



# NUCLEAR EMULSIONS FOR WIMP SEARCH

## directional measurement

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*on behalf of the NEWSdm Collaboration*

U.S. Cosmic Visions: New Ideas in Dark Matter

*University of Maryland, College Park*

23-25 March 2017

# LETTER OF INTENT

- Submitted to Gran Sasso Scientific Committee at the end of 2015

## NEWSdm Collaboration 70 physicists, 14 institutes

LNGS-LOI 48/15

### NEWS: Nuclear Emulsions for WIMP Search Letter of Intent (NEWS Collaboration)

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#### SOUTH KOREA

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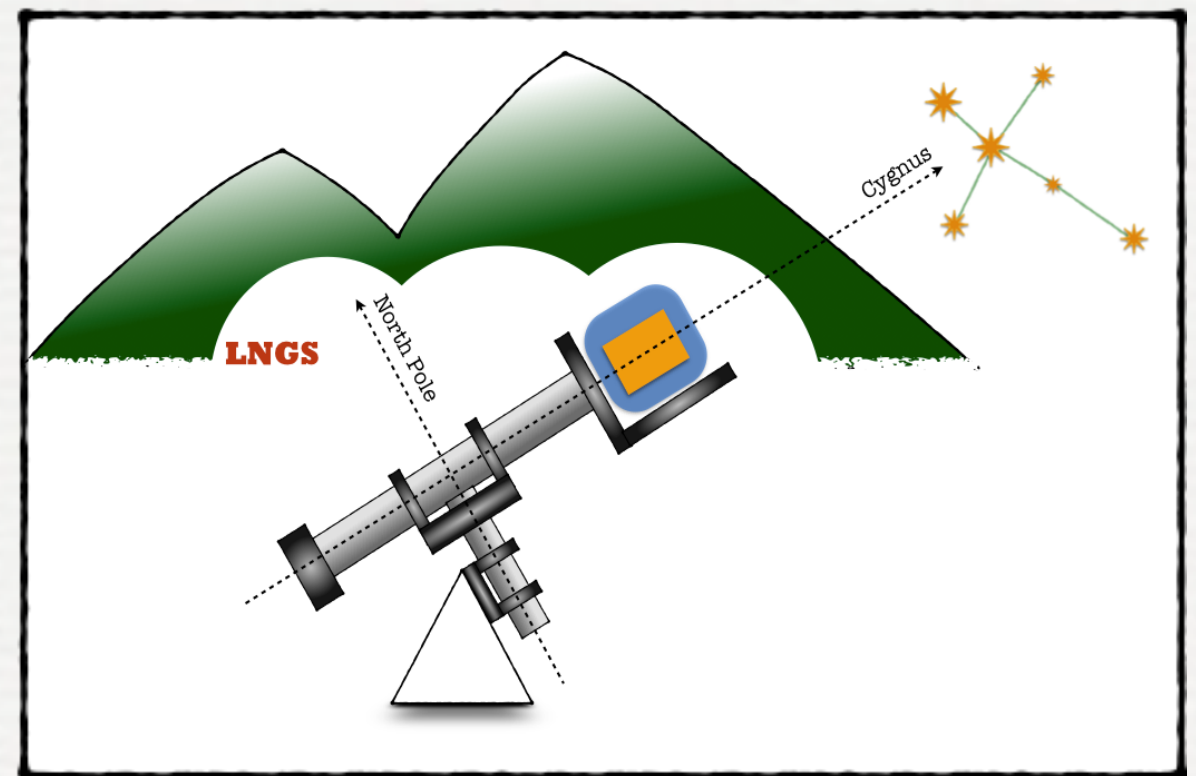
<https://arxiv.org/pdf/1604.04199.pdf>

[news-dm.lngs.infn.it](https://news-dm.lngs.infn.it)



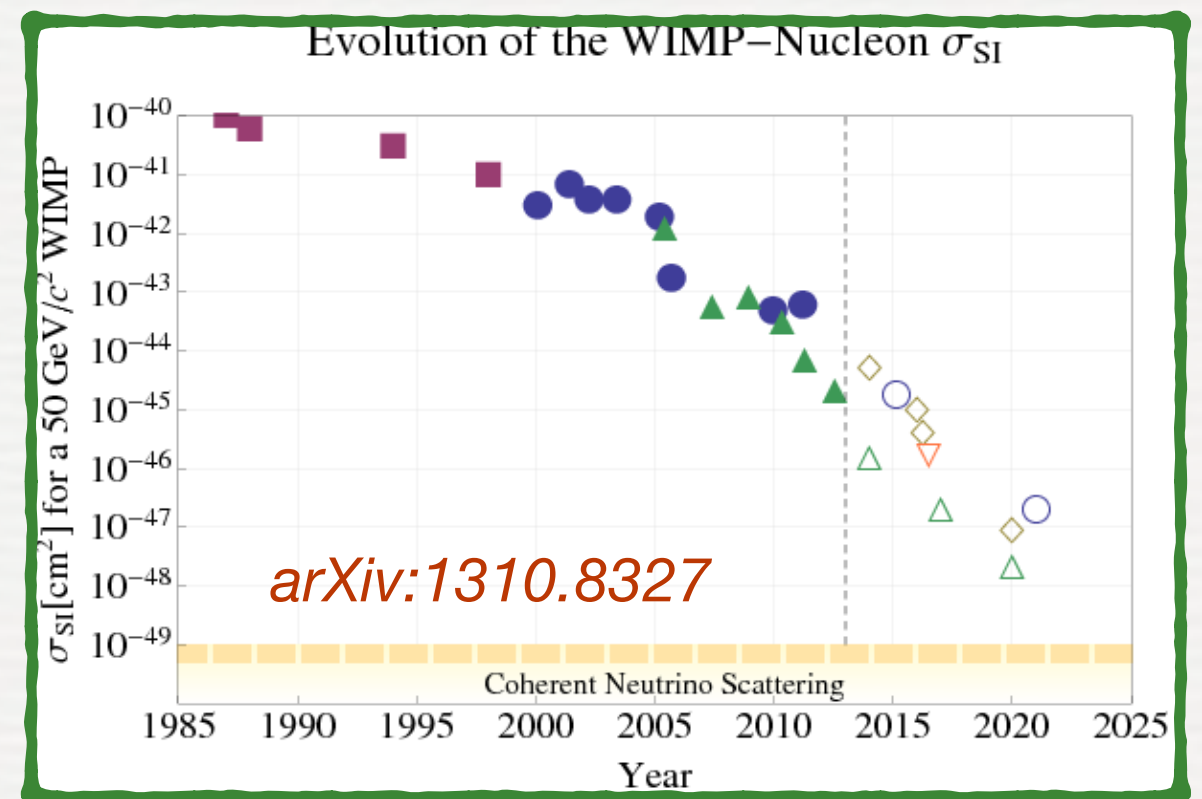
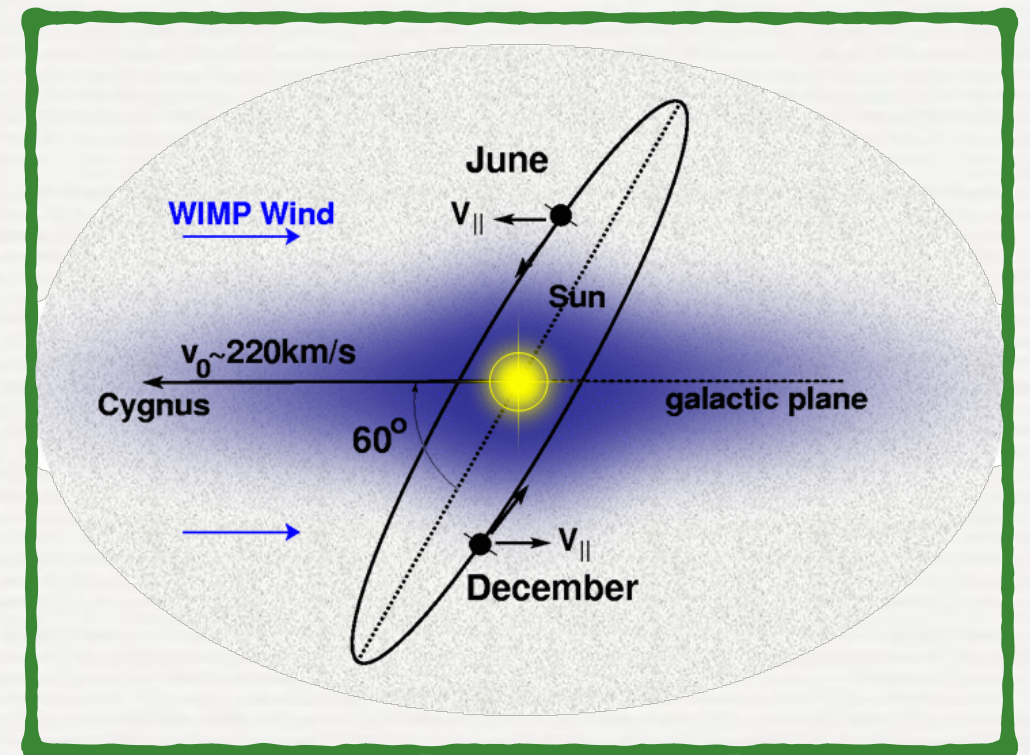
# OUTLINE

- The NEWSdm idea:
  - a novel approach to *directional* detection of Dark Matter
- High Resolution Nuclear Emulsions: NIT
- Detection principle
- Sensitivity
- Current status of the experiment
- Conclusions and perspectives



# POWER OF DIRECTIONALITY

- Impinging direction of DM particle is (preferentially) opposite to the velocity of the Sun in the Galaxy, i. e. from Cygnus Constellation
- Unambiguous proof of the galactic origin of Dark Matter
- Unique possibility to overcome the “neutrino floor”, where coherent neutrino scattering creates an irreducible background





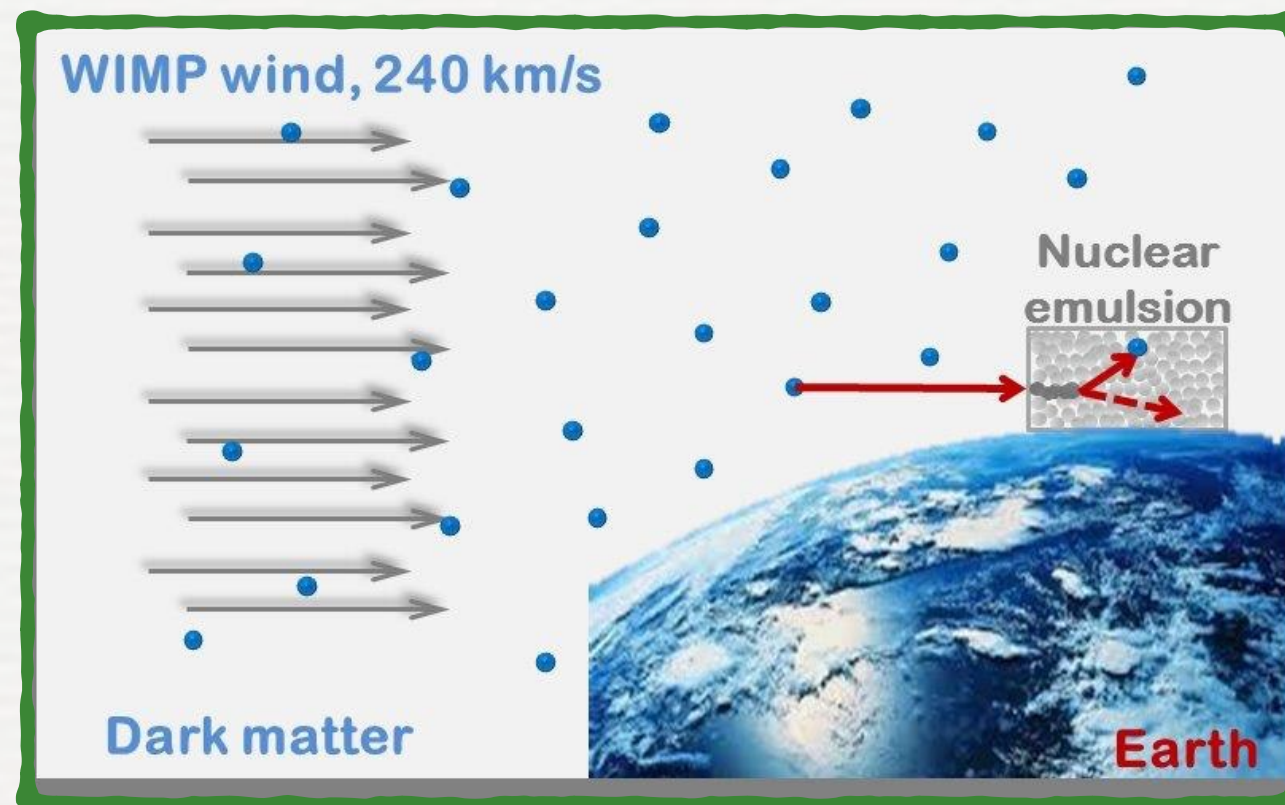
# DIRECTIONAL APPROACH

Use solid target:

- Large detector mass
- Smaller recoil track length  $O(100 \text{ nm})$

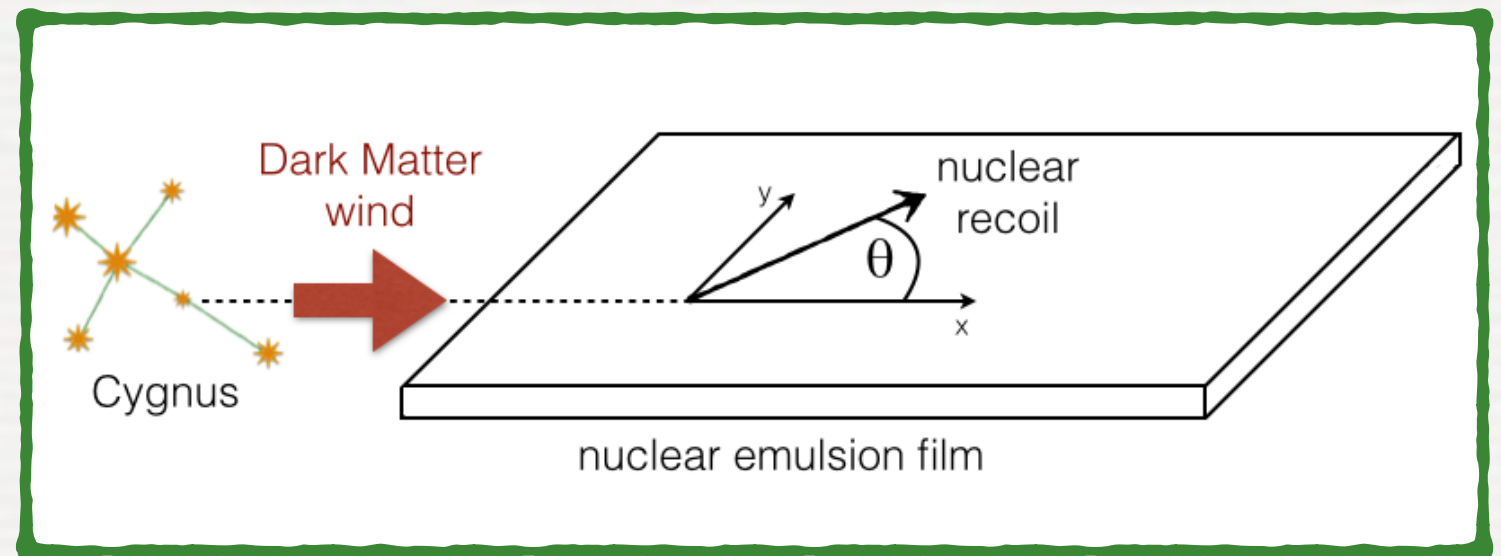
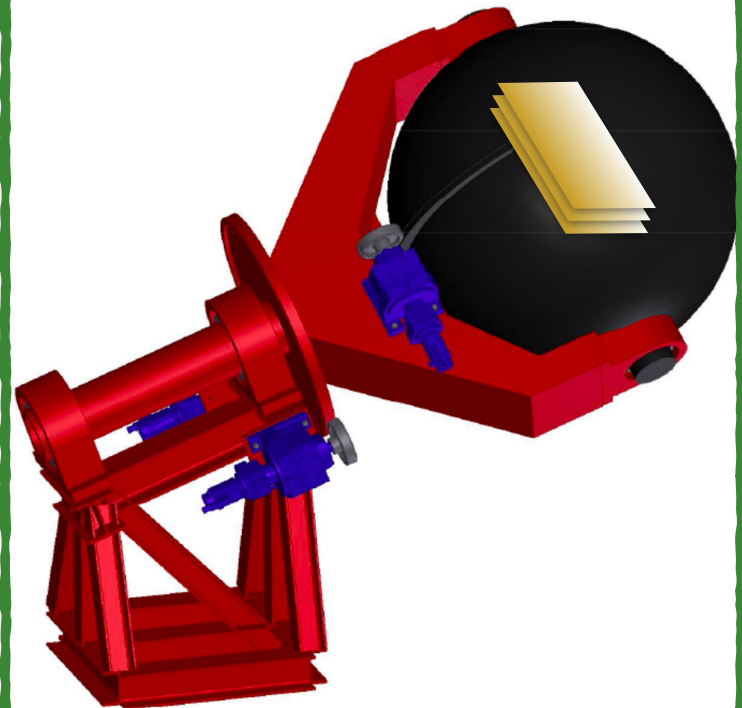
→ very high resolution tracking detector

**Nuclear Emulsion based detector  
acting both as target and tracking device**



# THE NEWSdm PRINCIPLE

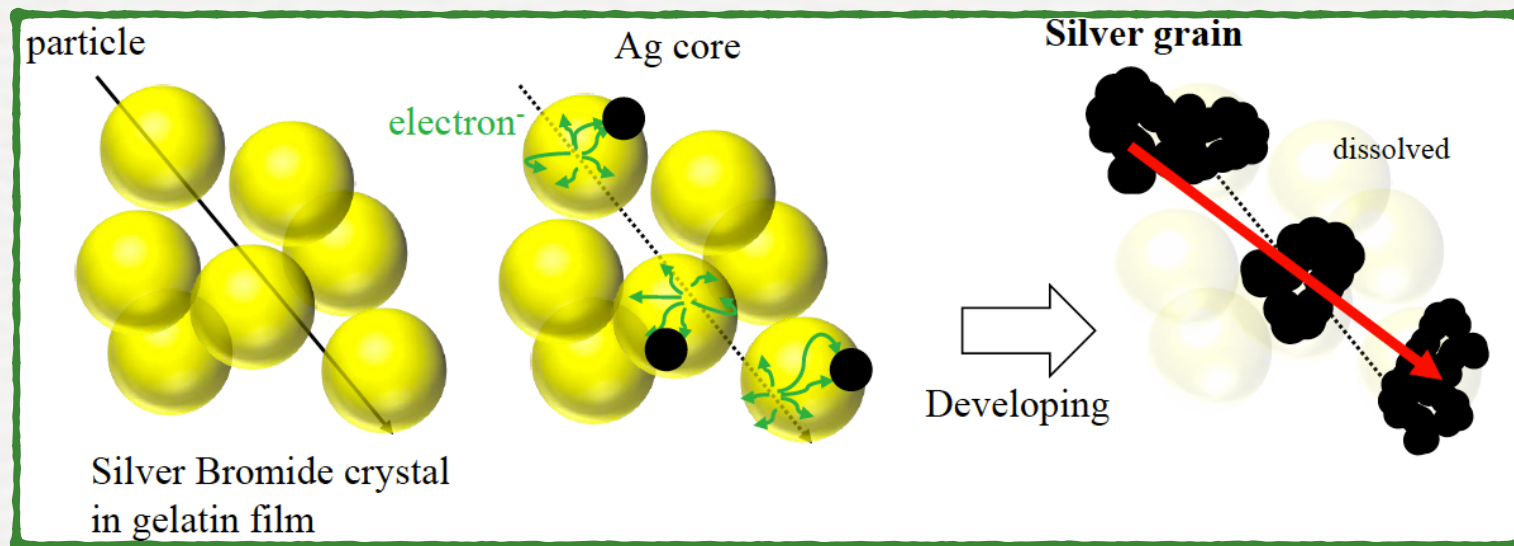
Equatorial Telescope



- **Aim:** detect the direction of **nuclear recoils** produced in WIMP interactions
- **Target:** nanometric nuclear emulsions acting both as target and tracking detector
- **Background reduction:** neutron **shield** surrounding the target
- **Fixed pointing:** target mounted on **equatorial telescope** constantly pointing to the Cygnus Constellation
- **Location:** Underground Gran Sasso Laboratory

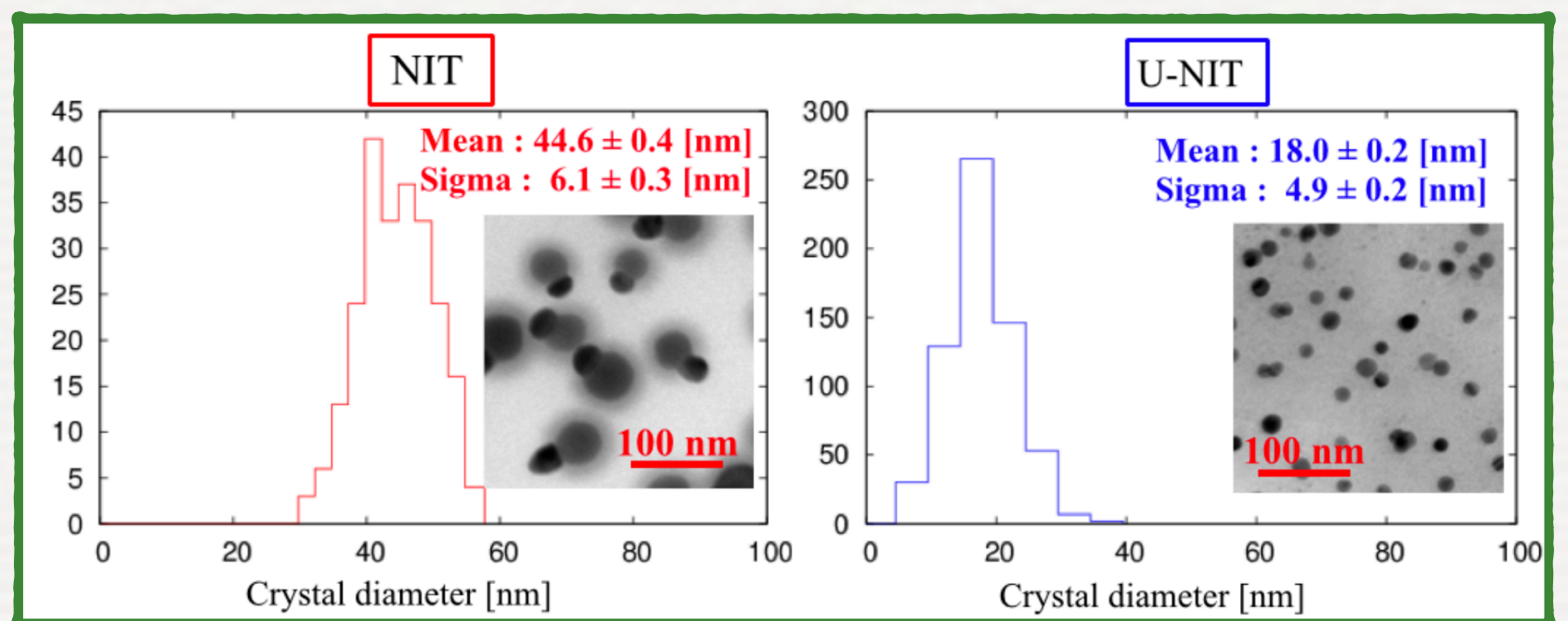
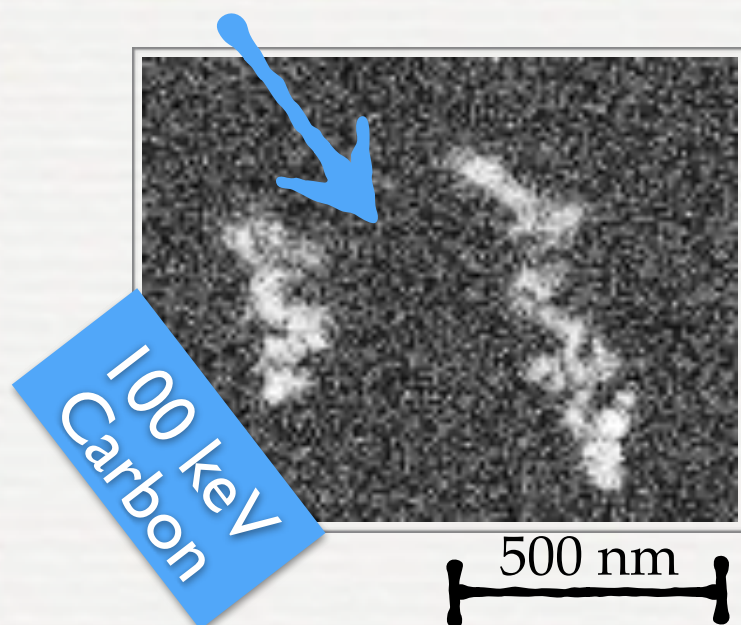


# NIT: NANO EMULSION IMAGING TRACKERS



A long history, from the discovery of the **Pion (1947)** to the discovery of  $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillation in appearance mode (**OPERA, 2015**)

- Nuclear emulsions: AgBr crystals in organic gelatine
- Passage of charged particle produce *latent image*
- Chemical treatment make Ag grains visible
- New kind of emulsion for DM search
- Smaller crystal size



# NIT EMULSIONS

Constituent	Mass Fraction
AgBr-I	0.78
Gelatin	0.17
PVA	0.05

(a) Constituents of nuclear emulsion

AgBr-I: sensitive elements

Organic gelatine: retaining structure

PVA to stabilise the crystal growth

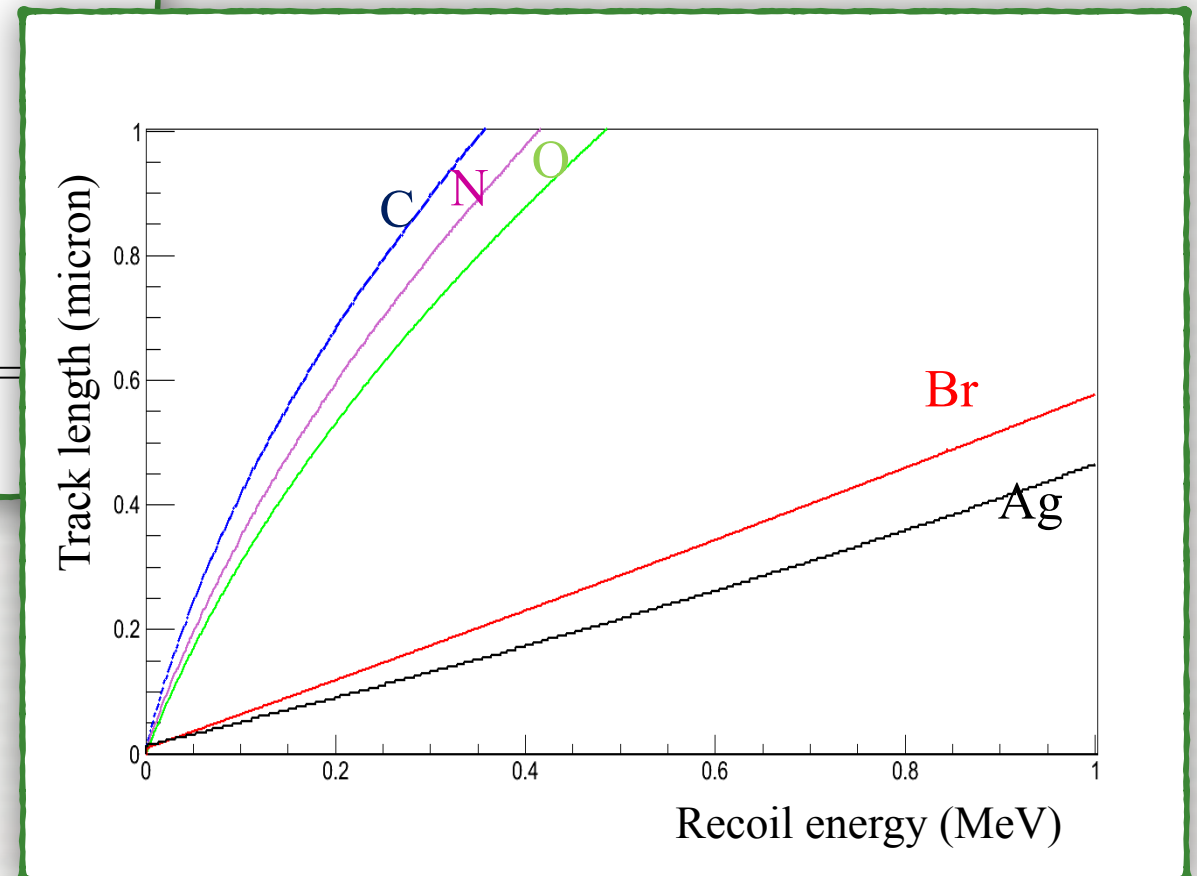
Each nucleus gives a different contribution to the overall sensitivity

heavy nuclei

light nuclei

Element	Mass Fraction	Atomic Fraction
Ag	0.44	0.12
Br	0.32	0.12
I	0.019	0.003
C	0.101	0.172
O	0.074	0.129
N	0.027	0.057
H	0.016	0.396
S	0.003	0.003

(b) Elemental composition



Lighter nuclei

(longer range at same recoil energy)



Sensitivity to low WIMP mass



# READOUT TECHNOLOGY

# TRACK IDENTIFICATION

- Challenge: detect tracks with lengths comparable / shorter than optical resolution
- Strategy: two-steps approach

## STEP 1

### CANDIDATE IDENTIFICATION WITH OPTICAL MICROSCOPES

Pros: Fast scanning profiting of the improvements driven by the OPERA experiment, dedicated measurement stations in each lab

Limit: Resolution with standard technologies ~ 200 nm

## STEP 2

### CANDIDATE VALIDATION

**X-ray microscope**

Pros: High resolution ~ 50 nm or better

Cons: extremely slow and not convenient (need an external lab)

**New technology with optical microscopes**

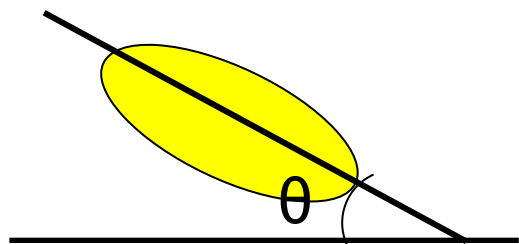


# READOUT STRATEGY

## STEP 1: CANDIDATE IDENTIFICATION

- Scanning with **optical microscope** and **shape recognition analysis**
- Automatic selection of candidate signals by optical microscopy
- Selection of clusters with elliptical shape: major axis along track direction
- Background: spherical cluster
- Resolution 200 nm (one order of magnitude better than the OPERA scanning system), scanning speed 20 cm<sup>2</sup>/h

Test using 400 keV Kr ions



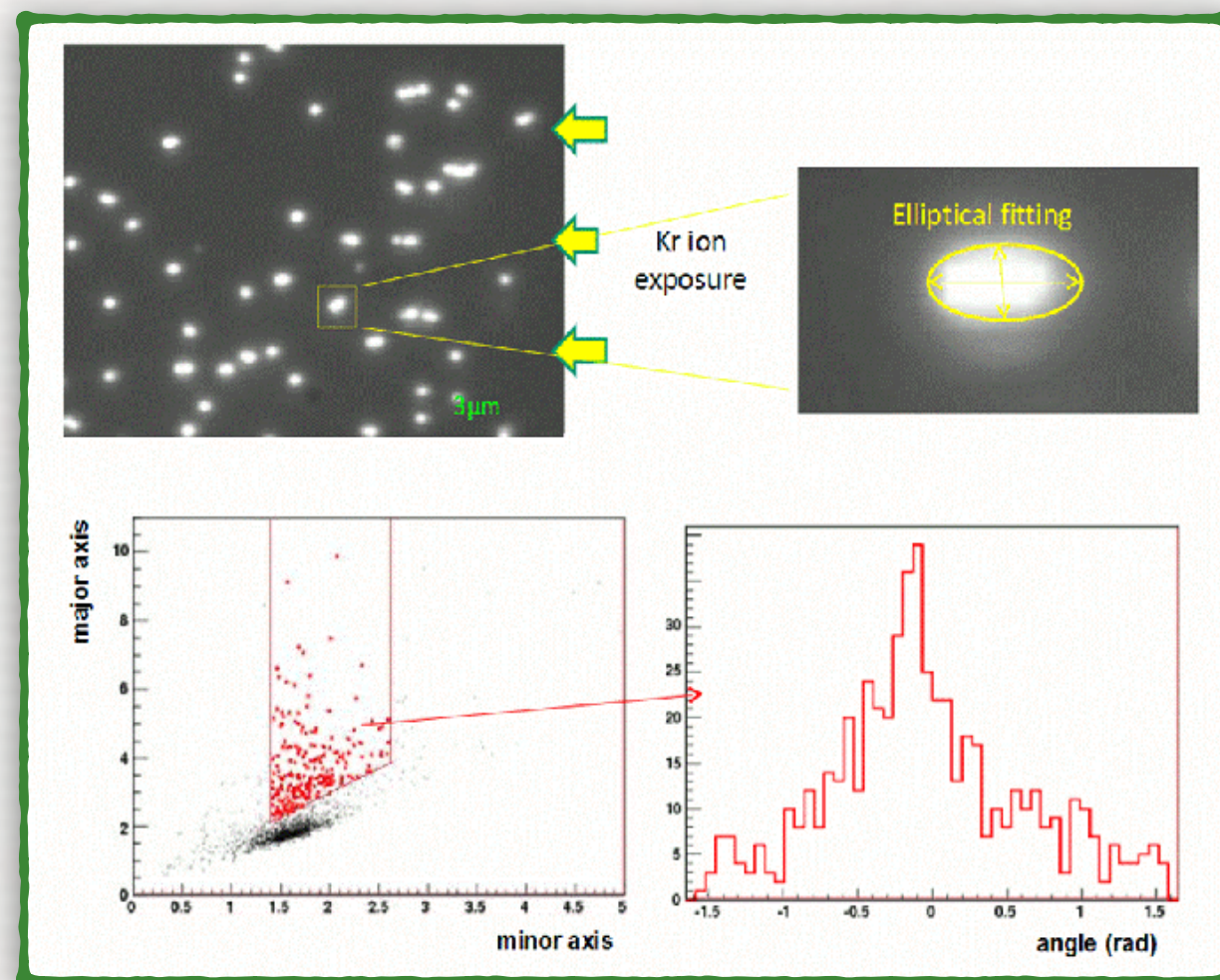
Direction detected!

*Nucl.Instrum.Meth. A680 (2012) 12-17*

OVERALL ANGULAR RESOLUTION

$$\sigma^2 = \sigma^2_{\text{intrinsic}} + \sigma^2_{\text{scattering}}$$

$$\sigma = 360 \text{ mrad}$$

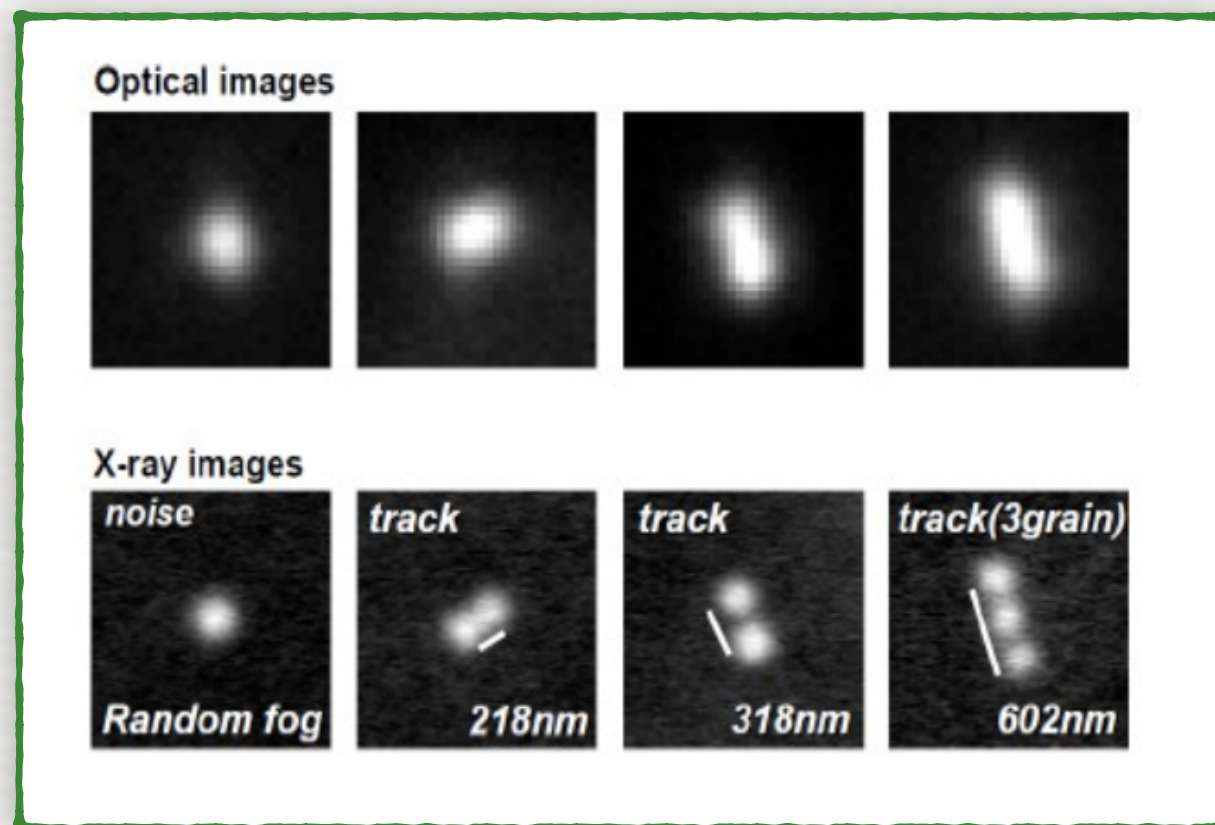




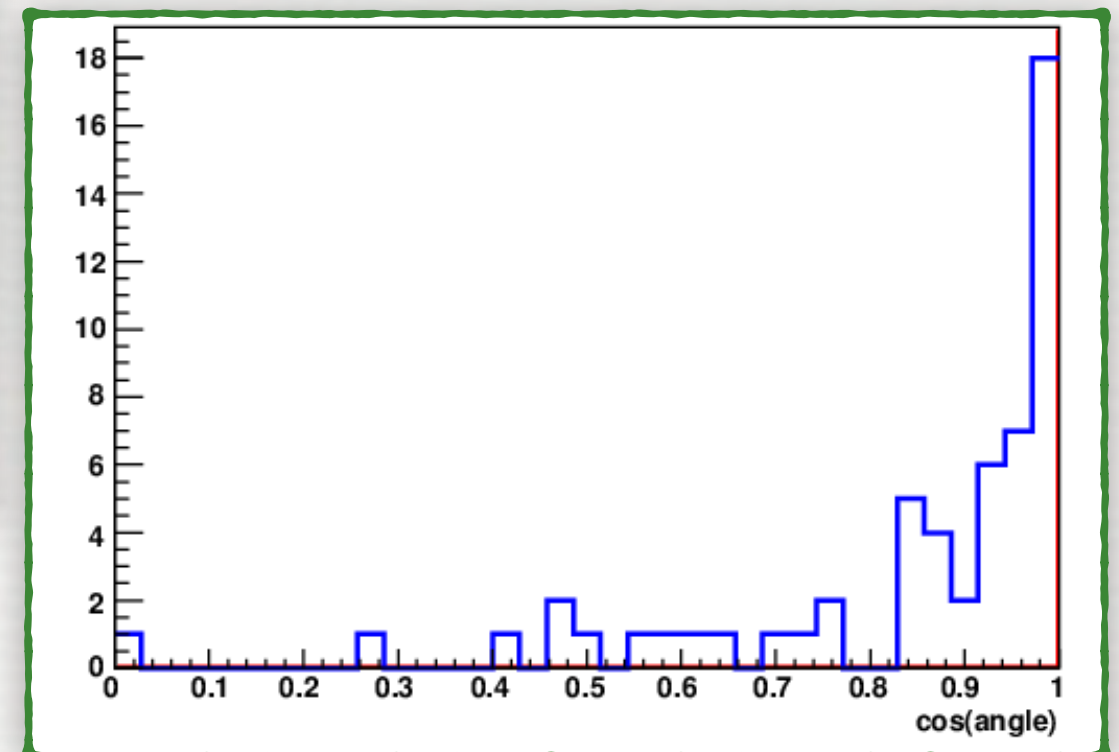
# READOUT STRATEGY

## STEP 2: CANDIDATE VALIDATION

- Scanning with X-ray microscope of preselected zones
- Pin-point check at X-ray microscope of candidate signals selected by optical readout.
- Resolution  $\sim 30$  nm



Matching Efficiency  
**99%**  
(572 / 579)

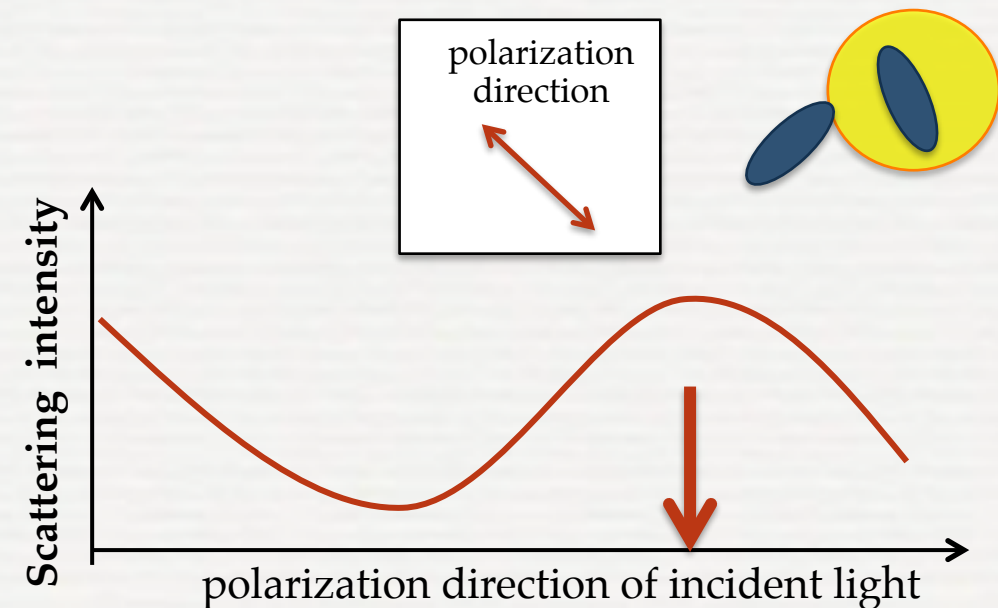
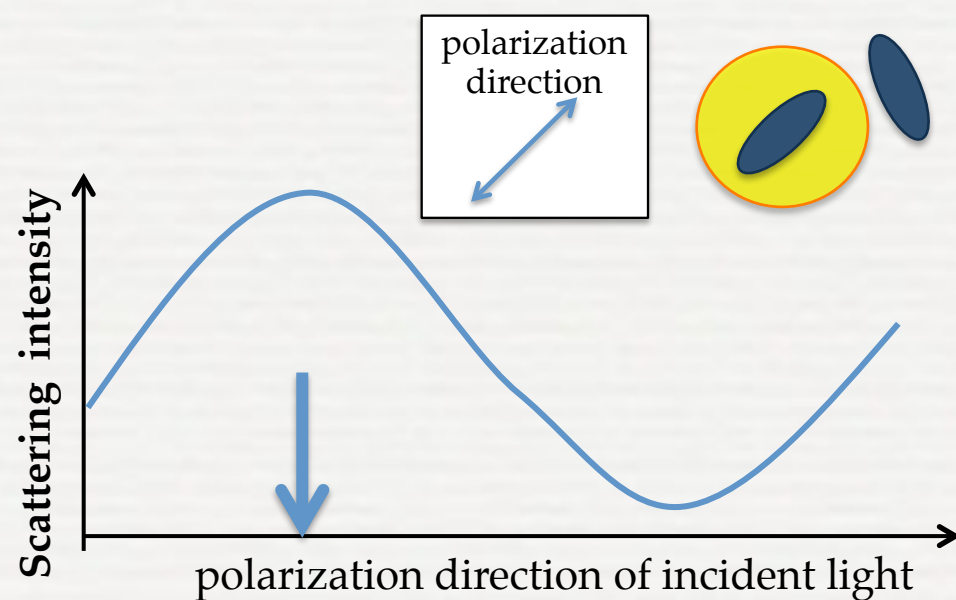


- Slow analysis speed
- Need of external X-ray guns



# RESONANT LIGHT SCATTERING

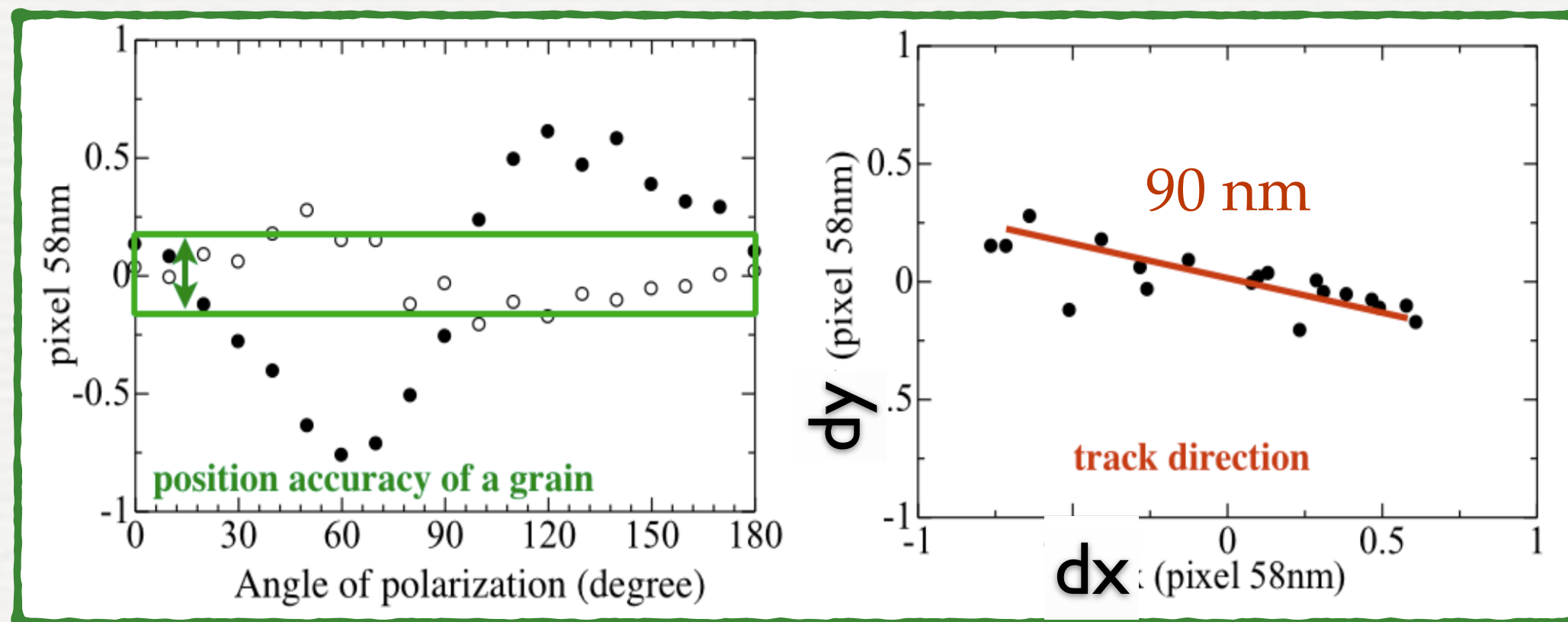
- Occurring when the light is scattering off a nanometric metallic (silver) grains are dispersed in a dielectric medium (*Applied Phys Letters 80 (2002) 1826*)
- Sensitive to the shape of nanometric grains: when silver grains are **not spherical**, the resonant response depends on the polarization of the incident light.
- Each grain is emphasized at different polarization values



- Taking multiple measurements over the whole polarization range produces a displacement of the barycenter of the cluster

# NANOMETRIC TRACK RECONSTRUCTION

- Application of resonant light scattering to an elliptical cluster
- Measure the displacement of cluster barycentre as a function of polarization angle (dx, dy)



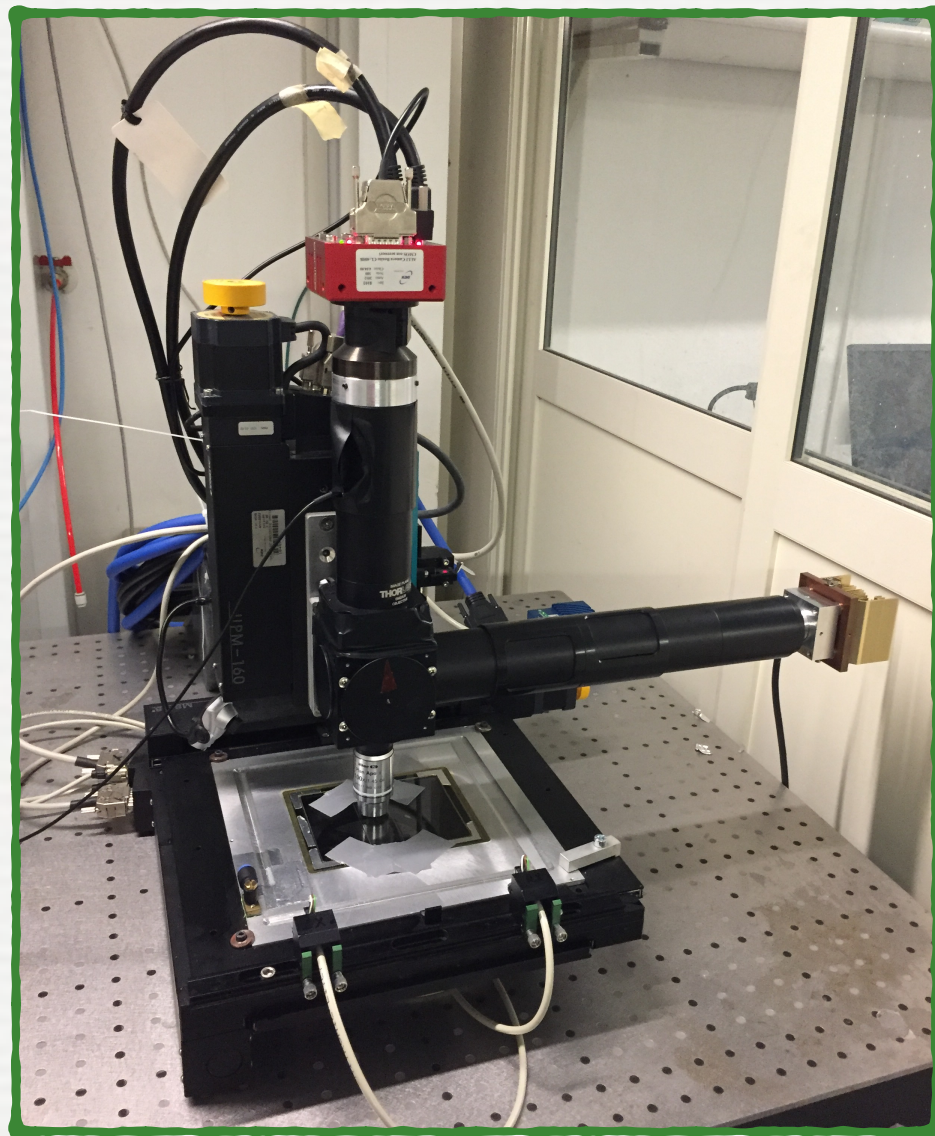
- Measurement of track slope and length

*beyond* optical resolution

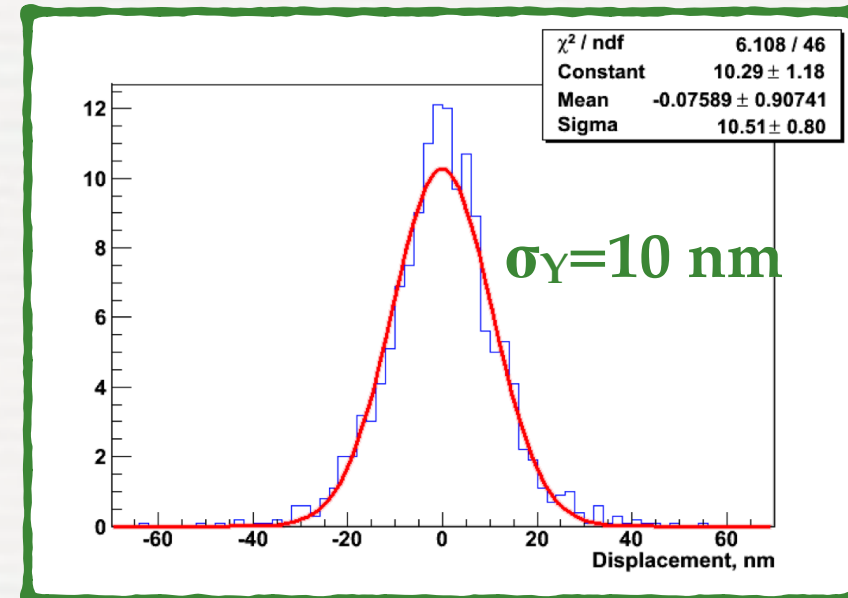
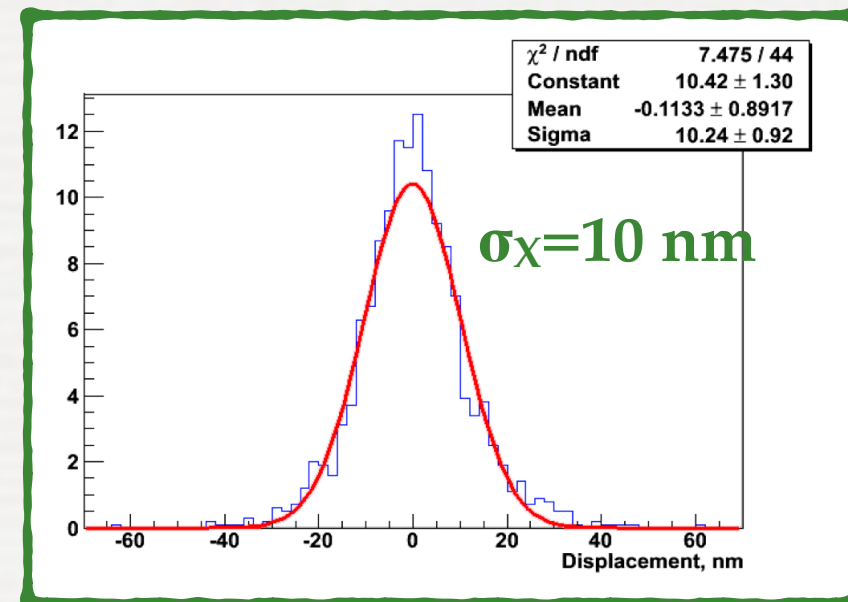


# POSITION ACCURACY

- Optical microscope assembled



- Exploiting resonant light effect



Unprecedented accuracy of **10 nm** achieved on both coordinates

**Breakthrough**



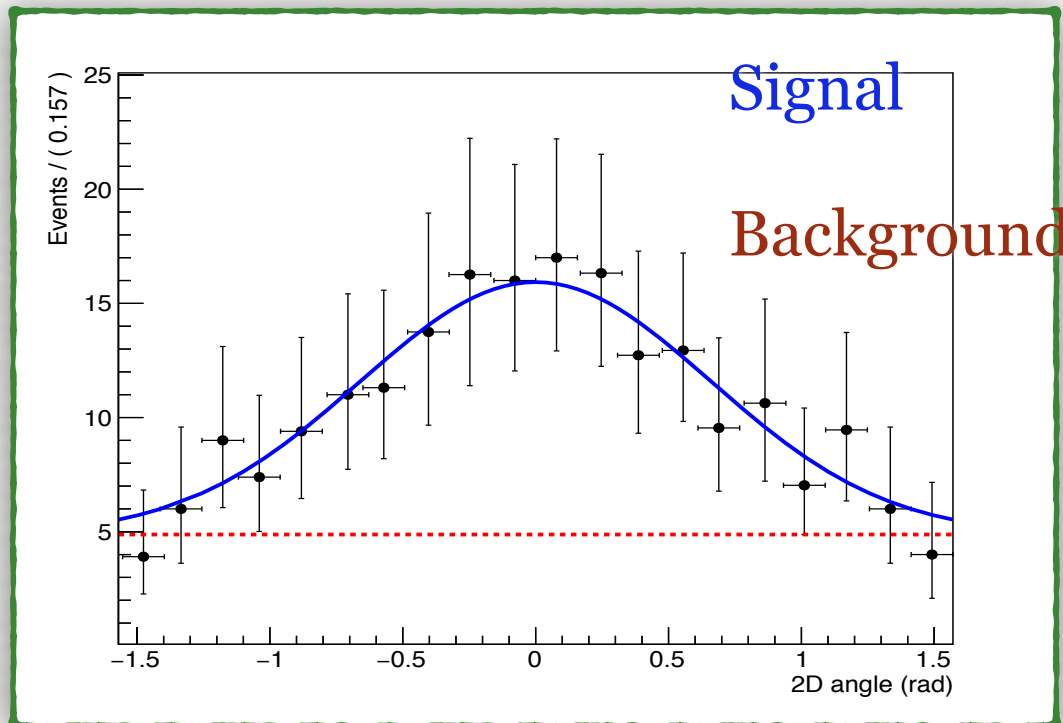
# NEWSdm SENSITIVITY



# EXPLOIT DIRECTIONALITY

- Evaluation of upper limit and sensitivity based on the profile likelihood ratio test

- Mass= 10 kg
- Exposure time = 10 years
- $N_{\text{background}} = 100$
- Threshold = 100 nm

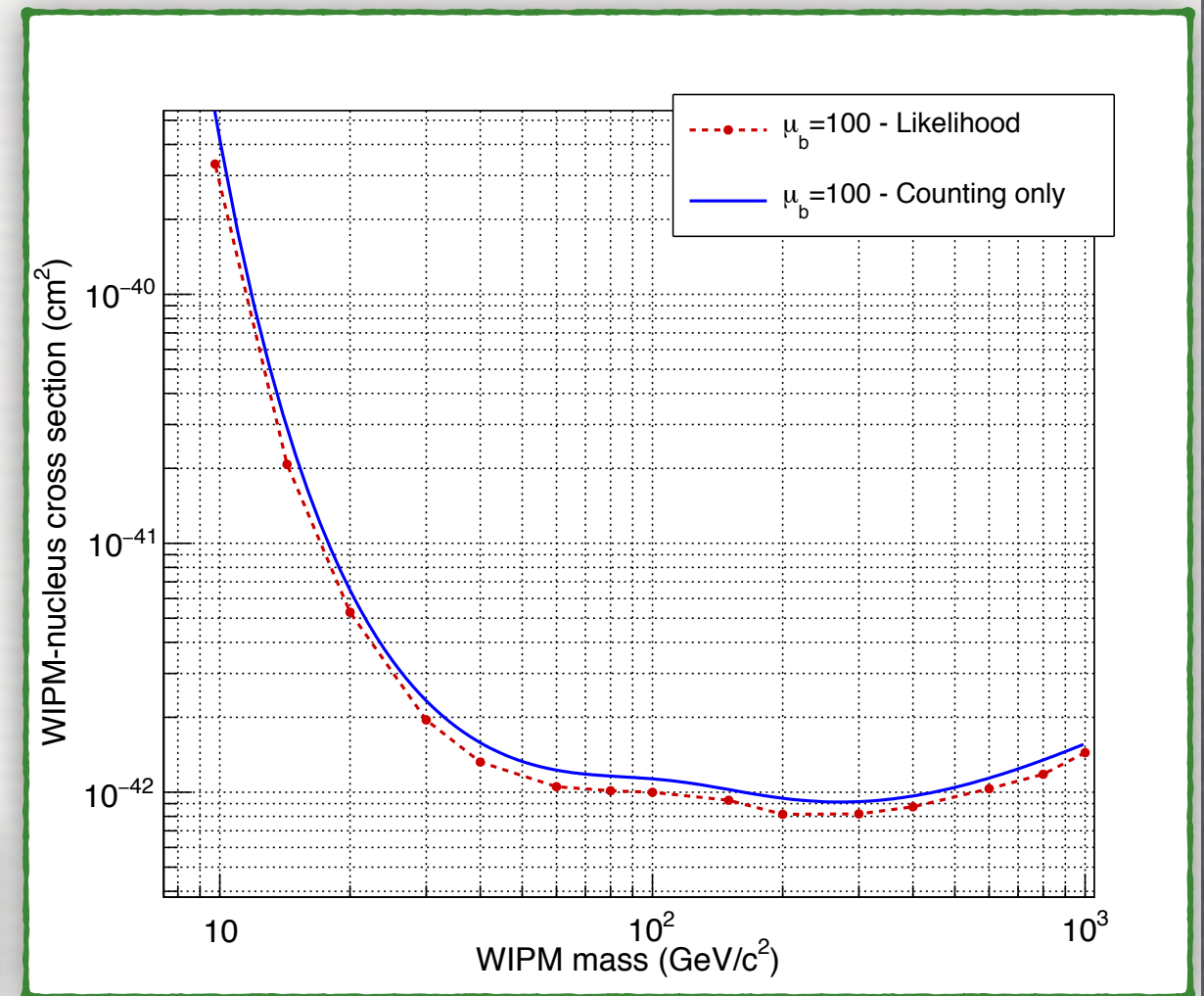


- Likelihood function

expected number of WIMP events  $\mu_\chi$       expected number of background events  $\mu_b$       signal pdf  $f_\chi$       background pdf  $f_b$

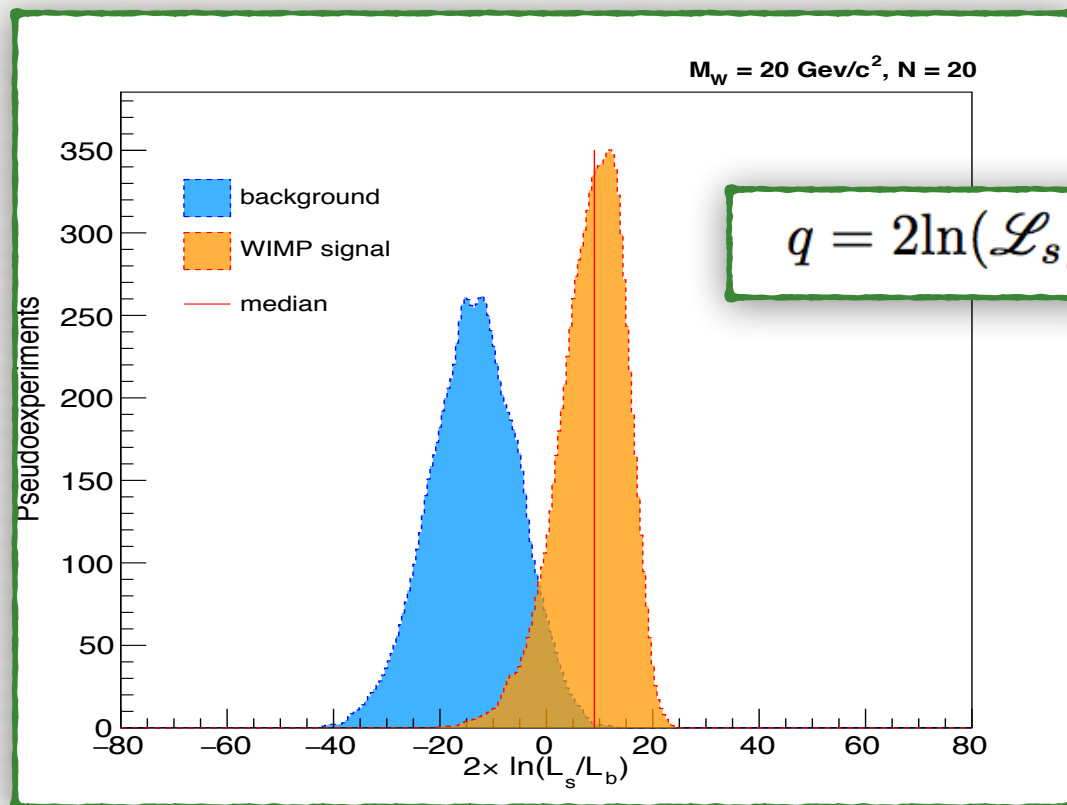
$$\mathcal{L}(\sigma_{\chi-n}, R_b) = \frac{e^{-(\mu_\chi + \mu_b)}}{N!} \times \prod_{i=1}^N [\mu_\chi f_\chi(\vec{q}_i; t_i) + \mu_b f_b(\vec{q}_i)]$$

total number of observed events  $N$       set of observables  $\vec{q}_i$

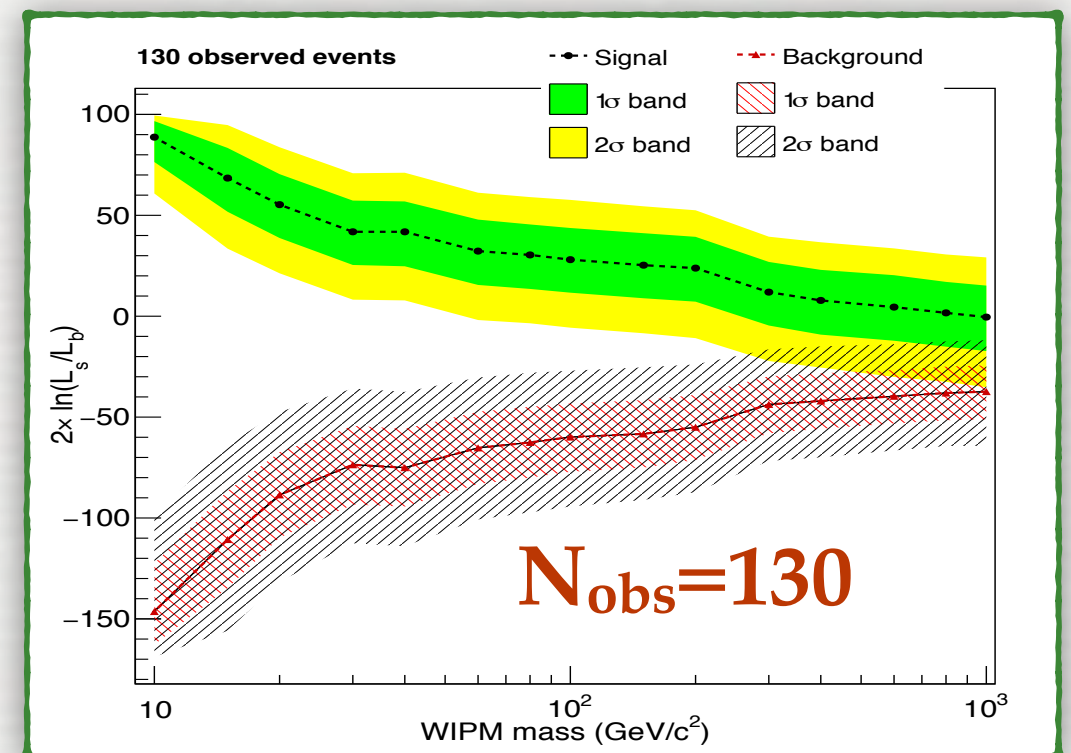
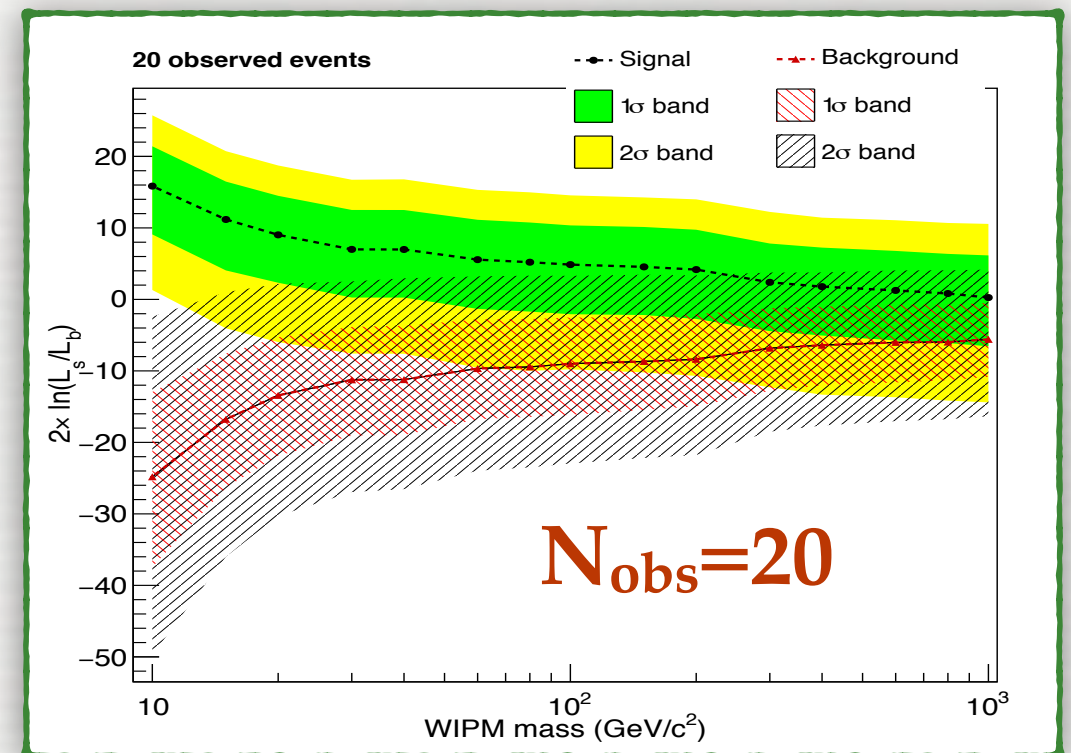


# WIMP SIGNAL IDENTIFICATION

- Test anisotropy of observed signal
- Unambiguous proof of WIMP origin of recoil signal
- Signal/background hypothesis separation



- 20 events required to prove that data are not compatible with background at  $3\sigma$  CL for  $M_W < 20 \text{ GeV}/c^2$
- 130 events give  $3\sigma$  CL in the whole WIMP mass range





# TOWARDS NEUTRINO FLOOR

- Discrimination based on measurement of recoil direction
- Unique possibility to search for WIMP signal beyond “neutrino floor”

Neutrino coherent scattering indistinguishable from WIMP interactions

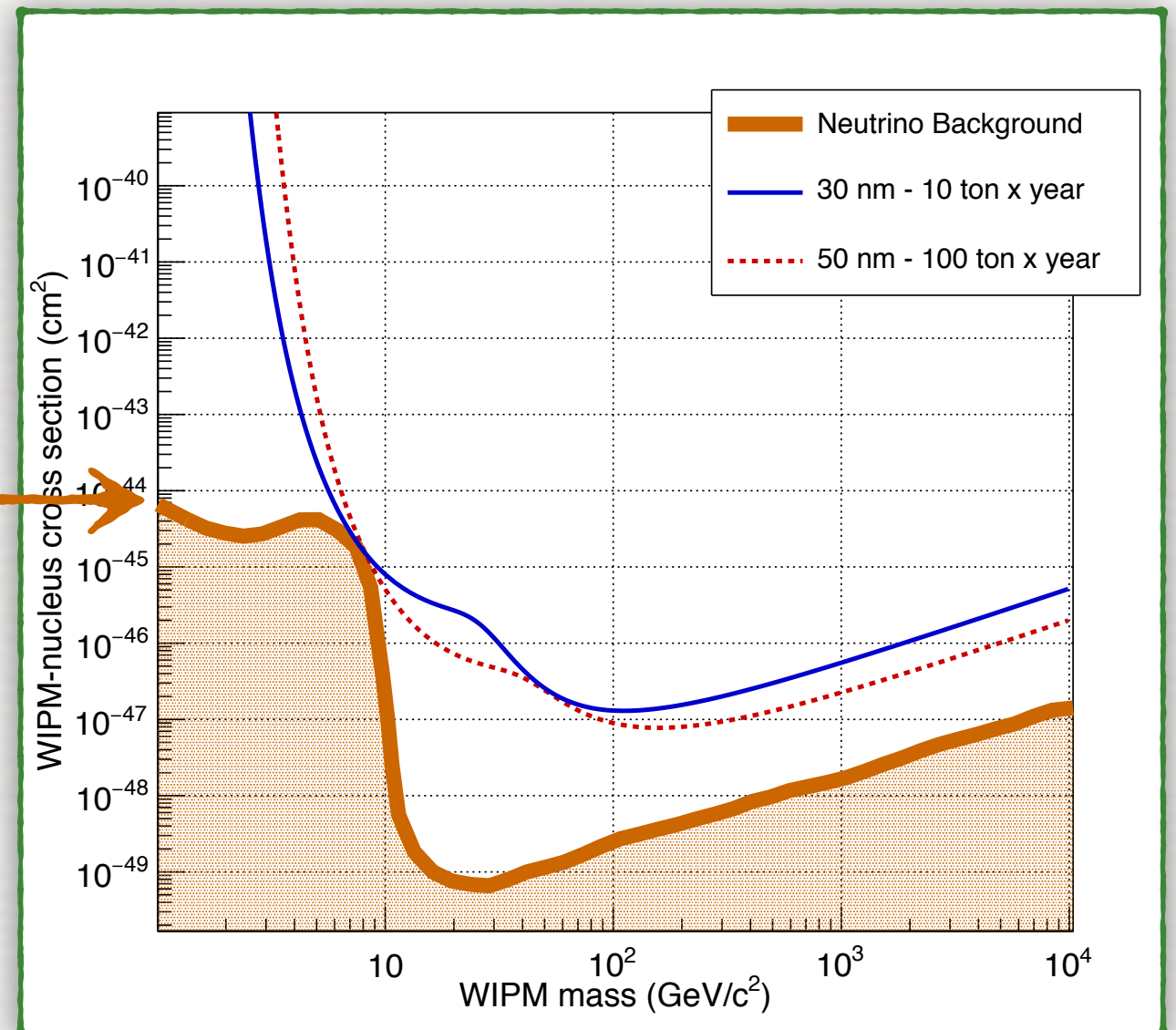
*Phys.Rev.D89 (2014) no.2, 023524*  
(Xe/Ge target)

## REQUIREMENTS

- Larger mass scale detector
- Reduction of track length threshold

The neutrino bound is reached with:

- ➔ 10 ton x year exposure if 30 nm threshold
- ➔ 100 ton x year exposure if 50 nm threshold



# CURRENT STATUS OF THE EXPERIMENT

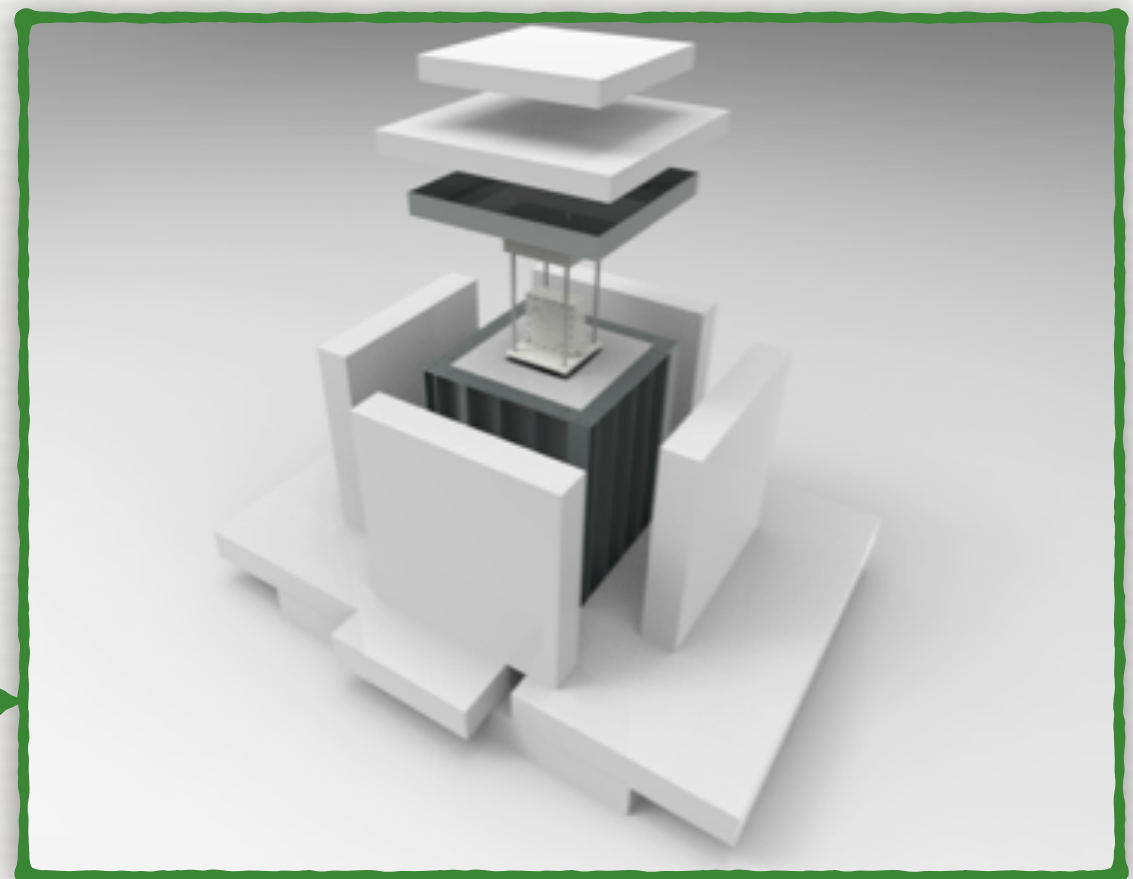


# TECHNICAL TEST

- **Aim**: measure the detectable background from environmental and intrinsic sources and validate estimates from simulations
- Confirmation of a negligible background will pave the way for the construction of a **pilot experiment** with an exposure on the **kg year** scale
- Pilot experiment will act as a **demonstrator** to further extend the mass range
- **Experimental setup**:
  - shield from environmental background
  - cooling system to ensure required temperature to NIT emulsions

Polyethylene slabs 40 cm-thick -  
absorb environmental and  
cosmogenic neutrons

Lead bricks 10 cm-thick - absorb  
environmental photons





# TECHNICAL TEST



- Installed in Underground Gran Sasso INFN Laboratories in March 2017



# CONCLUSIONS

- A novel approach for **directional Dark Matter searches** is proposed in NEWSdm
- Use of fine-grained **nuclear emulsion** as target and tracking system
- Breakthrough in readout technologies to go beyond optical resolution
- Neutron background from intrinsic radioactivity negligible up to  $\sim 10$  kg year
- Prepare a kg scale (pilot) experiment as a demonstrator of the technology
- Aim: large mass scale detector to go beyond “neutrino floor”
- Status:
  - Letter of Intent submitted to LNGSC in 2015
  - Technical test in progress
  - TDR in preparation



\*part of the Collaboration when  
test started in LNGS



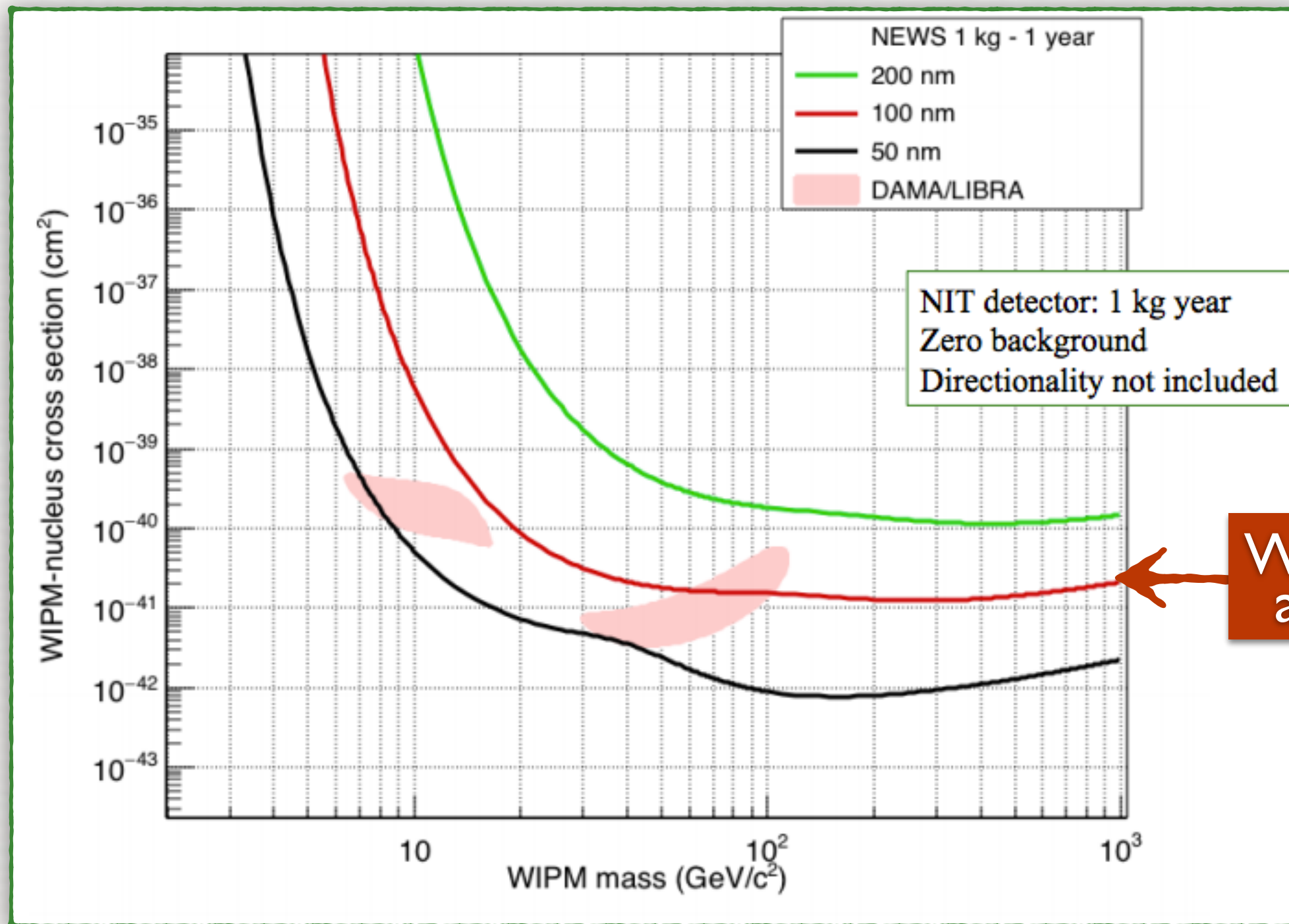
THANK YOU FOR YOUR ATTENTION



# BACKUP SLIDES

# EXCLUSION PLOT

- Pilot experiment: 1 kg year





# BACKGROUND STUDIES

# BACKGROUND STUDIES

- Measurement of intrinsic radioactivity: neutrons

Nuclide	Contamination [ppb]	Activity [mBq/Kg]
Gelatine		
<sup>232</sup> Th	2.7	11.0
<sup>238</sup> U	3.9	48.1
PVA		
<sup>232</sup> Th	< 0.5	< 2.0
<sup>238</sup> U	< 0.7	< 8.6
AgBr-I		
<sup>232</sup> Th	1.0	4.1
<sup>238</sup> U	1.5	18.5



<sup>238</sup>U: 1.87 ppb (23.1 mBq/kg)

<sup>232</sup>Th: 1.26 ppb (5.1 mBq/Kg)

Background yield from the intrinsic radioactive contamination of NIT:

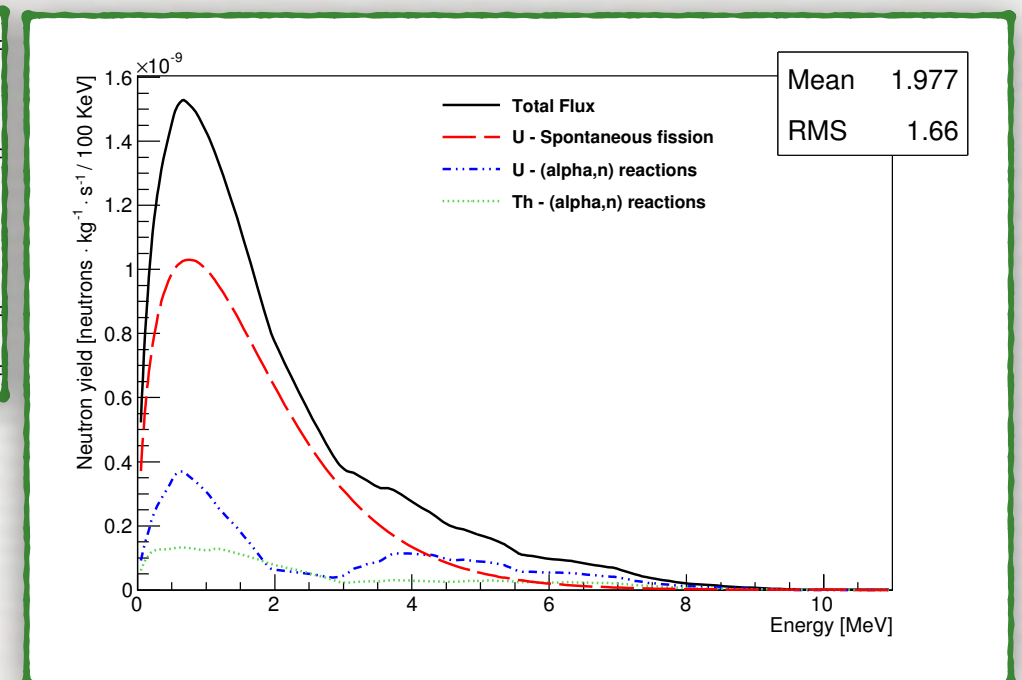
**~1.2 n/kg year**

Process	SOURCES simulation [n · kg <sup>-1</sup> · y <sup>-1</sup> ]	Semi-analytical calculation [n · kg <sup>-1</sup> · y <sup>-1</sup> ]
(α, n) from <sup>232</sup> Th chain	0.12±0.04	0.10±0.03
(α, n) from <sup>238</sup> U chain	0.27±0.08	0.26±0.08
Spontaneous fission	0.79±0.24	0.82±0.24
Total flux	1.18±0.35	1.18±0.35

From simulation: detectable neutron induced background

$\epsilon \sim 1\% \rightarrow \sim 0.01 \text{ n/kg year}$

Neutron background from intrinsic radioactivity negligible up to ~10 kg year



NEWSdm Collaboration

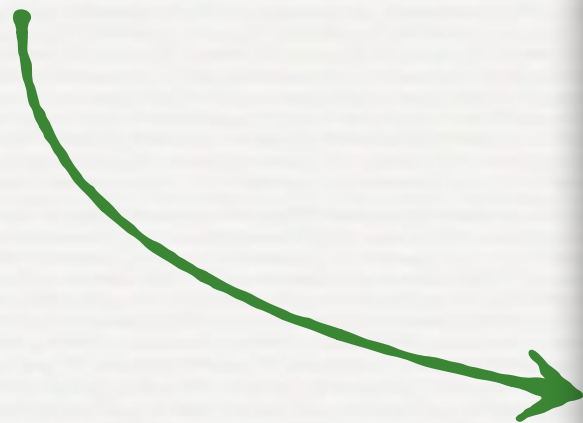
Astroparticle Physics 80 (2016) 16



# BEYOND OPTICAL RESOLUTION

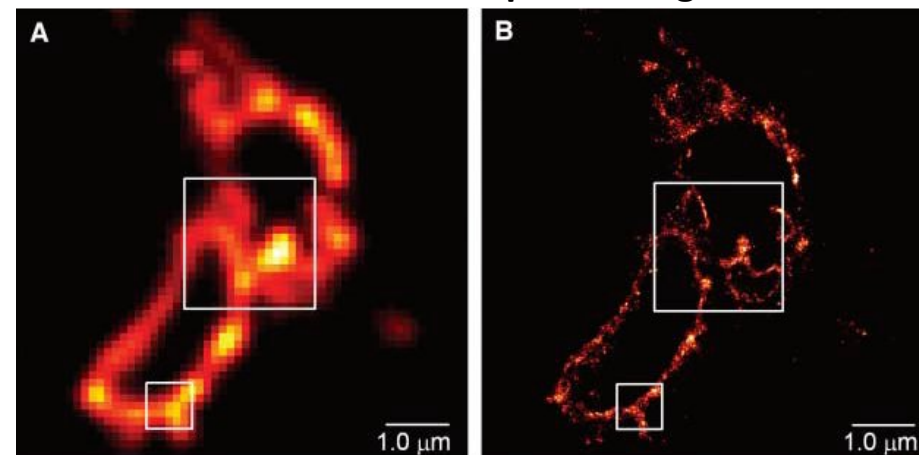
## OPTICAL MICROSCOPES

- New technologies



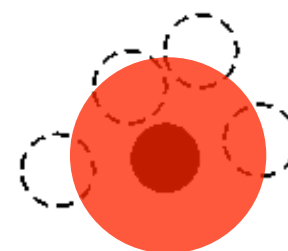
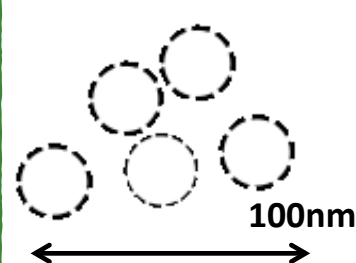
### Imaging beyond the optical resolution 2014 Nobel Prize in Chemistry

COS-7 cell optical images



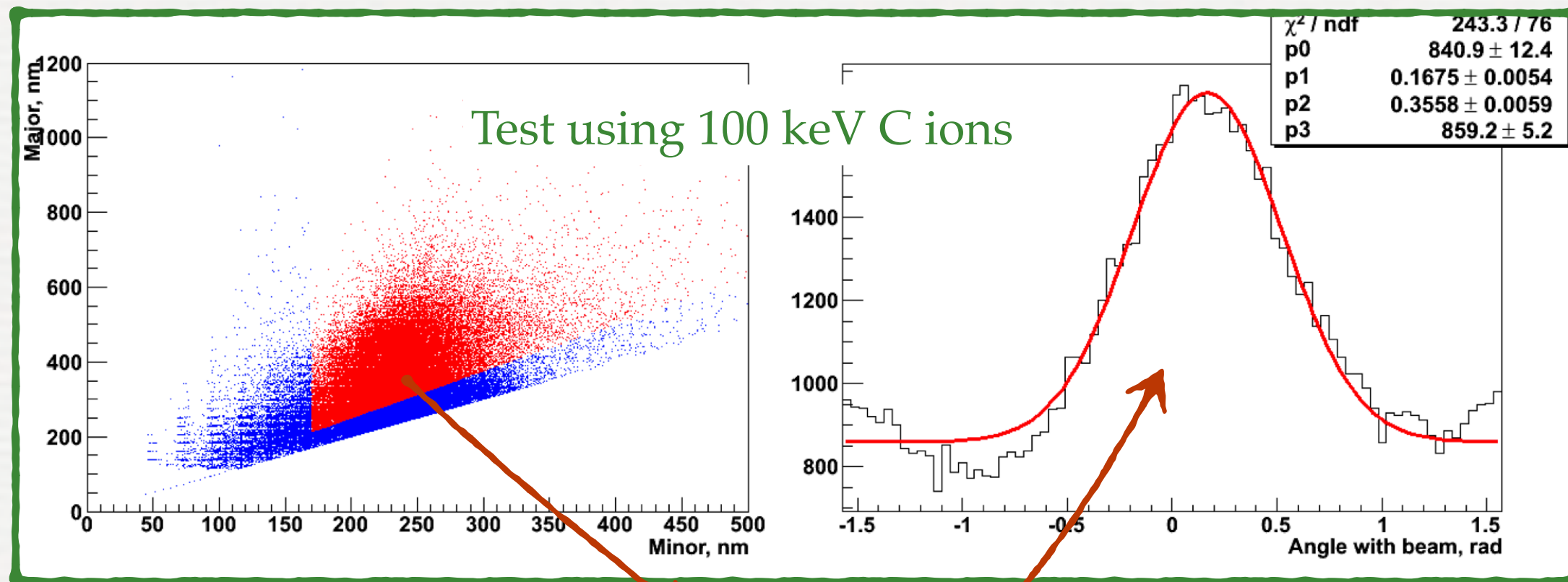
**Fluorescent molecule**

Eric Betzig *et al.*, Science 313, 1642 (2006)



Using fluorescence

# SELECTION OF TRACKS WITH SHAPE ANALYSIS



## SIGNAL SELECTION

- Major axis / minor axis > 1.25
- minor axis > 170 nm

## OVERALL ANGULAR RESOLUTION

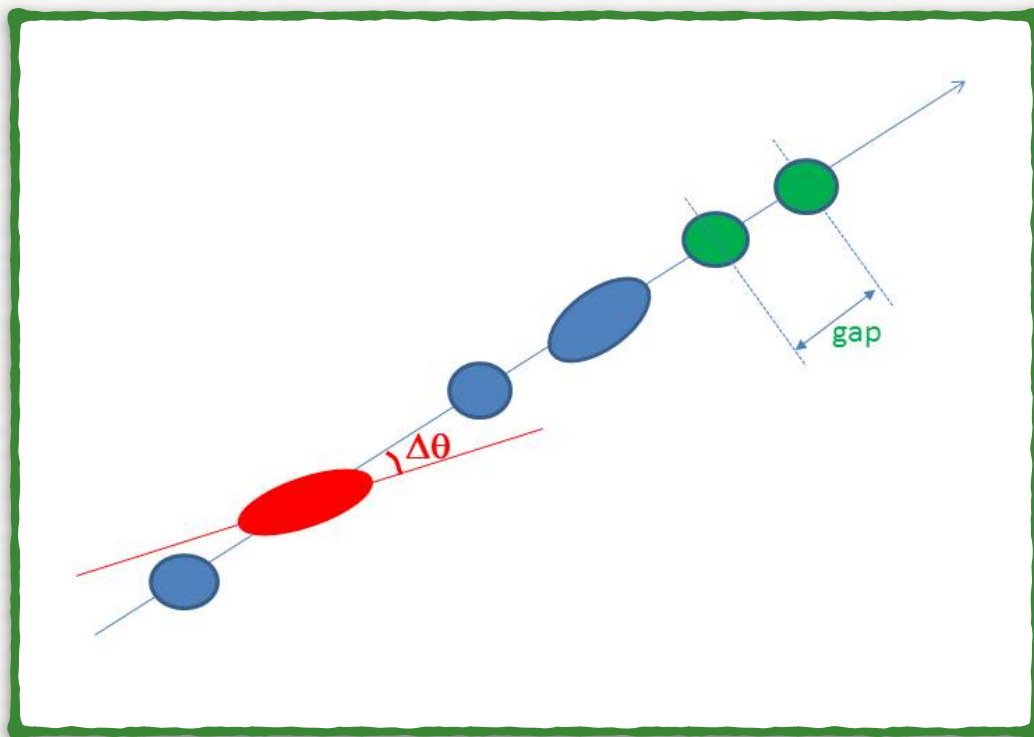
$$\sigma^2 = \sigma^2_{\text{intrinsic}} + \sigma^2_{\text{scattering}}$$

$$\sigma = 360 \text{ mrad}$$



# INTRINSIC ANGULAR RESOLUTION

- Neutron test beam sample: exposure at FNS (Japan)
- Compare clusters with elliptical ( $e > 1.1$ ) shape with the proton recoil direction
- Scattering contribution negligible



INTRINSIC ANGULAR RESOLUTION

$$\sigma = 235 \text{ mrad} = 13^\circ$$

