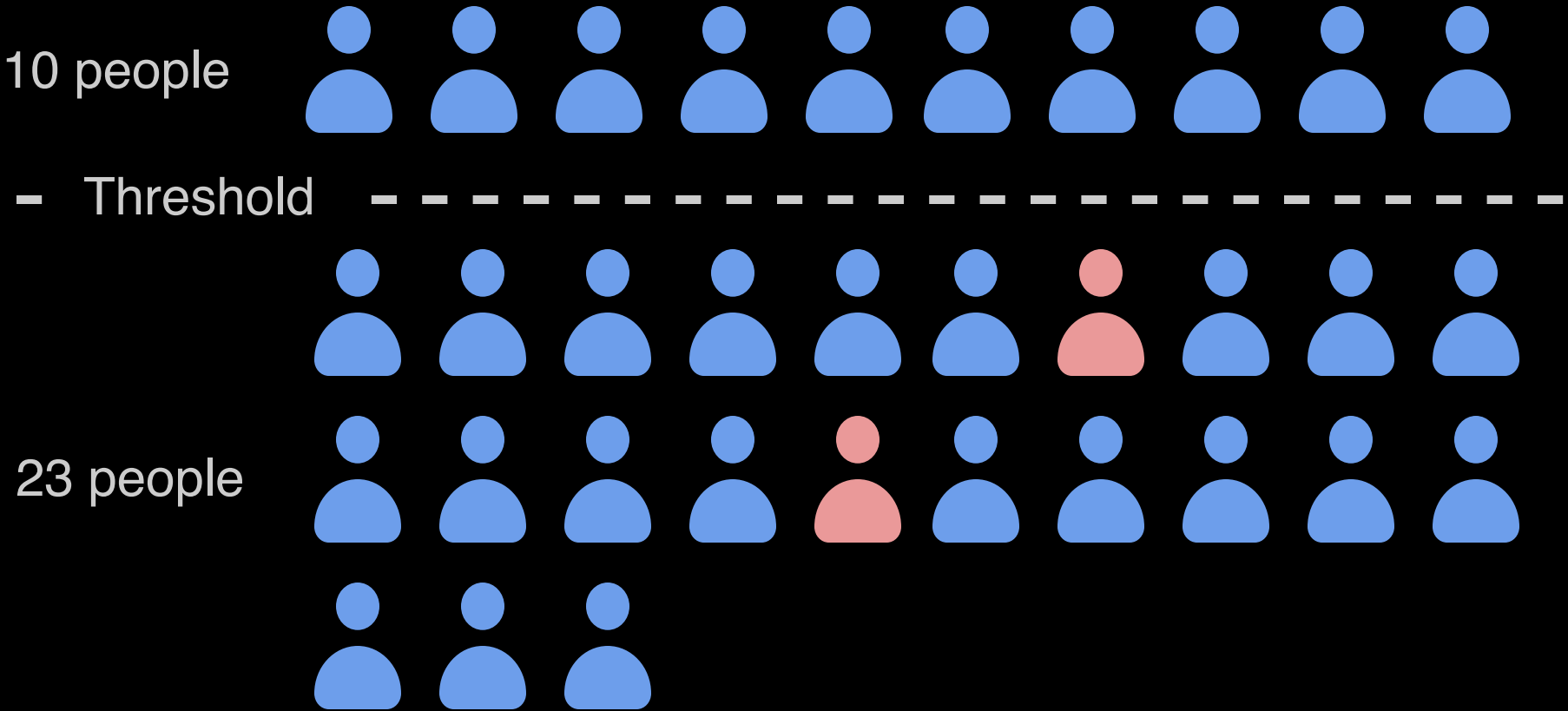


YST et al. in prep


The birthday paradox explained



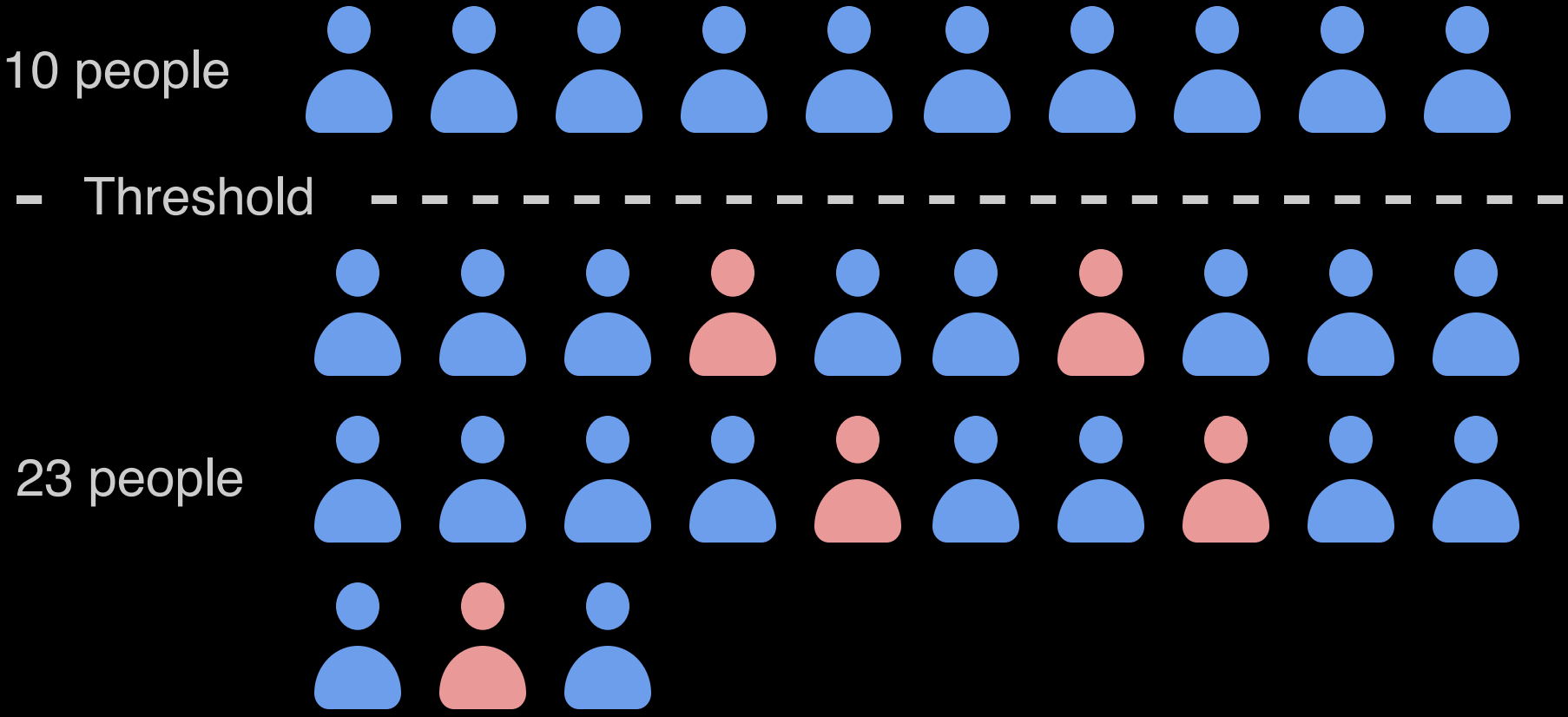
Infer the number of days through the number of pairs




Days in a year = 365

Sharing the same birthday 

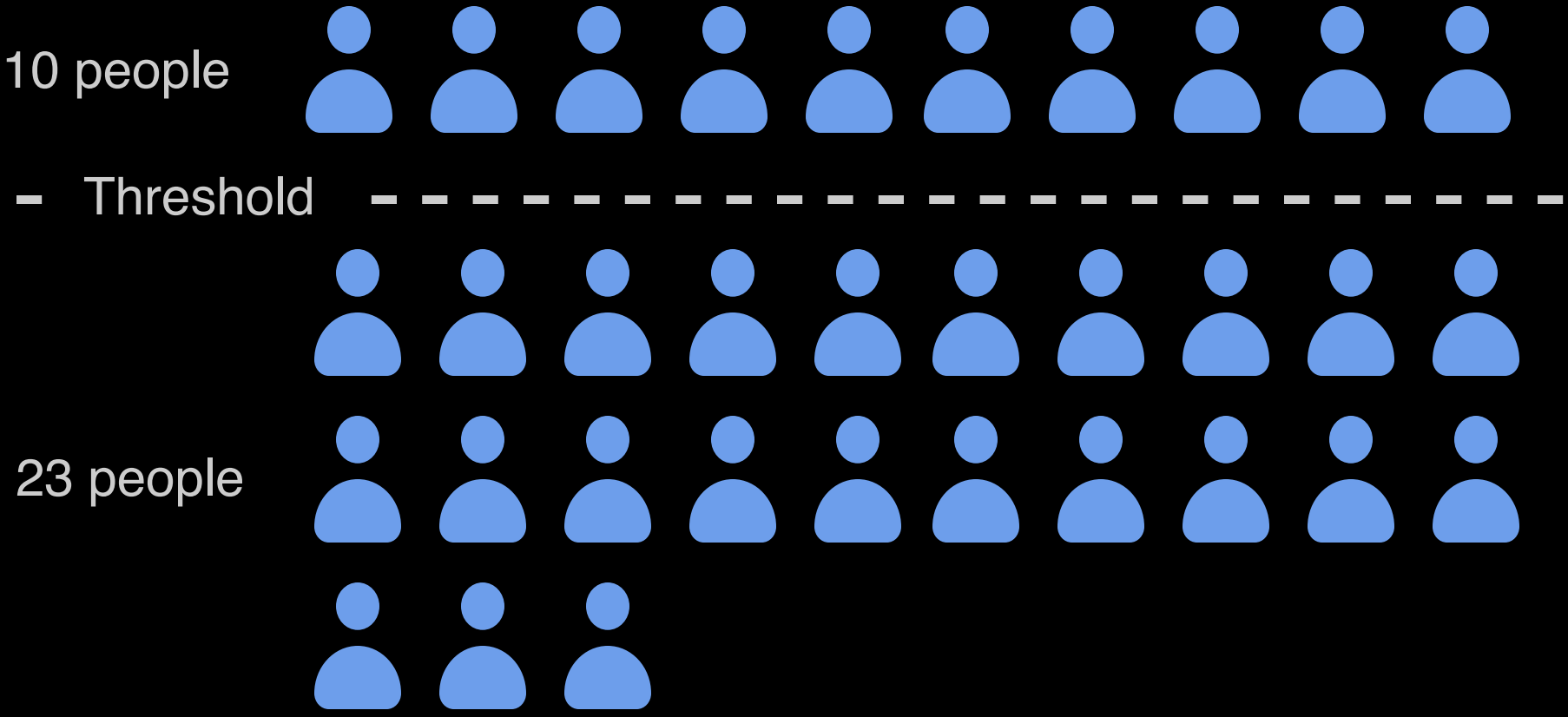
Infer the number of days through the number of pairs






Days in a year = 100 ↓

Sharing the same birthday 

Infer the number of days through the number of pairs

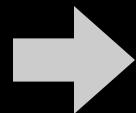
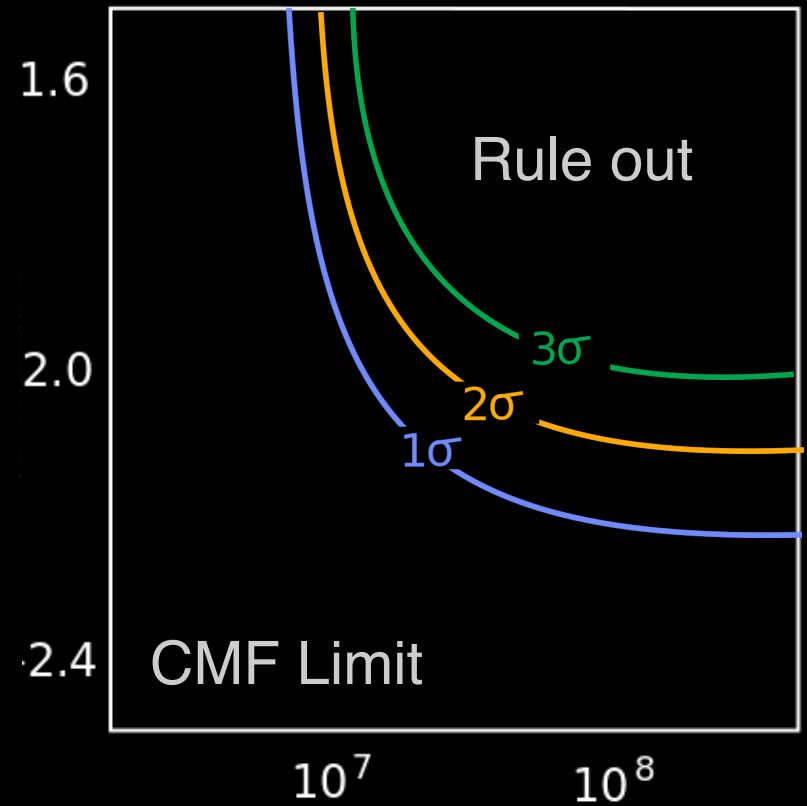
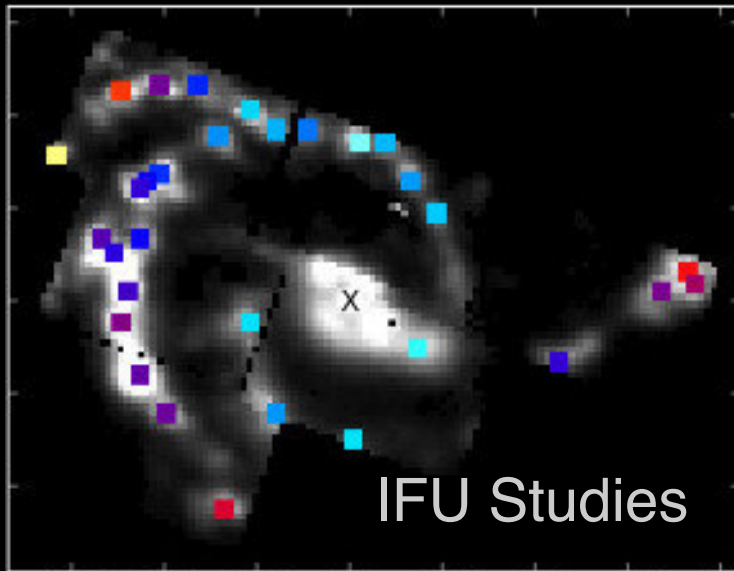


Days in a year = 900  

Sharing the same birthday 

Elemental-elemental correlation

Spatial-elemental correlation



How about kinematics (Gaia)

Exploit the kinematic-chemical correlation



Gaia : 10^9 stars

Simulating orbits of
a billion stars

Harshil Kamdar (student), Conroy, YST

Including kinematics can "purify" contaminations of chemical clumps

