



The WA105 collaboration



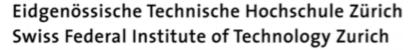
demonstrate the capabilities of the dual phase technology at the kton scale







































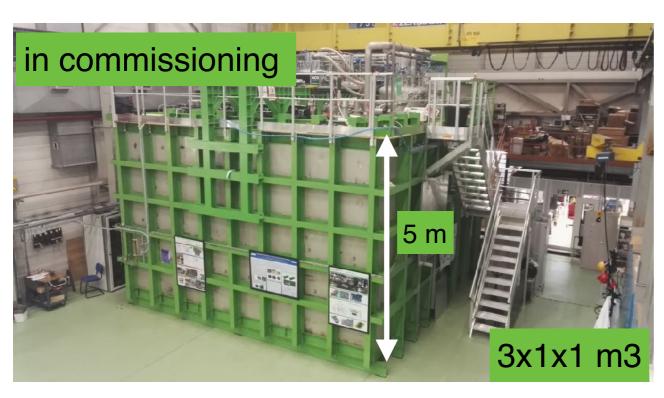








same technology, two scales, different goals





in construction

Common aspects

- ✓ LEMs and anode: design, purchase, cleaning and QA
- √ chimneys, FT and slow control sensors
- √ membrane tank technology
- ✓ Accessible cold front-end electronics and DAQ system
- √ amplification in pure Ar vapour on large areas

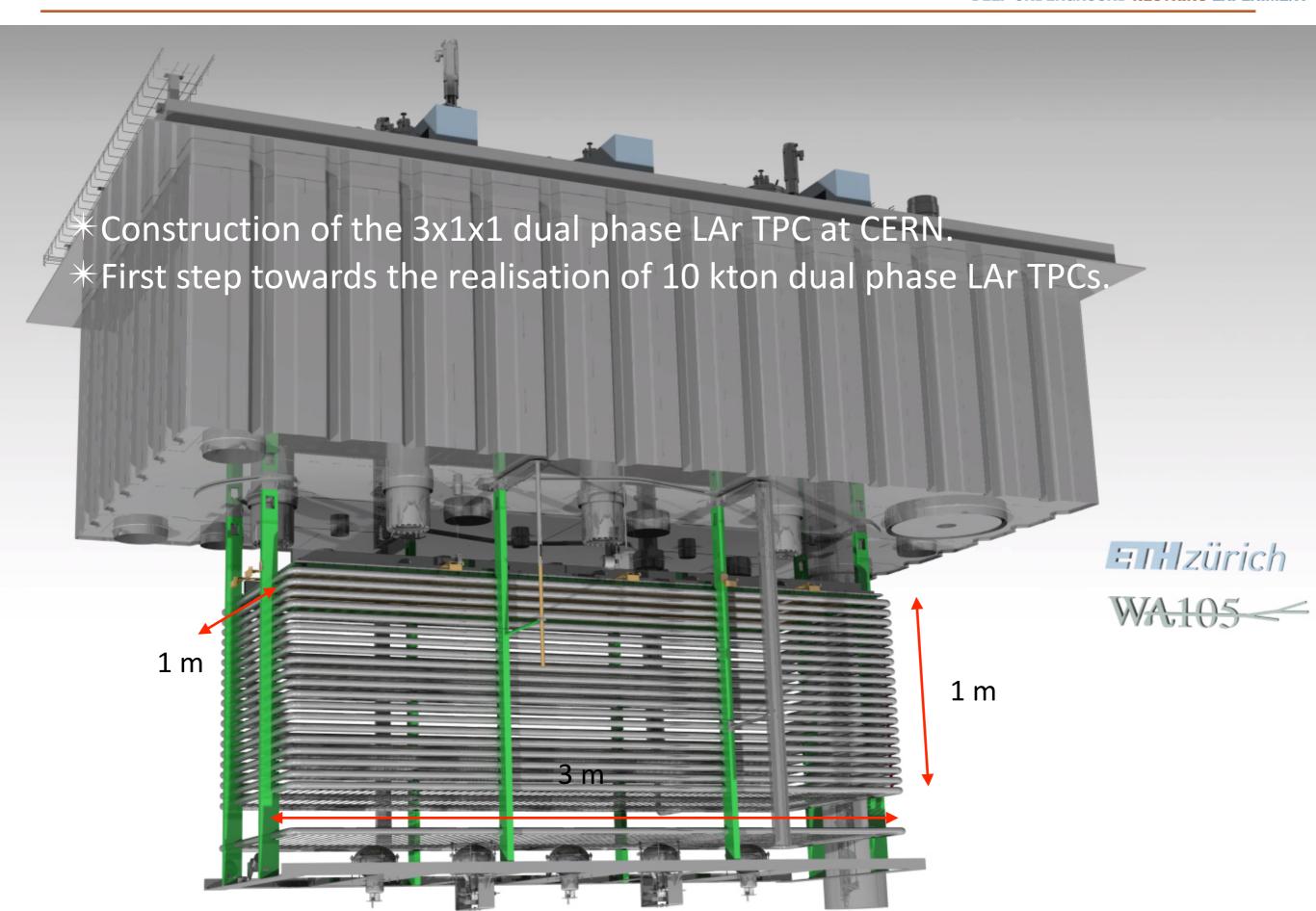
3x1x1 m3

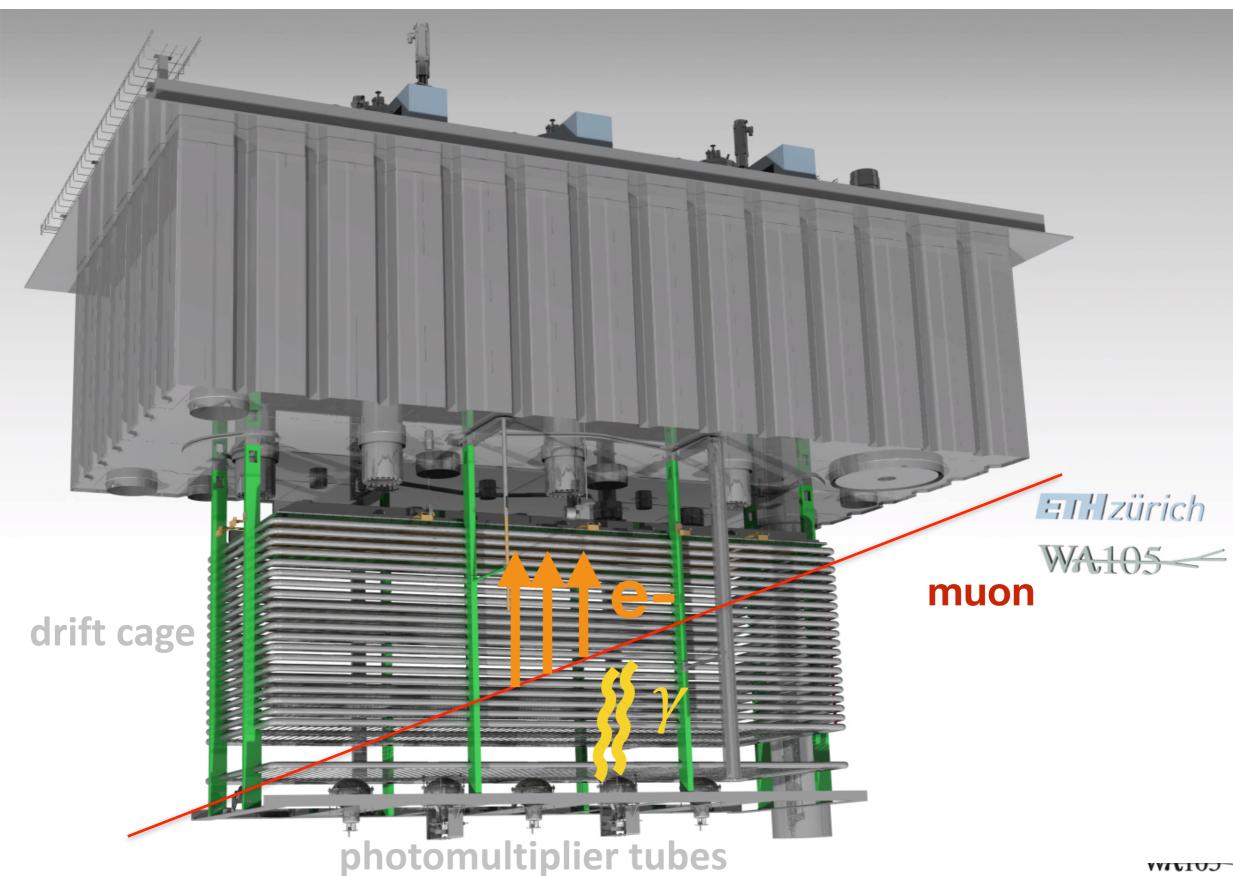
- √ First GTT constructed cryostat for LAr
- ✓ Fully engineered versions of many detector components with preproduction and direct implementation (installation details and ancillary services)
- ✓ First overview of the complete system integration: set up full chains for Quality Assessment, construction, installation and commissioning
- ✓ Anticipate legal and practical aspects related to procurement, costs and schedule verification
 - short term data taking with cosmics

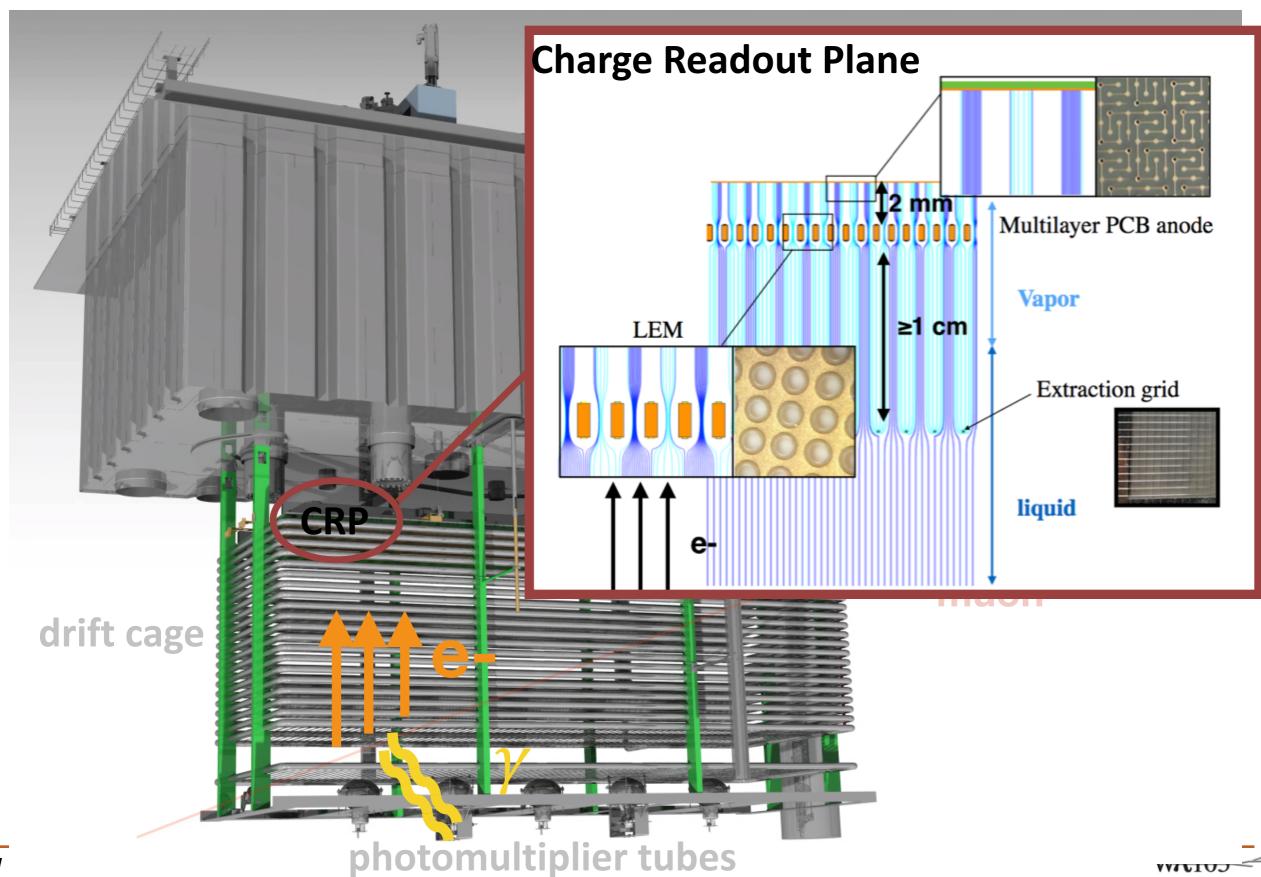
protoDUNE-DP

- ✓ Large hanging field cage structure
- ✓ Very high voltage generation and guiding
- √ Large area charge readouts
- √ long drift (e- diffusion, purity, etc..)
- √ test beam data (calibration, reconstruction, fully contained events, x-sections, etc...)
- ✓ Long term stability of UV scintillation light readout
- √ underground construction method



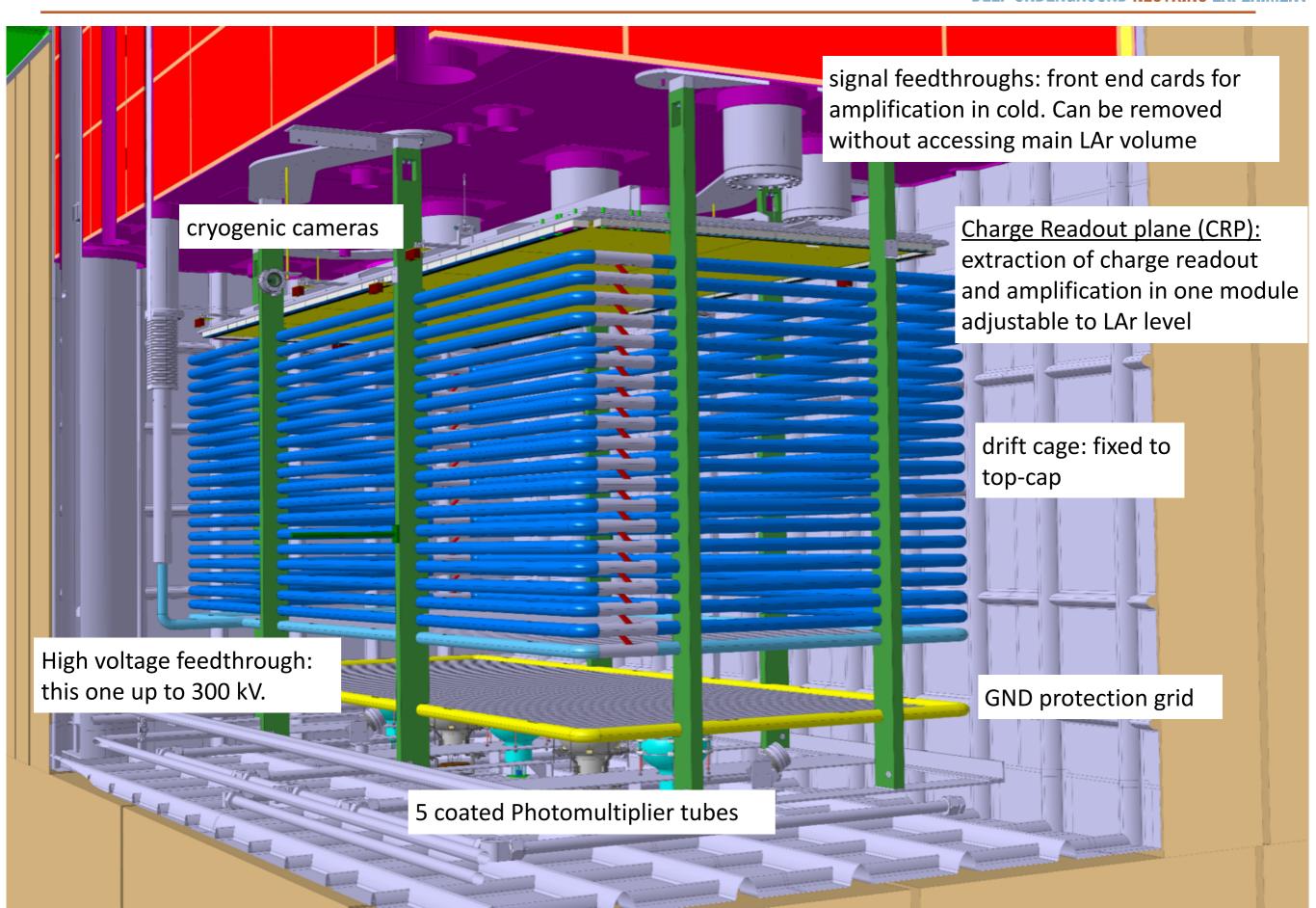






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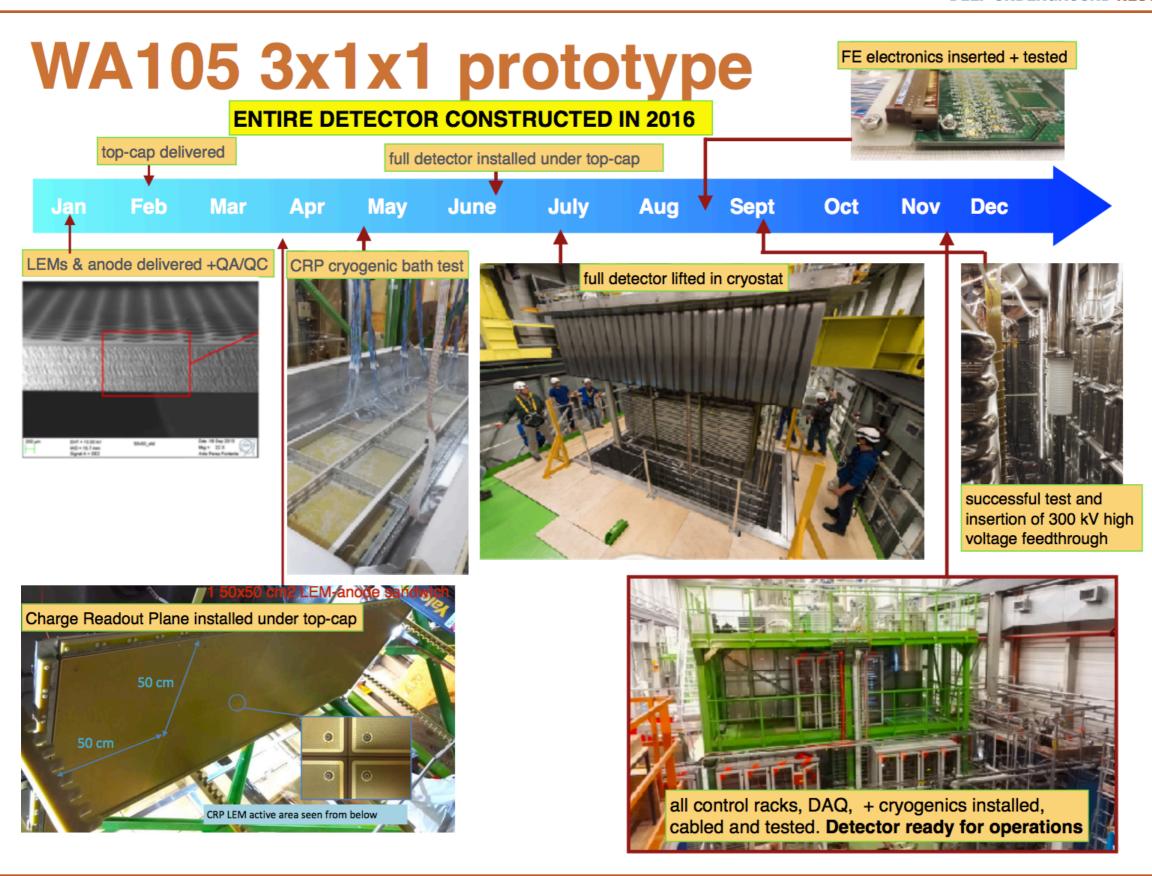






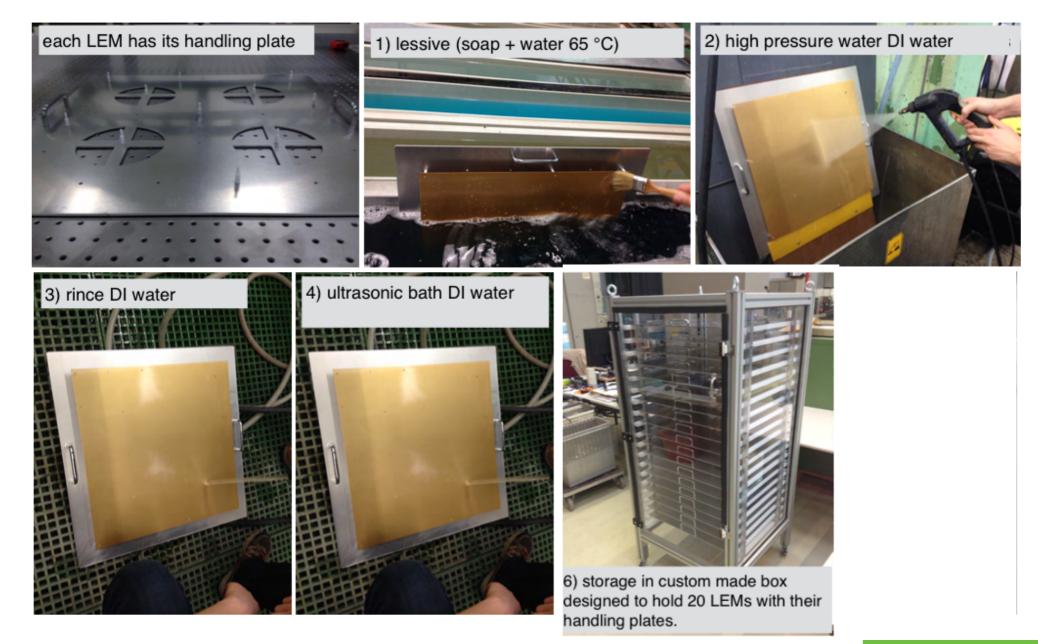


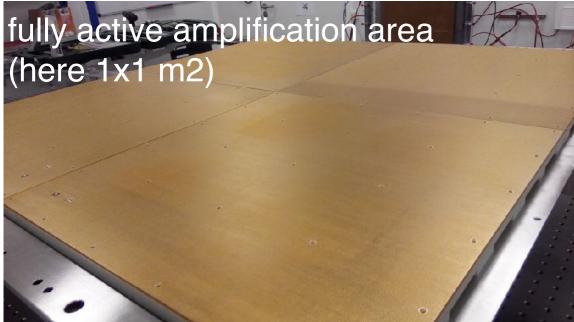
PART 1 DETECTOR INSTALLATION **MARCH-JULY 2016**

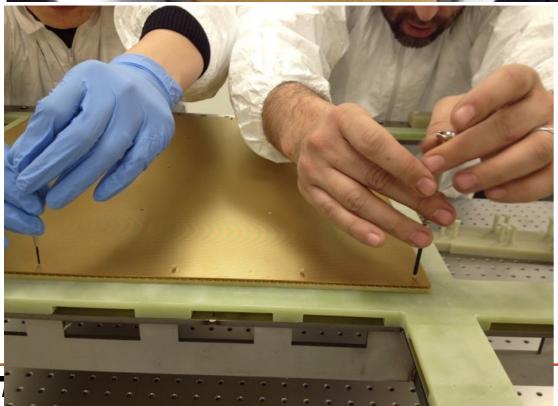


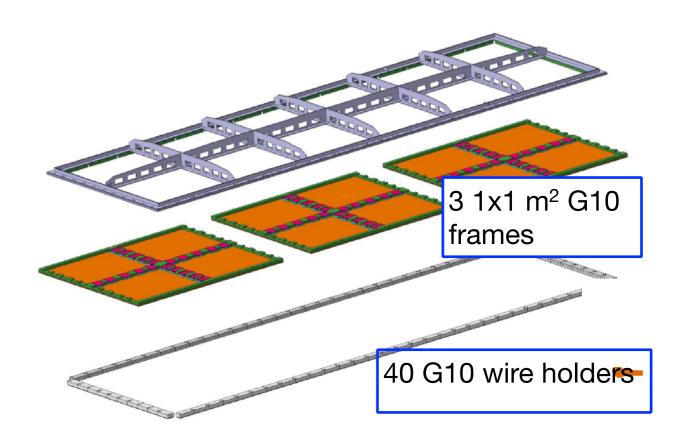
LEM cleaning

- Cleaning procedure is rather simple and straightforward. ~10 mins per LEM. Doesn't need specific facility.
- HV testing should be done in controlled environment (T,P and RH)
- no stringent requirements on storage, should be in a controlled environment that's about it.





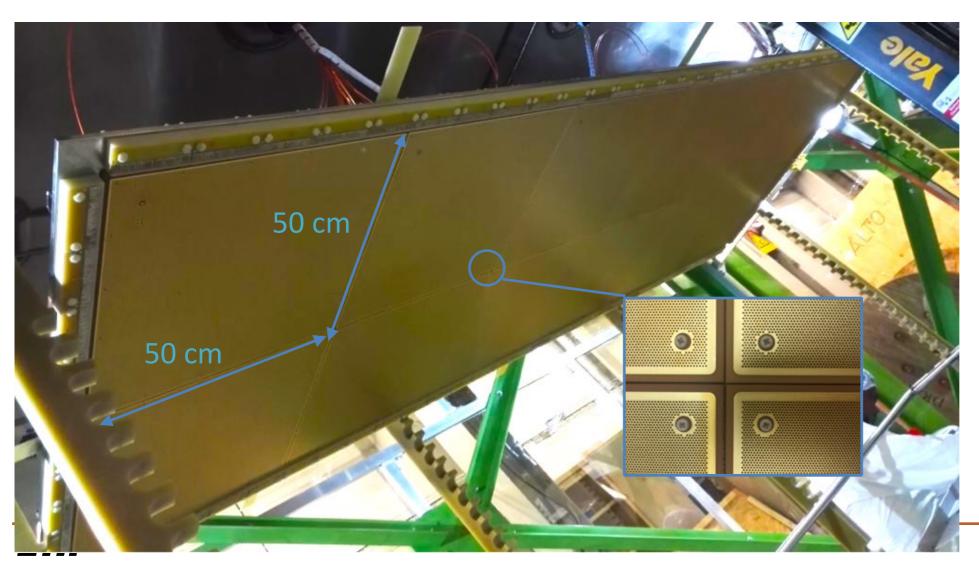


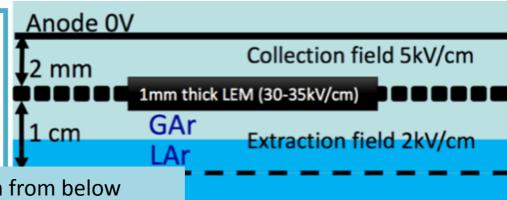




Charge Readout Plane

- * fully active 3x1 m2 amplification and readout adjustable to LAr level.
- * All components industrially fabricated with most of the QA/ QC performed by the companies.
- * mechanical tolerances validated in warm CRP LEM active area seen from below temperature in open cryogenic baths.
- * Assembly is straightforward and quick (~2 people, 2 days)

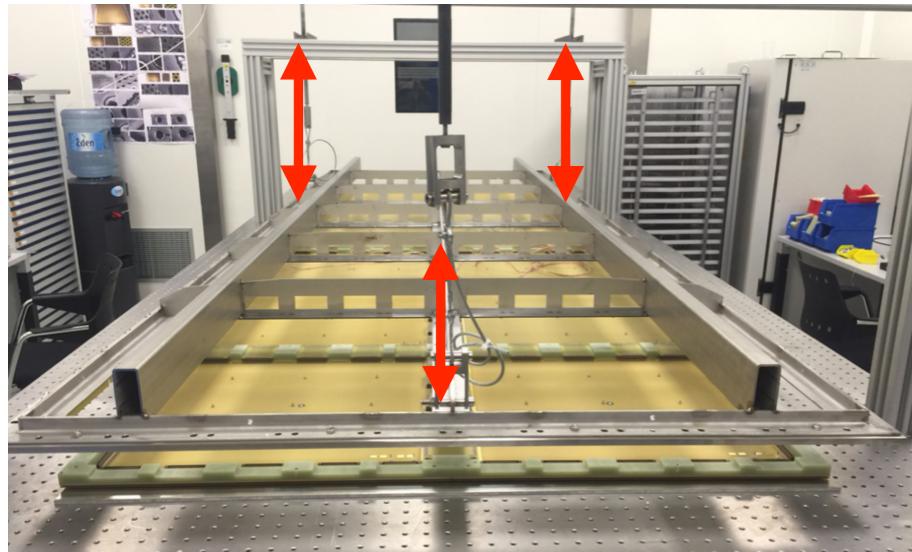






LEM + anode sandwich

extraction grid



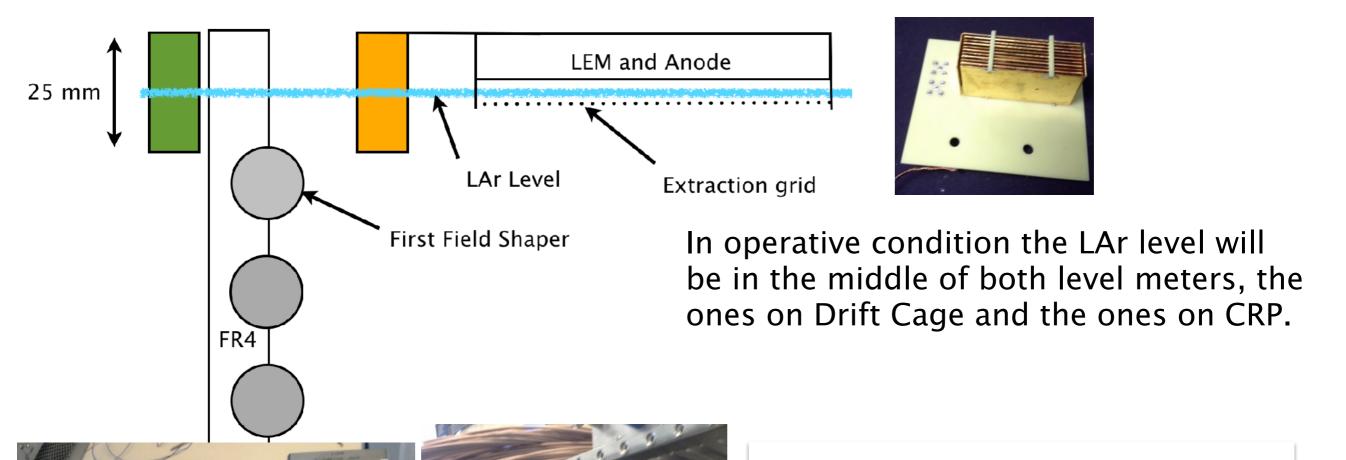






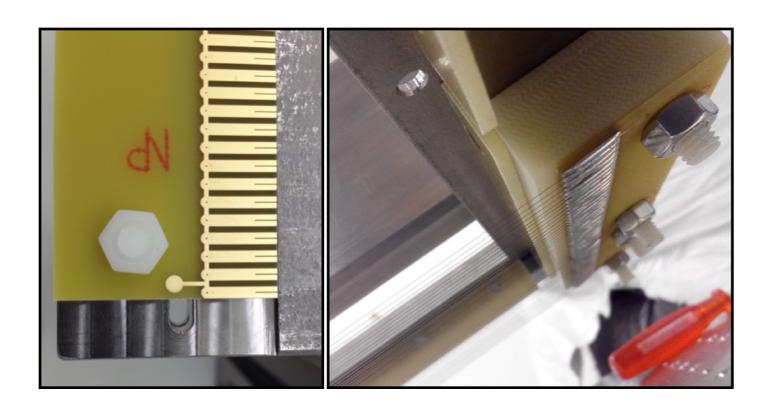
suspended by 3 ropes coupled to motors on top-cap. Precision of motors 100 um over 4 cm.
8 capacitive level meters readout the LAr level with similar precision





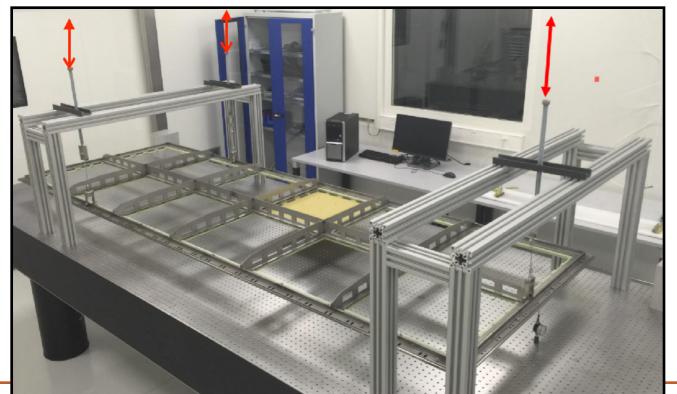
- A new NIM design has been done with:
 - 4 Channels
 - improved 0-10 V interface to NI Racks
 - improved filters on voltage rails and output
- 5 Boards NIM size are in production
- The assembly of the entire system will take place in the next 2 weeks, including calibration.
- Aso be a test bench for the 6x6x6 Level meter system





wire spacing.

wire: SS 100 micron diameter. spacing between each wire 3.125 mm. Each wire is precisely positioned on soldering PCBs in 220 um grooves. Precision obtained on wire spacing: <100 micron.



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planarity.

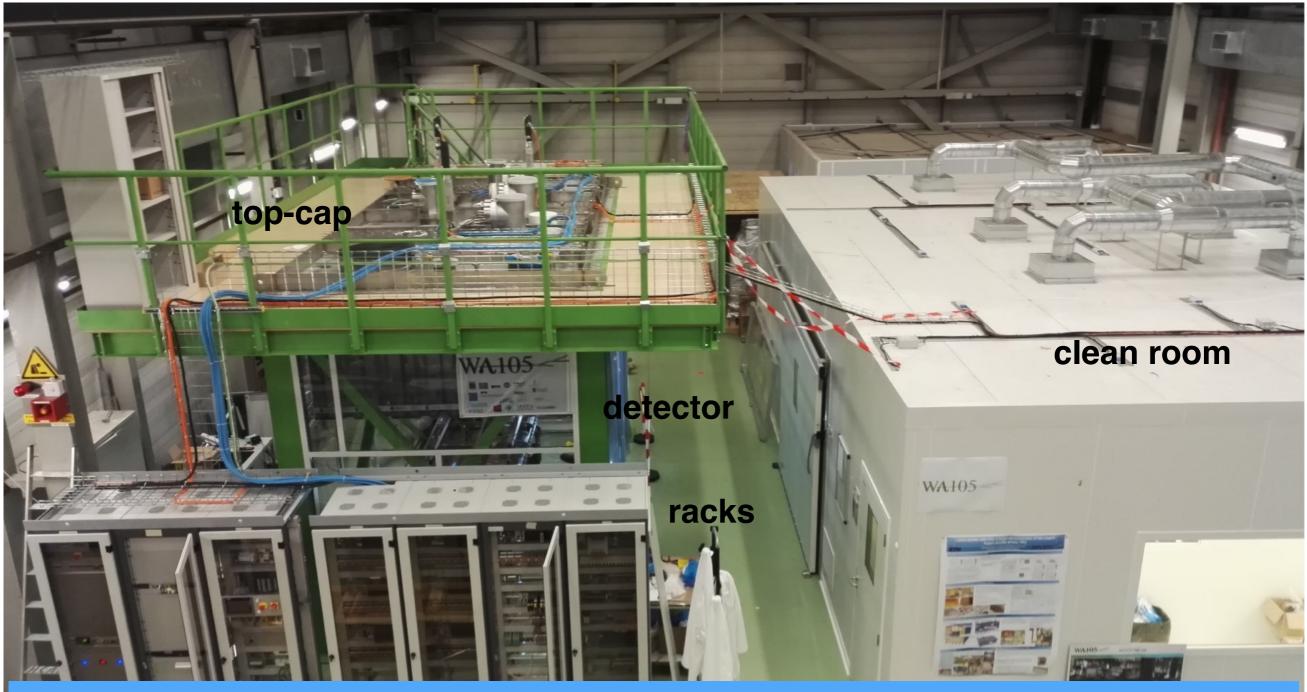
LEM-LAr 5 mm distance. Check that the mounted frame is within tolerance (planarity ±1 mm)

The horizontal geometry of the 3x§ m² CRP allows to perform **cryogenic test in open bath**. Already demonstrated for the 3x1x1.

What we checked:

- monitor expected shrinkage with photogrammetric measurements.
- extraction grid robustness.





The installation of the detector under the top-cap essentially started in May. By end of June it was complete and most parts were tested.

Thanks to the large collaboration effort during those 2 months

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WA105









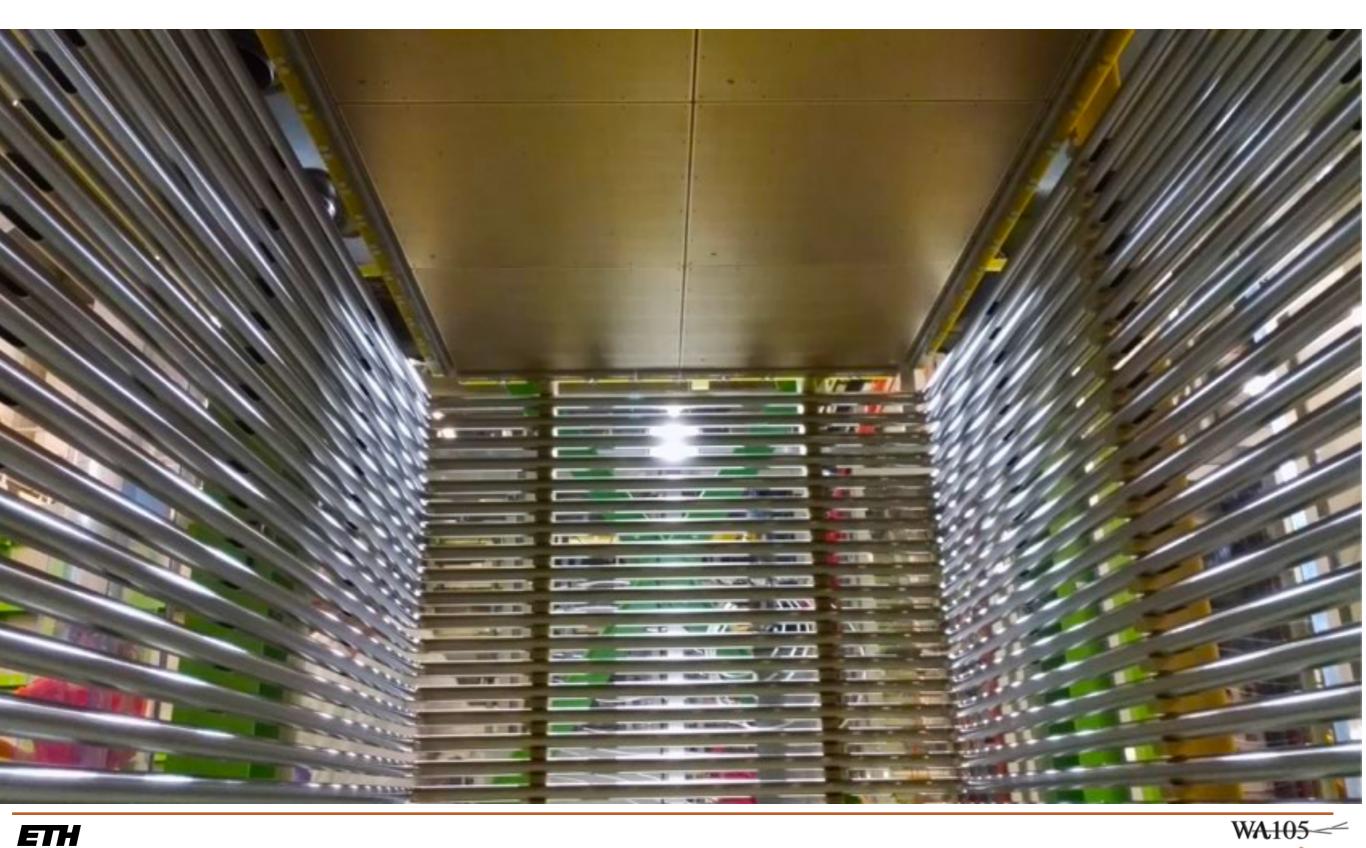






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Feedthroughs

All feedthroughs operational and tested over the past year. Same to be installed in pDUNE-DP

4 signal chimneys

3 slow control and medium voltage

3 CRP suspension

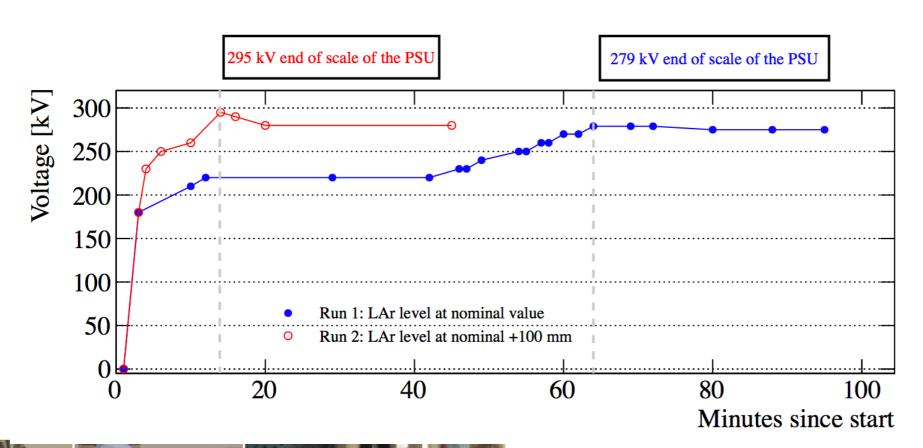
1 High voltage tested at 300 kV, operated <u>a</u>t ~50 kV



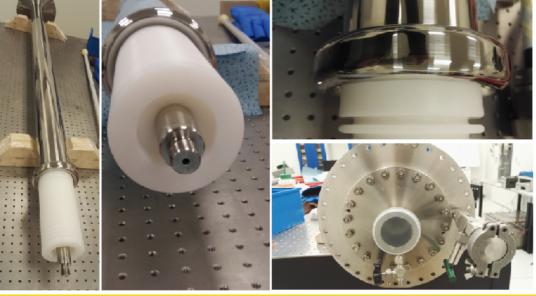












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Design successfully tested in dedicated setup up to the end of the scale of the Heinzinger PSU. <u>About 295 kV.</u>

JINST 12 P03021 arXiv:1611.02085



The insertion of the detector was fast about 2 hours and no problems were encountered. This was thanks to CERN support and professionalism.

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Inspection and measurements inside cryostat

DEEP UNDERGROUND NEUTRINO EXPERIMENT







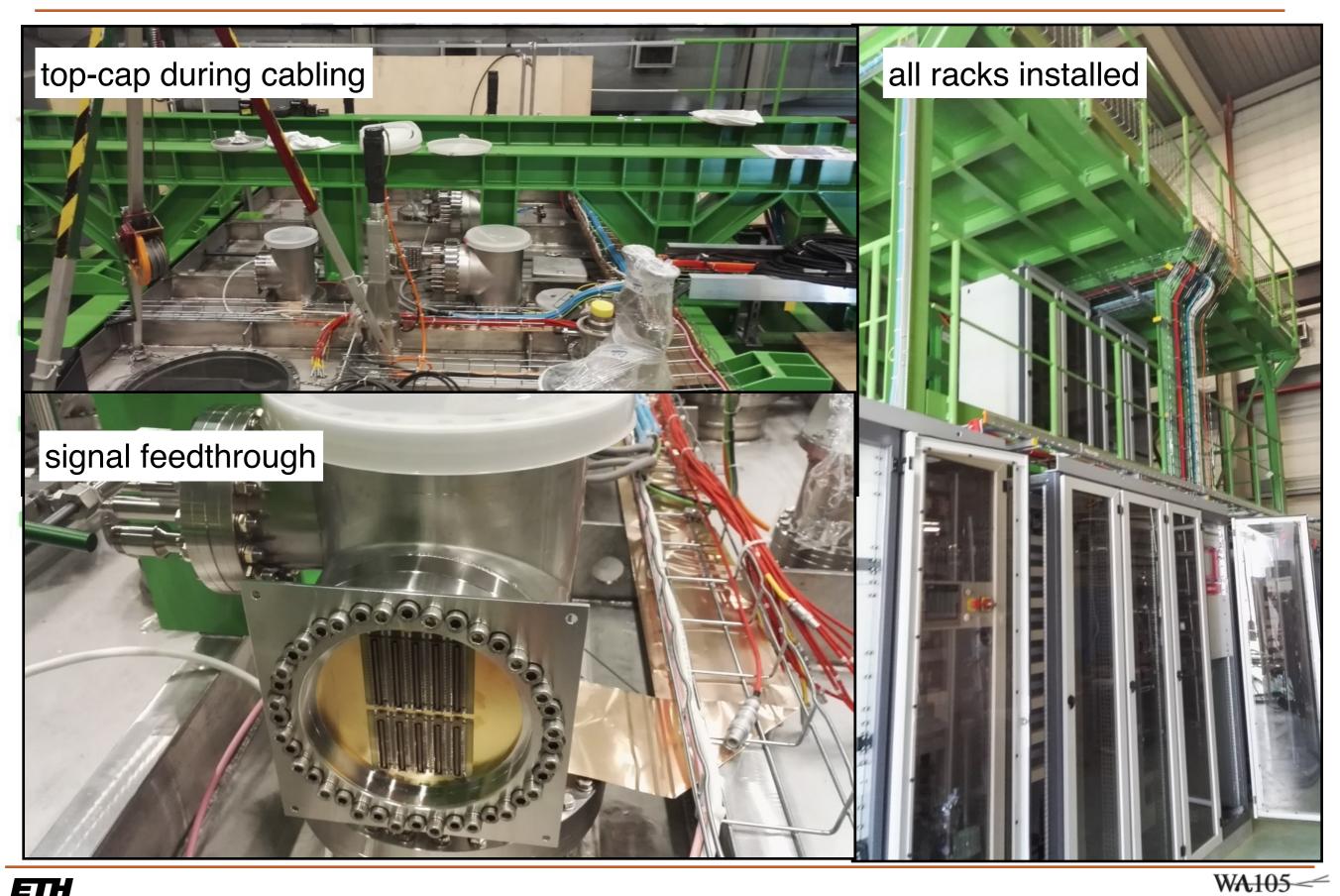


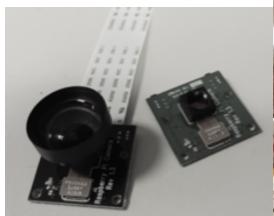


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PART 2 CRYOGENICS & COMMISSIONING AUG 2016 - APRIL 2017

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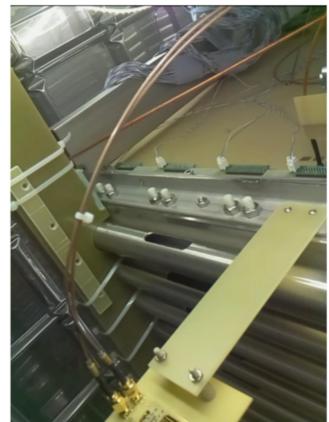
cameras tested in LAr for many days without degradation of image.



5 cameras placed in different strategic areas of the detector. Main purpose to have visual check of the LAr level.



Wa105cam0: - On top 7/7/2018HV feedhrougth



Wa105cam1:

- On top
- Ar level

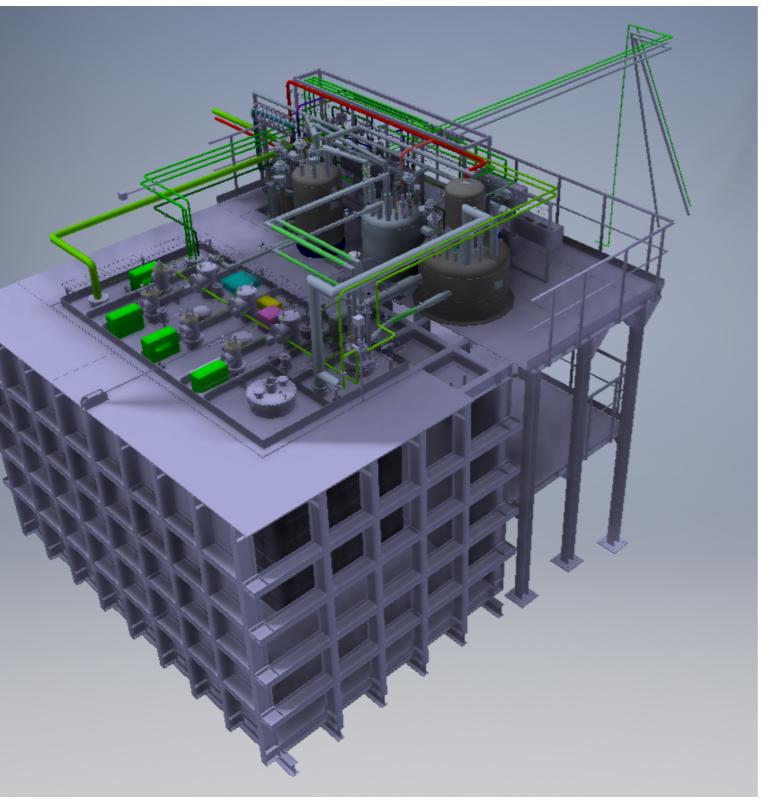


Wa105cam2:

- On top
- Ar level



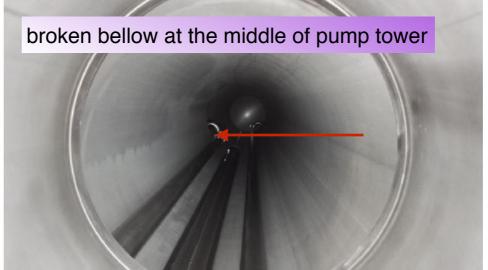
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- Cold piping (LAr+ LN2 lines, valve boxes, liquid purification,...) Sept 19th- Oct 13th
- Warm piping (gas argon purification) system, chimney purges, ..) Oct-Nov
- Control system Sept-Nov
- Start of gas argon piston purge Jan 24th
- Start of cool down Feb 27th



















Piston purge
4 warm gas lines each with 3 openings of 12 mm ø.
total flow rate during piston purge ~4 l/s

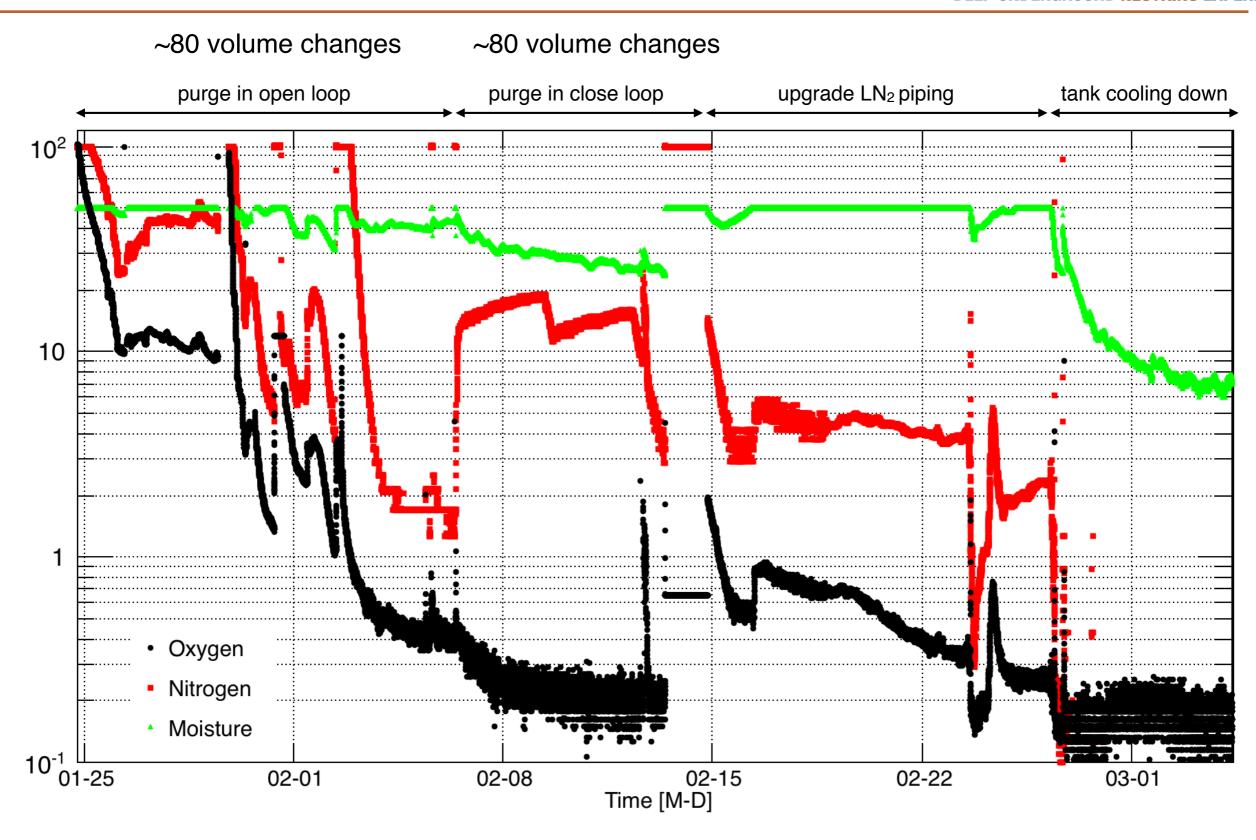
Cool down: 4 sprays mixture of LAr and GAr for slow and uniform cool down. Nominal flows: 300 K GAr 500 I/m 87 K LAr 21 I/h



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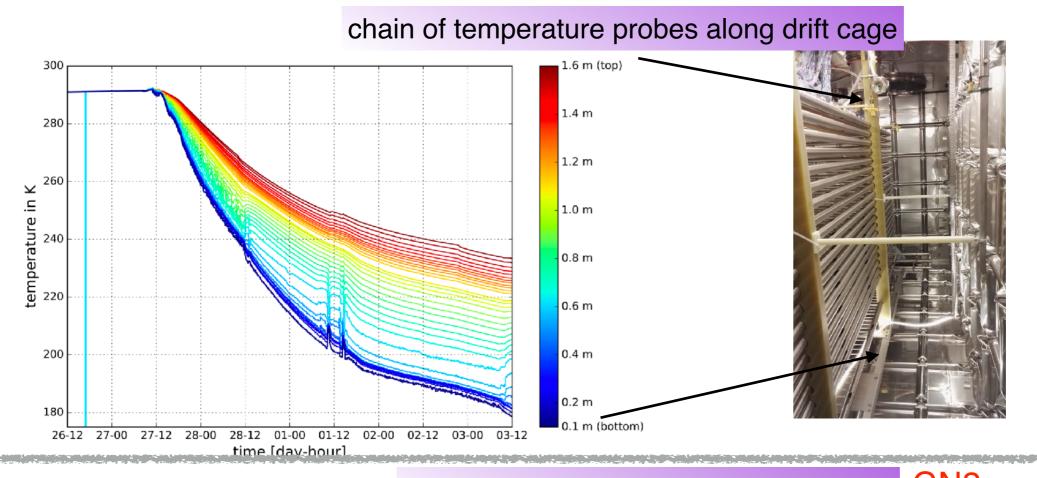
Piston purge & evolution of impurities

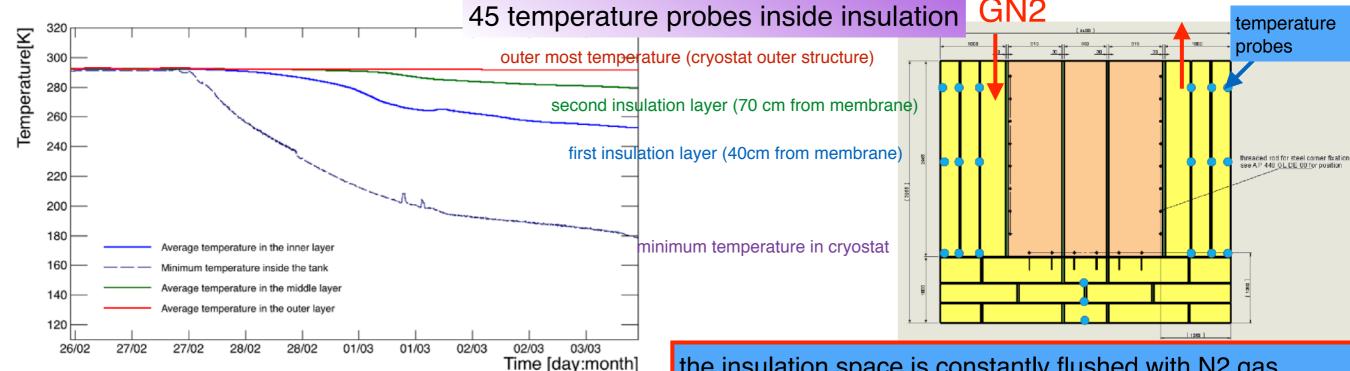


Impurity (ppm)

All sensors operational - some data during cool down

DEEP UNDERGROUND **NEUTRINO** EXPERIMEN



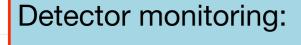


ETH

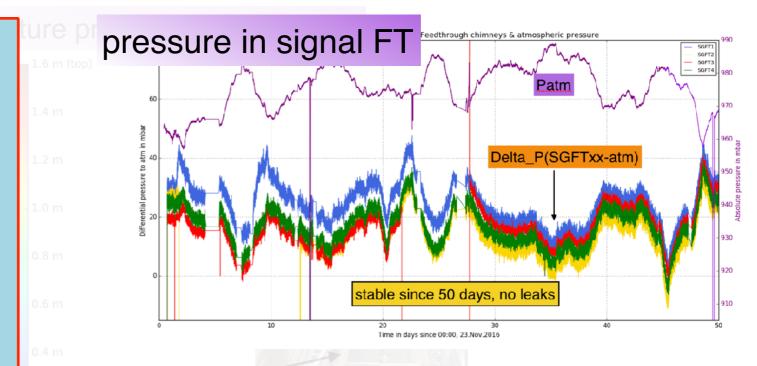
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the insulation space is constantly flushed with N2 gas.

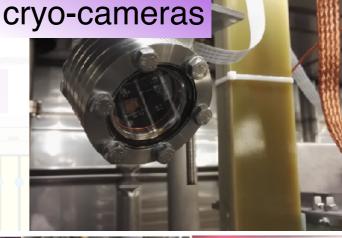
A bubbler at the output garanties a few mbar overpressure w/respect to the atmosphere

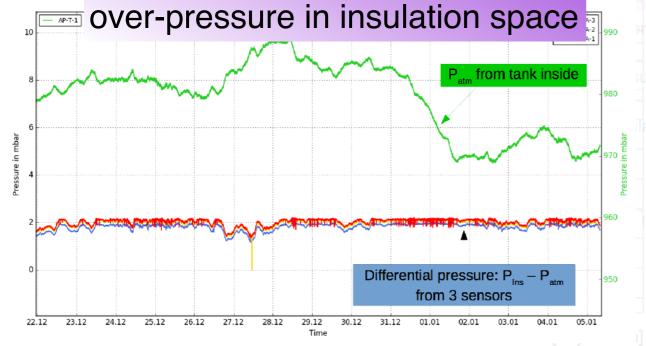


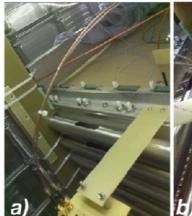
- >150 temperature probes
- 20 pressure probes
- 30 HV channels
- 1 300 kV HV channel
- Purity monitors (Gas + liquid)
- 15 level meters
- 5 cryogenic cameras









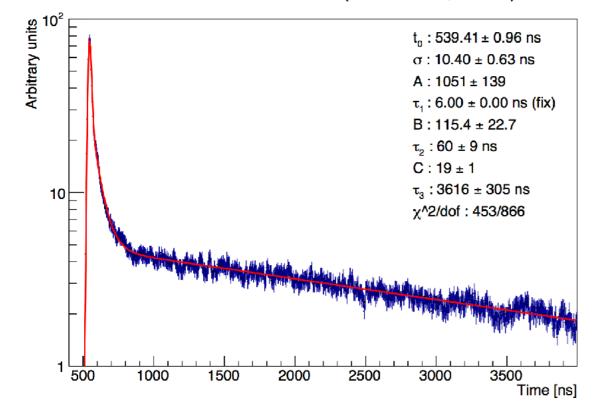








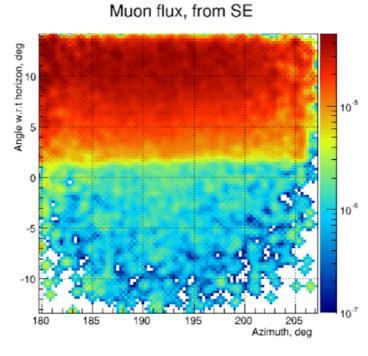
Scintillation time in GAr (1000 mBar, 215 K)



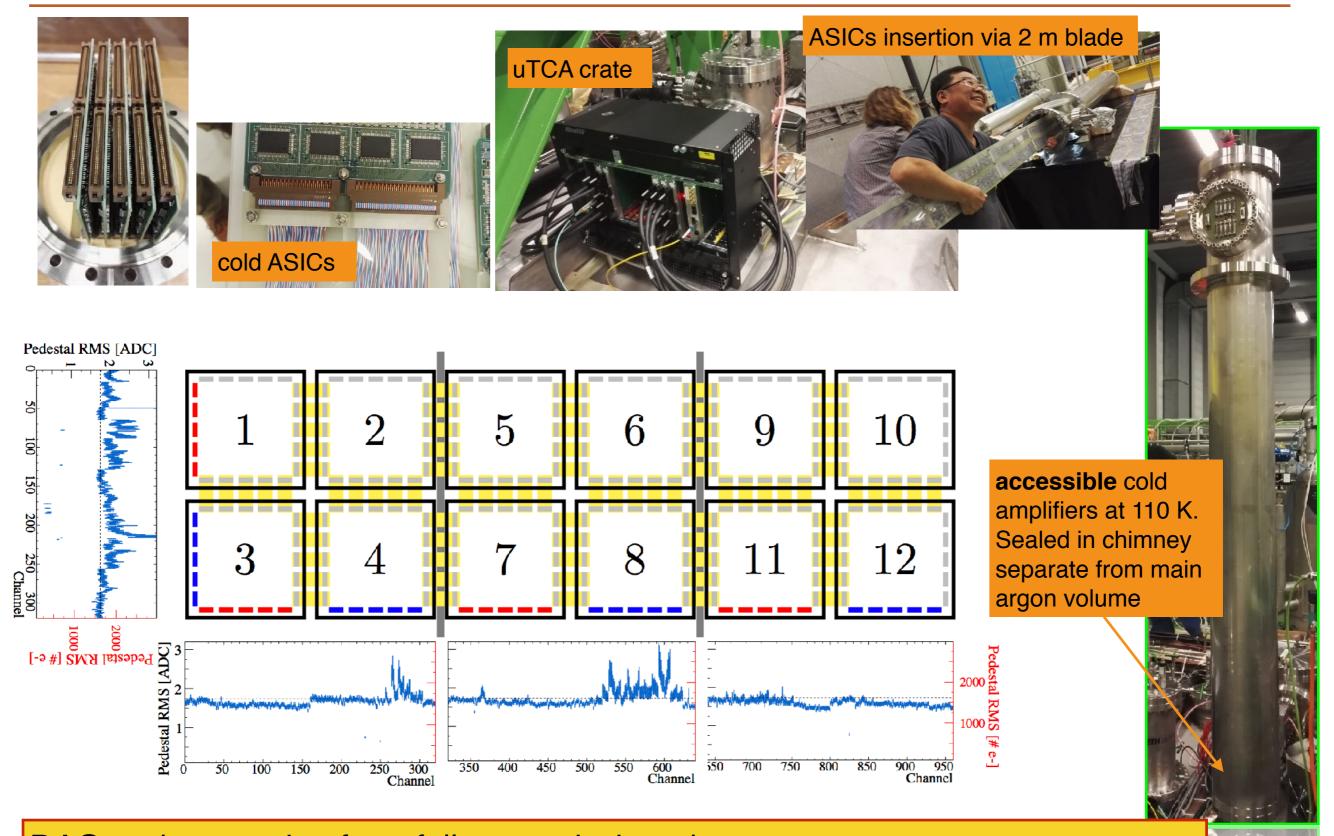
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Muon flux, from NW Azimuth, deg



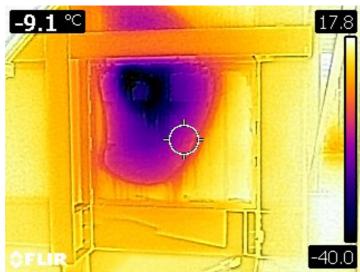




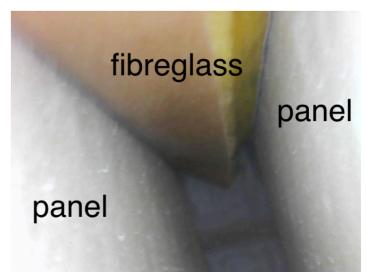
DAQ and computing farm fully commissioned noise at room temperature stable at about 1'500~ electrons

Cryostat- issues during cool down









- Since Feb. 28th two attempts at cooling down the cryostat have been made. Both have been interrupted due to the presence of multiple cold spots (some <-40°C) on the outer structure.
- The exact source of the cold spots as well as the solution to fix the issue is currently under investigation by CERN and GTT. So far the following has been understood:
 - A cold leak from the inner membrane has been excluded.
 - In some locations the outer structure was drilled and missing layers of fibreglass sheets were found. Those holes have been filled with expandable foam which cures the problem locally.
 - Next week a GTT team is coming to CERN and should be in position to solve the general issue based on results from a full simulation of the gas convection arising from the free spaces between the insulation blocks.



- Most of what has been shown in the slides has been designed to match the scale of protoDUNE. From that perspective, the installation and commissioning steps of the 3x1x1 has provided many valuable inputs
- The complete assembly of the 3x1x1 detector including cabling and DAQ commissioning took about 6 months.
- Some delay on the operation has been accumulated during cryogenic installation and commissioning phase and more recently due to the defect in the insulation which resulted in an abort of the cool down. The reason seems now understood and repaired. We are currently filling the cryostat.
- Although cosmic tracks have not yet been acquired, large experience has been gained for protoDUNE-DP design, installation and commissioning.

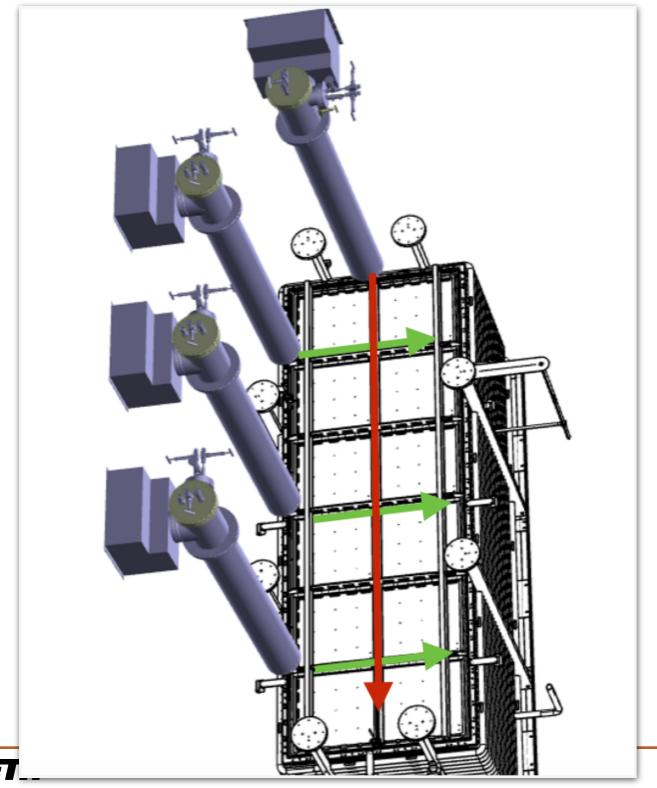
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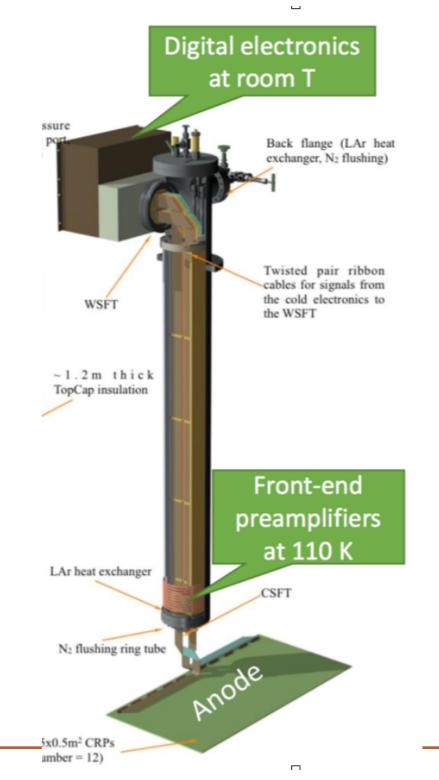
THANK YOU

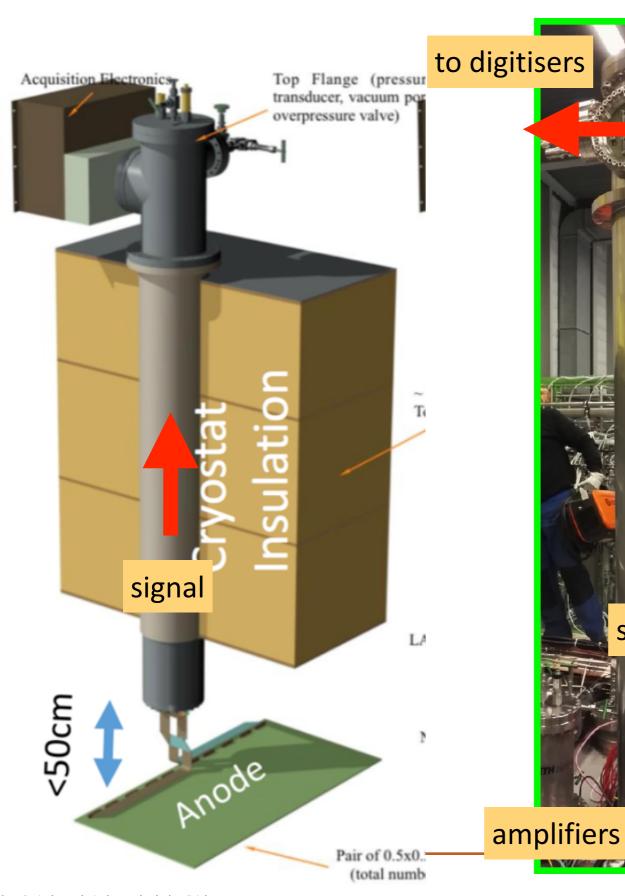


Signal feedthrough

2 view readout 3 chimneys for the 1 m strips. 1 chimney for the 3m long view.



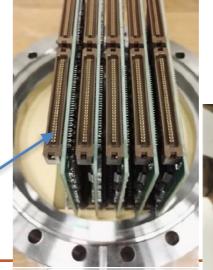




amplifiers <u>accessible during</u> <u>operations</u>



amplifiers inside closed volume. Close to anodes, ~110 K



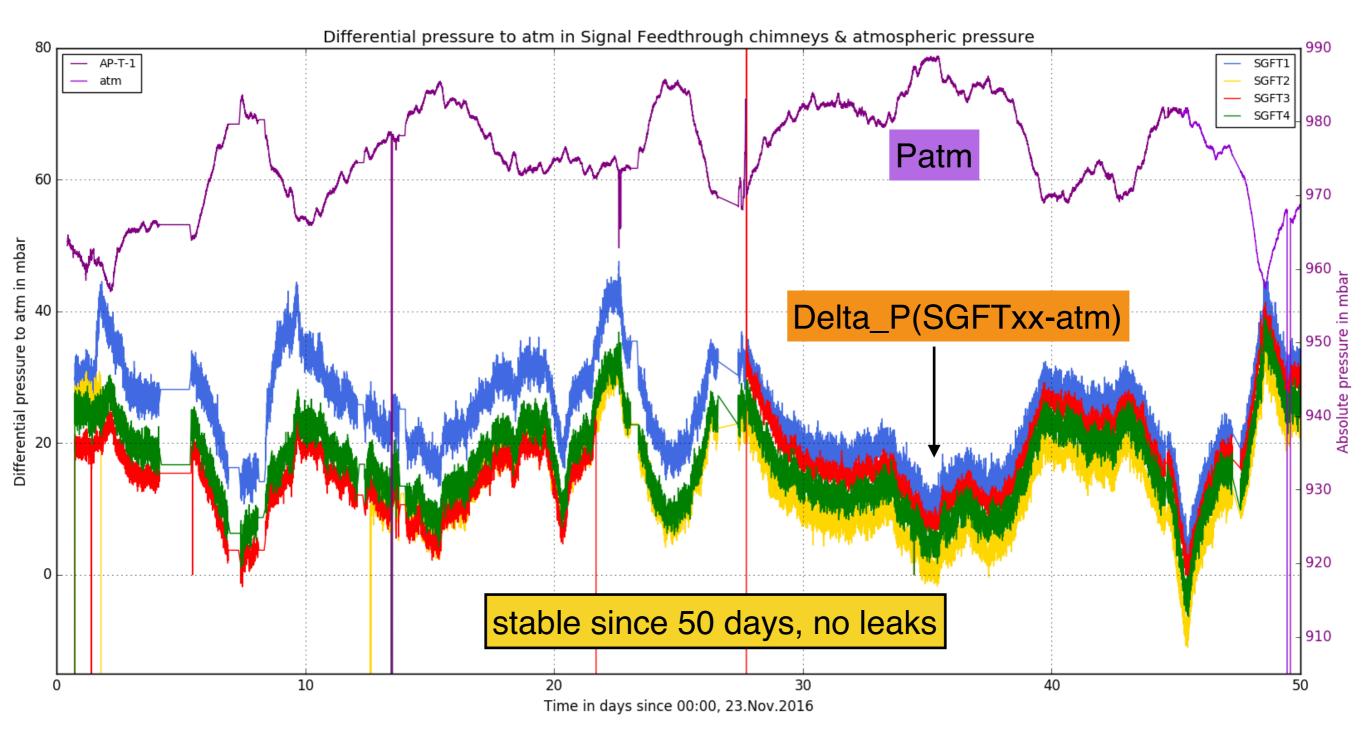
4 ASICs per board



signal

Signal chimneys

Independent volumes (to allow replacement of preamps). Pumped to remove air and sealed under GN2 with ~20 mbar overpressure W/ respect to Patm.





1. charge signal multiplied and collected on low capacitance anode strips



2. signal guided to cold amplifiers by group of 32 channels

inside detector (not accessible)

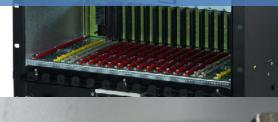
3. signal amplified by ASICS in cold

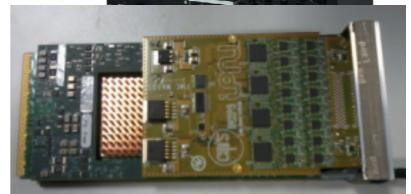


4. signal brought outside by vacuum tight custom designs PCB flanges



5. signal digitised by 12it in AMC arms in uTCA crates



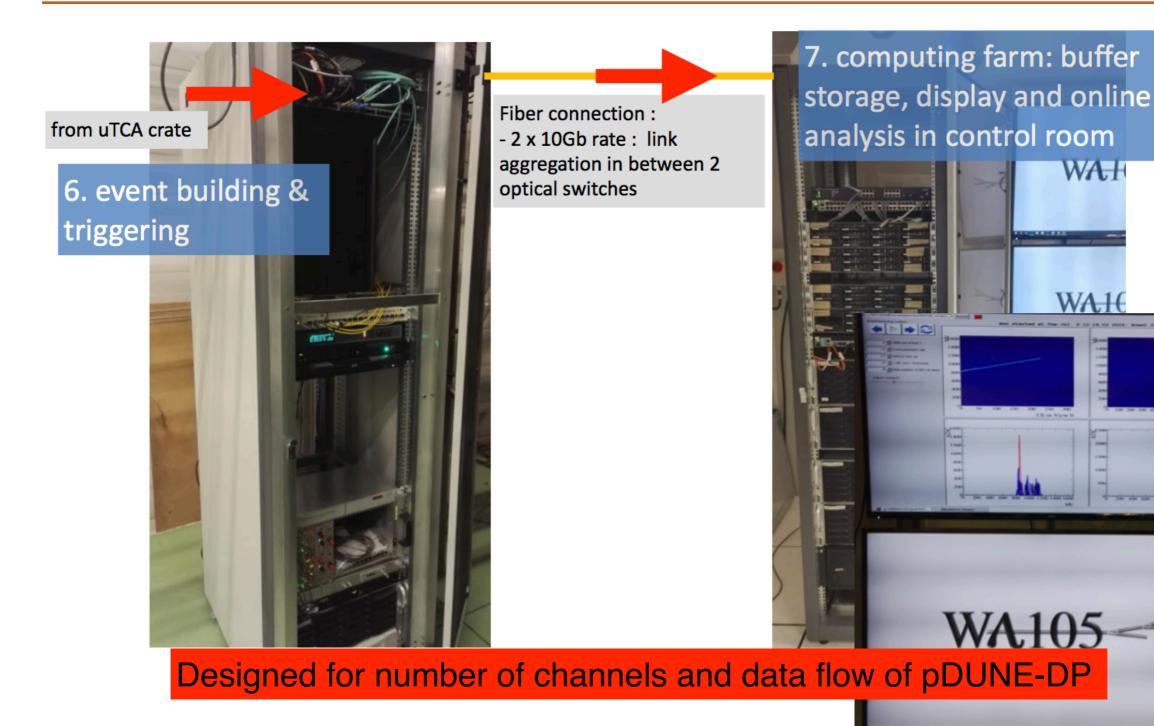




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signal feedthrough (accessible)

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D. Autierio (IPNL) pDUNE-DP plenary

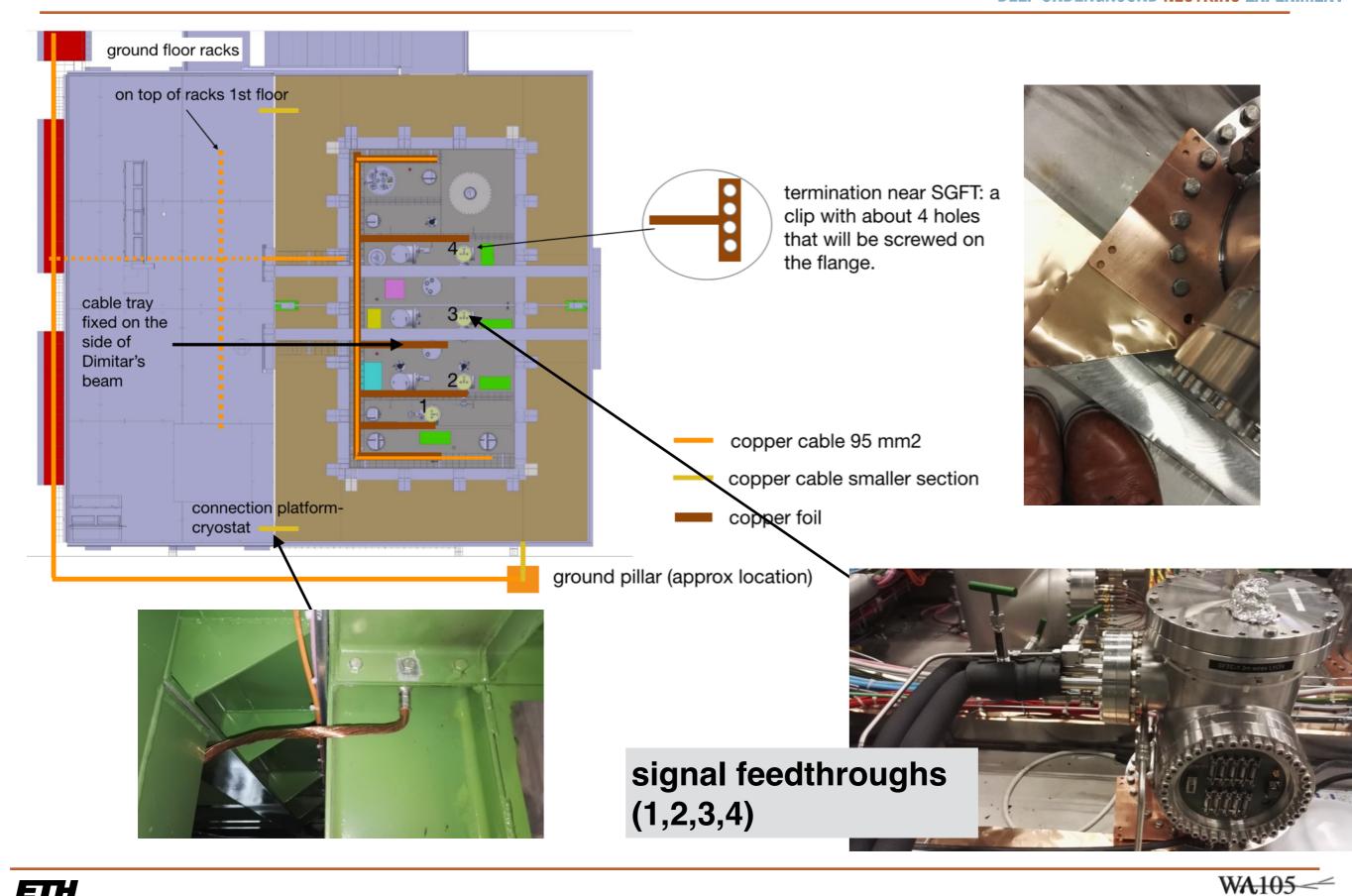


8. Send via 10 Gb CERN Network to computing center (EOS, etc..)



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Grounding



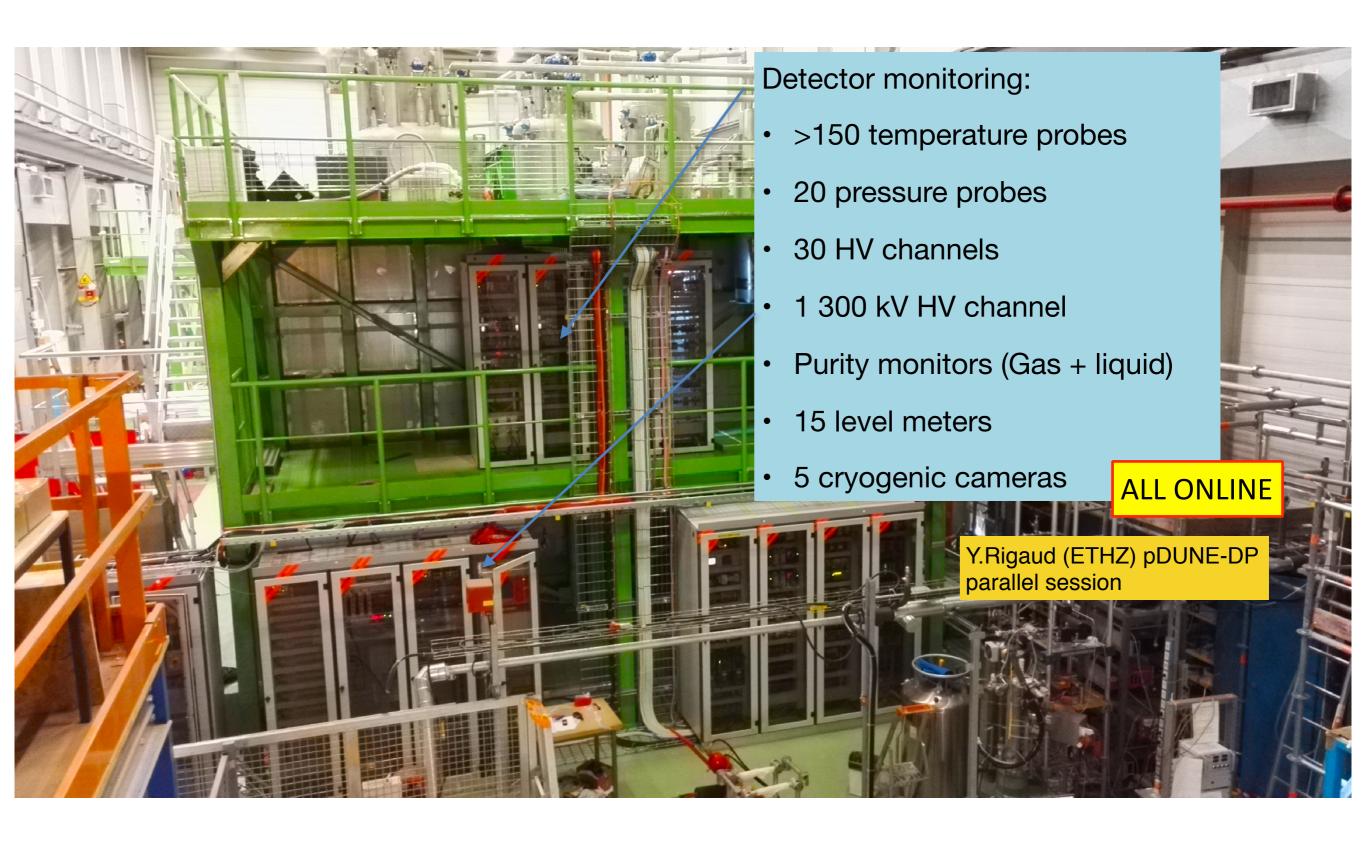
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Moved from "Construction" to "Operations" mode (conveners Laura M.B, Laura M.)

- Triggering scheme, cosmic ray trackers operational
- Cryogenic warm piping finished. Cryostat ready for purge.
- DAQ + pulsing system operational.
- · First data, cross check of all channels (continuity, noise and cross
 - talk) with DAQ
- All sensors ready and monitored
- Sealing of Manhole
- Start of GAr piston purge



Ready for detector operation

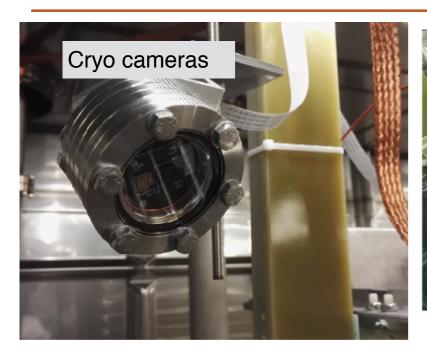


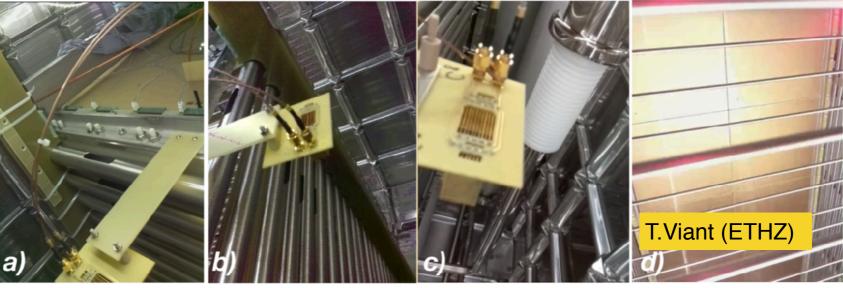
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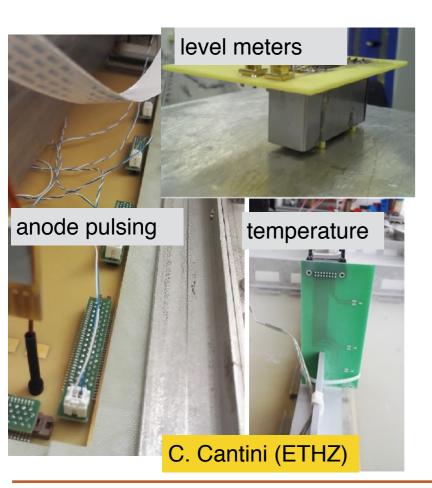
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Monitoring





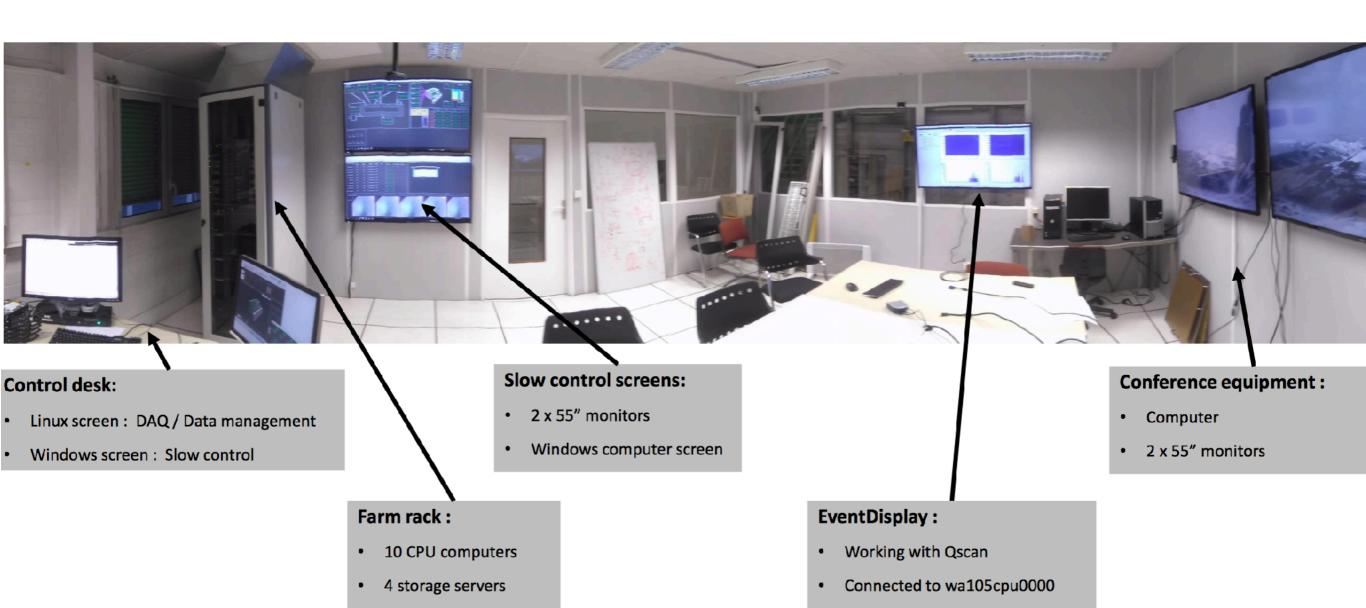
ALL RUNNING









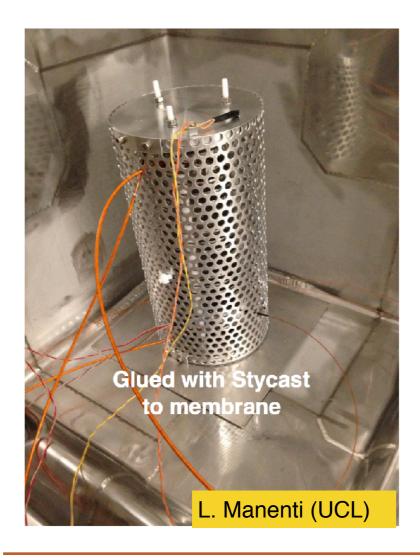




Managed from the control desk

Switches

Purity monitor Installed

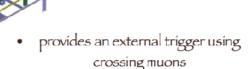






- 4 scintillating modules
- 2 modules at each short side of the cryostat





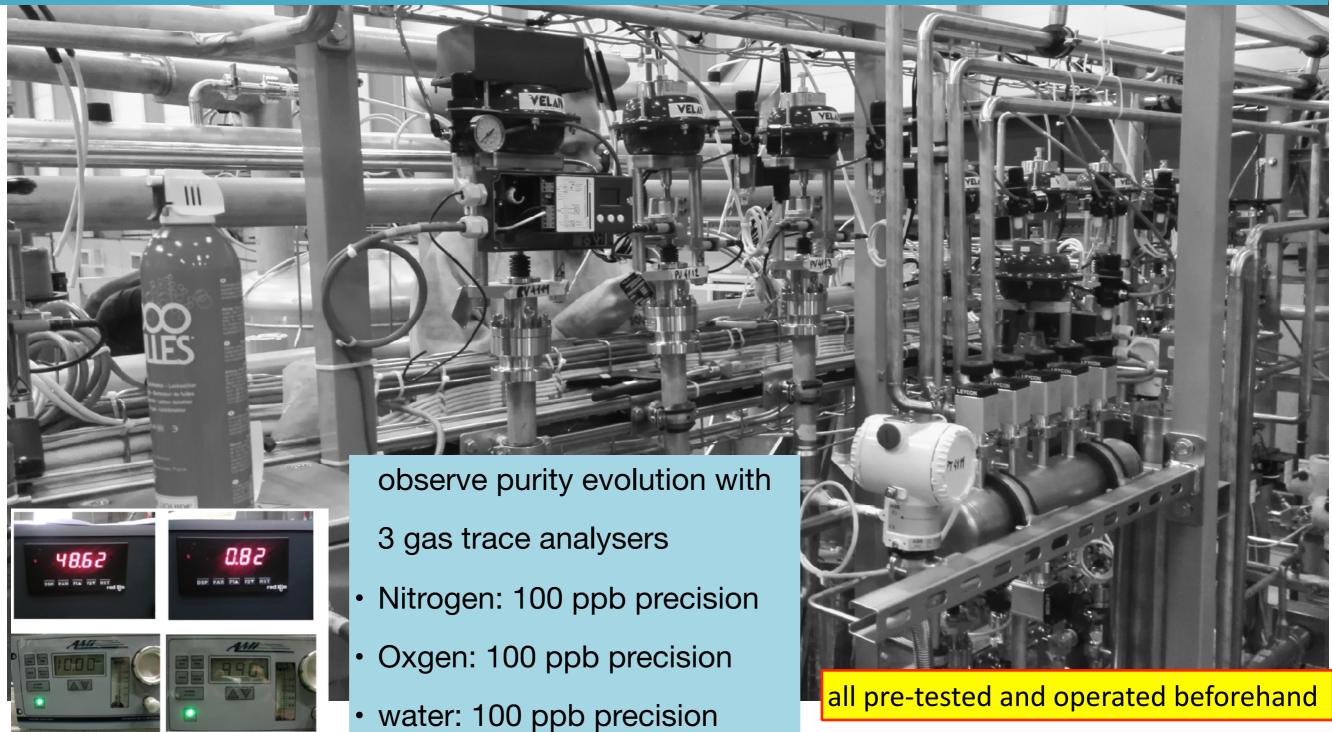
muon tracking reconstruction



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Remove contaminants by flushing gas argon and recirculating and filtering in closed loop.

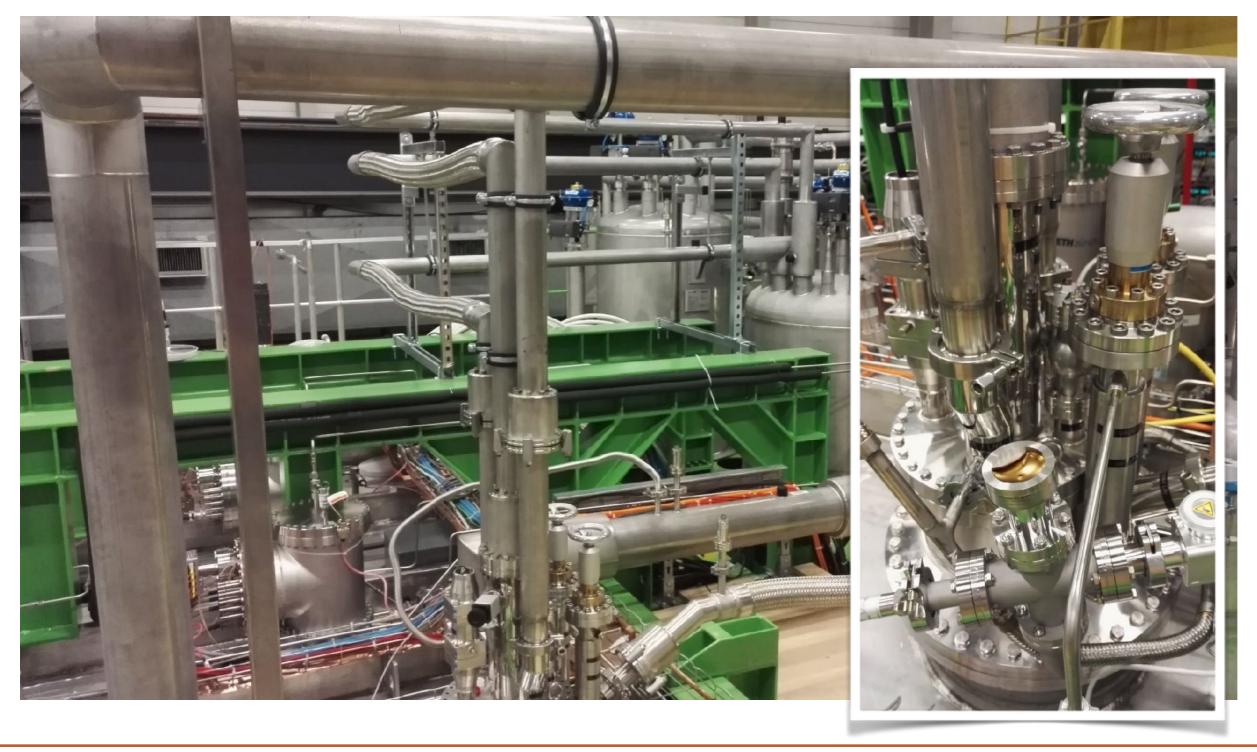




On December 19th the Manhole was sealed and GAr was flushed to pressurise the tank...

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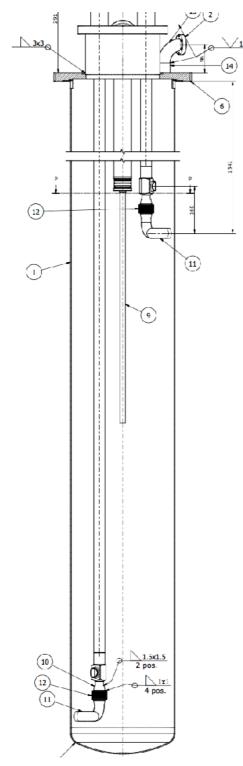
Large leaks we found on both flanges of the pump tower about 30 flanges on top-cap (+ piping)

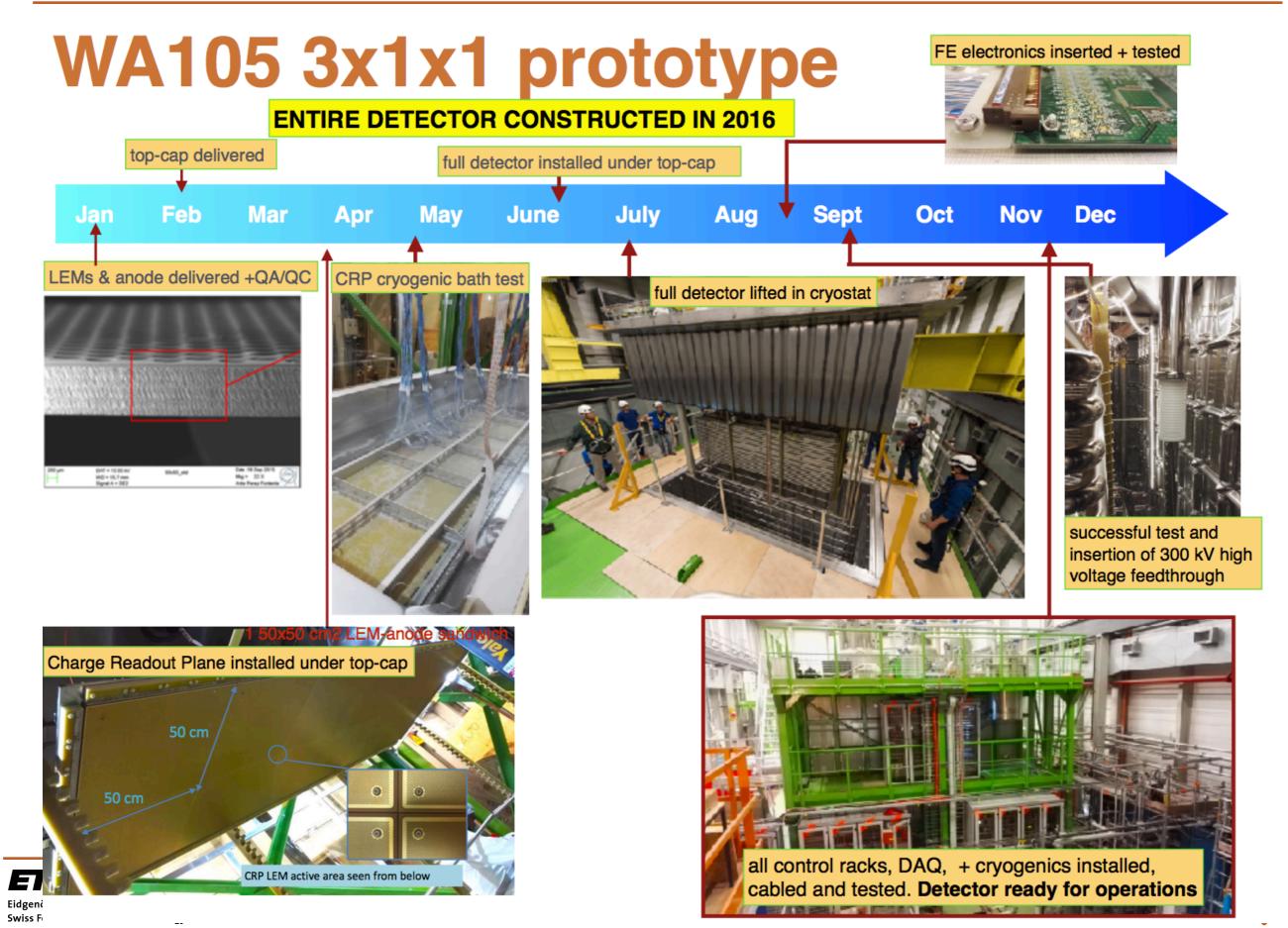


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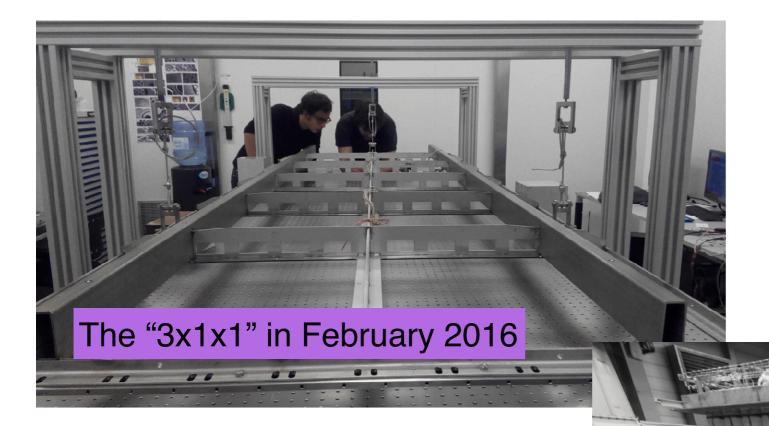


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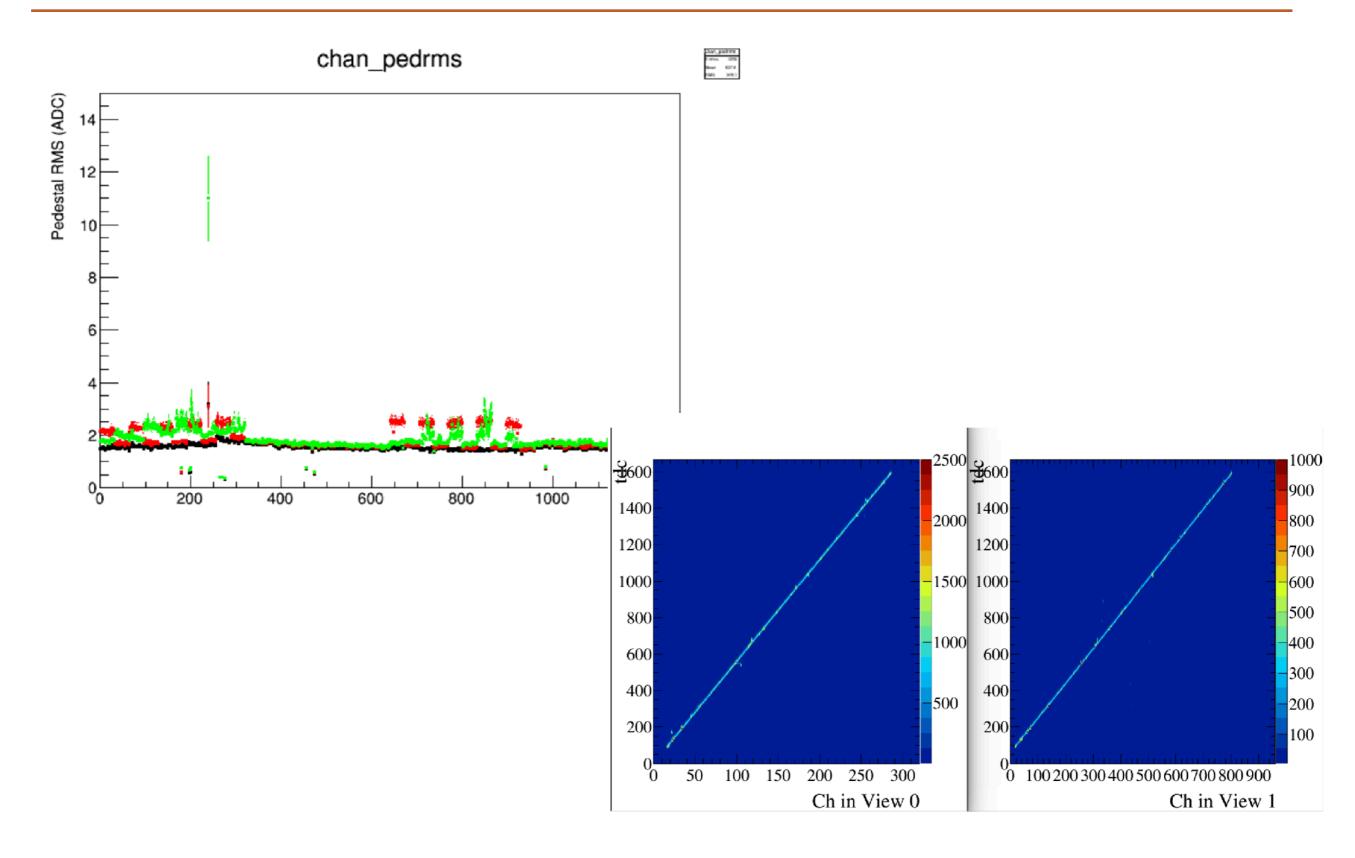




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6 months later.



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AND THE SAME FOR PROTODUNE-DP!

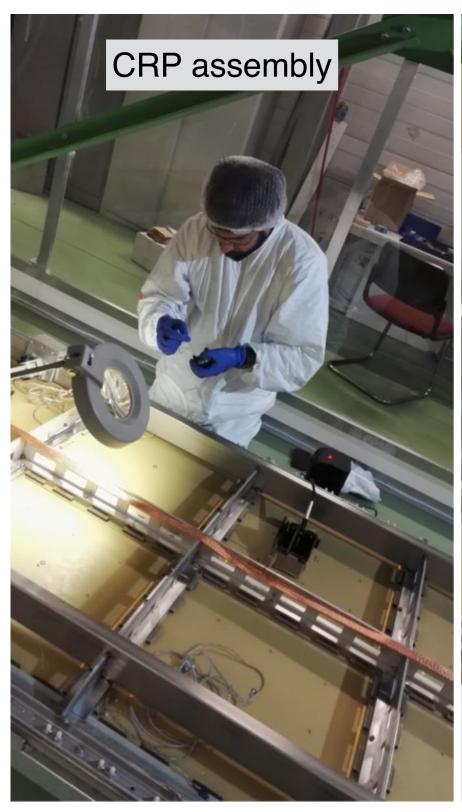
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See D. Autiero plenary + parallel session Wednesday









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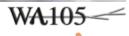


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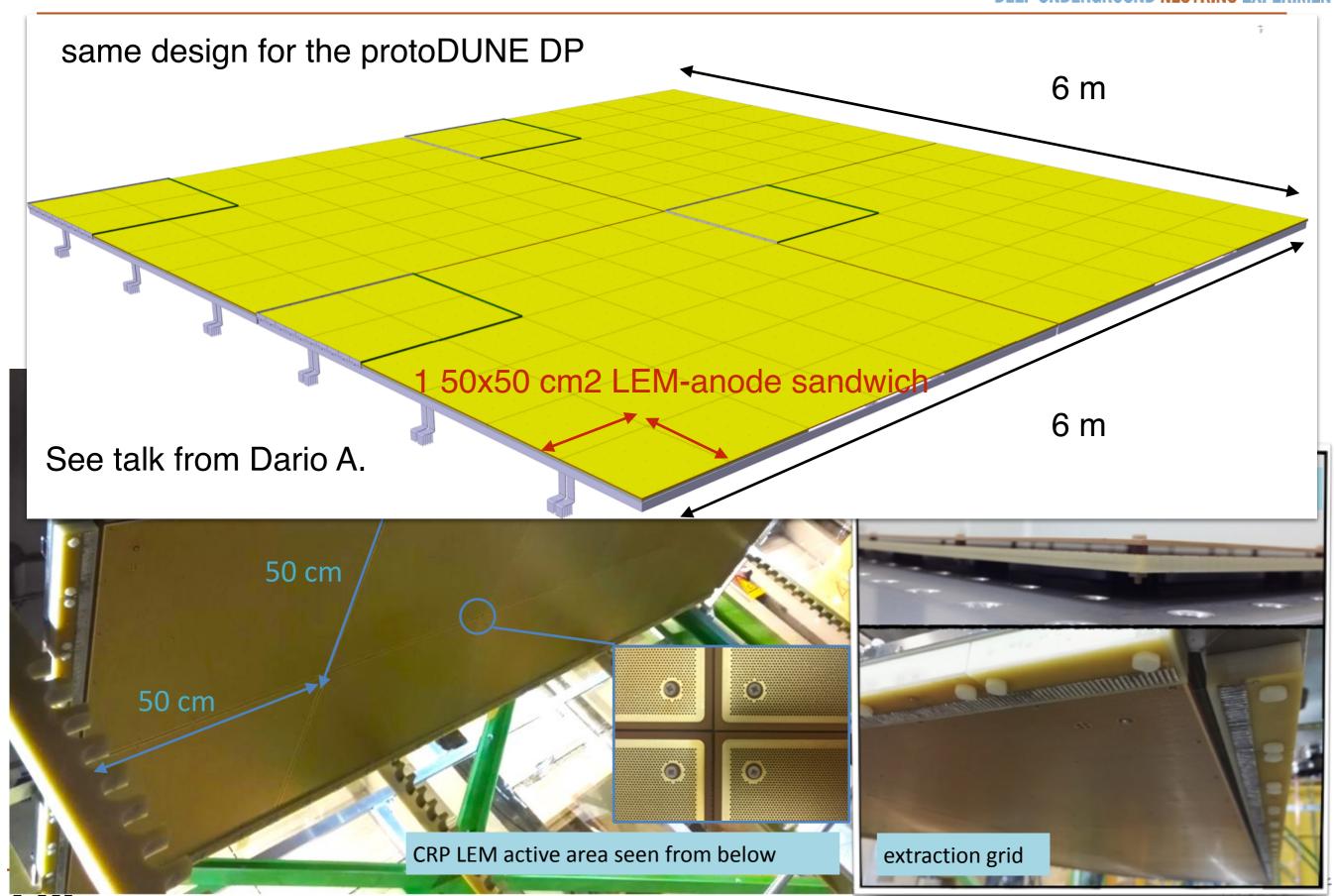




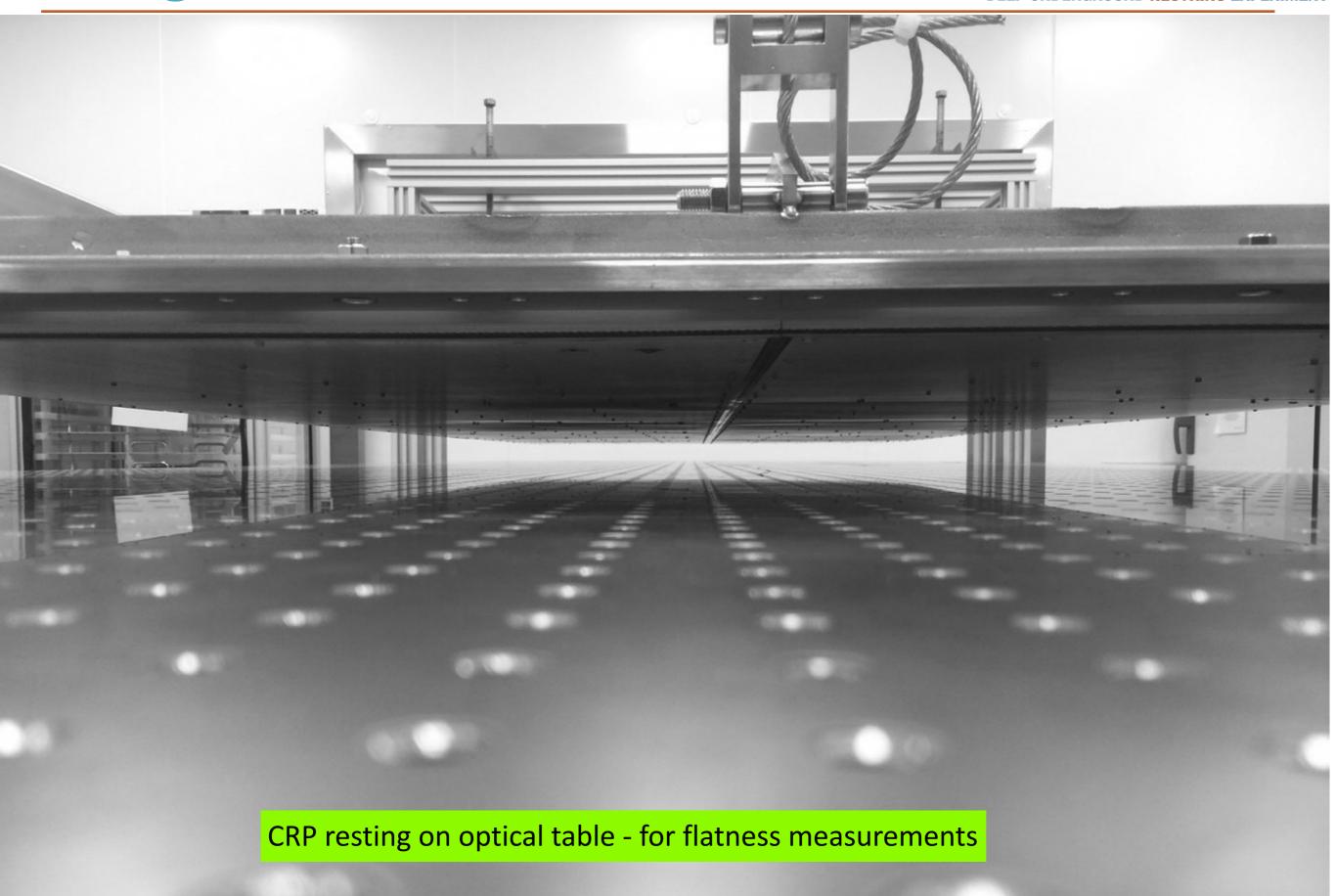


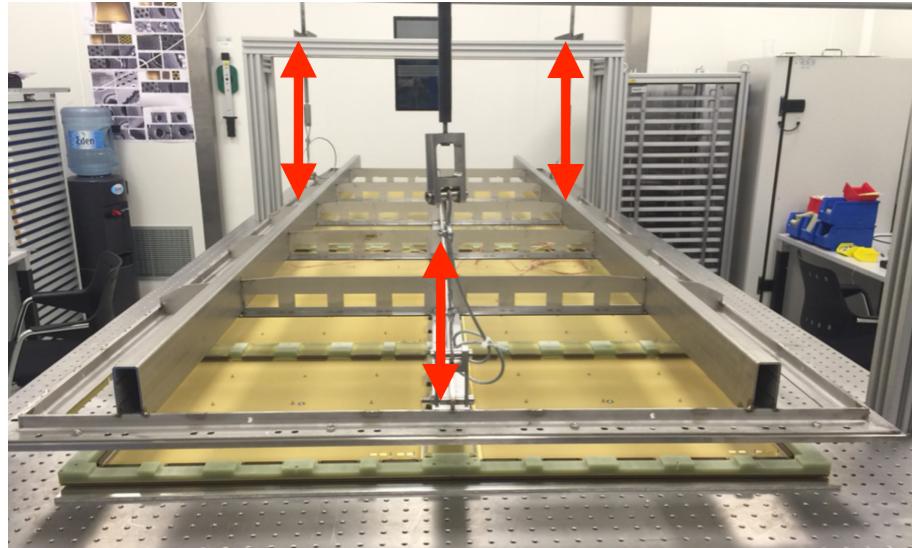


ETH



Charge Readout Plane











suspended by 3 ropes coupled to motors on topcap. Precision of motors 100 um over 4 cm. 8 capacitive level meters readout the LAr level with similar precision



Inspection and measurements inside cryostat

DEEP UNDERGROUND NEUTRINO EXPERIMENT

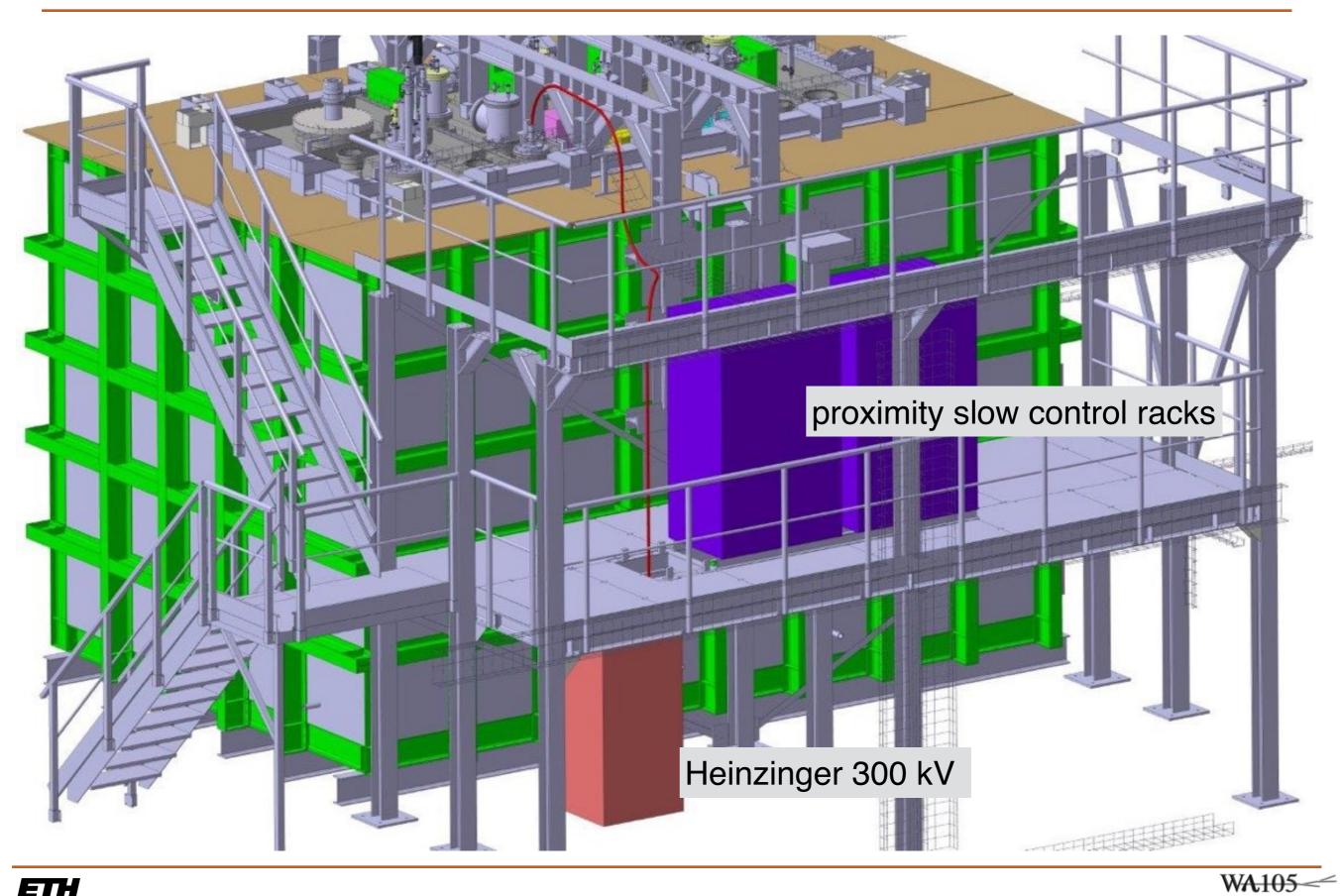








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10 Gb CERN Network for computing center (EOS, etc..)

~30 m

~10 m

Fibber

connection from uCTA to digitised signal fro proximity uTCA on signal rack feedthroughs

proximity DAQ rack

Fibber connection: - 2 x 10Gb rate : link aggregation in between 2 optical switches

protoDUNE DP Project Readiness Review 24 April 2017

computing farm



Proximity rack composed of:

- event-builder (46 TB storage server)
- 1 switch HP Procurve 6600 24xg → 2x10Gb
- White rabbit unit
- Cosmic counters crate/computer
- 4 x microTCA crate linked by fibber channel

~10 m

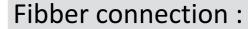


connection from uCTA to digitised signal fro proximity rack

proximity DAQ rack

10 Gb CERN Network for computing center (EOS, etc..)

~30 m



- 2 x 10Gb rate : link

aggregation in

betwee Farm rack composed of:

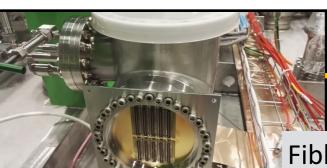
switche - 16 cpu modules with 16 cores inside

 4 storages servers : total space 192TB

- 1 switch HP Procurve 6600 24xg → 2 x 10 Gb

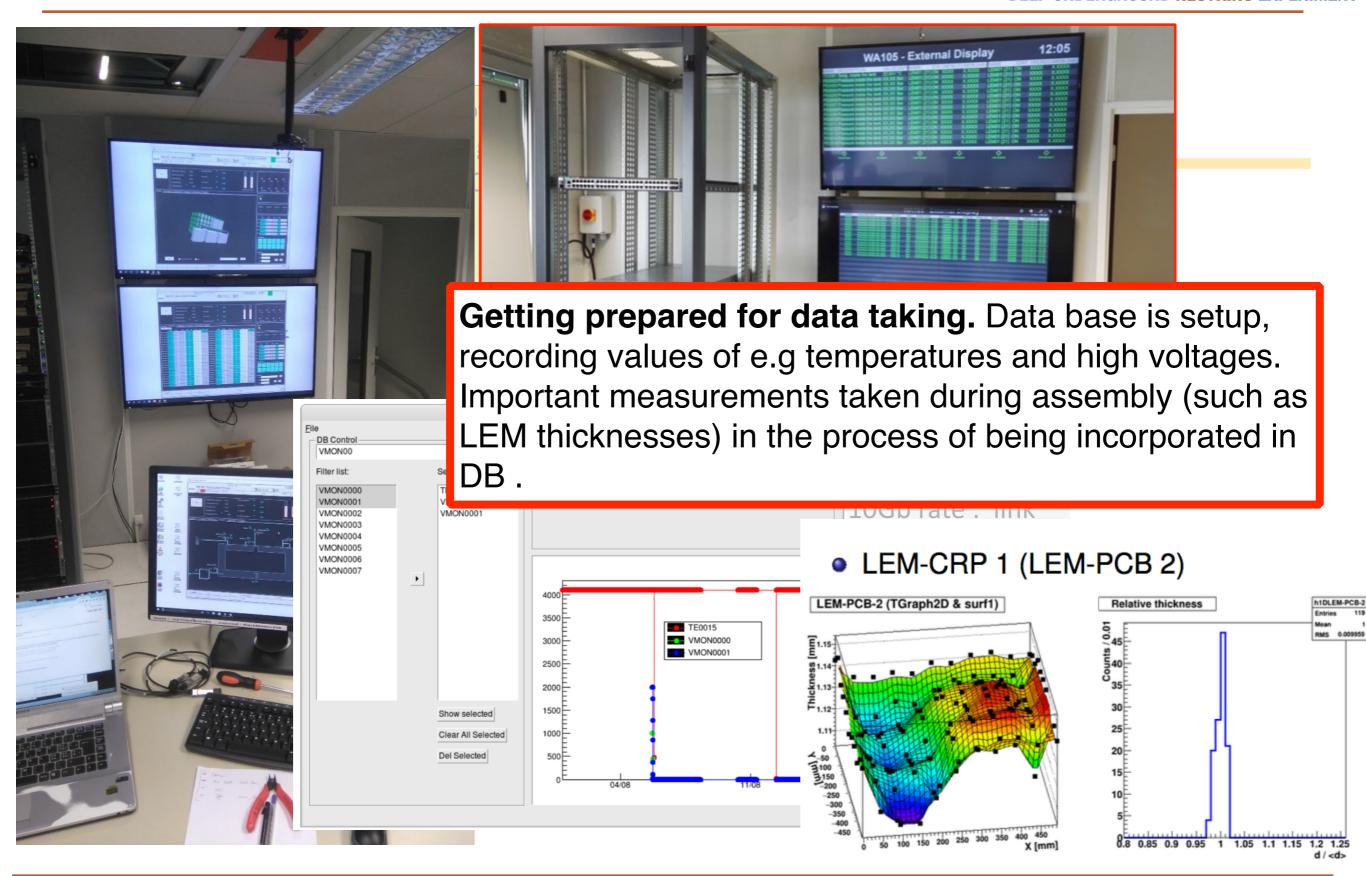
- 1 switch HP Procurve 6600 48g 4 xg

computing farm

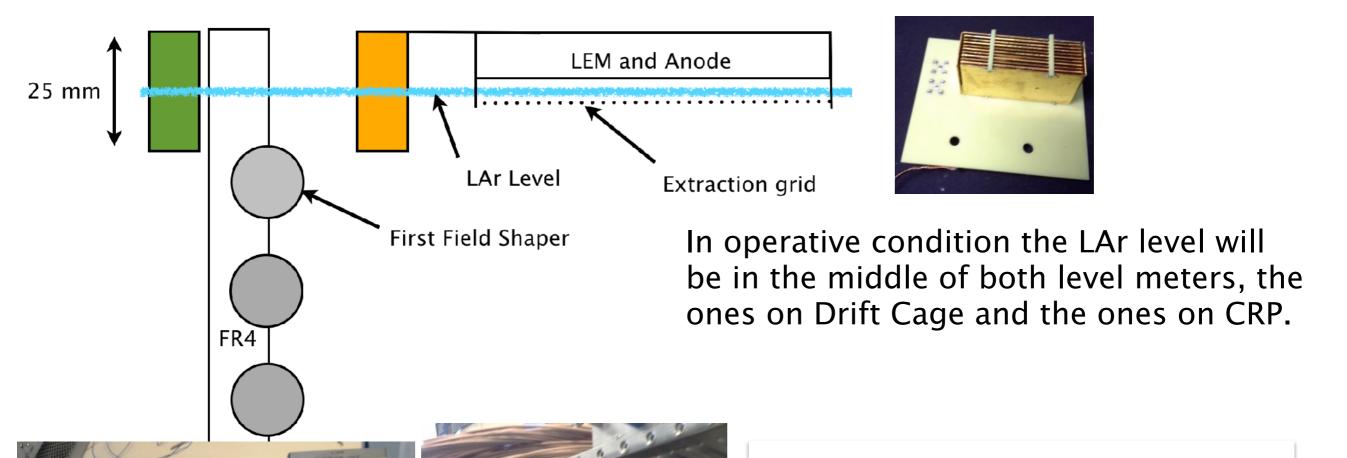


uTCA on signal feedthroughs





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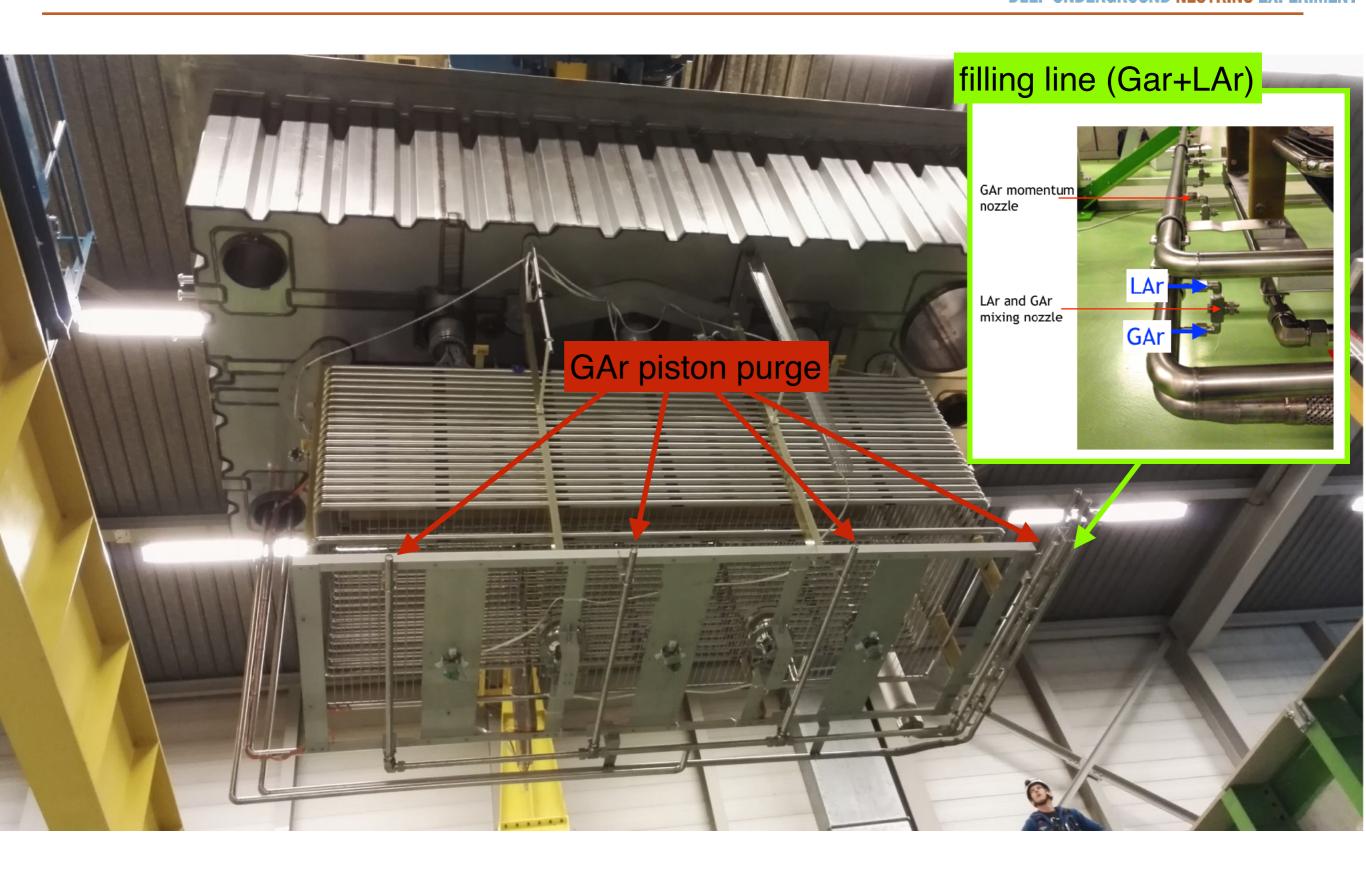


- A new NIM design has been done with:
 - 4 Channels
 - improved 0-10 V interface to NI Racks
 - improved filters on voltage rails and output
- 5 Boards NIM size are in production
- The assembly of the entire system will take place in the next 2 weeks, including calibration.
- Aso be a test bench for the 6x6x6 Level meter system



Cryogenics-internal piping

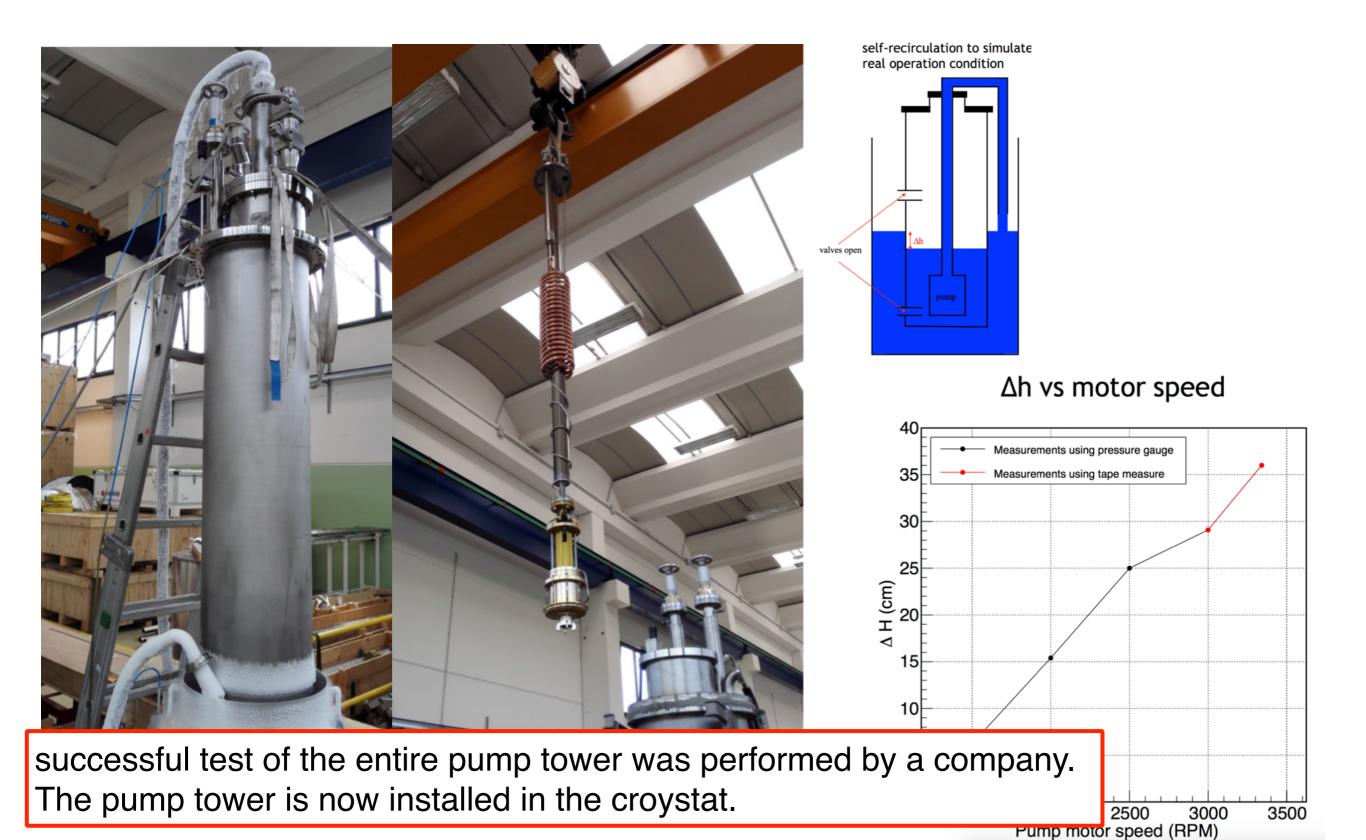
Sebastien Murphy ETHZ





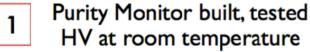


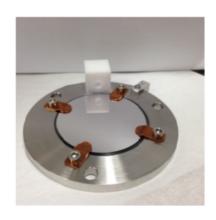
WA105



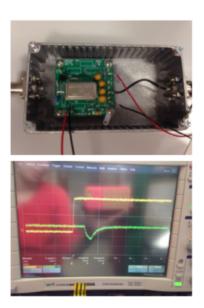
WA105



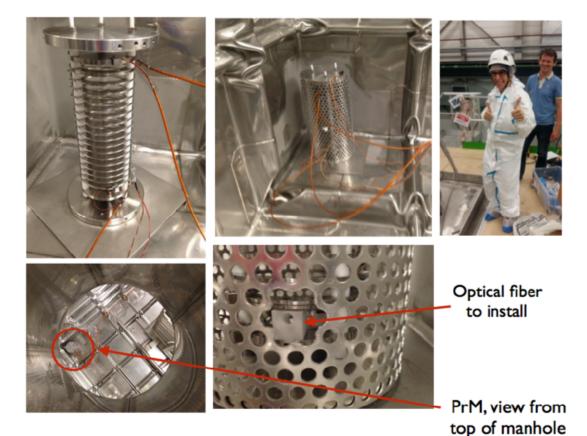




Tested PHC for no peeling in LAr



3 Charge amplifier tested

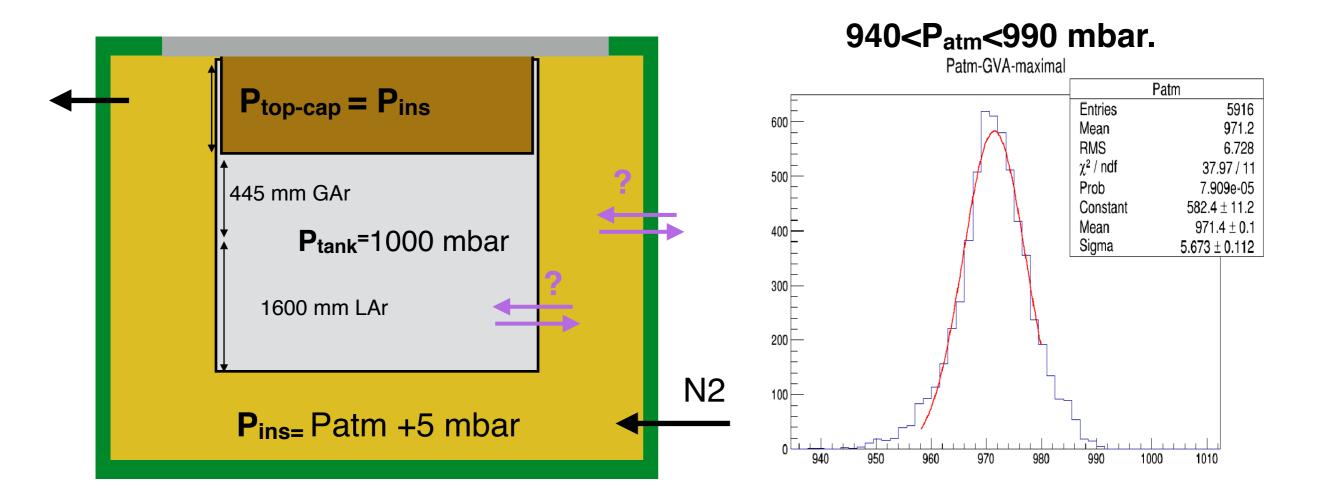


- 4
- PrM successfully installed in WA105
- Still to install optical fibre (1st week of October)
- Test planned when filling starts

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The cryostat - leak checking

Sebastien Murphy ETHZ



Want to keep Ptank constant at 1000 mbar independent from atmospheric pressure variations. Important to check tightness of the entire cryostat:

- 1. No leak from insulation space to inner tank above 1e-9 mbar I/s (to guaranty Ar purity)
- 2. leak-rate from atm to insulation space low enough to keep Pins= Patm +5 mbar