

Thoughts on a plastic scintillator tracker option

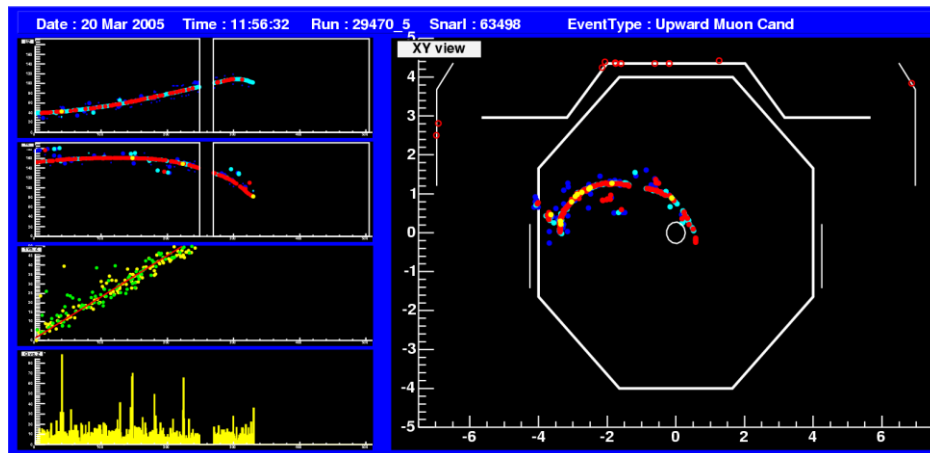
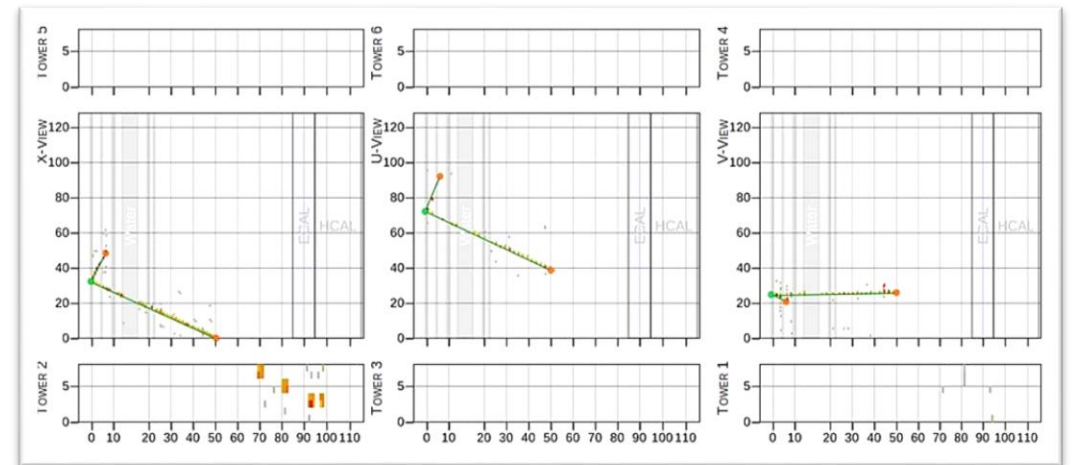
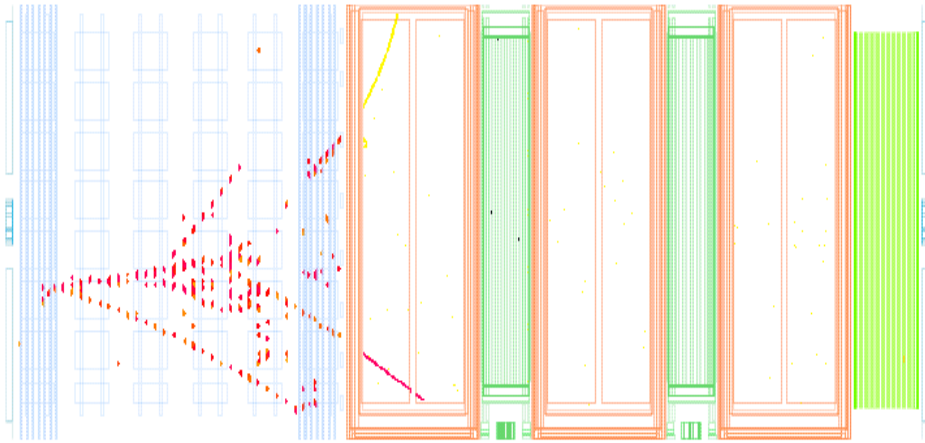
DUNE Near Detector Workshop

FNAL

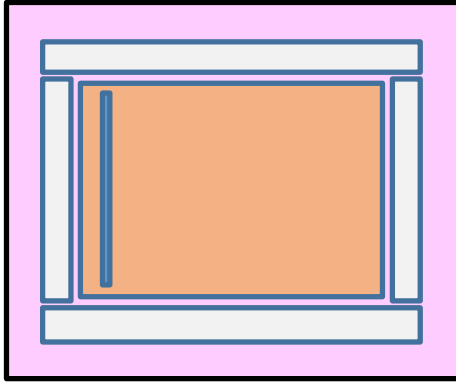
March 27-29, 2017

S. Manly

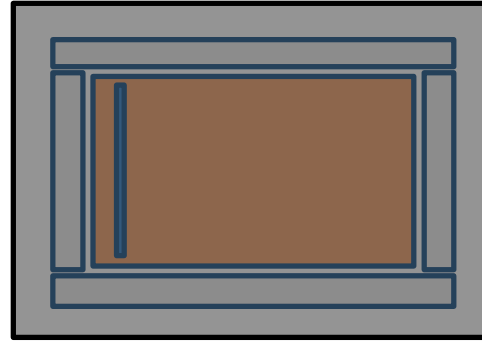
University of Rochester



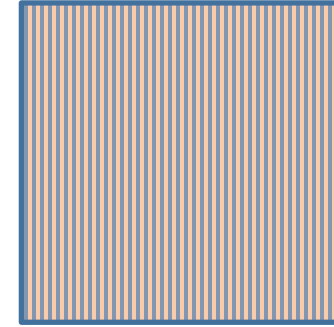
Some possibilities to consider



Scintillator tracker surrounded by ecal with plane of high pressure Ar gas tubes and a B-field

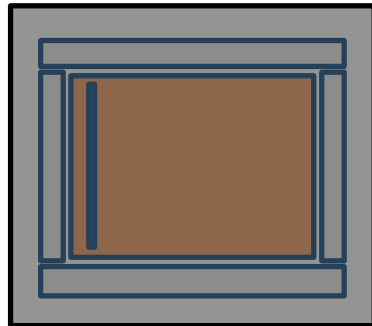


Unmagnetized black box tracker/ecal

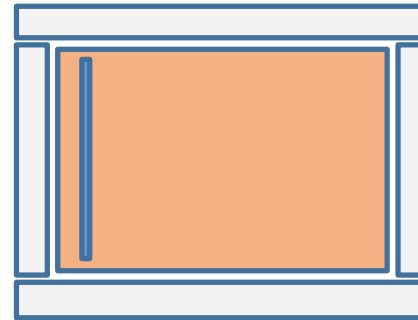


Magnetized muon spectrometer with active planes of scintillator readout

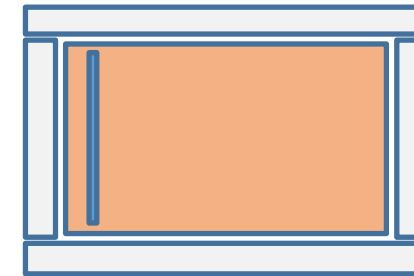
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Black box ND with Unmagnetized scintillator "outrigger(s)" off-axis to constrain beam direction and focusing errors

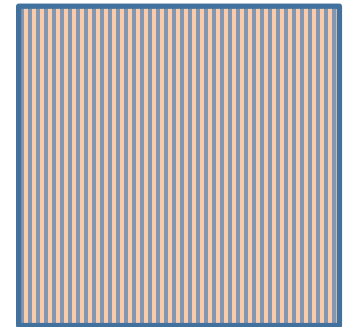


Unmagnetized scintillator tracker/ecal



Unmagnetized scintillator tracker/ecal

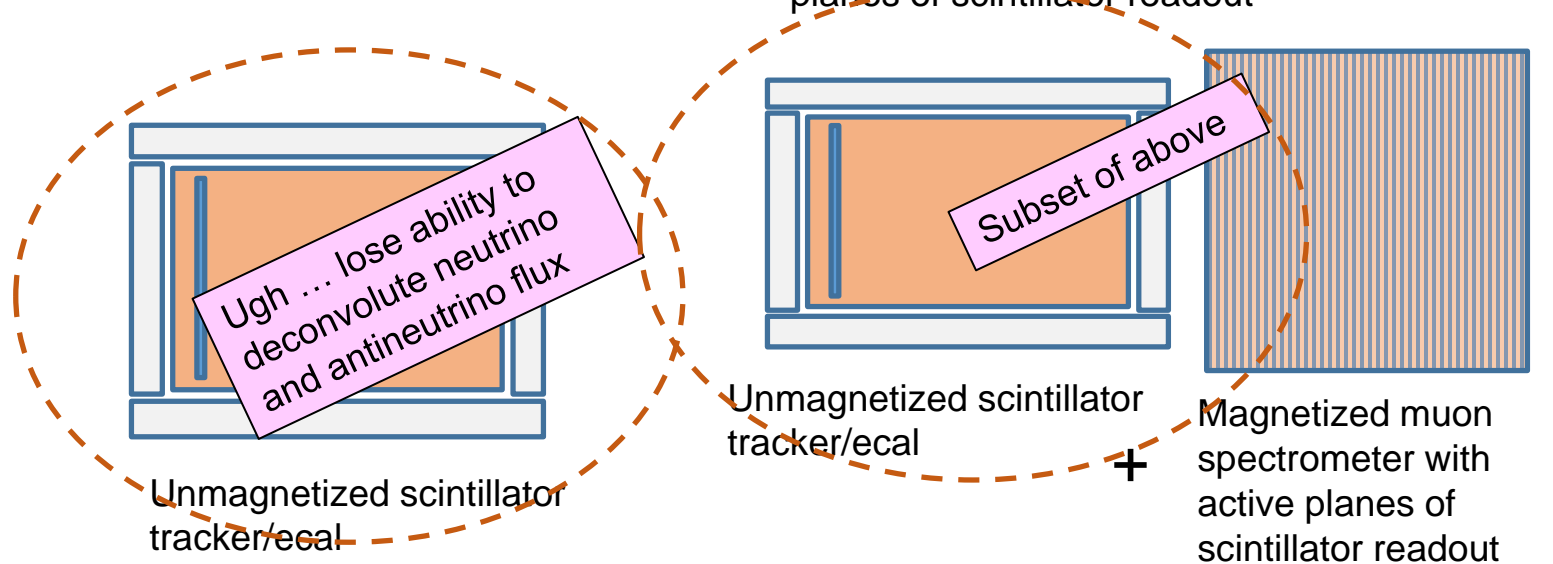
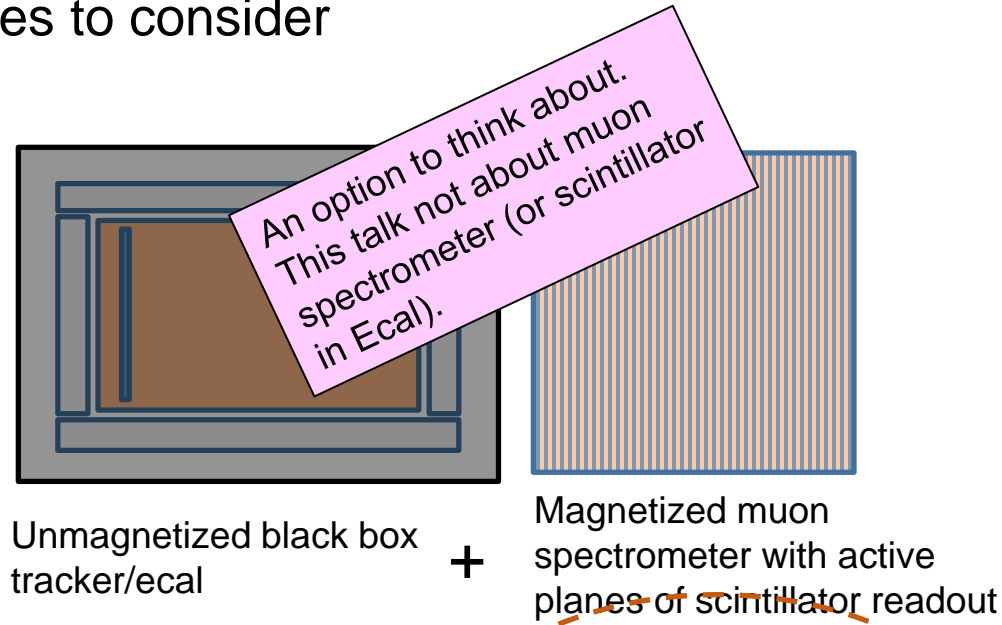
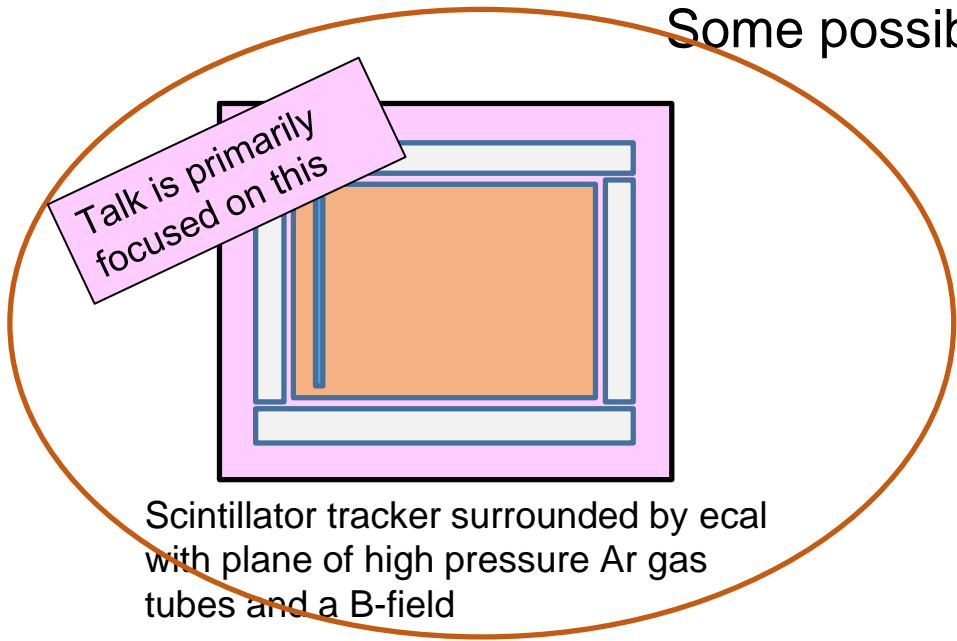
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Magnetized muon spectrometer with active planes of scintillator readout



Some possibilities to consider

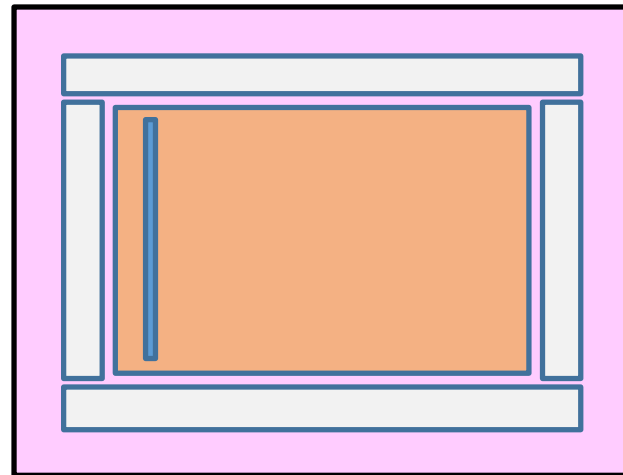


Black box ND with Unmagnetized scintillator "outrigger(s)" off-axis to constrain beam direction and focusing errors

What we need from ND

Flux measurement:
CCnumu (+other three types),
nu-e scat, low-nu, inverse
muon decay, maybe others.

**Detector reconstruction
systematics** (best if ND has Ar
target and is similar to FD)



Program to improve xsec errors and constraints
(want a versatile and powerful detector with good
vtx/tracking/energy resolution and ability to select
topologies. Want full angular coverage.)

Beam monitoring
(high stats, post-target, post-horn)

**DUNE is a precision
exploration experiment and
needs to be ready for surprises**

Systematics from nuclear effects
(Ar and perhaps a program of
studying diff. nuclear targets)

**Need ND that CAN be built
technically with the financial and
people resources we can pull
together**

Naïve grade expectations for scintillator tracker option

A-

Flux measurement:
CCnumu (+other three types)
nu-e scat, low-nu, inverse
muon decay, maybe others

Done by MINERvA
and T2K

**Detector reconstruction
systematics** (best if ND has Ar
target and is similar to FD)

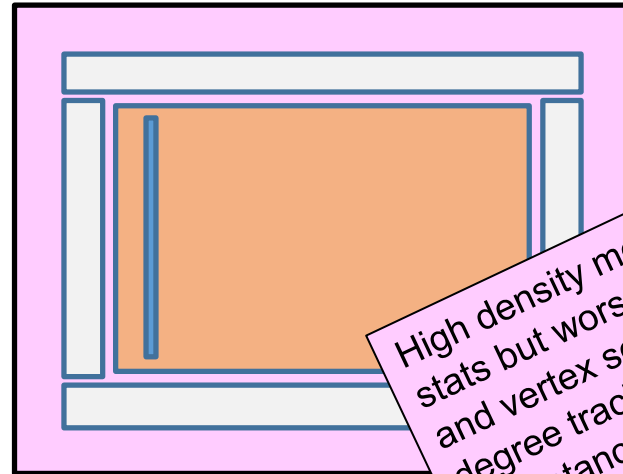
D

Can use layers of
Ar targets (similar
to FGT) but
detector not at all
like LArTPC

A

Beam monitoring
(high stats, post-target, post-horn)

Done by MINERvA
and T2K



High density means good
stats but worse resolution
and vertex tracking. 90
degree tracking
acceptance hole.

Program to improve xsec errors and constraints
(Want a versatile and powerful detector with good
vtx/tracking/energy resolution and ability to select
topologies. Want full angular coverage.)

B

**DUNE is an precision
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B

Systematics from nuclear effects
(Ar and perhaps a program of
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B

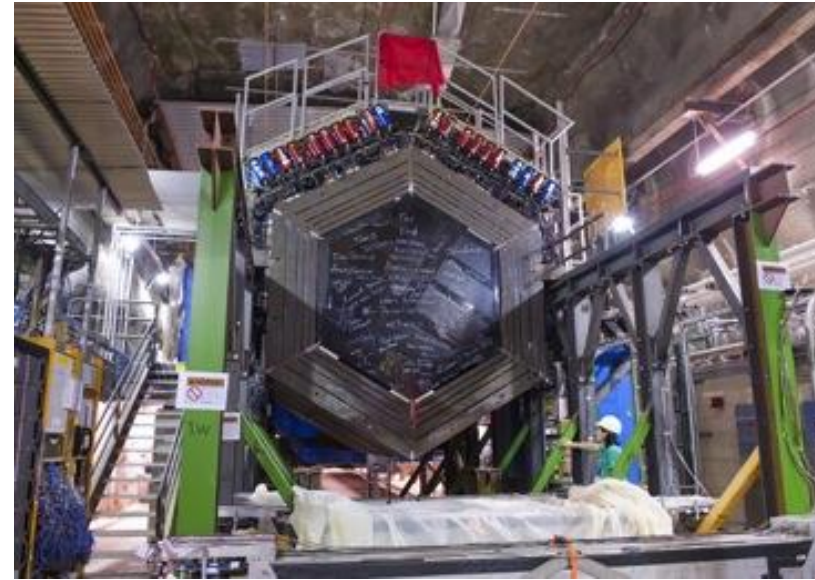
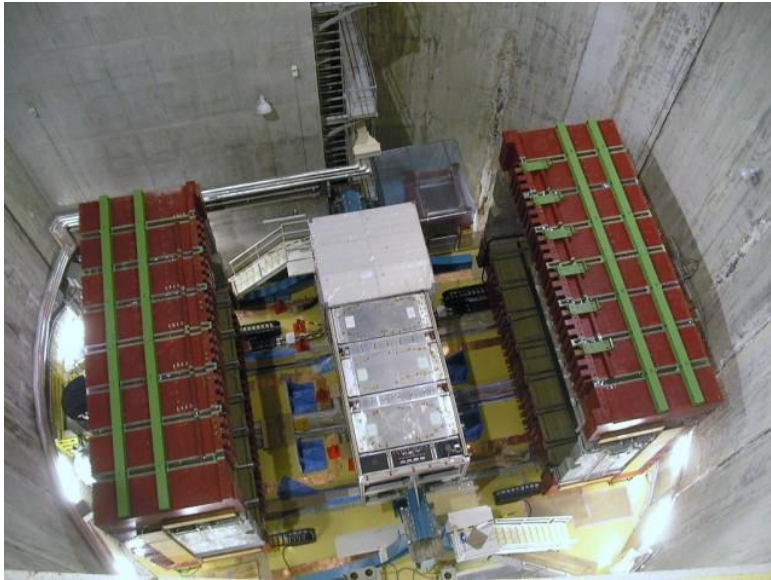
**Need ND that CAN be built
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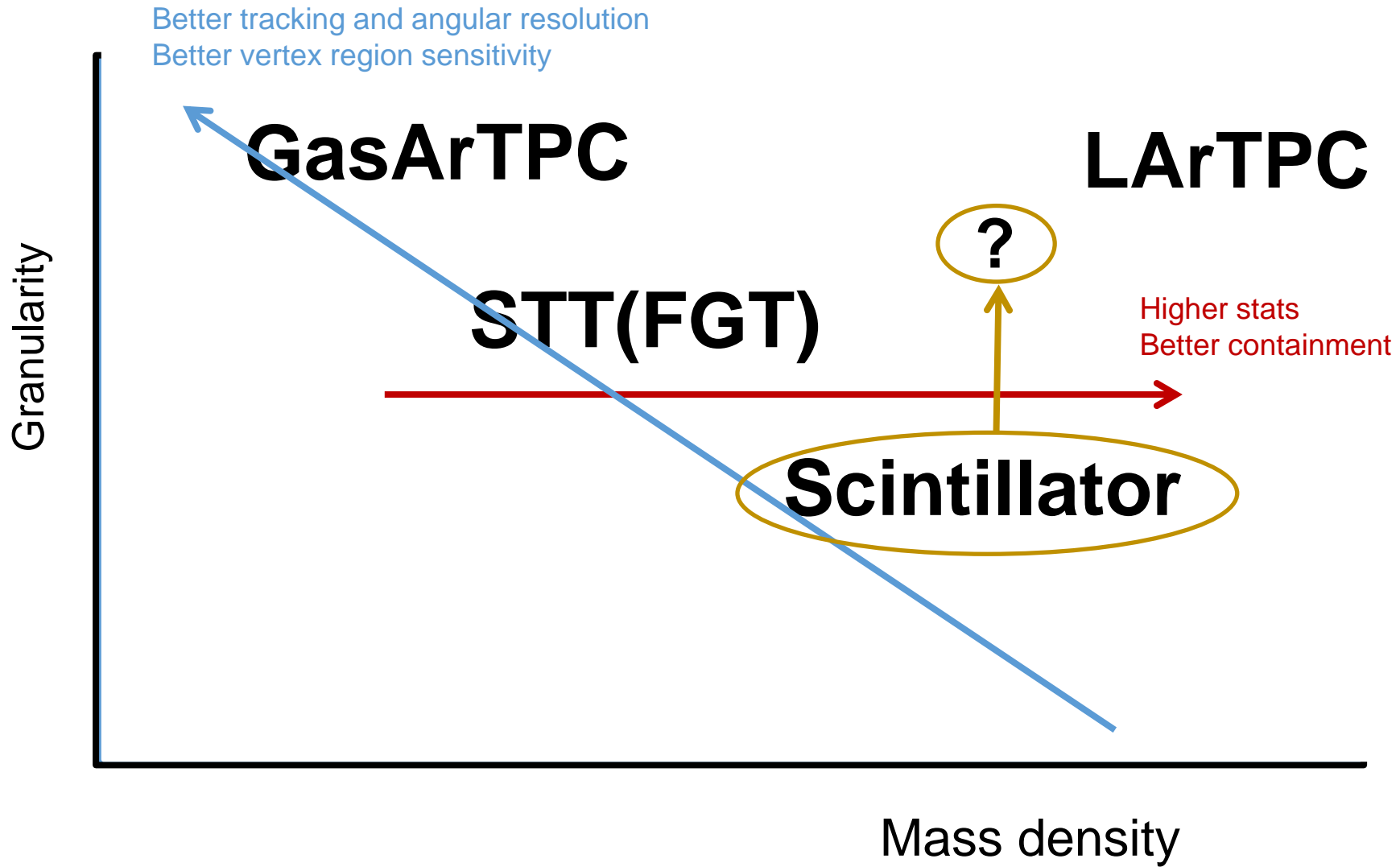
A

Similar scale
detectors built
before, tracker part
probably on "cheap"
end of options?

Can use layers of Ar
and other targets as
done by MINERvA
(similar to FGT)

Recent scintillator tracker existence proofs.
What have we learned?
What can we do better (and where not)?



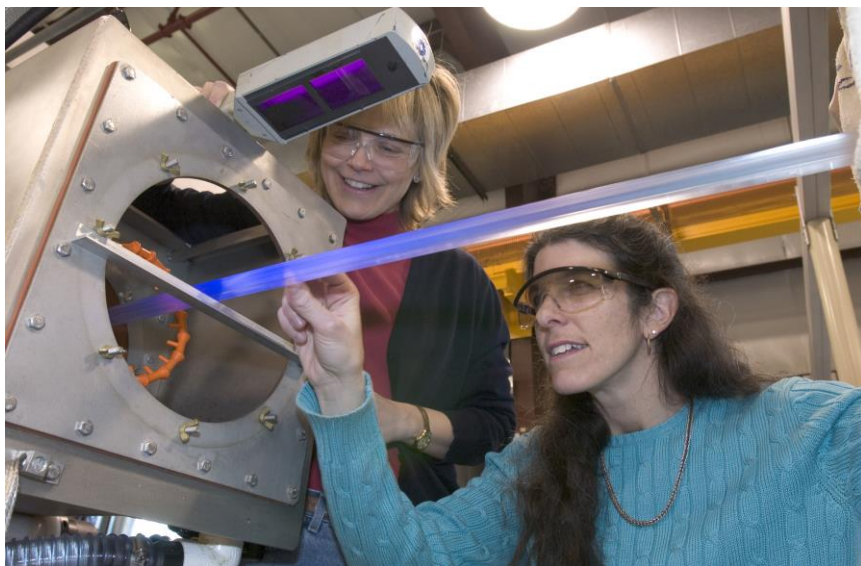
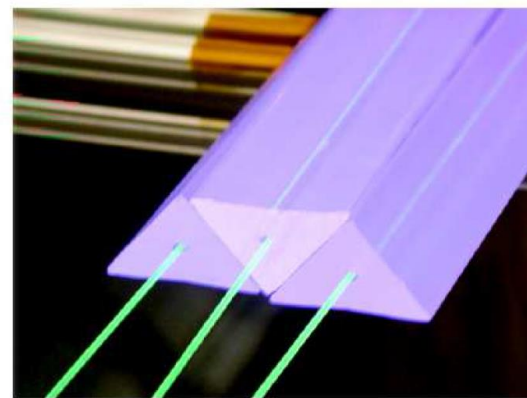


MINERvA (and T2K P0D) scintillator

1.7 cm



3.3 cm



Extrusion Line Facility at Fermilab

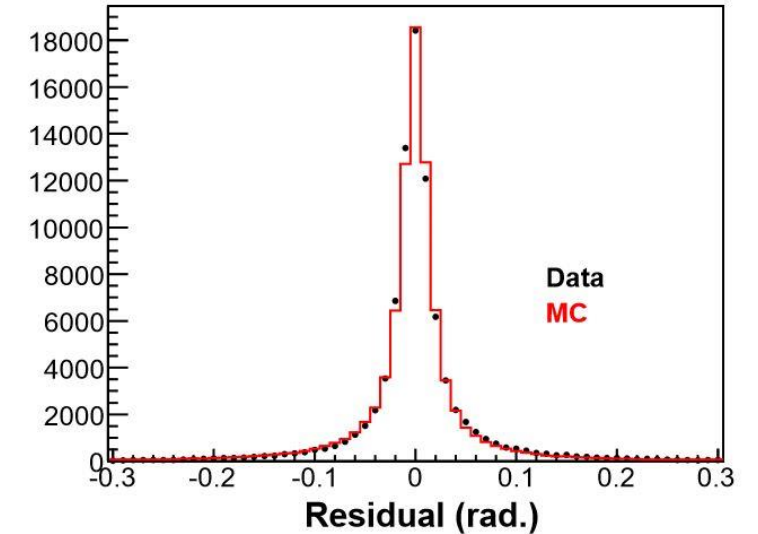
Anna Pla-dalmu: Probably can make extrusions down to 1 cm.
Have done 1 cm² xsec with 2.5-3 mm hole before.

For a few plots in this talk, assume triangles with
1.5 cm base and 0.77 cm height

Charged track angular resolution performance

$$\mathcal{G}_{rms} \Big|_{multiscatt} = \frac{0.015}{p} \sqrt{\frac{t}{X_o}}$$

Option	~thickness	Xo	Θ _{rms} @ 1 GeV/c
Minerva-like	7 cm (4 hits)	40 cm	6.3 mrad
Scint. With smaller strips	3.1 cm (4 hits)	40 cm	4.2 mrad
LArTPC	1 cm?	14 cm	4 mrad
STT	16 cm (4x and 4y hits)	5.5 m	2.5 mrad
	8 cm (~2x and ~2y hit)		1.3 mrad
GasArTPC	1 cm?	12.6 m	0.04 mrad



MINERvA angular resolution from rock muon study (from detector NIM)

Charged track dp/p performance

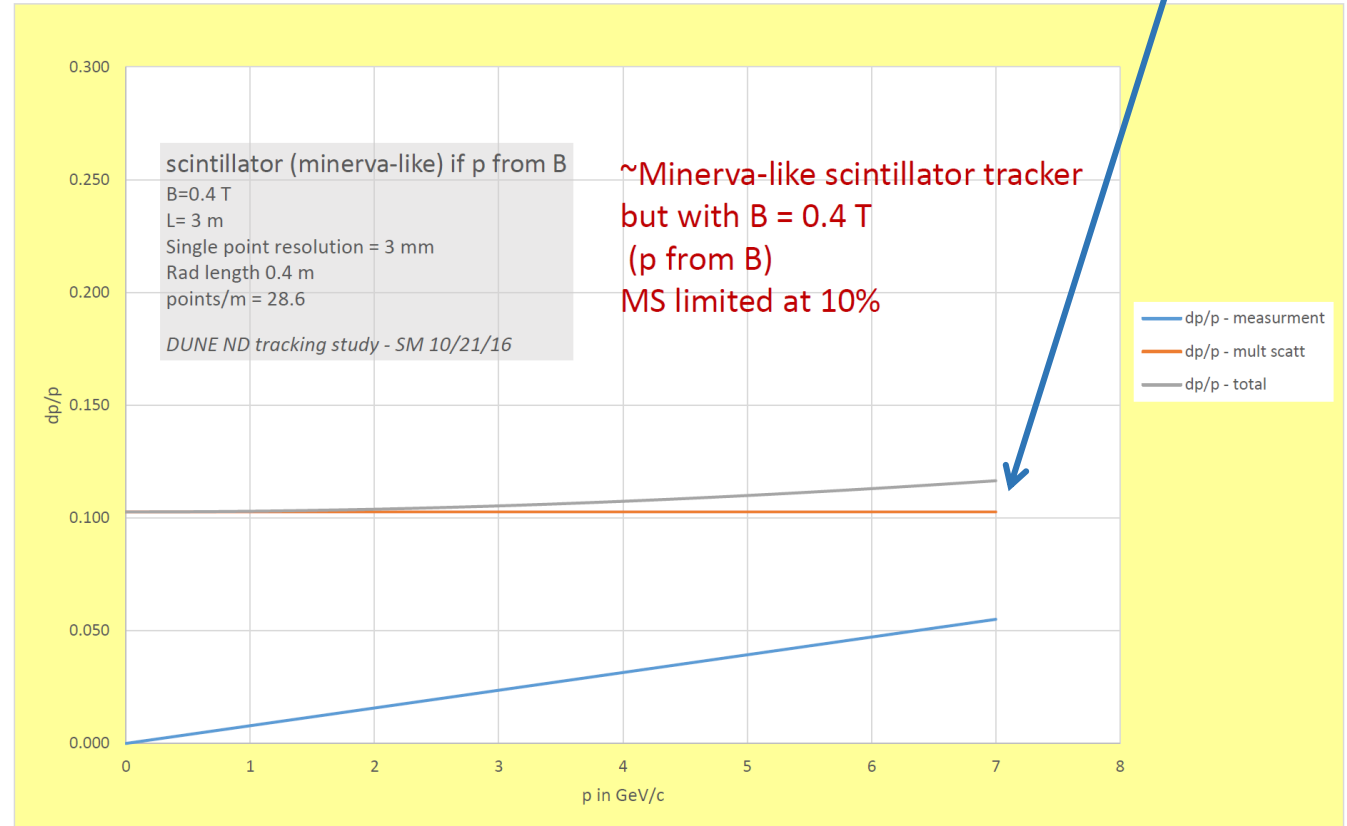
$$\left(\frac{\delta p}{p}\right)^2 = \left(\frac{\sigma_s p}{0.3BL^2} \sqrt{\frac{720}{N+4}}\right)^2 + \left(0.045 \frac{1}{B\sqrt{LX_o}}\right)^2$$

Multiple scattering dominated. Expect to see relatively little improvement with smaller transverse xsec strips.

All options are multiple scattering limited

$$\frac{\delta p}{p}$$

FGT (straw tube tracker 0.4 T B)	~3%
HPLArTPC (0.4 T B)	~2%
LArTPC (1 T B)	~7%
MINERvA-ish with 0.4 T B	~10%



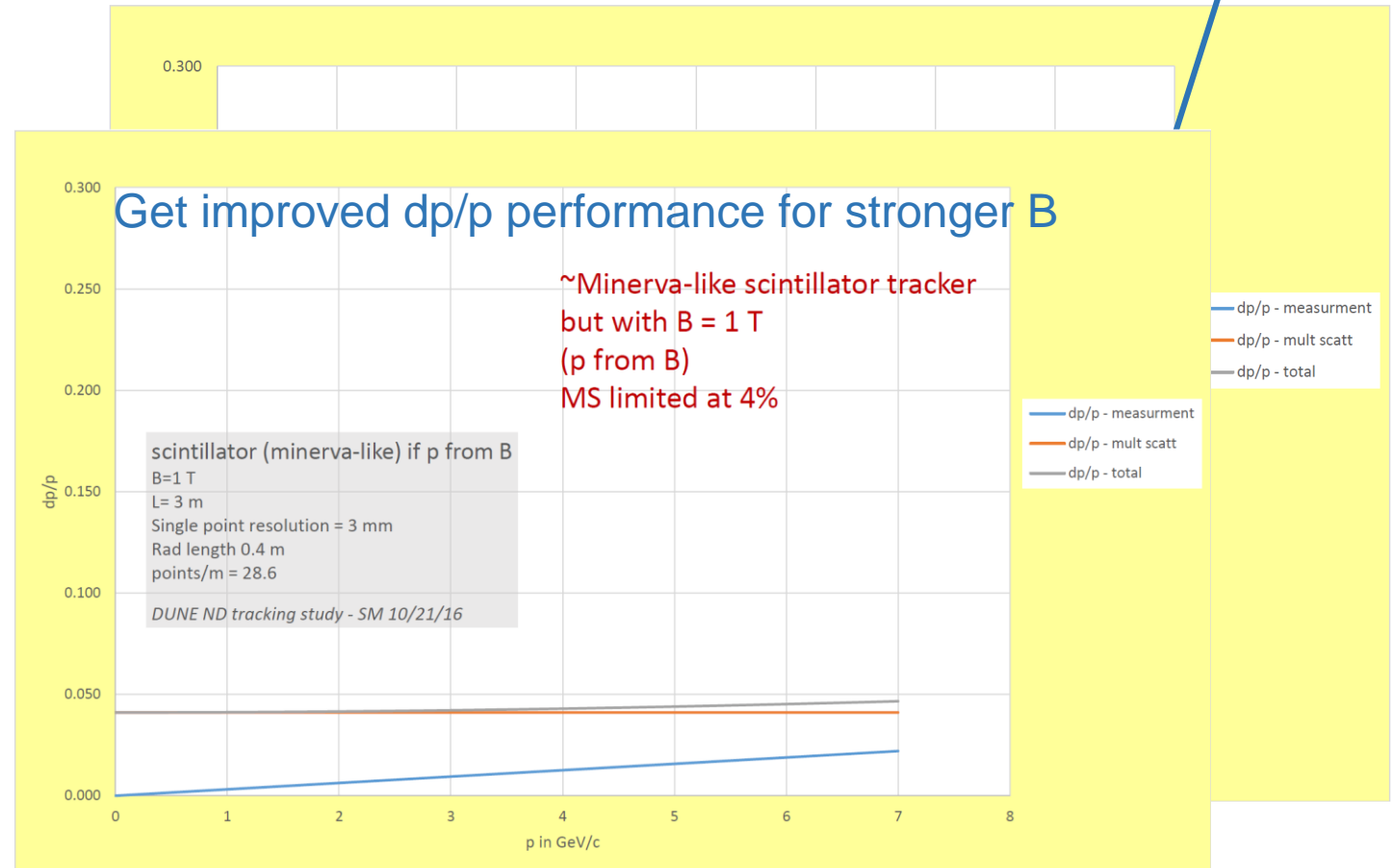
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MINERvA-ish with 0.4 T B	~10%
Scint. Tracker with 1 T B	~4%

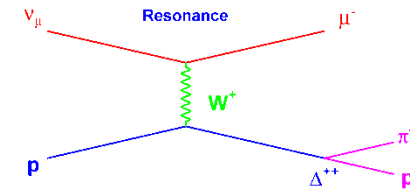
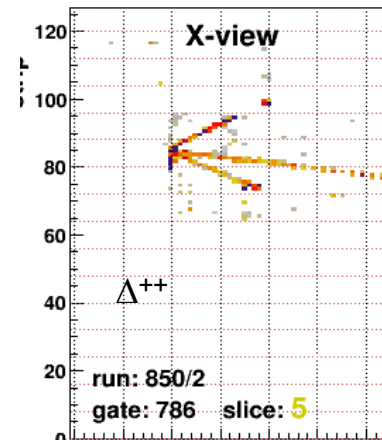
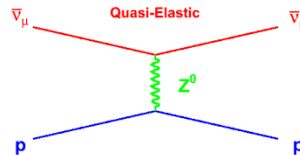
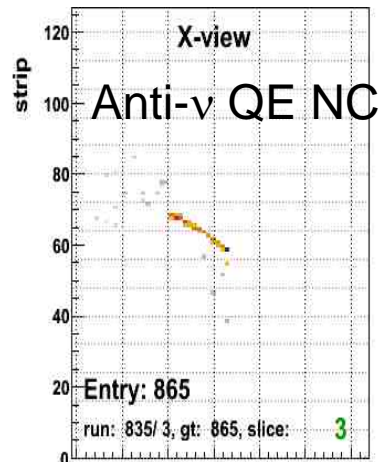
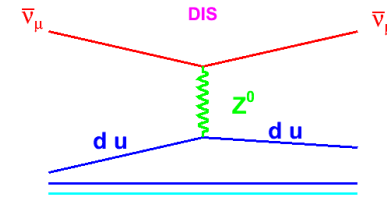
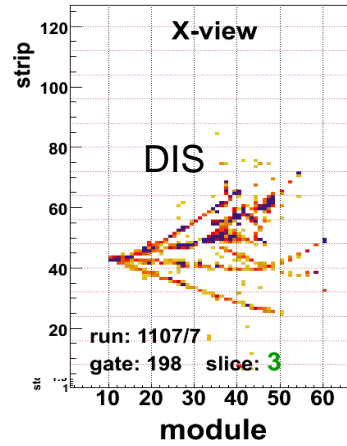
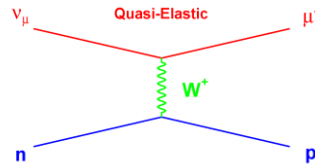
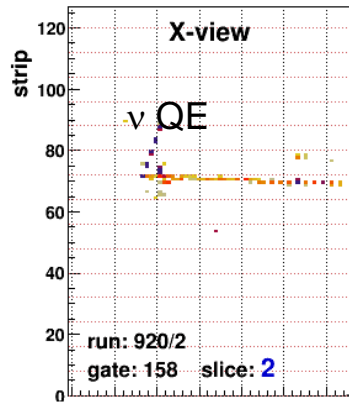


MINER ν A Events

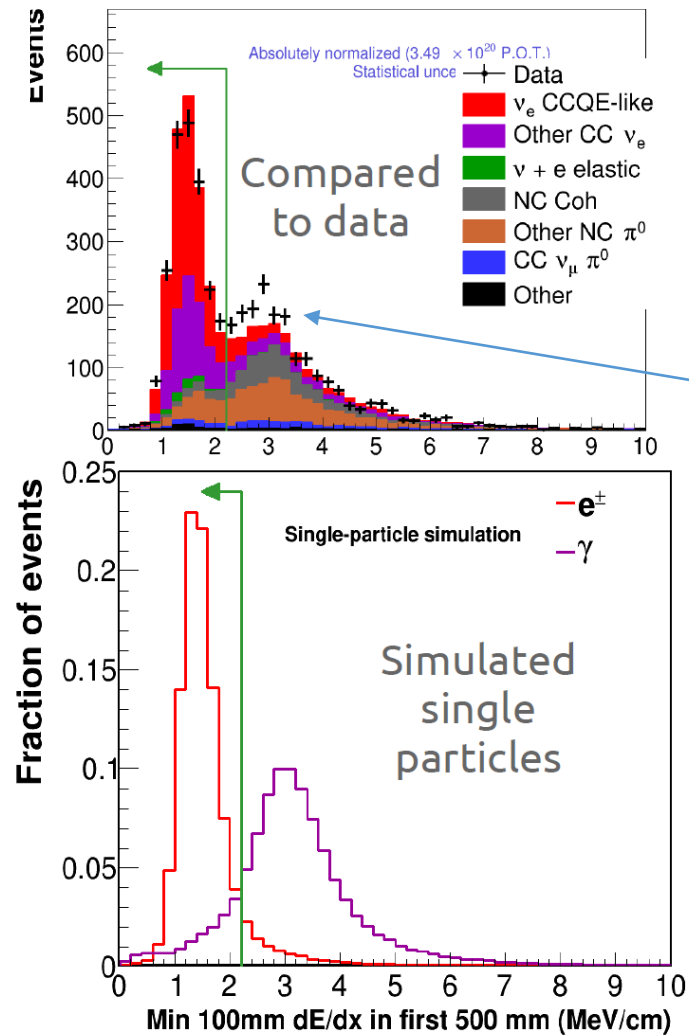
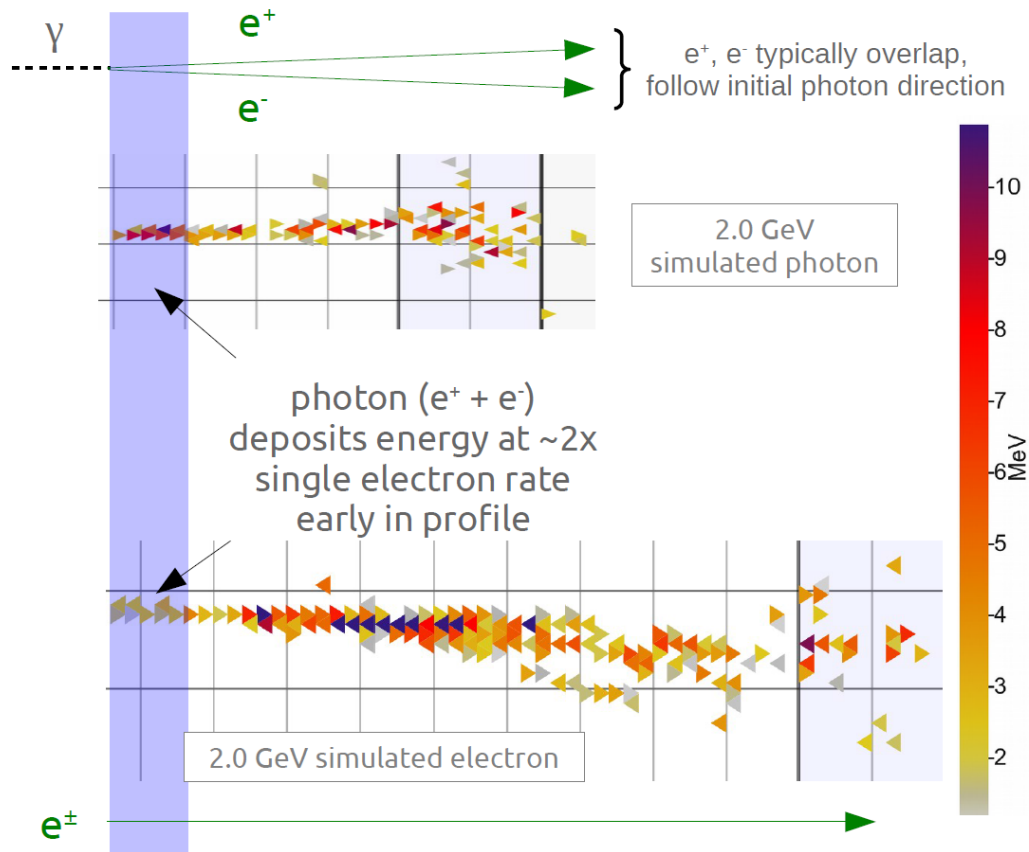
(X view)

Can see that scintillator tracker is able to select topologies.

Expect this/vertex sensitivity/two-track separation to improve with smaller strips



MINER ν A γ -e separation

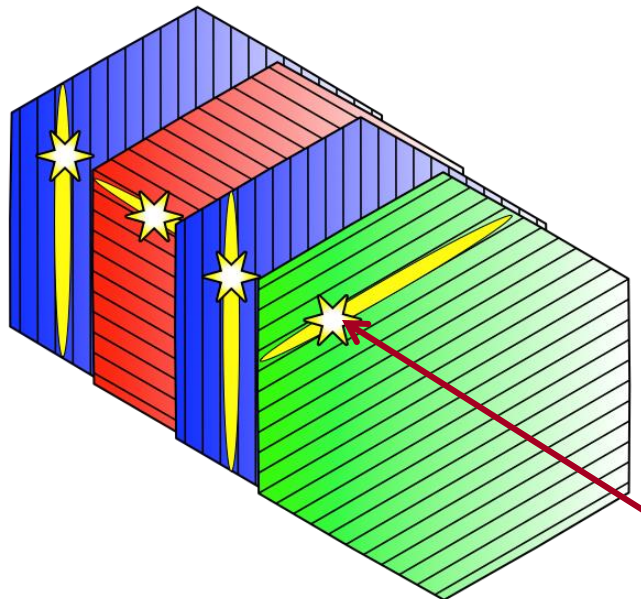
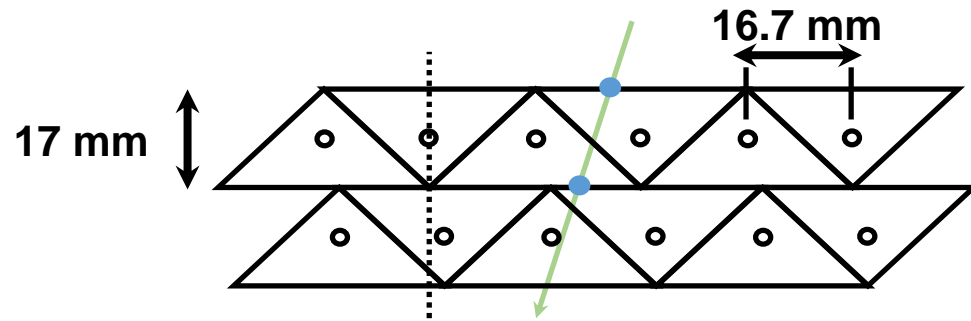


MC nicely validated with Michels and photons from pizeros

J. Wolcott

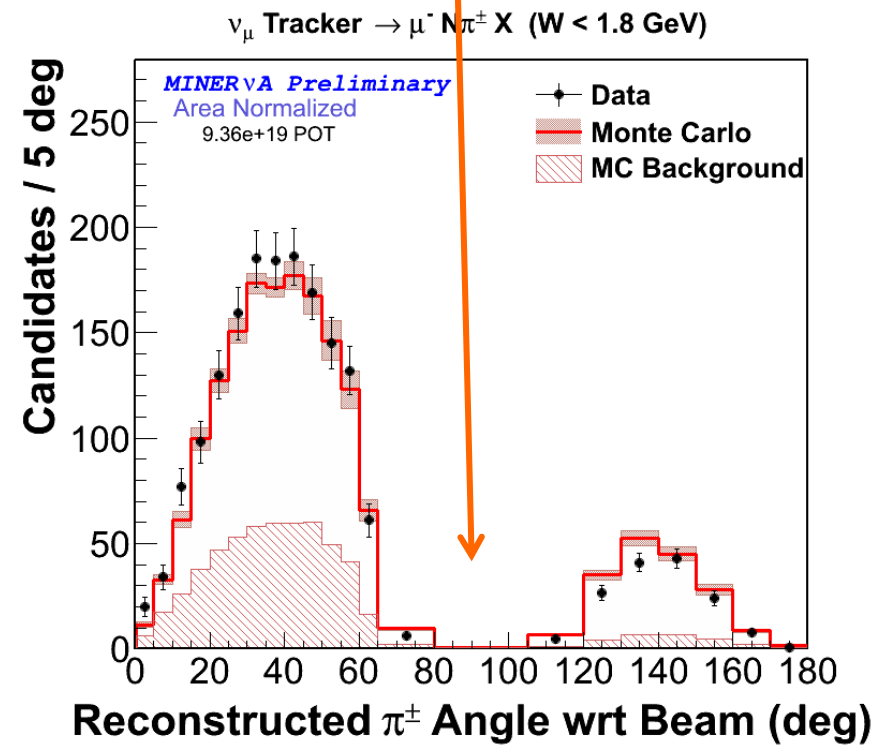
Might improve somewhat with more granularity

MINERvA planes

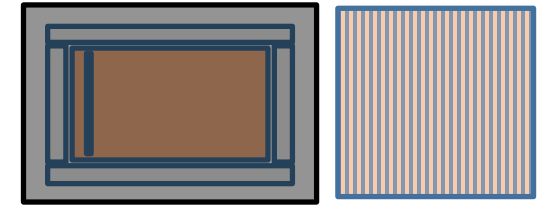


Hurts acceptance and identification by topology.
Less vertex region phase space covered.

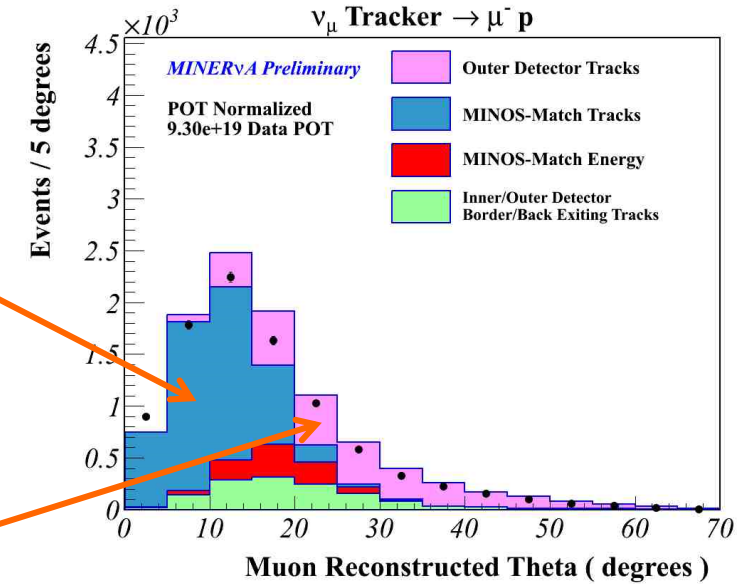
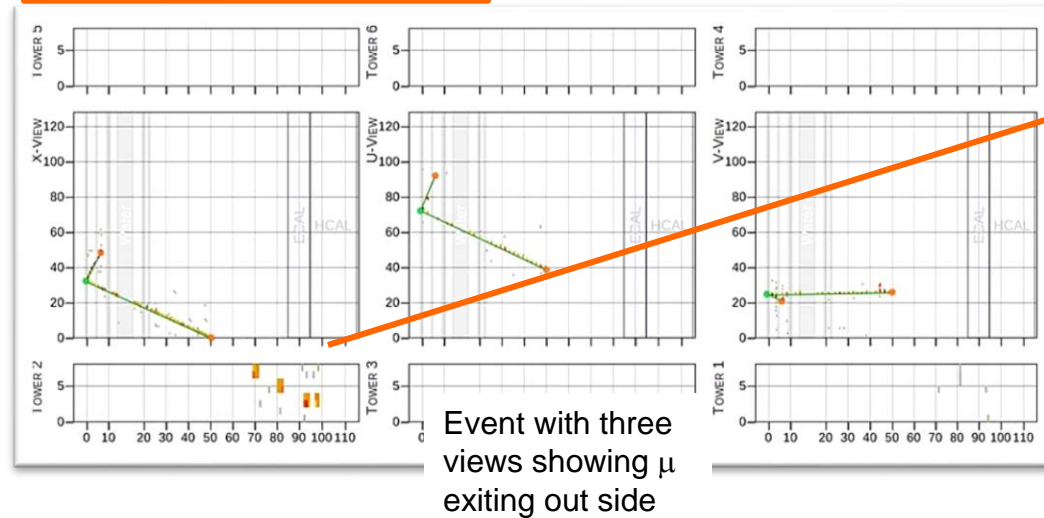
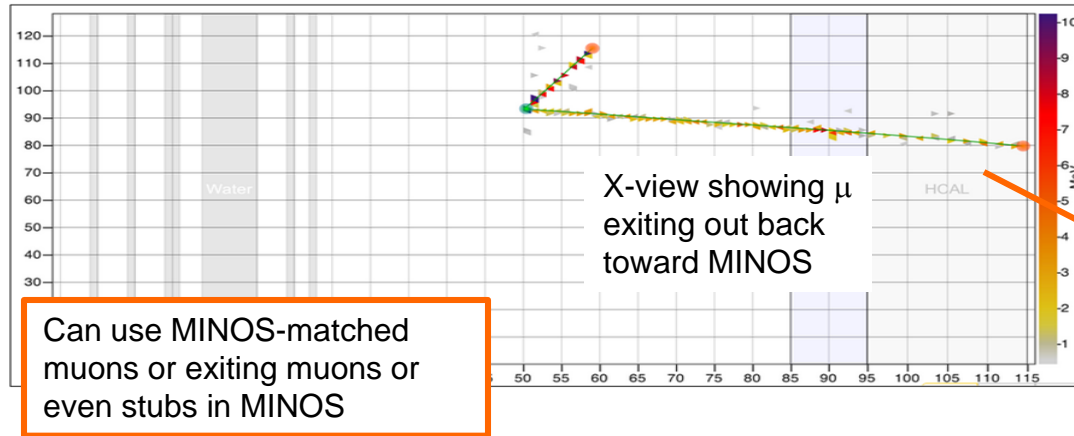
Construction of tracker gives a hole in reconstruction at 90°



For scintillator tracker (or other tracker) if used in non-magnetized target plus muon spectrometer configuration

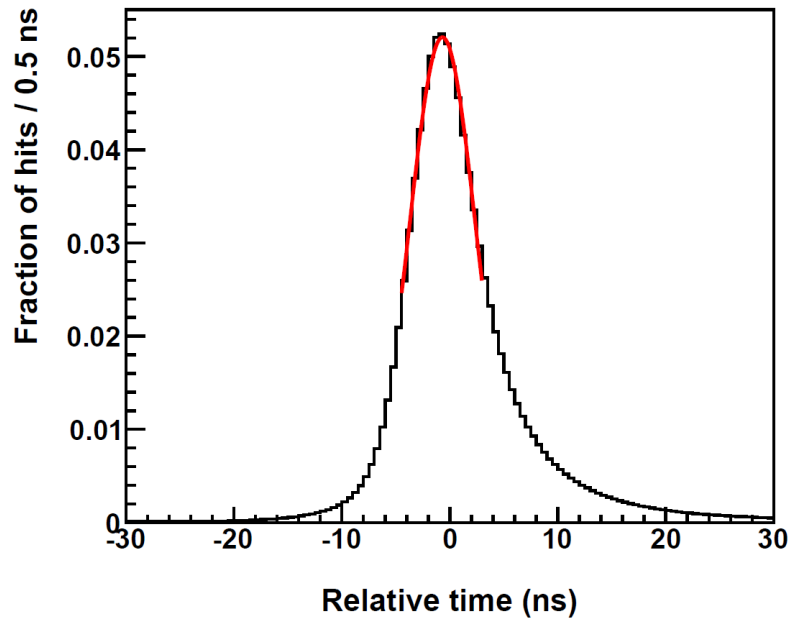


2-track CCQE analysis



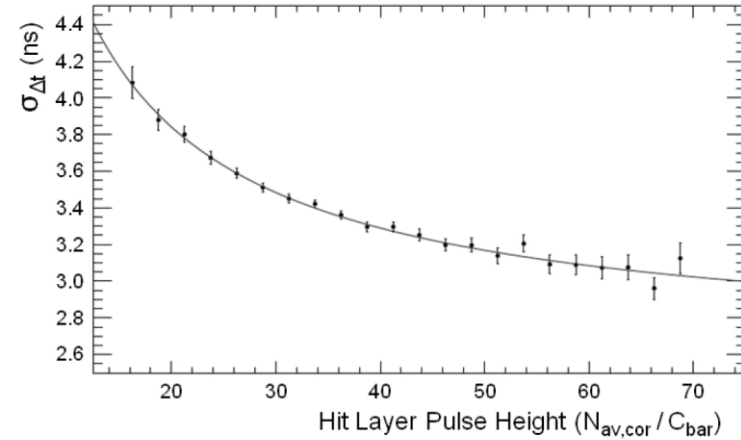
Limited angular acceptance for momentum-analyzed muons causes sculpting in angular and Q^2 acceptance.

Scintillator detectors have fast timing – few ns

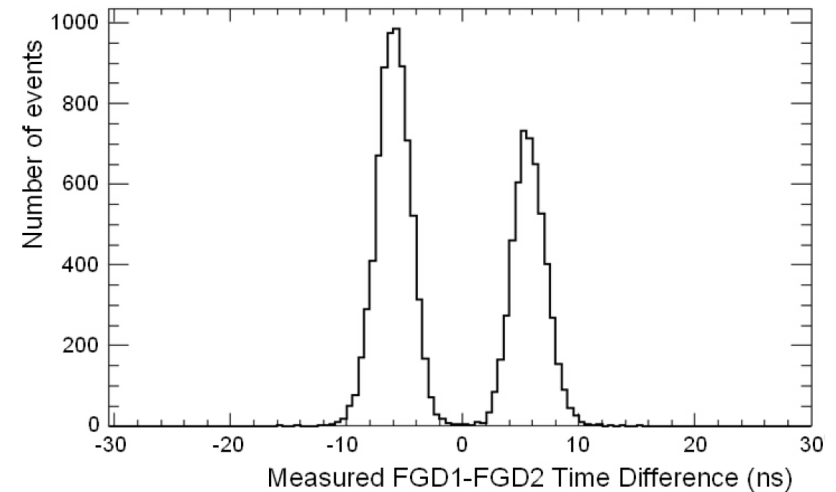


MINERvA calibrated hit time along rock muon tracks relative to the truncated mean calibrated hit time

From MINERvA NIM



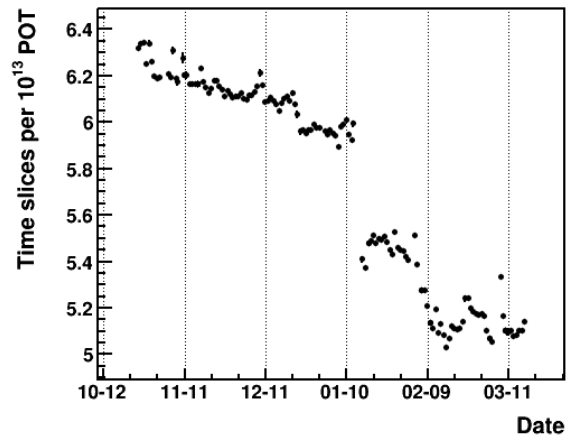
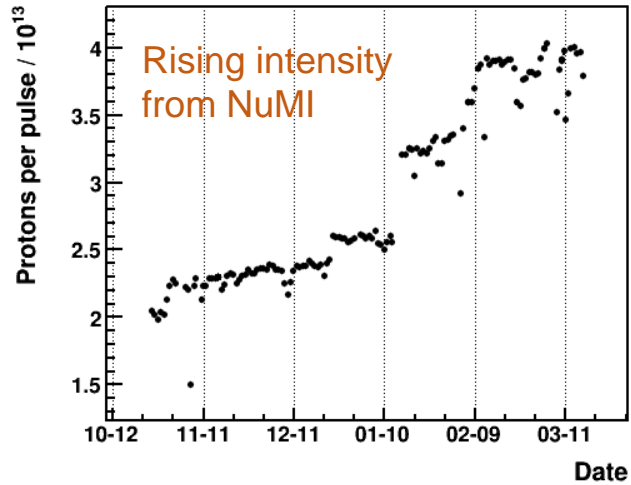
T2K FGD spread of timing residuals between a hit layer and a reference layer as function of hit layer charge



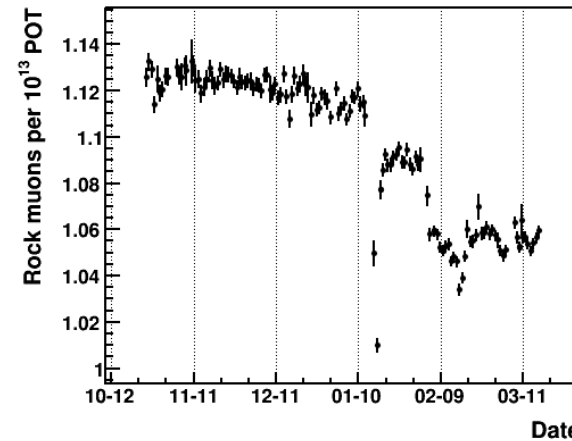
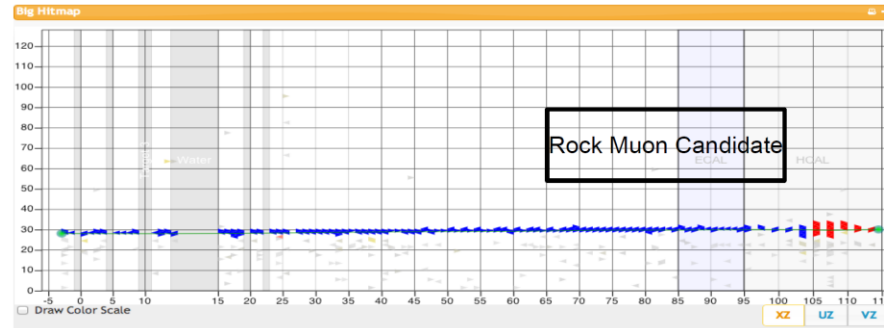
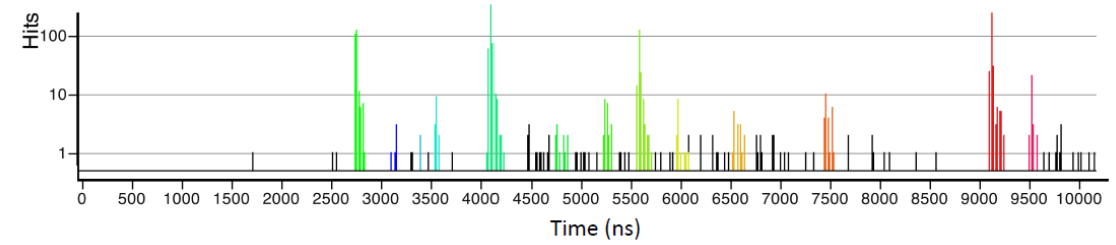
FGD1-FGD2 time difference for cosmics passing through both FGD detectors

From T2K ND280 NIM

MINERvA sees intensity effects

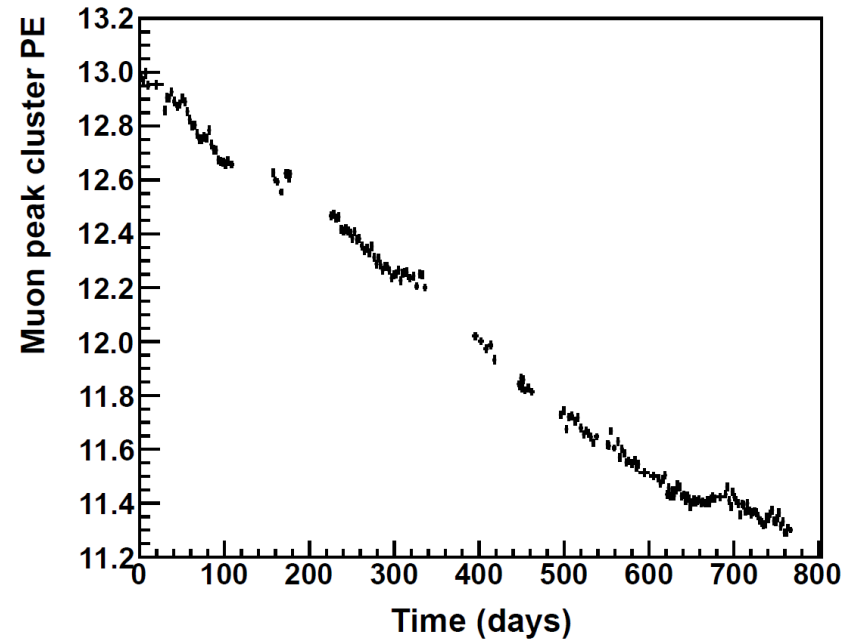


Number of temporally separated bits of activity in the detector drops – more activity means overlap increases within the time resolution of our algorithm



Number of reconstructed rock muons traversing the detector drops – extra activity and overlaps causes drop in track reconstruction efficiency

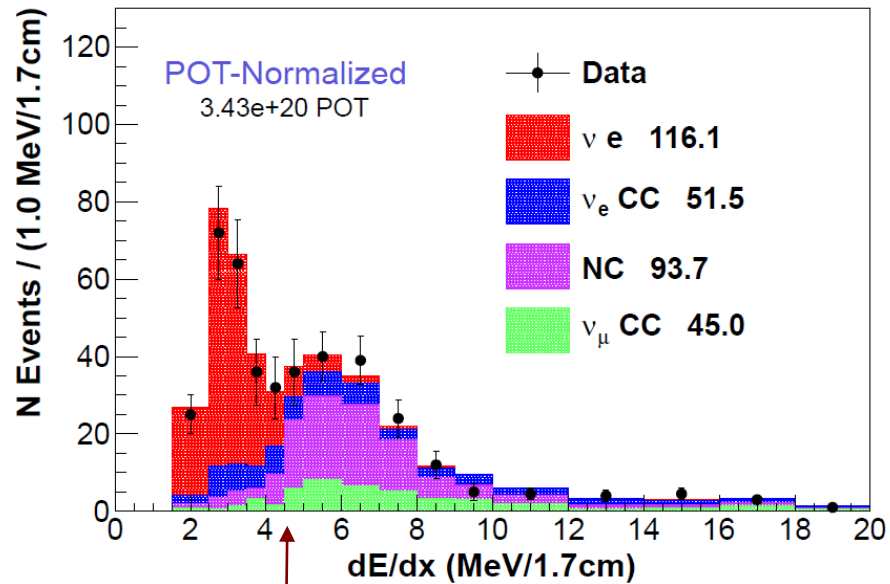
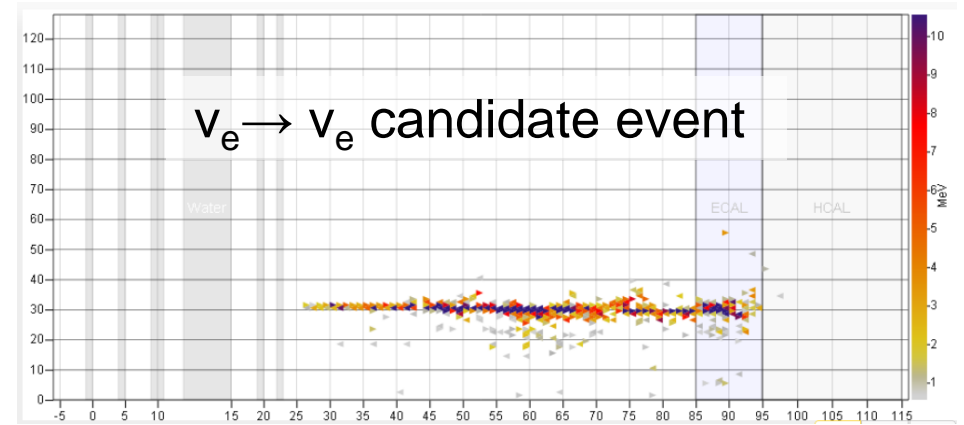
Scintillator ageing leads to slow light loss with time



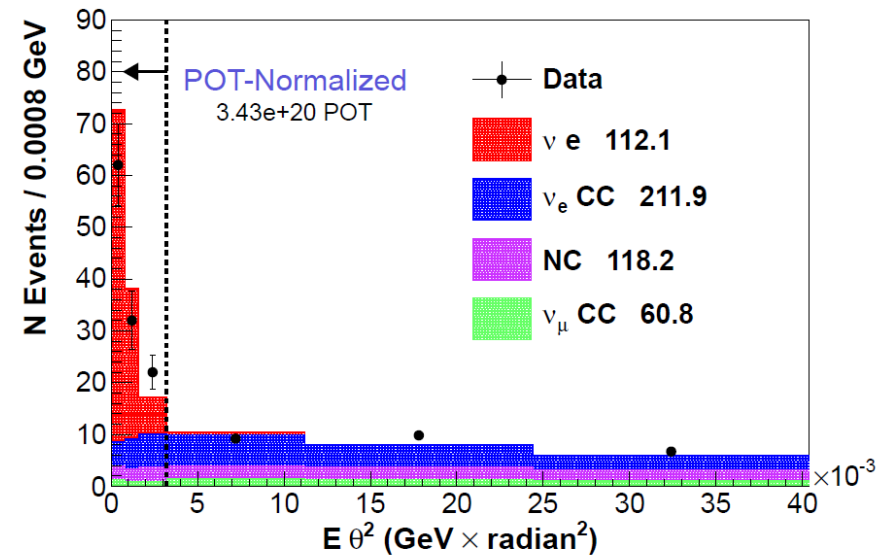
From MINERvA NIM

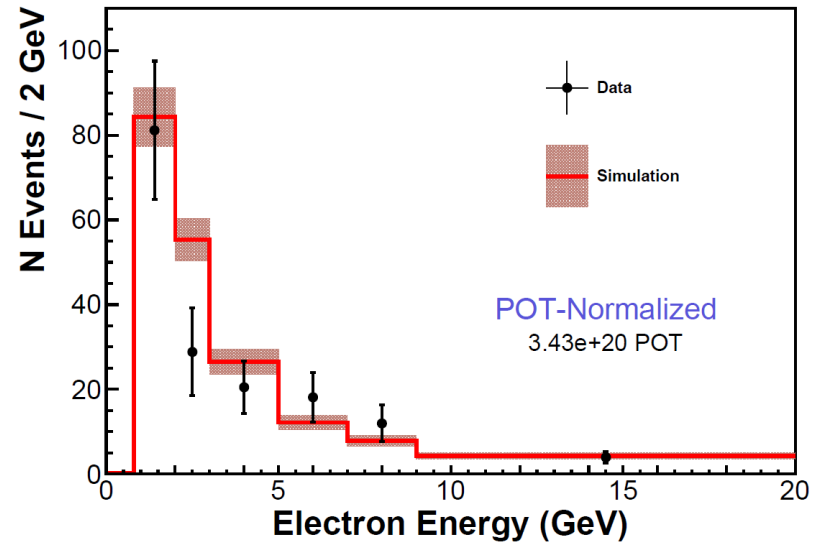
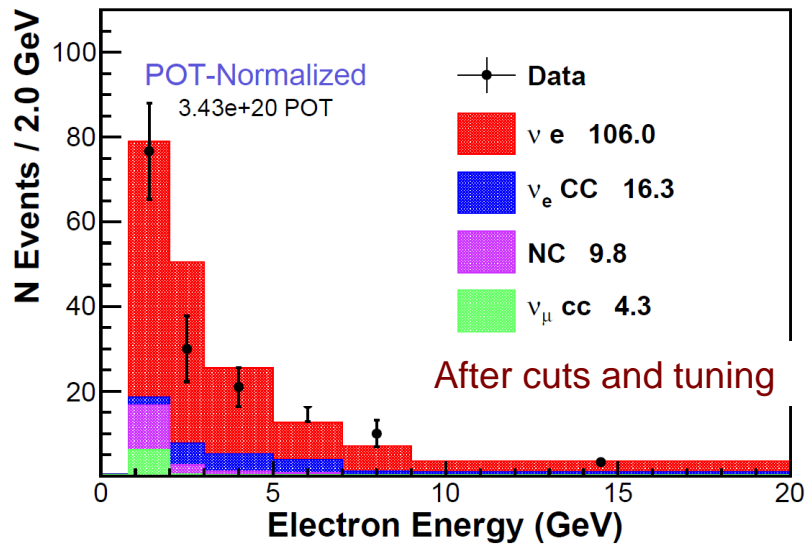
First MINERvA detector module was completed in 2006, detector completed in 2009-2010. T2K P0D and FGD detectors installed in 2009.

MINERvA and T2K experience can give insight on the limiting errors



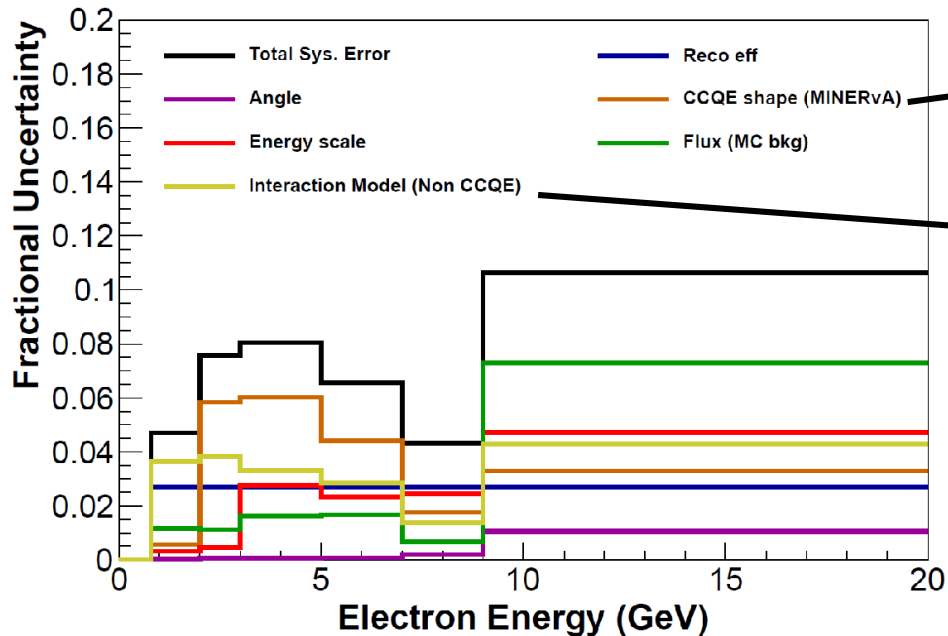
Demand dE/dx near start of track be
<4.5 MeV/1.7 cm
Removes most of NC photon
background





After the tuned background subtraction and efficiency correction (systematic errors shown in the MC, data errors statistical only)

Systematic Error on Data - MC bkg



Not relevant for DUNE

Some improvement for DUNE perhaps

Used to constrain the flux and reduce the error envelope on the flux

EM energy scale – data and Michel comparison

Reconstruction efficiency – extracted from muon tracking efficiency looking at muons in MINOS and MINERvA (these electrons are seeded in the same way)

Plastic scintillator tracker option

Pros

- High mass density – high statistics
- Good containment
- Can put in layers of Ar targets (and other targets)
- Fast – can take high rate
- Decent segmentation – can isolate different topologies
- Can separate e from gamma
- Tracking and angular resolution fairly good

Cons

- Not at all like FD
- Ar target layers possible, but those will not be useful for 1-track topologies
- Angular hole in acceptance at 90 degrees

Probably a cheap-ish option

No real attempt at costing.

A very rough guess based on a few T2K P0D & MINERvA numbers and number of channels is O(\$10M) for tracker (not including ECAL or magnet or Ar target planes)

Backup slides

Single point measurement
resolution (m)

$$\left(\frac{\delta p}{p}\right)^2 = \left(\frac{\sigma_s p}{0.3BL^2 \sqrt{\frac{720}{N+4}}}\right)^2 + \left(0.045 \frac{1}{B\sqrt{LX_o}}\right)^2$$

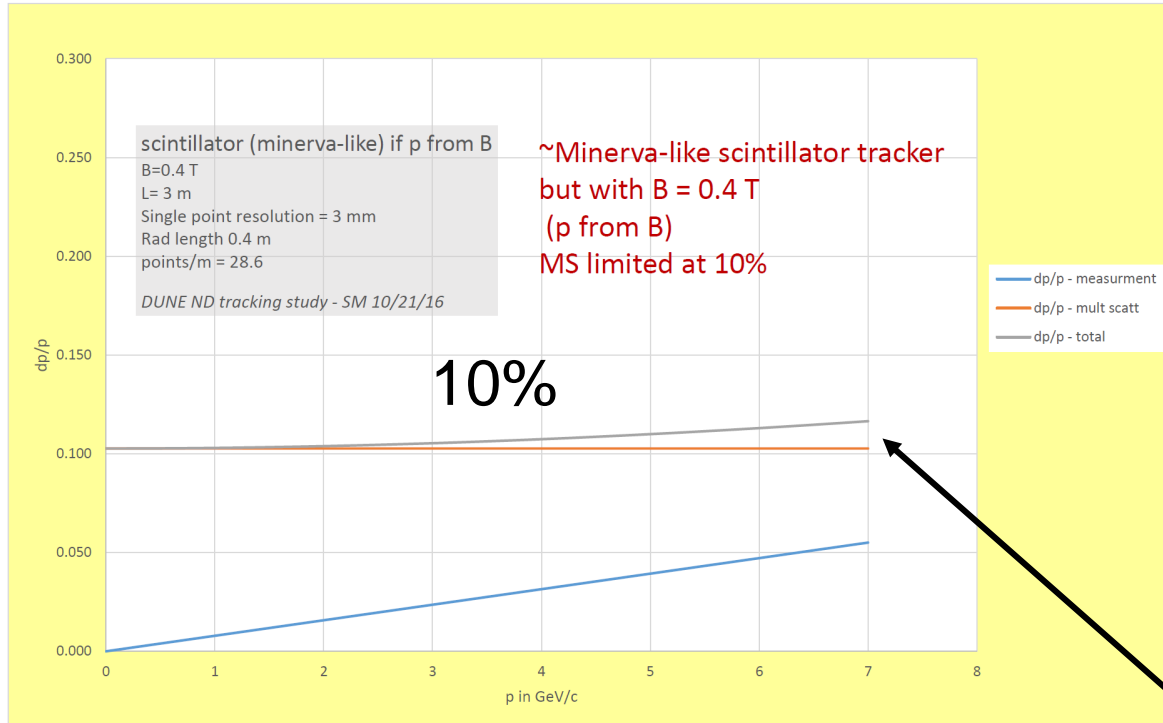
The diagram shows the equation with several annotations:

- A red arrow points from the text "Single point measurement resolution (m)" to the σ_s term in the first term of the equation.
- A blue arrow points from "Track momentum in GeV/c" to the p in the numerator of the first term.
- A blue arrow points from "B field strength in T" to the B in the denominator of the first term.
- A blue arrow points from "Track length in m" to the L in the denominator of the first term.
- A grey arrow points from "Number of measurement points" to the N in the denominator of the first term.
- An orange double-headed arrow labeled "Measurement term" spans the width of the first term.
- An orange double-headed arrow labeled "Multiple scattering term" spans the width of the second term.
- A purple arrow points from "Radiation length of medium (m)" to the X_o in the denominator of the second term.

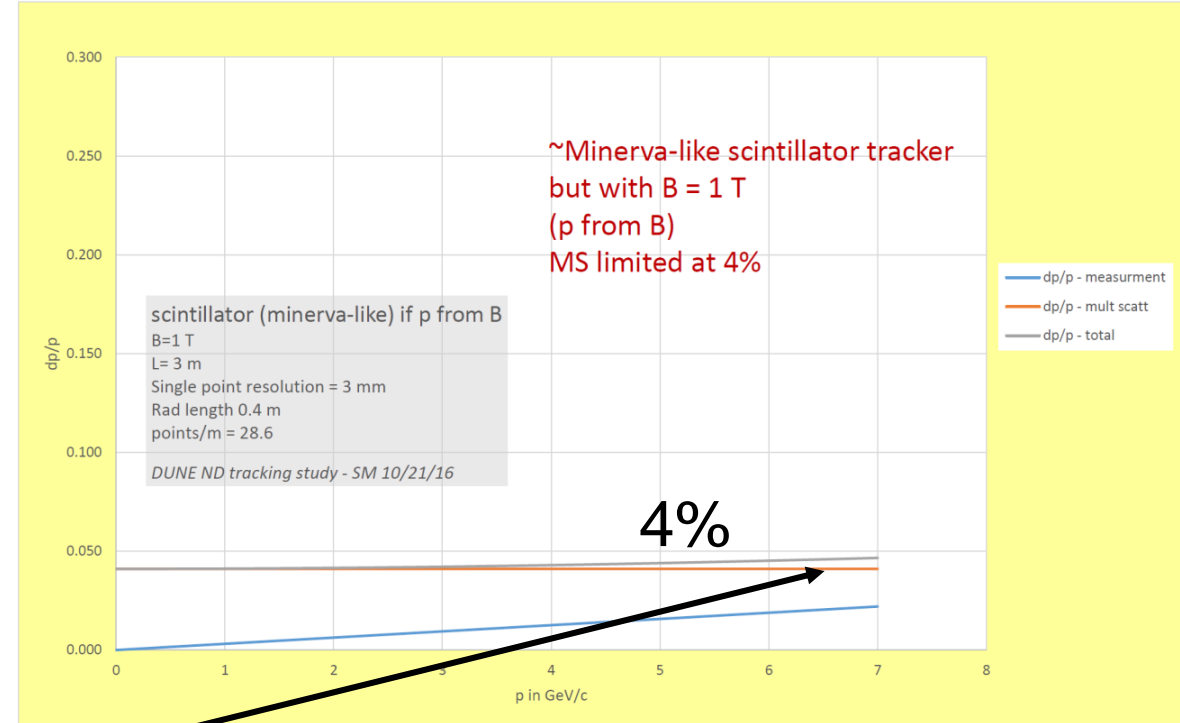
From:
 Gluckstern NIM 24 (1963) 381
 Michael Moll, presentation at CERN on tracking on 15 May 2011

dp/p Performance

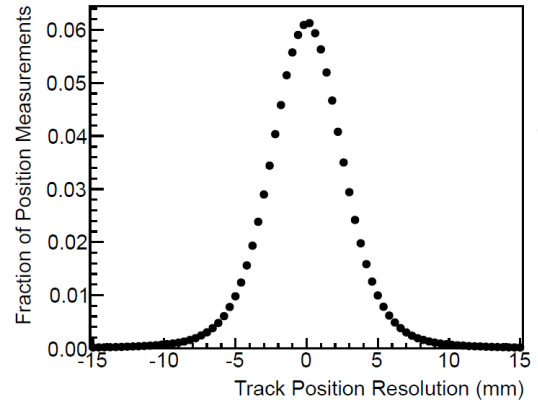
B = 0.4 T



B = 1.0 T

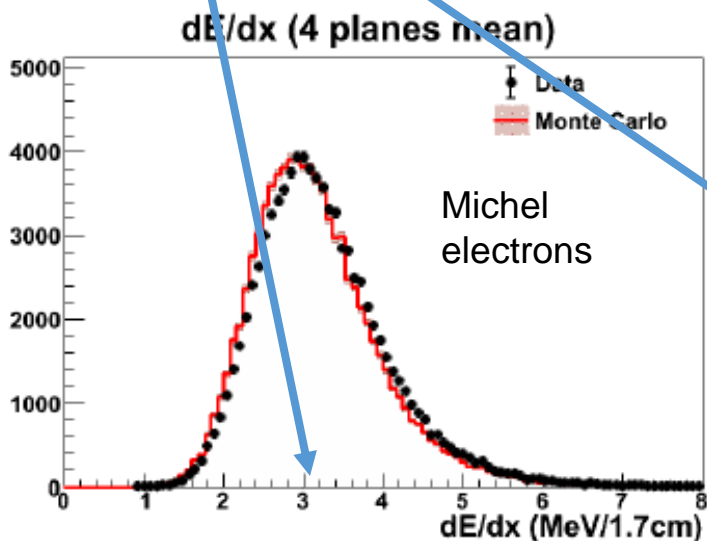
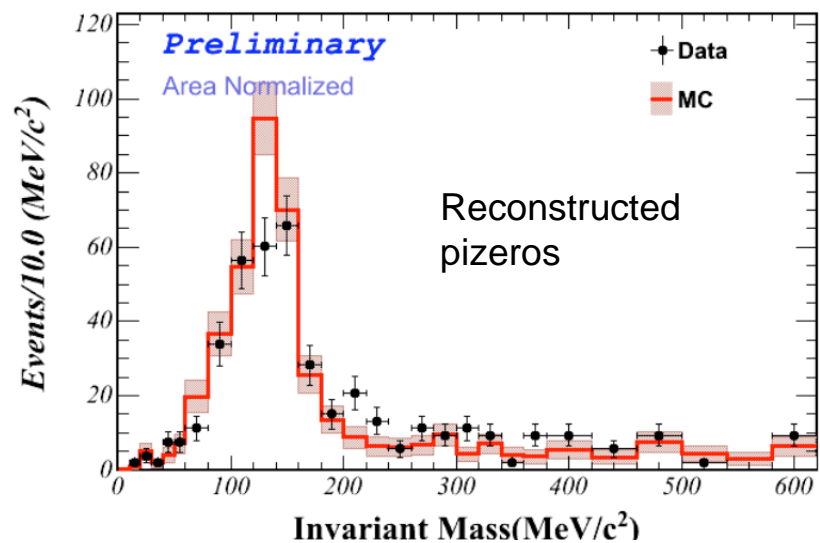
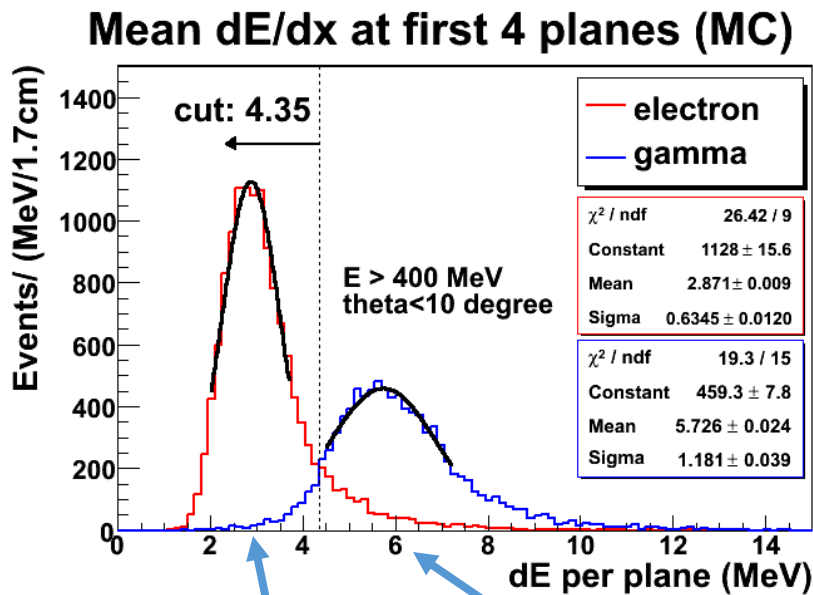


Multiple scattering dominated. Expect to see relatively little improvement with smaller transverse xsec strips.

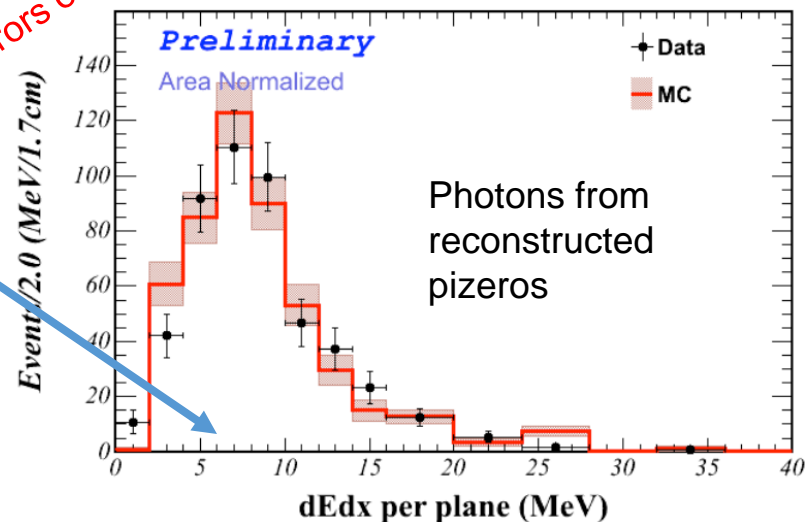


Resolution of fitted positions along a track relative to the measured cluster position for a sample of data rock muons, RMS is 3.1 mm

Michels, pizeros and e/ γ separation

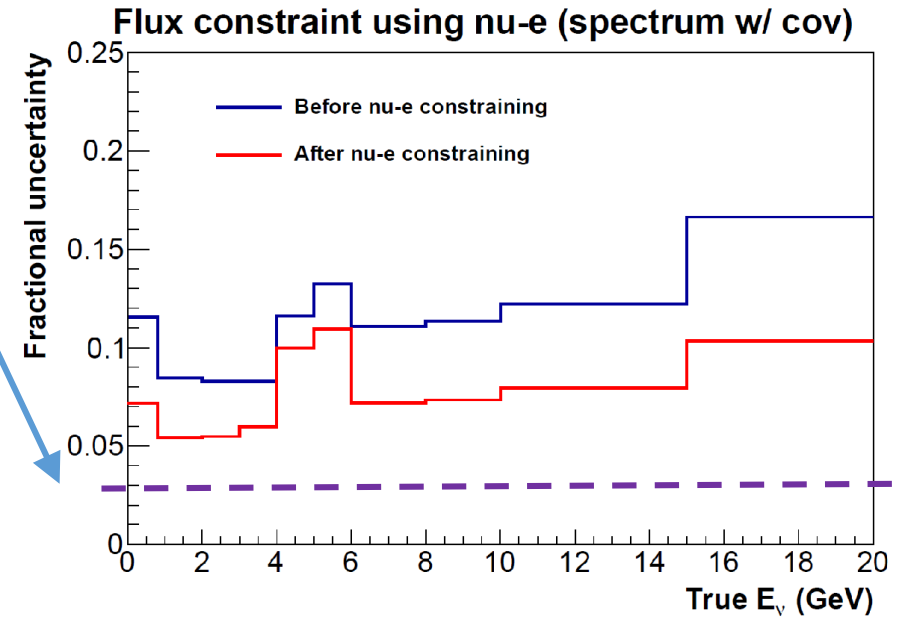
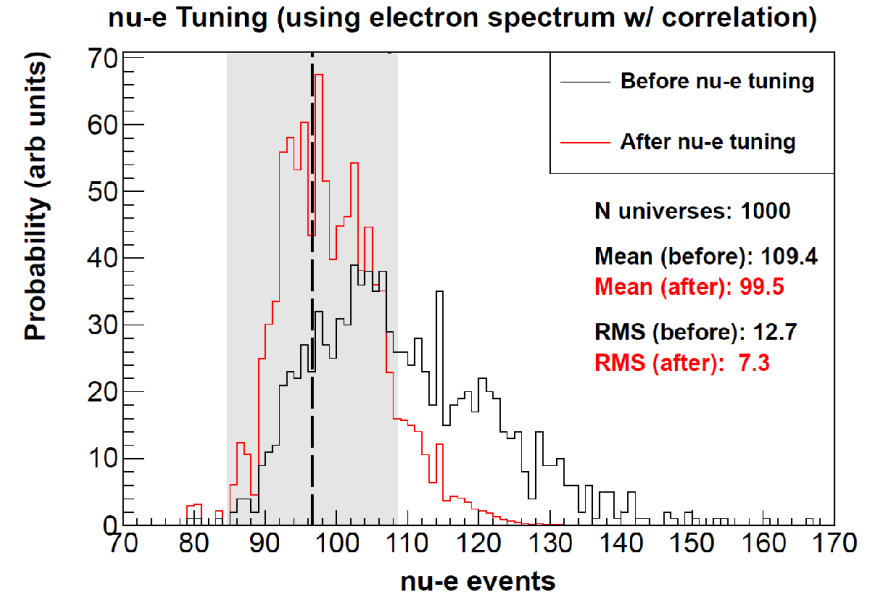
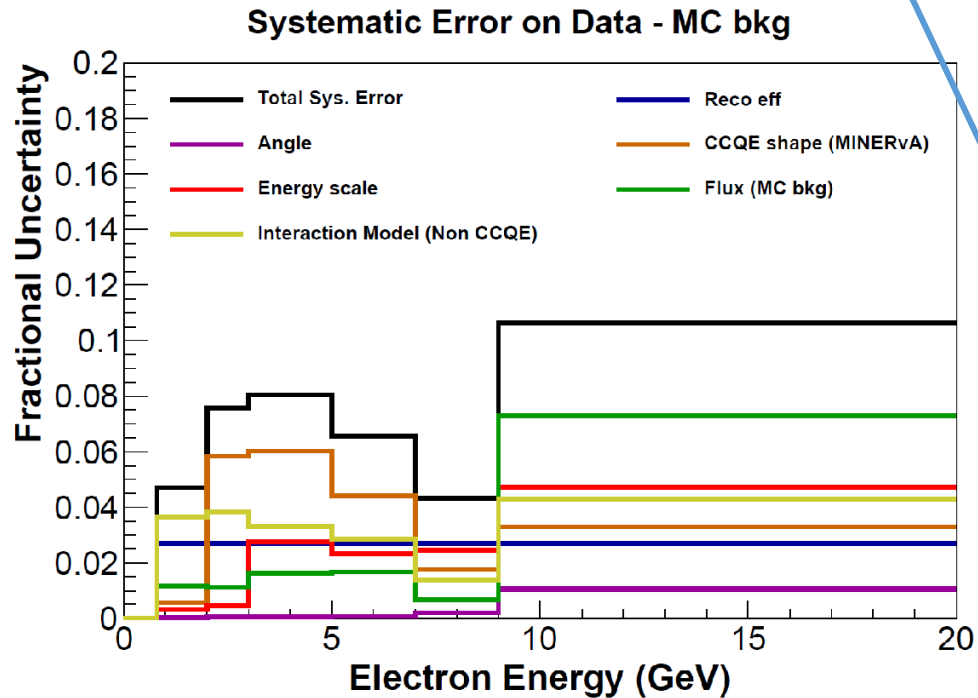


Statistical errors only

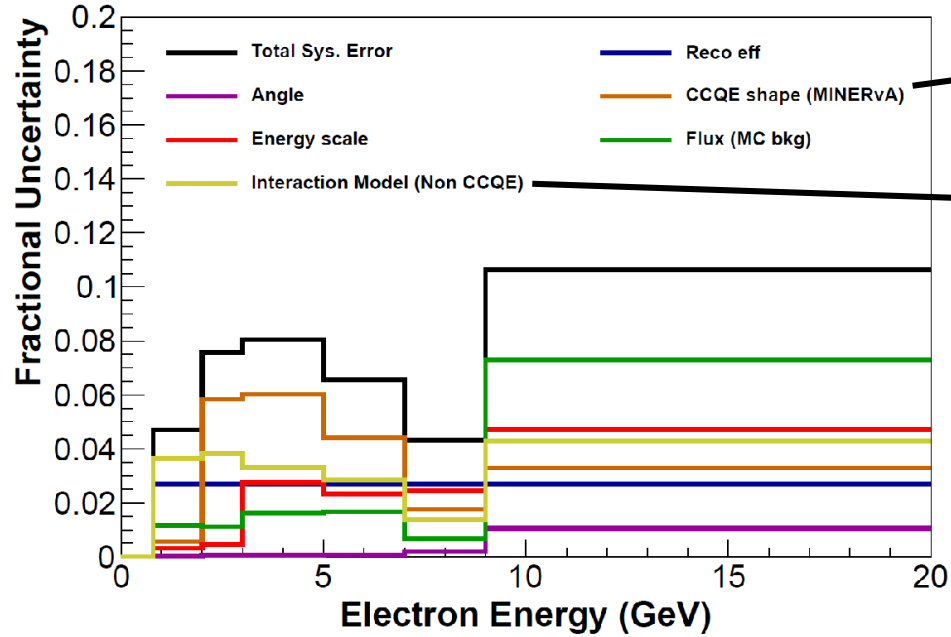


CDR (physics 6-93) states absolute neutrino flux known to <3% in 0.5-10 GeV neutrino energy range

Neutrino-electron scattering state-of-art
MINERvA PRD



Systematic Error on Data - MC bkg



Not relevant for DUNE

Some improvement for DUNE perhaps

- DUNE – higher stats (can use non-Ar part of FGT)
- Premium on EM energy recon and electron angular resolution (for background rejection)
- EM energy scale error important (want test beam probably)
- Full analysis should be done to see how a 3% error in number of events as function of electron energy translates into error in absolute flux as function of neutrino energy

- Aggressive but perhaps not crazy

Low ν method

Claim gives absolute shape of numu and numubar flux to 1-2% precision for $1 < E_\nu < 50$ GeV.

Gives FD/ND(Ev) to 1-2% precision.

MINERvA reports flux shape from Low nu with lower limit $E_\nu = 2$ GeV due to worries about sensitivities to mis-modeling for $\nu < 300$ MeV.

MINERvA normalizes to NOMAD data in the 9-12 GeV region with a 3.6% uncertainty. Statistical error in the normalization bin (9-12 GeV) blows up the normalization error. FD/ND should be much smaller.

For antineutrinos lack of good normalization data has MINERvA normalizing to GENIE tuned to world average neutrino xsec

Needs study. (1-2% seems optimistic. 3-4% likely doable)
DUNE ND might have better stat error. But not clear to me there is a better normalization at higher energy to use. Not sure how plays out in the ratio.

