

NOvA Short-Baseline Tau Neutrino Appearance Search

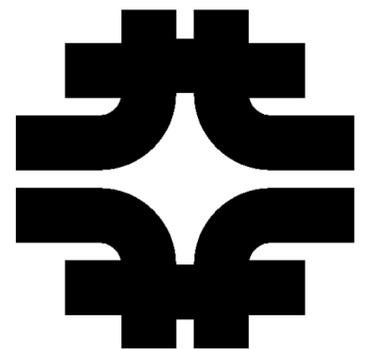
Rijeesh Keloth

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For The NOvA Collaboration

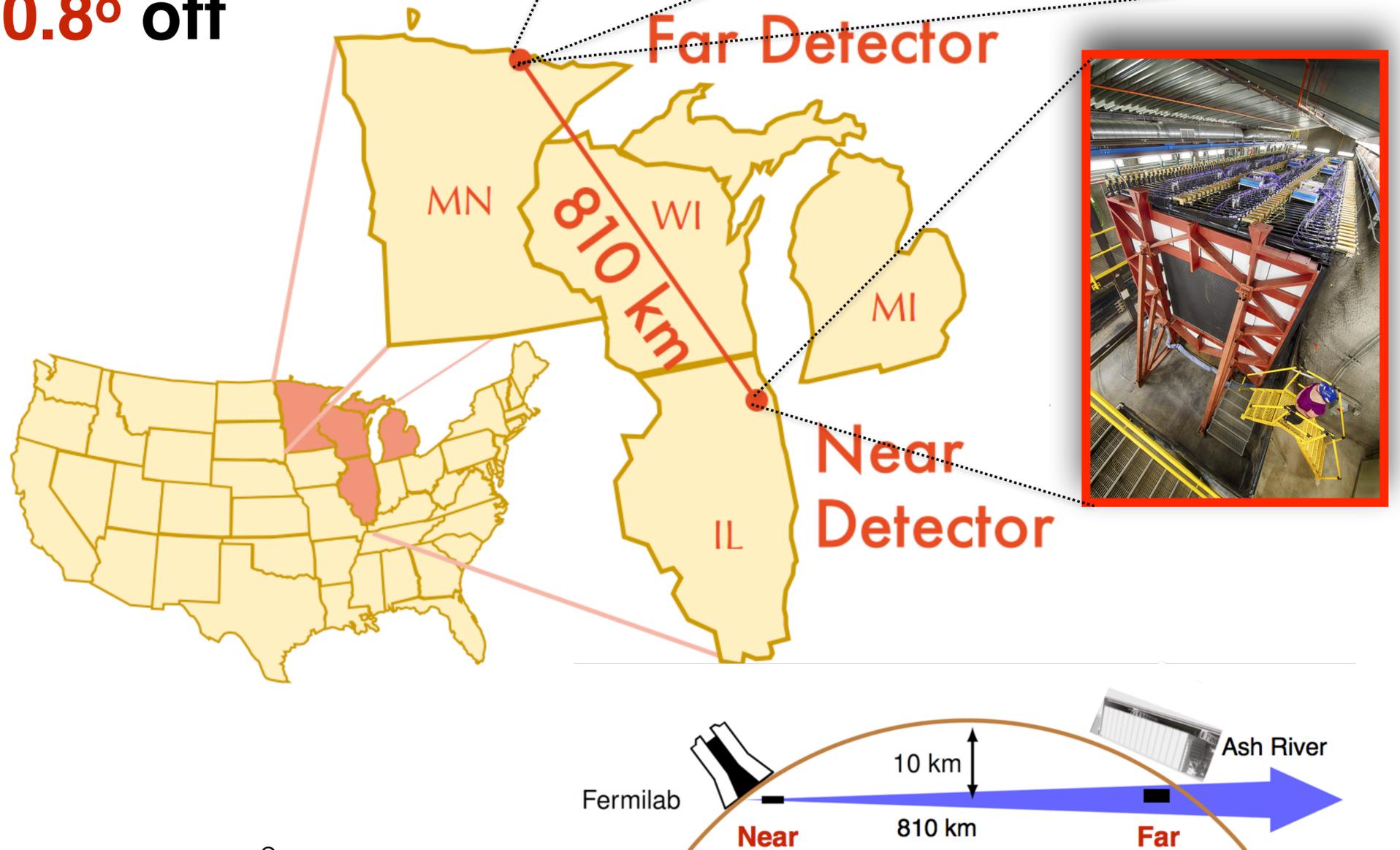


New Perspectives 2017
05 June 2017, Fermilab



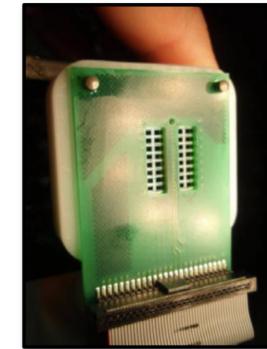
NOvA Experiment

- ◆ **Off-axis** long-baseline neutrino oscillation experiment
- ◆ Narrow-band beam peaked at 2 GeV as detectors are located **0.8°** off NuMI beam axis
- ◆ **Near Detector:**
 - 105 m underground
 - 1 km from target
 - 0.3 kton
- ◆ **Far Detector:**
 - Located on surface
 - 810 km from target
 - 14 kton



NOvA Detectors

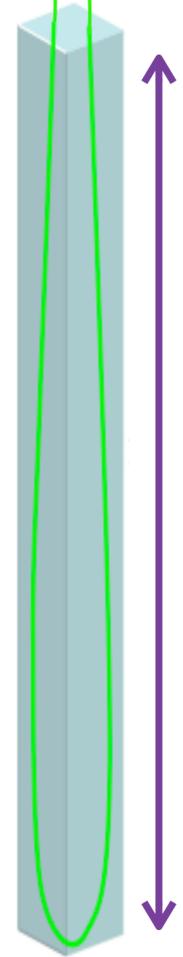
- ◆ **Functionally identical detectors made of planes of extruded PVC cells, alternating horizontal and vertical planes to provide 3D tracking.**
- ◆ **Each cell filled with liquid scintillator, light collected in wavelength shifting fibers and fed into avalanche photodiode.**



32-pixel APD

A NOvA cell

To APD



4.1 m

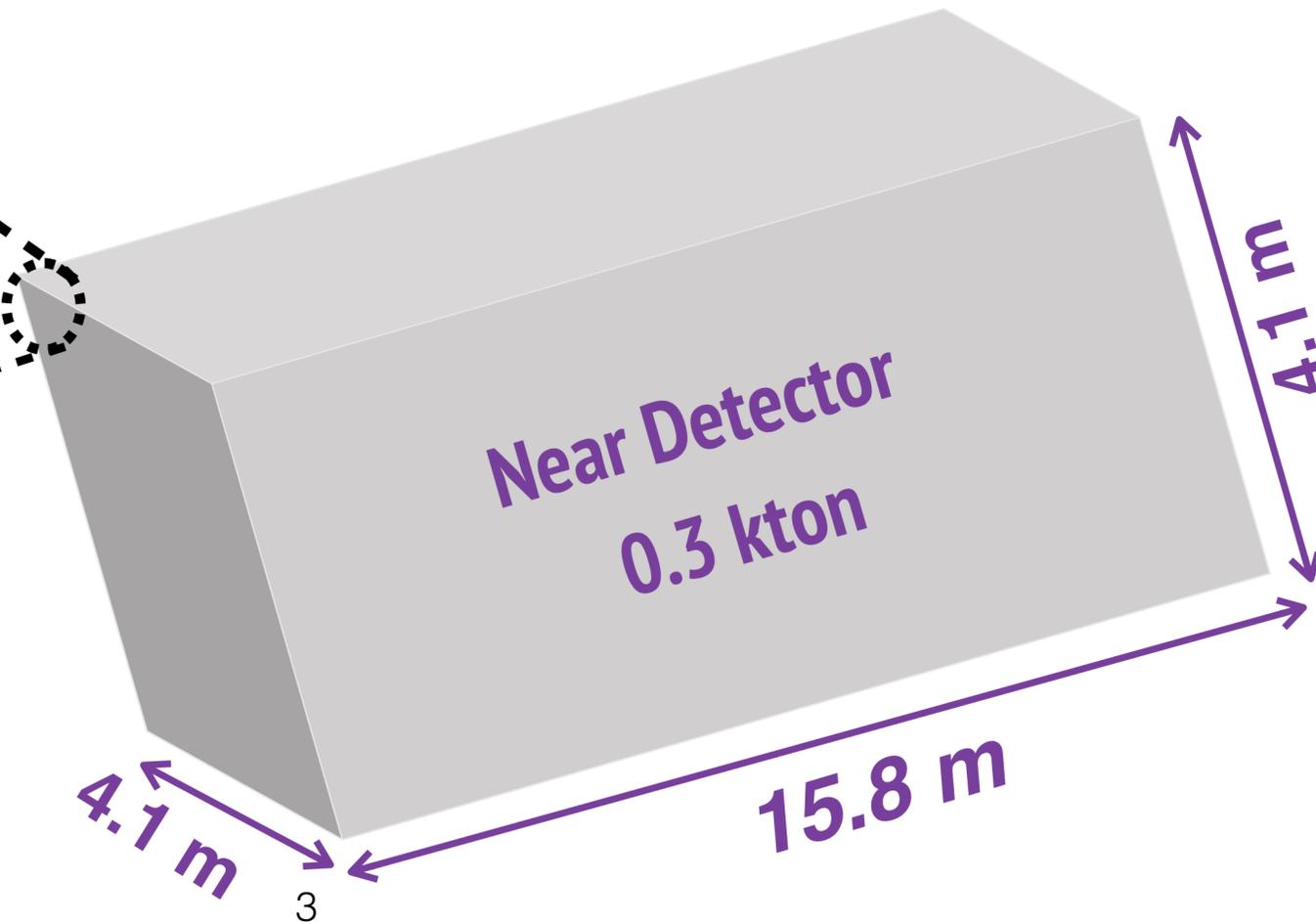
4 cm x 6 cm

Near detector:

Fine-grained,
low-Z, highly-active
tracking calorimeter

→ 20k PVC cells

→ 65% active by mass

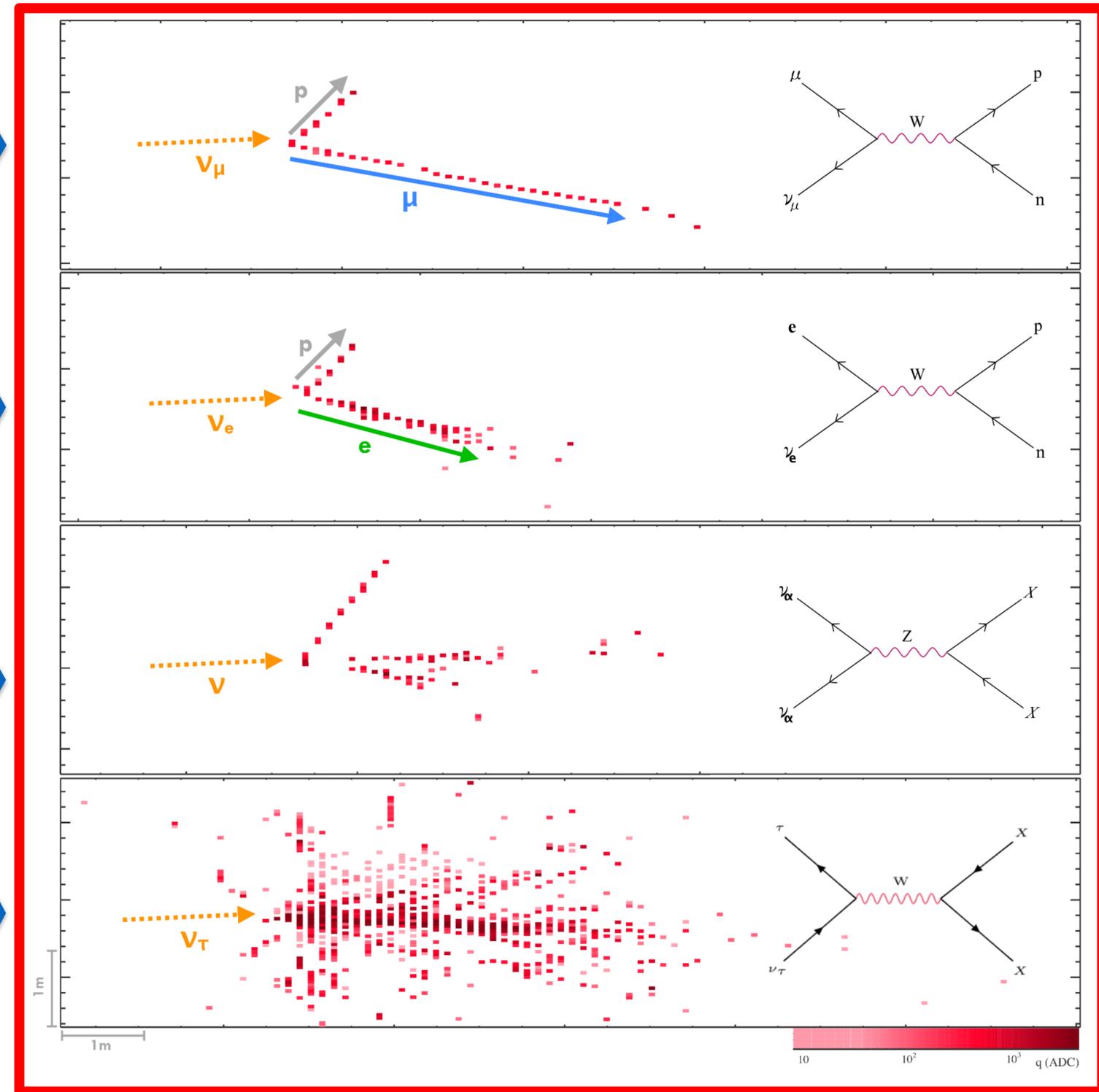
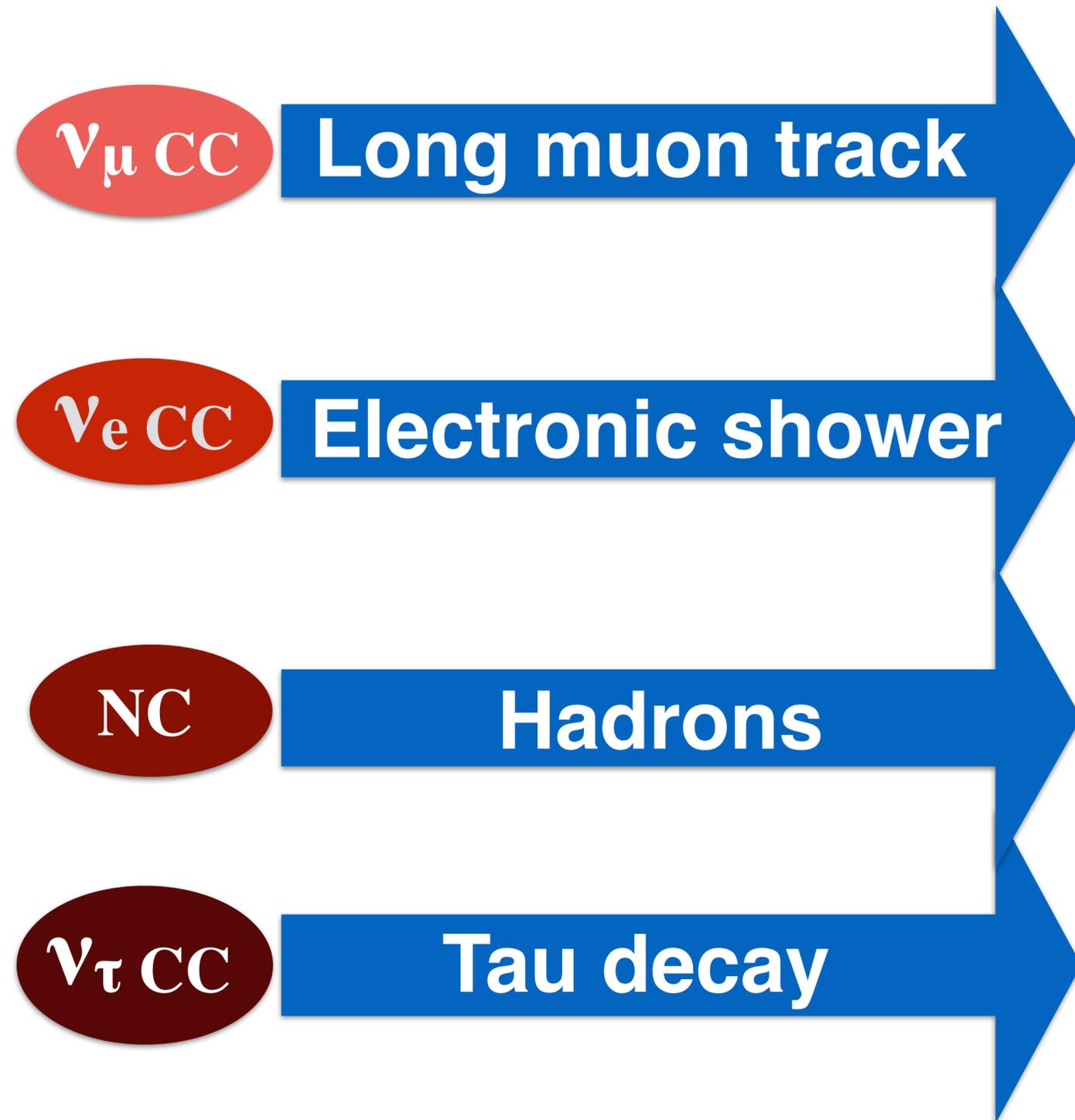


4.1 m

15.8 m

Near Detector
0.3 kton

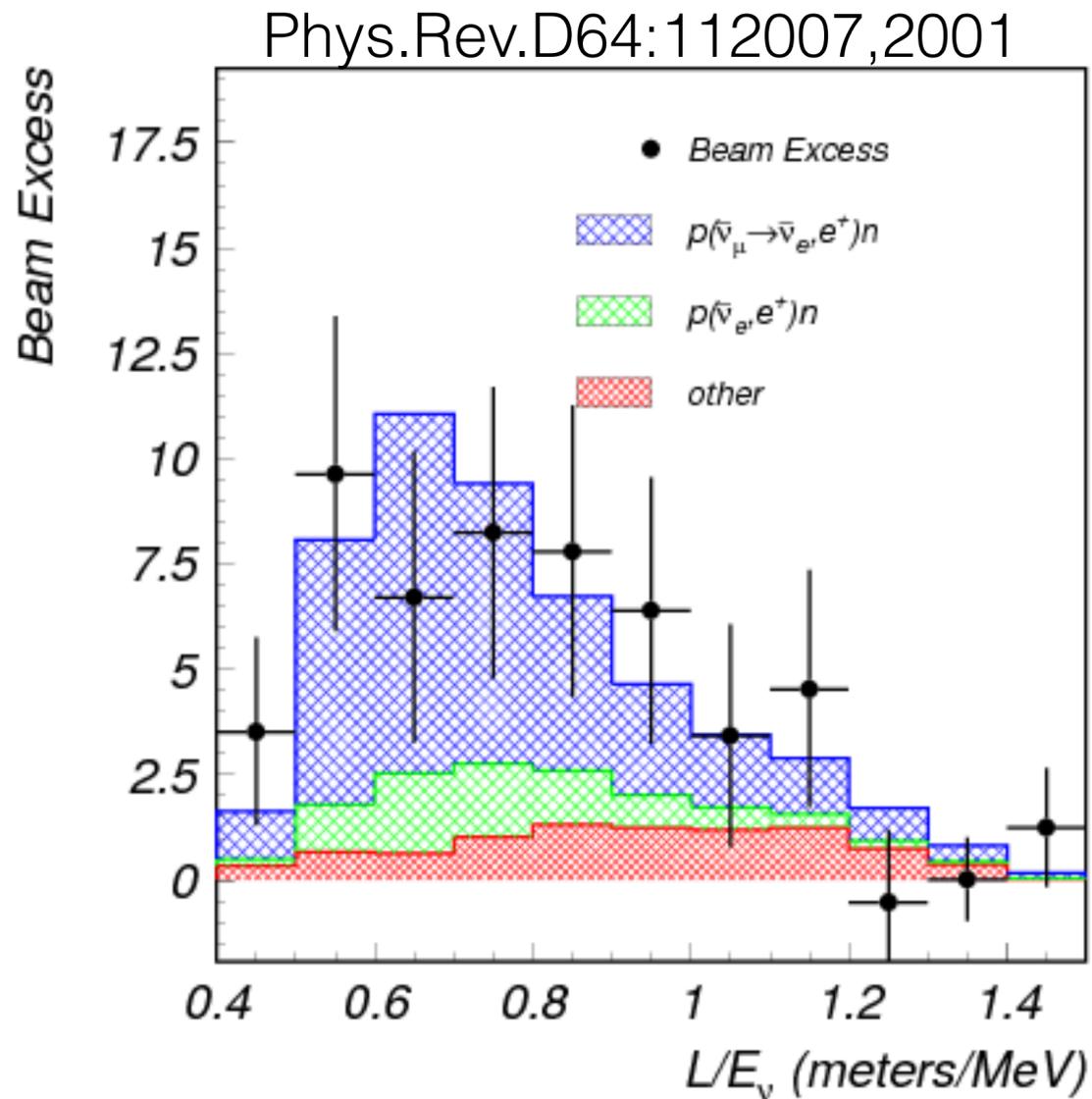
Neutrino Interactions in NOvA Detectors



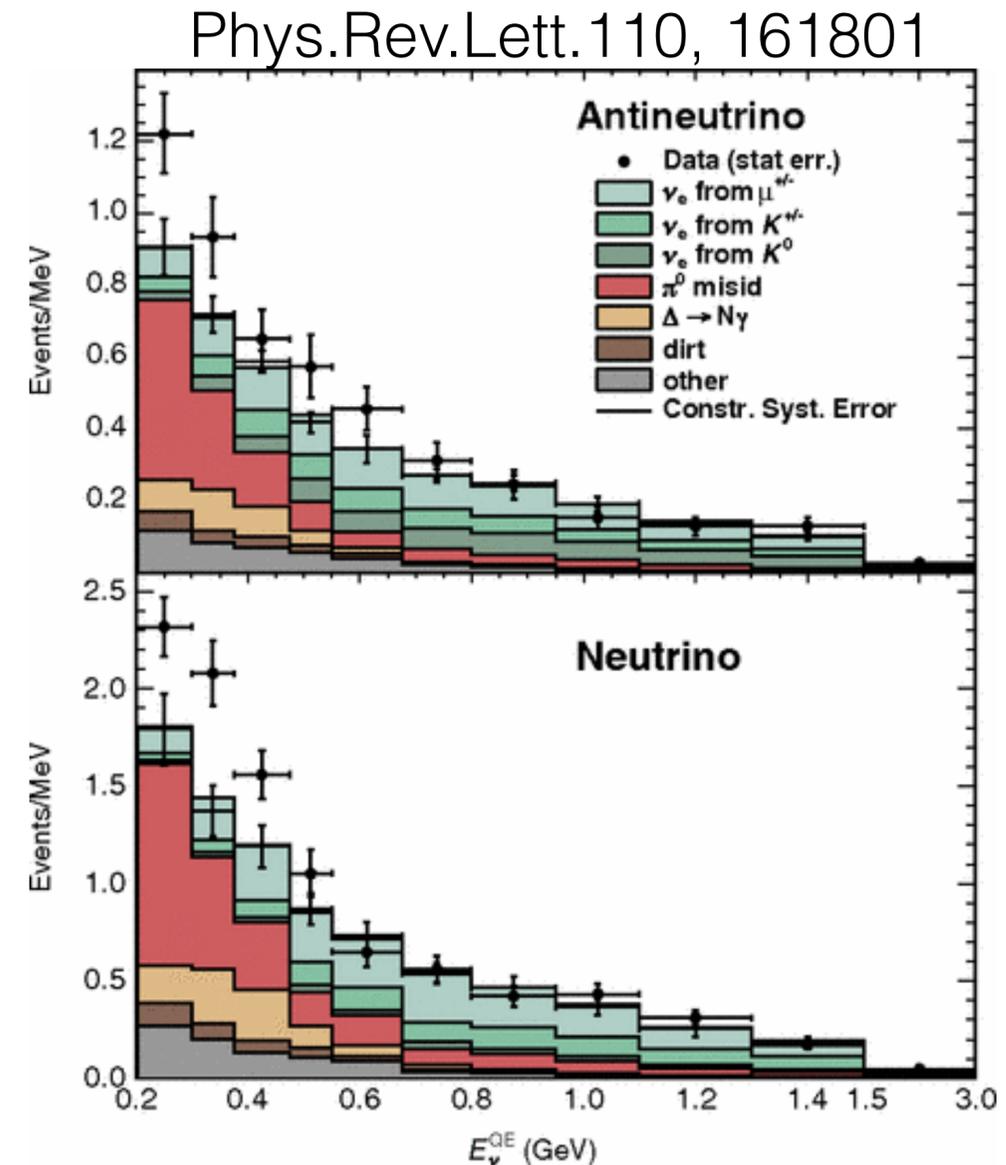
Short-Baseline Oscillations

Why we are interested in looking for sterile neutrinos?

- LSND and miniBooNE reported a $\bar{\nu}_e$ excess in anti-neutrino mode
- Is this evidence that **sterile neutrinos** exist?



LSND Beam Excess

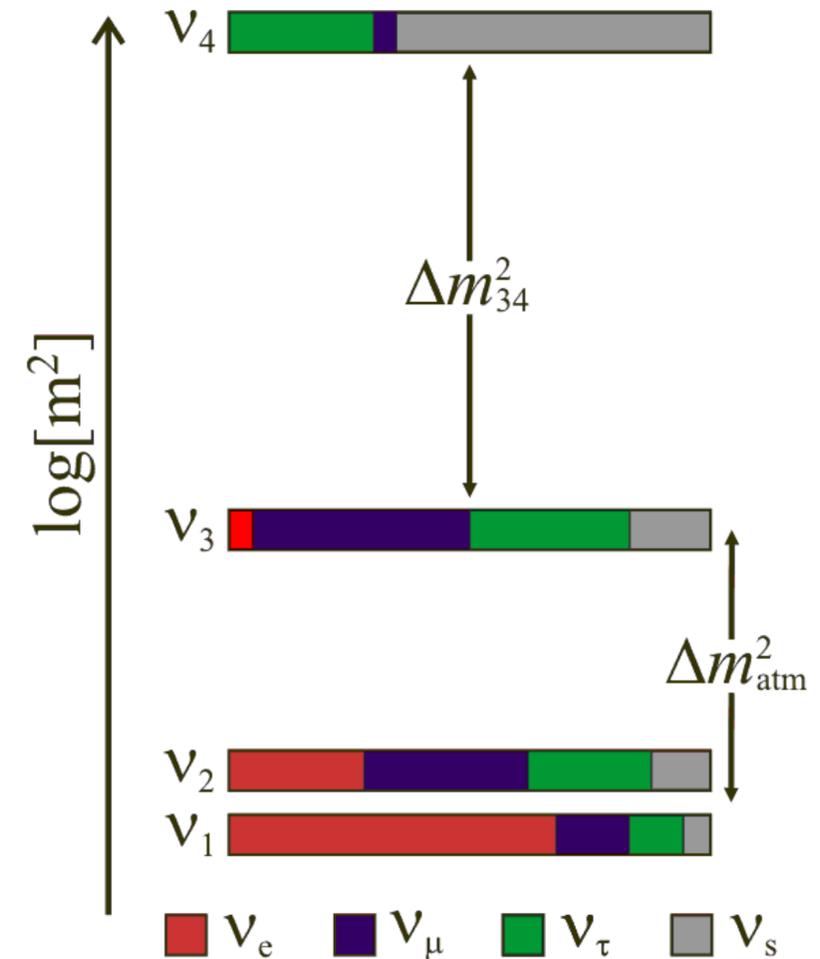


MiniBooNE Beam Excess

Sterile Neutrino Oscillations

3+1 Oscillation Model

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{pmatrix}$$



- ◆ Looking for sterile neutrino oscillations using a two-flavor approximation of the 3+1 model, governed by a single large mass-squared splitting

Two-flavor Oscillation Model

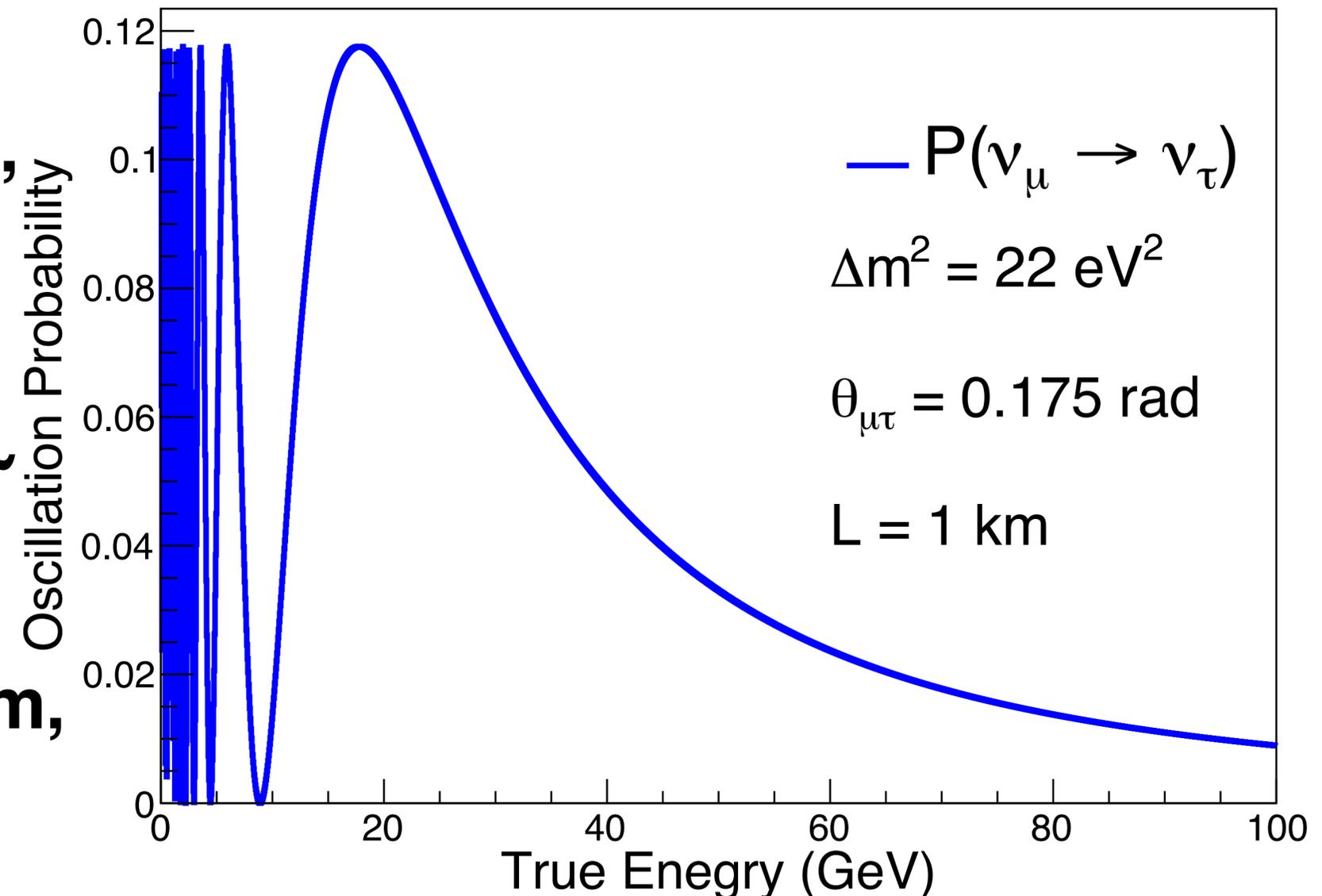
Two-Flavor $\nu_\mu \rightarrow \nu_\tau$ oscillations probability (SBL Approximation)

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta_{\mu\tau} \sin^2(1.27 \Delta m^2 L / E_\nu)$$

NOvA Simulation

- ◆ Appearance of ν_τ at the NOvA ND, due to oscillations with mass squared difference between the active and sterile neutrino, $\Delta m^2 \sim 10 - 100 \text{ eV}^2$

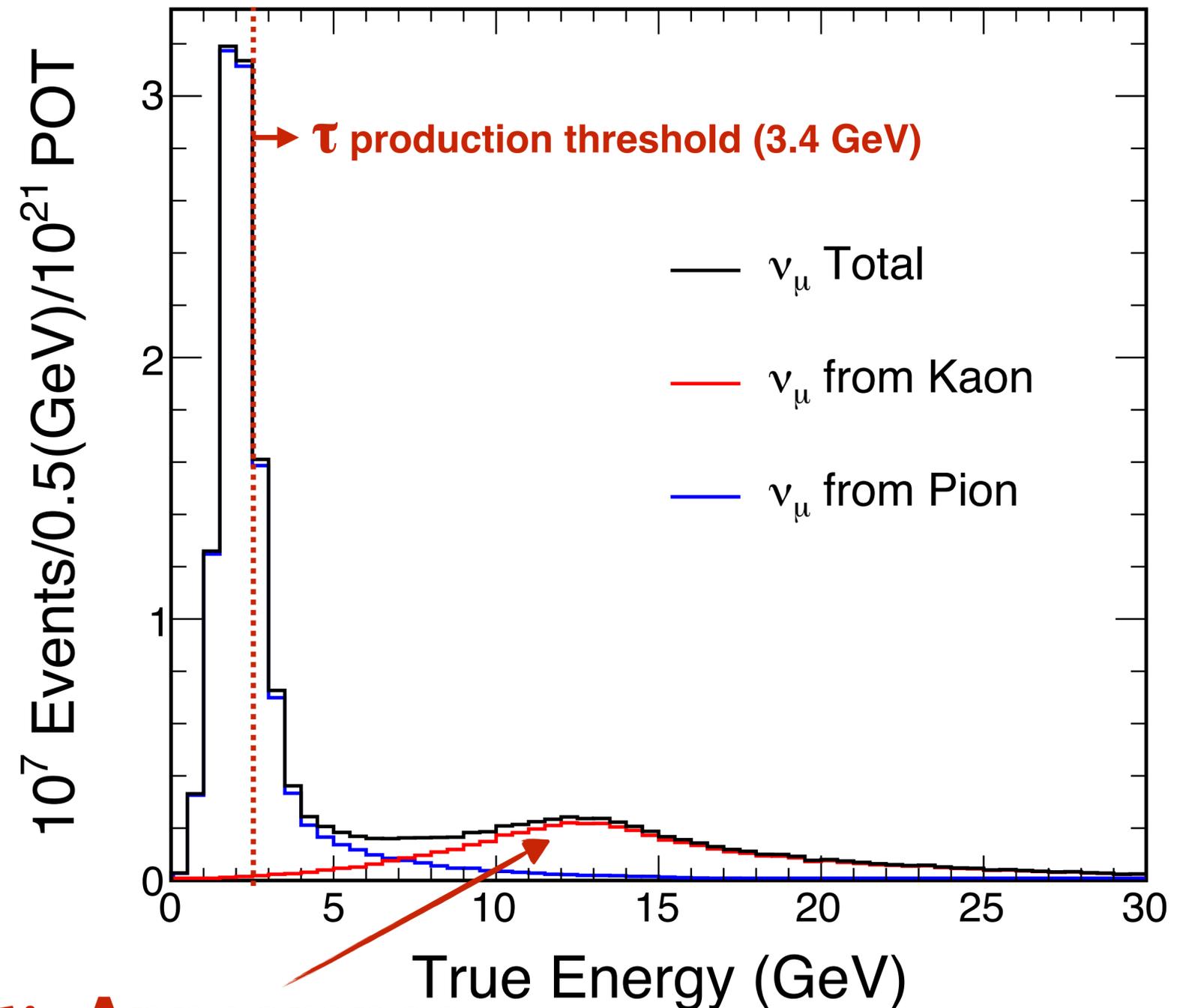
- ◆ Oscillations while traveling $\sim 1 \text{ km}$, from Source to Near Detector



SBL ν_τ Appearance in NOvA

NOvA Simulation

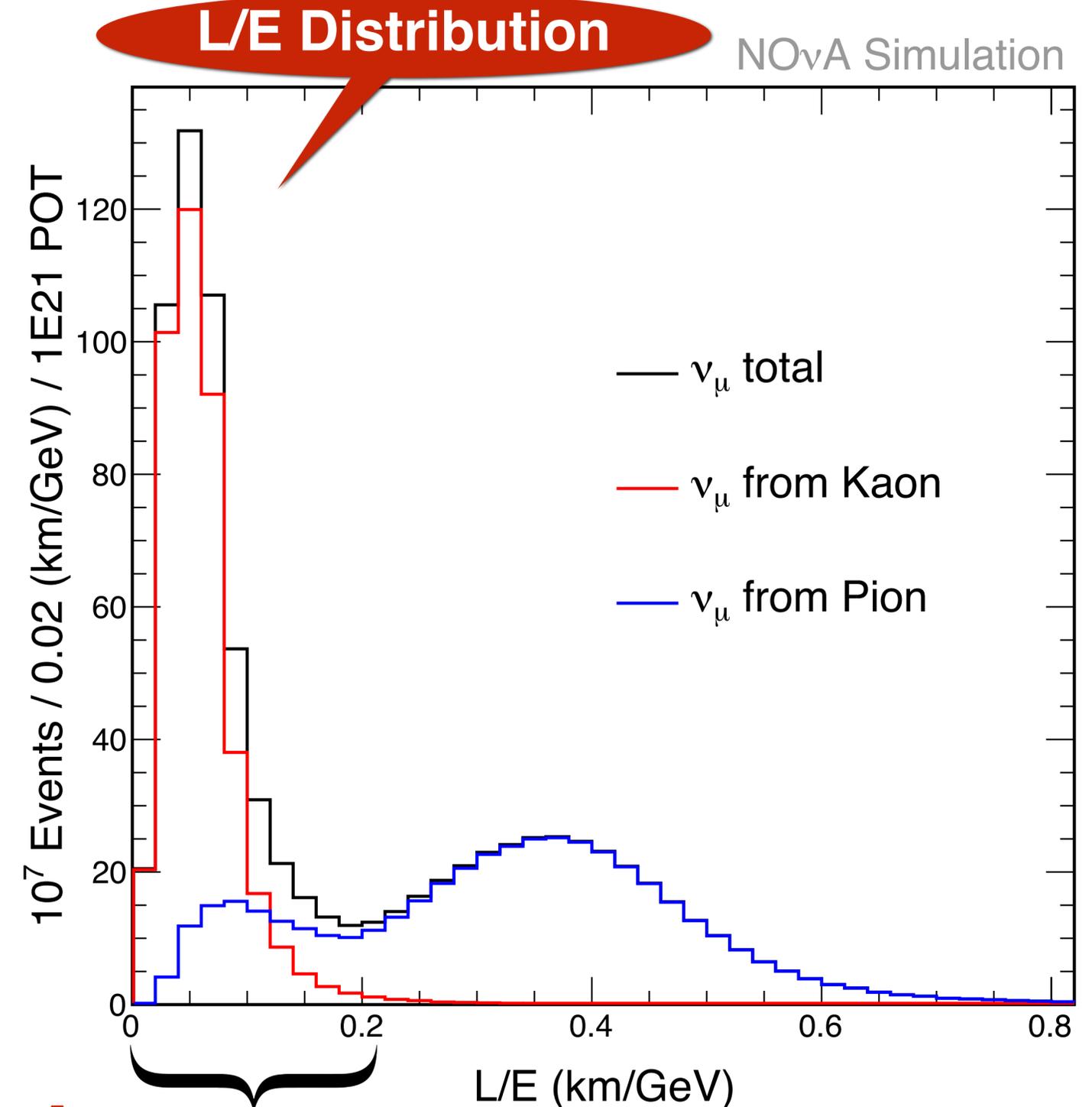
- ◆ Neutrinos in narrow-band beam peaked at 2 GeV predominantly created by pion decays
- ◆ But τ production threshold is 3.4 GeV
- ◆ Focusing on high-energy neutrinos resulting from kaon decays to search for SBL $\nu_\mu \rightarrow \nu_\tau$ oscillations



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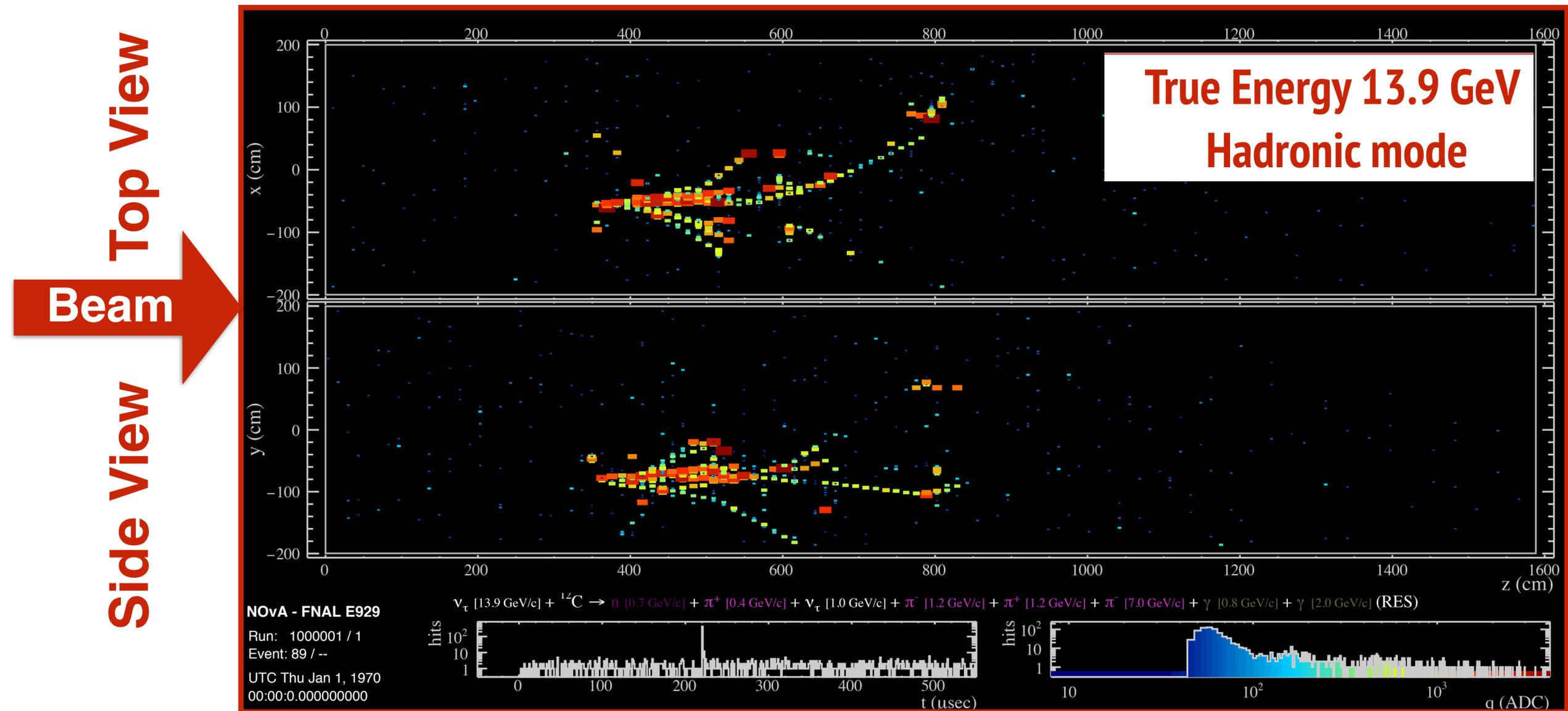
SBL ν_τ Appearance

LSND sensitive region(0.4 -1.4)

ν_τ Detection in NOvA Near Detector

- ◆ ν_τ CC interactions in the ND occur primarily at greater than ~ 5 GeV
- ◆ High-multiplicity hadronic systems originating from nuclear scattering.
- ◆ In addition we see the decay products of the outgoing τ lepton in the final state.
- ◆ Categorize ν_τ CC interaction using τ decay modes

Simulated ν_τ CC in NOvA Near Detector

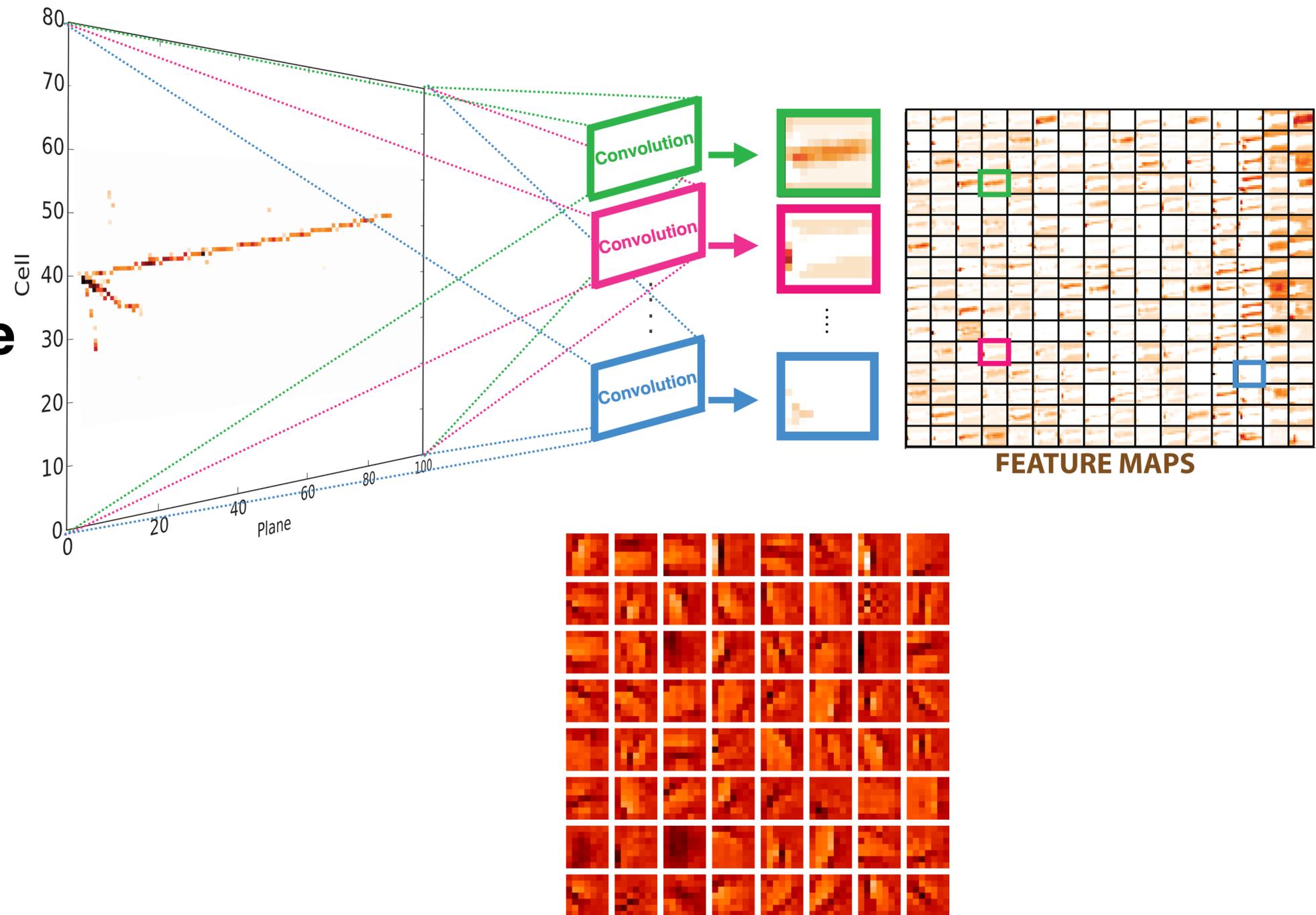


- **Hadronic:** one outgoing ν_τ and typically one or three pions
- **Leptonic:** two outgoing neutrinos and an electron or muon.

Convolutional Neural Network Based PID

- ◆ Use the tools from the computer vision community to classify neutrino interactions.
- ◆ **Convolutional Visual Network (CVN)**, based on the **GoogLeNet convolutional neural network (CNN)** architecture.
- ◆ Treat each event as an image, and pass these images through layers of learned filters to extract the features.

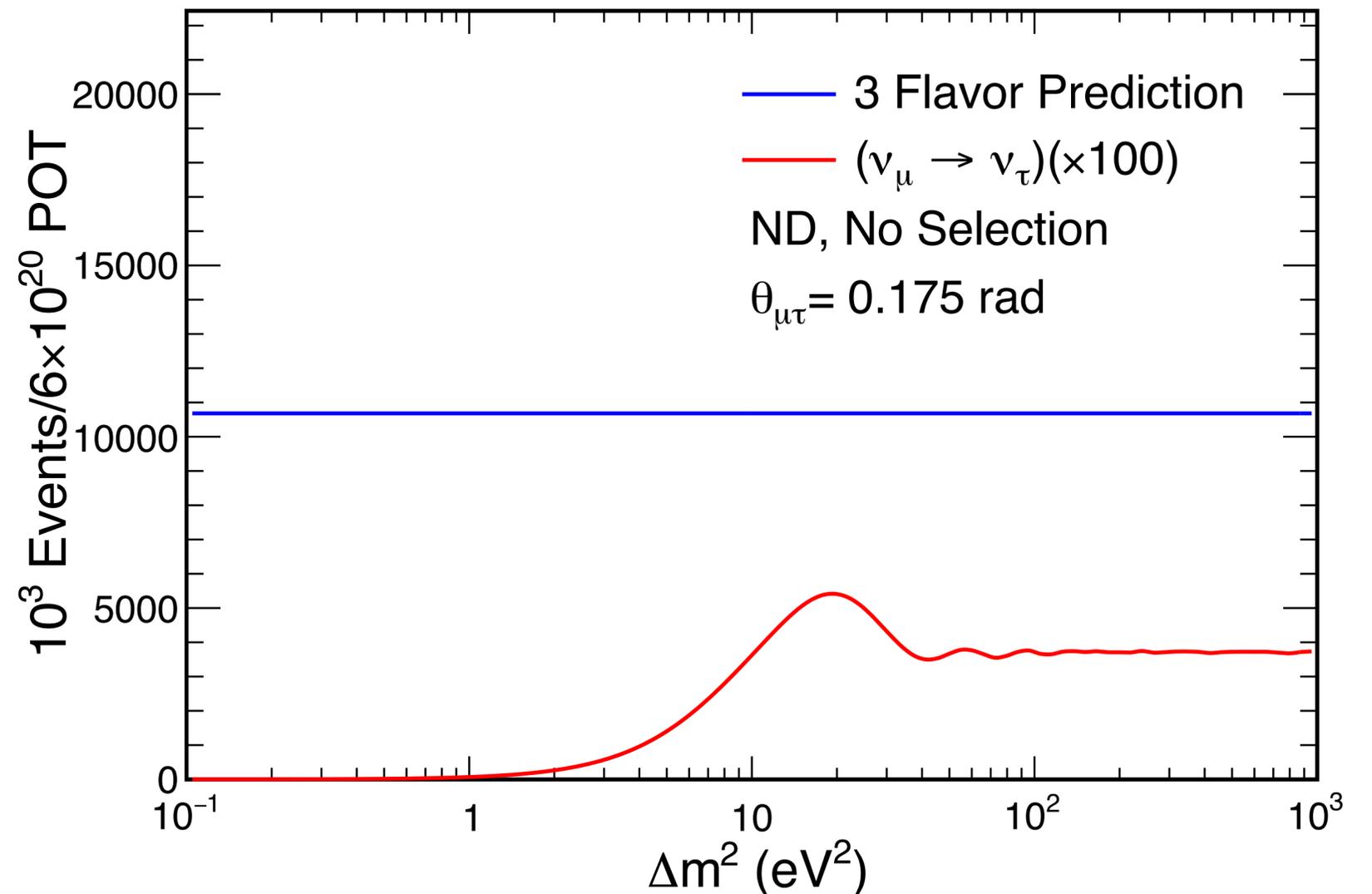
A. Aurisano et al., 2016 JINST 11 P09001



Signal Selection

NOvA Simulation

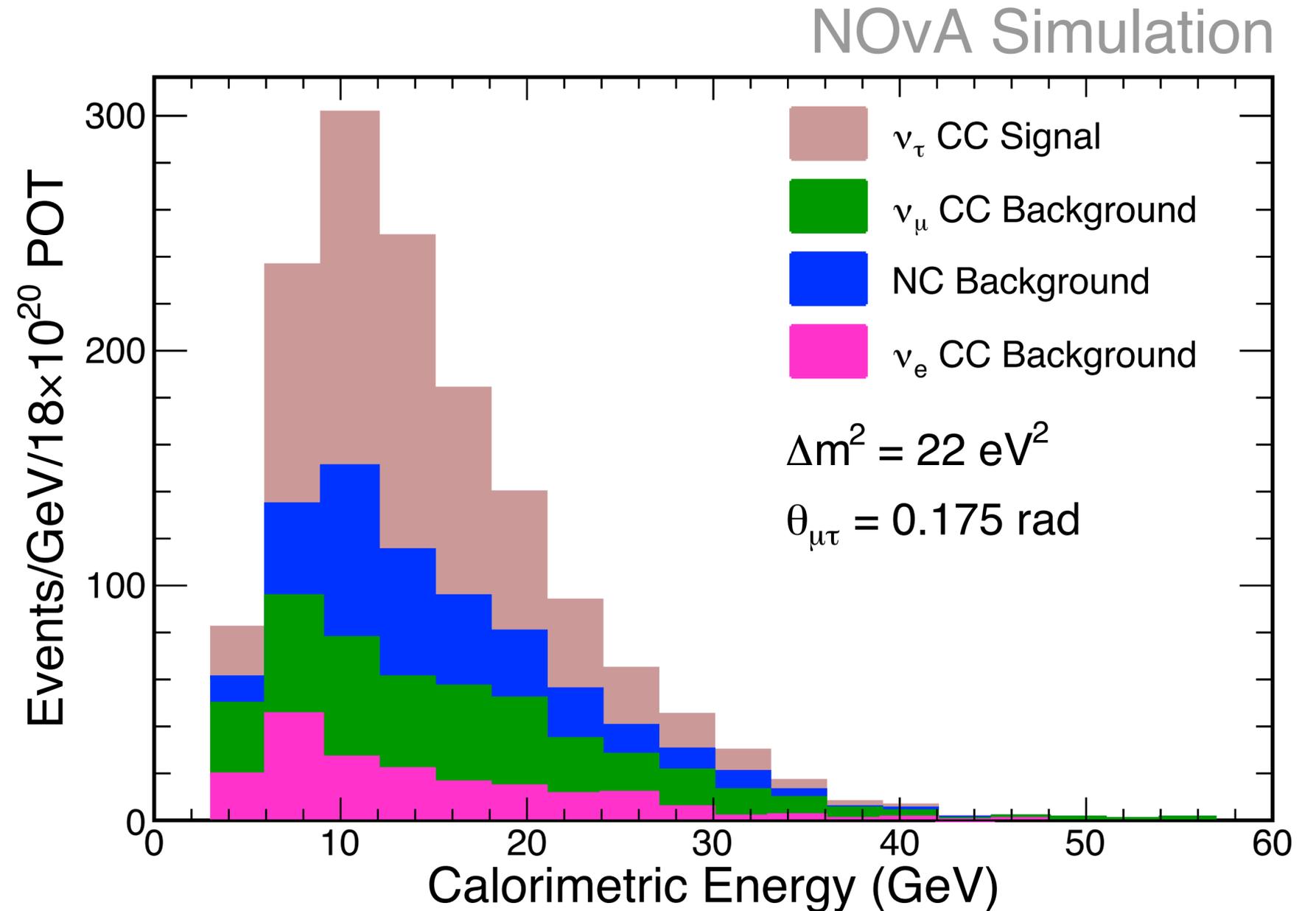
- ◆ **CVN is used as a primary selector**
- ◆ **Preselection: Cuts to remove the detector external activity**



The ν_τ CC signal is maximized for $\Delta m^2 = 22 \text{ eV}^2$ corresponding to small L/E for muon neutrinos originating from kaon decays.

ν_τ CC Signal Prediction

◆ **Signal and Background predictions for three nominal years of NOvA running (18×10^{20} POT)**

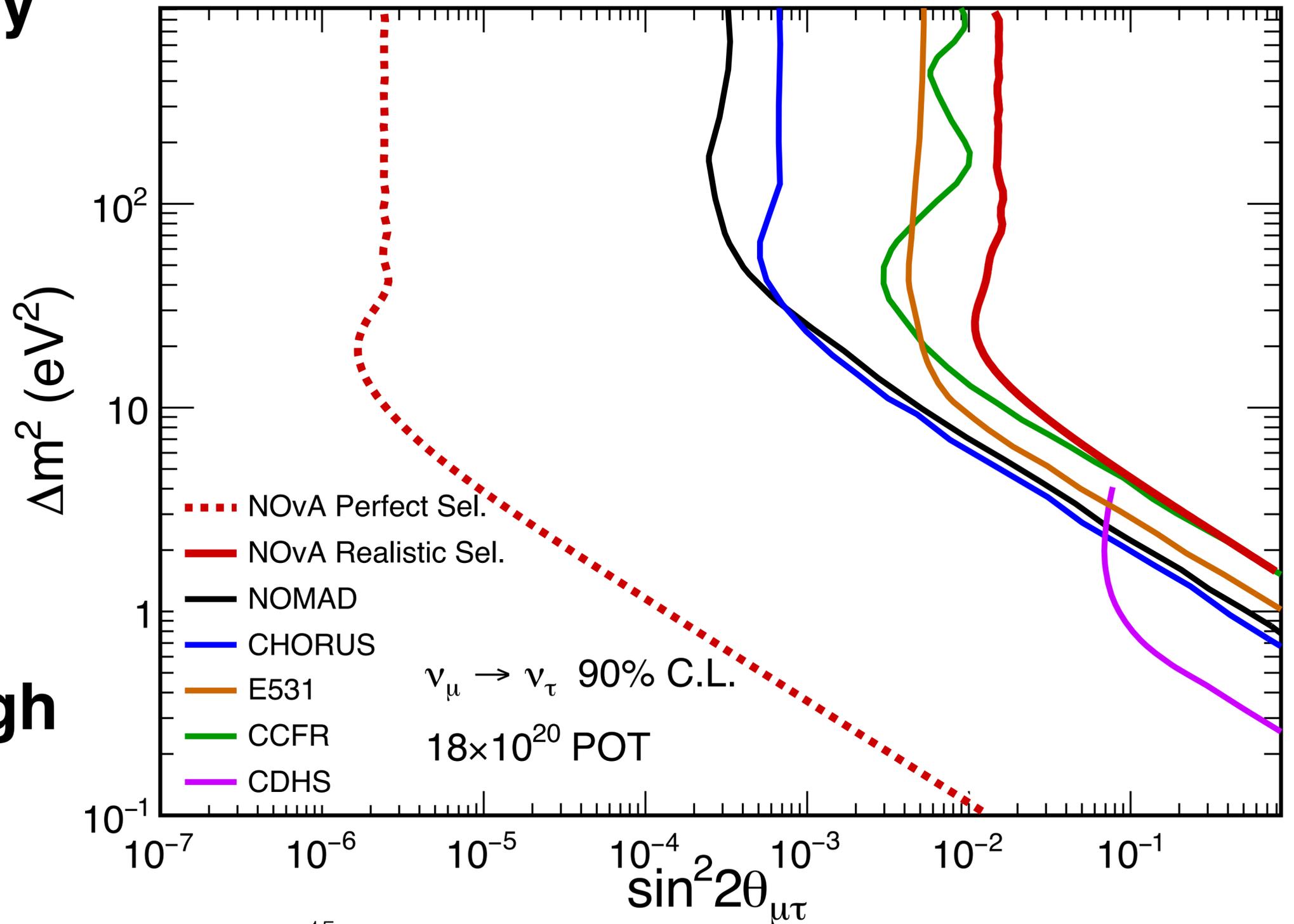


Calorimetric energy distribution of ν_τ CC signal and background stacked plot

NOvA Sensitivity

NOvA Simulation

- ◆ **Stats-only sensitivity plot**
- ◆ **NOvA sensitivity competitive with previous experimental limits**
- ◆ **Much room for improvement through optimization of τ selection**



Summary

- ◆ **Search for sterile neutrinos using $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillations in NOvA is very promising.**
- ◆ **Preliminary sensitivity shows NOvA will be competitive with previous experiments after 3 years of running.**
- ◆ **Started assessing effects of systematic uncertainties**
- ◆ **Future work will include modifying the PID by categorizing ν_{τ} CC according to τ decay modes**

Stay tuned for the upcoming results...



www-nova.fnal.gov

NOvA Collaboration



Argonne, Atlantico, Banaras Hindu University, Caltech, Cochin, Institute of Physics and Computer science of the Czech Academy of Sciences, Charles University, Cincinnati, Colorado State, Czech Technical University, Delhi, JINR, Fermilab, Goiás, IIT Guwahati, Harvard, IIT Hyderabad, U. Hyderabad, Indiana, Iowa State, Jammu, Lebedev, Michigan State, Minnesota-Twin Cities, Minnesota-Duluth, INR Moscow, Panjab, South Carolina, SD School of Mines, SMU, Stanford, Sussex, Tennessee, Texas-Austin, Tufts, UCL, Virginia, Wichita State, William and Mary, Winona State

Backup

3+1 Oscillation Model

Sterile neutrino Oscillations(SBL Approximation)

- $\nu_\mu \rightarrow \nu_\tau$ oscillations using 3+1 model is sensitive to θ_{34} and θ_{24}

$$\sin^2 2\theta_{\mu\tau} = 4 |U_{\mu 4}|^2 |U_{\tau 4}|^2$$

$$|U_{\tau 4}|^2 = \cos^2 \theta_{14} \cos^2 \theta_{24} \sin^2 \theta_{34}$$

$$|U_{\mu 4}|^2 = \cos^2 \theta_{14} \sin^2 \theta_{24}$$

$$\sin^2 2\theta_{\mu\tau} = \cos^4 \theta_{14} \sin^2 2\theta_{24} \sin^2 \theta_{34}$$

3+1 Oscillation Model

Sterile neutrino Oscillations(SBL Approximation)

(3+1) Model ν_e Appearance Probability²

$$P_{\nu_{\mu}^{(-)} \rightarrow \nu_e^{(-)}}^{\text{SBL},3+1} = 4|U_{\mu 4}|^2|U_{e 4}|^2 \sin^2 \frac{\Delta m_{41}^2 L}{4E} = \sin^2 2\theta_{\mu e} \sin^2 \frac{\Delta m_{41}^2 L}{4E}$$

Where

$$\sin^2 2\theta_{\mu e} \equiv 4|U_{\mu 4}|^2|U_{e 4}|^2$$

and

$$\sin^2 2\theta_{\mu e} = \sin^2 \theta_{24} \sin^2 2\theta_{14}$$

(3+1) Model ν_{μ} Disappearance Probability

$$P_{\nu_{\mu}^{(-)} \rightarrow \nu_{\mu}^{(-)}}^{\text{SBL},3+1} = 1 - 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2) \sin^2 \frac{\Delta m_{41}^2 L}{4E} = 1 - \sin^2 2\theta_{\mu\mu} \sin^2 \frac{\Delta m_{41}^2 L}{4E}$$

Where

$$\sin^2 2\theta_{\mu\mu} \equiv 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2)$$

and

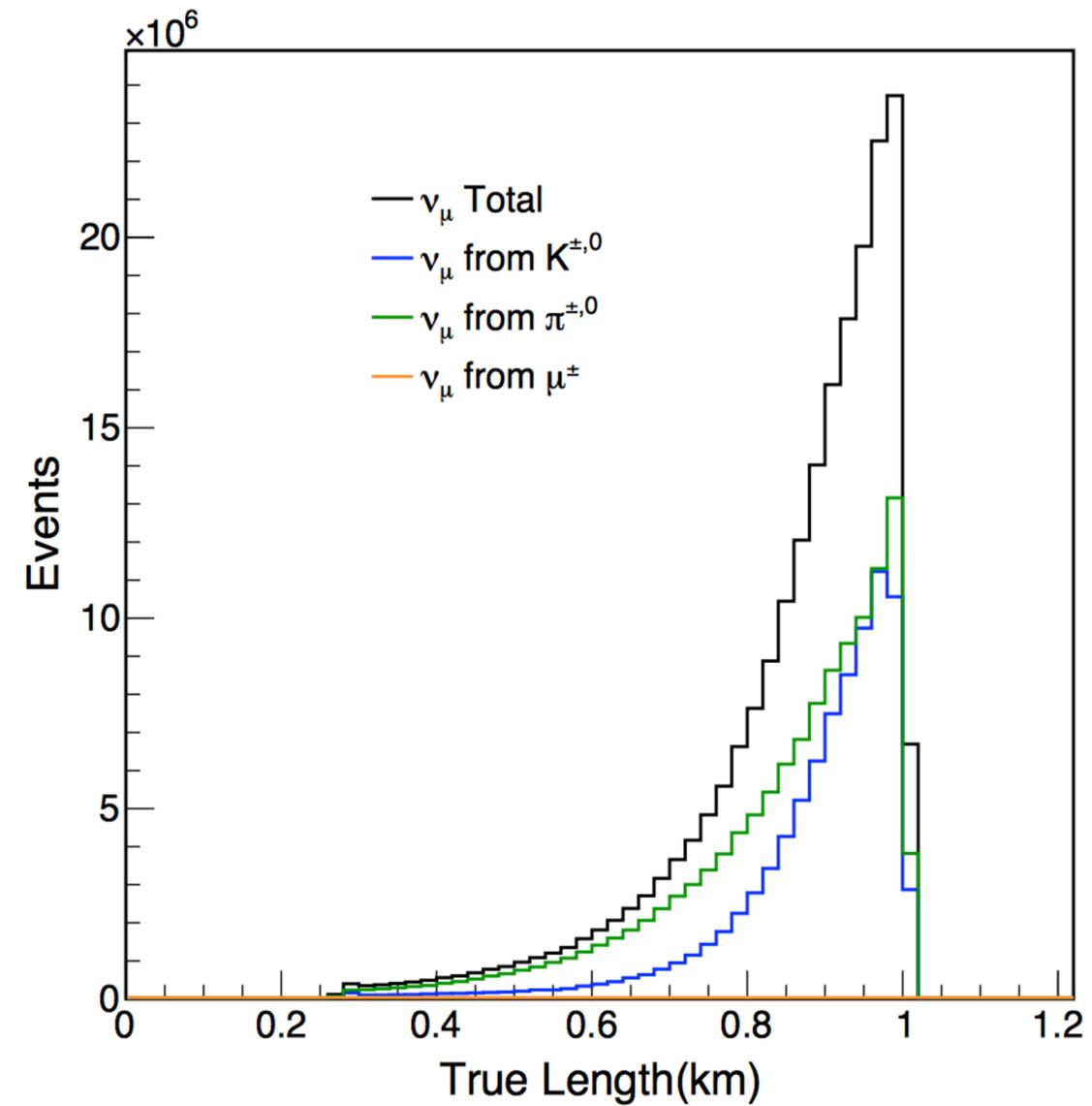
$$\sin^2 2\theta_{\mu\mu} = \cos^2 \theta_{14} \sin^2 \theta_{24}$$

²Sterile Neutrino Oscillations: The Global Picture, arXiv: 1303.3011

Other Sterile Neutrino Searches in NOvA

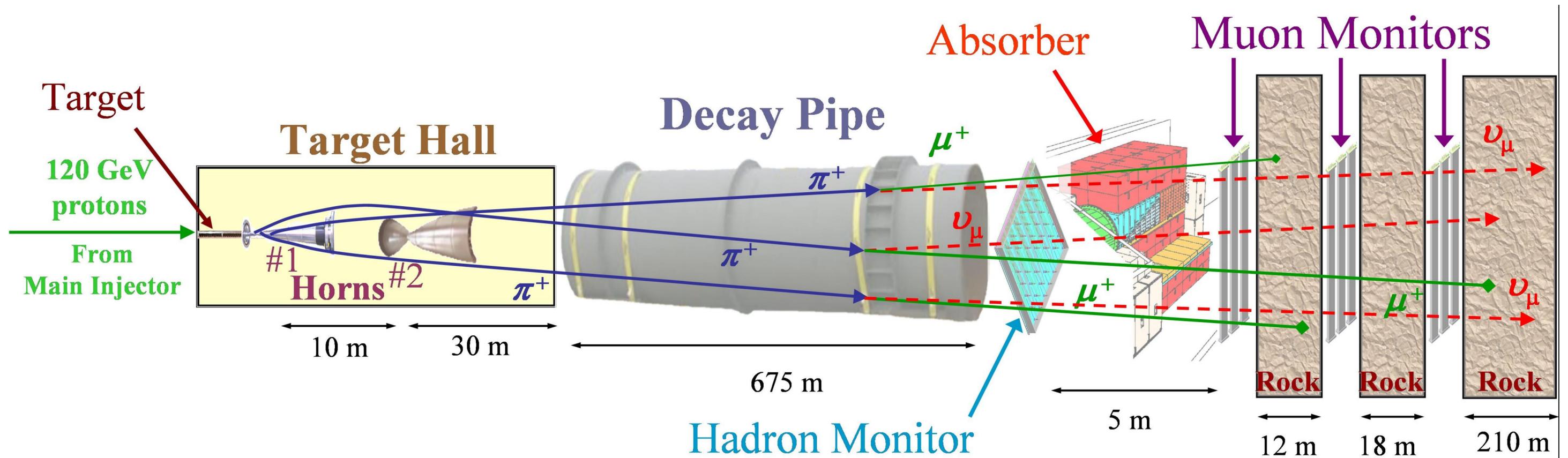
- **Short-Baseline ν_μ disappearance and ν_e appearance search**
 - ◆ Disappearance in muon neutrinos and appearance in electron neutrinos at the NOvA ND, due to oscillations with mass squared difference between the active and sterile neutrino, $\Delta m^2_{41} > 1\text{eV}^2$
- **Long-Baseline NC disappearance search**
 - ◆ Neutral-current interactions are not affected by three-flavor oscillations but mixing between active and sterile neutrinos will reduce the rate of NC events at the FD

True Length



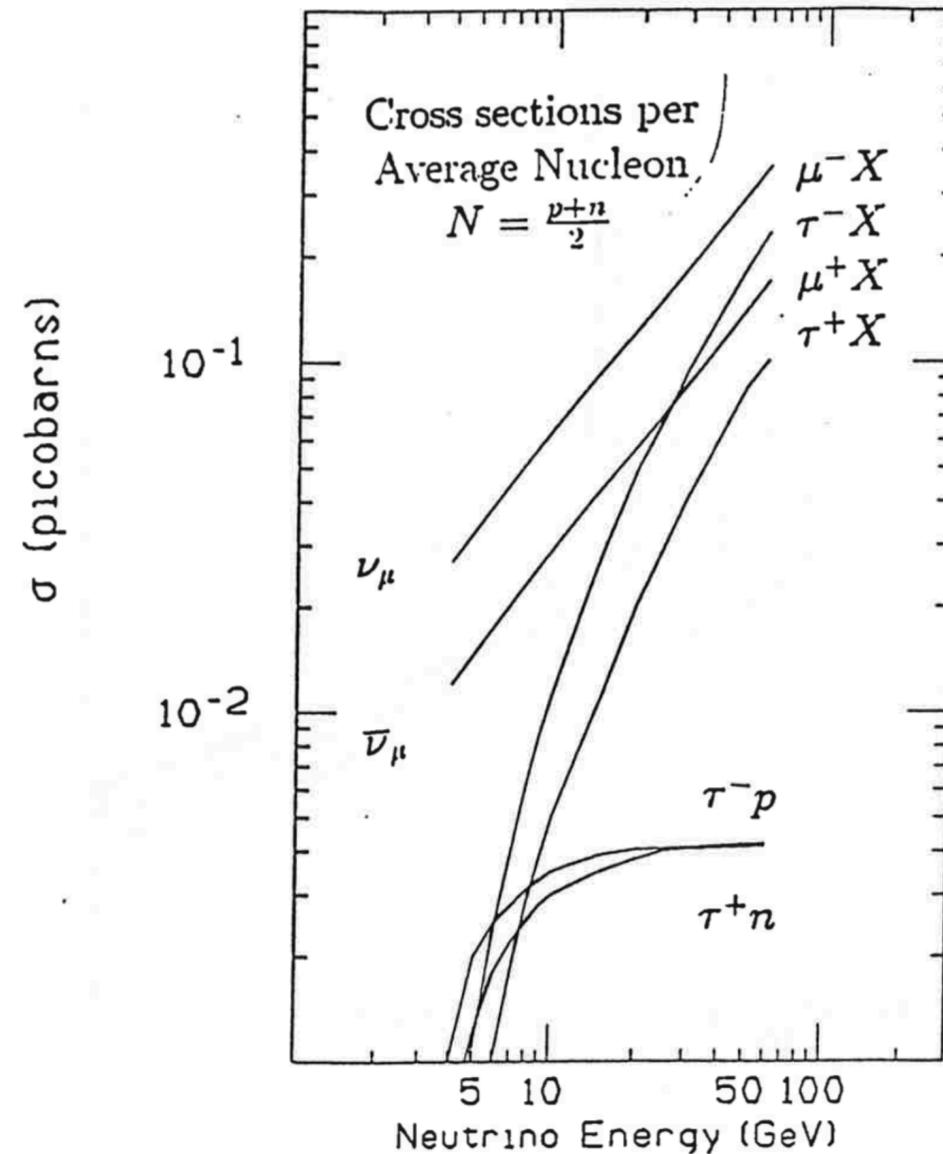
Distance travelled by neutrinos interacting in NOvA near detector

NuMI Beam



Charged Current Cross Sections

Charged current cross sections



Ref: Progress Report and Revised P822 Proposal to
Fermilab, 12 October 1993.
FERMILAB-PROPOSAL 0822
[http://lss.fnal.gov/archive/test-proposal/0000/
fermilab-proposal-0822.pdf](http://lss.fnal.gov/archive/test-proposal/0000/fermilab-proposal-0822.pdf)

Figure 10: Calculated ν_τ cross sections