

# ***NDK-High-E & Sim/Reco***

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# NDK-High-E

- ❖ **Nucleon Decay**
- ❖ **Atmospheric Neutrinos**
- ❖ **Cosmogenic**

*One person's signal is another person's background*

- ❖ **Nucleon Decay**

Atmospheric Neutrinos

Cosmogenic

- ❖ **Cosmogenic**

Calibration w/ cosmic-ray muons

- ❖ **Atmospheric Neutrinos**

Atmospheric neutrinos provide a complementary analysis approach to beam neutrinos, and can help resolve ambiguities in beam-only analyses

# Proton Decay Analysis

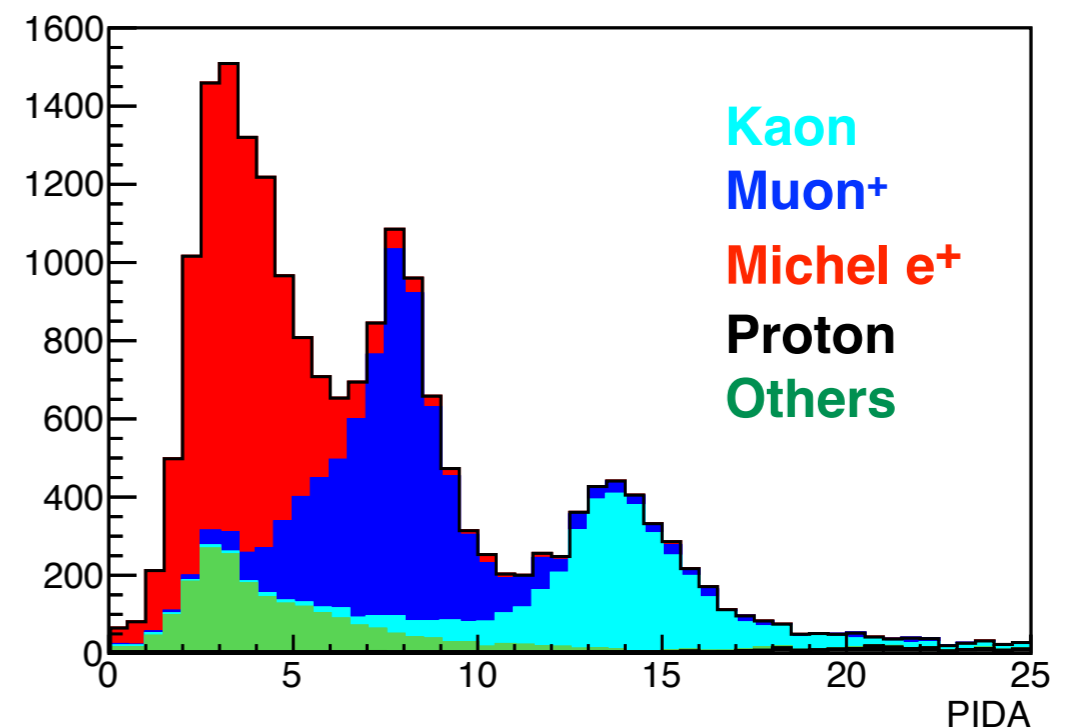
Analysis

Far Detector Task Force Report

Requirement	$p \rightarrow \bar{\nu} K_{\mu 2}^+$ signal efficiency (%)	Atmospheric $\nu$ background rate (Mton <sup>-1</sup> ·year <sup>-1</sup> )
None	100.0	$2.9 \times 10^5$
Kaon tracking efficiency	61.8	N/A
Kaon and muon ID	38.0	$9.2 \times 10^3$
Not shower-like	30.7	$1.0 \times 10^3$
Vertex-muon separation	23.2	$1.2 \times 10^2$

*Far from what we promised in CDR*

We would focus only on  $K \rightarrow \mu^+ + \nu_{\mu}$  events and atmospheric neutrino events as background

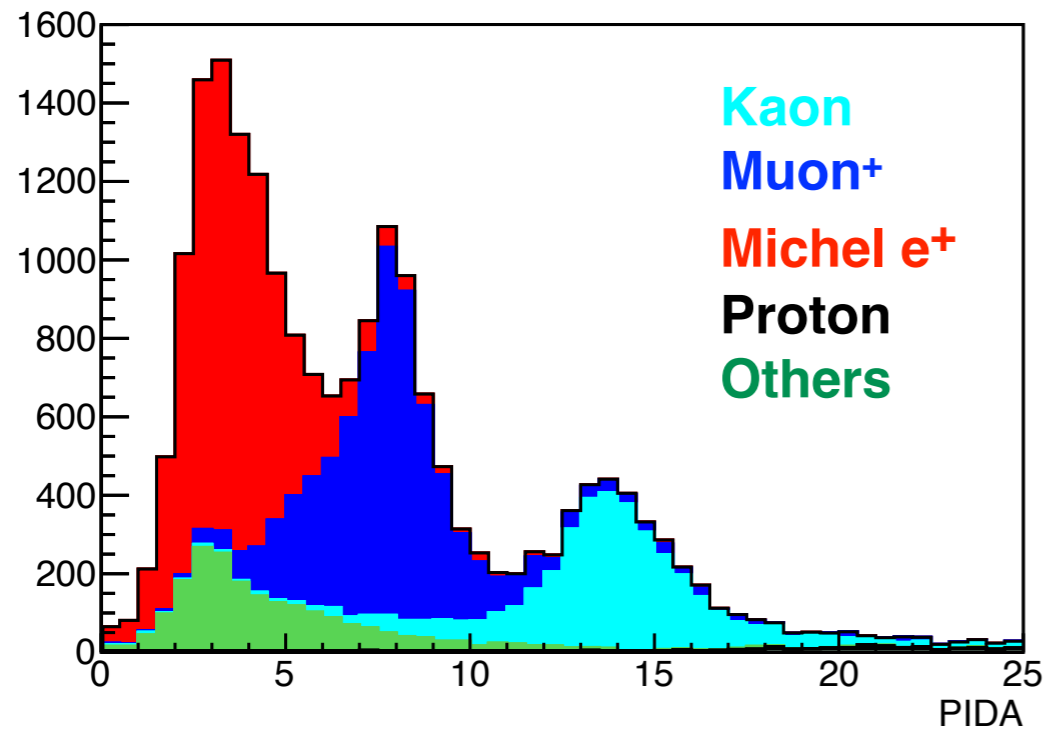




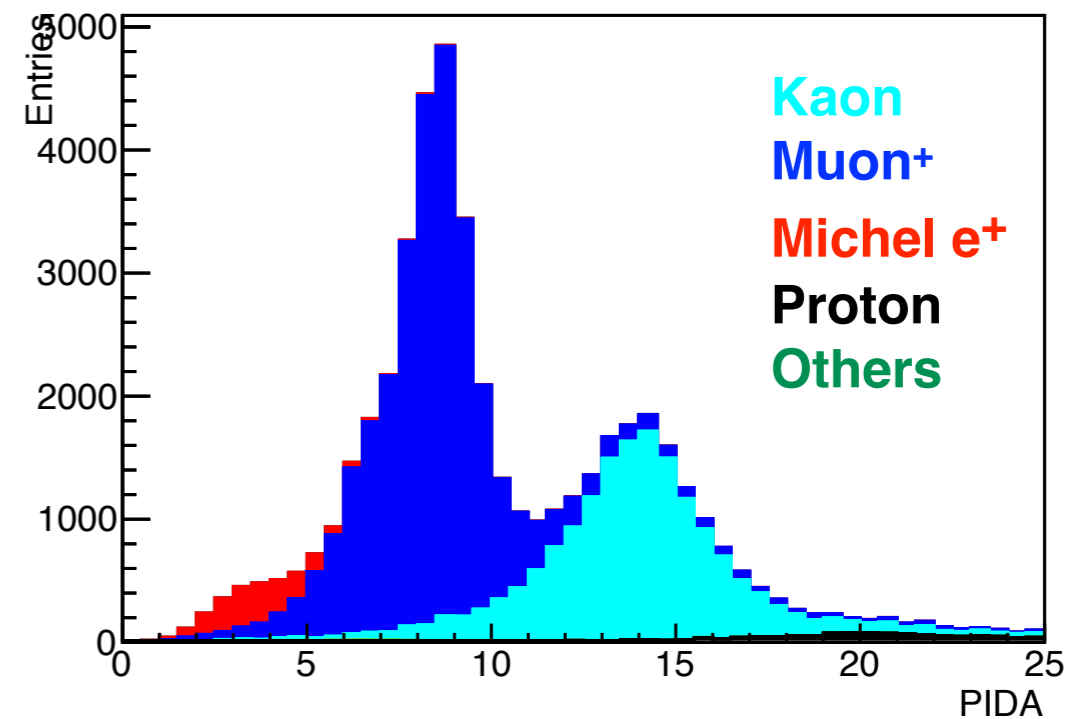
# Proton Decay Analysis

Only on  $K \rightarrow \mu^+ + \nu_\mu$  events

Before using CNN



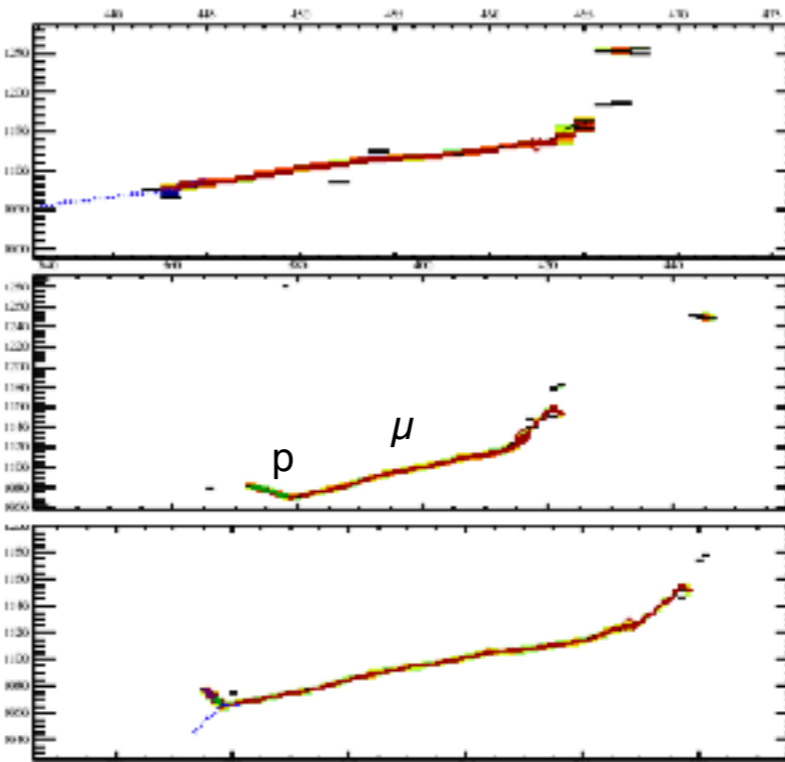
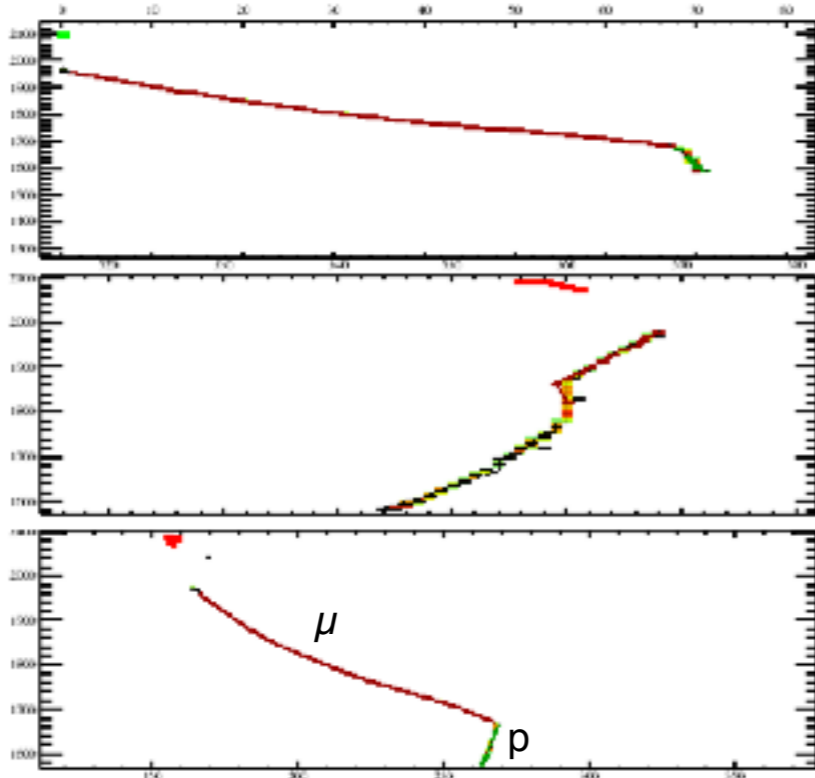
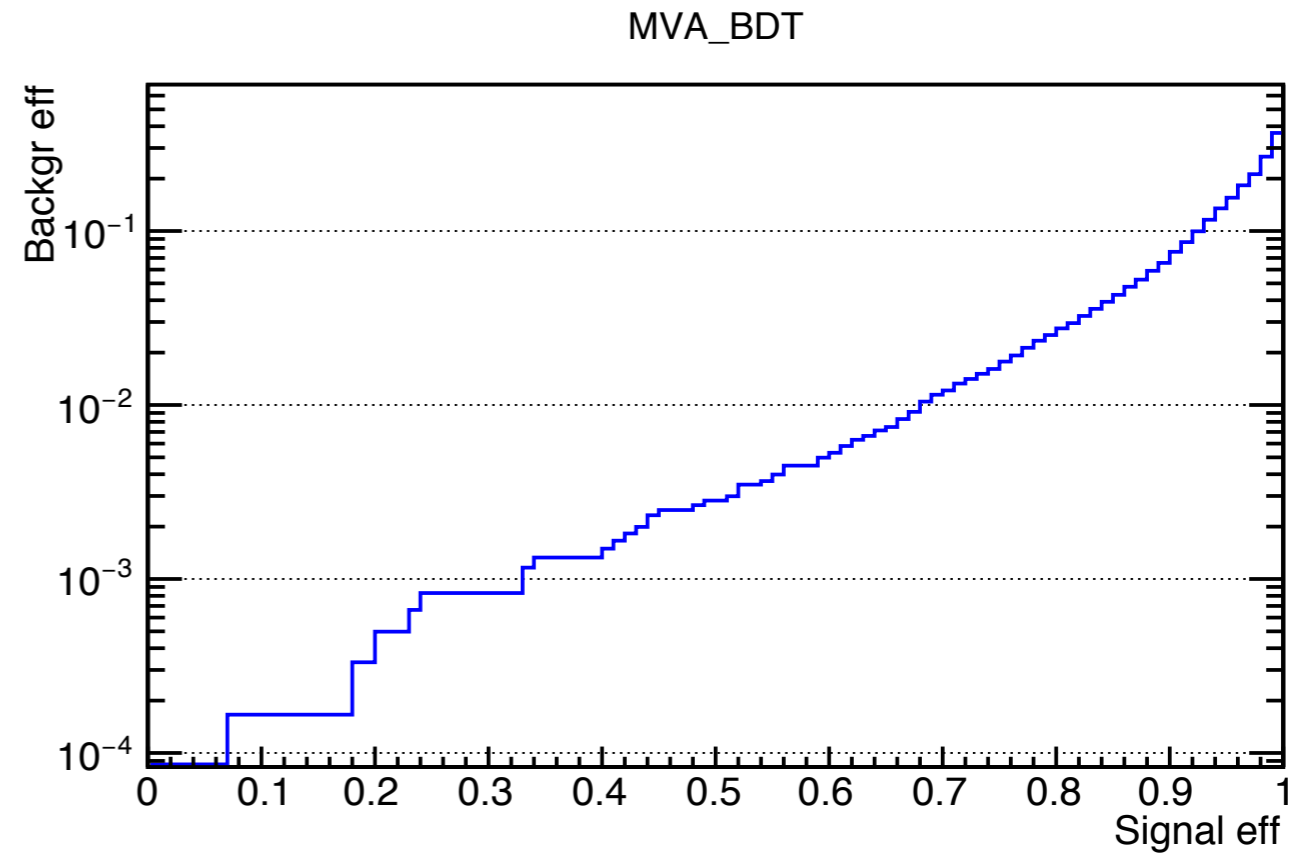
After using CNN



# MVA for Proton Decay

## Boosted Decision Tree (BDT)

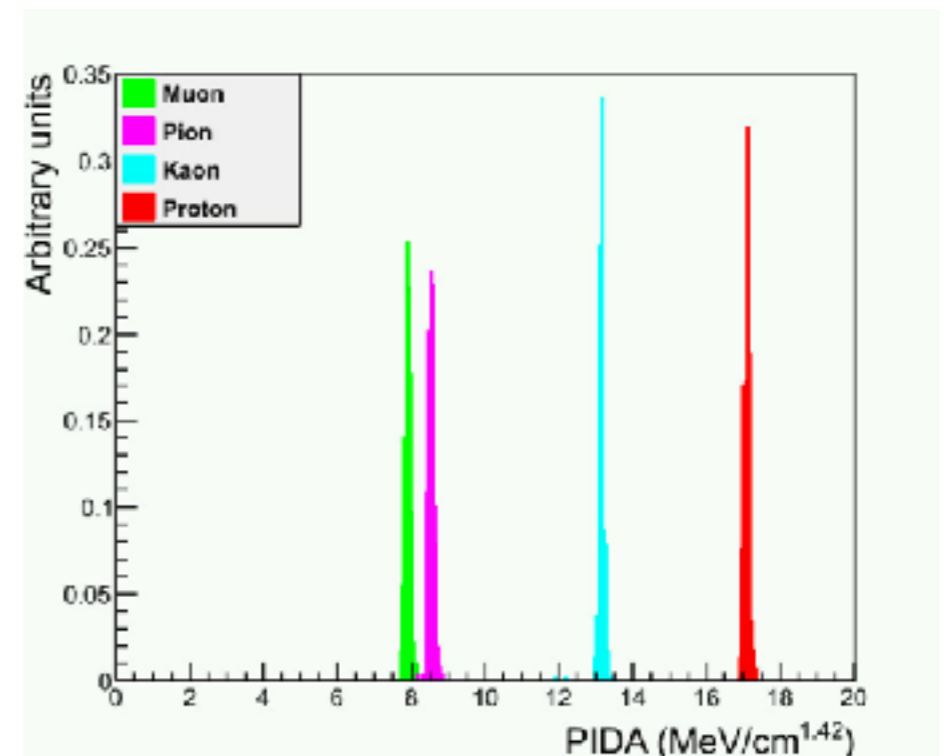
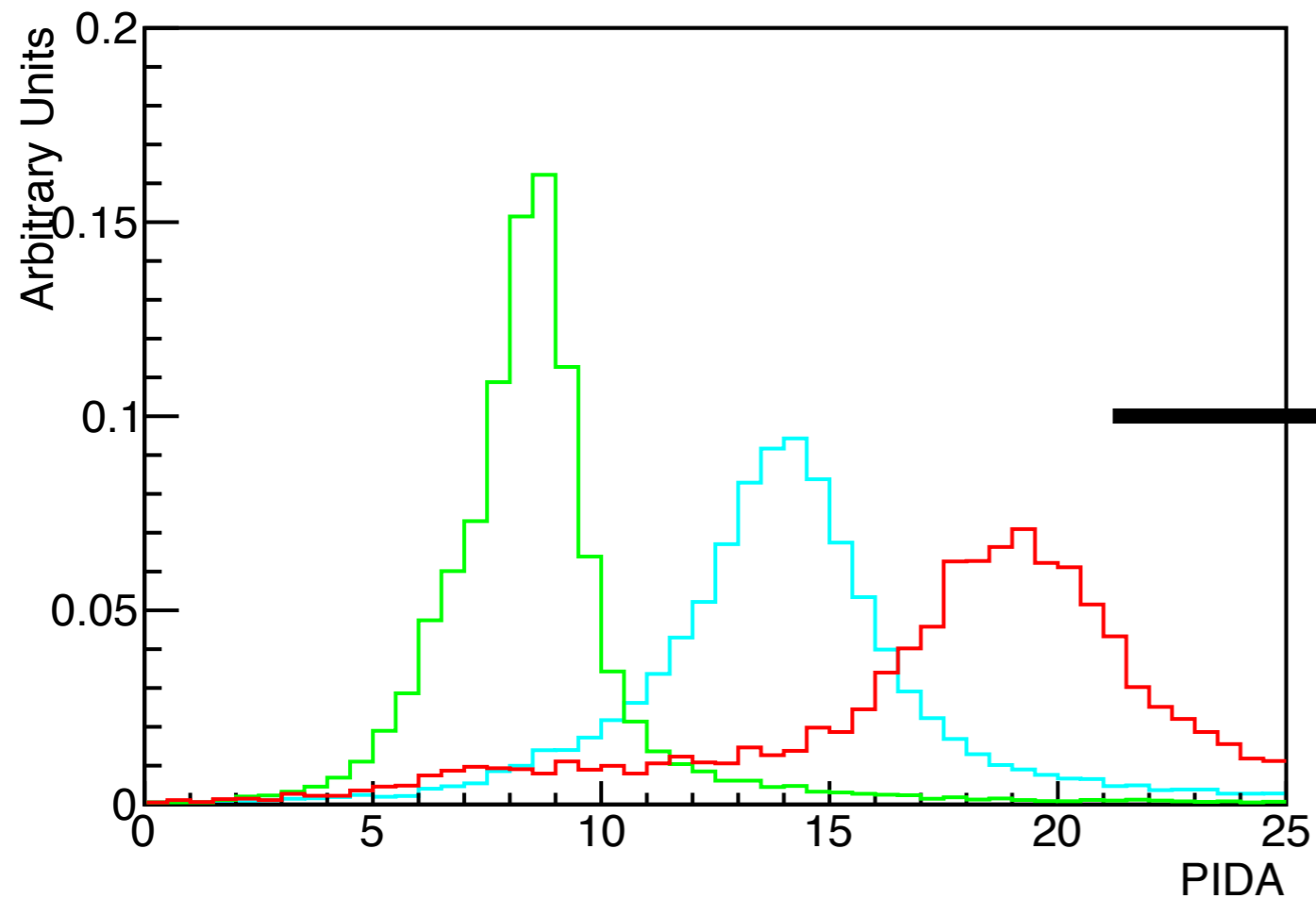
Signal	Background Mton <sup>-1</sup> year <sup>-1</sup>
100%	$2.9 \times 10^5$
<b>50%</b>	<b>783</b>
20%	95



Background events are  $\nu_\mu$  CCQE

# How we move forward?

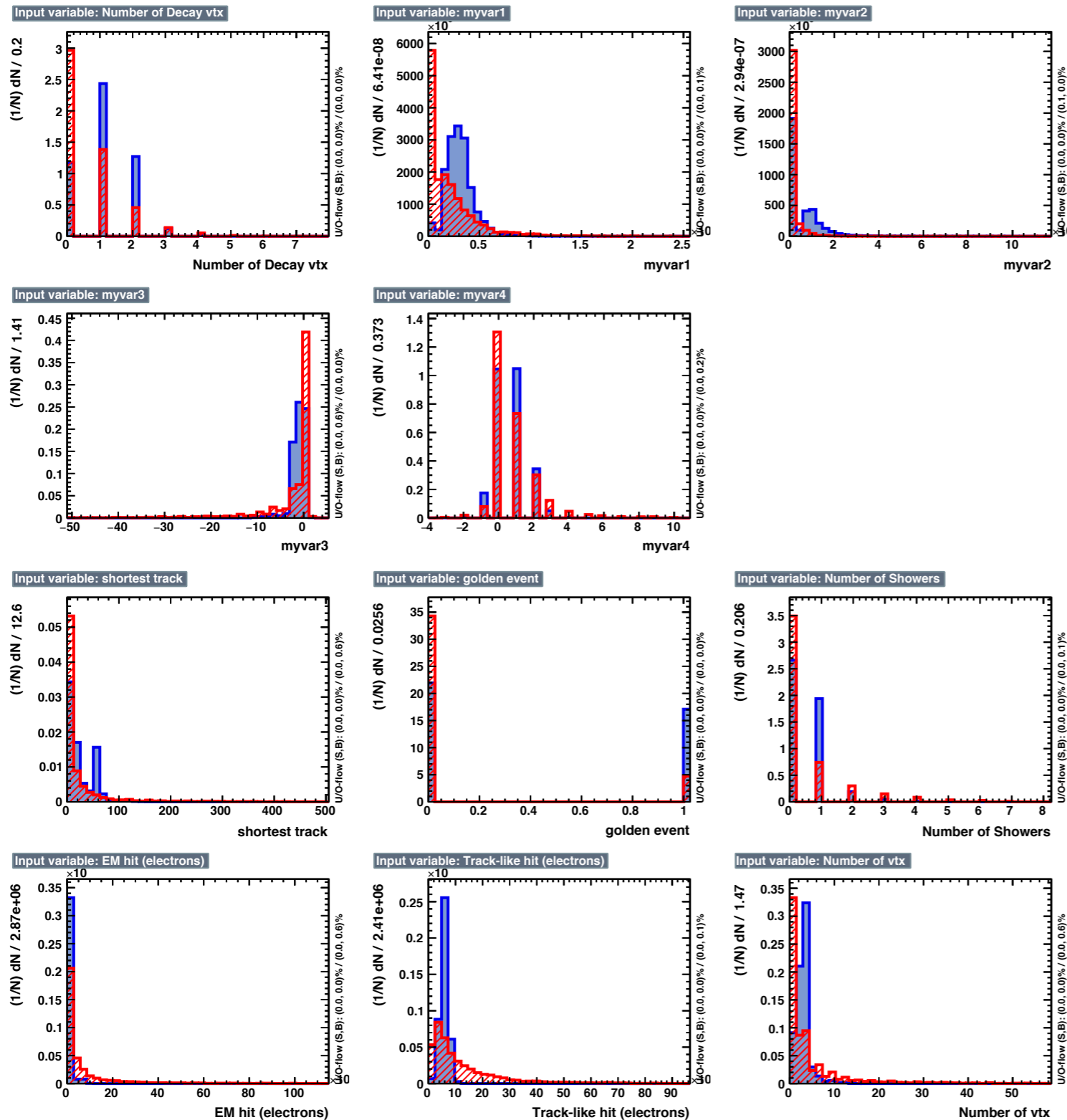
track-like hits(CNN) + PMA (linecluster)



The End



# MVA for Proton Decay



Signal:

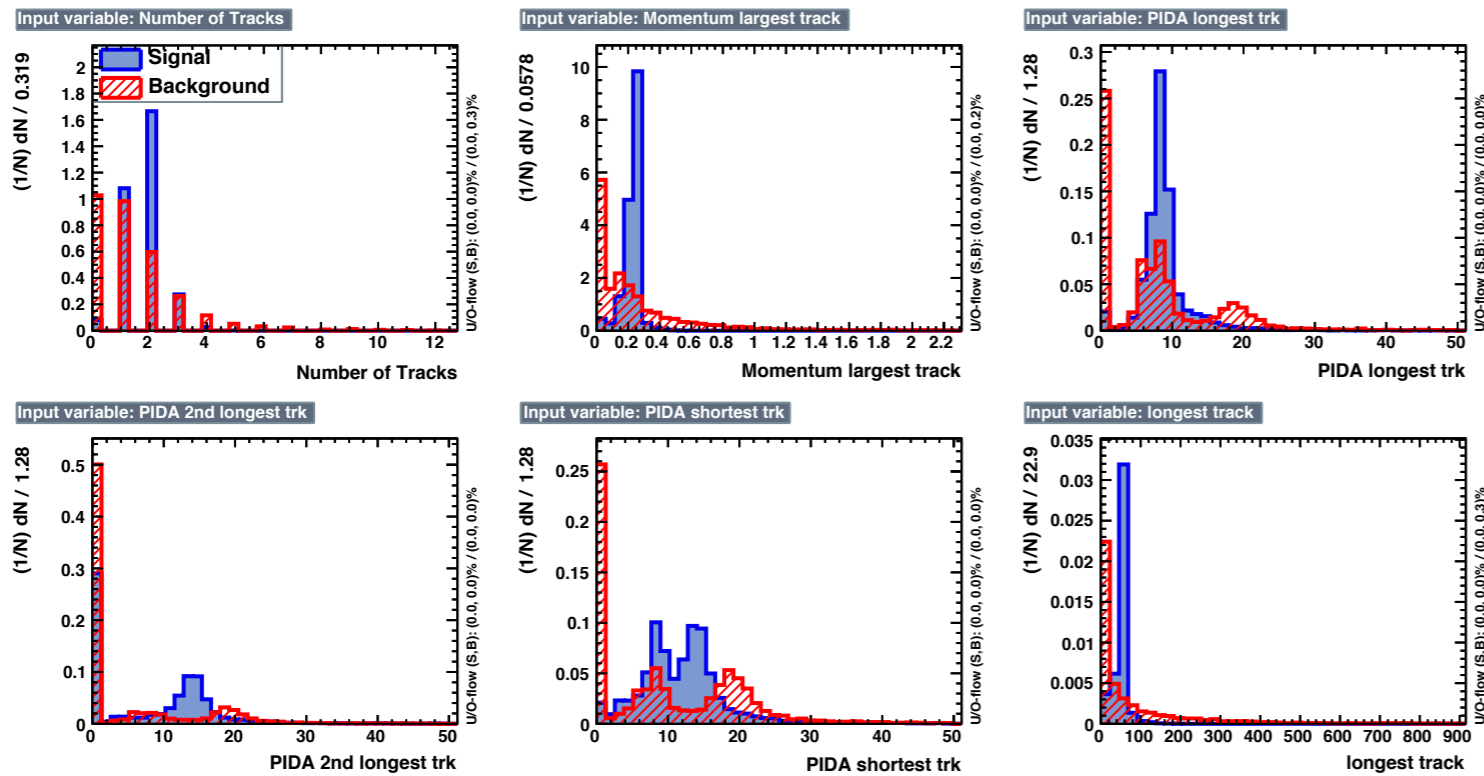
$K \rightarrow \mu^+ + \nu_\mu$  events

Background:

Atmospheric events

- Number of dcy vtx
- golden event
- Track-like hits
- EM-like hits
- Number of showers
- Total shower energy
- N tracks/trk-like hits
- N showers/em-like hits
- N trks - N vtx

# MVA for Proton Decay



Signal:

$K \rightarrow \mu^+ + \nu_\mu$  events

Background:

Atmospheric events

- Number of tracks
- PIDA
- Track length
- P by range