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# NuSTEC at the Short Baseline Neutrino programme

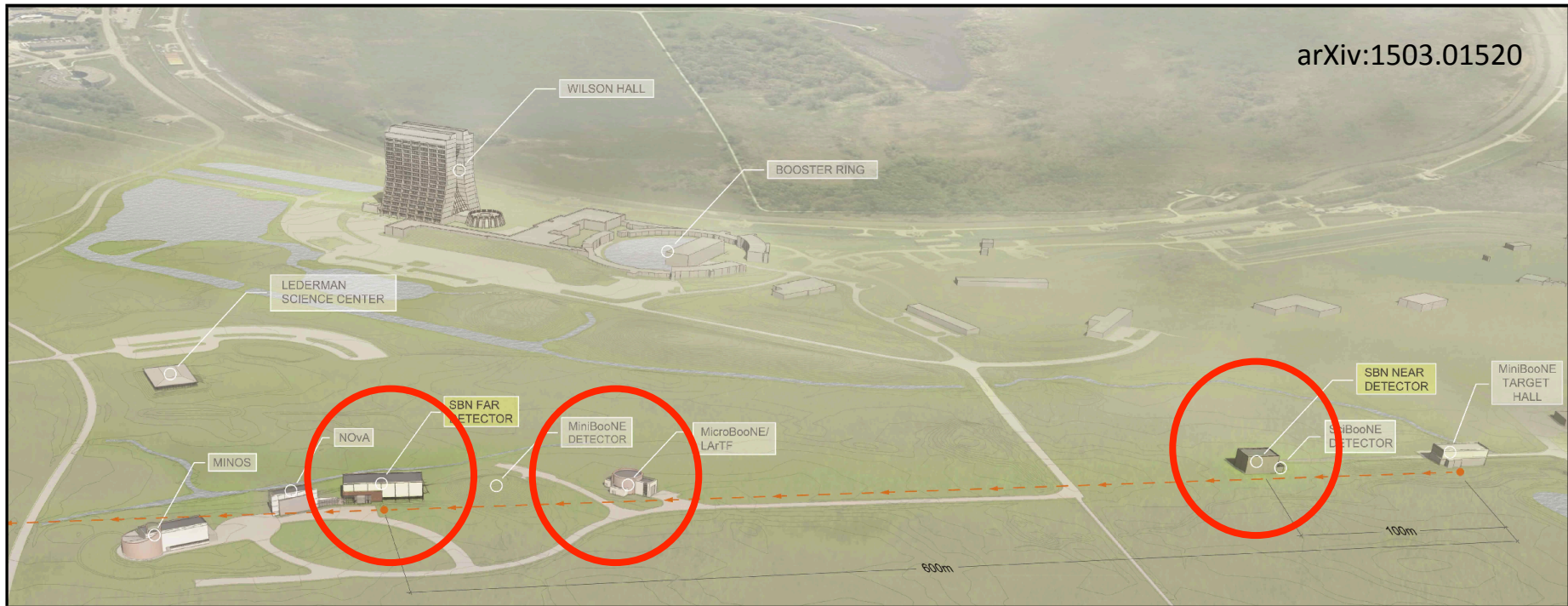
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NuSTEC board meeting

16<sup>th</sup> November 2017

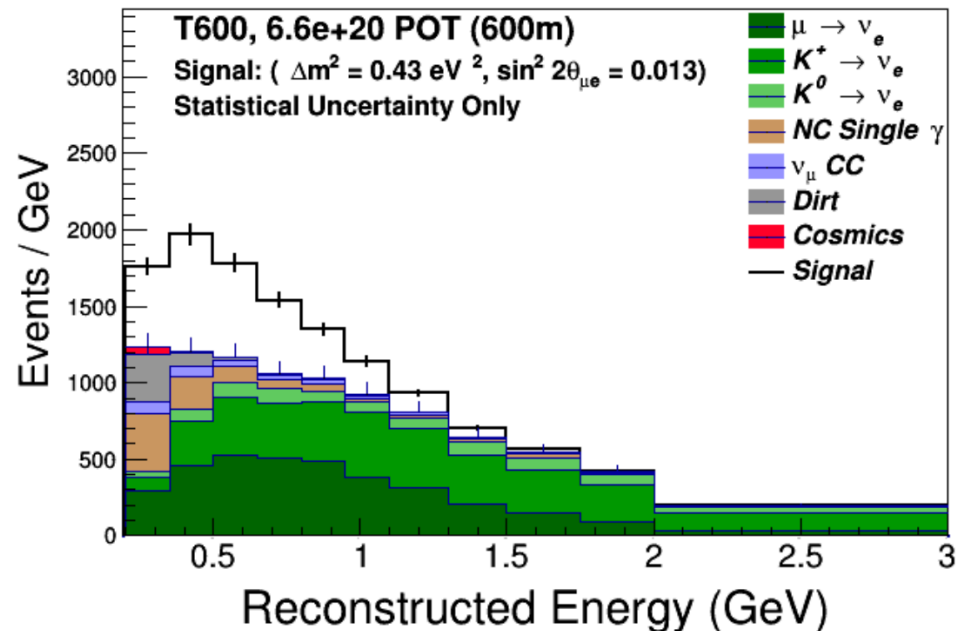
# SBN summary

- Three LArTPCs at 100, 450, and 600m (approx) along the BNB beamline
- BNB is a wide band beam with peak at 600MeV



# Oscillation searches

- Using near, intermediate, and far detectors, search for  $\nu_\mu$  disappearance and  $\nu_e$  appearance
- Try to avoid strong model dependence using calorimetric energy reconstruction
  - Of course, this does introduce different model dependencies
  - And new difficulties!
- Potential oscillation signal at ICARUS (far detector)
- Intrinsic and other backgrounds constrained at near detector



# Key detector challenges

- Long drift time, and a high cosmic rate (no overburden)
- Photon conversion length is 14cm
  - Worries about merging cosmogenic activity in neutrino reconstruction
- Detector effects such as space charge, lifetime, recombination, diffusion
- Large fraction of muons escape detector – need to use momentum estimation from multiple coulomb scattering
  - Angular dependence to momentum resolution

# Low energy protons

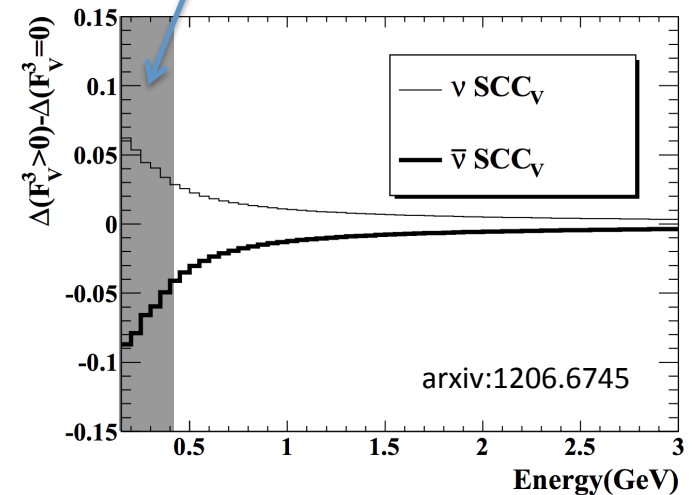
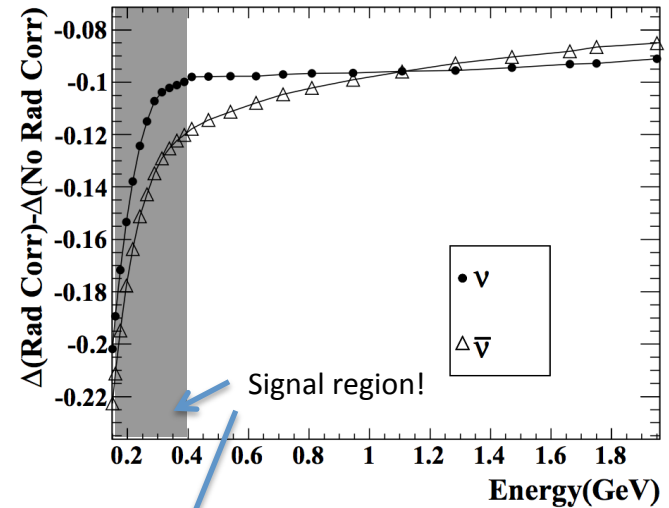
- We are pushing our proton reconstruction threshold down from  $\sim 400\text{MeV}/c$  to  $\sim 200\text{MeV}/c$ 
  - Not just yet, but work has started. MicroBooNE is already highly competitive with MINERvA
- Fairly strong angular-dependent efficiency
- Lost protons (and neutrons) means a model-dependent energy correction
- Theory models that retain the **hadronic degrees of freedom** would be very welcome!

# Extrapolation from light nuclei

- Most recent data at these energies is on hydrocarbon targets
- Models **built for carbon extrapolated to argon**
  - Argon is bigger – more FSI etc
  - Different p/n ratio – np/pp correlations?
- Advice on this extrapolation, or models built for argon, would be awesome!

# muon-electron differences

- Best sensitivity – joint fit of all  $\nu_\mu$  and  $\nu_e$  samples at 3 detectors
- Appearance signal expected at low energy (<500MeV)
  - Mass difference leads to different coverage of **energy/ momentum transfer**
  - **Radiative corrections**
  - **mass terms in form factors** (second class currents, pseudoscalar form factor changes)



# What we bring to the table - data!!!

- SBND in particular will have huge statistics
- We can provide very precise topological flux-integrated measurements

Inclusive	5,389,168
0 $\pi$	3,814,198
0 $\pi$ + 0p	27,269
0 $\pi$ + 1p (> 20 MeV)	1,629,252
0 $\pi$ + 2p (> 20 MeV)	1,150,368
0 $\pi$ + 3p (> 20 MeV)	413,956
0 $\pi$ + >3p (> 20 MeV)	396,212
1 $\pi^+$ + X	942,555
1 $\pi^-$ + X	38,012
1 $\pi^0$ + X	406,555
2 $\pi$ + X	145,336
$\geq 3\pi$ + X	42,510

- We are going to need help to understand all the data we have!
- What is the most useful thing to measure?
- What is the best way of presenting this?

More detector interactions in one month than historical dataset

Hyperon Production Expectations (CC + NC)  
(GENIE Expectations)

	1 Month	1 Year	3 Years
$\Lambda^0$ Production	200	2,600	8,000
$\Sigma^+$ Production	125	1,500	4,500



# Summary

- SBN wants:
  - Models that predict proton spectra
  - Advice on extrapolating from lower mass nuclei
  - Calculations of corrections that impact the  $\nu_\mu/\nu_e$  cross section ratio
- In exchange for:
  - Large amounts of high-precision, high-statistics data!