

## CRTs installation & performance

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WA105 collaboration meeting  
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<http://lbnodemo.ethz.ch:2500/3x1x1/35> :

1. feb-v3-0-1.pdf: description of FE electronics and communication
2. sbnd-crt-part1,2: technical design note for SBND CRT (90% compatible)
3. 3D CAD model for 3x1x1

16 strips per module

Strip width: 112 mm

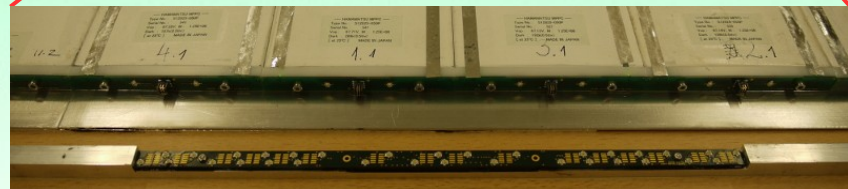
Strip length: 1755mm

Module size:

1.8x1.8m

Aluminum case

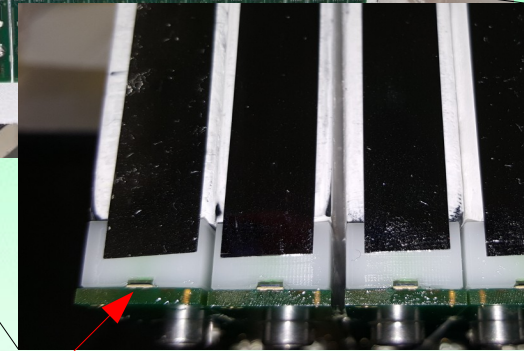
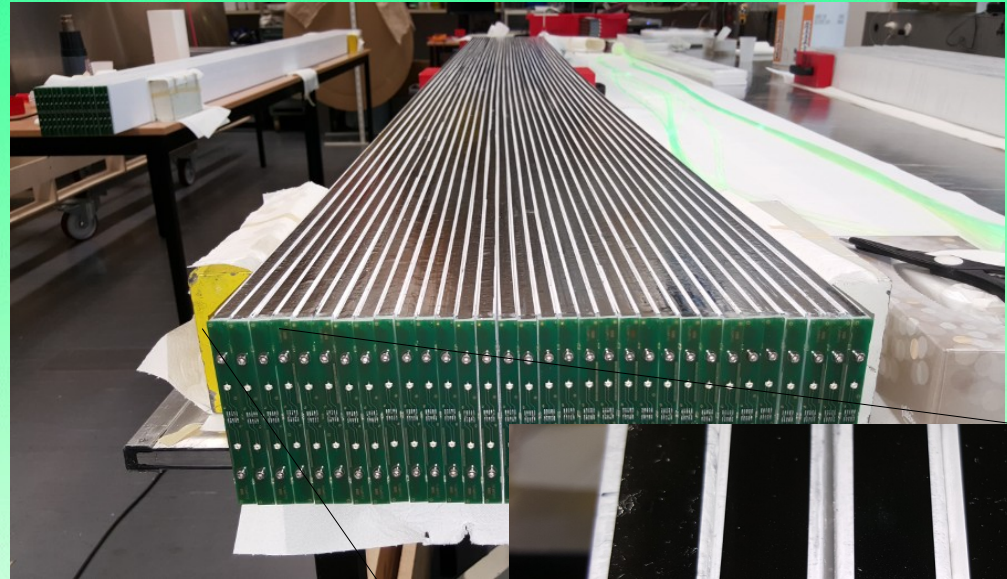
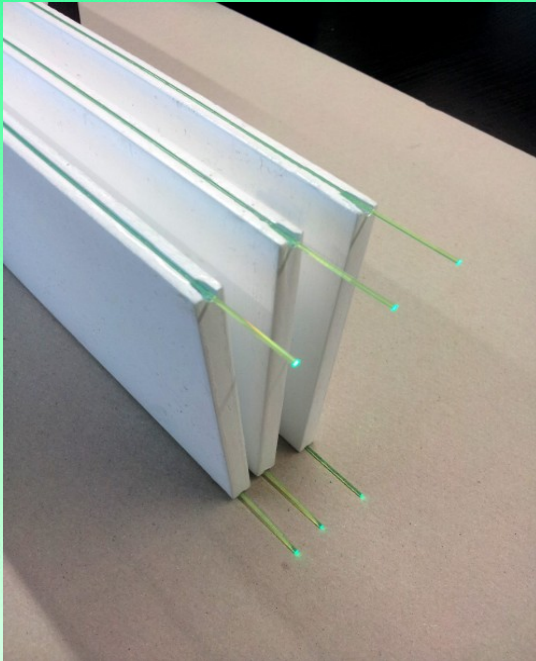
(2 mm thick)



Feed-through PCB

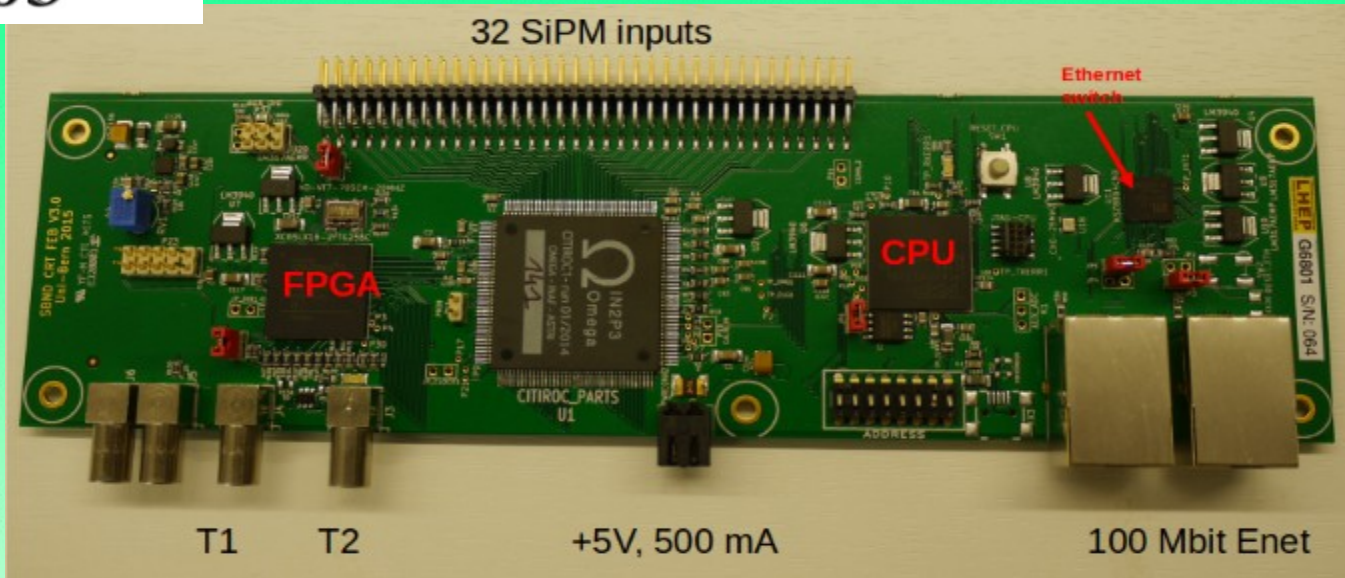


32 coax cables inside

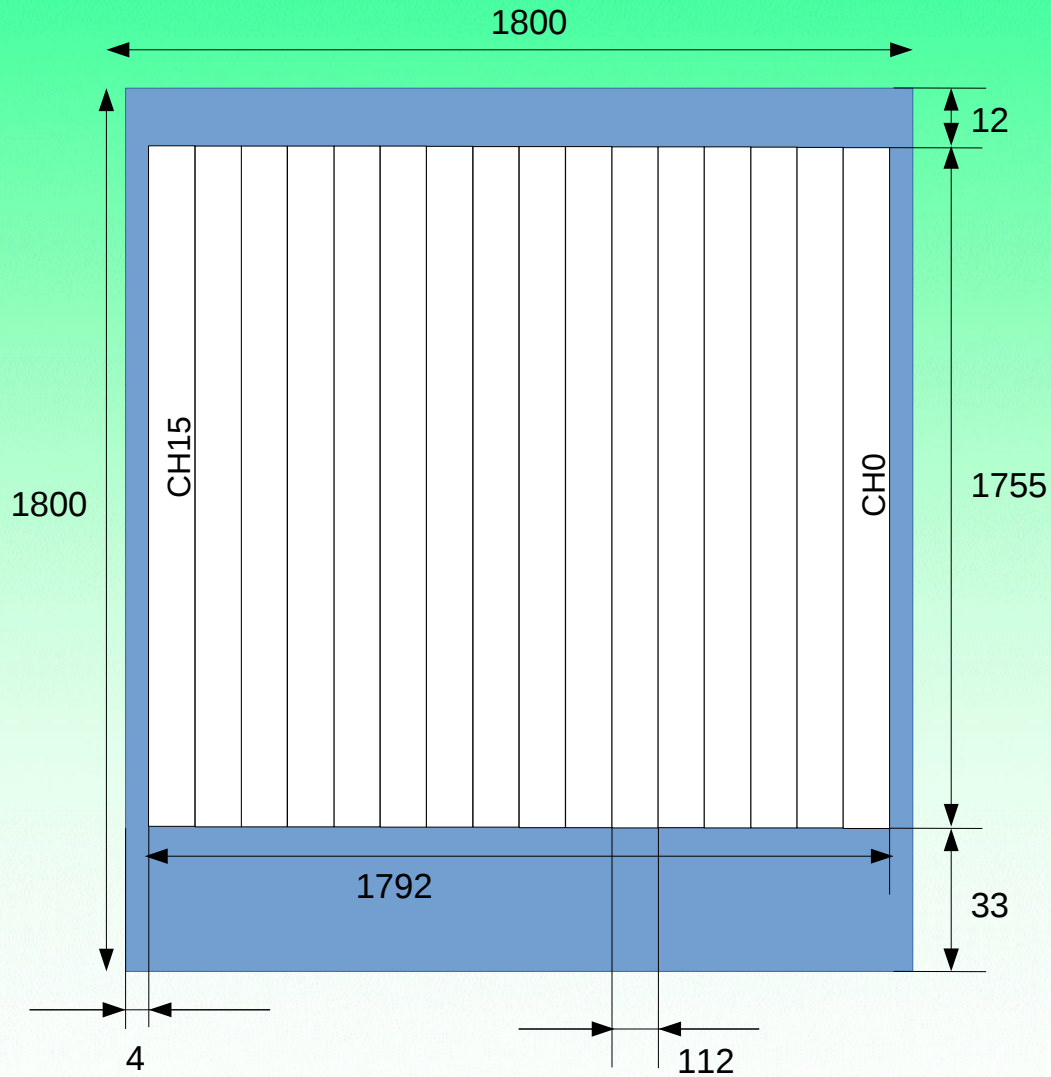


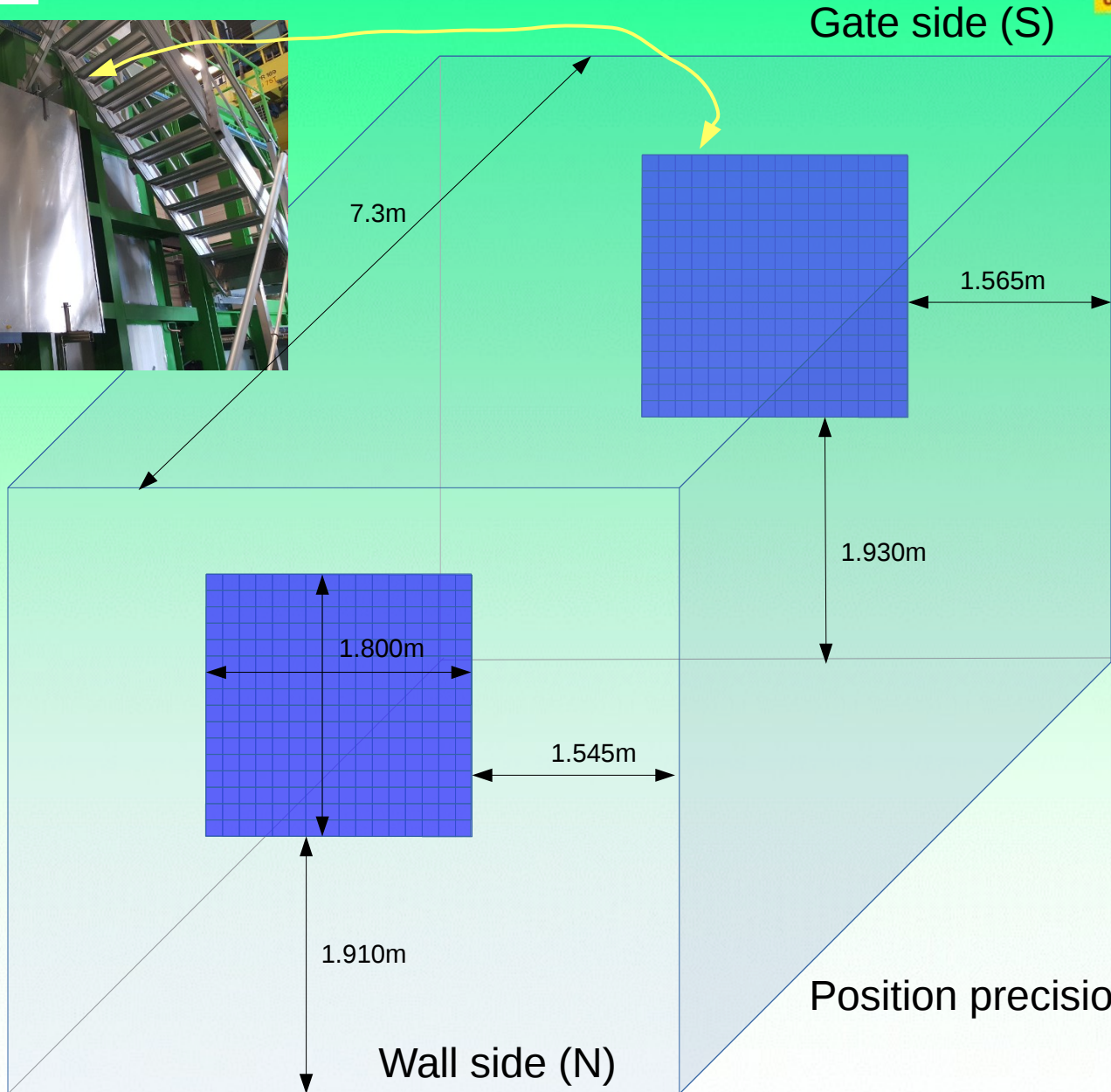
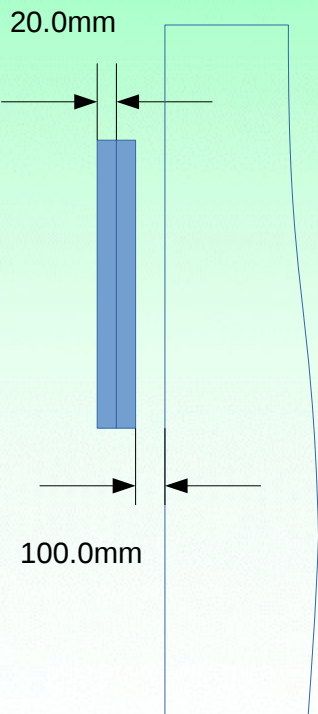
Scintillator: USMS-03 (PS+PTP+POPOP)  
Reflective surface (UNIPLAST technology)  
WLS fibers: Kuraray Y11(200)MS, 1mm diameter  
Optical glue: ESA 7250 polysiloxane compound  
SiPM: Hamamatsu S12825-050P

2 SiPMs per strip

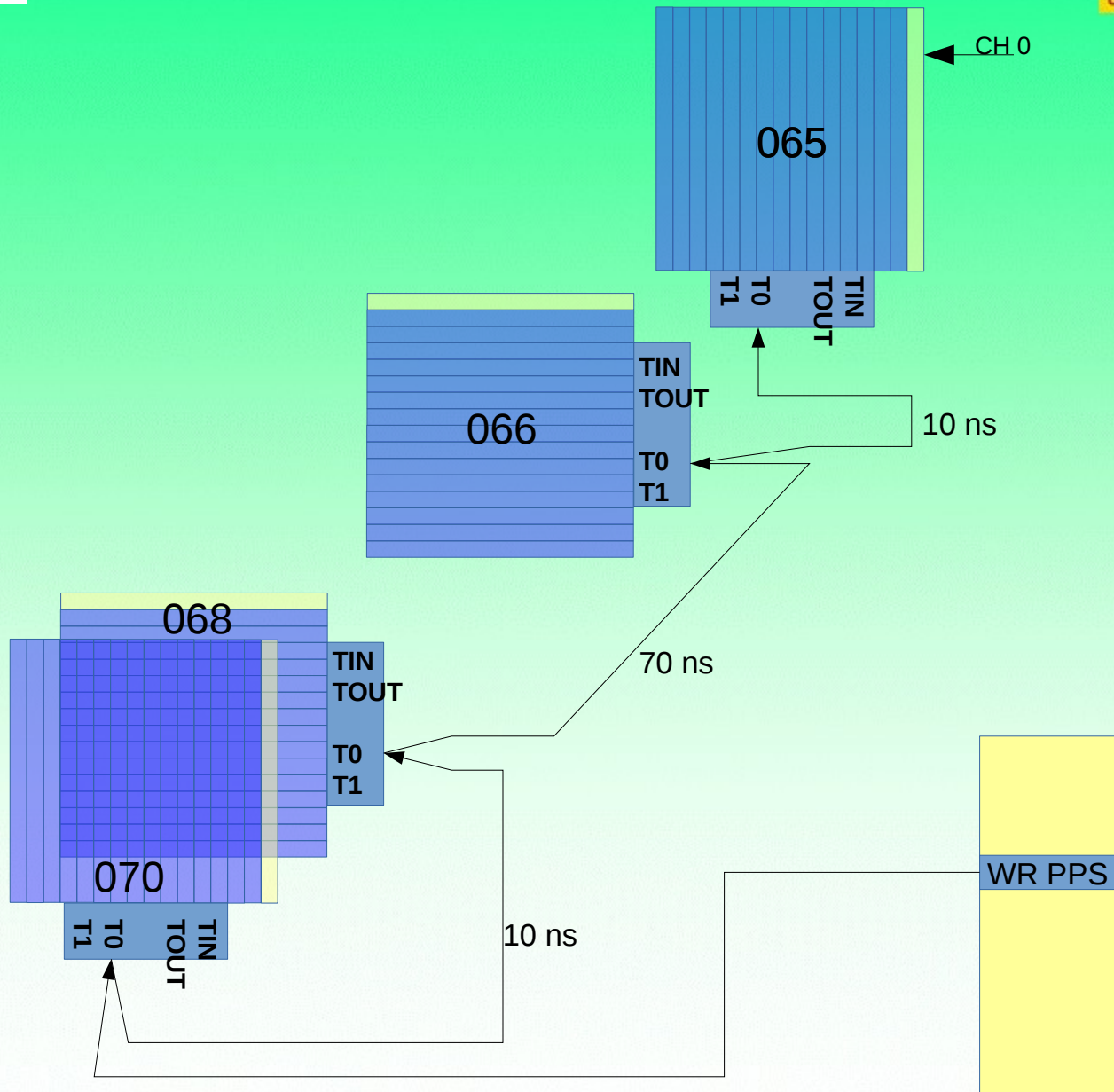


50 used in uBooNE, 140 in SBND, commercialized by CAEN (A1702)





Position precision  $\pm 3$  mm

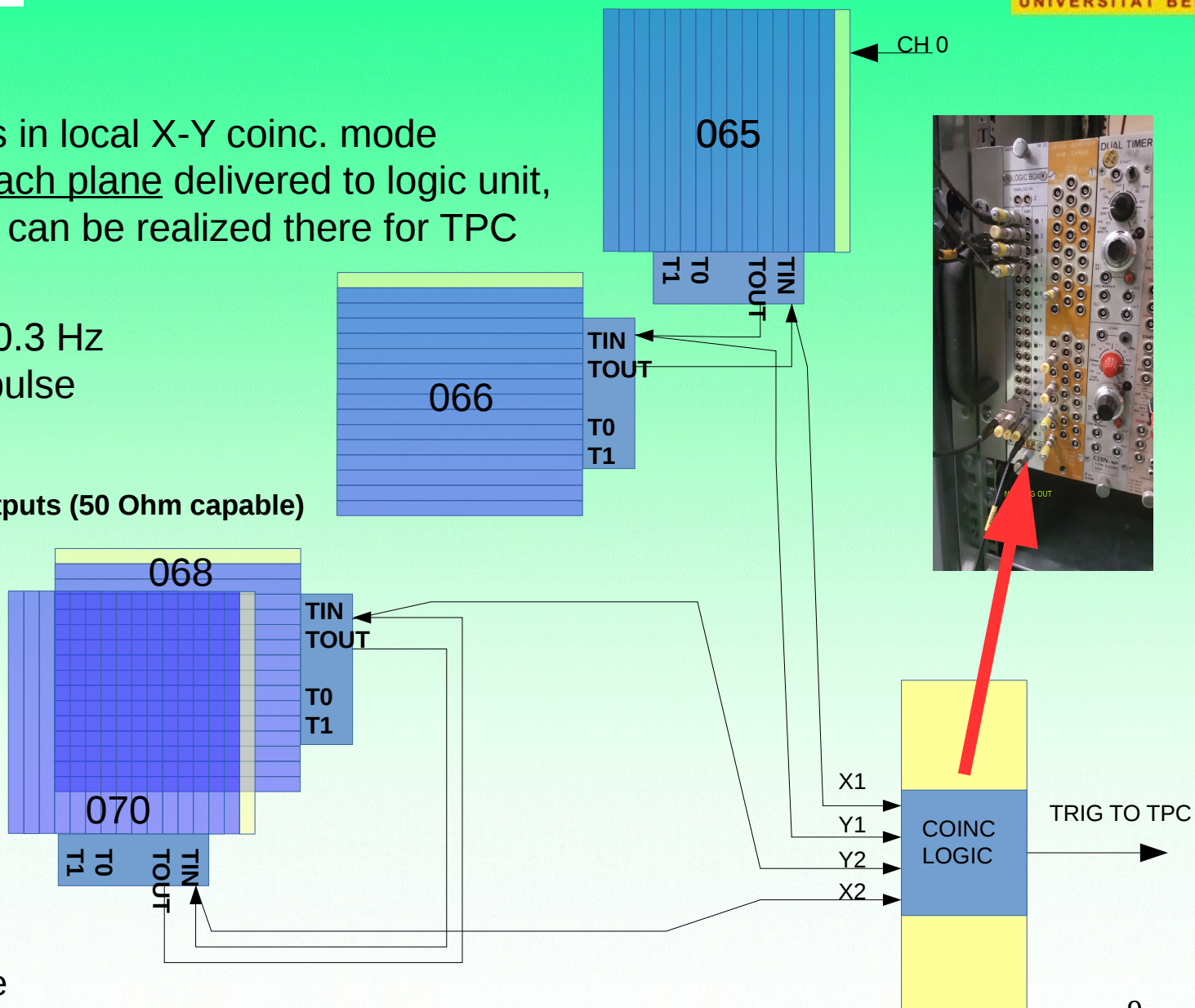




Each pair works in local X-Y coinc. mode  
Triggers from each plane delivered to logic unit,  
any coinc. logic can be realized there for TPC

Trigger rate ~ 0.3 Hz  
500 ns - long pulse

2 x NIM outputs  
2 x 3.3LVCMOS outputs (50 Ohm capable)



Configurable:  
channels can  
be active/inactive

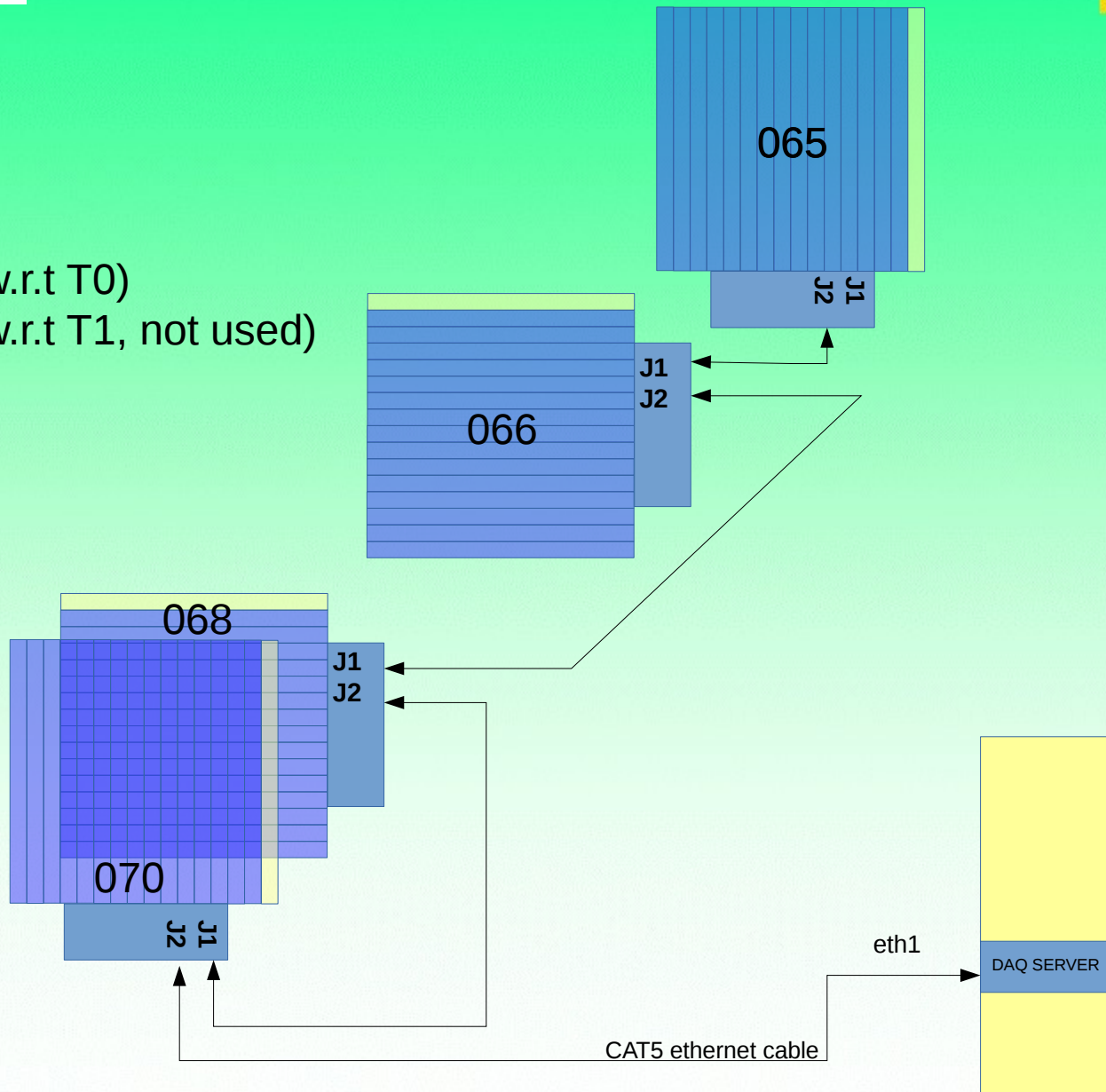
Event:

flags

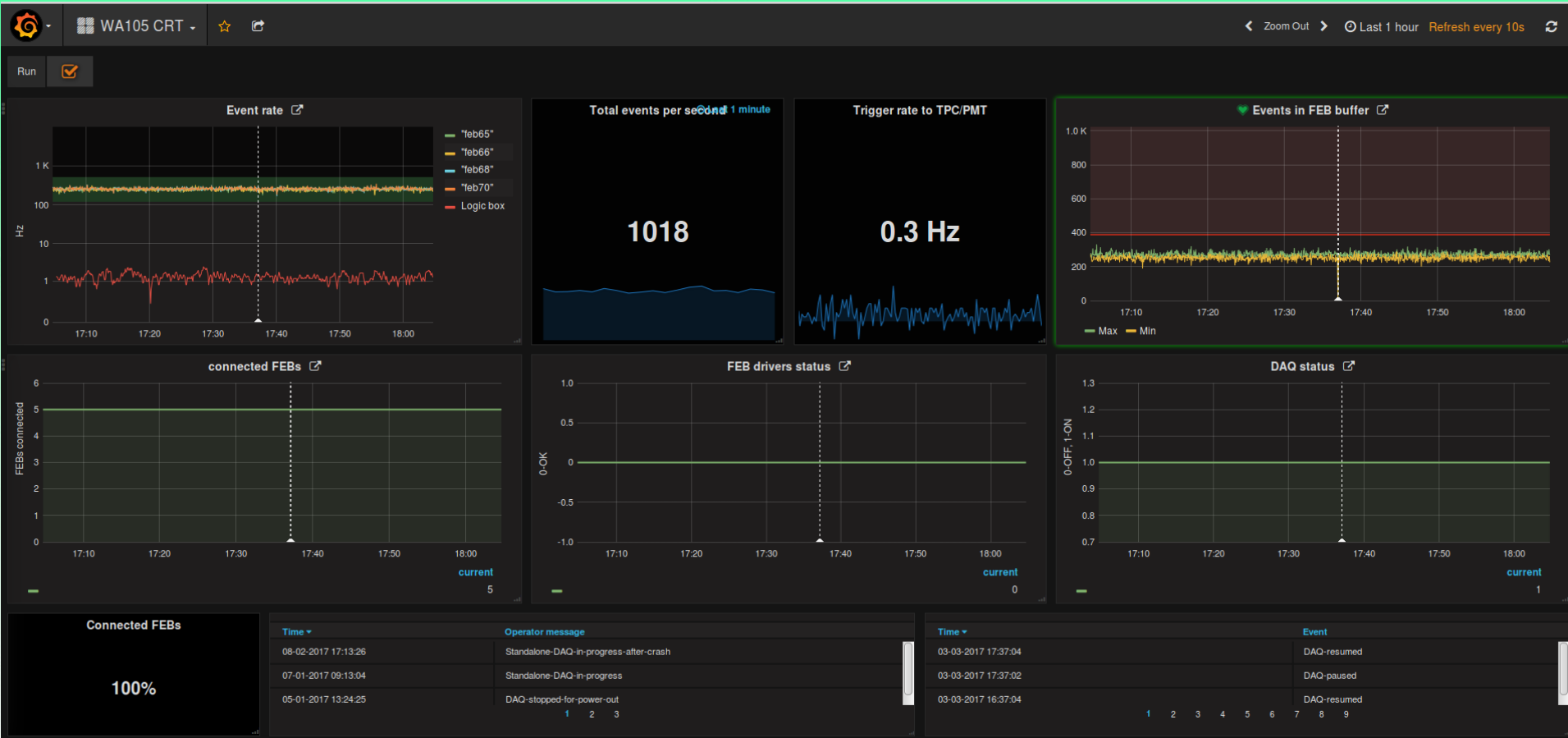
ts0, ns (time w.r.t T0)

ts1, ns (time w.r.t T1, not used)

ADC[32]

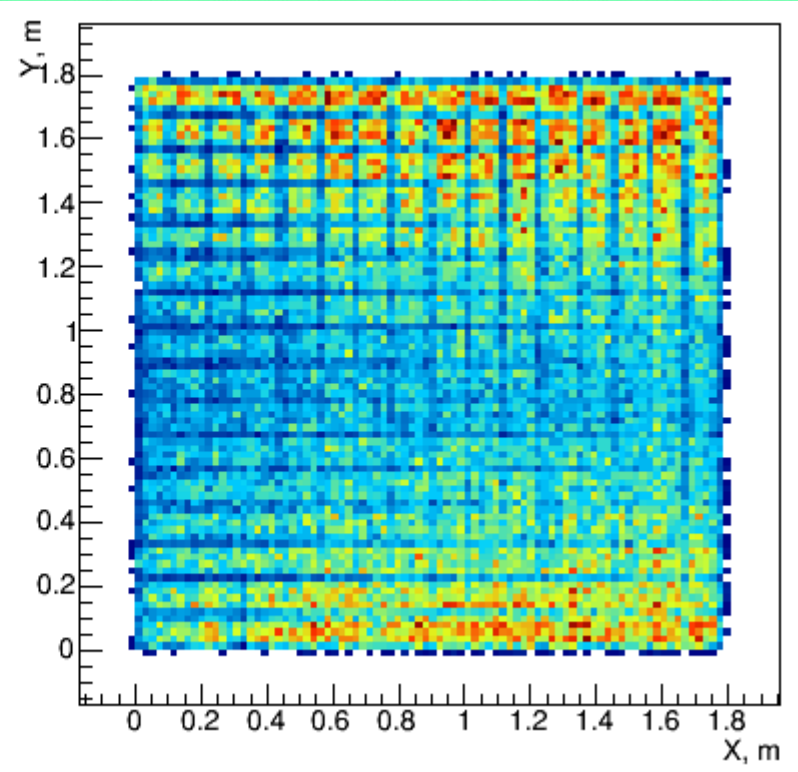
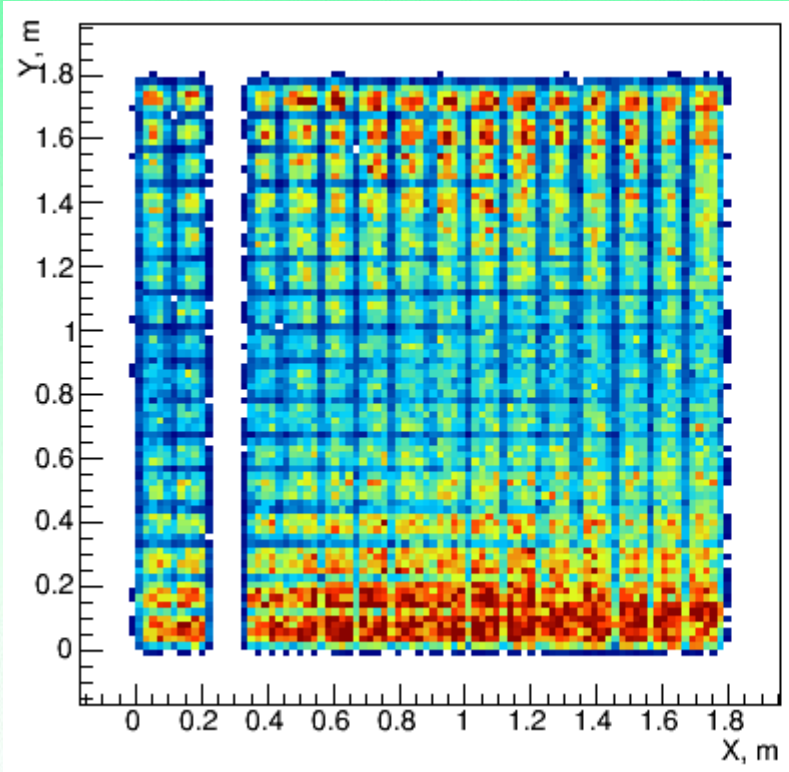


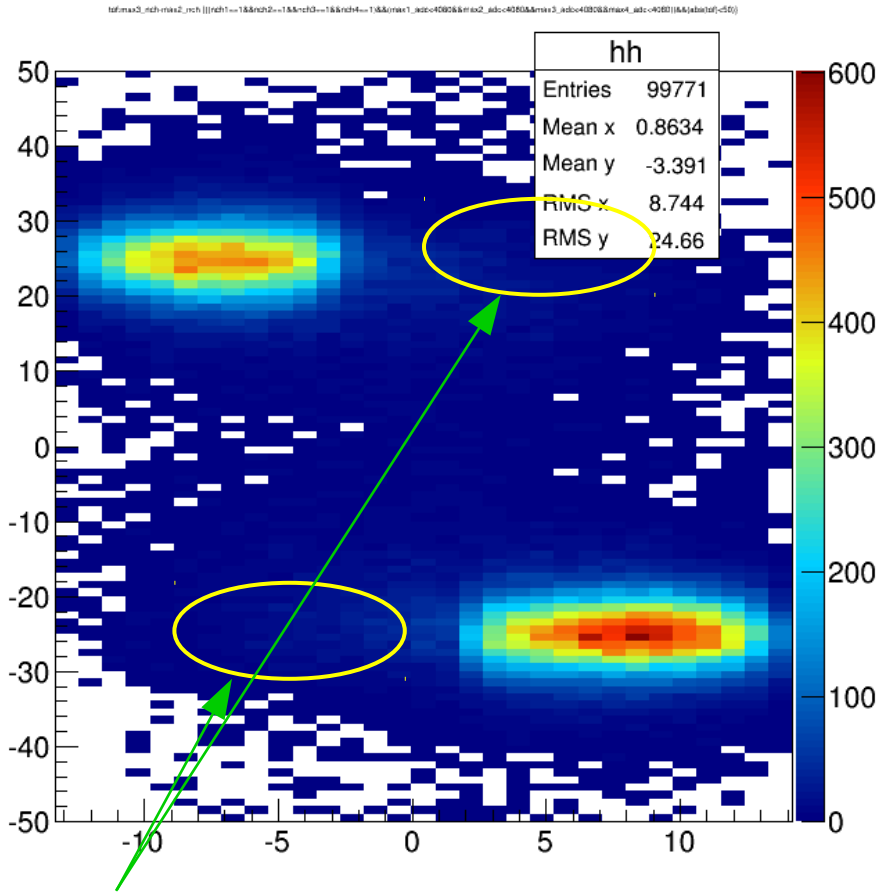
# CRT Slow Monitor (based on Grafana)



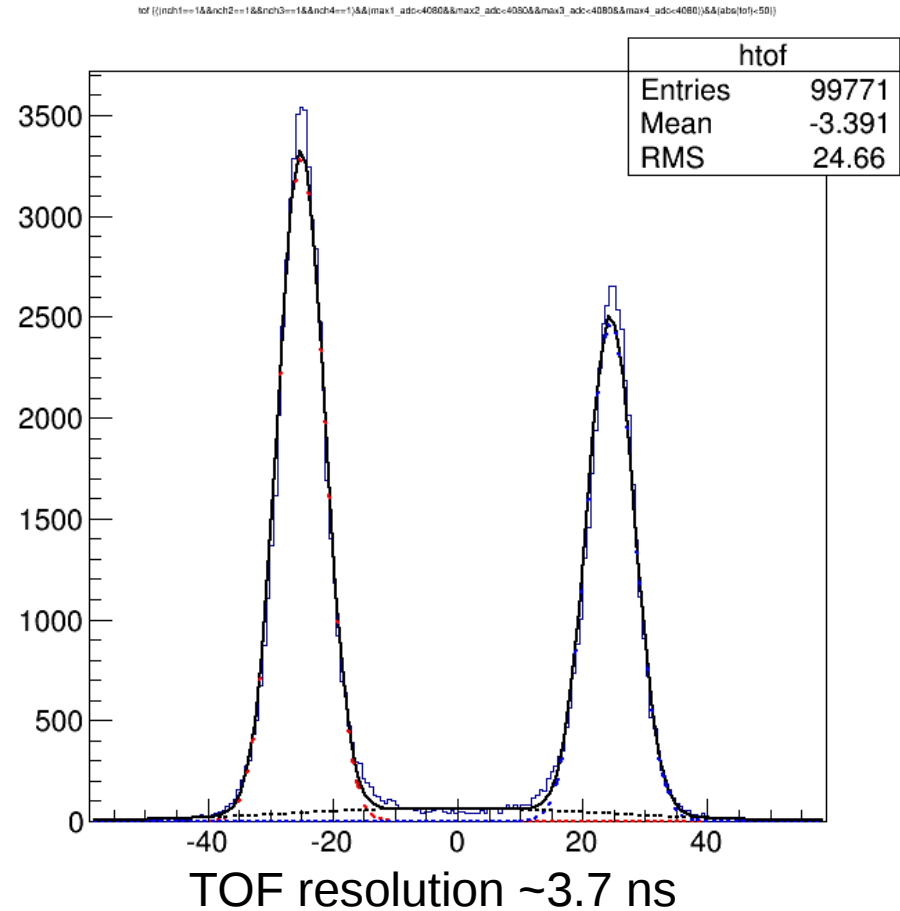
available at <http://wa105crt.cern.ch:3000>

Hit map in two planes (muons in both directions summed up)





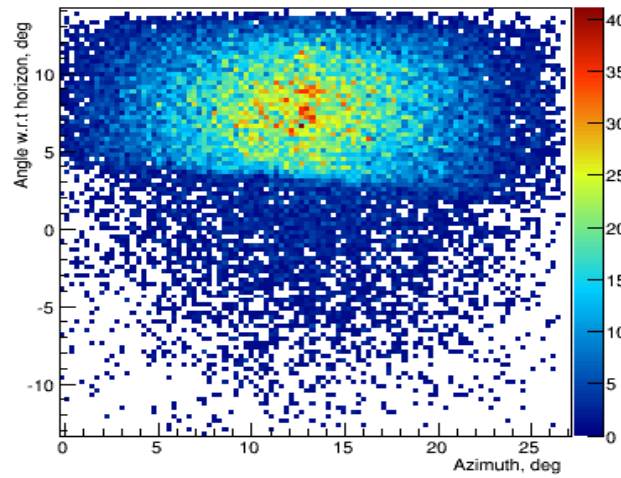
Upward-going muons



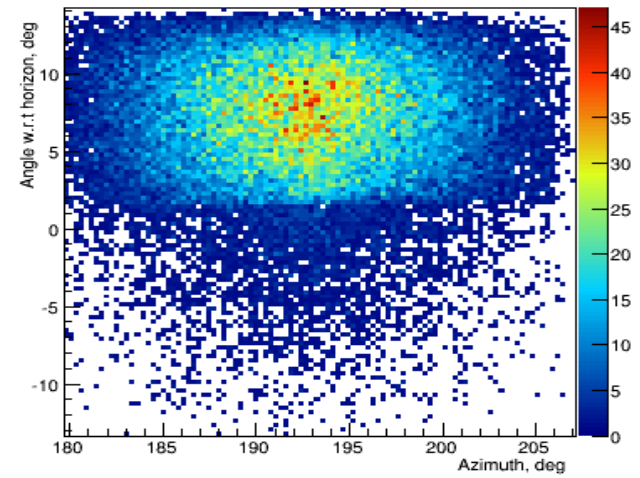
Detected "good" muon flux:  
 N->S direction: ~0.9 muons/min  
 S->N direction: ~1.0 muons/min

Raw events

Detected muons, from NW



Detected muons, from SE

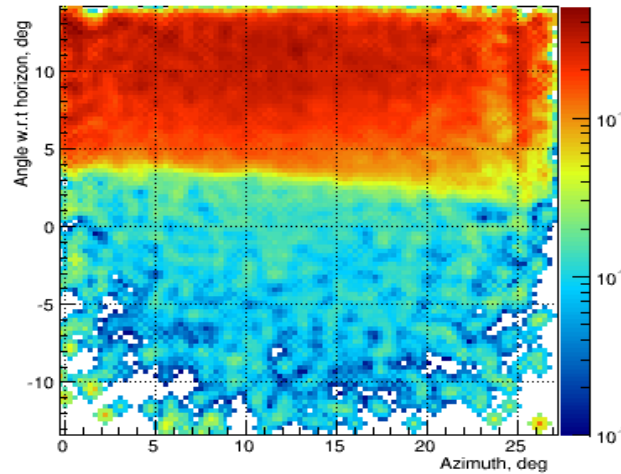


Normalized to:

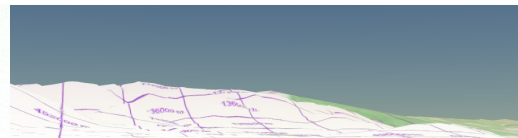
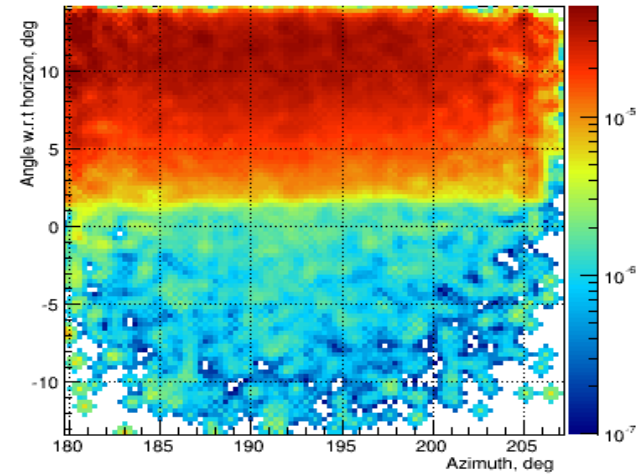
1. acceptance
2. run time

(flux in  $\text{s}^{-1}\text{sr}^{-1}\text{cm}^{-2}$ )

Muon flux, from NW



Muon flux, from SE



## CRT trigger for TPC/PMT

1. Hardware trigger (NIM/LVTTL pulse, 500 ns)
2. Event data available at Zero-MQ socket at TCP port 5600  
request- response logic
3. Event matching by sequence + crosscheck by event time stamp

## Conclusions

1. Cosmic Ray Tagger for  $3 \times 3 \times 1$  is installed and operative
2. Configurable geometry - possible to select narrow solid angles
3. Trigger provided to the experiment
4. DAQ is deployed, Slow Monitor is operative
5. Extensive data on the triggered event is stored: time, coordinates, pulse heights
6. Interesting data being collected
7. Ready for “wet&cold” commissioning



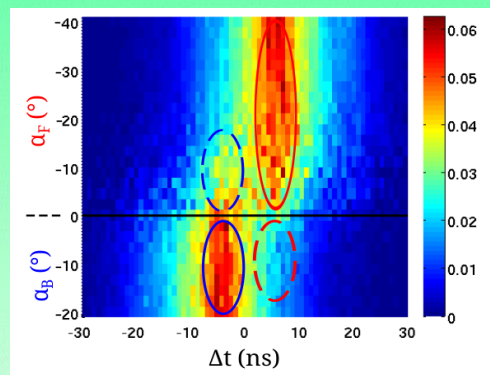
# Backup slides

# Upward going muons at surface ( literature refs )

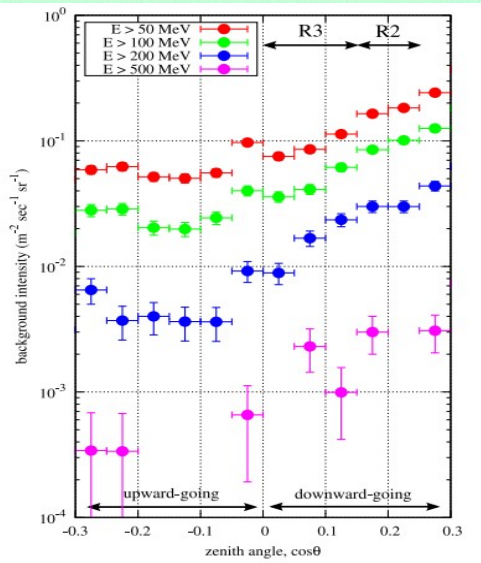
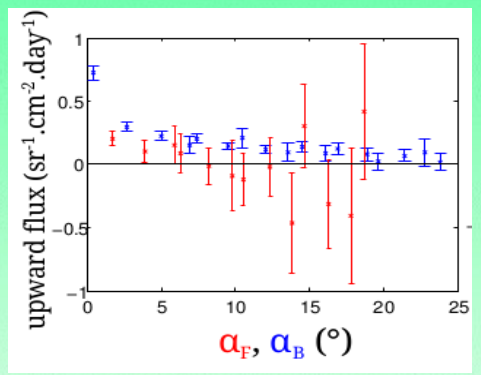
- Lack of information or not very reliable measurements
- PDG review on cosmic rays says almost nothing about flux near horizon
- “Effects of upward-going cosmic muons on density radiography of volcanoes”  
K. Jourde et al., arXiv:1307.6758v1

**Question remains:  
Low or high energy?  
(Scattered or through-going?)**

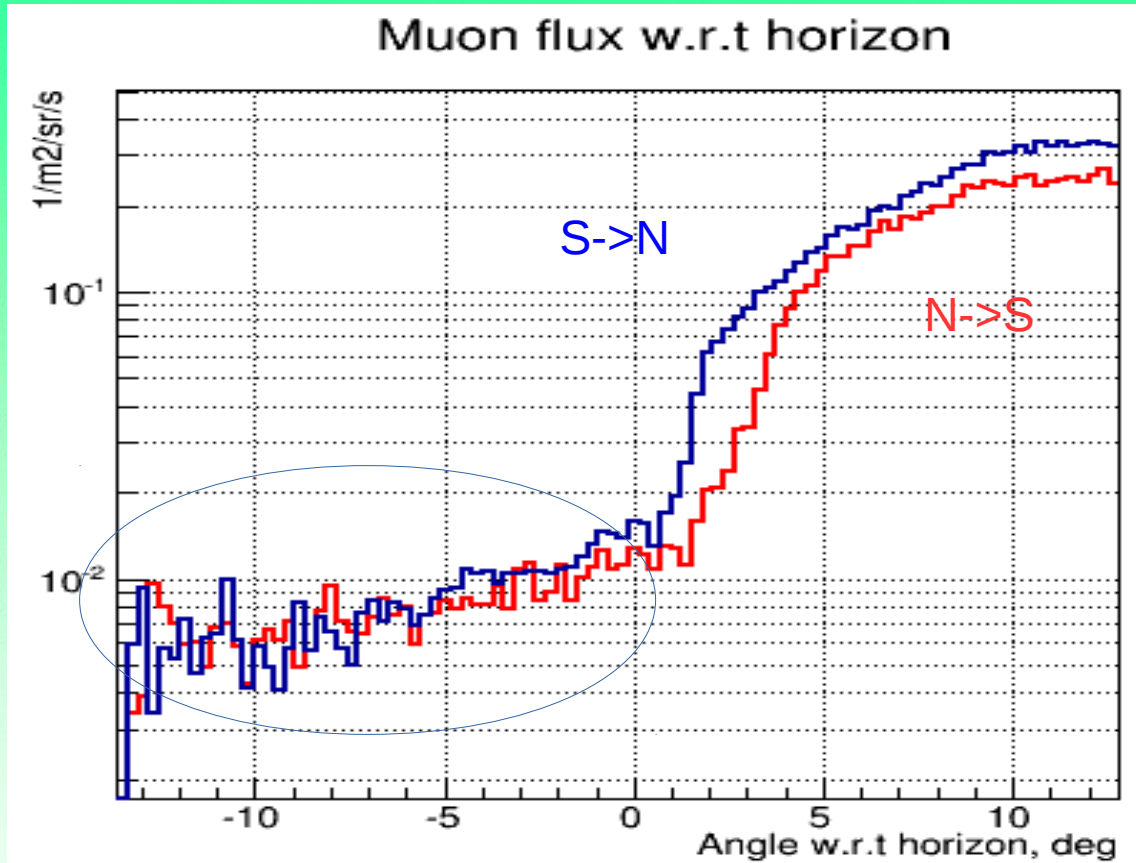
- Monte Carlo simulation for background study of geophysical inspection with cosmic-ray muons,  
R. Nishiyama et al.,  
Geophys. J. Int. (2016) 206, 1039–1050



**Figure 4.** TOF distribution for the SMTOMO data set shown as normalized histograms as a function of zenith angle. The horizon is represented by the dashed line. The blue and red solid ellipses respectively show the backward ( $\alpha_B < 0$  and  $\Delta t < 0$ ) and forward ( $\alpha_F < 0$  and  $\Delta t > 0$ ) events corresponding to the downward fluxes. The dashed ellipses show events corresponding to upward-going muons from forward (red ellipse,  $\alpha_B < 0$  and  $\Delta t > 0$ ) and backward (blue ellipse,  $\alpha_F < 0$  and  $\Delta t < 0$ ).



**Figure 6.** Zenith angle dependence of the calculated flux of the background particles for kinetic energy above 50, 100, 200 and 500 MeV (protons, electrons and muons are added).  $\cos \theta < 0 (> 0)$  indicates downward (upward) going particle.



### Proposal:

measure energy spectrum of particles near (below) horizon with the TPC data.  
 (if fast- delta ray count, delta ray spectrum, shower reco; if slow - MCS,  $dE/dx$ )  
 Clarify “through-going vs scattered” dilemma.