U.S. Cosmic Visions: New Ideas in Dark Matter

University of Maryland March, 23rd 2017

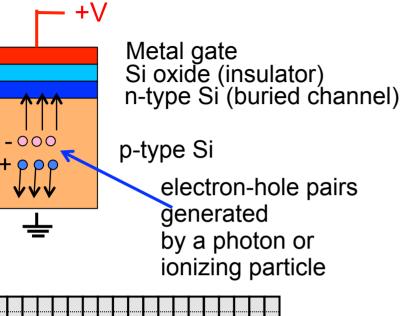
DAMIC and a kg-size CCD experiment

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(Photo image: particle tracks in a DAMIC CCD)

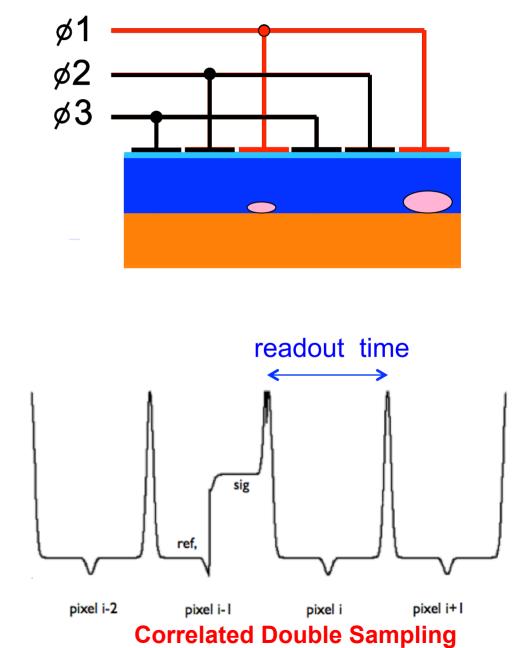
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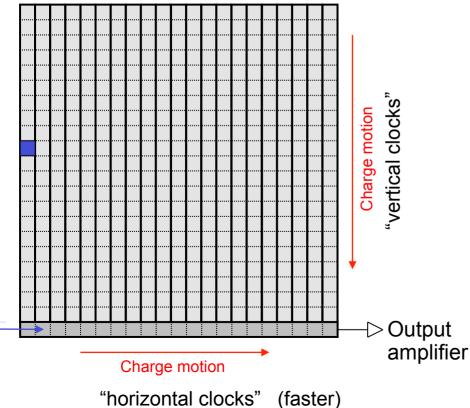
Metal-Oxide-Semiconductor capacitor

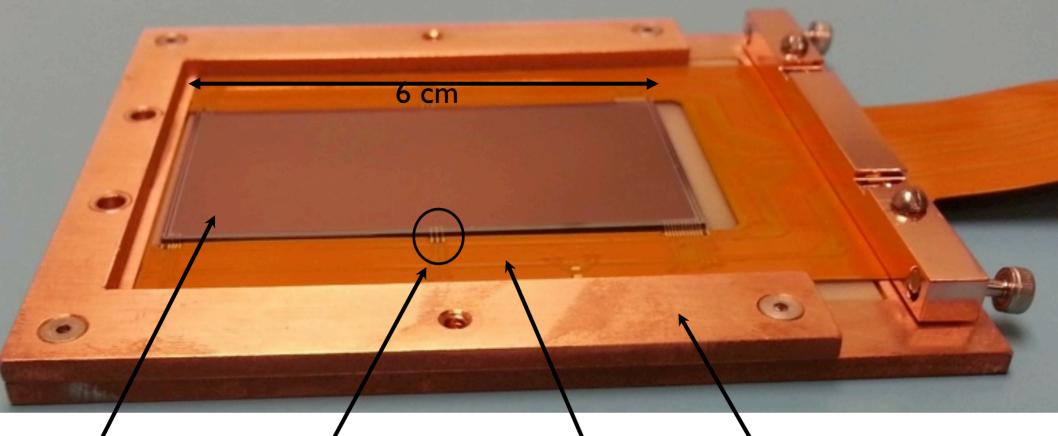


CCD principle

Moving charge from pixel to pixel







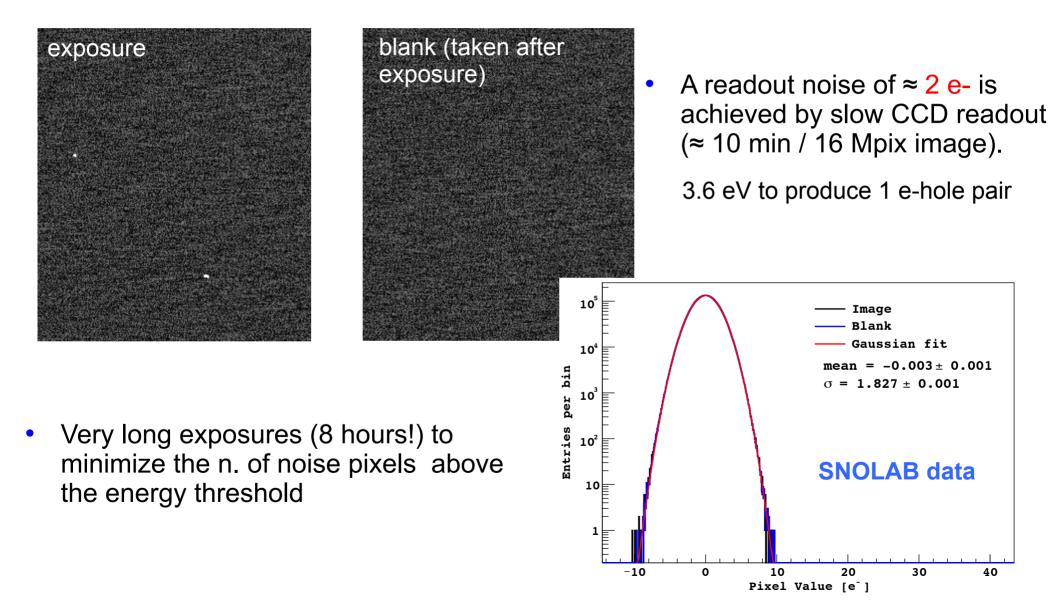
CCD 2k x 4k Wire bonds Clocks, Bias, Copper frame and Signal cable

1) Sizable mass (high resistivity, thick CCDs designed by LBNL) A DAMIC CCD has an active area of 6 cm x 6 cm, 16 Mpixel (each 15 μ m x 15 μ m) and a record thickness of 675 μ m for a total of 5.9 g mass

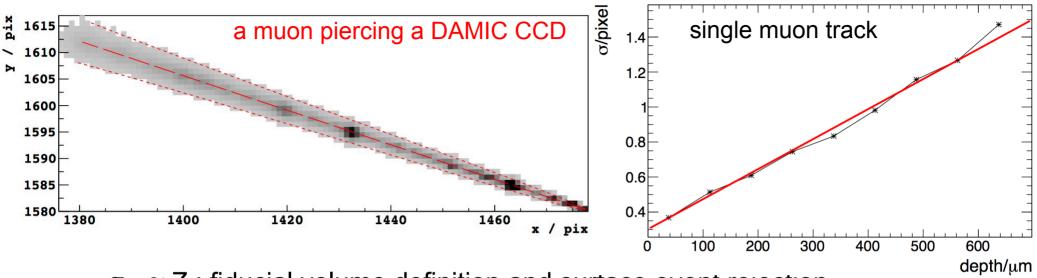
DAMIC100 currently taking data at the SNOLAB underground laboratory

2) Unprecedented low energy threshold

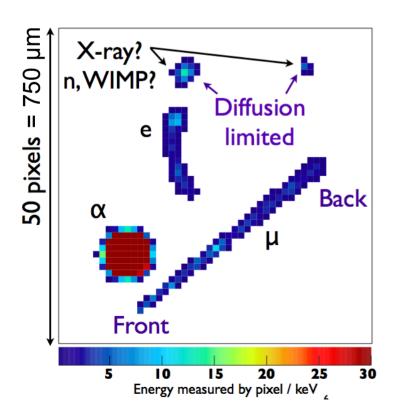
 Negligible noise contribution from dark current fluctuations (dark current < 0.001 e/pixel/day with CCD cooled at 120 K). Readout noise dominant contribution.



4) Unique spatial resolution: 3D position reconstruction and particle ID

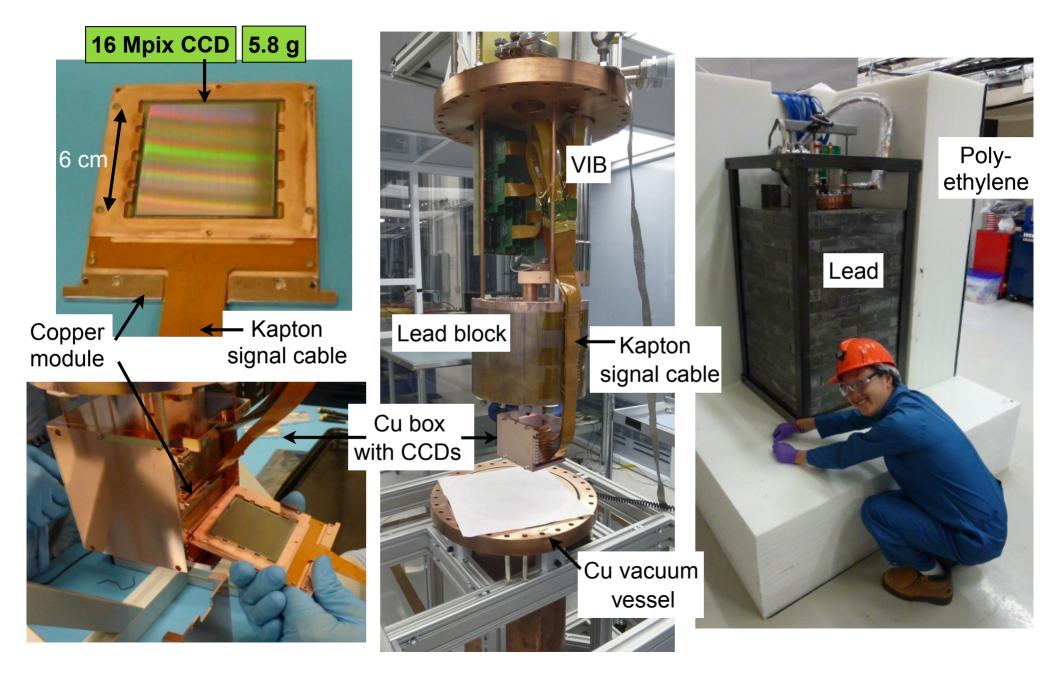


 $\sigma_{xv} \approx Z$: fiducial volume definition and surface event rejection



- "Worms": straggling electrons
- Straight tracks: minimum ionizing particles
- MeV charge blobs: alphas
- Diffusion-limited clusters: low-energy X-rays, nuclear recoils
- CCD spatial resolution provides a unique handle to the understanding of the background

DAMIC @ SNOLAB





Measurement of radioactive contamination in the high-resistivity siliconCCDs of the DAMIC experimentJINST 10 (2015) P08014

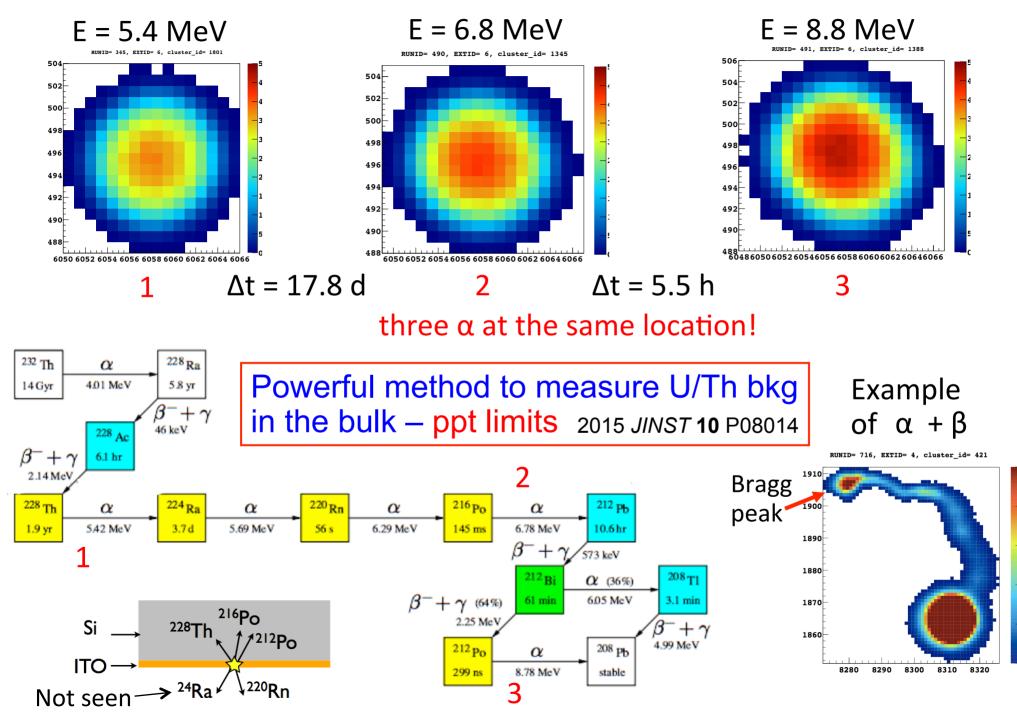
Search for low-mass WIMPs in a 0.6 kg day exposure of the DAMIC experiment at SNOLAB Phys. Rev. D 94, 082006 (2016)

First direct detection constraints on eV-scale hidden-photon dark matterwith DAMIC at SNOLABarXiv:1611.03066 accepted by Phys. Rev. Lett.

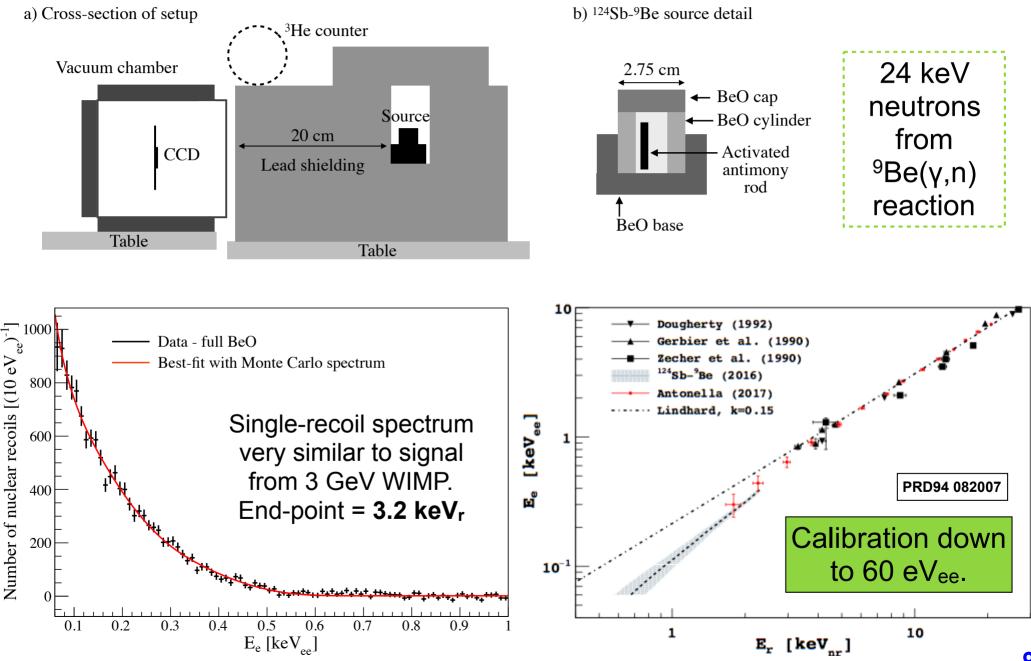
Measurement of the ionization produced by sub-keV silicon nuclear recoilsin a CCD dark matter detectorPhys. Rev. D 94, 082007 (2016)

Antonella: A nuclear-recoil ionization-efficiency measurement in silicon at low energies arXiv:1702.00873

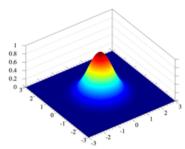
Radiogenic backgrounds



Nuclear recoil calibration

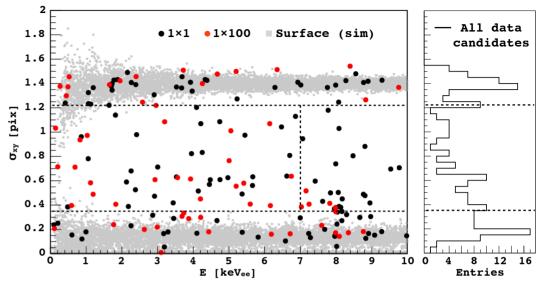


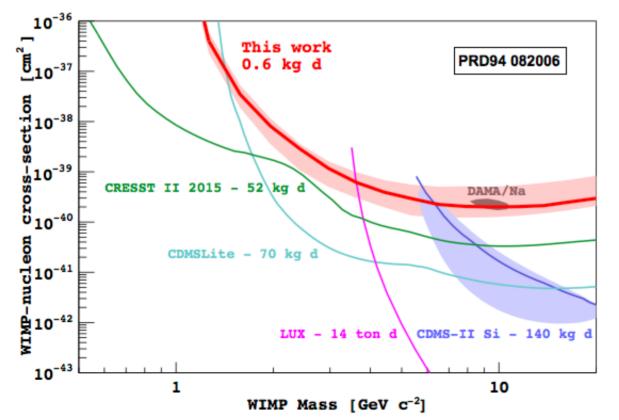
WIMPs search



Measure **E** and σ_{xy} for every cluster event.

 $\sigma_{xy} \approx$ proportional to depth of interaction in the bulk silicon



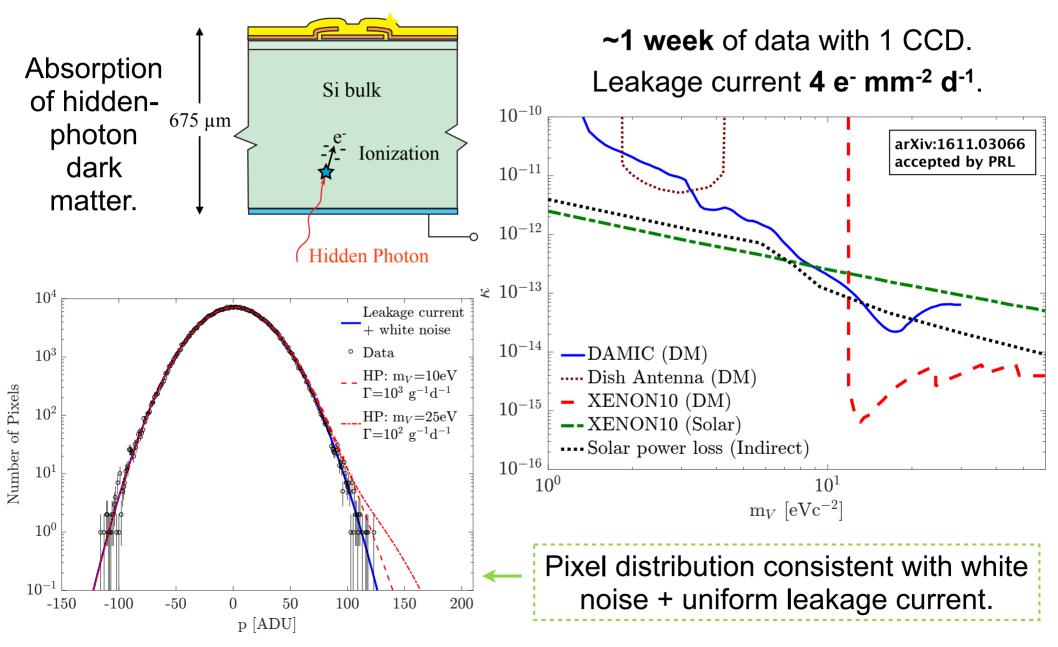


limited exposure taken during R&D phase (bkg. ≈ 30 dru)

demonstration of DAMIC sensitivity to low-mass Dark Matter

NOTE: current bkg. ≈ 5 dru

Hidden photon search



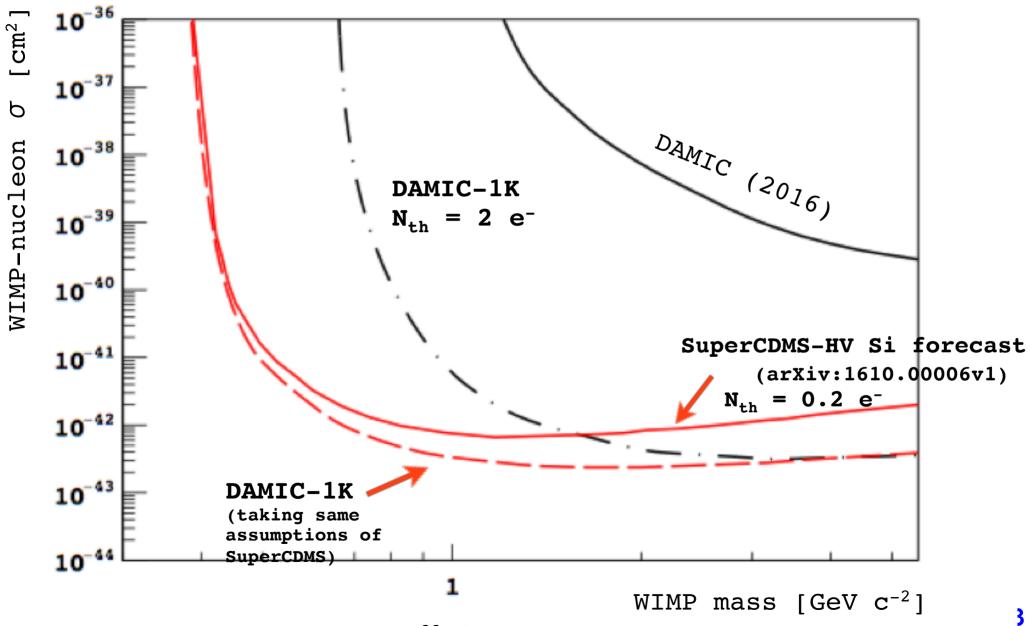
DAMIC now

- Already achieved radioactive background (5 dru) and low-noise (<2 e-) performance for a larger detector.
- Stack of 16 Mpix CCDs: DAMIC100 in current SNOLAB vacuum vessel and shielding.
- Installation took place in January, results with ≈ 10 kg day of data expected in 2017.
- Ongoing R&D for thicker, larger-area CCDs for a lower-noise, lower-background kg-size detector.

DAMIC-1K

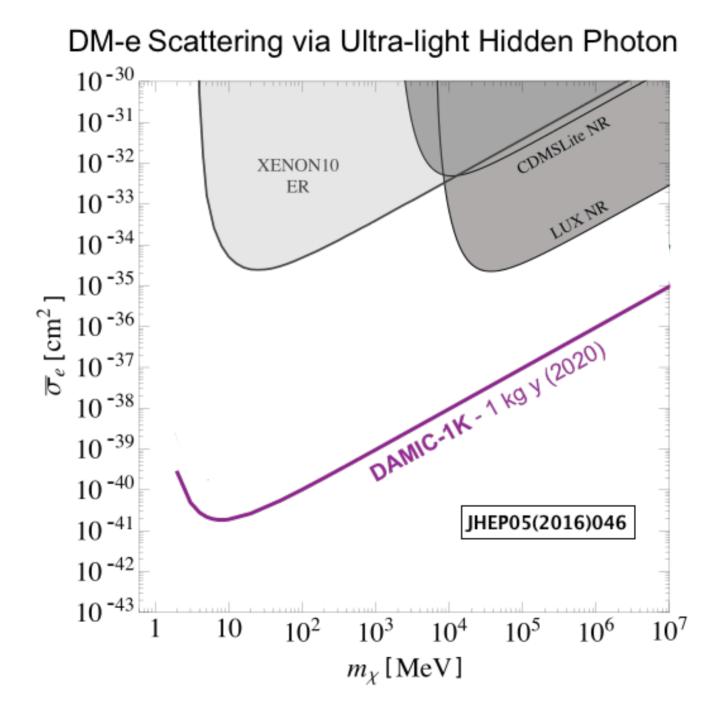
- A kg-size experiment with 0.1 dru background and ≤ 2e- threshold
- To lead the exploration of WIMPs and dark sector candidates in the low-mass DM parameter space

DAMIC-1K and WIMPs



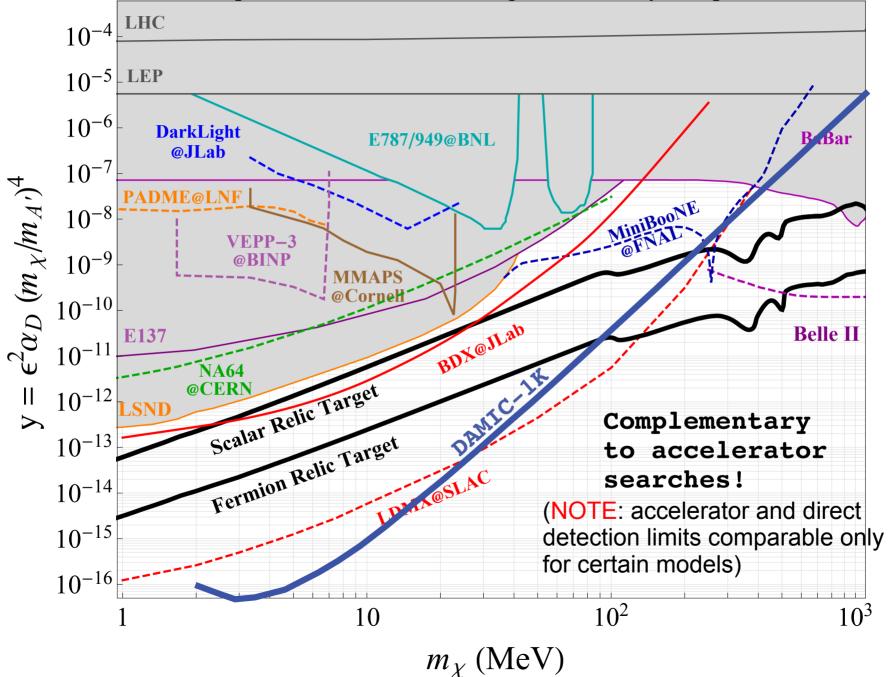
DAMIC-1K not limited by ³²Si bkg.

DAMIC-1K and dark sector



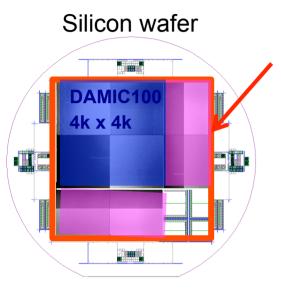
DAMIC-1K and dark sector

All Experiments (Kinetic Mixing + Elastically Coupled DM)



DAMIC-1K technical challenges

• A kg-size DAMIC can be built with the existing technology



6k x 6k pixels, 1 mm thick

≈ 20 g / CCD

≈ 50 CCDs / 1 Kg

DALSA has confirmed the feasibility fabrication of these larger and thicker CCDs

• Background

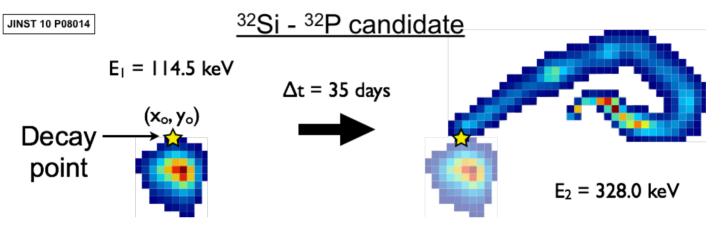
from a few dru to a fraction of dru.

external bkg.: improved design, materials (e.g. electroformed copper), strict procedures (silicon storage underground, radon, surface contamination)

internal bkg.: cosmogenic ³²Si and tritium

DAMIC-1K background

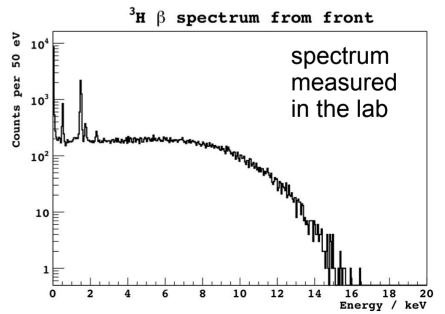
 Cosmogenic ³²Si rate will be accurately measured by the current detector at SNOLAB



≈ 1 dru (dominant bkg. in SuperCDMS); rejected in DAMIC-1K by spatial correlations

• Tritium expected to be the dominant bkg. for DAMIC-1K.

<u>A measurement of its rate may be</u> <u>within reach of the current DAMIC</u> <u>detector at SNOLAB</u> (so far only estimates are used for forecasts)



DAMIC-1K sub-e⁻ noise

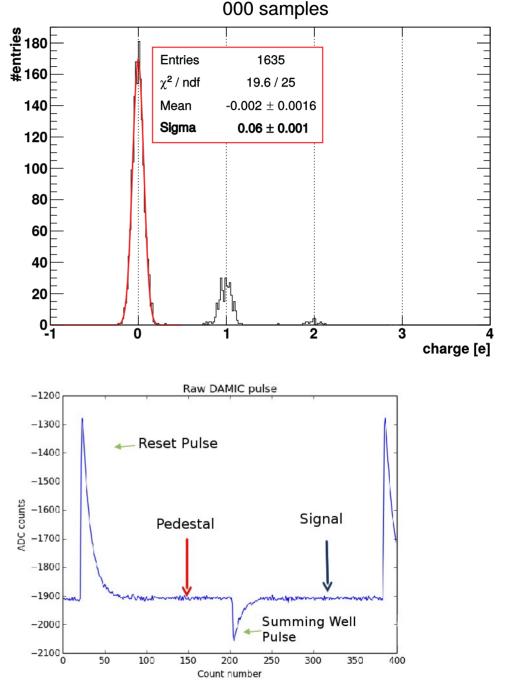
Skipper readout

Already achieved on a small size CCD (LBNL design; LDRD at Fermilab) See J. Tiffenberg SENSEI talk

Digital filtered CDS

does not require a new CCD design, could already significantly improve sensitivity of DAMIC100





The DAMIC-1K experiment

DAMIC Collaboration

CAB, Fermilab, LPNHE, SNOLAB, U Chicago, U Michigan, UNA, UNAM, UFRJ, U Zürich

Several new groups interested in joining DAMIC-1K (PNNL, Denmark, France, Germany, Spain)

- Timing
- 2017-2018 R&D for sub-e⁻ noise/background measurements; DAMIC100 results; finalize detector design
- 2018-2019 components validation (test at SNOLAB)
- 2019-2020 Construction
- Budget
 - \leq 3 M\$ (construction)

Conclusions

- In the last two years DAMIC has established the CCD technology as a competitive technique for the search of low-mass Dark Matter particles
- DAMIC-1K, a kg-size CCD detector with low background and subelectron noise, will explore a new large parameter space, scrutinizing the WIMPs paradigma, as well as dark sector candidates with sensitivity comparable to accelerator searches (for certain classes of models)
- The DAMIC-1K detector is an incremental step of proven technologies (larger size CCD, sub-electron noise). It will work as specified.
- There are strong synergies between DAMIC-1K and SuperCDMS Si: measurement of the quenching factor down to the ionization threshold; ³²Si "depleted" silicon: DAMIC spatial coincidence technique unique tool to measure residual background.