

K. ZUREK

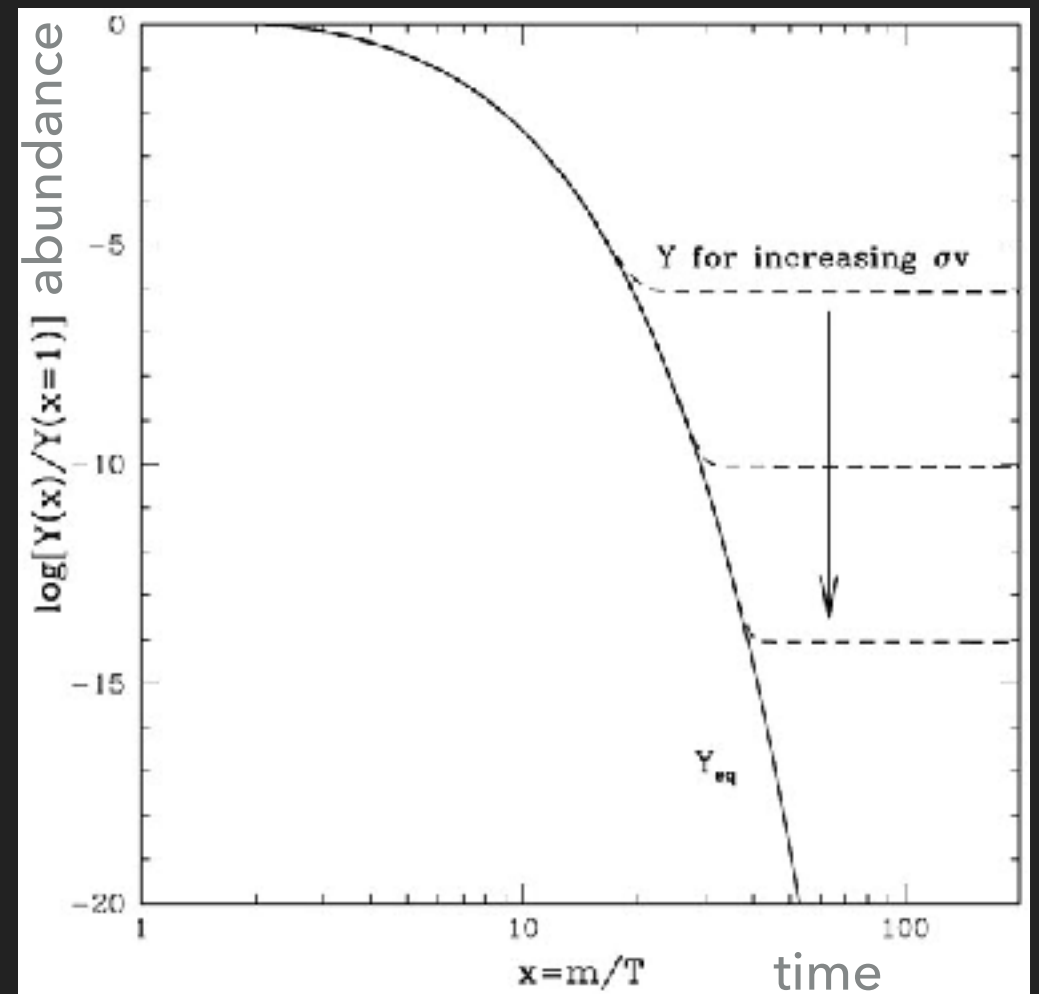
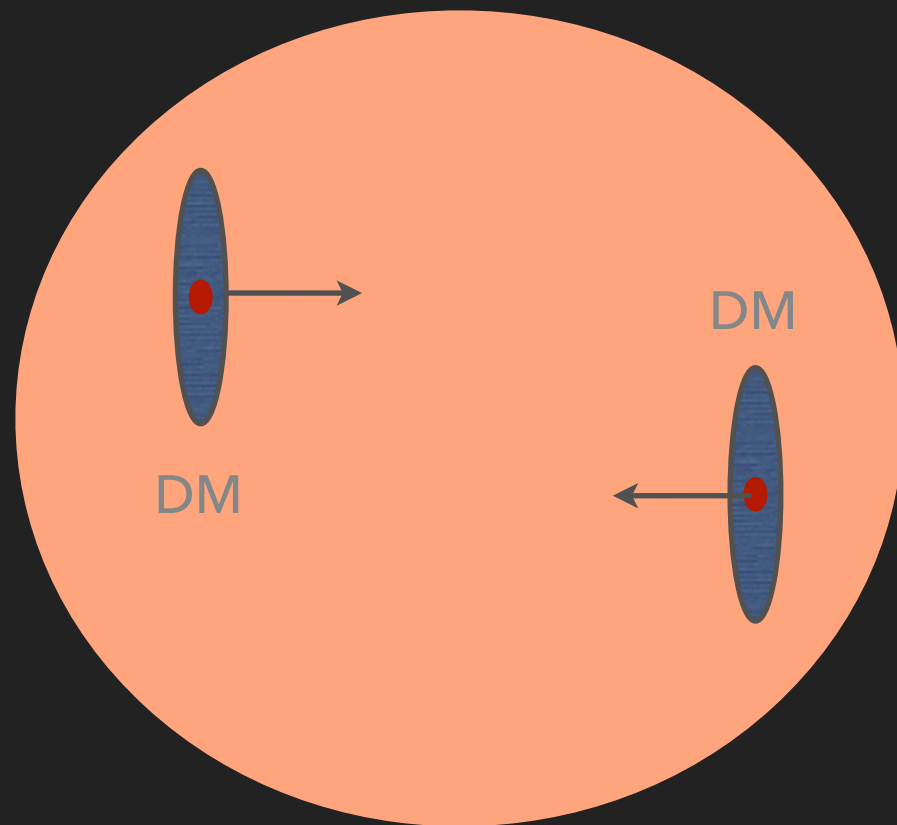
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Leveraging the many faces (and phases) of matter

# QUANTUM MATERIALS & DARK MATTER DETECTION

# NEW DIRECTIONS IN DARK MATTER THEORY

- ▶ Old paradigm: weak scale dark matter (with relic density fixed by freeze-out)



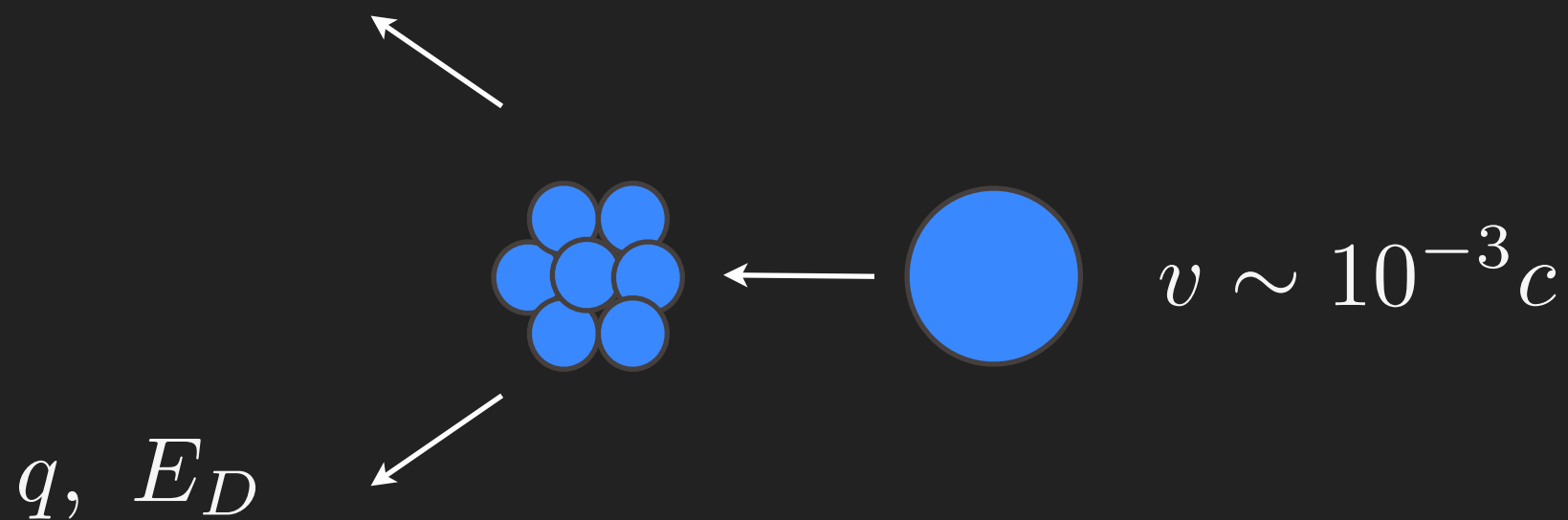
Kolb and Turner

$$n\langle\sigma v\rangle = H(T_{fo})$$

$$\implies \langle\sigma v\rangle \simeq \frac{1}{(20 \text{ TeV})^2} \simeq \frac{g_{wk}^4}{4\pi(2 \text{ TeV})^2}$$

## DIRECT DETECTION GOLD STANDARD

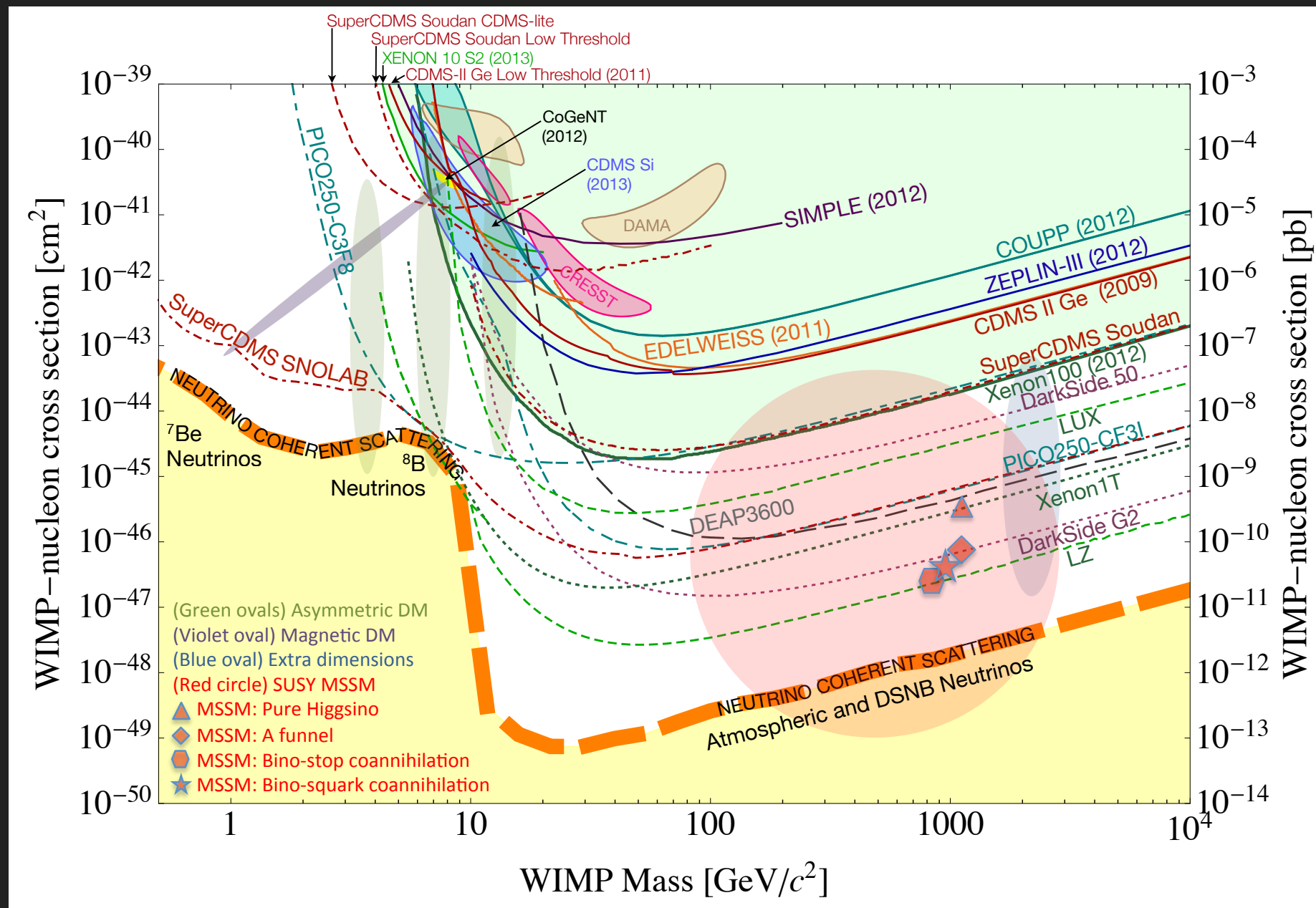
- ▶ Nuclear recoil experiments; basis of enormous progress in direct detection



$$\implies 2\mu_N v = q_{\max} = \sqrt{2m_N E_D} \quad \mu_N \equiv \frac{m_N m_X}{m_X + m_N}$$

$$v \sim 300 \text{ km/s} \sim 10^{-3}c \implies E_D \sim 100 \text{ keV} \quad \text{for 50 GeV target}$$

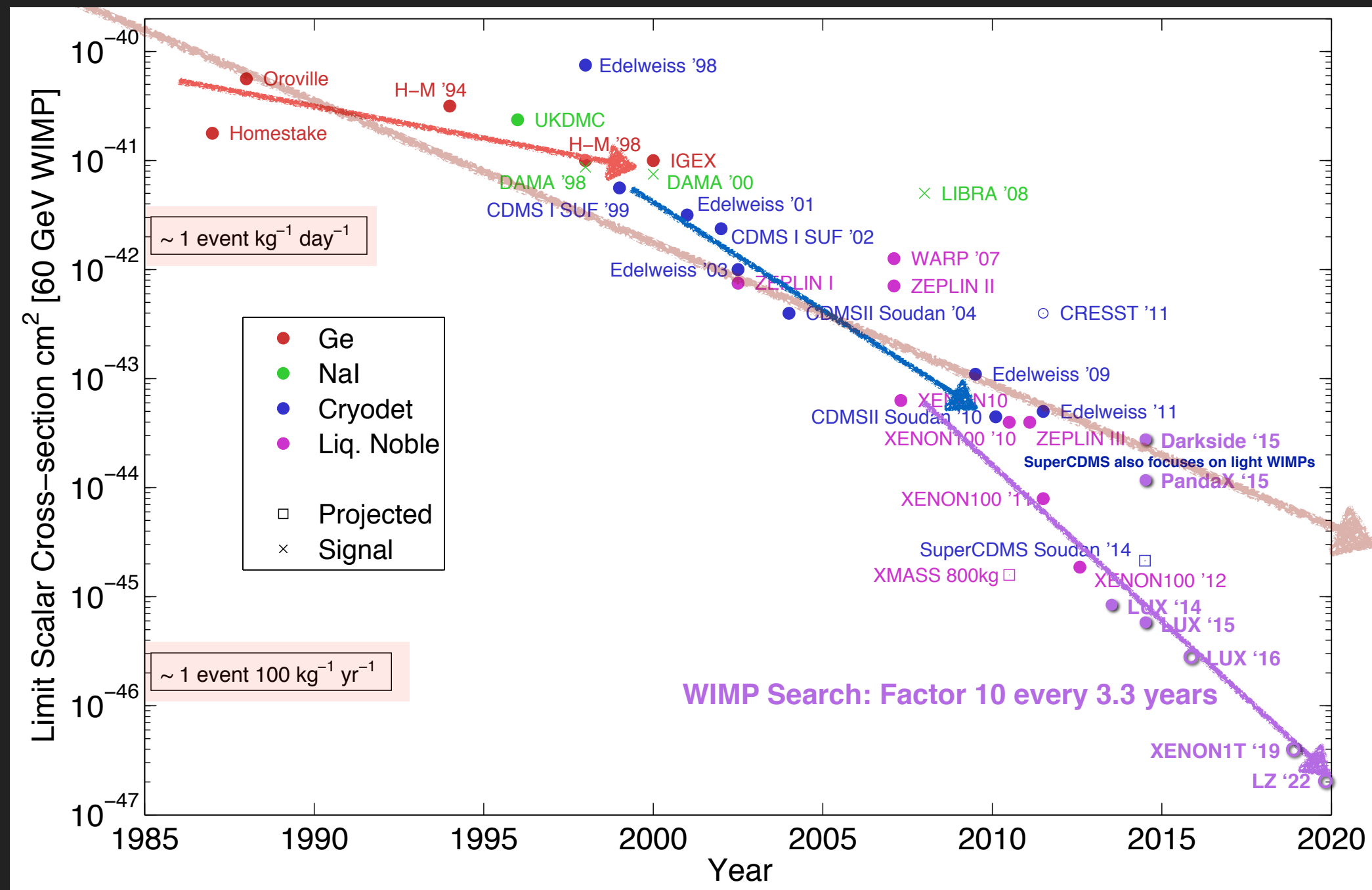
# DARK MATTER AND CLASSICAL BILLIARD BALLS



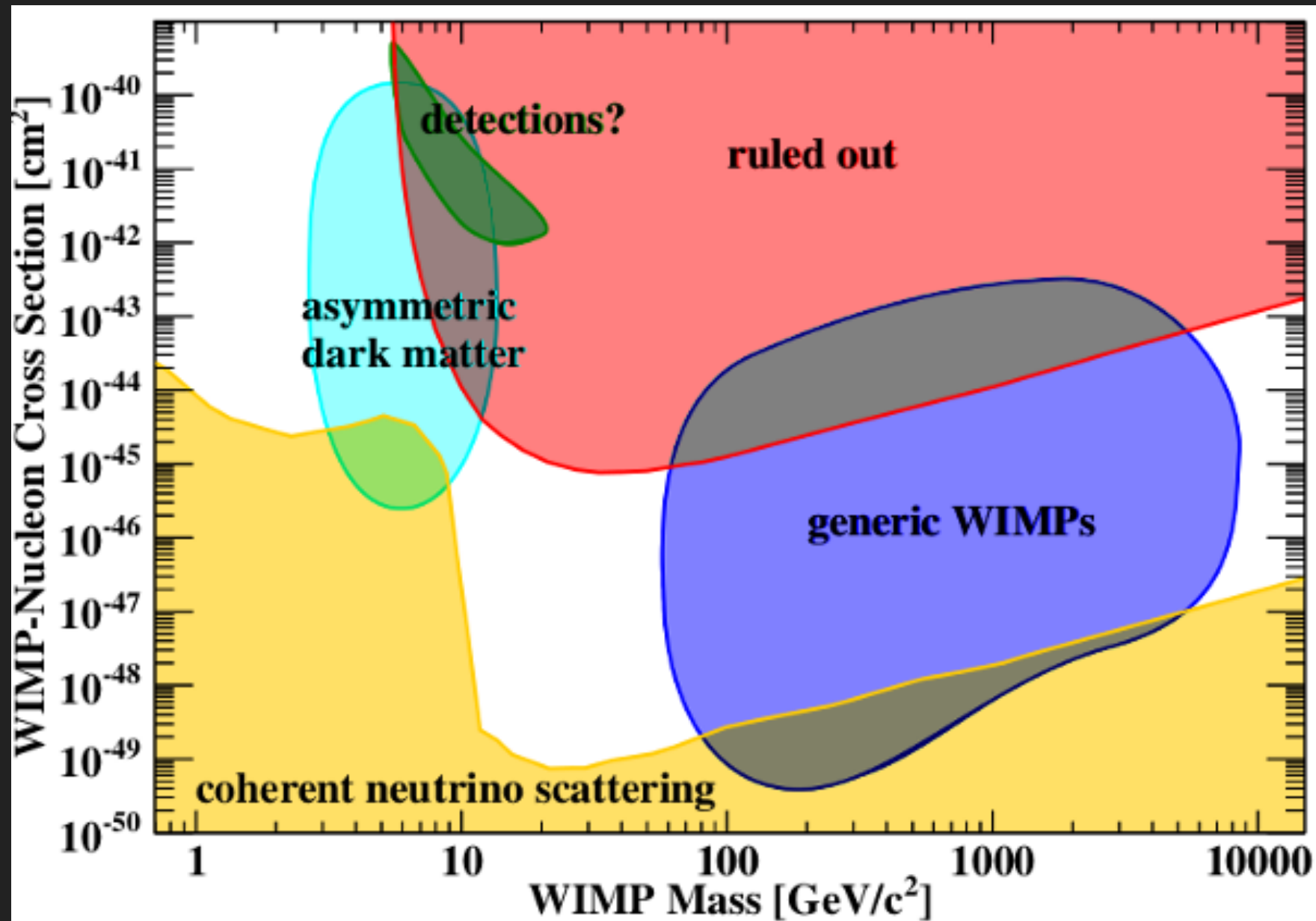


# DARK MATTER MOORE'S LAW

Factor of 10 every 6.5 years

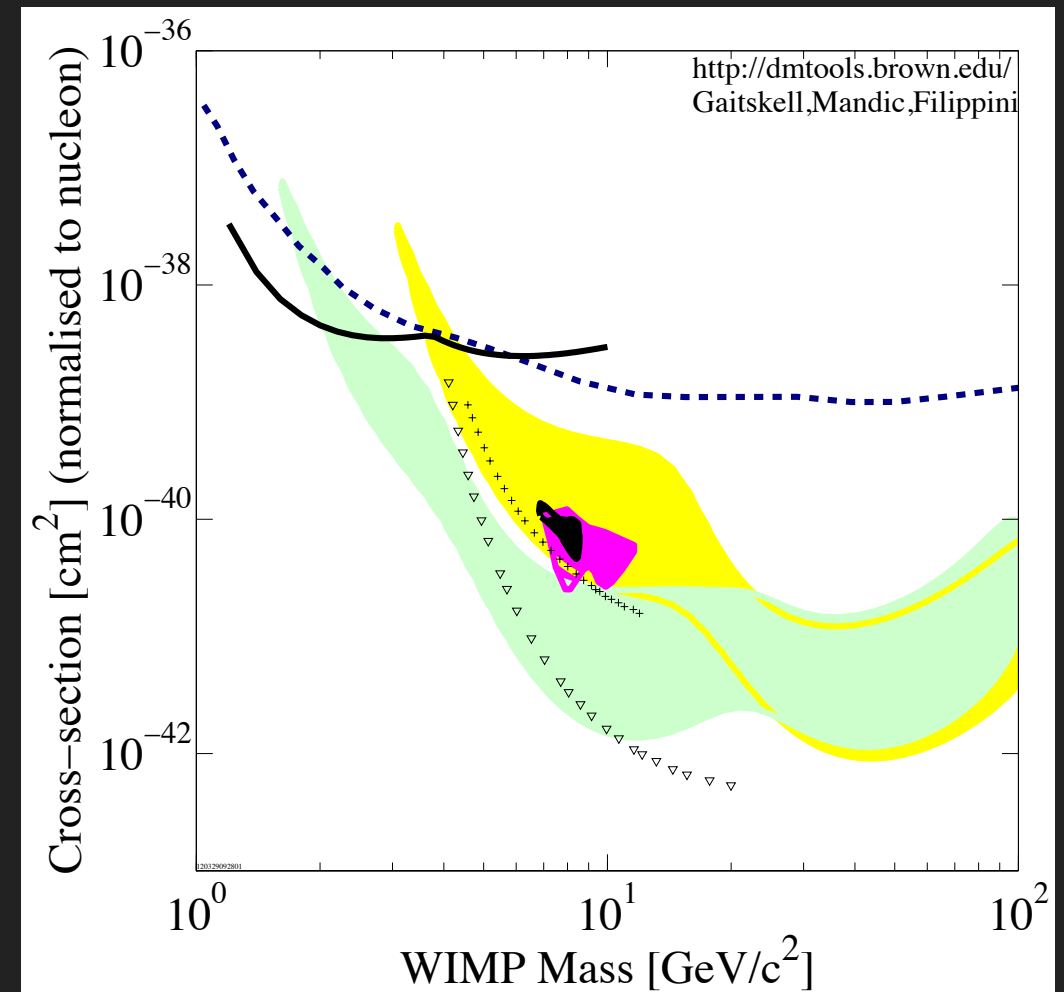


# THEORY TARGETS



## EARLY EFFORT : DAMIC CCD AT FNAL

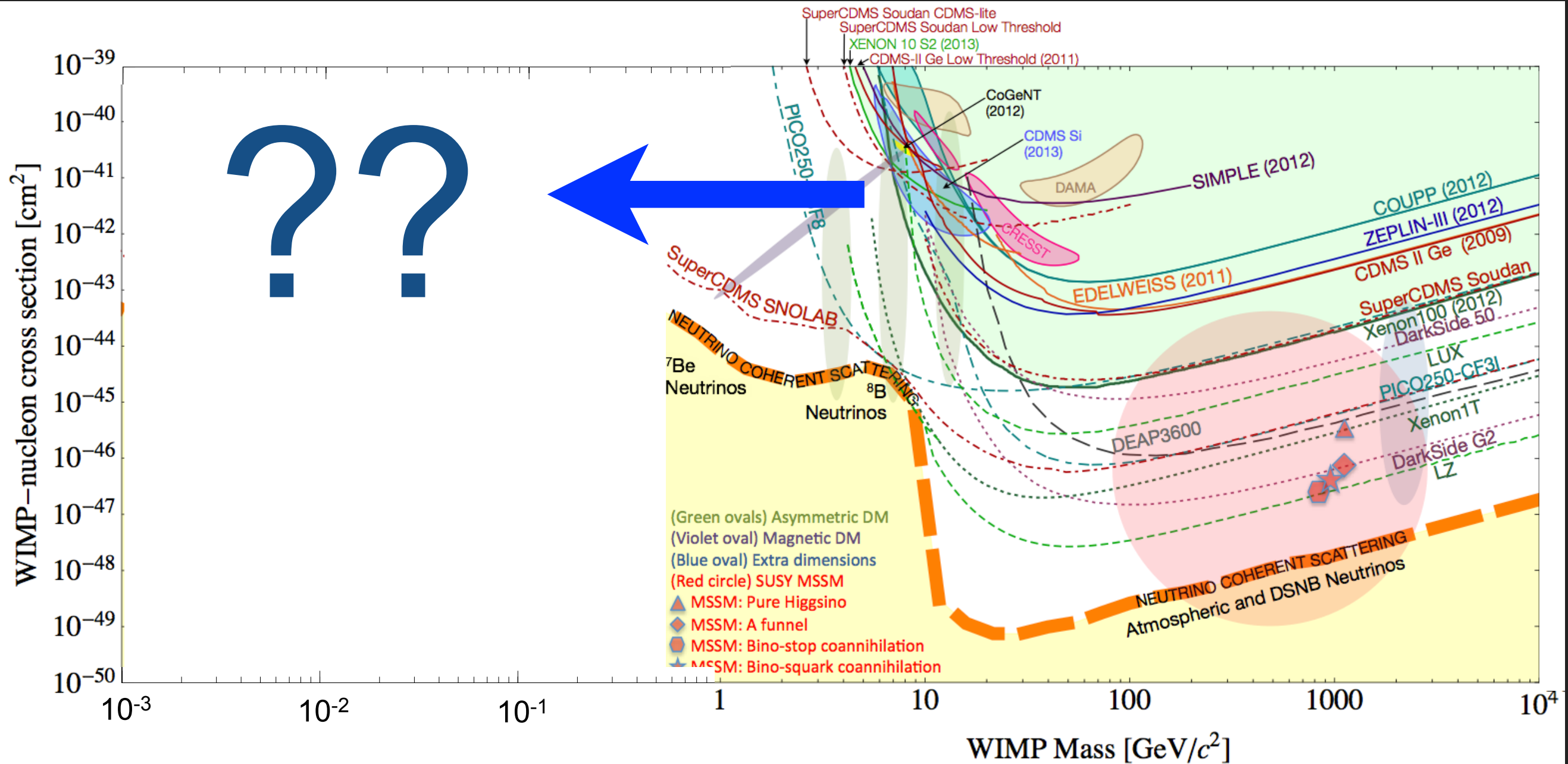
- ▶ Detecting DM Whispers dependent on dark counts and read-out noise
- ▶ 40 eV threshold, nuclear recoils



DAMIC collaboration, 1105.5191

## DIRECT DETECTION GOLD STANDARD

??



# TOWARDS LIGHT DARK MATTER

Dark Matter May Reside in a Hidden Sector

Standard Model



Connector



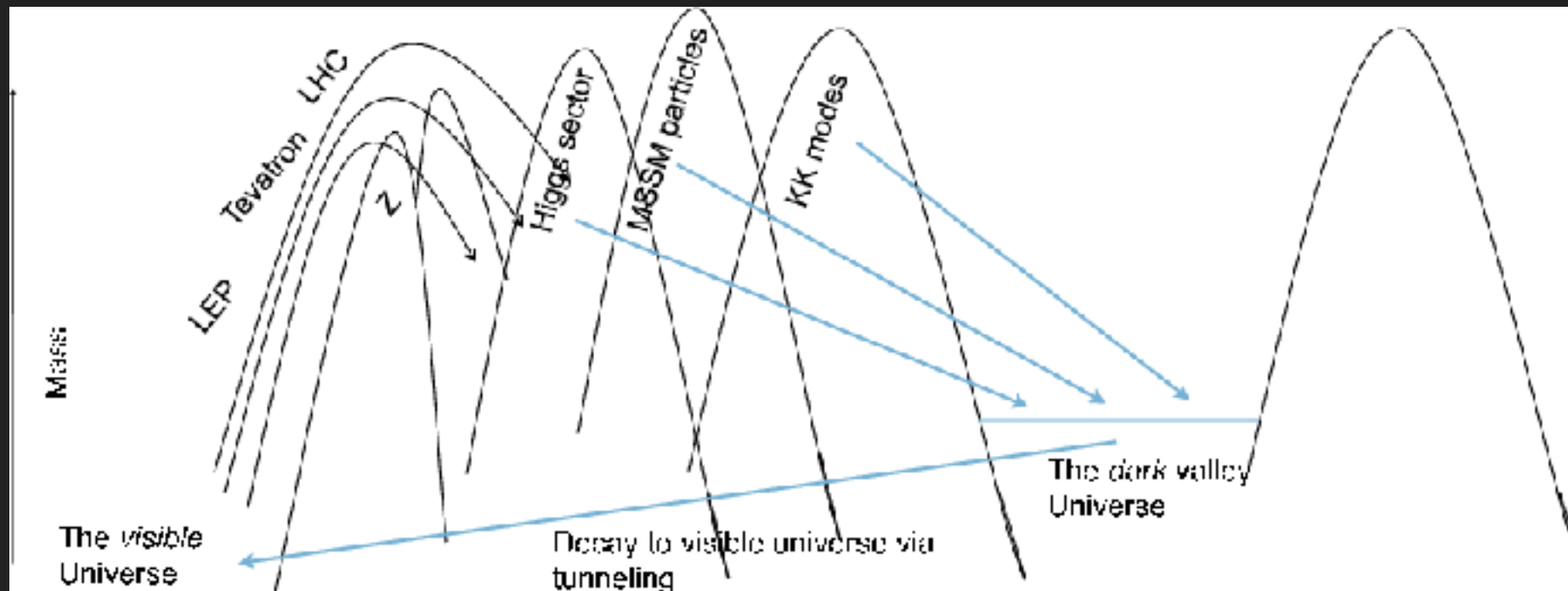
Dark Matter

e.g. a stable dark pion

no weak force

$$\pi_v^+ \pi_v^- \rightarrow \pi_v^0 \pi_v^0$$

$$\pi_v^0 \rightarrow b\bar{b}, \gamma\gamma$$



## NUCLEAR RECOILS

- Kinematic penalty when DM mass drops below nucleus mass

$$E_D = \frac{q^2}{2m_N} \quad q_{\max} = 2m_X v$$



$$E_D \gtrsim \text{eV} \Leftrightarrow m_X = 300 \text{ MeV}$$

even though  $E_{\text{kin}} \gtrsim 300 \text{ eV}$

### NEXT UP: ELECTRON

- ▶ More bang for the buck if DM lighter than 1 GeV

$$E_D = \frac{q^2}{2m_e} \qquad q_{\max} = 2m_X v$$

- ▶ Allows to extract all of DM kinetic energy for DM MeV and heavier

$$E_D \gtrsim \text{eV} \leftrightarrow m_X = 1 \text{ MeV}$$

## ELECTRONS IN MATERIALS

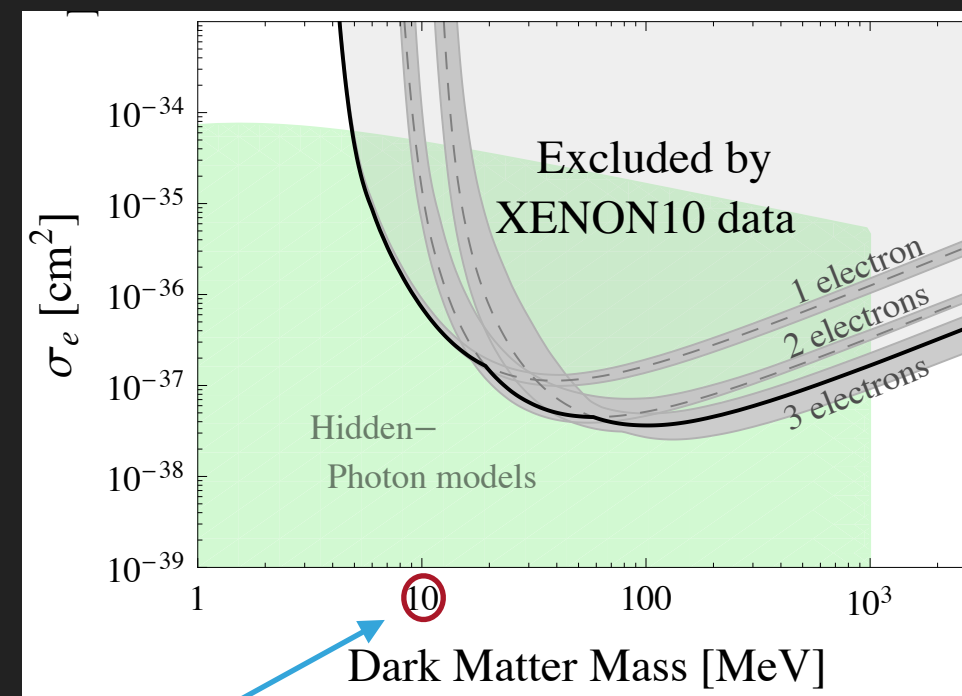
- In insulators, like xenon

Ionize electron

- In semi-conductors, like Ge, Si

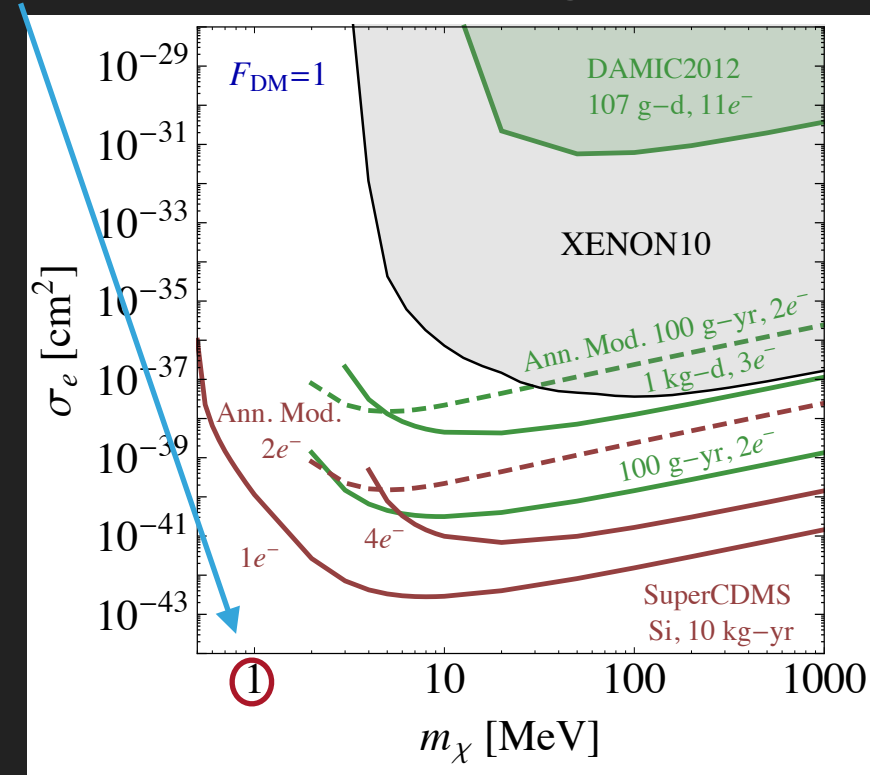
Excite electron to conduction band

P. Sorensen et al 1206.2644



Gap = DM Kinetic Energy

Essig et al 1509.01598

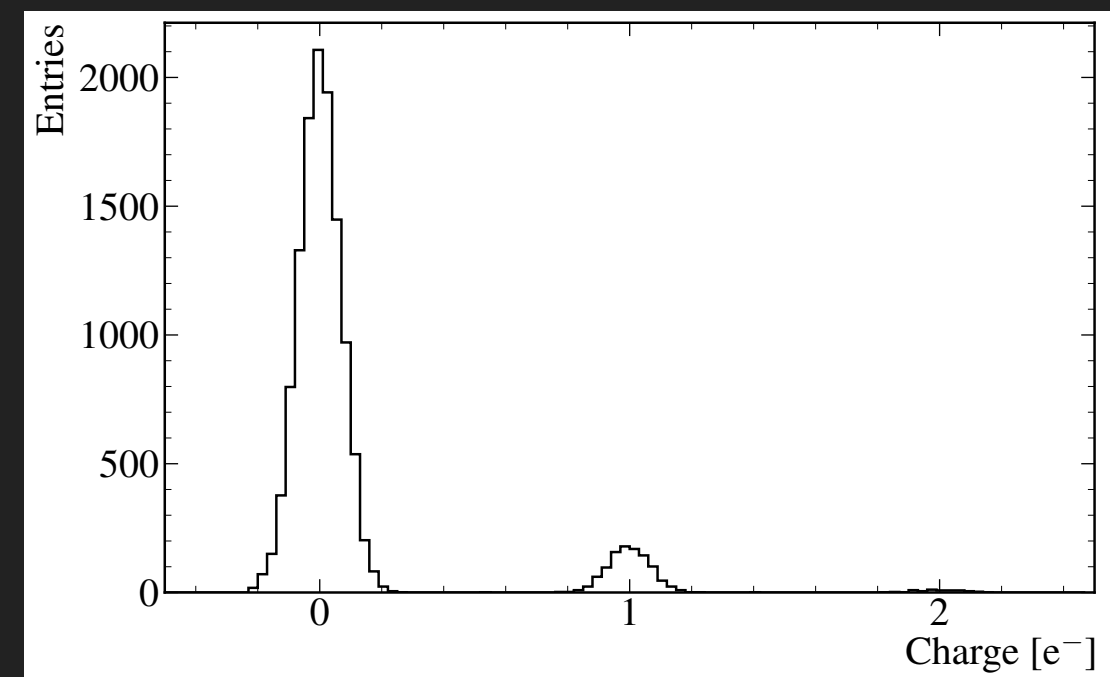




# SENSEI AND SKIPPER CCD'S

- ▶ DAMIC utilized sensitivity to charge to place constraints on DM
- ▶ Fundamentally limited by noise
- ▶ More noise = less sensitivity to DM Whispers
- ▶ Improved Read-out

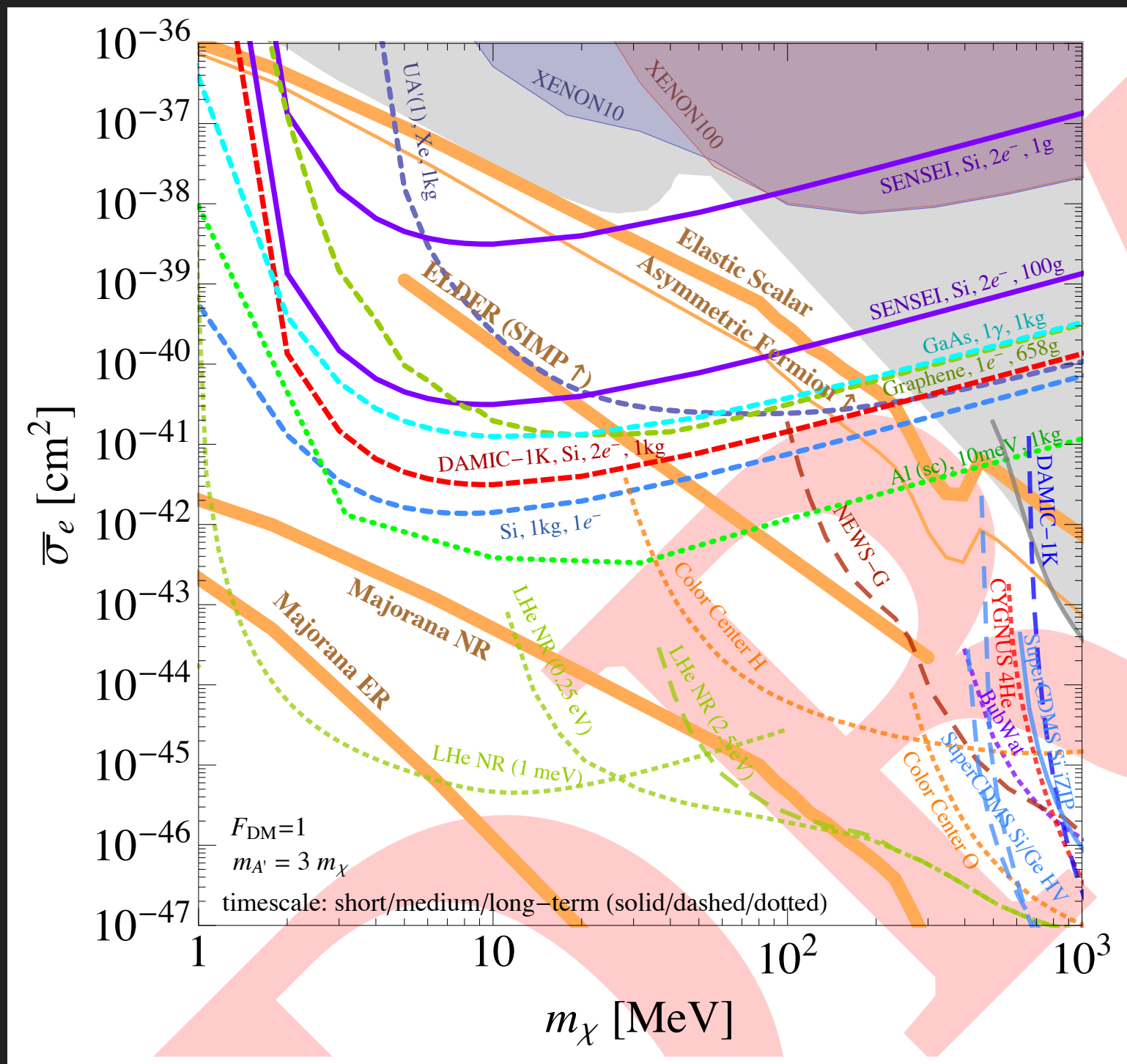
LDRD led by Javier Tiffenberg



SENSEI, 1706.00028

RMS = 0.068 e/pix

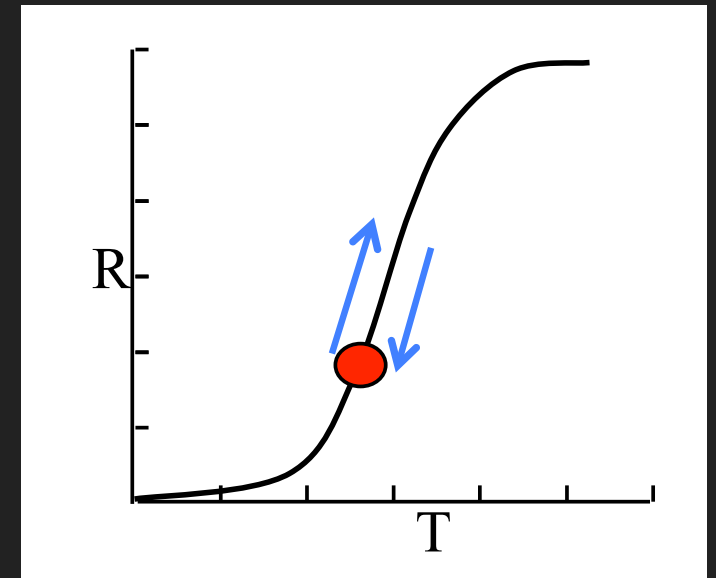
## SENSEI AND SKIPPER CCD'S



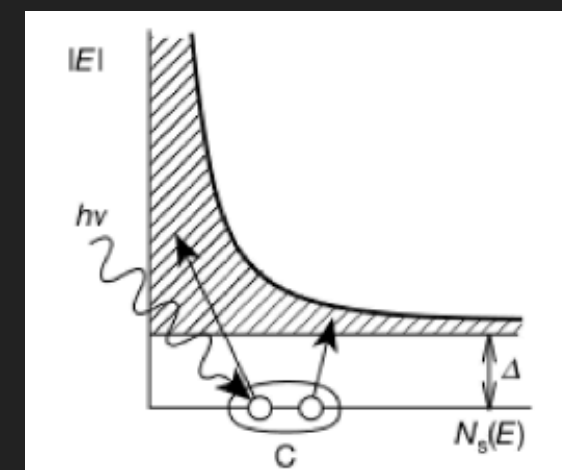
## QUANTUM DEVICE R&D

- ▶ In addition to suitable target (quantum phases of matter), need quantum devices capable of measuring small energy deposits
- ▶ Superconducting devices that measure single quanta
- ▶ Single infrared or microwave photon detectors, e.g. Aaron Chou LDRD

Transition Edge Sensor calorimeter

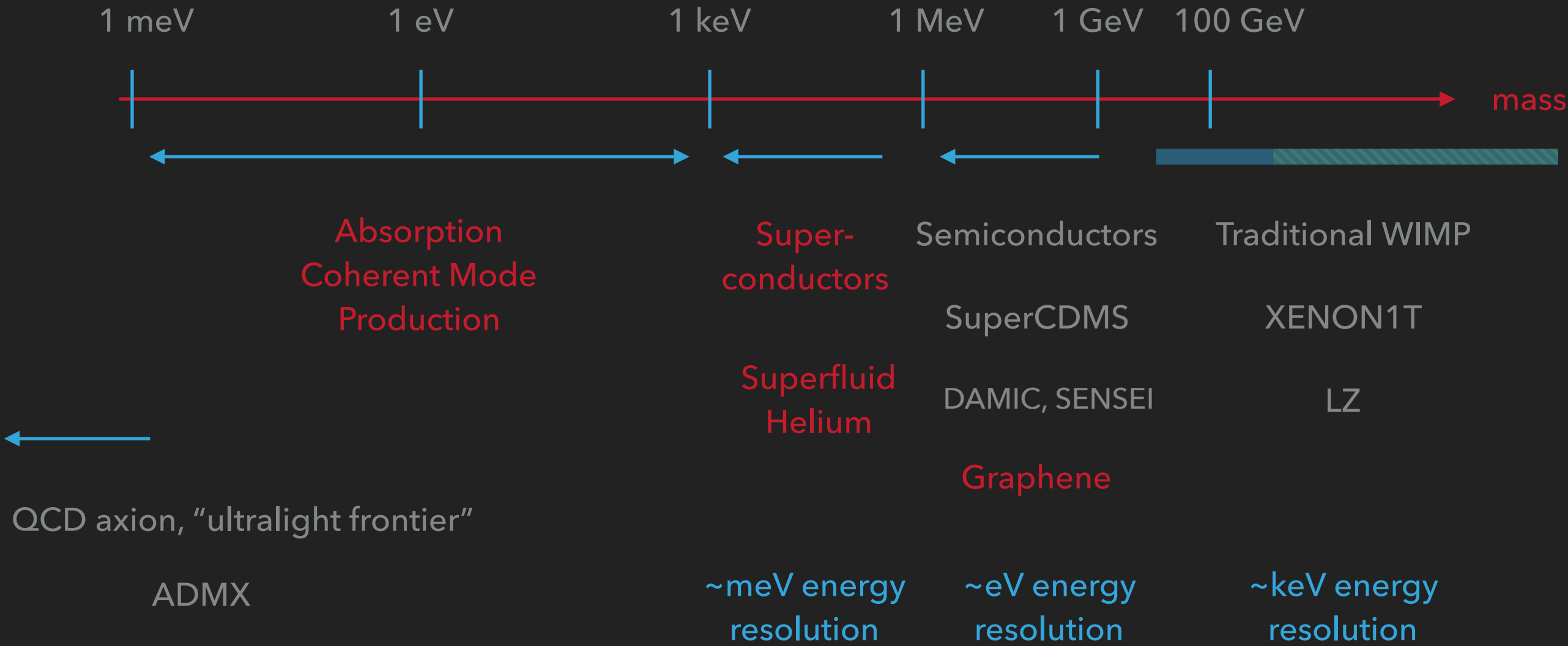


Microwave Kinetic Inductance Device



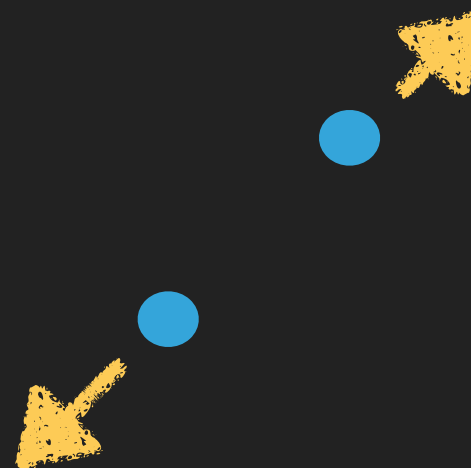
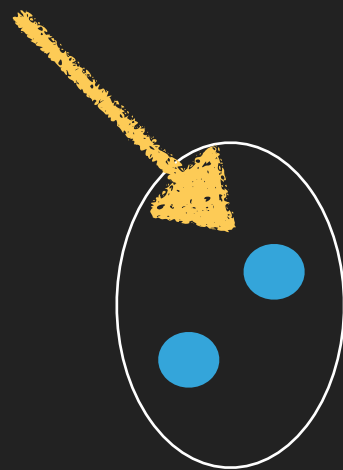
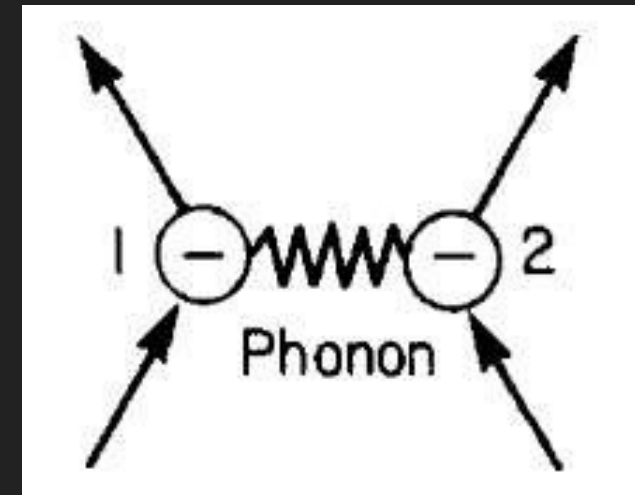
See W. Wester talk

# DARK MATTER LANDSCAPE



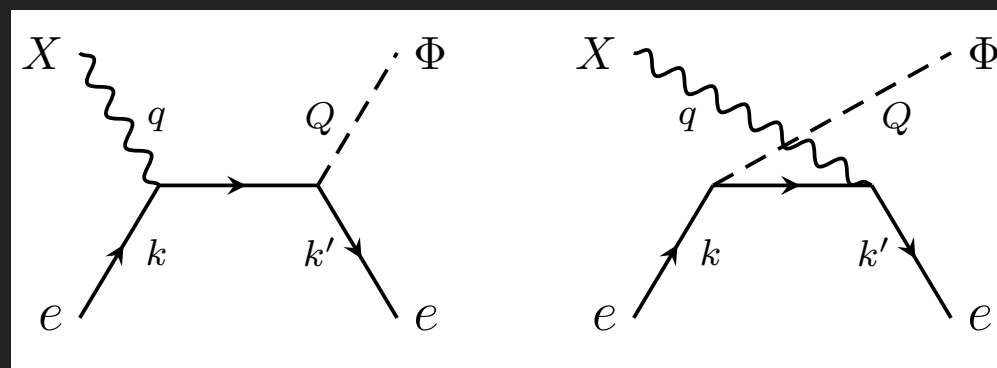
## E.G. SUPERCONDUCTORS

- ▶ Free electrons succumb to collective dynamics
- ▶ Typical gap  $\Delta \simeq 0.3 \text{ meV}$



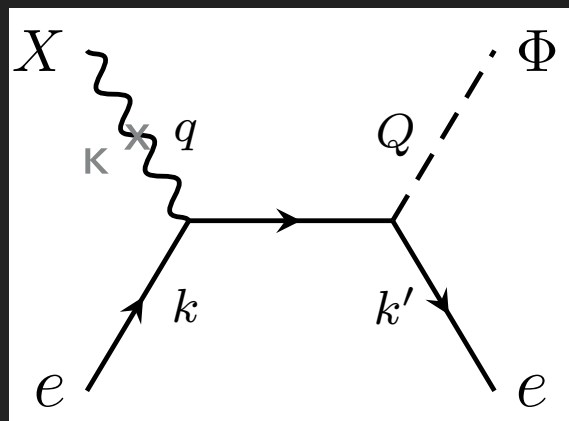
## ABSORPTION — SUPERCONDUCTORS

- ▶ Can we absorb ultralight DM particles on electrons in a superconductor?
- ▶ Seems not – basic energy and momentum conservation
- ▶ Take advantage of collective modes! i.e. phonons

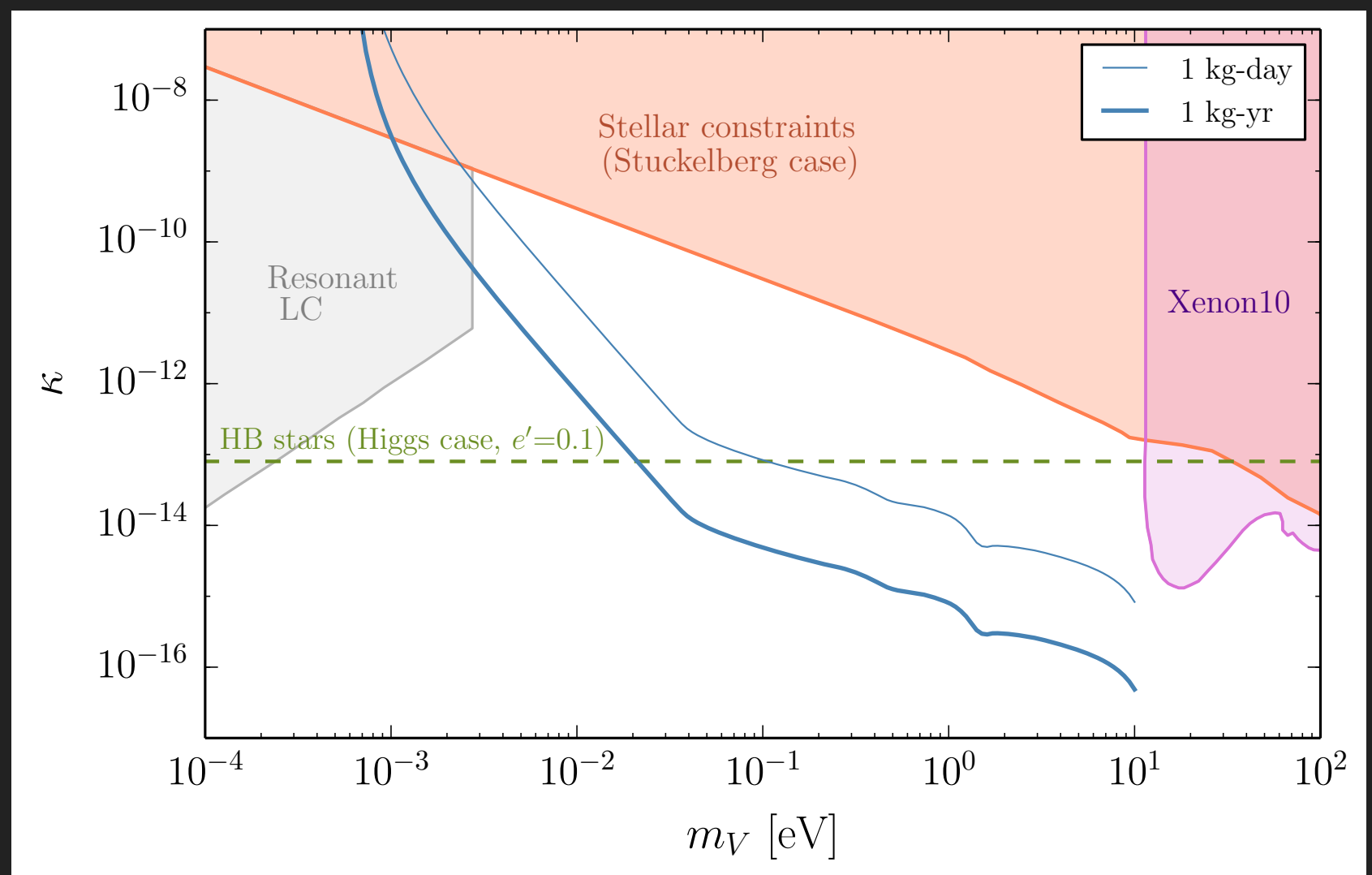


$$\mathcal{H} = \int d^3y_{ph} \phi \bar{\psi} \psi = \frac{1}{\sqrt{V}} \sum_{\vec{k}} \sum_{\vec{k}'} \frac{C_{ph} |\vec{Q}|}{\sqrt{\rho}} \frac{1}{\sqrt{2E_Q}} (c_{\vec{Q}} + c_{-\vec{Q}}^\dagger) a_{\vec{k}'} a_{\vec{k}}$$

# ABSORPTION — SUPERCONDUCTORS

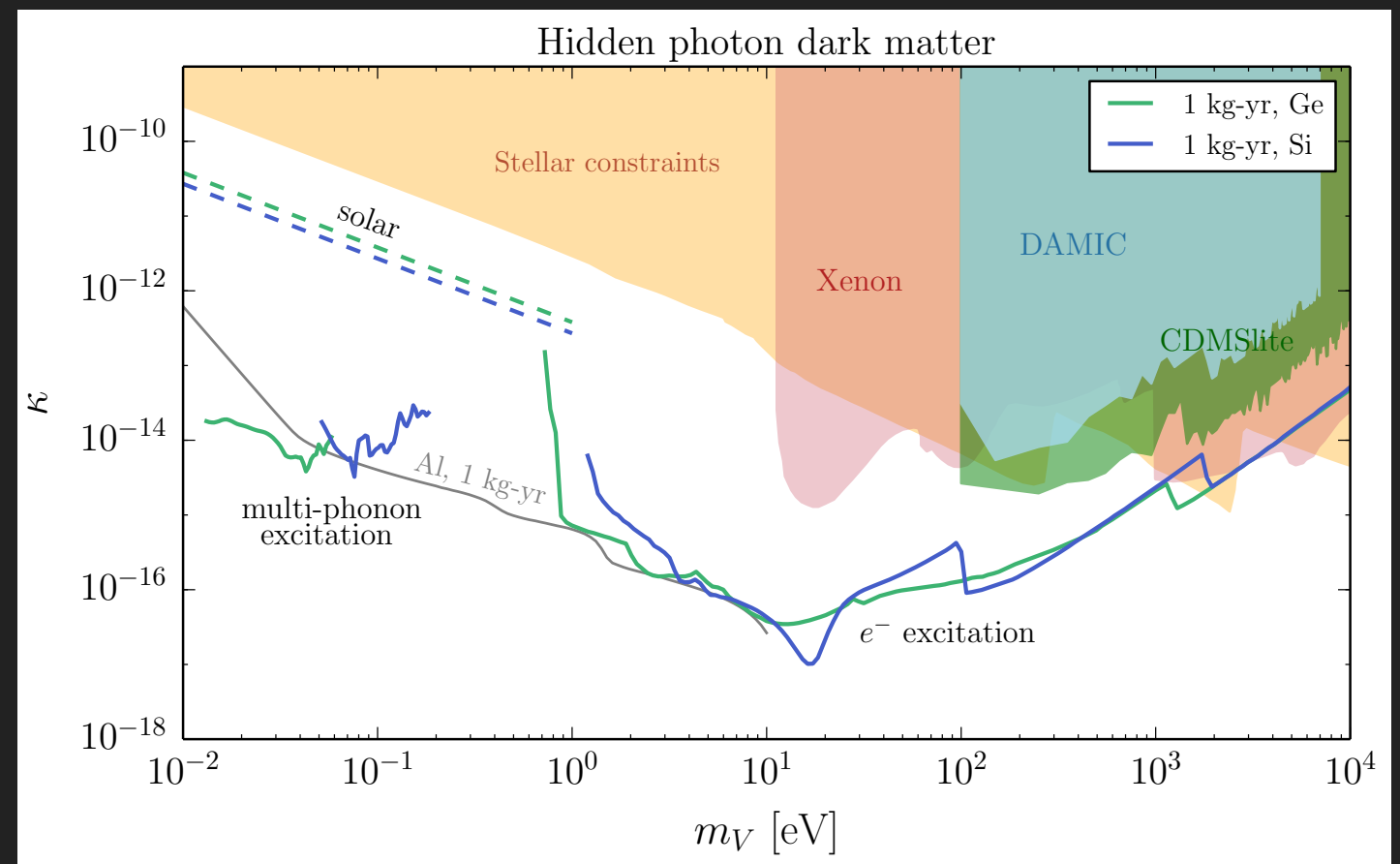
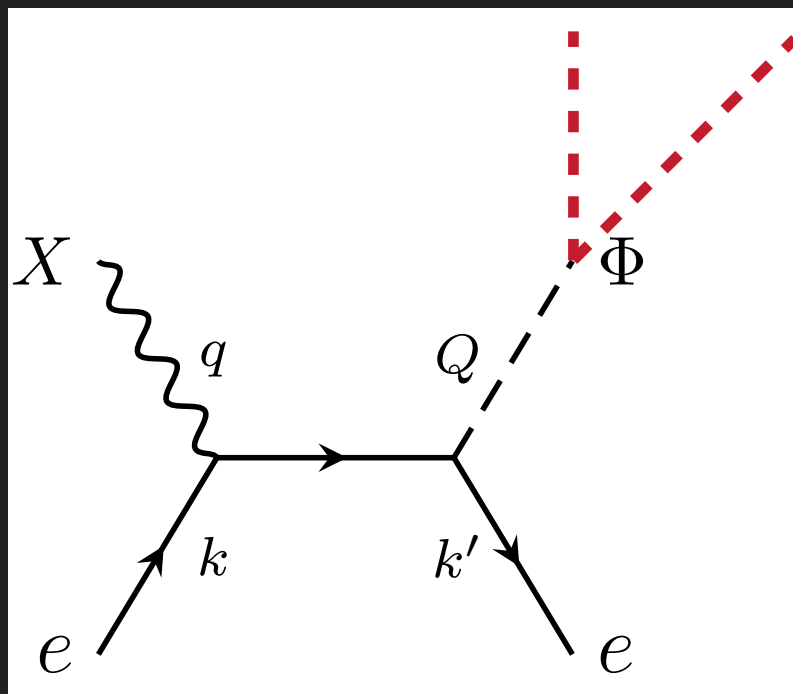


Dark Photon



## ABSORPTION — SEMICONDUCTORS

- Larger gap means sensitivity only to heavier particles ... but, there is a new process!





## HELIUM

- ▶ Superfluids are naturally insensitive to noise. A good light DM detector? In the context of ordinary nuclear recoils, yes, see e.g. 1605.00694
- ▶ To detect lighter DM, couple to phonon modes.
- ▶ Viable? At first glance – no

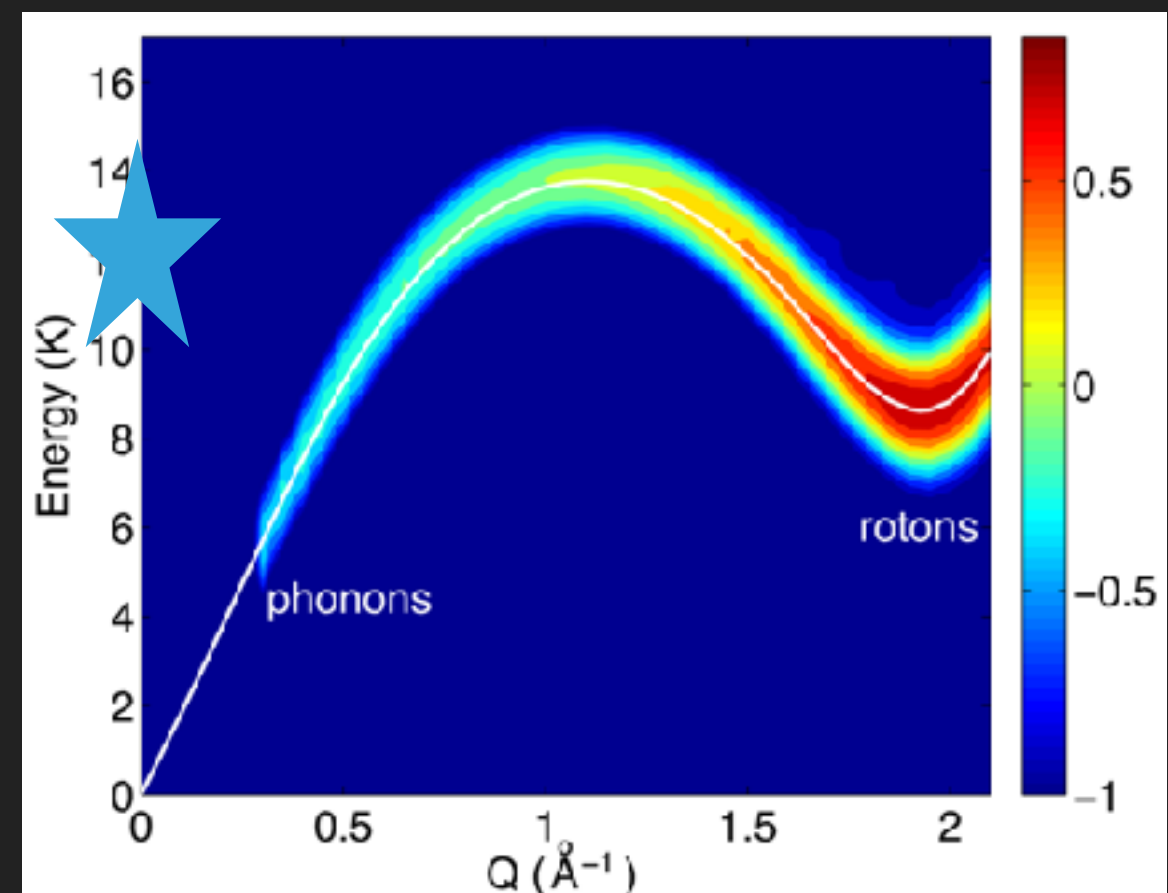
$$E_D \sim v_X q$$

vs

$$c_s \ll v_X$$

$$E_D \sim c_s q$$

- ▶ Next glance -- yes!



## HELIUM

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$$E_D \sim v_X q$$

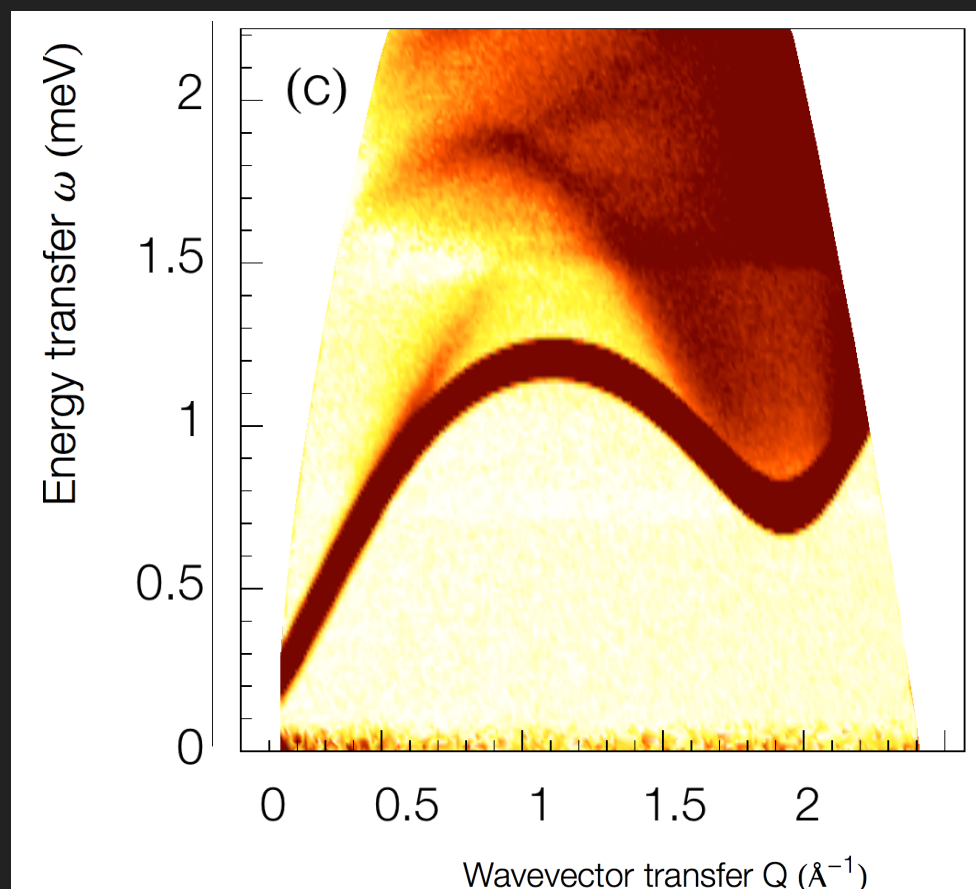
vs

$$c_s \ll v_X$$

$$E_D \sim c_s q$$

- ▶ Next glance -- yes!

Beauvois et al 1605.02638



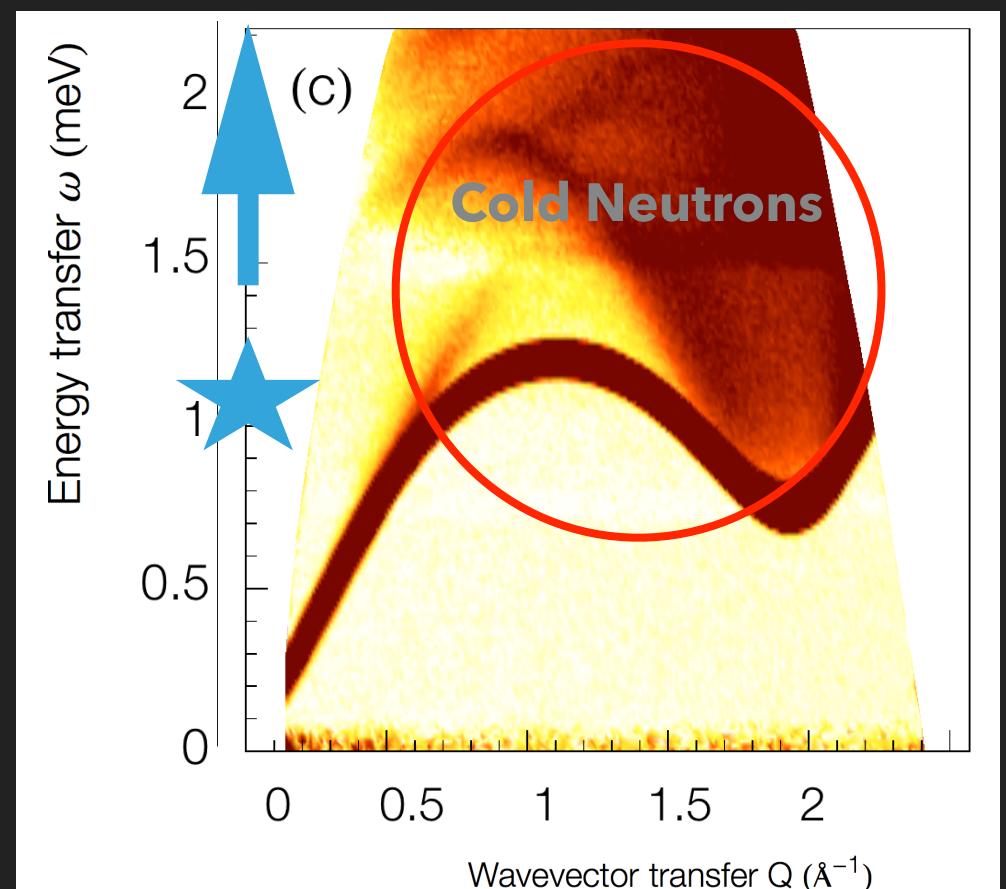
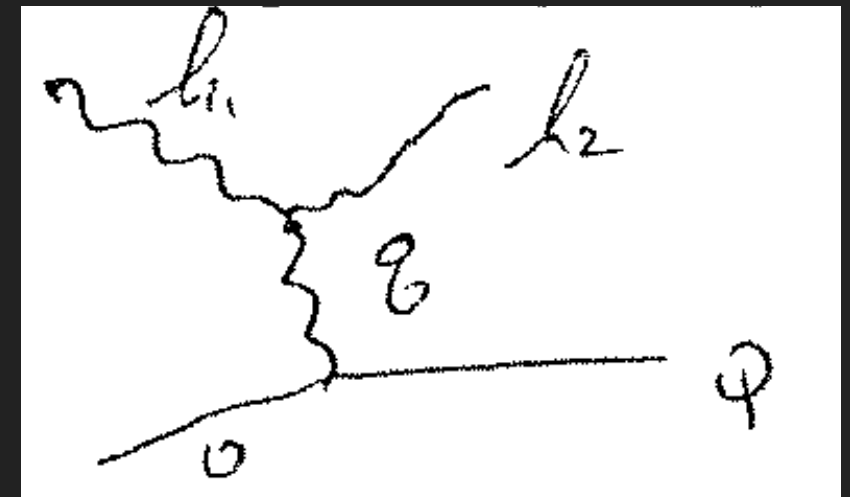
## MULTI-EXCITATIONS

- ▶ Calculated and observed for cold neutrons

$$V_3 = \int d^3r \left[ \frac{\vec{v} \cdot \vec{g}' \vec{u}_4}{2} - \frac{1}{3!} \frac{d}{d\vec{g}} \left( \frac{c^2}{\vec{g}} \right) (\vec{g}')^3 \right]$$

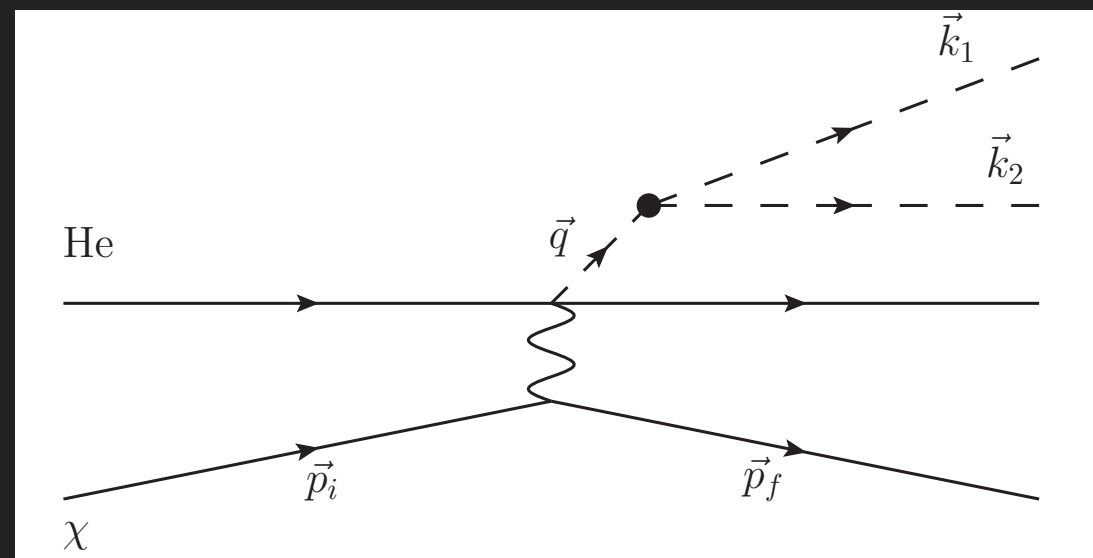
- ▶ However, this is in a very different kinematic regime
- ▶ No existing calculations in regime of interest

Internal note, R. Golub, 1977



## MULTI-EXCITATIONS

- ▶ emit back-to-back excitations to bleed off energy while conserving momentum



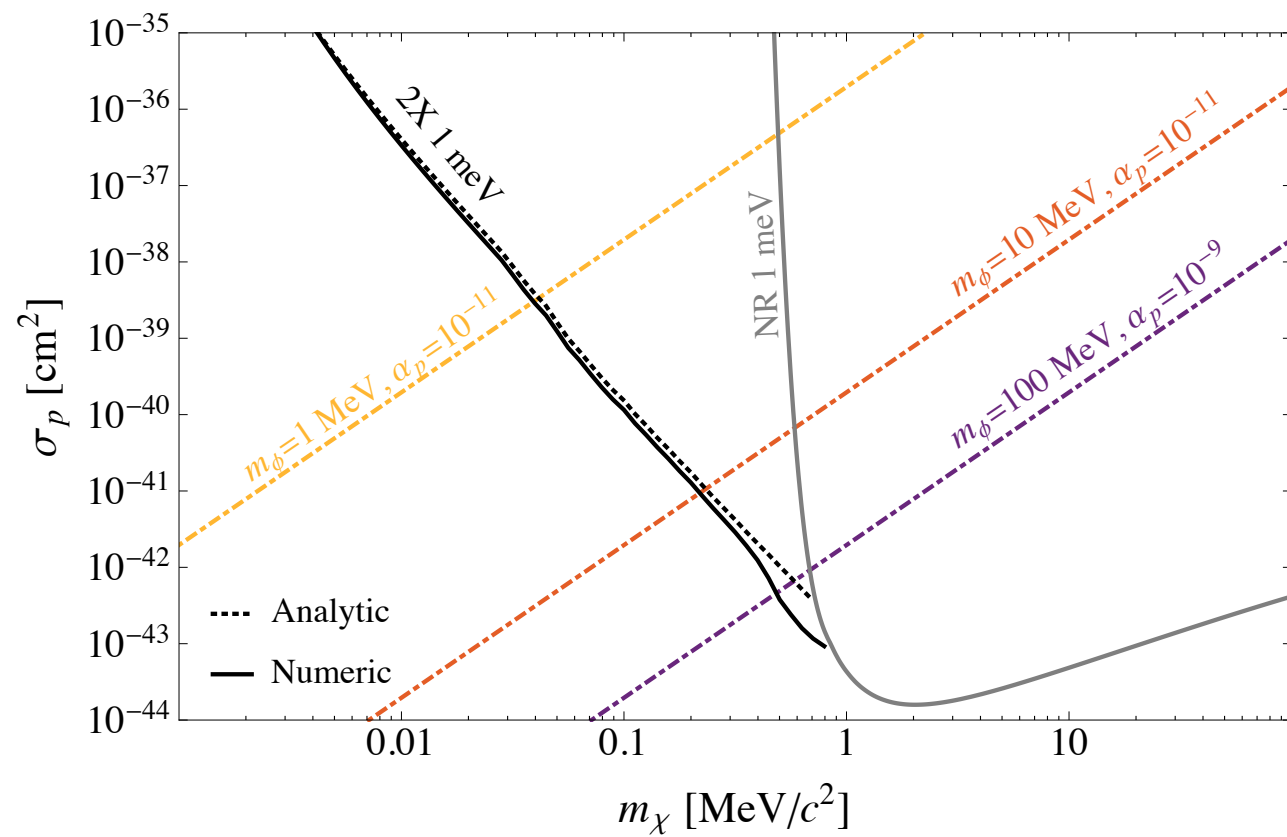
Schutz, KZ 1604.08206

- ▶ Quantize the fluid Hamiltonian, like SHO

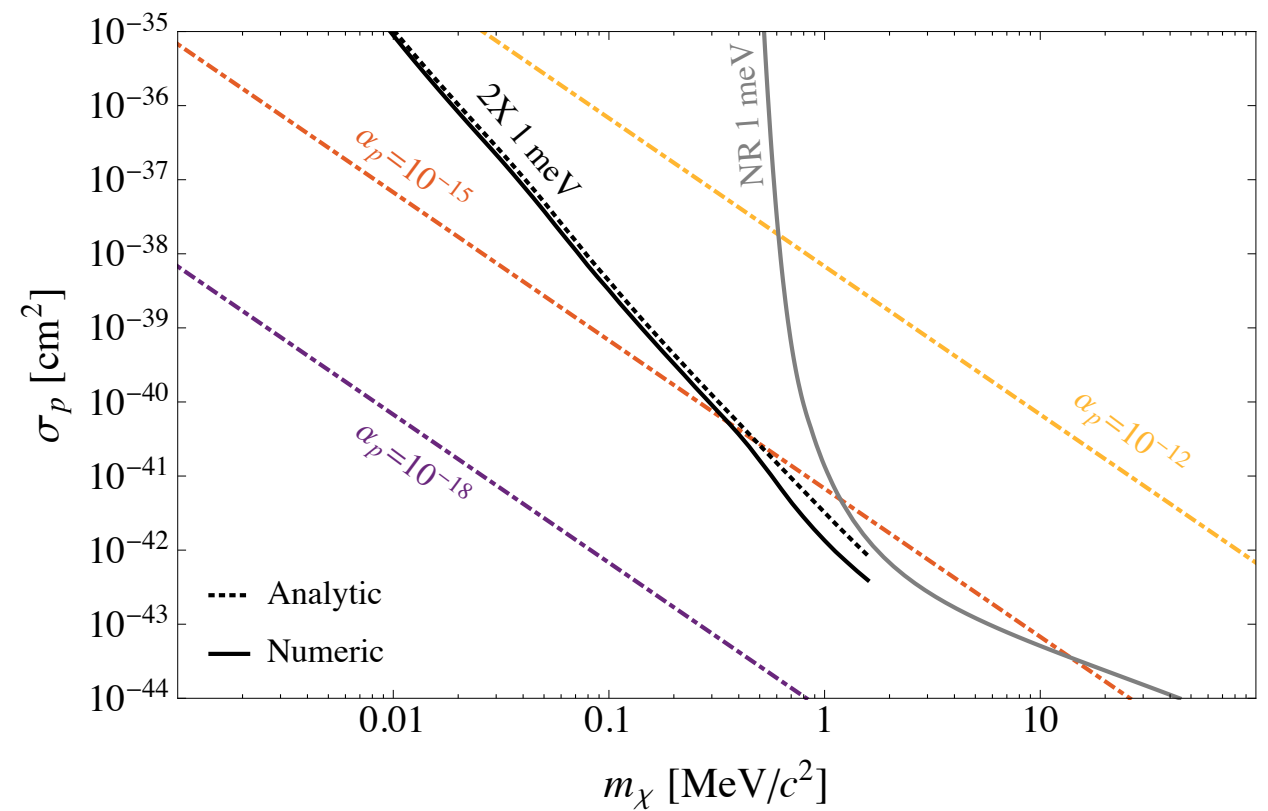
$$H_0 = \frac{1}{2} \sum_k \left( \rho_0 v_{\vec{k}} v_{-\vec{k}} + \phi(k) \rho_{\vec{k}} \rho_{-\vec{k}} \right) \quad m_{\text{He}}^2 S(k) = \langle \rho_k \rho_{-k} \rangle$$

# RESULTS

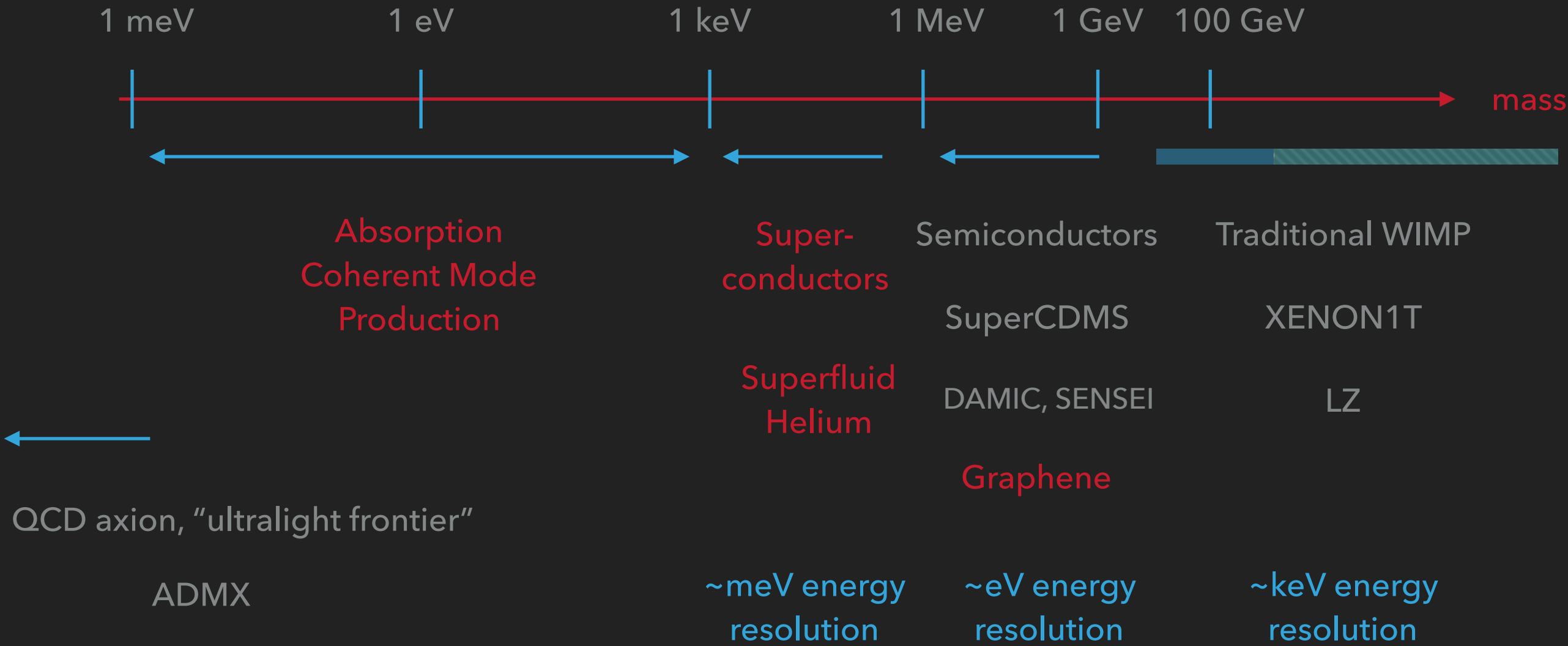
Sensitivity to DM via a Massive Mediator

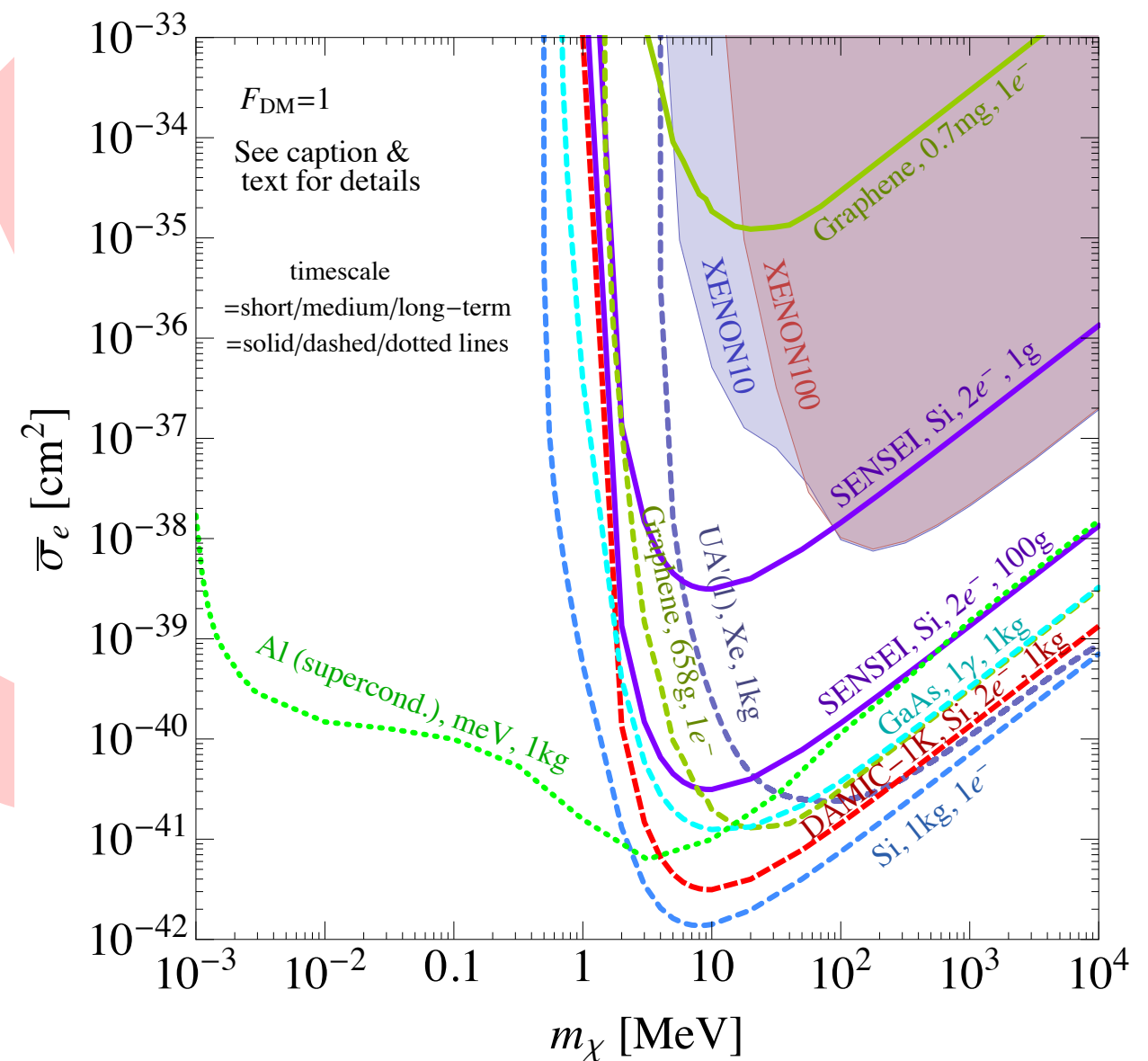


Sensitivity to DM via a Massless Mediator



# DARK MATTER LANDSCAPE

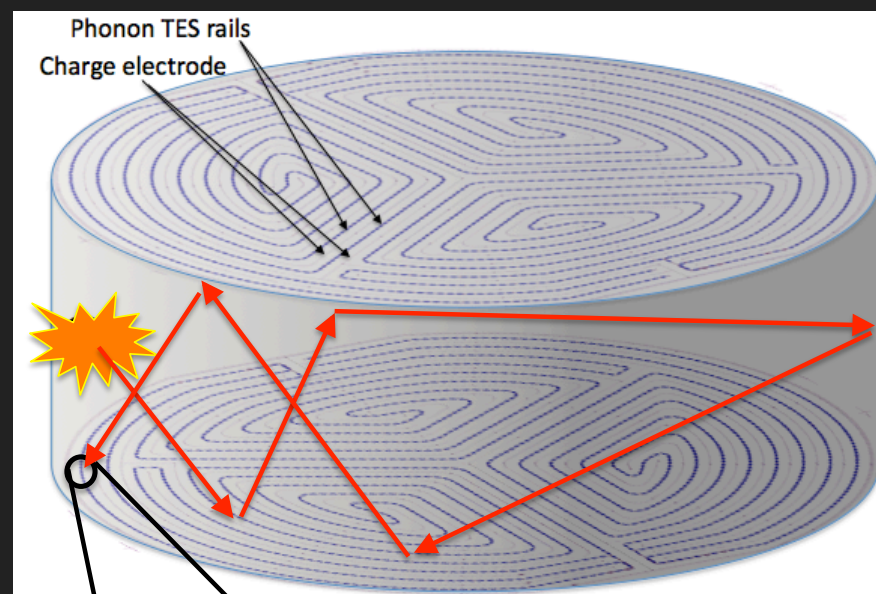






## ROAD FORWARD

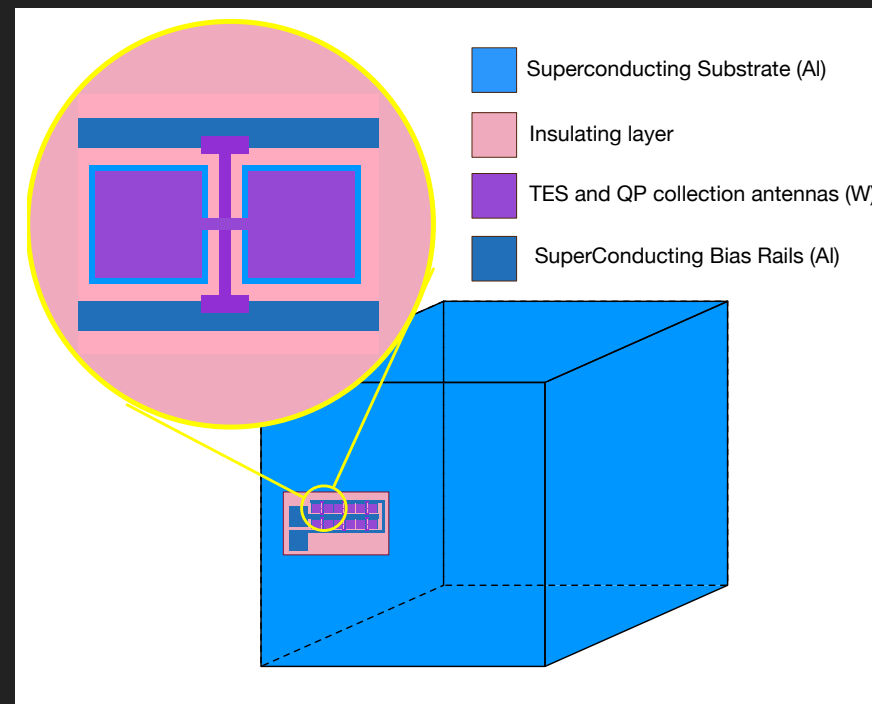
- Large part depends on better energy resolution sensors (TESs or KIDs); TESs or KIDs are portable to multiple targets



Semiconductors SuperCDMS

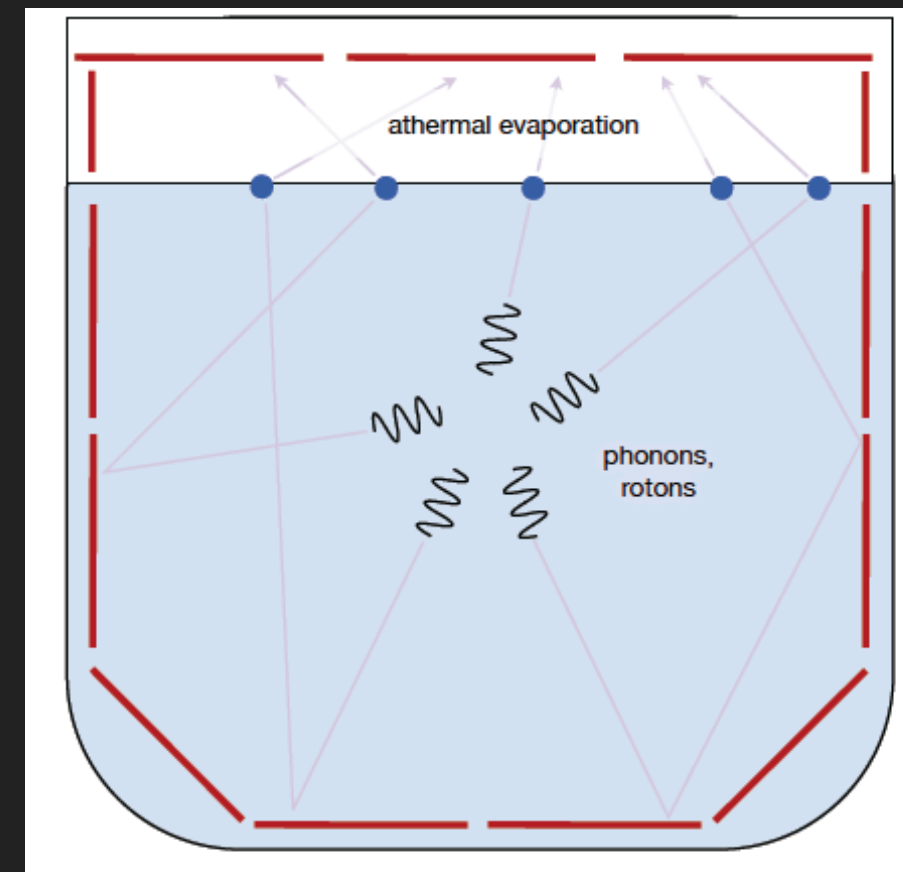
Current energy resolution:  $\sim 300$  eV

Goal:  $\sim 1$  eV



Superconductors

Goal:  $\sim 1$  meV



Superfluid Helium

Goal:  $\sim 1$  meV



# ROAD FORWARD

- ▶ New ideas for dark matter detection!
- ▶ Moving beyond nuclear recoils into phases of matter crucial to access broader areas of DM parameter space
- ▶ Target diversity essential. graphene, superconductors, semiconductors, helium ..... Weyl semi-metal
- ▶ Leverage progress in materials and condensed matter physics
- ▶ Realizing experimental program is 5-10+ years into future
- ▶ Nine orders of magnitude increased sensitivity in mass
- ▶ Long view necessary!