

Quantum computing technology for particle detection

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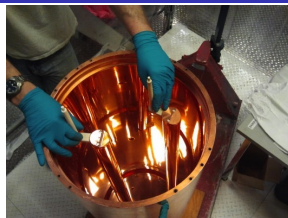
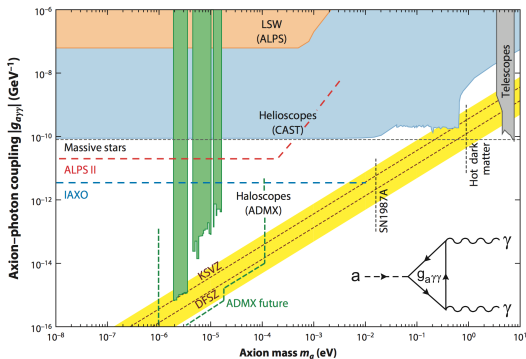
Detectors for Science Working Group Pre-Meeting 2

25 April 2017

A brief introduction to QCD axions

- Locally coherent oscillation of QCD vacuum angle θ about the CP-conserving minimum.
- $\theta(x, t) = \theta_{\max} e^{i(kx - m_a t)}$, with $\theta_{\max} \sim 10^{-19}$
- While WIMPs are point-like and suited to bolometry, sub-eV axions are field-like and require unique detector strategies.

QCD axions: well-motivated, but the mass is not well-constrained



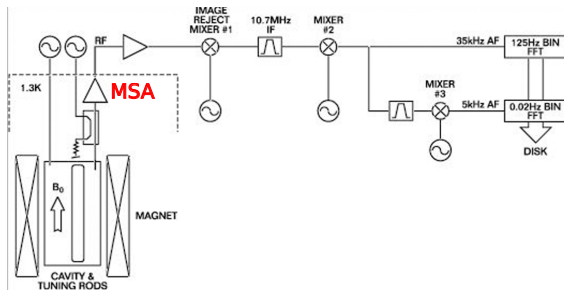
ADMX 500 MHz - 1 GHz
“haloscope”

- $\mathcal{L}_{a\gamma\gamma} = g_{a\gamma\gamma} \mathbf{E} \cdot \mathbf{B}$
- $P_{a\gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_0}{m_a} B_0^2 V C_{nml} Q_L \sim 10^{-23} \text{ W}$
- Current haloscope experiments looking at $\leq 2 \text{ GHz}$ and $\sim 6 \text{ GHz}$.
- SN1987A give us an upper limit $m_a \sim 250 \text{ GHz} \dots$

Signal integration is nontrivial.

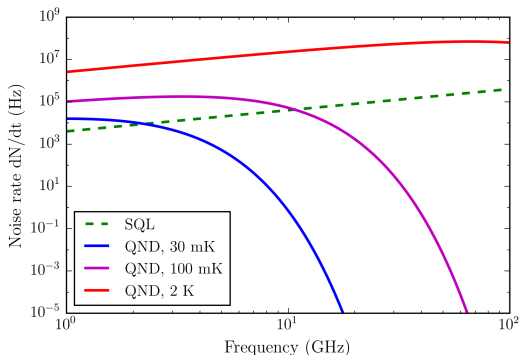
- $P_{a\gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_0}{m_a} B_0^2 V C_{nm\ell} Q_L$
- To boost signal power, you can:
 - increase Q_L (but $Q_0 = 10^6 \sim$ axion linewidth, and superconductors are not easy to use in $B_0 \sim 10$ T)
 - increase $B_0^2 V$
 - increase form factor $C_{nm\ell}$ via clever cavity design
- Alternative approach: lower noise rates!

Axion Dark Matter eXperiment layout



- Cavity operates at $O(100)$ K
- Superconducting electronics package (microstrip squid amplifier) in a field-compensated region
- Warm electronics: heterodyning, etc.
- This approach is ok up to $O(10)$ GHz, beyond which we run into the **single quantum limit**.

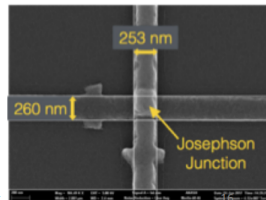
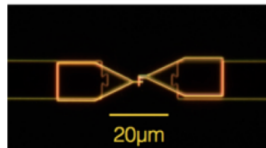
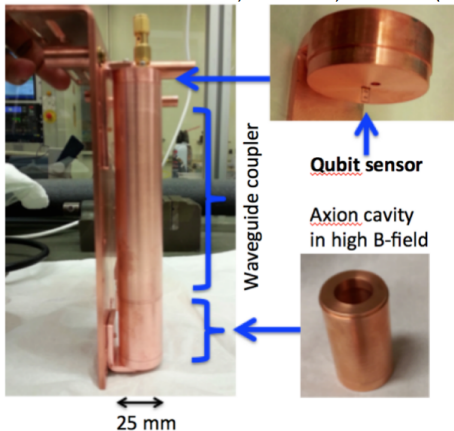
We need qubits to beat noise-limited linear amplifiers.



- Linear amplifiers suffer from irreducible zero-point noise (Caves, Phys. Rev. D, 1982).
- Quantum nondemolition measurements possible w/ qubits (Schuster *et al.*, Nature, 2007)

10 GHz band: QND detector with artificial atom based on 3D superconducting qubit

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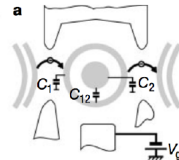


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Aaron S. Chou, Fermilab Scientists Retreat

Possible single photon detectors in the 100 GHz band

- **Tiny low heat capacitance TES bolometers** placed inside cavities, reducing threshold from 1 eV down to 10^{-3} eV
 - Unlike WIMP detectors, we do not need large volume since we want quantum efficiency to be tiny to maintain high cavity quality factor.
 - Unlike CMB detectors we do not need to sink 10 pW DC power
 - Possible collaboration with Flavio Gatti (Genova) funded by Marie Curie staff exchange grant
- QND or direct photon absorption using artificial atoms implemented using **quantum dots**
 - Single photon absorption induces transition between Landau levels
 - Transition frequencies are continuously tunable via external magnetic field
 - Single electron transistors become single photon detectors



Komiyama, et.al, Nature (2000)

$$200 \text{ MHz} \lesssim m_a \lesssim 200 \text{ GHz}$$

Our LDRD supports a mK-class test stand.

- He³-He⁴ dilution refrigerator
 - Integrated $\mathcal{O}(10)$ T solenoid
 - Siting is in progress.
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- We are currently searching at $\mathcal{O}(1)$ GHz.
 - Our QC LDRD effort gives us access to $\mathcal{O}(10)$ GHz band.
 - We have the beginnings of a plan to reach 100 GHz.
 - All this work can be accommodated by our test stand.