



HAPPY 50TH ANNIVERSARY..!

# SEARCH FOR STERILE NEUTRINOS AT THE NOVA NEAR DETECTOR

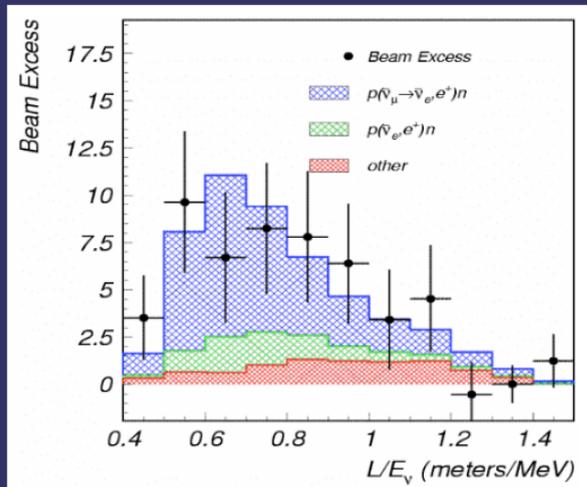
**SIVA PRASAD K**  
UNIVERSITY OF HYDERABAD

**FOR THE NOVA COLLABORATION**  
NEW PERSPECTIVES | FERMILAB  
JUNE 05, 2017

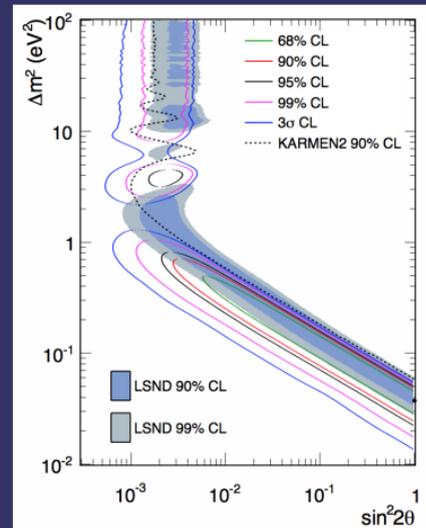
# LSND/MINIBoONE EXCESS

- The excess can be interpreted as "Sterile Neutrinos" with  $\Delta m^2 \geq 1 \text{ eV}^2$ .
- In the 2-flavor neutrino oscillation model, the excess can be expressed as

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \frac{1.27 \Delta m^2 L}{E}$$



A. Aguilar et al., Phys. Rev. D 64, 112007 (2001).



# (3+1) MODEL

- For our analysis, we consider (3+1) model. The PMNS matrix is given by

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{pmatrix}$$

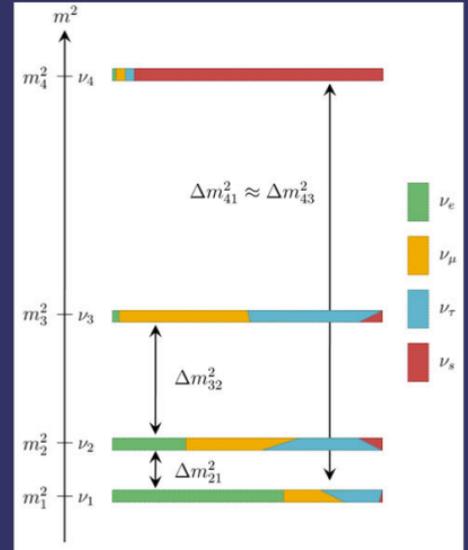
- $\nu_e$  Appearance Probability

$$P_{\nu_\mu \rightarrow \nu_e}^{(-) \text{SBL}, 3+1} = 4|U_{\mu 4}|^2|U_{e4}|^2 \sin^2 \frac{\Delta m_{41}^2 L}{4E} = \sin^2 2\theta_{\mu e} \sin^2 \frac{\Delta m_{41}^2 L}{4E}$$

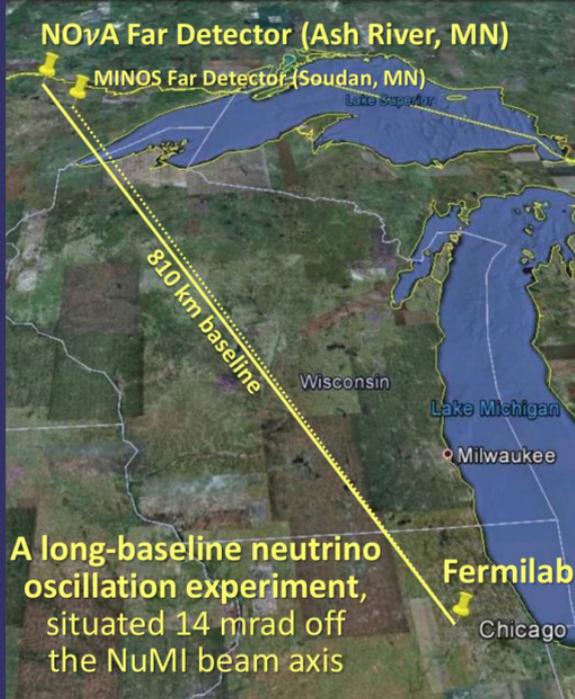
- $\nu_\mu$  Disappearance Probability

$$P_{\nu_\mu \rightarrow \nu_\mu}^{(-) \text{SBL}, 3+1} = 1 - 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2) \sin^2 \frac{\Delta m_{41}^2 L}{4E} = 1 - \sin^2 2\theta_{\mu\mu} \sin^2 \frac{\Delta m_{41}^2 L}{4E}$$

$$\sin^2 2\theta_{\mu e} = \sin^2 2\theta_{14} \sin^2 \theta_{24} \quad \text{and} \quad \sin^2 2\theta_{\mu\mu} = \cos^2 \theta_{14} \sin^2 \theta_{24}$$



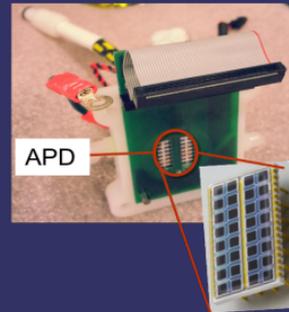
# NOvA EXPERIMENT



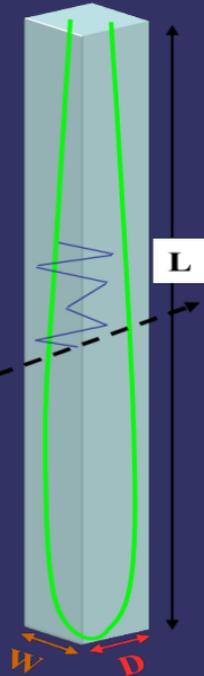
- NOvA (NuMI Off-Axis Electron Neutrino Appearance) is an accelerator based neutrino oscillation experiment.
- Near Detector
  - 1 km from source, 100 meter deep underground
  - 0.3 kton in mass.
- Far Detector
  - 810 km from Fermilab on surface.
  - 14 kton in mass.
- This analysis is done with the near detector alone.

# NOVA DETECTORS

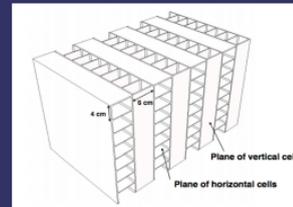
- NOvA detectors are made of plastic PVC cells which are glued together in vertical and horizontal planes.
- Each cell is filled with mineral oil based scintillator with the liquid scintillant.
- Each cell has a Wavelength Shifting Fiber inside to carry light photons to the Avalanche Photodiode read out.



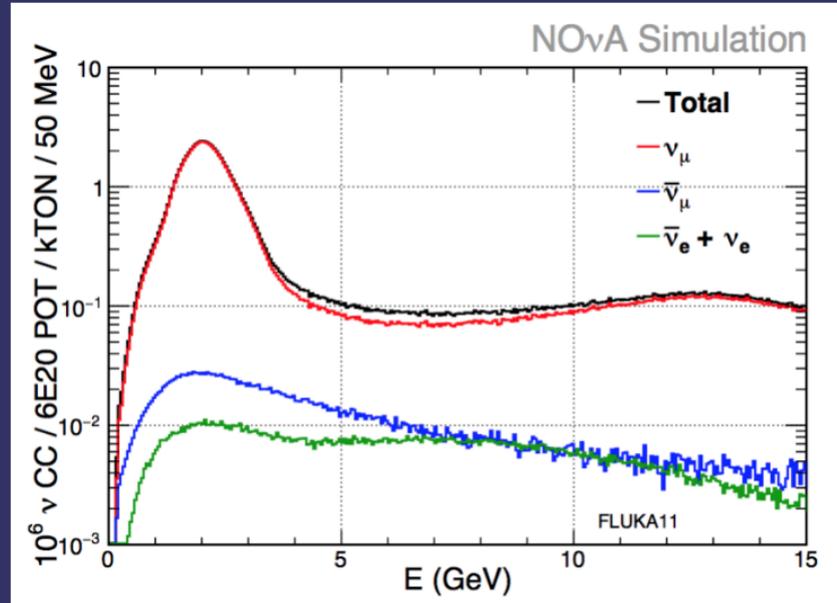
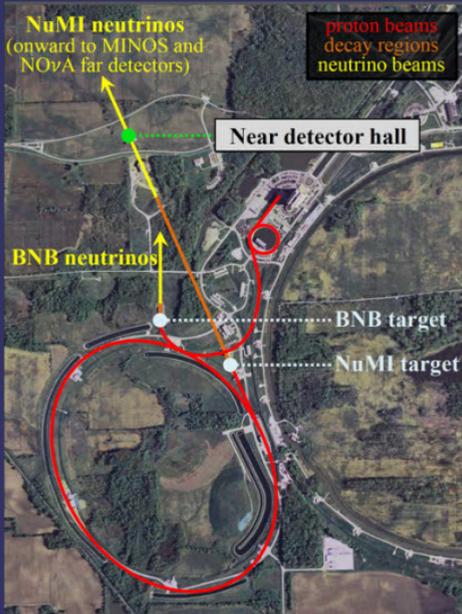
To 1 APD pixel



typical  
charged  
particle  
path

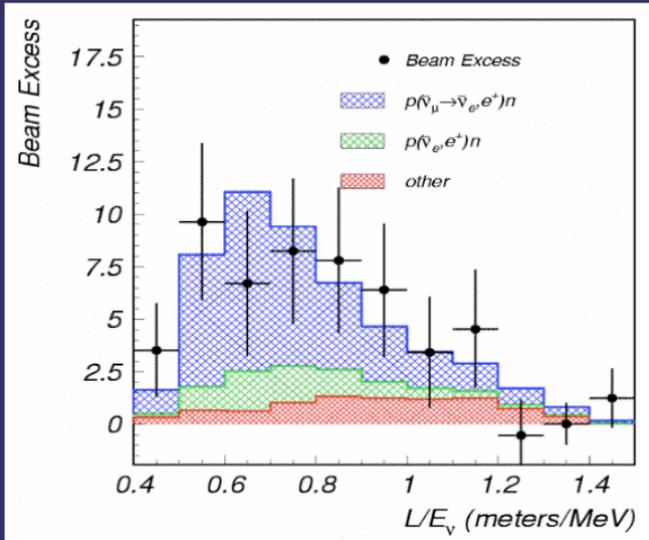


# NUMI BEAM IN NOvA

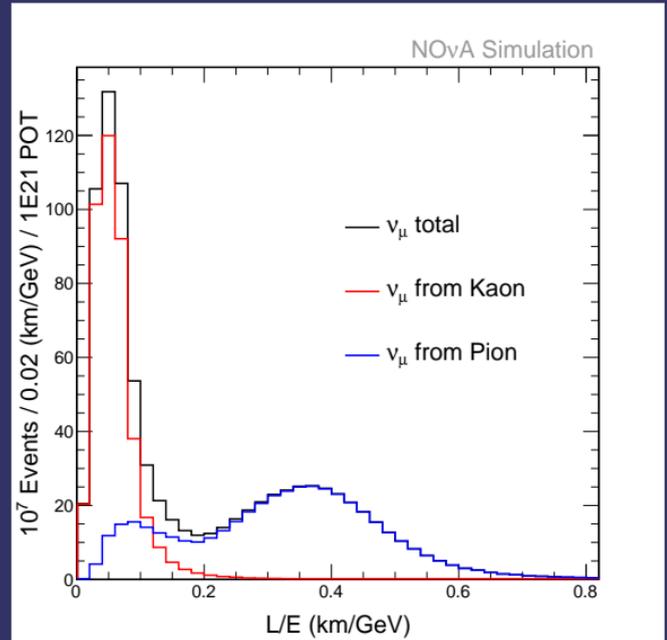


- Both detectors use the same NuMI beam which peaks at 2 GeV, at 14.6 mrad off-axis.
- NuMI beam composition:  $\approx 98\% \nu_\mu$ 's and  $\approx 2\% \nu_e$ 's.
- NOvA is operating at it's designed NuMI power, 700 kW.

# WHAT NOVA CAN DO?

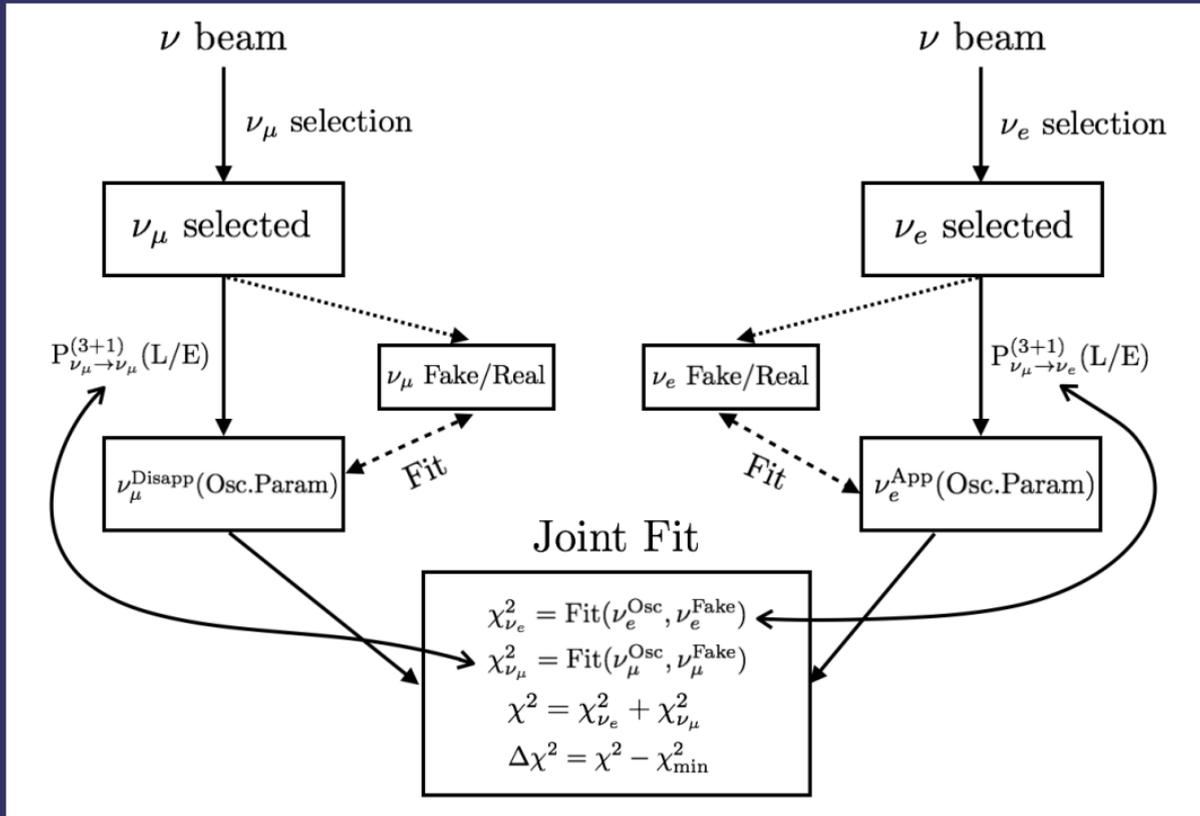


LSND excess with  $L/E$



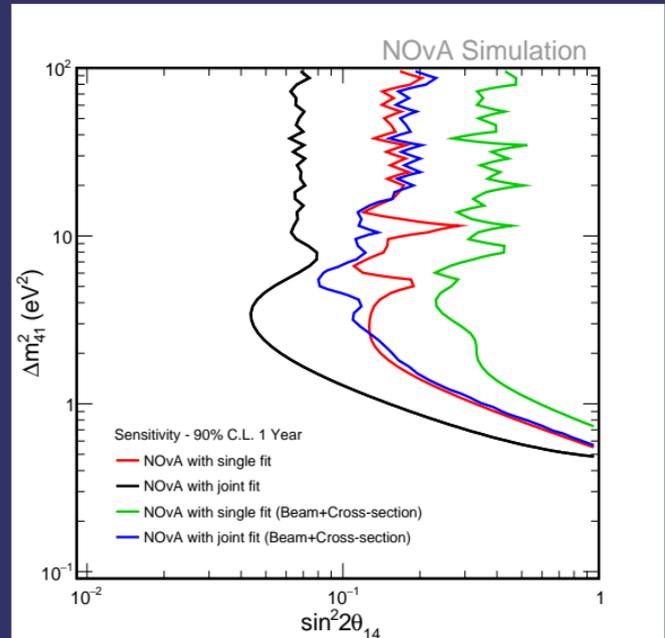
NOvA Near Detector is well placed to probe the Short-Baseline Neutrino Oscillations with  $\Delta m^2 \geq 1\text{eV}^2$ .

# ANALYSIS STRATEGY



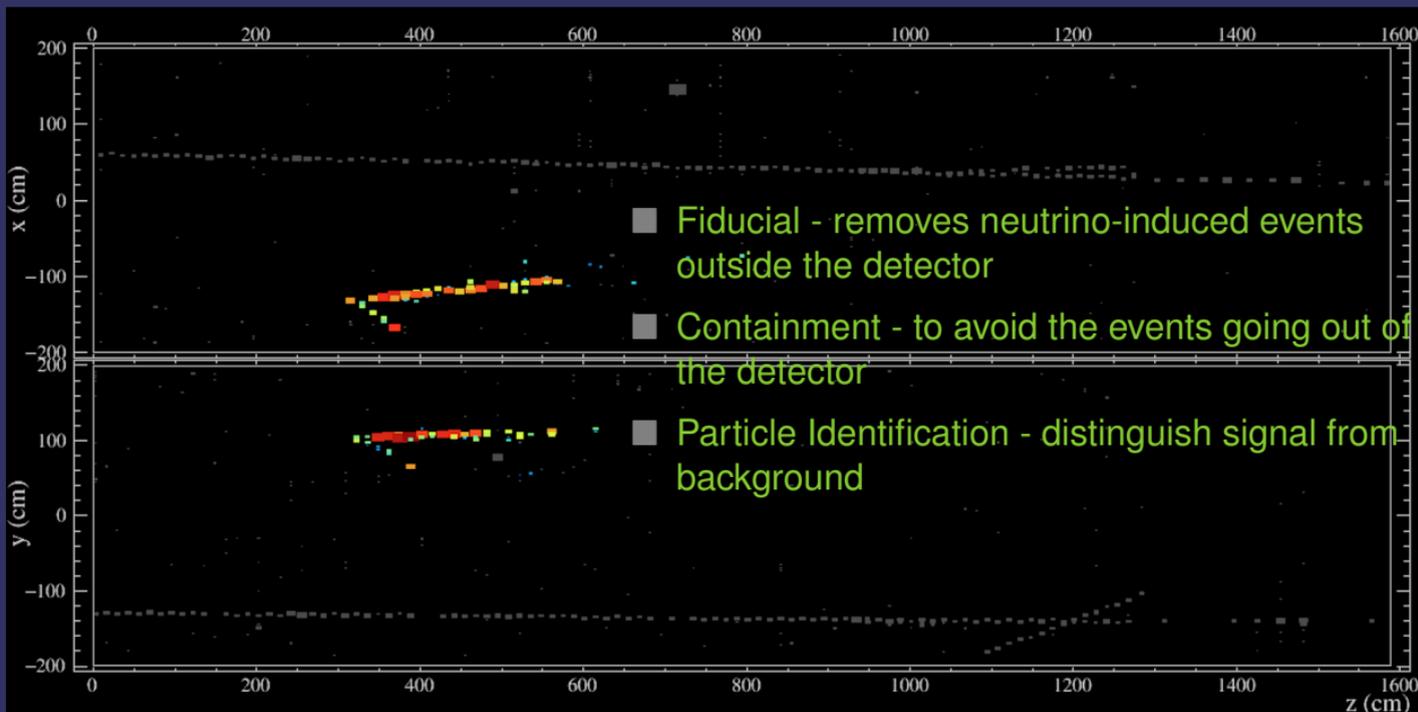
# JOINT FIT METHOD SENSITIVITIES

- Comparison of joint fit sensitivities with the independent  $\nu_\mu$  and  $\nu_e$  fits.
- Joint Fit allows for the partial cancellation of systematic uncertainties, significantly improving sensitivity of the analysis.



$$\sin^2 2\theta_{\mu e} = \sin^2 2\theta_{14} \sin^2 \theta_{24}.$$

# EVENT SELECTION



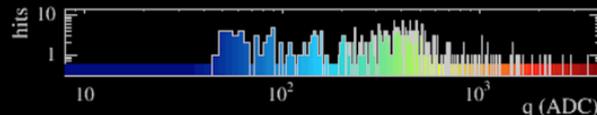
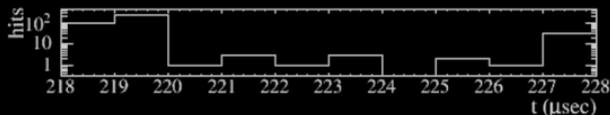
NOvA - FNAL E929

Run: 10586 / 10

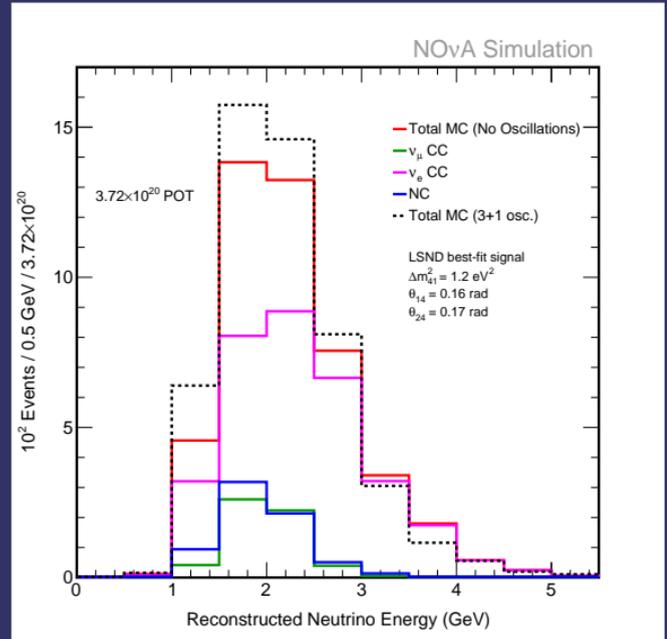
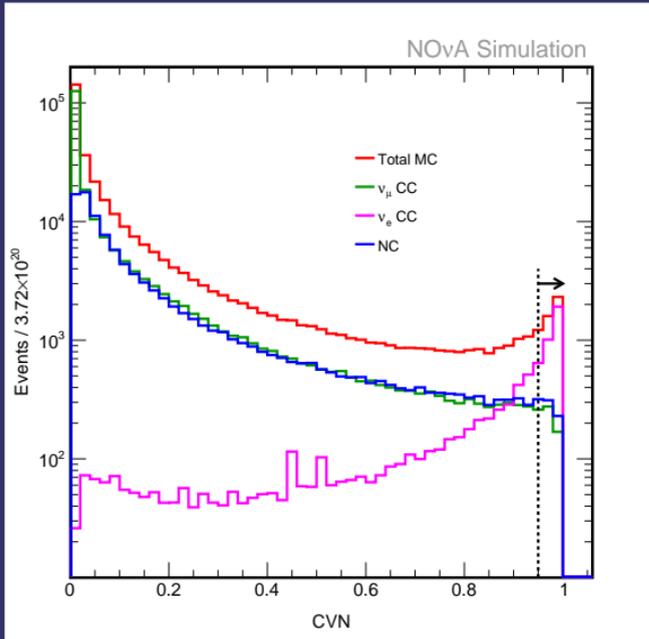
Event: 676464 / --

UTC Wed Nov 26, 2014

10:03:6.472595040

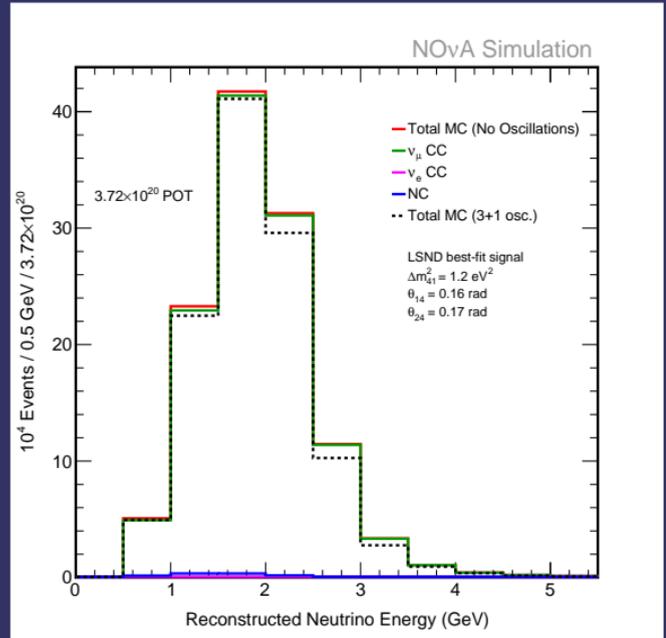
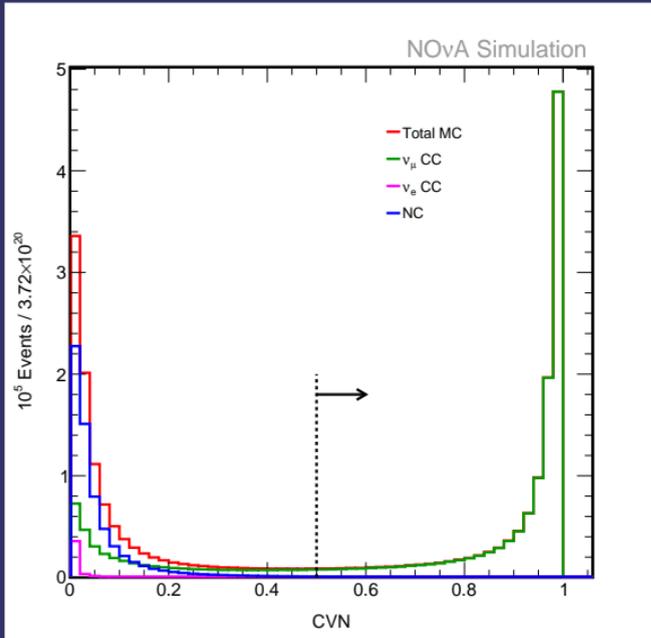


# EVENT SELECTION - $\nu_e$



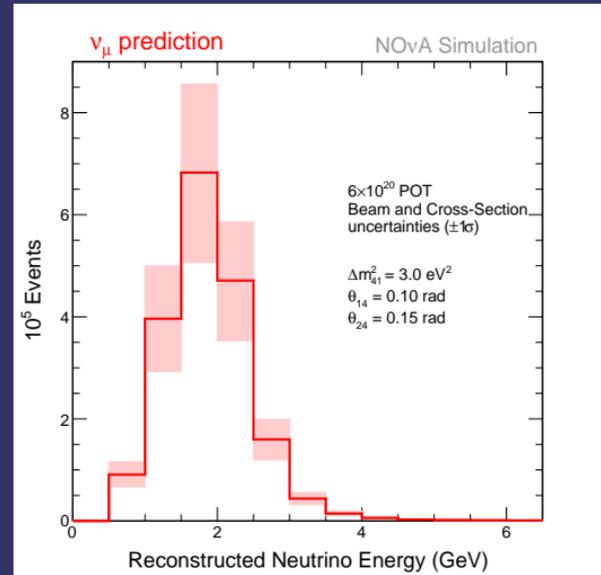
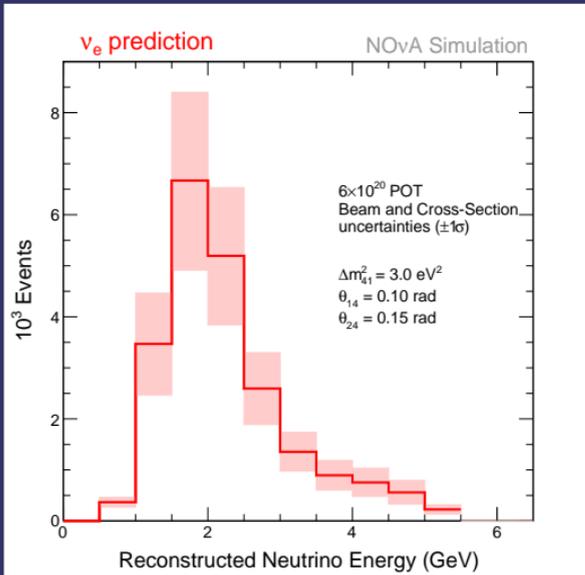
Selection efficiency of  $\nu_e$ 's with CVN > 0.95 is 49% and purity is 86%.

# EVENT SELECTION - $\nu_\mu$



Selection efficiency of  $\nu_\mu$ 's with CVN > 0.5 is 58% and purity is 96%.

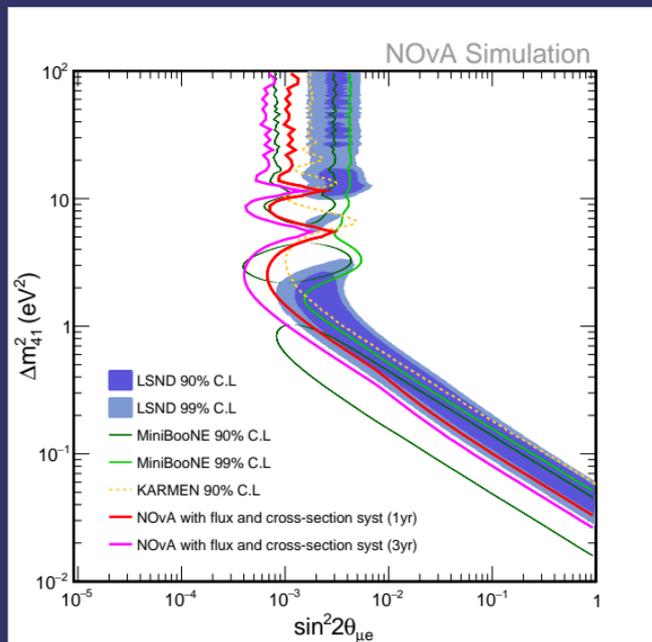
# PREDICTION WITH SYSTEMATICS



- Predictions for both  $\nu_e$  (left) and  $\nu_\mu$  (right) after including the flux and cross-section uncertainties added in quadrature at  $\pm 1\sigma$ .
- Any correlated systematics will be cancelled in the joint fit, e.g cross-section systematics.

# SENSITIVITY - $\theta_{\mu e}$

- Fitting simultaneously both  $\nu_e$  and  $\nu_\mu$  prediction with  $\nu_e$  and  $\nu_\mu$  selected MC spectrum with no (3+1) oscillations.
- Fit for  $\theta_{\mu e}$ , profiling  $\theta_{14}$ ,  $\theta_{24}$  and  $\theta_{34}$ , and keeping all other parameters fixed.
- Included flux and cross-section uncertainty into the fit.



$$\sin^2 2\theta_{\mu e} = \sin^2 2\theta_{14} \sin^2 \theta_{24}.$$

# SUMMARY

- NOvA is well positioned to search for the sterile neutrinos with  $\Delta m^2 \geq 1 \text{ eV}^2$  at its baseline with the NuMI beam.
- Joint fit provides significant improvement in sensitivities over the individual fits.
- Current NOvA sensitivities are competitive with the on-going and future experiments like Short-Baseline Near Detector program at Fermilab etc.
- Stay tuned for the exciting results.

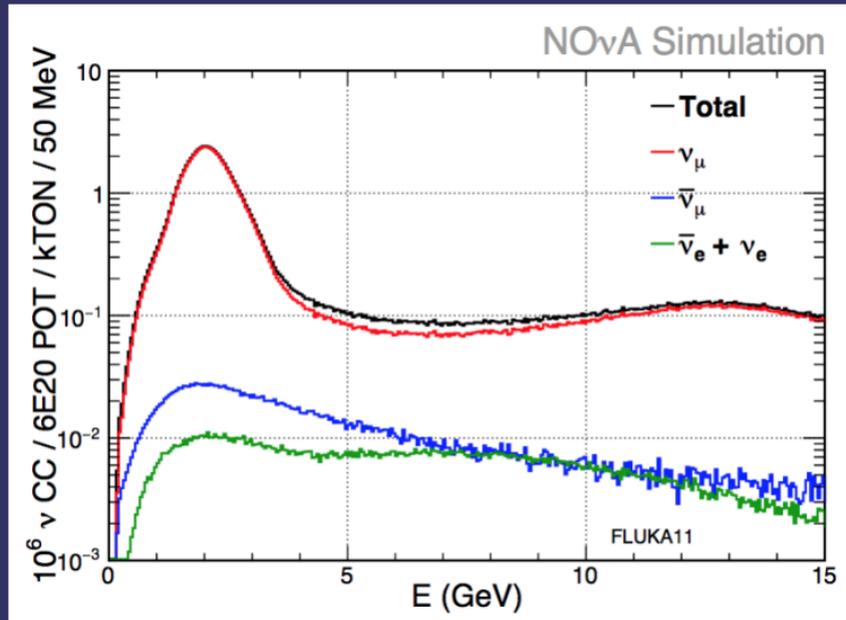
# THANK YOU



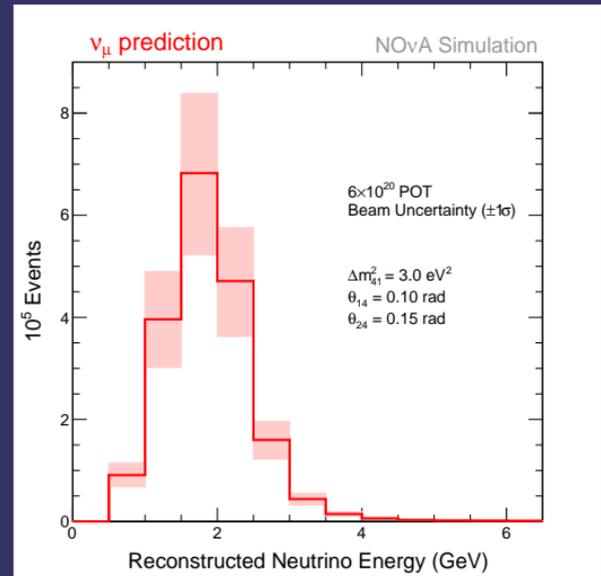
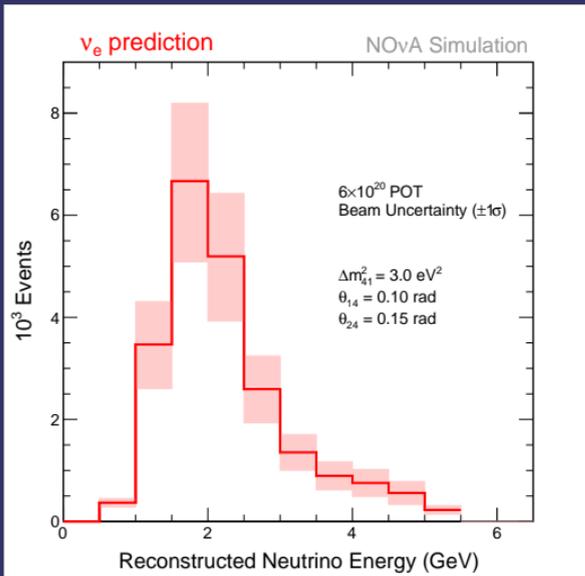
[www-nova.fnal.gov](http://www-nova.fnal.gov)

# BACKUP

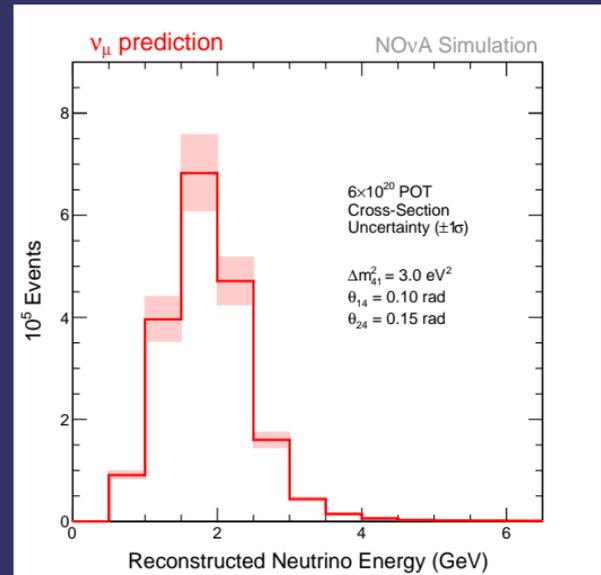
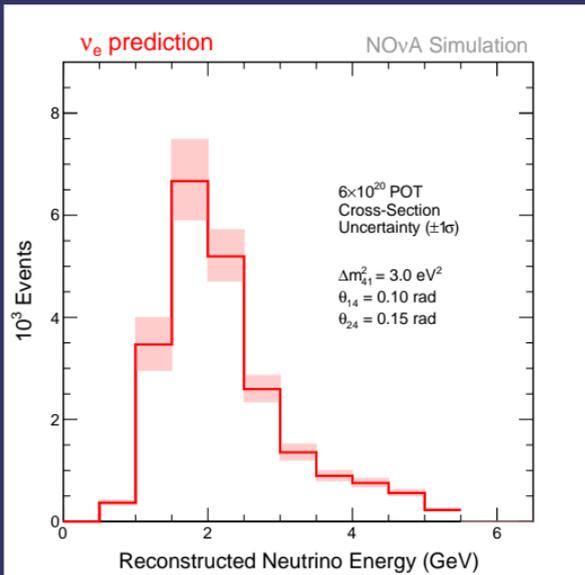
# NUMI BEAM COMPOSITION - FHC



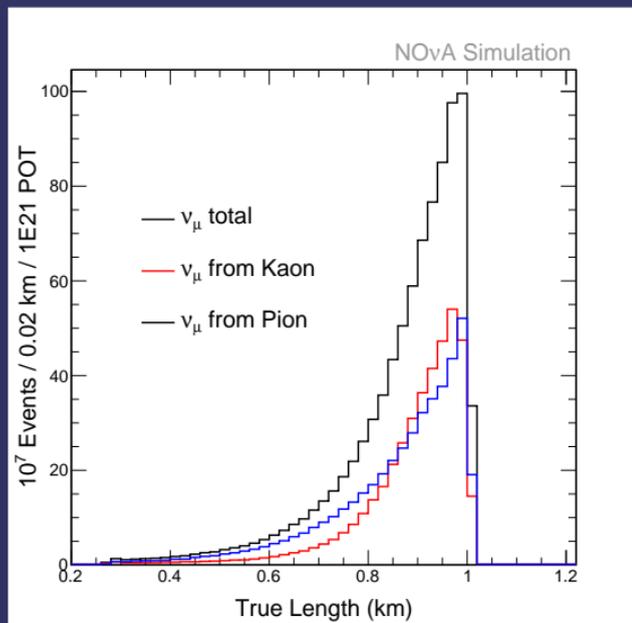
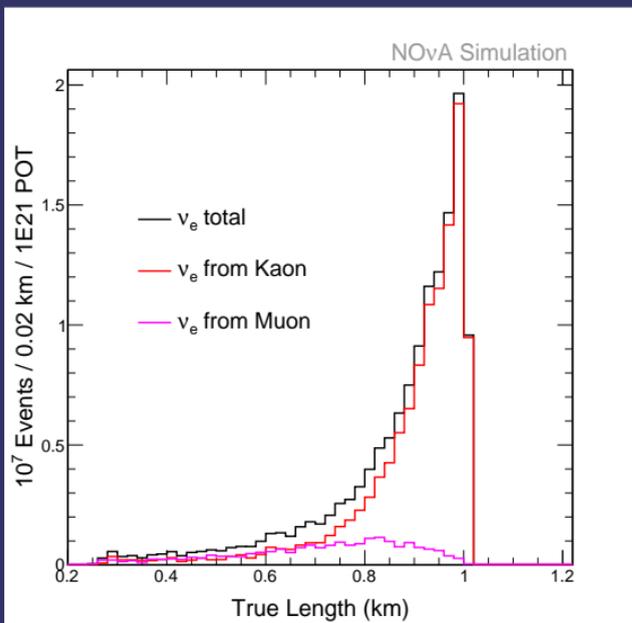
# PREDICTION WITH FLUX SYSTEMATICS



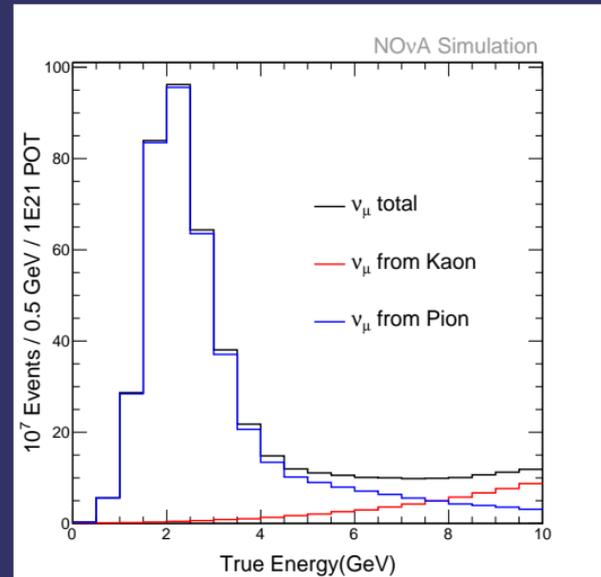
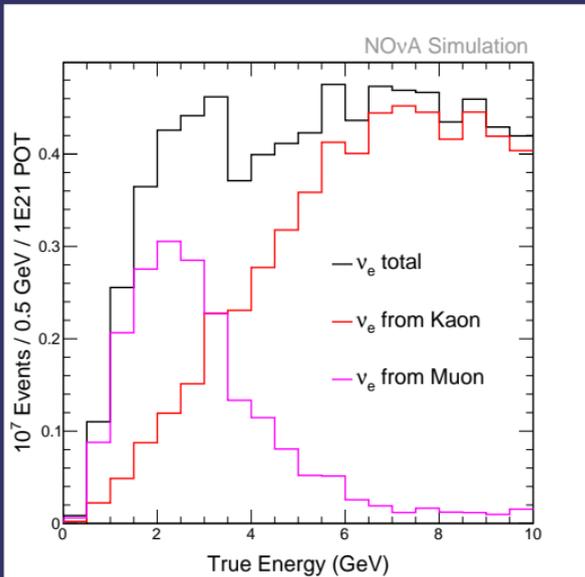
# PREDICTION WITH CROSS-SECTION SYSTEMATICS



# TRUE LENGTH

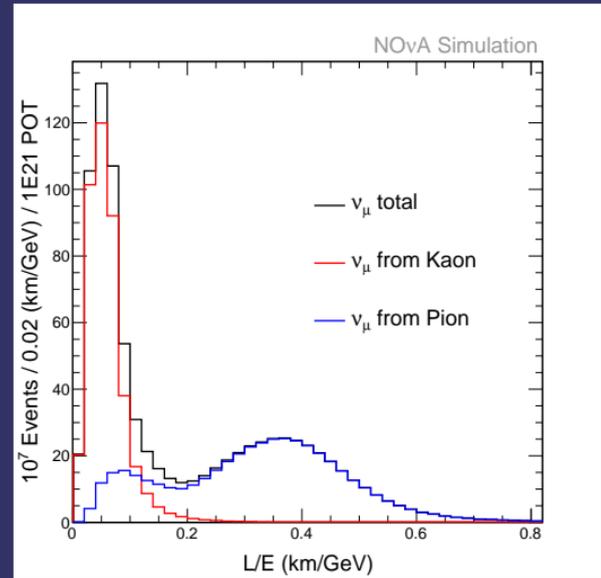
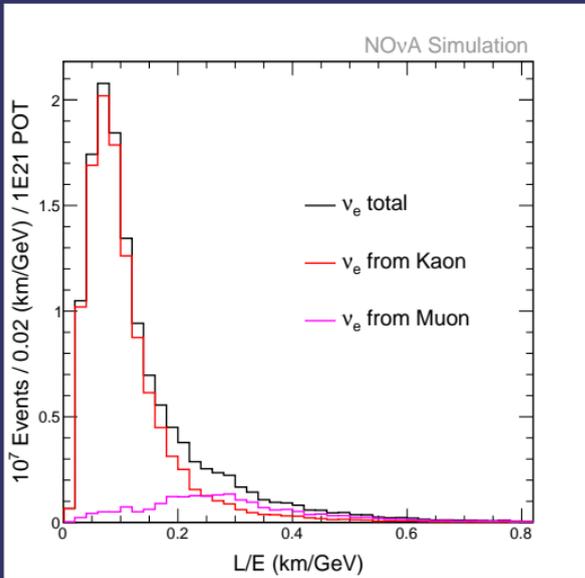


# TRUE ENERGY



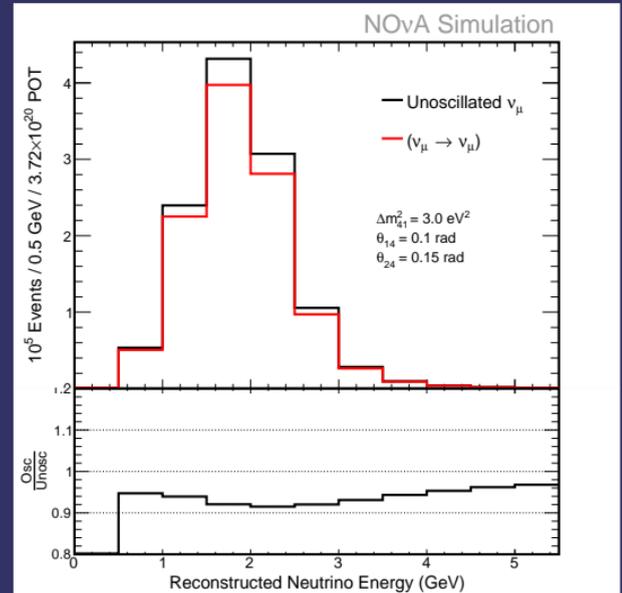
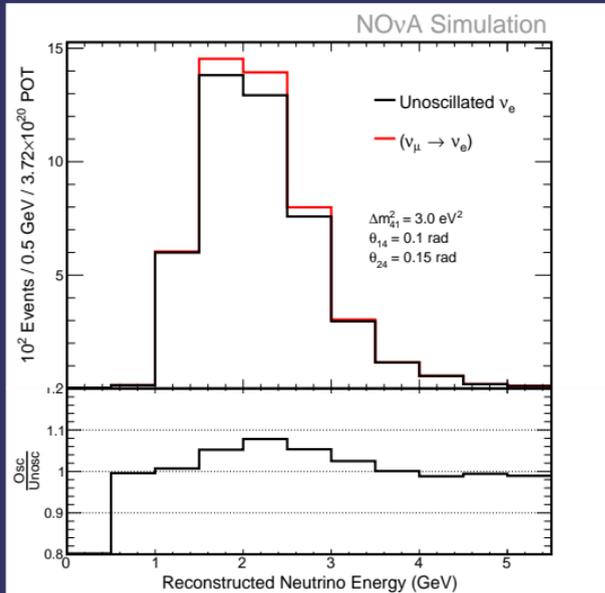
- Distribution of true energy of neutrinos break down by their parents in the near detector.

# TRUE LENGTH OVER TRUE ENERGY



- Distribution of true length over true energy of neutrinos break down by their parents in the near detector.

# SIGNAL PREDICTION



- Predictions, Left:  $\nu_e$  prediction and Right:  $\nu_\mu$  prediction, after applying the  $\nu_e$  and  $\nu_\mu$  selection respectively, for the oscillation parameters  $\Delta m_{41}^2 = 3.0 \text{ eV}^2$ ;  $\theta_{14} = 0.1(\text{rad})$ ;  $\theta_{24} = 0.15(\text{rad})$ .

# SELECTION

## $\nu_e$ selection

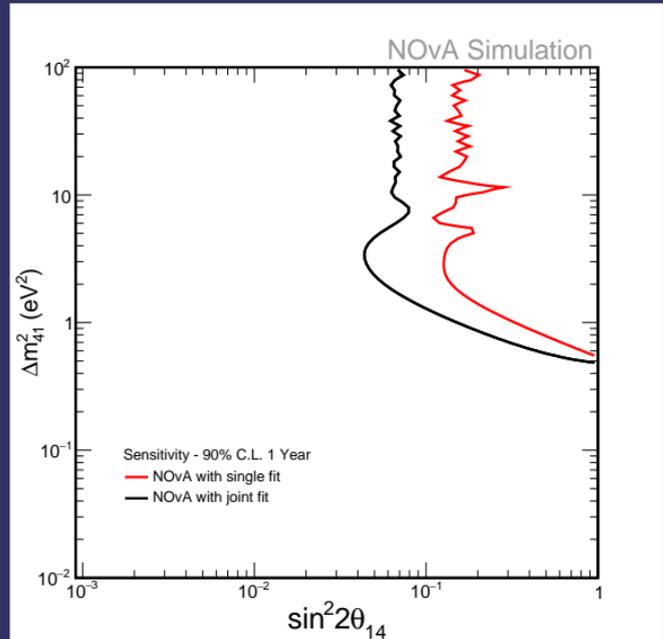
- DQ: flasher cut (hits per plane  $\leq 8$ )
- Reco: each slice has a vertex and 1 3D FuzzyK prong and one shower
- Vertex containment:  $\text{abs}(vX) < 140$ ,  $\text{abs}(vY) < 140$ ,  $100 < vZ < 700$
- Energy: Slice calorimetric energy  $< 5$  GeV
- Loose containment and  $n_{\text{planestofront}} > 6$
- Hits:  $20 < n_{\text{hits}} < 200$
- Length:  $140 < \text{longest fuzzyk prong} < 500$  cm
- Gap: Shower start to vertex gap  $< 100$  cm
- LID  $> 0.95$

## $\nu_\mu$ selection

- kNumuQuality
- kNumuContainND
- kNumuNCRej (ReMID  $> 0.75$ )

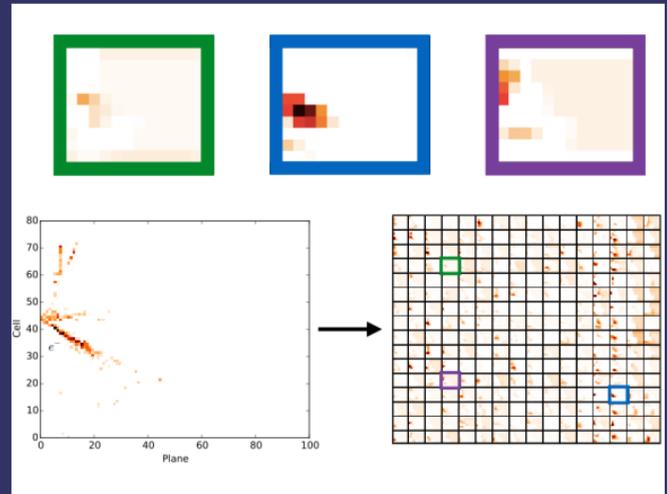
# $\text{SIN}^2 2\theta_{14}$ VS $\Delta M_{41}^2$

- Single Fit: Performed fit for only  $\nu_e$  appearance.
- Joint Fit: Performed fit for both  $\nu_e$  appearance and  $\nu_\mu$  disappearance.
- Fitting  $\nu_e/\nu_\mu$  prediction with  $\nu_e/\nu_\mu$  selected MC spectrum with no (3+1) oscillations.
- Fit for  $\theta_{14}$ , profiling  $\theta_{24}$ , and keeping all parameters fixed.
- No systematics included.



# CONVOLUTIONAL VISUAL NETWORK

- This analysis is based on using the new algorithm for particle identification, Convolutional Visual Network.
- Based on machine learning and deep learning methods.
- No other information is needed except initial hit clustering and energy deposited in hits.
- All interactions with at least 15 hits are used for training.
- CVN achieves the figure of merit equivalent to collecting 30% more exposure



A. Aurisano et al., arXiv:1604.01444.

# CUT FLOW TABLE

Selection	Total MC	$\nu_\mu$	$\nu_e$	NC
No Cut	2.749e+08	2.279e+08	4.062e+06	4.289e+07
Data Quality	3.812e+07	3.366e+07	4.339e+05	4.029e+06
Fiducial	2.703e+06	1.998e+06	5.393e+04	6.516e+05
Containment	7.801e+05	4.525e+05	2.012e+04	3.075e+05
FrontPlanes	7.764e+05	4.505e+05	2.001e+04	3.059e+05
# Hits	7.242e+05	4.354e+05	1.654e+04	2.723e+05
Energy	7.213e+05	4.353e+05	1.469e+04	2.714e+05
ProngLength	5.96e+05	3.862e+05	1.406e+04	1.958e+05
Ptp	5.163e+05	3.412e+05	1.337e+04	1.617e+05
CVN > 0.95	7242	944.1	5171	1127

For  $\nu_e$  for 6e+20 POT.

Selection	Total MC	$\nu_\mu$	$\nu_e$	NC
No Cut	2.749e+08	2.279e+08	4.062e+06	4.289e+07
Data Quality	6.81e+07	6.212e+07	5.978e+05	5.382e+06
Containment	3.565e+06	2.484e+06	5.777e+04	1.022e+06
CVN > 0.5	1.908e+06	1.889e+06	365.4	1.894e+04

For  $\nu_e$  for 6e+20 POT.

# JOINT FIT METHOD

- The formalism to combine several measurements by performing a joint maximum-likelihood fit of different datasets. This analysis is done using log-likelihoods.

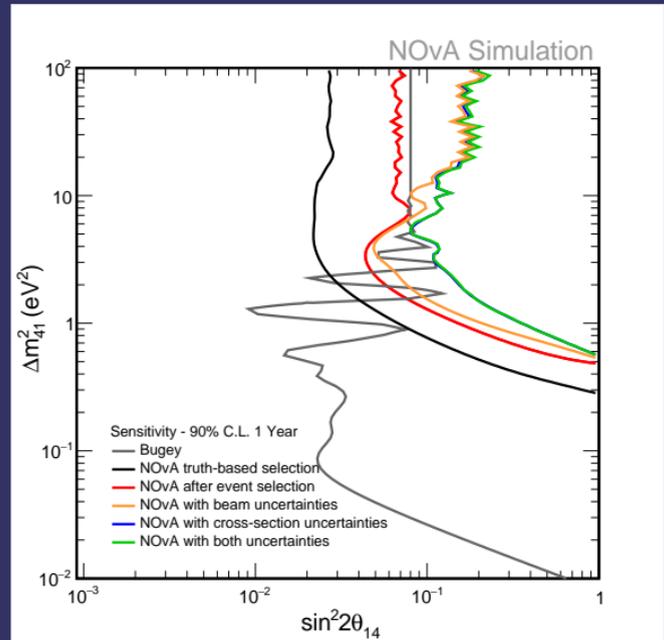
$$\mathcal{L}_{\text{Joint}} = \mathcal{L}_A + \mathcal{L}_B$$

where  $\mathcal{L}$  is Log-likelihood. A and B are datasets.

- Combine  $\nu_\mu$  disappearance and  $\nu_e$  appearance performing a joint fit for the datasets after applying  $\nu_\mu$  and  $\nu_e$  selection.

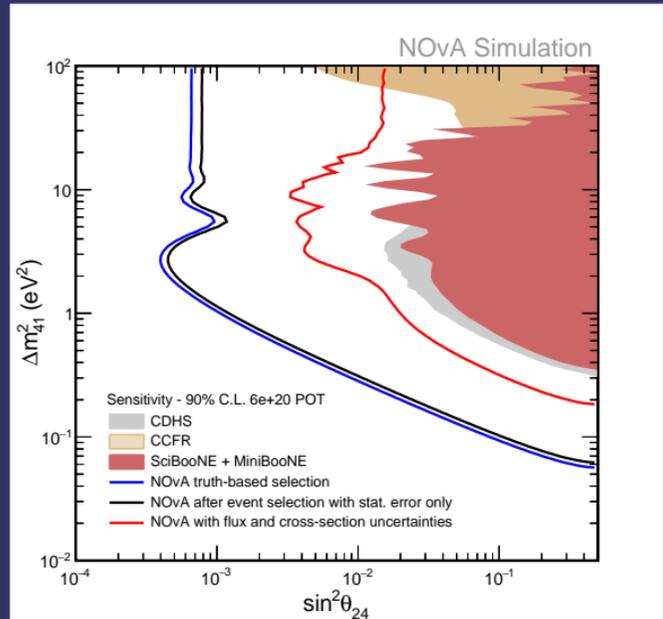
# SENSITIVITY - $\theta_{14}$

- Fitting  $\nu_e/\nu_\mu$  prediction with  $\nu_e/\nu_\mu$  selected MC spectrum with no (3+1) oscillations.
- Fit for  $\theta_{14}$ , profiling  $\theta_{24}$ , and keeping all parameters fixed.
- Included flux and cross-section uncertainty.
- With the improved selection and understanding of systematics, improves the sensitivity in a better way.



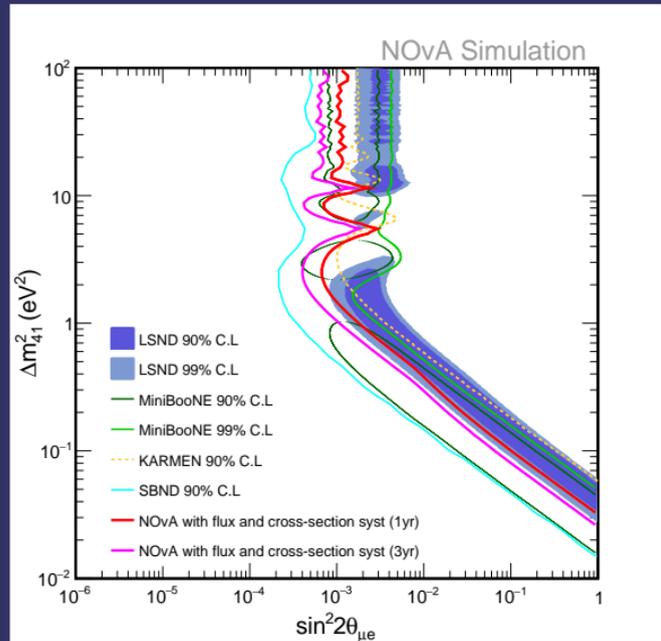
# SENSITIVITY - $\theta_{24}$

- Fitting  $\nu_e/\nu_\mu$  prediction with  $\nu_e/\nu_\mu$  selected MC spectrum with no (3+1) oscillations.
- Fit for  $\theta_{14}$ , profiling  $\theta_{24}$ , and keeping all parameters fixed.
- Included flux and cross-section uncertainty.
- With the improved selection and understanding of systematics, improves the sensitivity in a better way.



# SENSITIVITY - $\theta_{\mu e}$

- Fitting simultaneously both  $\nu_e$  and  $\nu_\mu$  prediction with  $\nu_e$  and  $\nu_\mu$  selected MC spectrum with no (3+1) oscillations.
- Fit for  $\theta_{\mu e}$ , profiling  $\theta_{14}$ ,  $\theta_{24}$  and  $\theta_{34}$ , and keeping all other parameters fixed.
- Included flux and cross-section uncertainty into the fit.



$$\sin^2 2\theta_{\mu e} = \sin^2 2\theta_{14} \sin^2 \theta_{24}.$$