

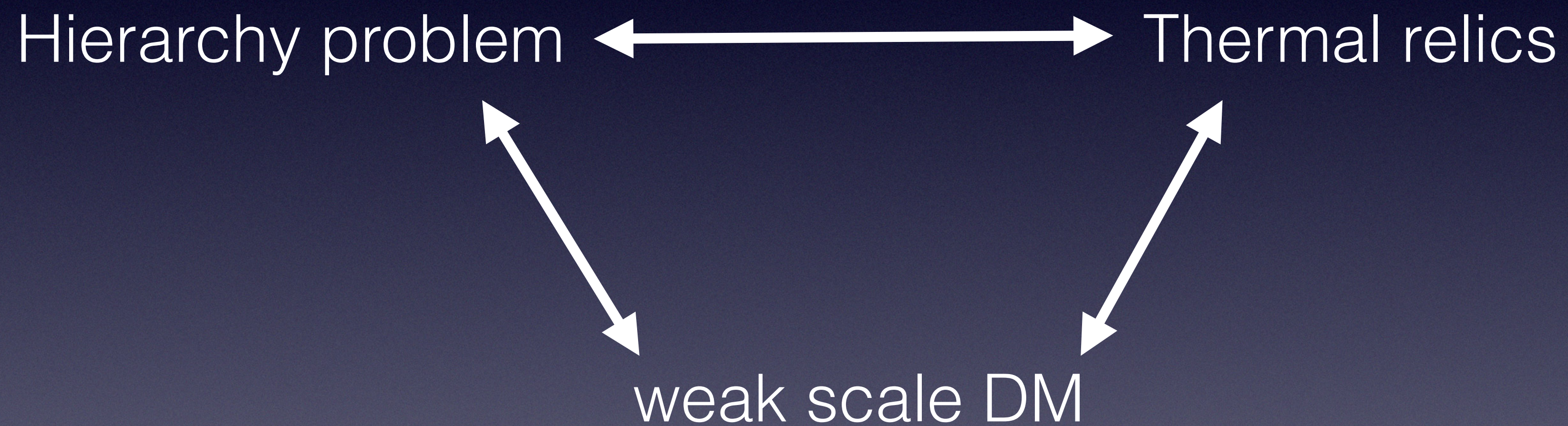
# New Priors in Hidden Physics

Neal Weiner  
Cosmic Visions Workshop  
March 23, 2017

# The era of strong priors 199x~2016

- Hierarchy problem
- Weak scale DM
- Questions of the SM (unification, neutrino mass, strong CP...)

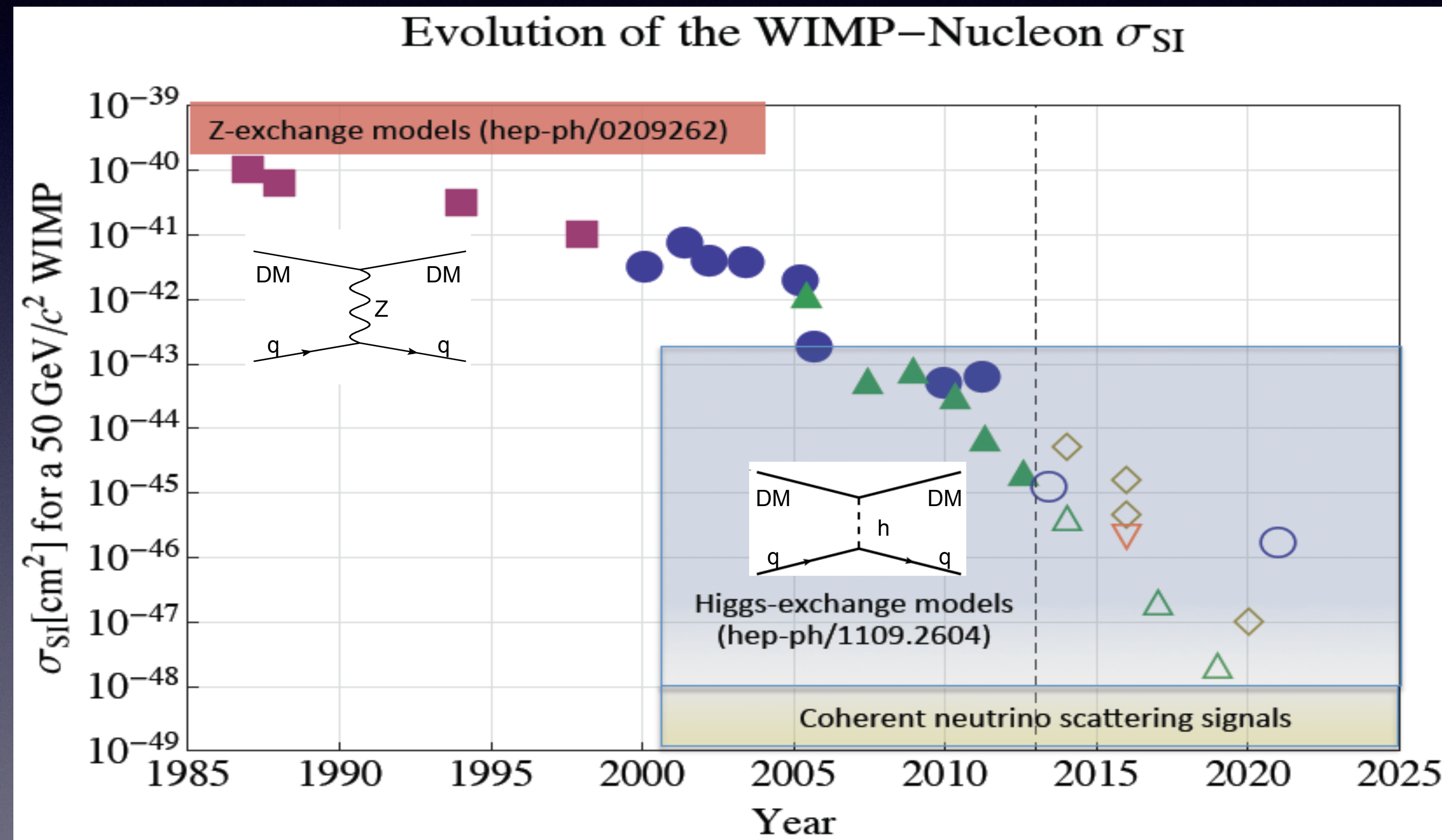
# Dark matter in the era of strong priors



“WIMP” miracle made (and makes) a motivating case

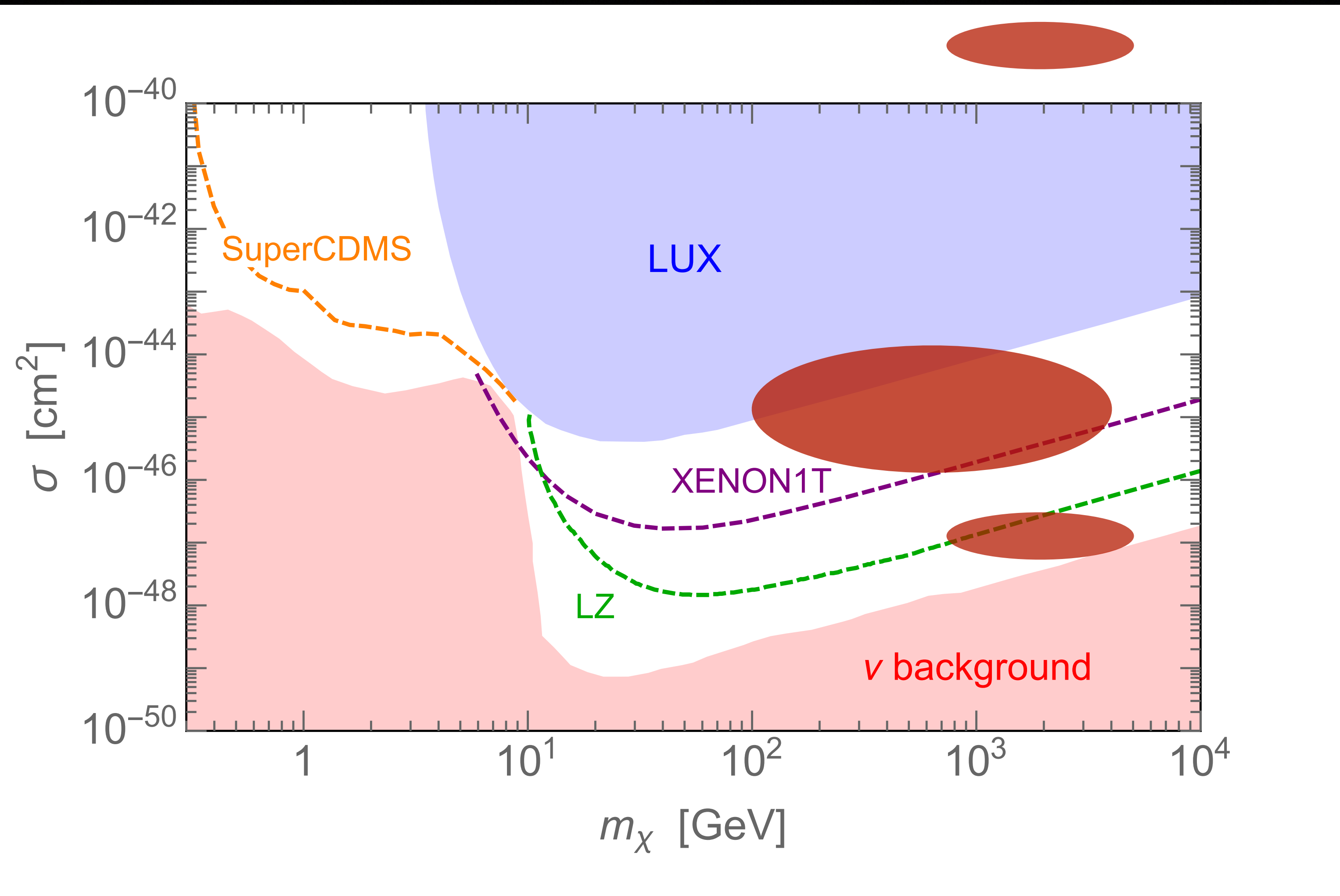
breaking down priors

# rapid progress in WIMP search



slide from J Feng

Plot from Josh Ruderman



No sign of WIMPs so far



# Neutrinos/GUTs

- Neutrino mass guess: few  $\times 10^{-3}$
- Neutrino mixing guess: small (CKM)
- Proton decay?



We have pursued scenarios under  
very strong assumptions

- Where do we go from here?

# Moving beyond the era of strong priors

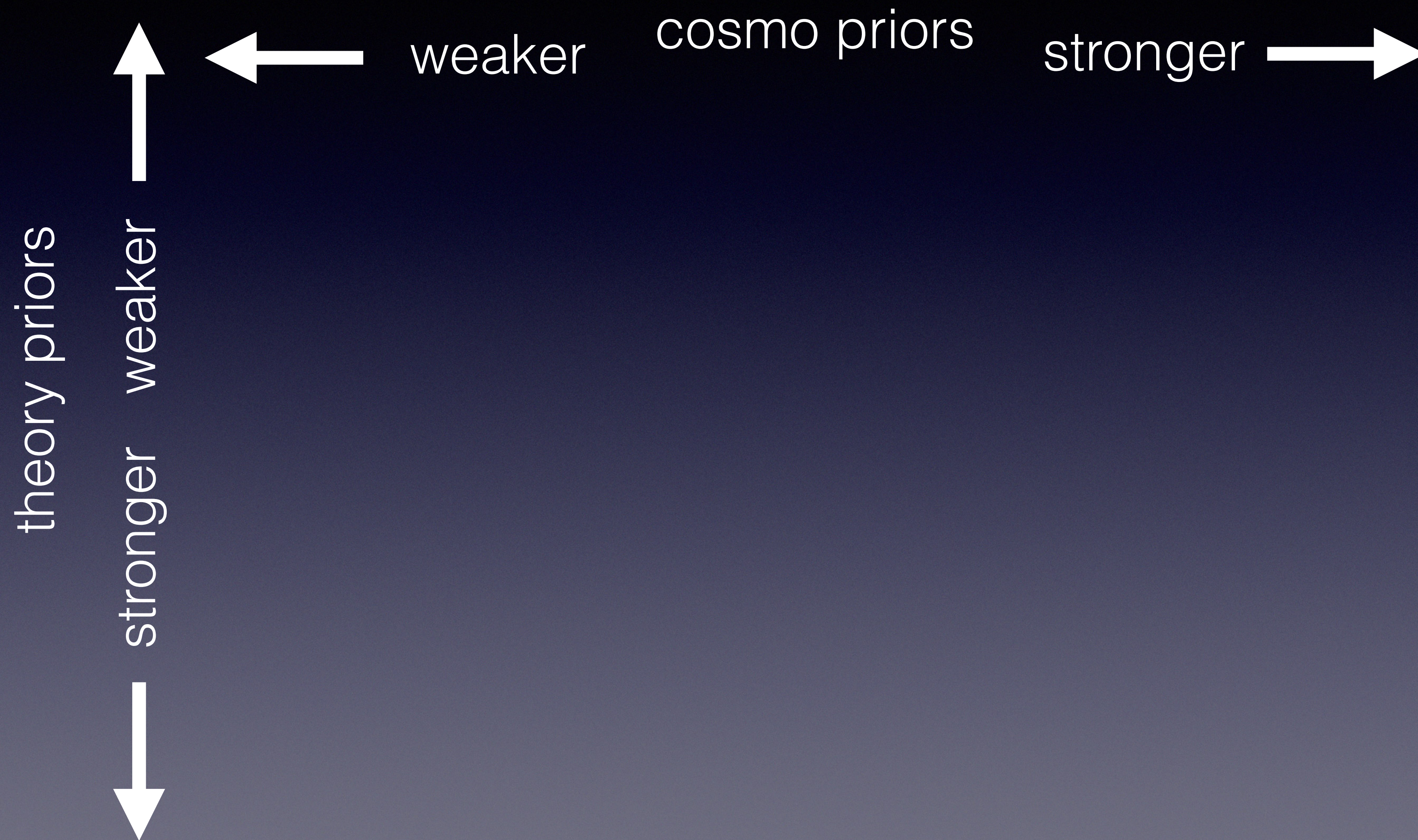
- No priors?
- Weak priors?
- New priors?

We know there is physics beyond the standard model (DM)!

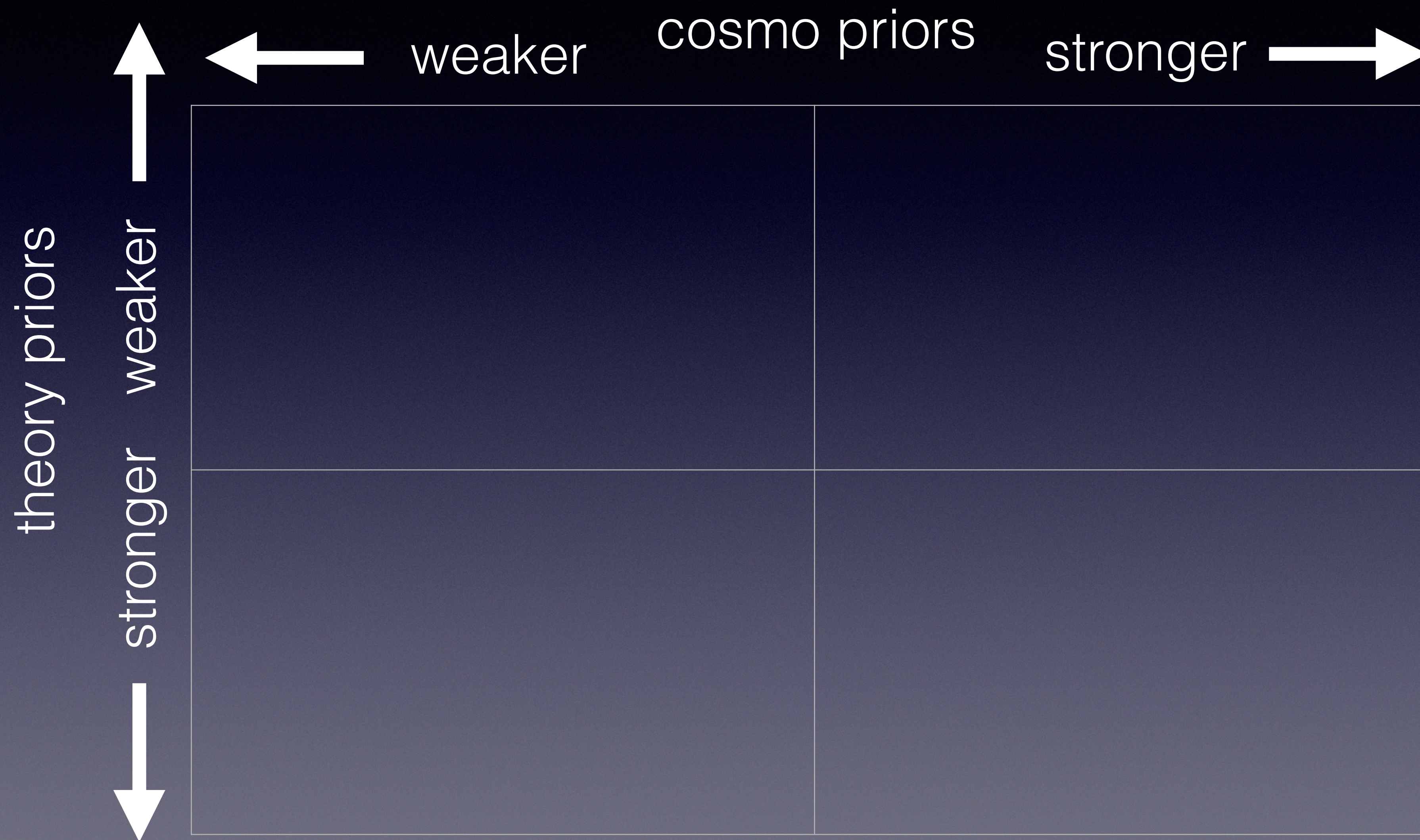
# Example: axion

- weak priors: EFT of scalars and pseudoscalars coupling to SM
- moderate priors: QCD axion
- strong priors: QCD axion + standard cosmology

# The Priorhedron



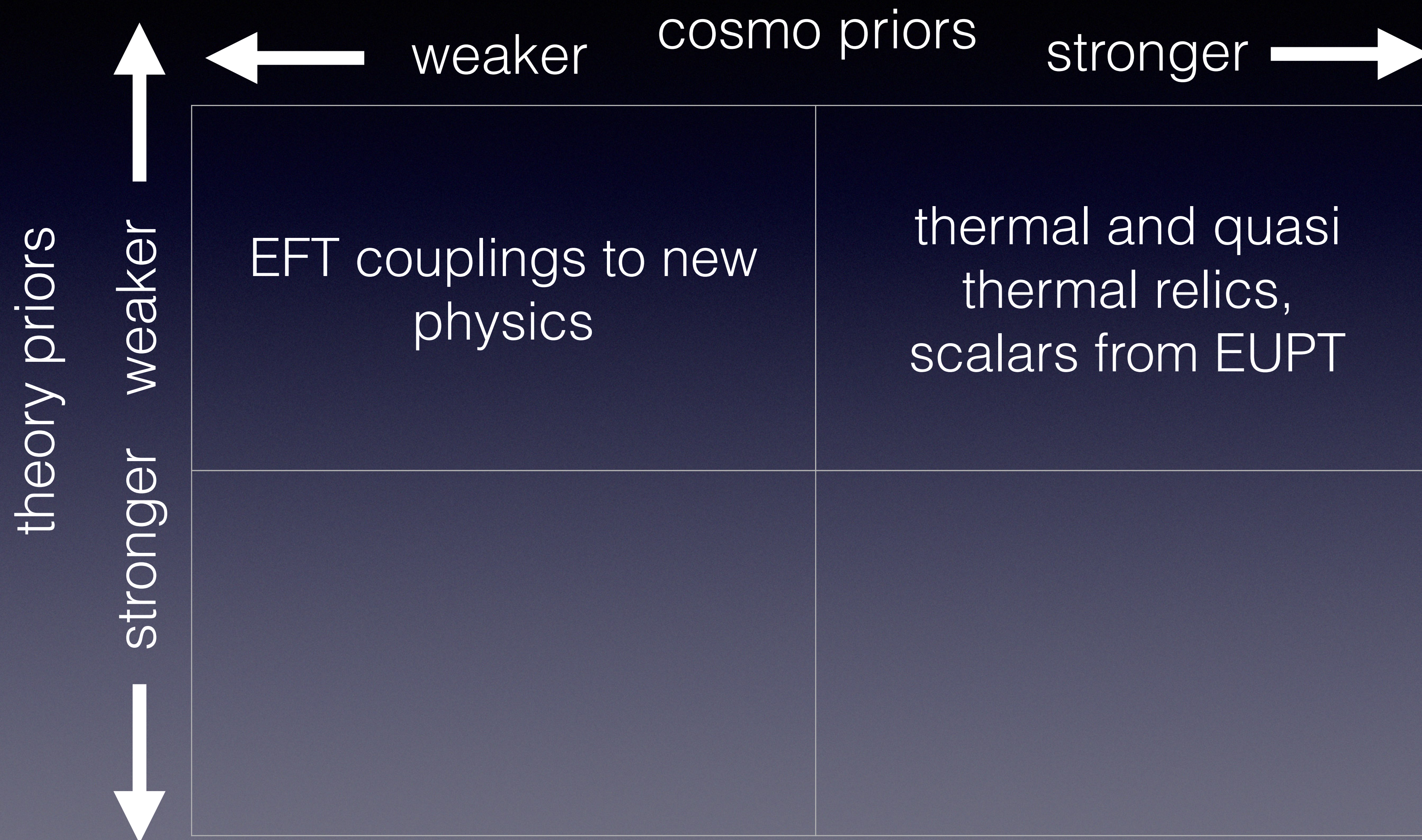
# The Priorhedron



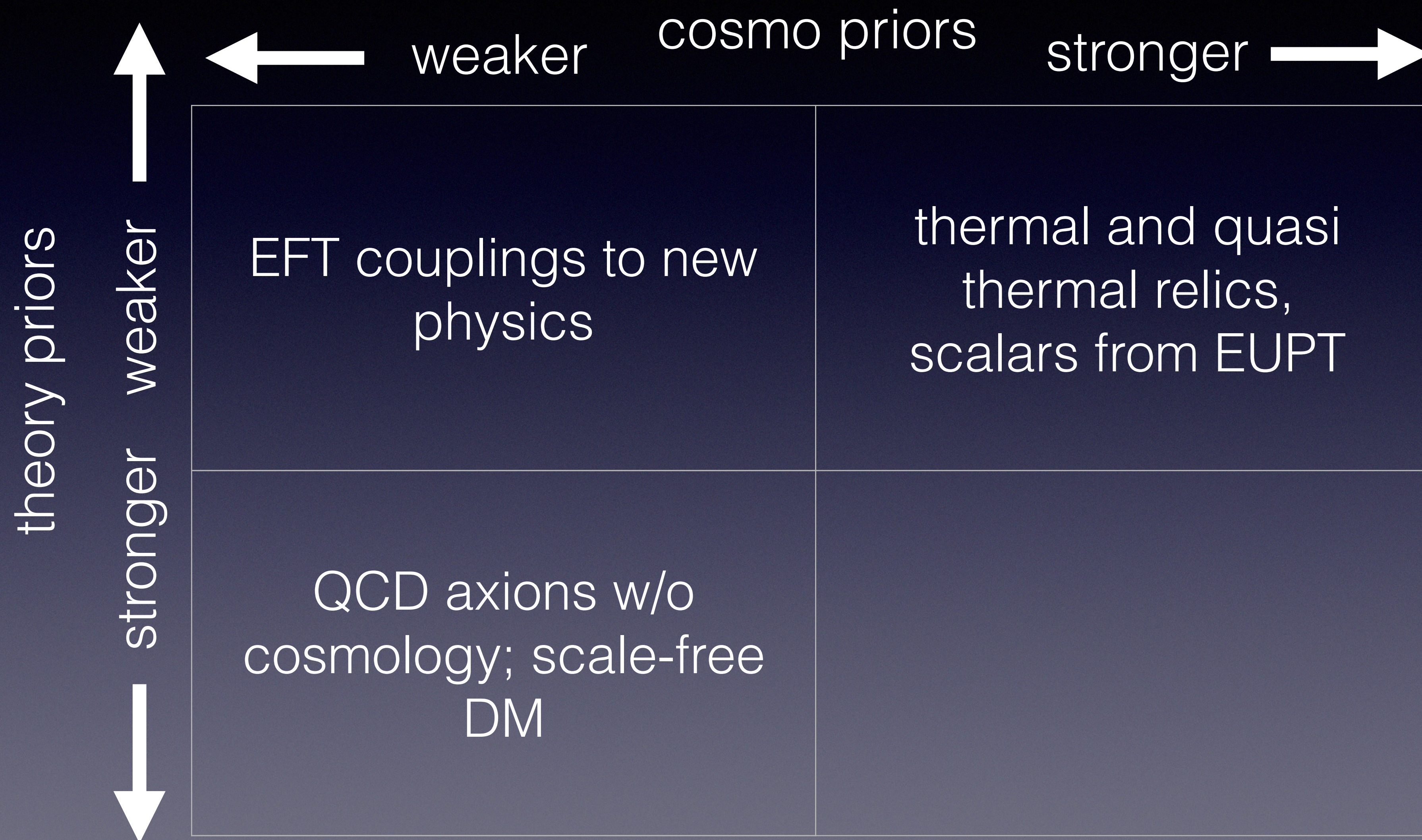
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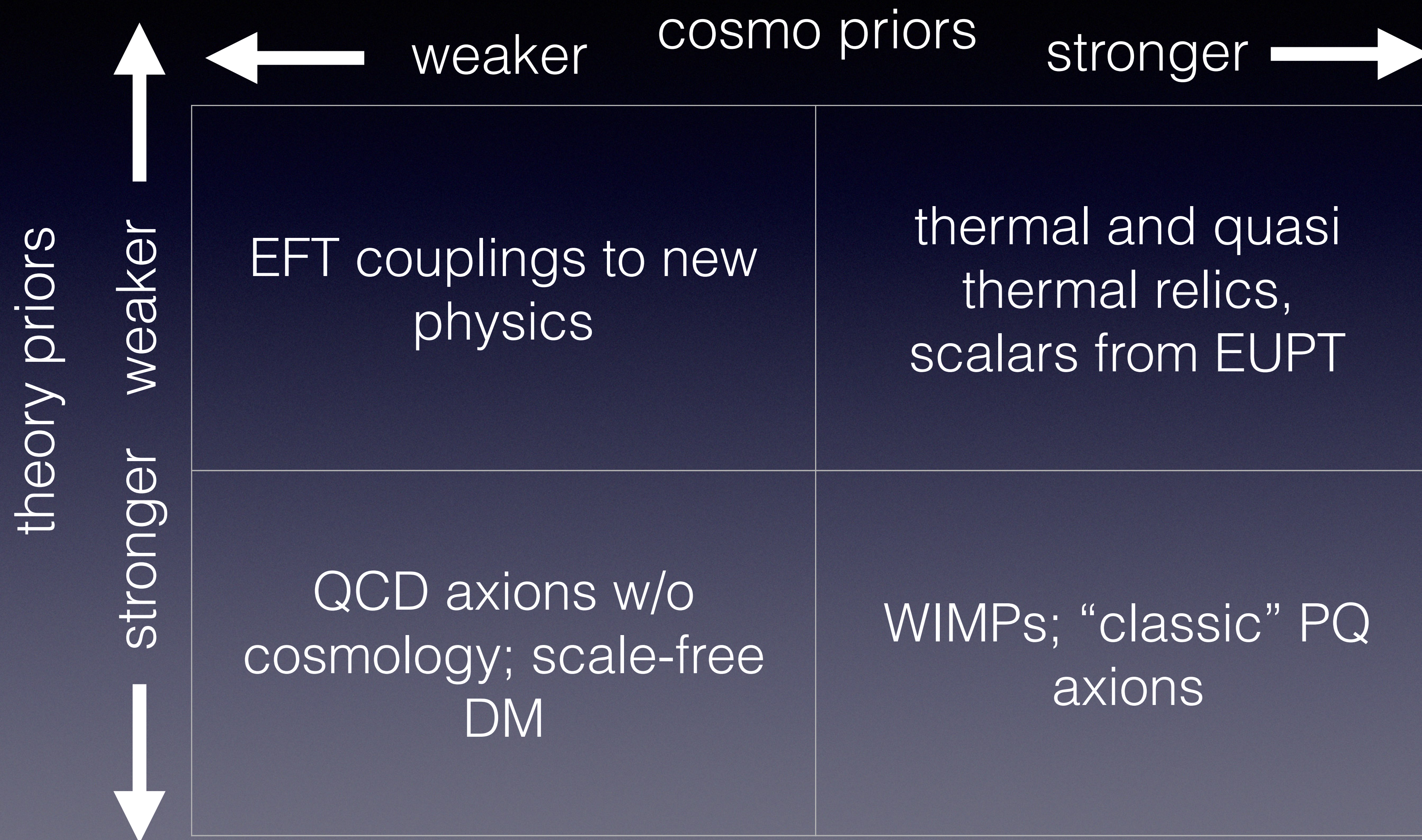


# The Priorhedron





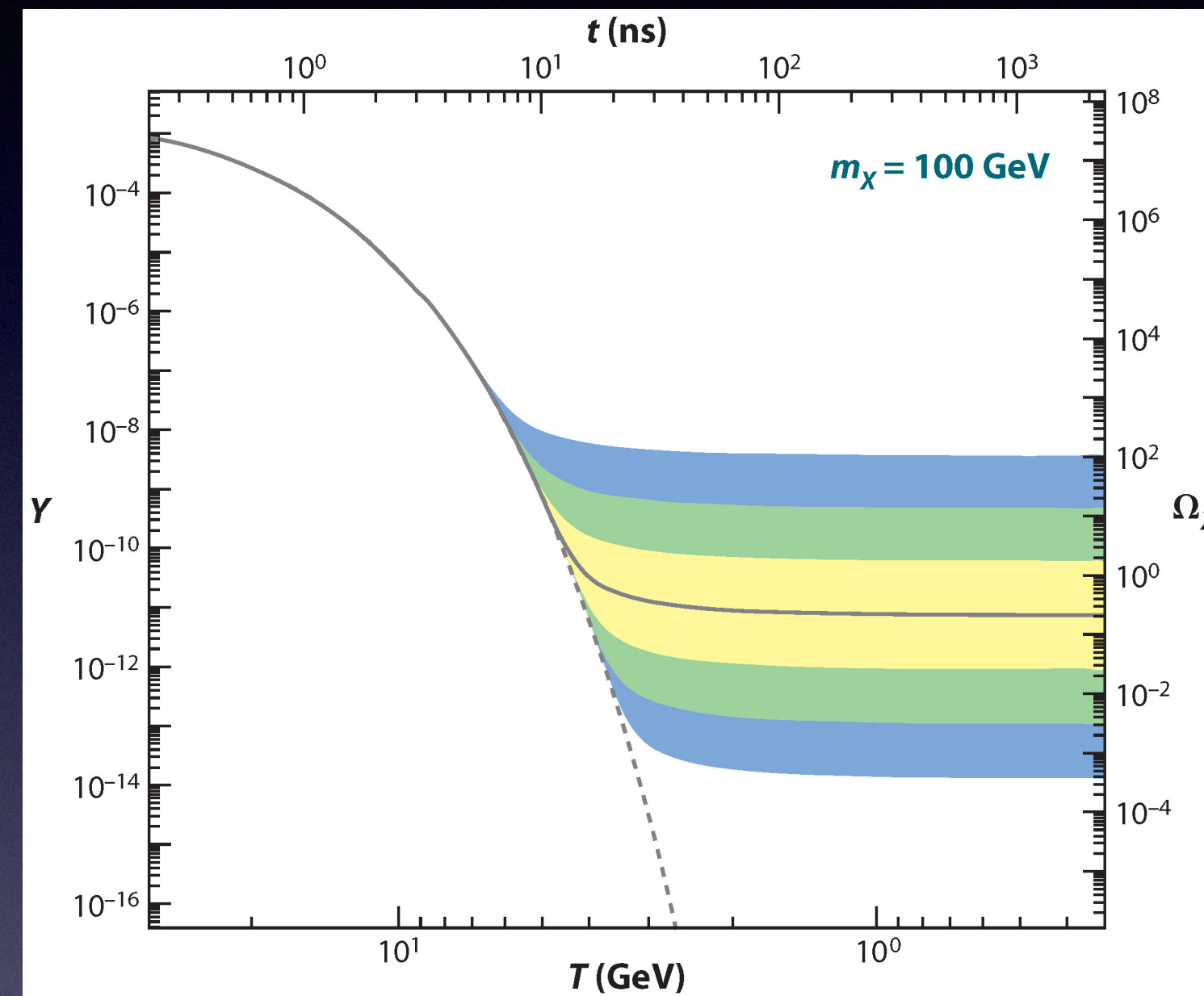
# The Priorhedron



# BSM in the era of moderate priors

- Opportunity to ask broader questions
- Can't simply be fishing expedition (although IMHO a certain amount of cost effective fishing is important)
- Take one step back on some prior axis and find target regions
  - e.g., consider a thermally connected particle
  - a broader class of axion like particles

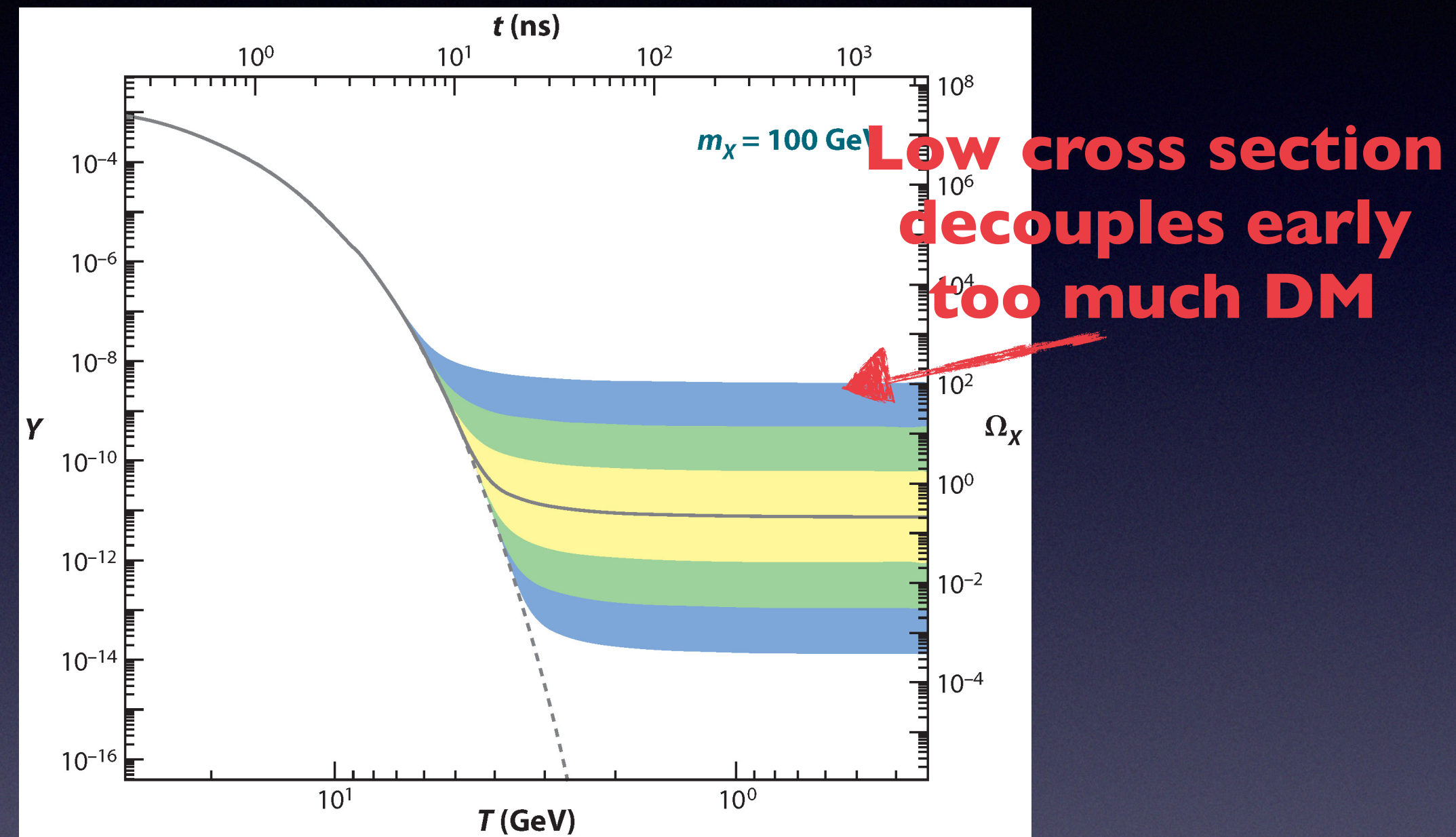
# The story of the WIMP



For a thermal relic, you learn precisely one number, namely the annihilation cross section

$$\begin{aligned} \langle \sigma v \rangle_{ann} &\approx 3 \times 10^{-26} \text{ cm}^3 \text{ sec}^{-1} \\ &\approx \frac{\alpha^2}{(200 \text{ GeV})^2} \end{aligned}$$

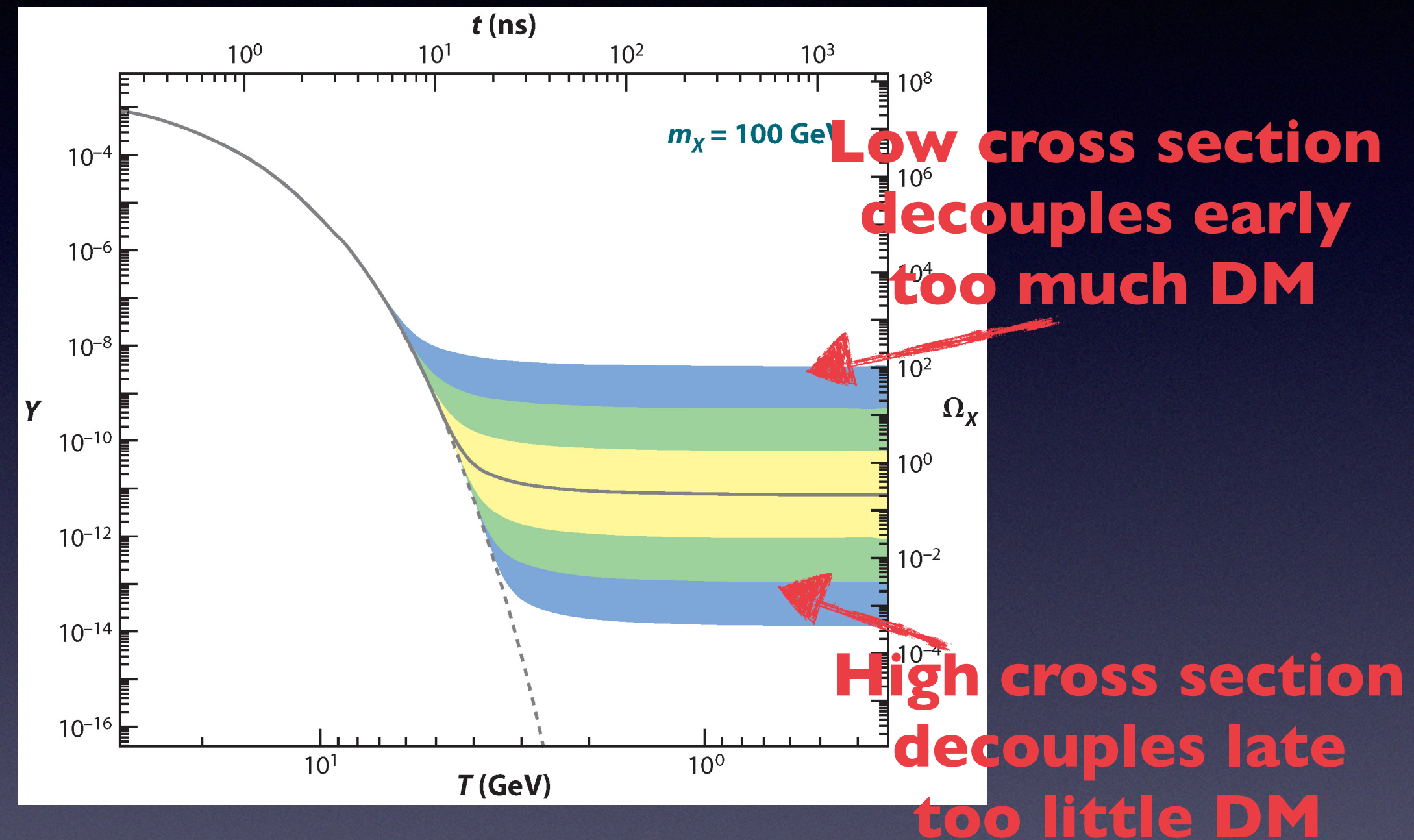
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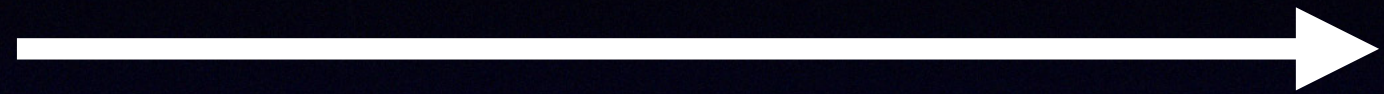
# Broadening the scope



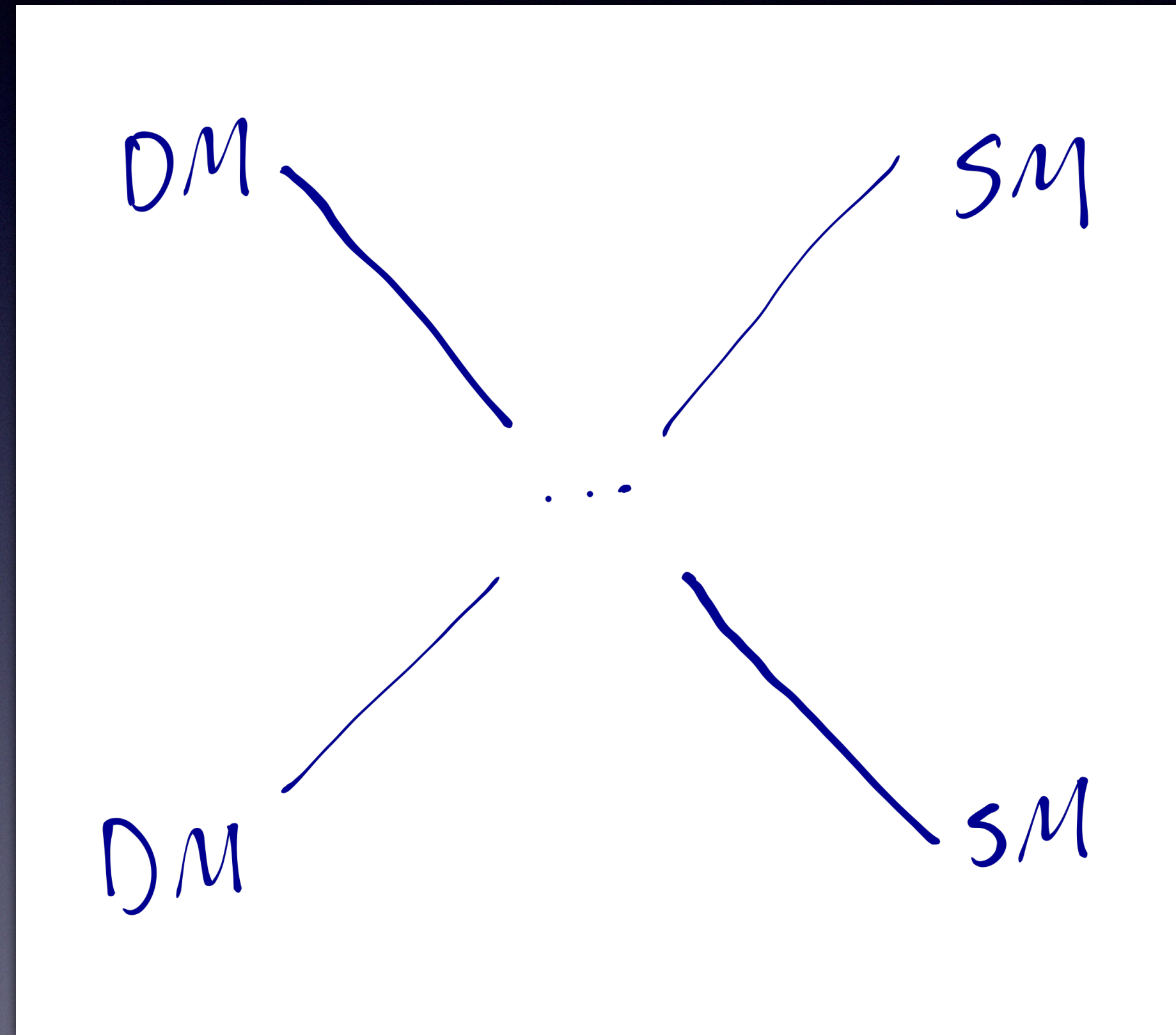
Huge range of possibilities from keV to GeV scale

# WIMP complementarity

cosmic rays (indirect)



nuclear recoils (direct)

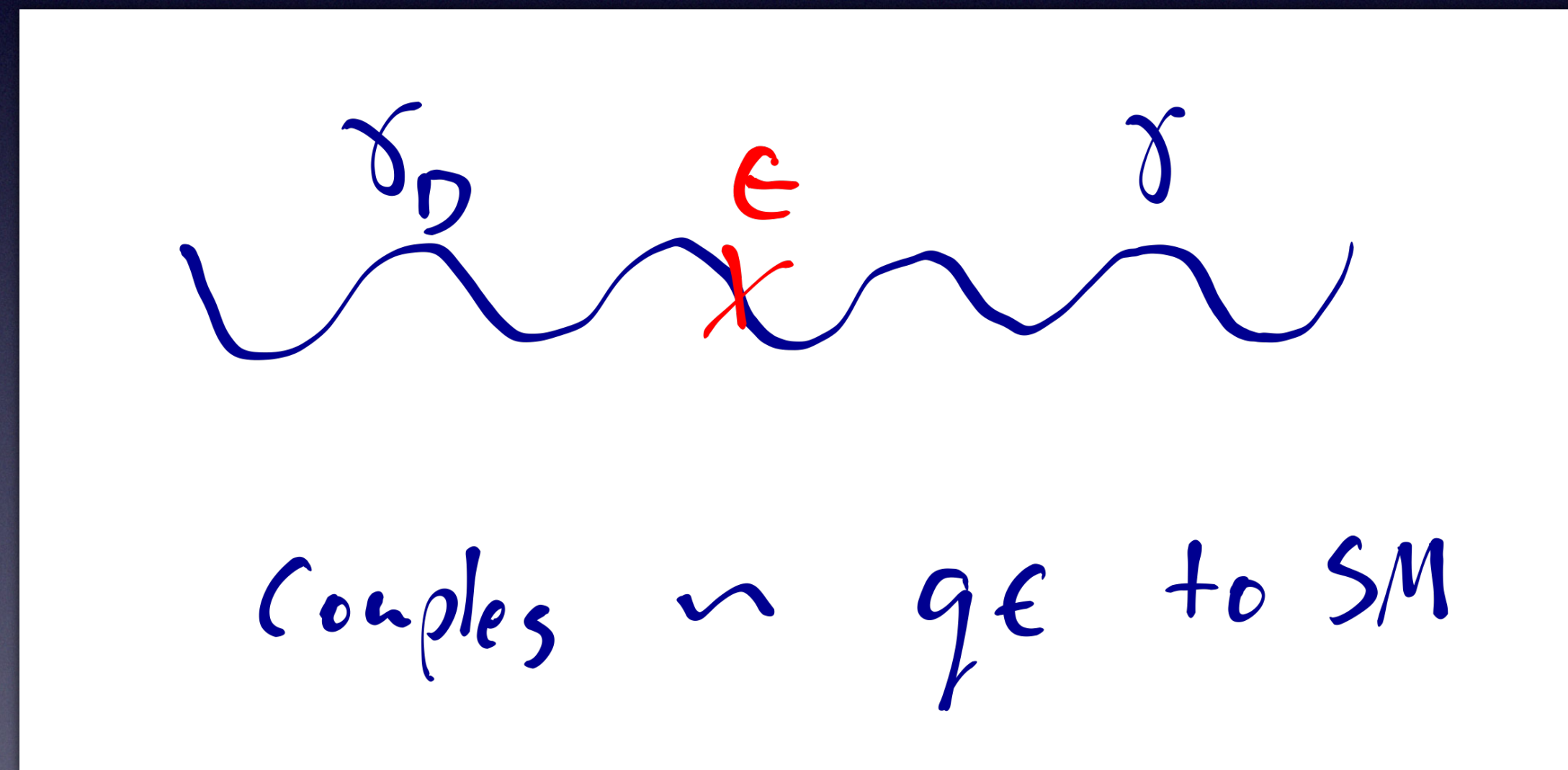


colliders (production)



# Coupling and decoupling a light particle

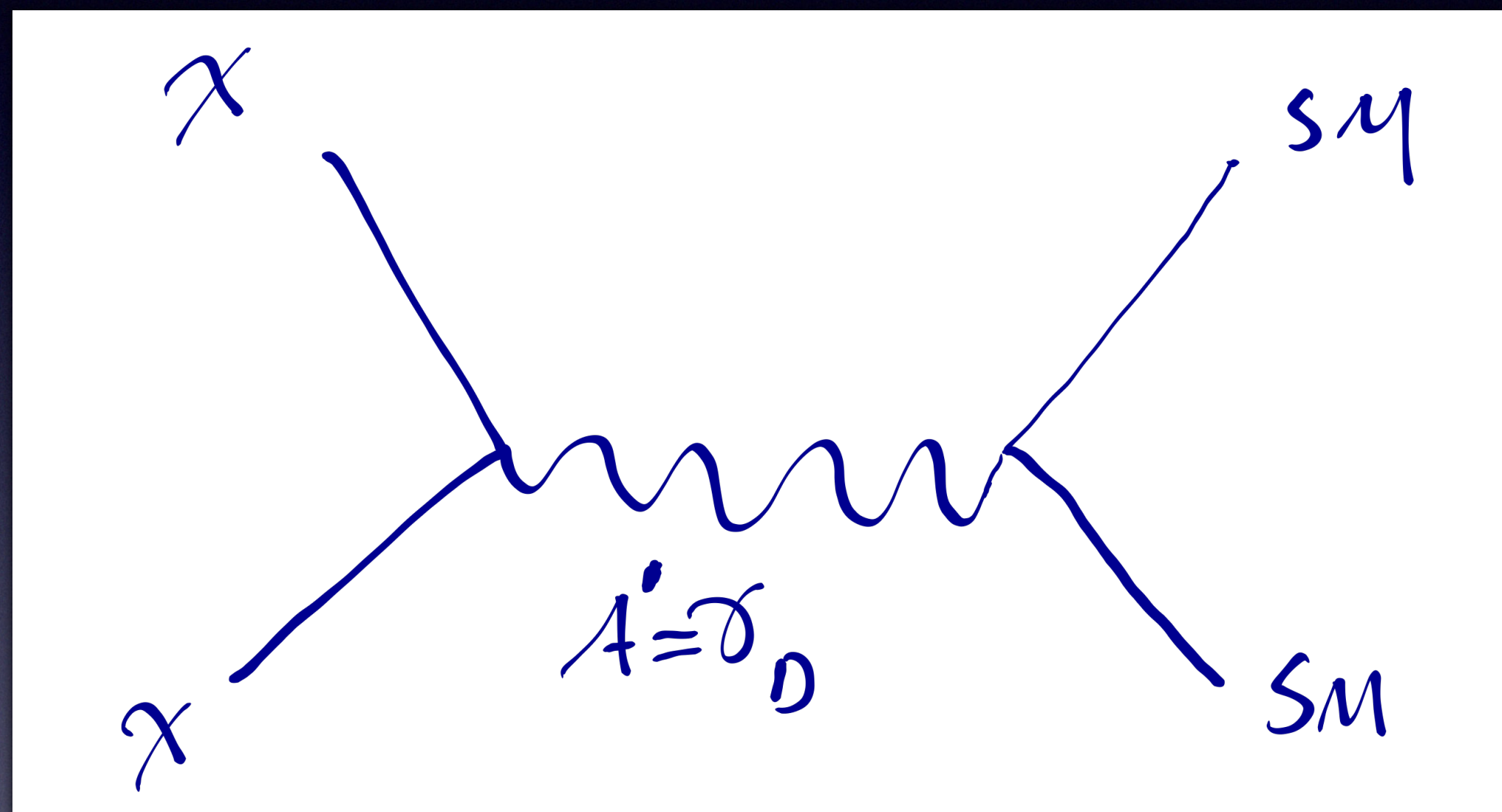
A light DM particle needs a new interaction to stay in equilibrium



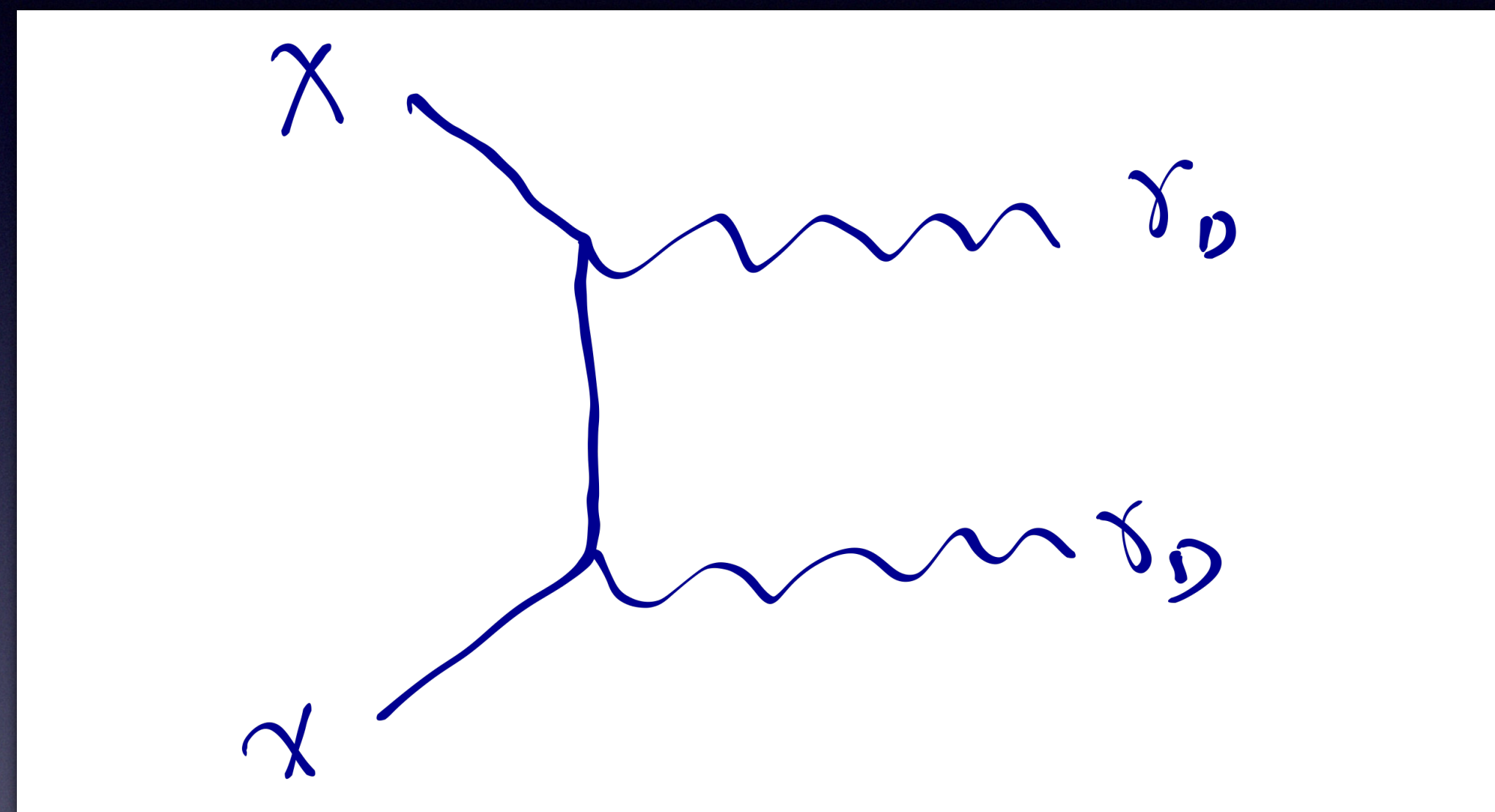
Simple example a “dark photon” - can naturally be very weakly mixed

Holdom; Boehm + Fayet

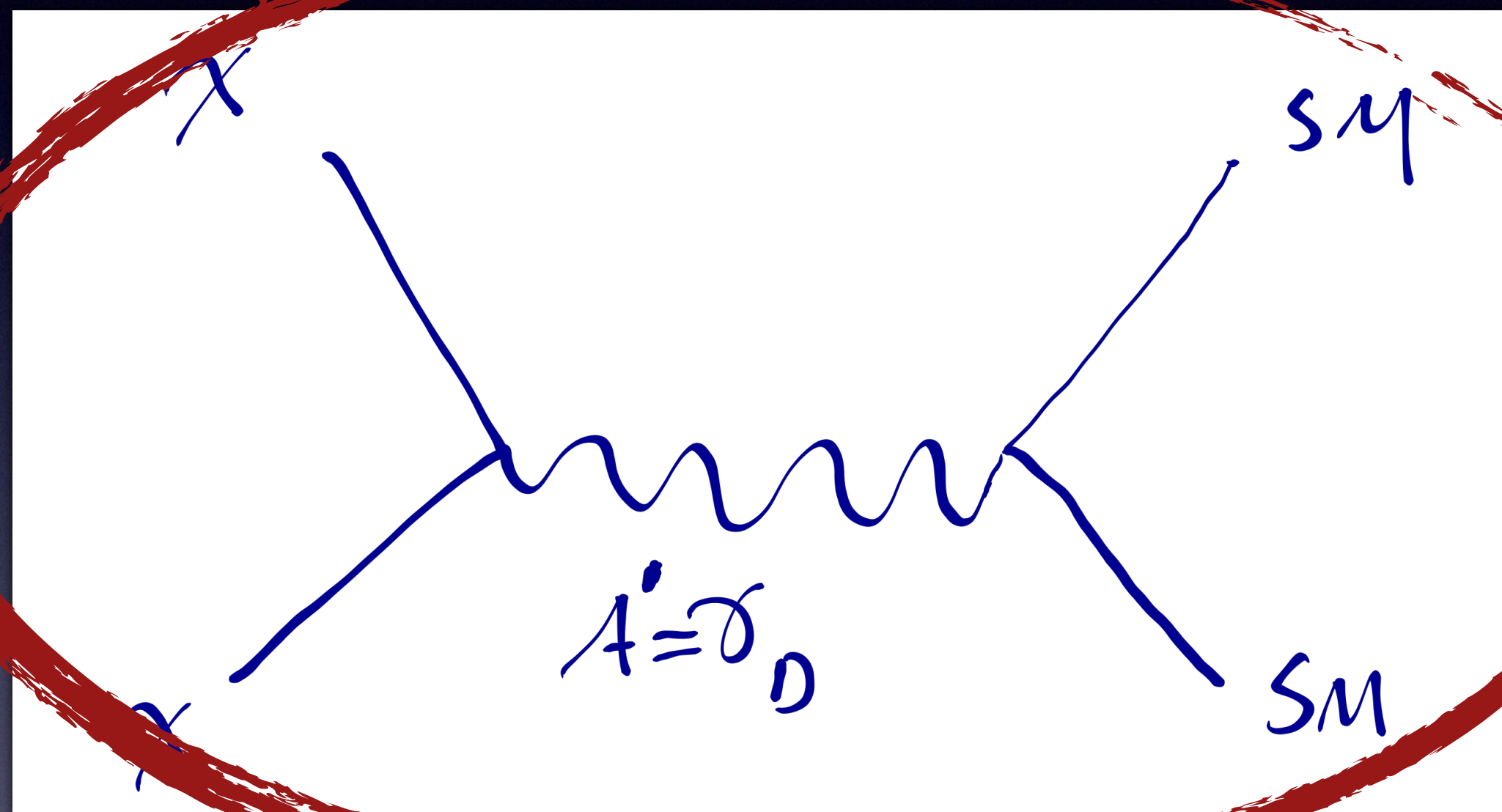




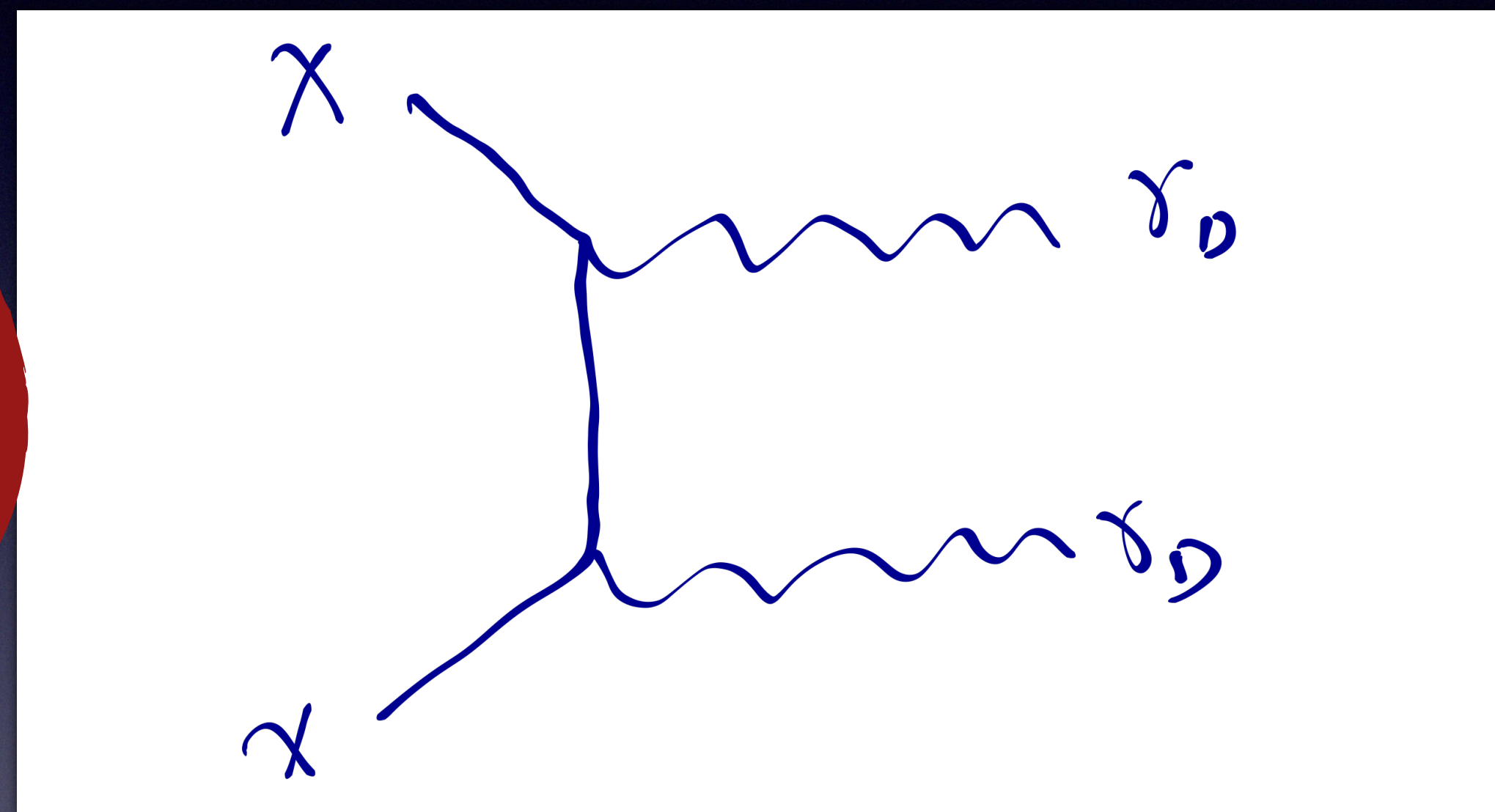
SM Annihilation



Hidden Sector Annihilation

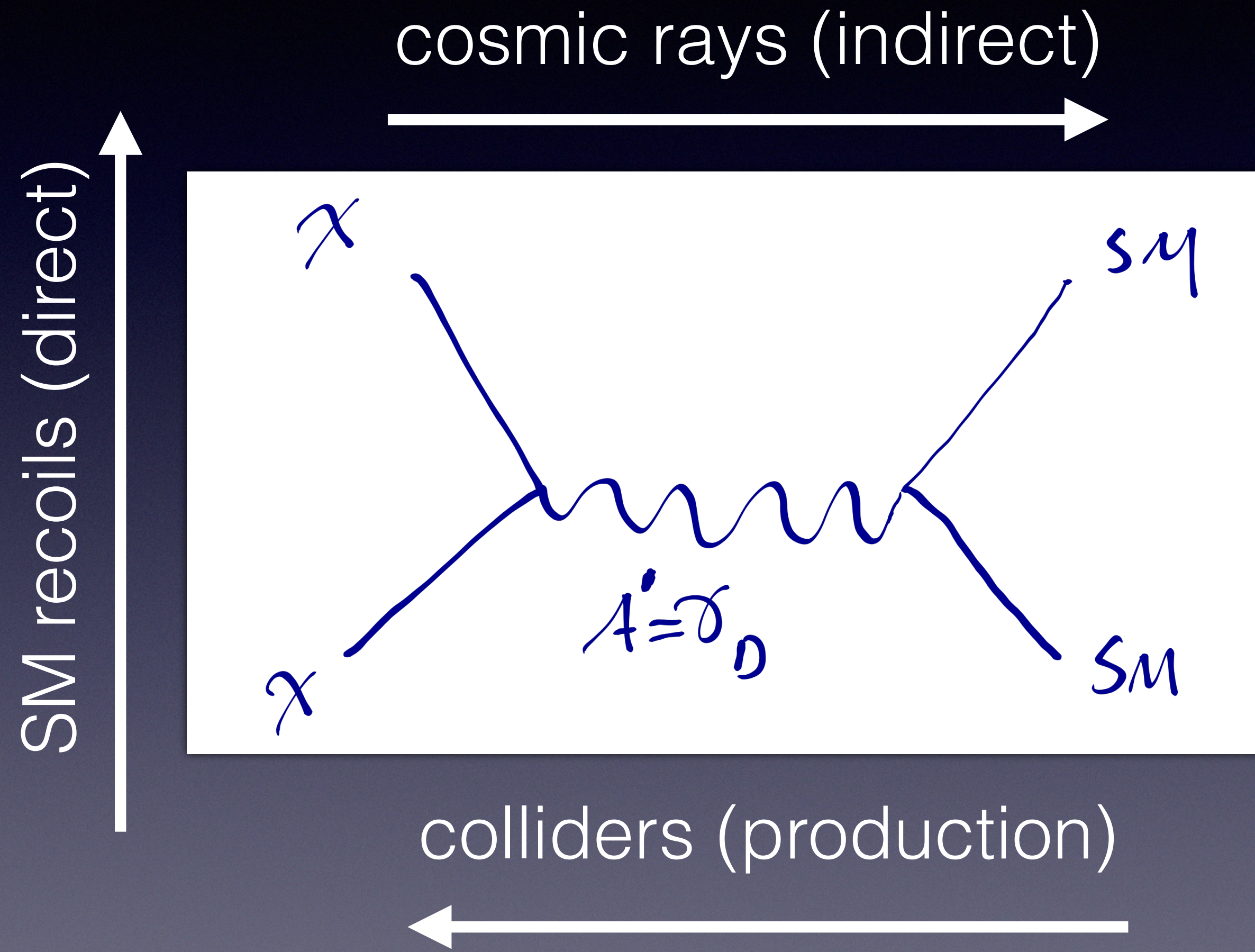


SM Annihilation



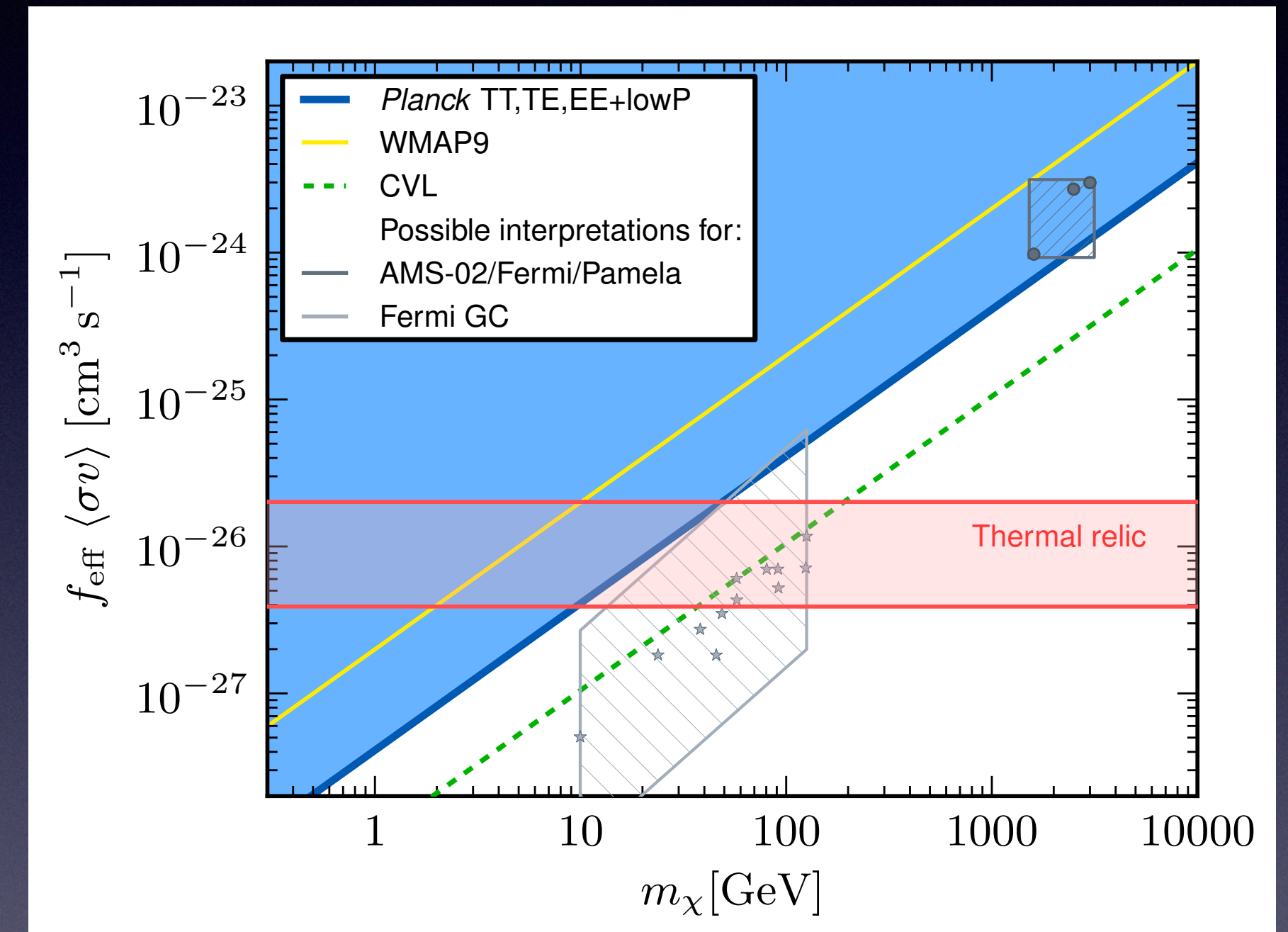
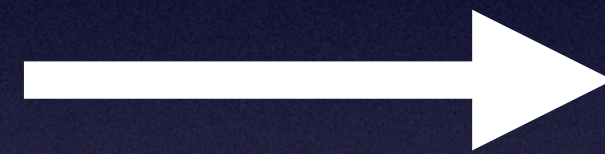
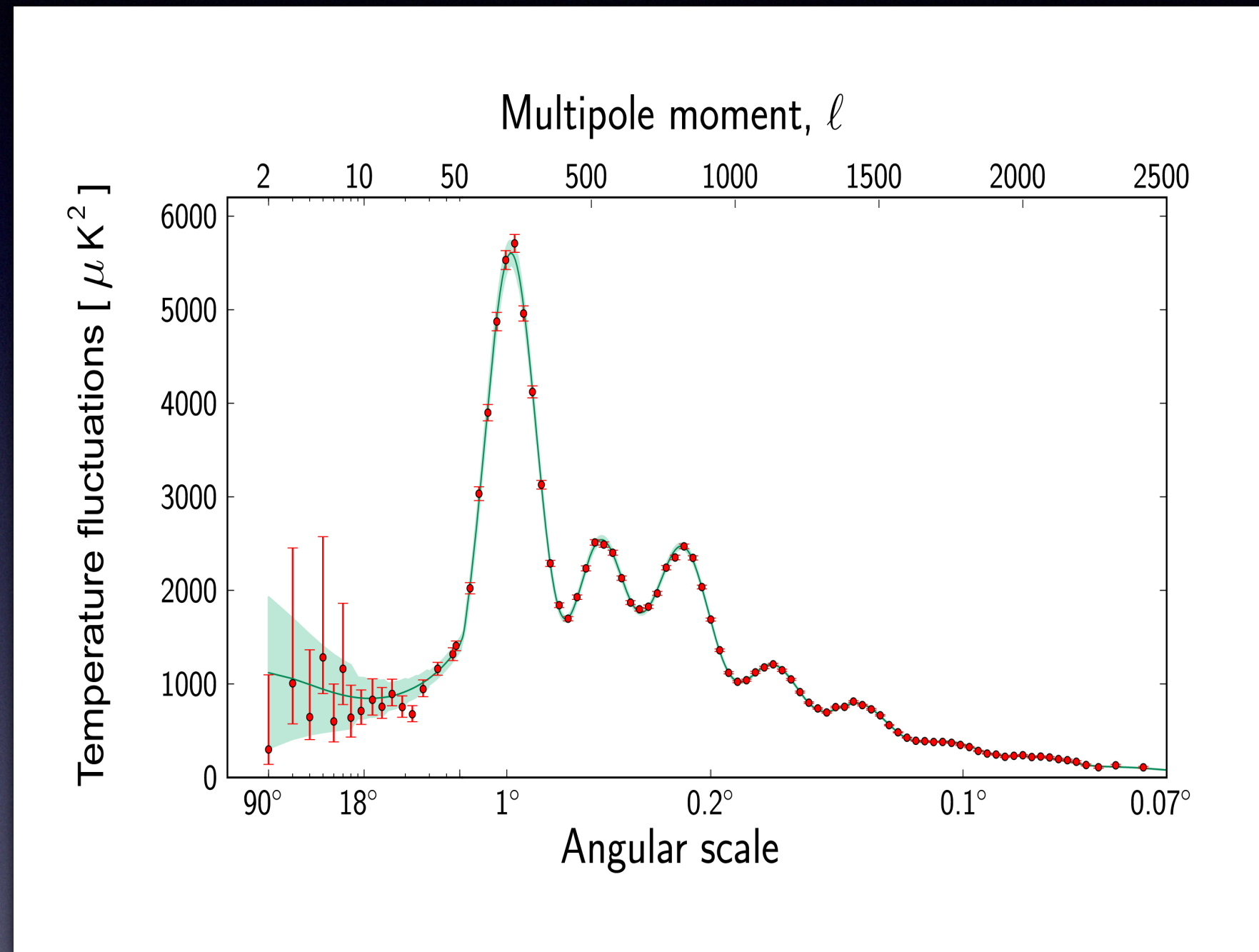
Hidden Sector Annihilation

# Complementarity “Classic”



Limited final states so complementarity is more robust

# Cosmology: already powerful



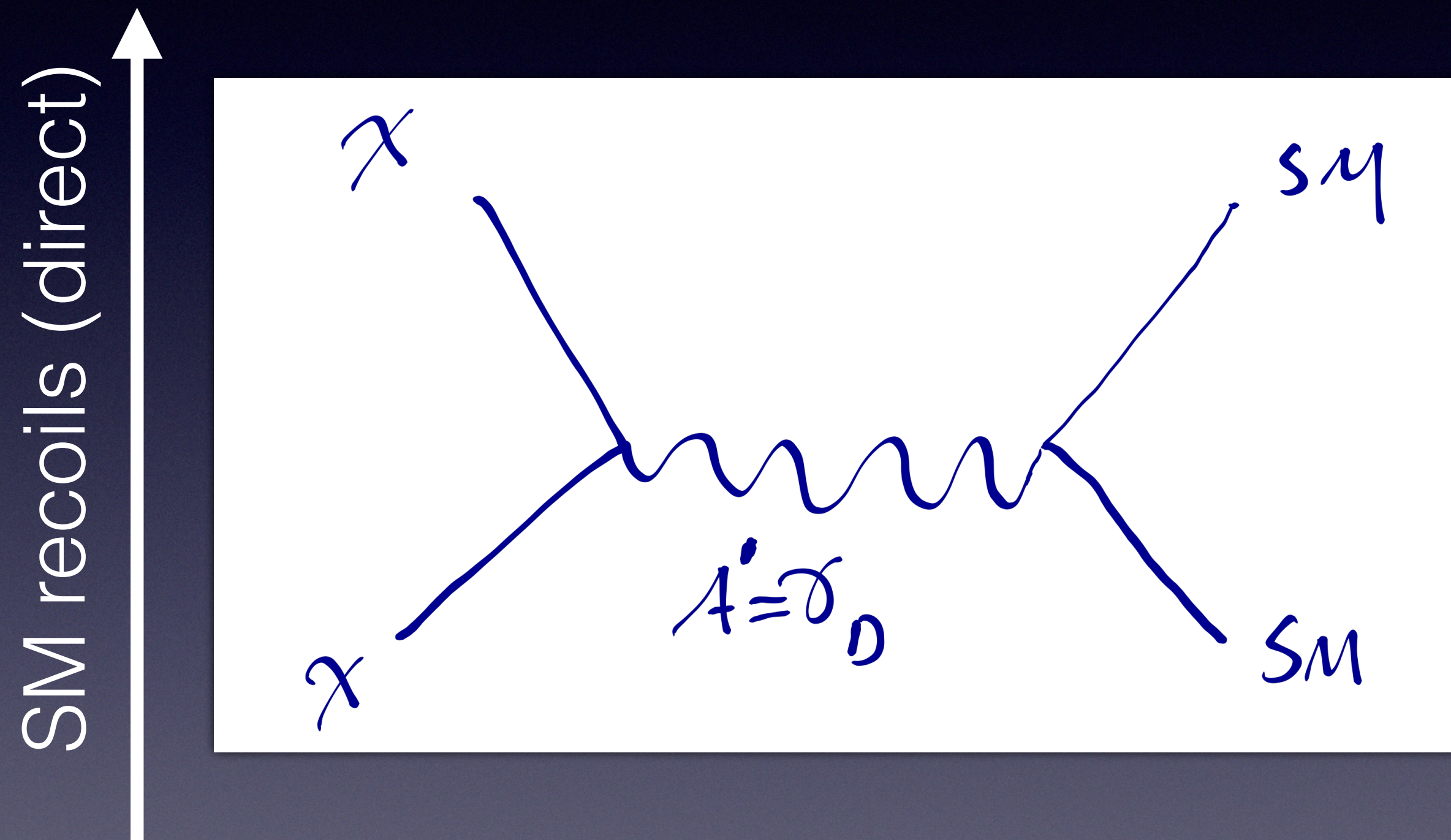
- CMB, LSS much more advanced than in 90's
- CMB constraints light relics more effectively

# A signal from $z=1100$

$$\frac{E}{v_d} \sim n_x^2 \langle \sigma v \rangle m_x$$
$$\sim \frac{(n_x^2 \langle \sigma v \rangle)}{m_x}$$

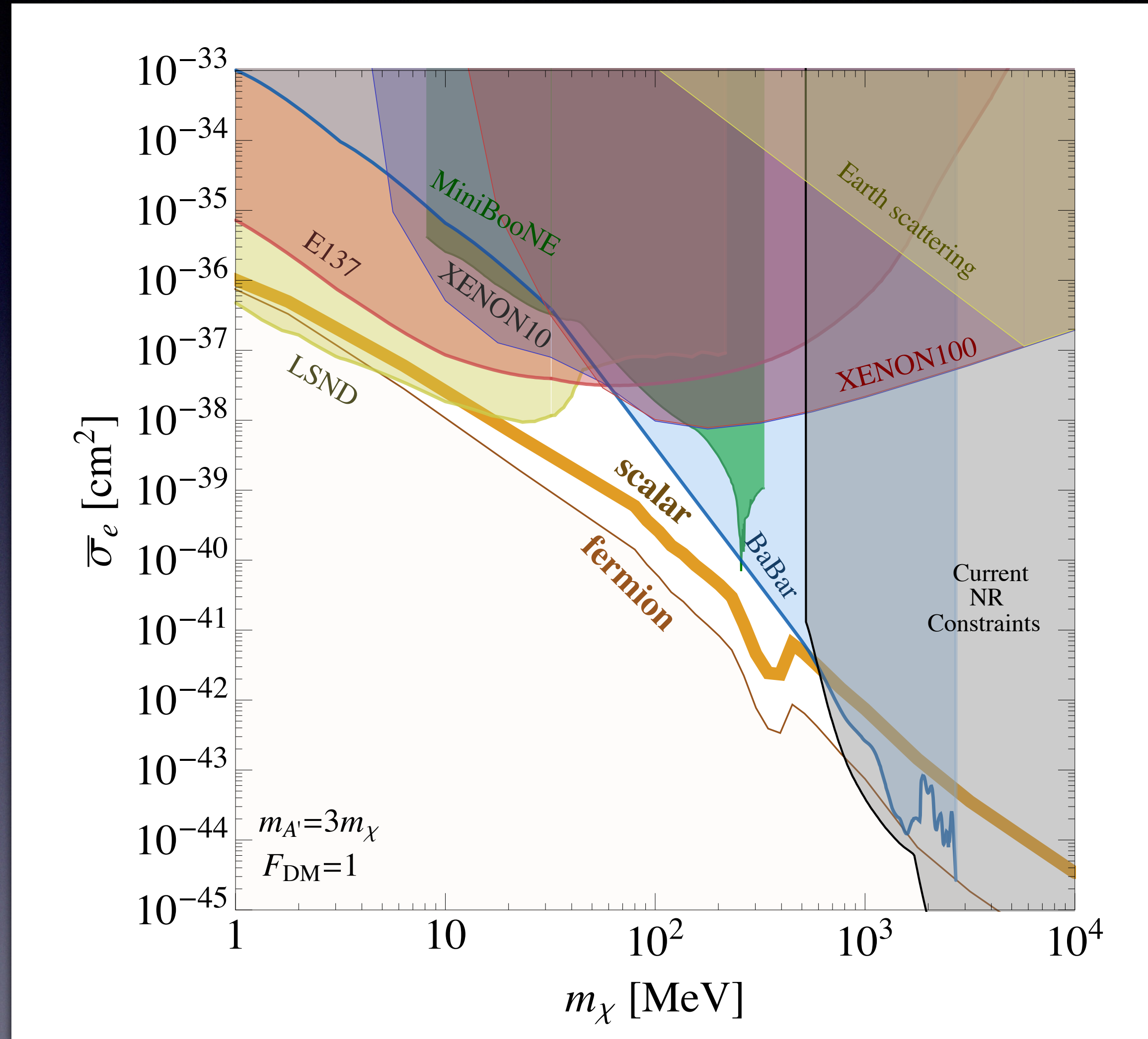
- Need to turn off annihilation at recombination
  - Annihilation is p-wave (velocity suppressed) [scalar]
  - Mass splitting between Majorana states [pseudo-Dirac fermion]

# Direct detection



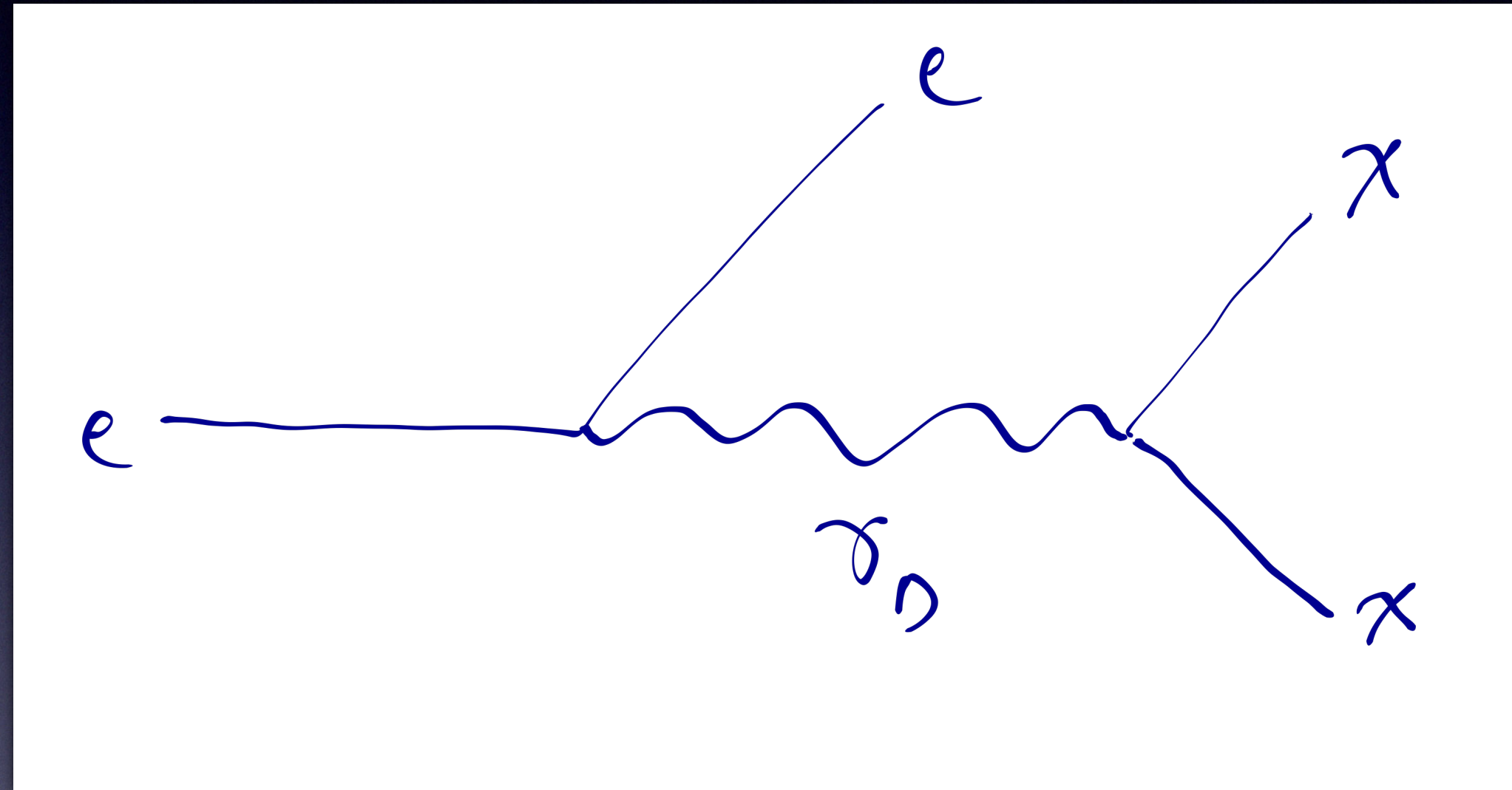
Not just nuclear anymore

# The Thermal Target



Plot from Essig

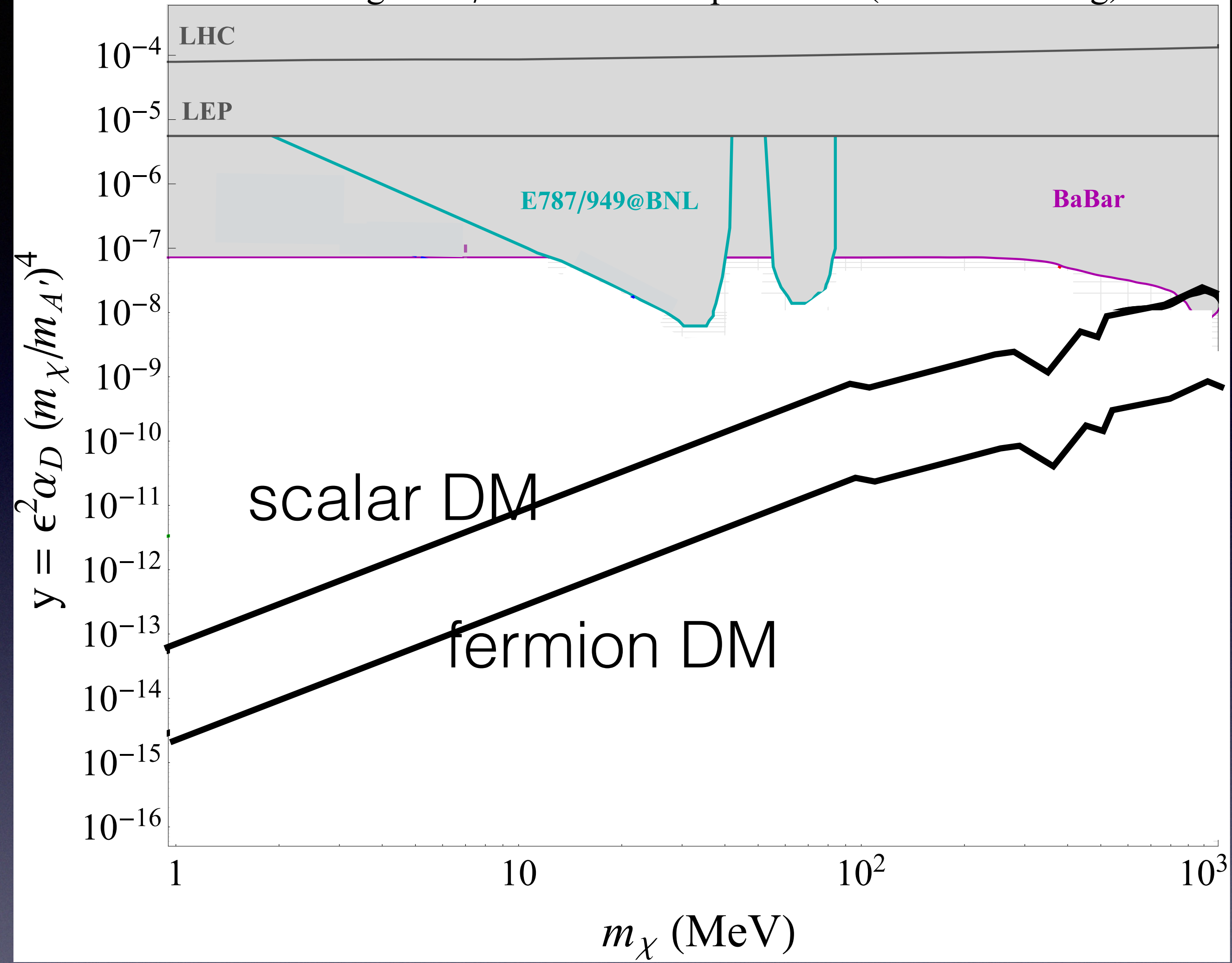
# “New” Complementarity



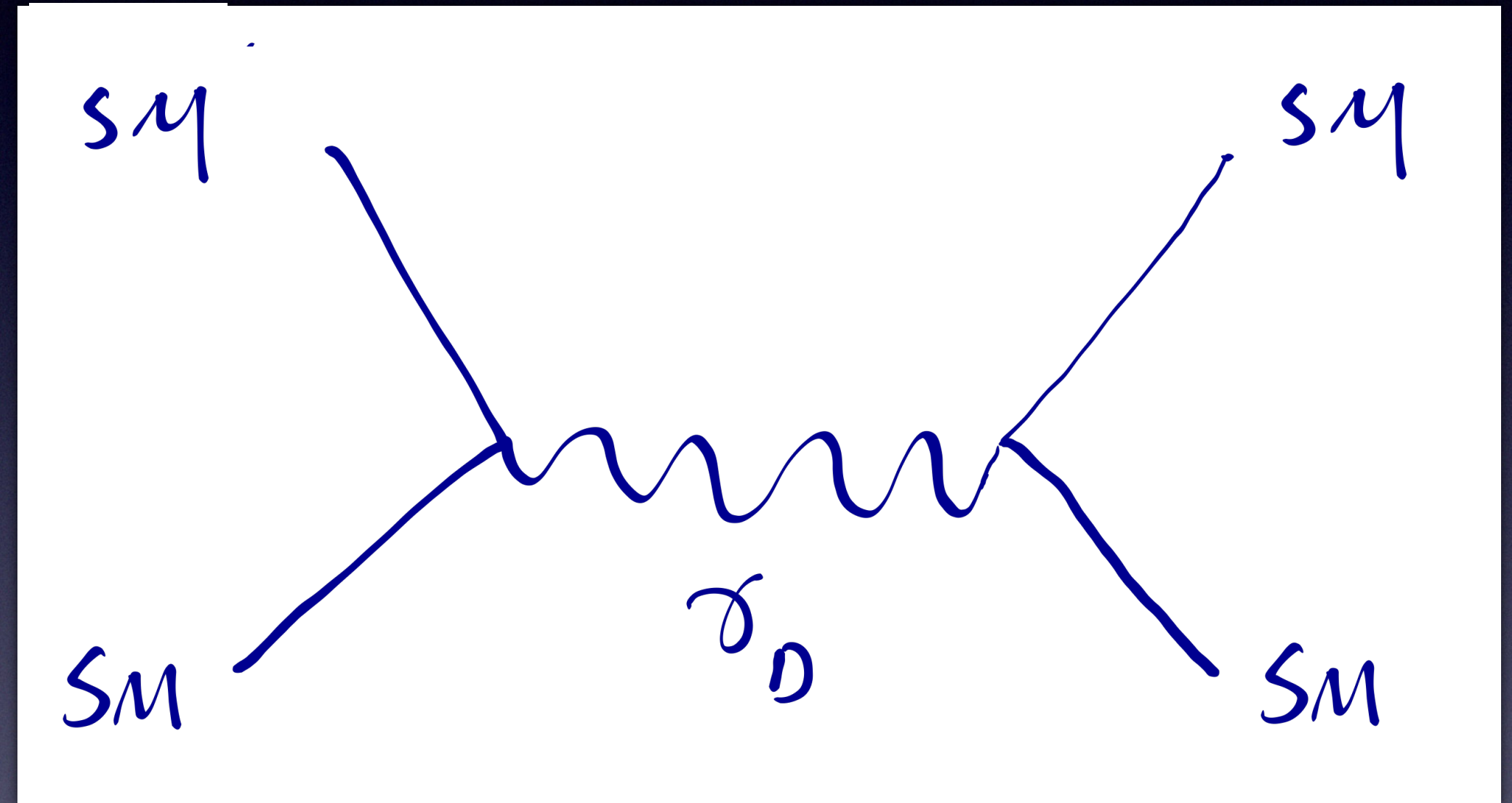
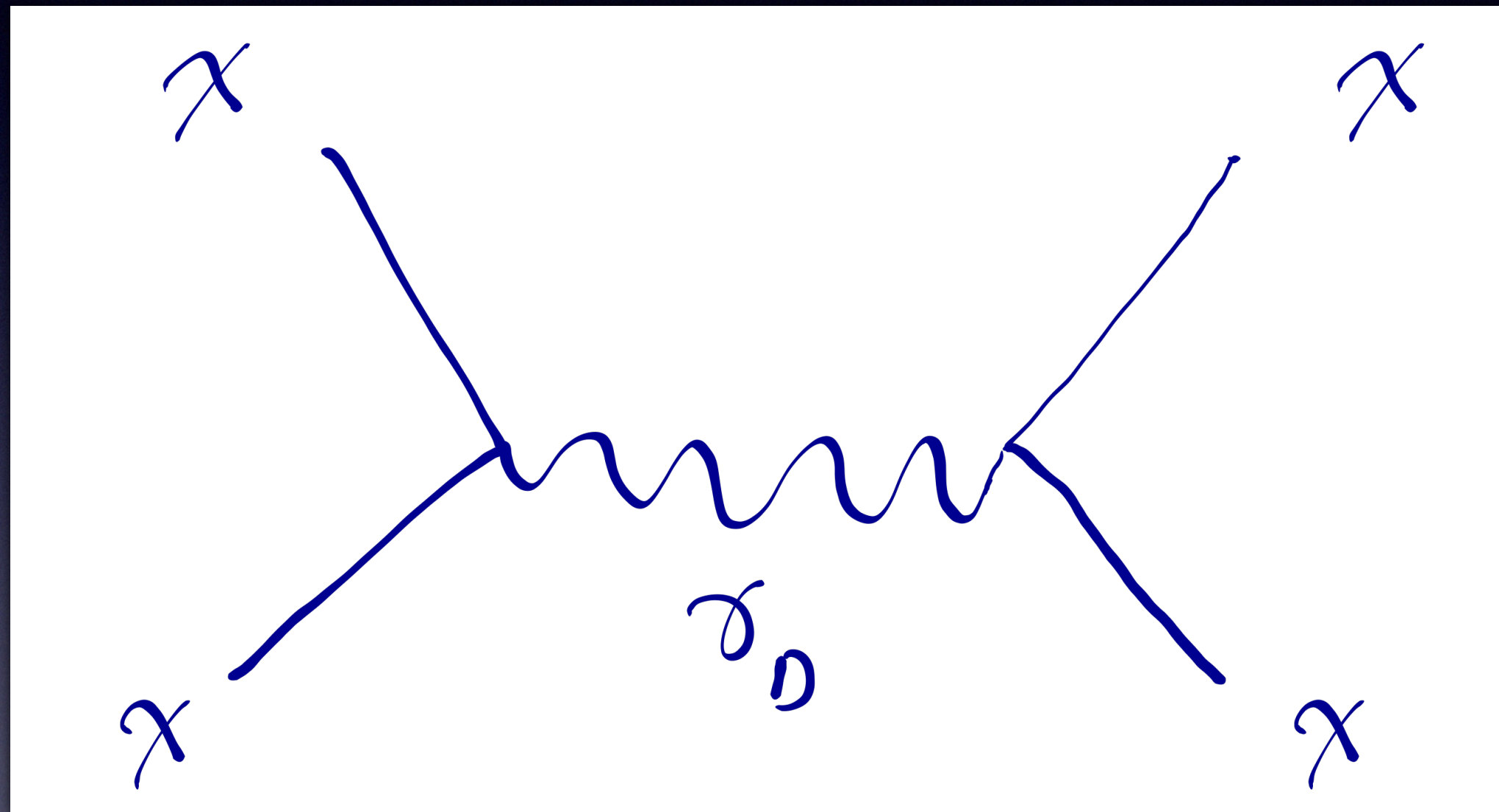
Parametrically linked tightly to thermal diagram



# Missing Mass/Momentum Experiments (Kinetic Mixing)



# “New” Complementarity



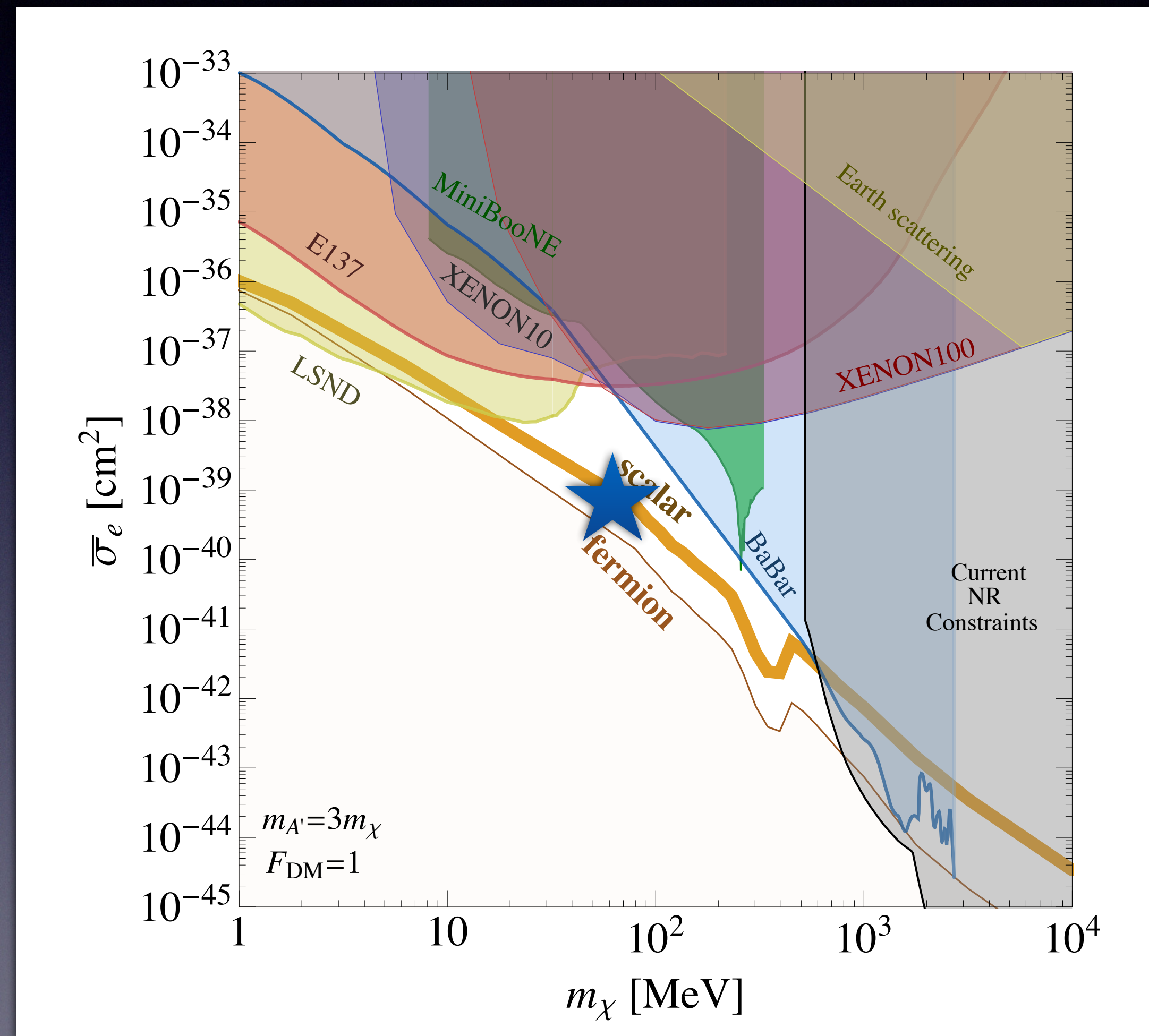
Parametrically linked more weakly to thermal diagram

DM self interaction? Anomalies like  $^8\text{Be}$ ?

# Asking general questions

- This needn't be *the* DM to be interesting
- Are there any particles leftover from “late” (sub TeV) thermal contact?

# Scaling of signals



# Scaling arguments

$$\rho \sim \frac{1}{\sigma} \text{ i.e. } \rho \sim \rho_0 \frac{\sigma_0}{\sigma}$$

direct signal stable

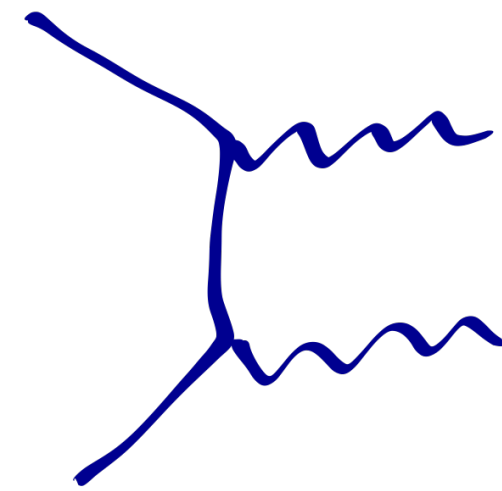
$$\text{but } \eta \sigma v \sim \frac{\rho}{m} \sigma v \sim \frac{\rho_0 \sigma_0}{m \sigma} \sigma v \sim \text{const}$$

$$\eta^2 \sigma v \sim \frac{\rho_0^2 \sigma_0^2}{m^2 \sigma^2} \sigma \sim \frac{1}{\sigma} \text{ decreases for larger } \sigma$$

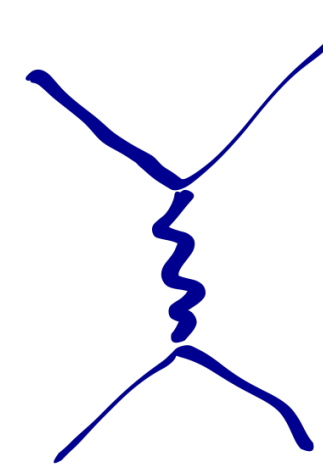
indirect signal decreases

The thermal target here is \*more\* robust

# Scaling arguments


$$\sim g_0^4$$

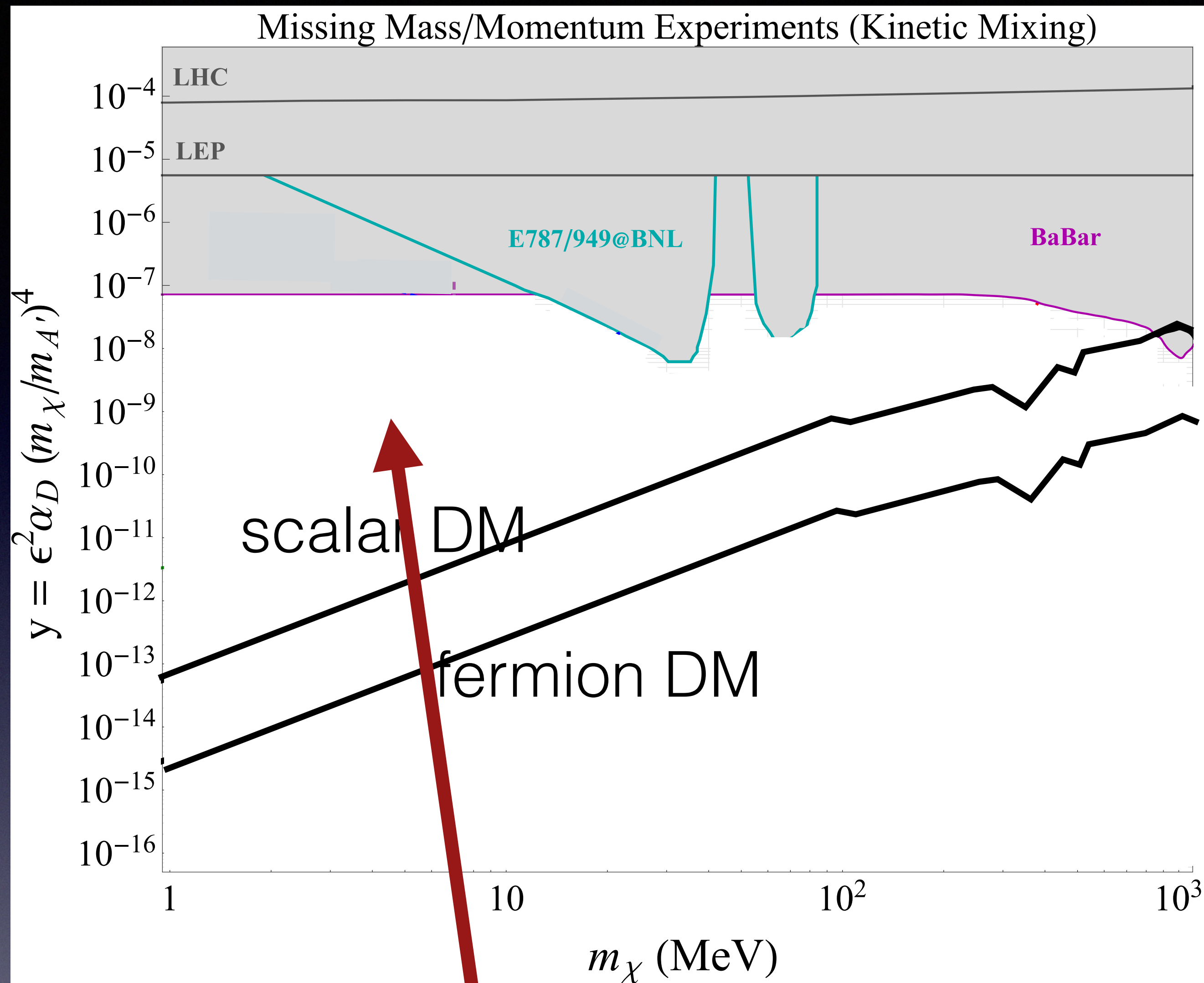
direct signal decreases


$$\sim g_0^2 \Rightarrow \rho\sigma \sim \frac{1}{g^2}$$

but  $\frac{\rho^2 \sigma v}{m^2} \sim \frac{1}{g^4}$

can turn off CMB & still scatter

indirect signal decreases faster



qualitative physics information here, too

# Moving away from priors changes questions

- Looking for the DM is critical, but finding cosmologically “relevant” hidden physics would be profound
- A WIMP was motivated to be \*the\* DM, but hidden physics may or may not be
- Broadening the question changes how we think about targets, how robust they are and *changes the implications of complementarity*



# Generalizations of thermal history

- Freeze-in (signal arises later)
- SIMPS (new “miracle” at 100 MeV)
- ELDERS
- Forbidden DM
- Dynamical DM
- ...
- Major theoretical advancement has been in rethinking “generic” thermal particles

*Still need labs for new ideas*

# Rebalancing our priorhedron

