## Approximate Methods for Nuclei II

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## Outline

#### 6. Approximate methods for nuclei (II) (2 hours).

The polarization propagator; RPA approach; RPA equations; many-body diagrams; meson exchange currents and 2p2h terms. Continuum RPA.

Mean field description of the nucleus and beyond: long range correlations, collective excitations and random phase approximation; interactions at low energy; short range correlations, meson exchange currents and their influence on genuine quasi-elastic and 2-nucleon knockout neutrino-induced processes.

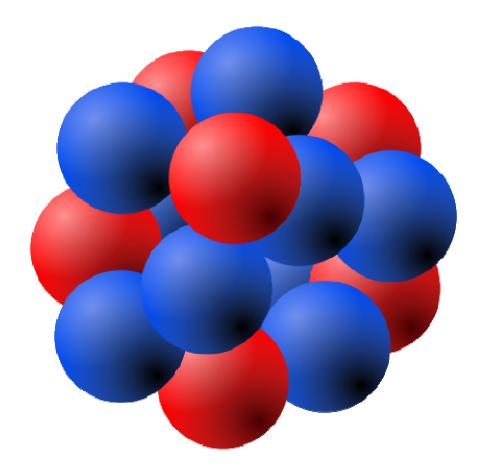
- Basic nuclear models
- Long range correlations : propagators, RPA and low energy processes
- Short range correlations
- Meson exchange corrections

## Modeling the nucleus

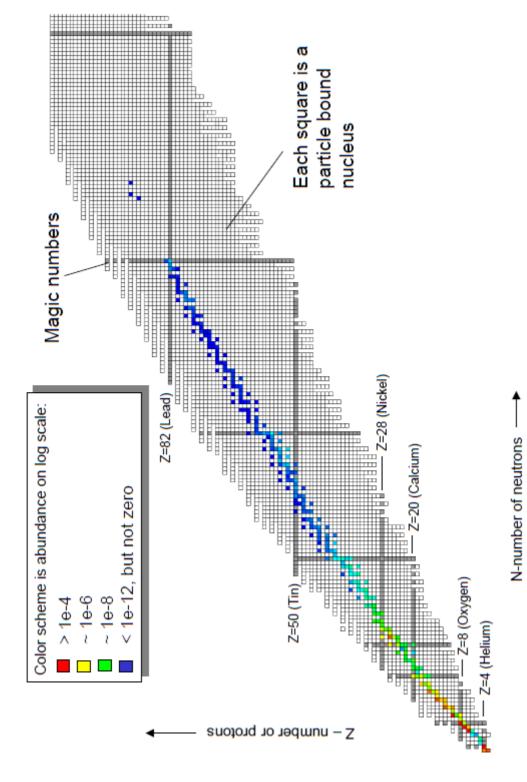
The nucleus is a mesoscopic system:

- -usually too big for few-body techniqes
- -usually too small for statistical methods

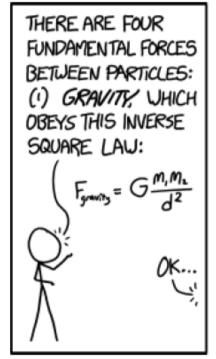
Nuclear physics is hard work!

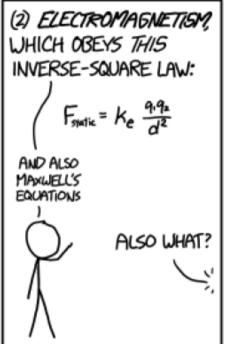


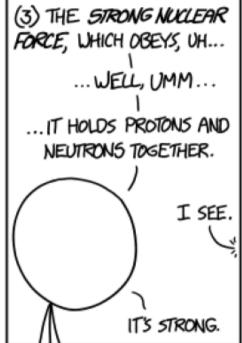
Abundances of nuclei on the chart of nuclides:

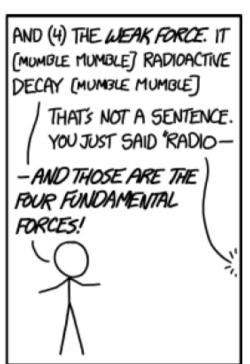


Nuclear dynamics is governed by the nuclear force = the residue of the strong interaction



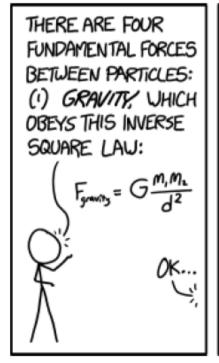


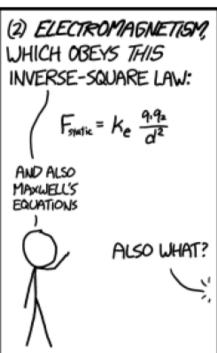


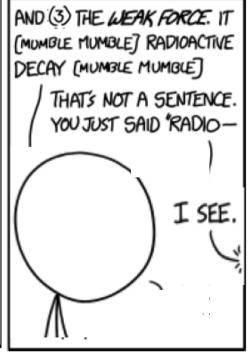


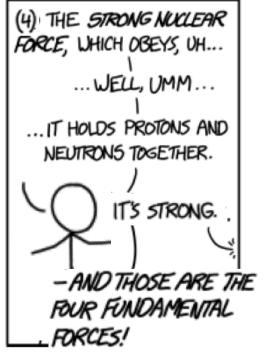
And (quite) a bit by the Coulomb interaction

- The nucleons inside the nucleus are bound via the nuclear force. This force is understood as a residual effect of the strong force, which is the force that binding quarks together to form the nucleons.
- To a large extent, the nuclear force can be understood in terms of the exchange of virtual light mesons such as the  $\pi$ ,  $\rho$  and  $\omega$  meson

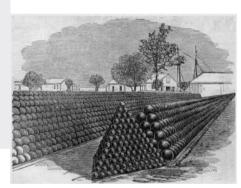






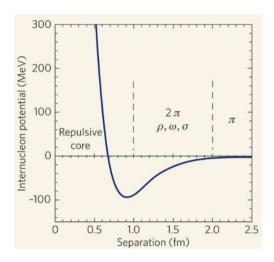


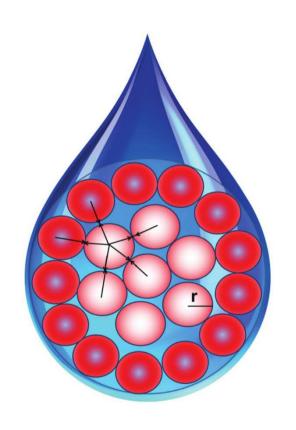
- Nuclear radius  $\approx 1.2A^{\frac{1}{3}}$  fm
- ► Nucleon is a diffuse system
  - lacksquare Hard core (repulsion) pprox 0.5 fm
  - RMS charge radius from (e,e') = 0.897(18) fm
- ▶  $0.07 \lesssim \text{NPF} \lesssim 0.42$ 
  - closest packing fraction of spheres  $\approx 0.74$
  - packing fraction of Argon liquid  $\approx 0.032$
  - packing fraction of Argon gas  $\approx 3.75 \cdot 10^{-5}$
- ► The nuclear medium is a rather dense quantum liquid



C. Colle, PhD, UGent 2017

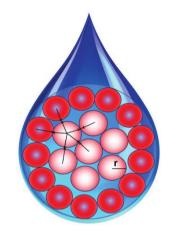






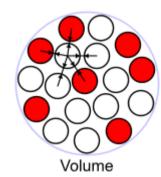
$$M(Z,N) = Z \, M_H + N \, M_n - BE$$
 Binding energy

Bulk description of the nucleus Liquid drop of uniform density



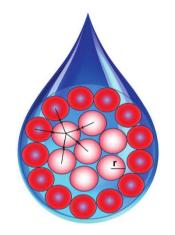
$$M(Z,N) = Z M_H + N M_n - BE$$

$$BE = C_1 A$$



#### **Volume term**

- Binding energy per nucleon is ~constant for stable nuclei over the whole mass table
- The nuclear force is short-range, attractive and saturated
- $V \propto A$  $R = r_0 A^{1/3} = 1.28 fm A^{1/3}$

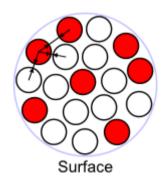


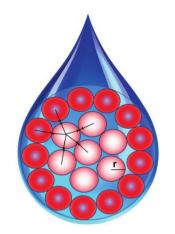
$$M(Z,N) = Z M_H + N M_n - BE$$

$$BE = C_1 A - C_2 A^{2/3}$$

#### Surface term

- Nucleons at the surface are not surrounded by other nucleons to attract
- Surface nucleons have reduced binding compared to those in the nuclear interior



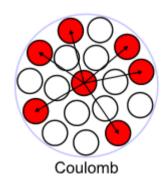


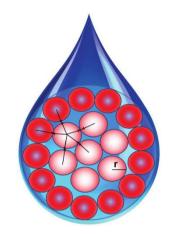
$$M(Z,N) = Z M_H + N M_n - BE$$

$$BE = C_1 A - C_2 A^{2/3} - C_3 \frac{Z^2}{A^{1/3}}$$

#### Coulomb term

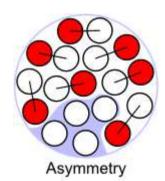
- Coulomb force is repulsive and long range
- This term reduces the number of protons in large nuclei





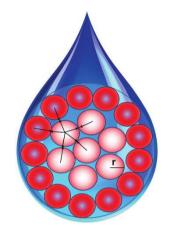
$$M(Z,N) = Z M_H + N M_n - BE$$

$$BE = C_1 A - C_2 A^{2/3} - C_3 \frac{Z^2}{A^{1/3}} - C_4 \left(\frac{A}{2} - Z\right)^2$$



### Symmetry term

- Pauli principle
- This term tends to keep the number of protons and neutrons equal

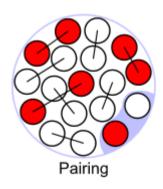


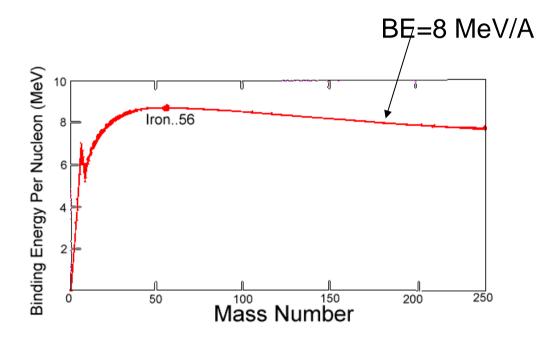
$$M(Z,N) = Z M_H + N M_n - BE$$

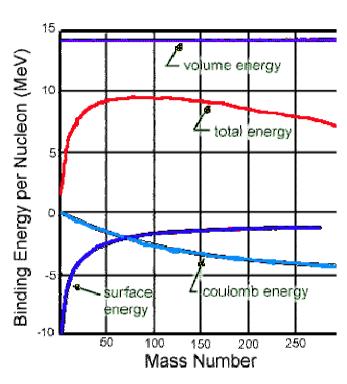
$$BE = C_1 A - C_2 A^{2/3} - C_3 \frac{Z^2}{A^{1/3}} - C_4 \left(\frac{A}{2} - Z\right)^2 + \delta_{N,Z}$$

### **Pairing term**

- Pairing force
- Protons and neutrons that are paired tend to have enhanced binding energy
- +  $\delta$  (even-even)
- + 0 (odd-even /even-odd)
- $-\delta$  (odd-odd)





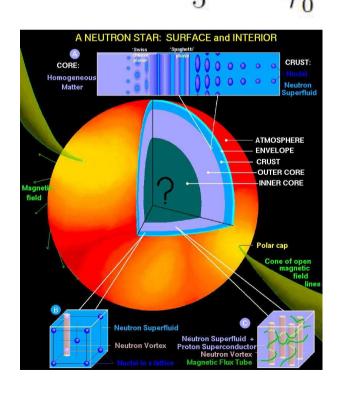


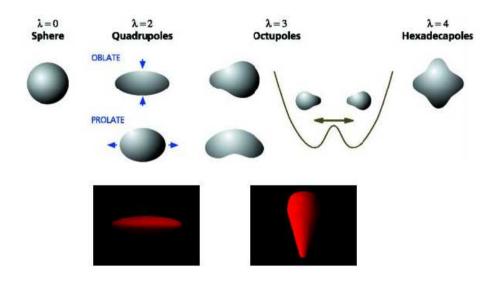
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## To boldly go ... Neutron Star Stability

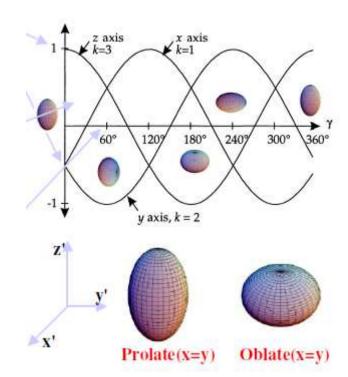
$$BE = C_1 A - C_2 A^{2/3} - C_3 \frac{Z^2}{A^{1/3}} - C_4 \left(\frac{A}{2} - Z\right)^2 + \delta_{N,Z} + \frac{3}{5} G \frac{M^2}{R} + \frac{3}{5} G \frac{M^2 A^{-1/3}}{R}$$

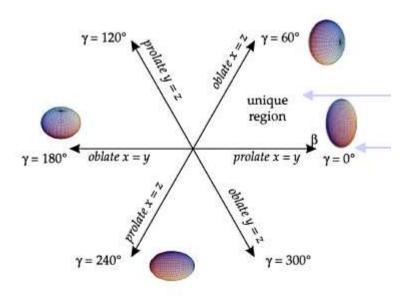
$$A \simeq 5 \times 10^{55}$$
 $R \simeq 4.3 \,\mathrm{km}$ 
 $M \simeq 0.045 M_{\odot}$ 





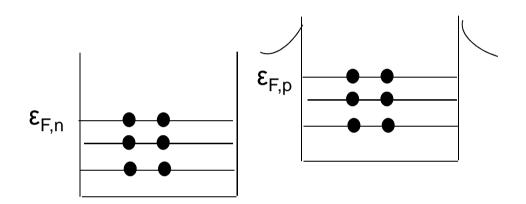
Octupolevibration vibrations





## Slightly less approximate: (still almost) no correlations: the Fermi gas model

### Easiest microscopic independent particle model



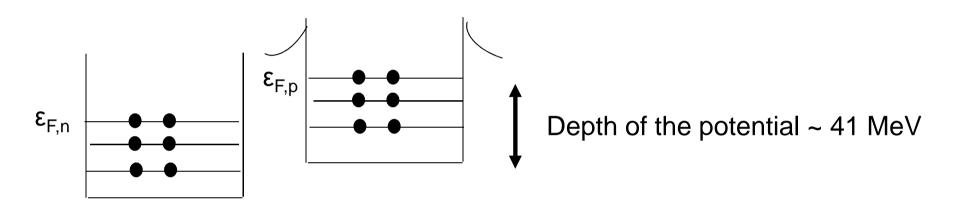
$$n = 2 \frac{V}{(2\pi\hbar)^3} \int_0^{p_F} d^3p = 2 \frac{V}{(2\pi\hbar)^3} \left(\frac{4}{3}\pi p_F^3\right)$$

$$\Rightarrow p_F = \hbar \left(\frac{3n\pi}{V}\right)^{1/3} \sim 247 \,\text{MeV/c}$$

$$E_F = \frac{p_F^2}{2m} \sim 33 \,\text{MeV}$$

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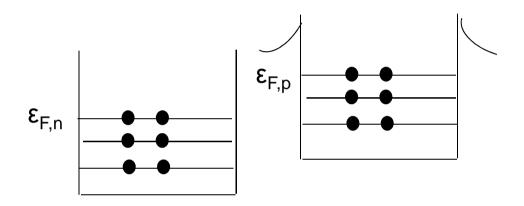
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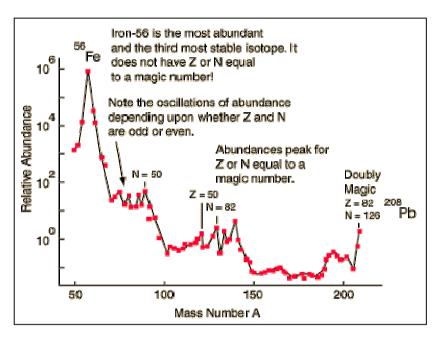
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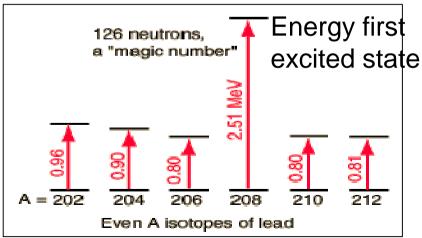
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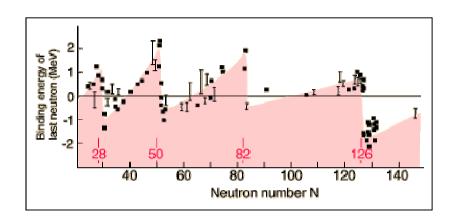


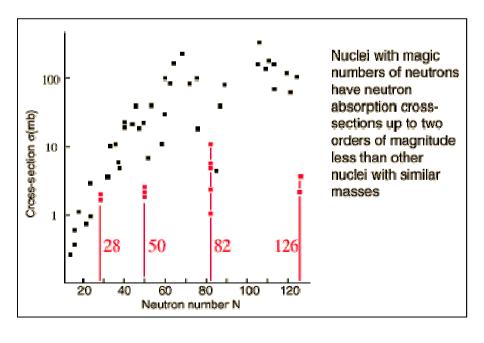
$$\langle E \rangle = \frac{\int_0^{E_F} E_{kin} d^3 p}{\int_0^{p_F} d^3 p} = \frac{1}{2m} \frac{\int_0^{E_F} p^2 d^3 p}{\int_0^{p_F} d^3 p} = \frac{3}{5} E_F \sim 20 \text{MeV}$$

# Not too approximate: already quite some correlations: the mean field model (or shell-model)





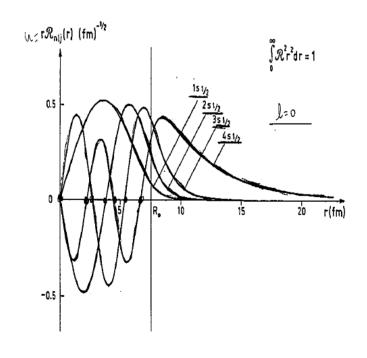




## Not too approximate: already quite some correlations: the mean field model (or shell-model)

→solve Schrödinger equation for nucleon in nuclear potential

### Harmonic oscillator potential

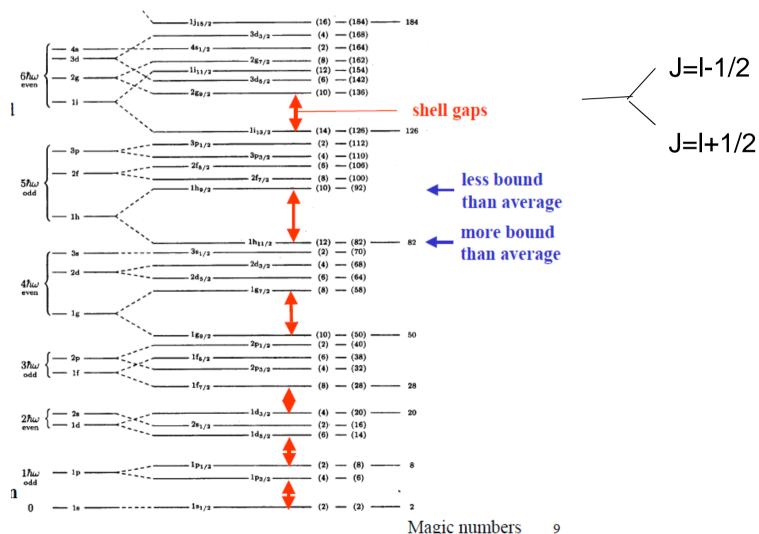


$$\varphi(\vec{r}) = u(r) Y_m^l(\Omega)$$
 spherical 
$$\frac{\sqrt{(0.4)}}{\sqrt{(1.2)}} \cdot \frac{\sqrt{(0.4)}}{\sqrt{(1.2)}} \cdot \frac{\sqrt{(0.4)}}{\sqrt{(0.4)}} \cdot \frac{\sqrt{(0.4)}}{\sqrt{(0$$

## Not too approximate: already quite some correlations: the mean field model (or shell-model)

## →Work harder : add spin-orbit term

$$\widehat{H} \to \widehat{H} + \xi(r)\overline{l} \cdot \overline{s}$$



### Independent particle picture ???

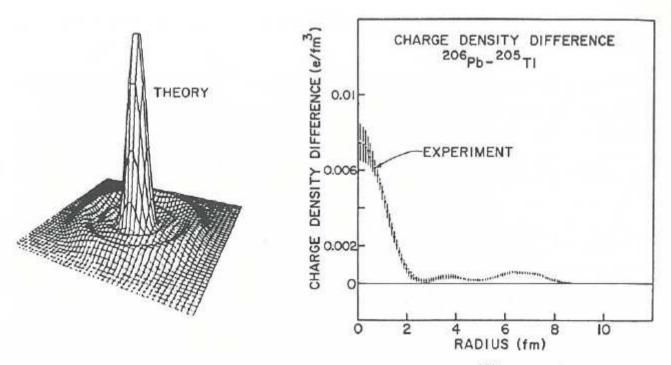


Fig. 3.18. The nuclear density distribution for the least bound proton in  $^{206}\text{Pb}$ . The shell-model predicts the last  $(3s_{1/2})$  proton in  $^{206}\text{Pb}$  to have a sharp maximum at the centre, as shown at the left-hand side. On the right-hand side the nuclear charge density difference  $\varrho_c$  ( $^{206}\text{Pb}$ ) $-\varrho_c$  ( $^{205}\text{Tl}$ ) =  $\varphi_{3s_{1/2}}^2$  (r) is given [taken from (Frois 1983) and Doe 1983)]

$$\rho(r) = \sum_{b < F} \varphi_b^*(r) \varphi_b(r)$$

## Not too approximate: already quite some correlations: the mean field model (or shell-model)

Woods-Saxon potential:

$$V(r) = V_0 \frac{1}{1 + e^{\left(\frac{r - R}{a}\right)}}$$

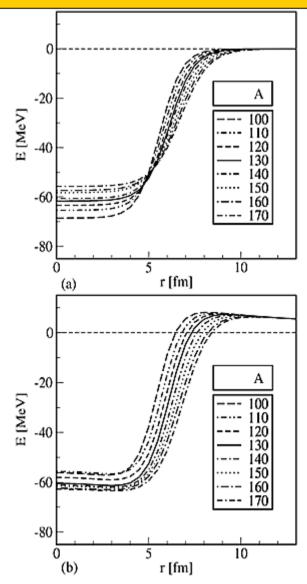


Figure 1. Woods-Saxon potential for neutrons (a) and protons (b) along the Sn isotopic chain.

#### The Hartree-Fock mean field

- Nucleons are moving independent from each other in a mean field
- How do we obtain a reliable and consistent mean field?

$$H = \sum_{i} T_{i} + \frac{1}{2} \sum_{i,j} V_{i,j}$$

$$H = \sum_{i} (T_{i} + U(r_{i})) + \left(\frac{1}{2} \sum_{i,j} V_{i,j} - \sum_{i} U(r_{i})\right)$$

$$H = \sum_{i} h_{0}(i) + H_{res}(i,j)$$

Residual interaction

### The Hartree-Fock recipe:

Nucleons fill up a number of orbitals and form a density that can be written in terms of the occupied states as :

$$\rho(r) = \sum_{b < F} \varphi_b^*(r) \varphi_b(r)$$

The potential at a position r', generated by the nucleon-nucleon two-body interaction V(r,r') is given by

$$U_{H}(r') = \sum_{b < F} \int \varphi_{b}^{*}(r) V(r, r') \varphi_{b}(r) dr$$

$$- \frac{\hbar^{2}}{2m} \Delta \varphi_{i}(r) + \sum_{b < f} \int \varphi_{b}^{*}(r') V(r, r') \varphi_{b}(r') dr' \cdot \varphi_{i}(r)$$

$$- \sum_{b < f} \int \varphi_{b}^{*}(r') V(r, r') \varphi_{b}(r) \varphi_{i}(r') dr' = \varepsilon_{i} \varphi_{i}(r)$$

$$-\frac{\hbar^2}{2m}\Delta\varphi_i(r) + U_H(r)\varphi_i(r) - \int U_F(r,r')\varphi_i(r')dr' = \varepsilon_i\varphi_i(r)$$

$$U_H(r') = \sum_{b < F} \int \varphi_b^*(r)V(r,r')\varphi_b(r)dr$$

$$U_F(r,r') = \sum_{b < F} \varphi_b^*(r')V(r,r')\varphi_b(r)$$

Hartree-Fock recipe :Start with an initial guess for either the average field or the wave functions and use V(r,r') to solve the coupled equations to obtain better values e.g.

$$U_{H(F)}^{(0)}(\mathbf{r})$$
  $U_{H(F)}^{(1)}(\mathbf{r})$   $U_{H(F)}^{(2)}(\mathbf{r})$ 
 $\downarrow$   $\downarrow$   $\downarrow$   $\downarrow$   $\downarrow$   $\downarrow$   $\varphi_i^{(0)}(\mathbf{r})$   $\varphi_i^{(1)}(\mathbf{r})$   $\varphi_i^{(1)}(\mathbf{r})$   $\varphi_i^{(2)}(\mathbf{r})$ 
 $\varepsilon_i^{(0)}$   $\varepsilon_i^{(1)}$   $\varepsilon_i^{(1)}$   $\varepsilon_i^{(2)}$ 

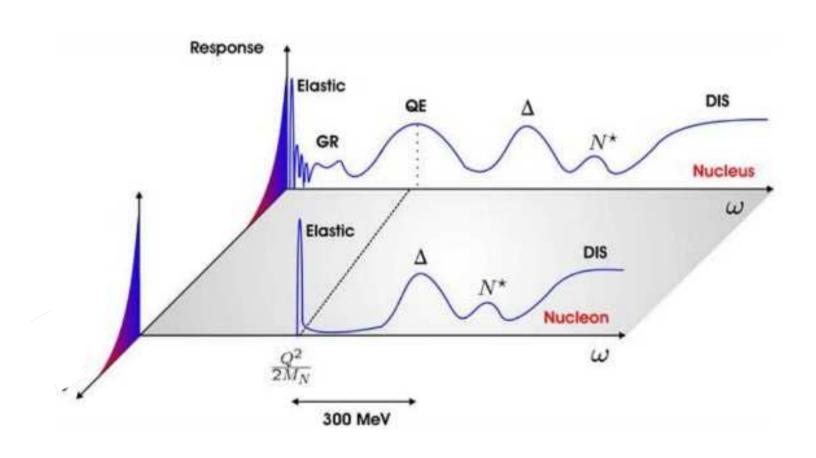
Wave function for the nucleus = Slater determinant

$$\Psi_{1,2,\ldots,A}(r_1,r_2,\ldots,r_A) = \frac{1}{\sqrt{A!}} \left| \begin{array}{ccc} \varphi_1(r_1) & \varphi_2(r_2) & \cdots & \varphi_A(r_A) \\ \varphi_2(r_1) & \varphi_3(r_2) & \cdots & \varphi_1(r_A) \\ \vdots & & & \end{array} \right|$$

$$E_0 = \sum_{i=1}^{A} \varepsilon_i$$

Antisymmetrization takes into account the Pauli principle

## Interactions ...

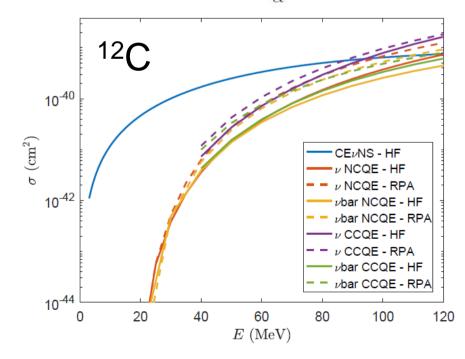


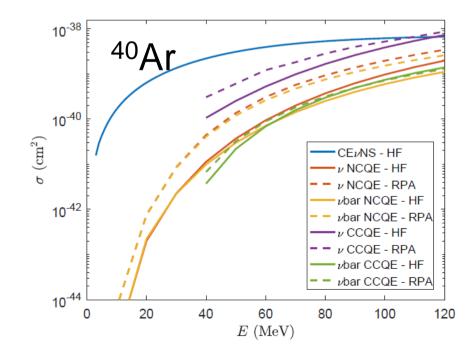
## Coherent (elastic) scattering: CEvNS

$$\frac{d\sigma}{dT} = \frac{G_F^2}{4\pi} Q_W^2 |F(-2MT)|^2 M \left(1 - \frac{T}{E} - \frac{MT}{2E^2}\right)$$

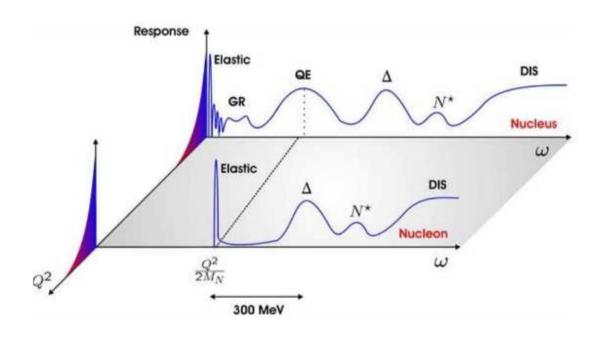
$$F(q^2) = \frac{4\pi}{Q_W} \int \left( (1 - 4\sin^2\theta_W) \rho_p(|\mathbf{r}|) - \rho_n(|\mathbf{r}|) \right) \frac{\sin(|\mathbf{q}||\mathbf{r}|)}{|\mathbf{q}||\mathbf{r}|} \mathbf{r}^2 d|\mathbf{r}|$$

$$\rho_{\tau}(|\mathbf{r}|) = \frac{1}{4\pi |\mathbf{r}|^2} \sum_{\alpha} (2j_{\alpha} + 1) R_{\alpha}^2(|\mathbf{r}|)$$





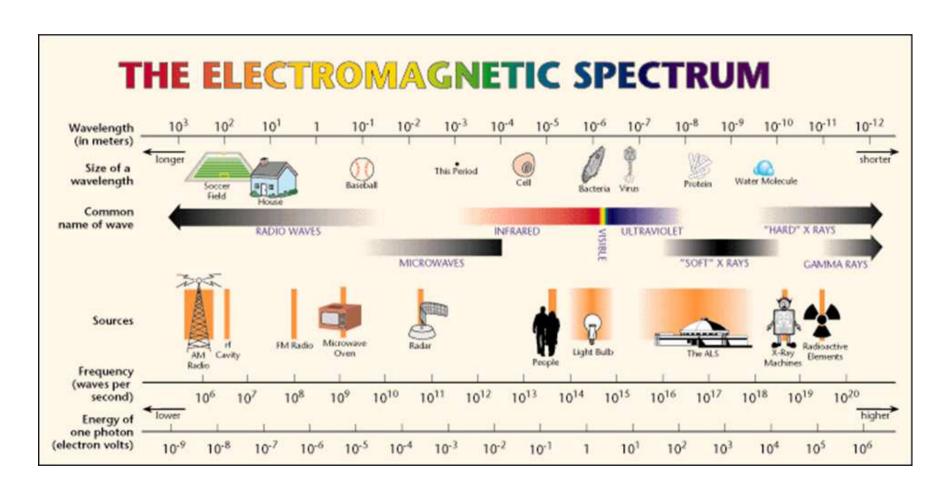
## The Random Phase approximation



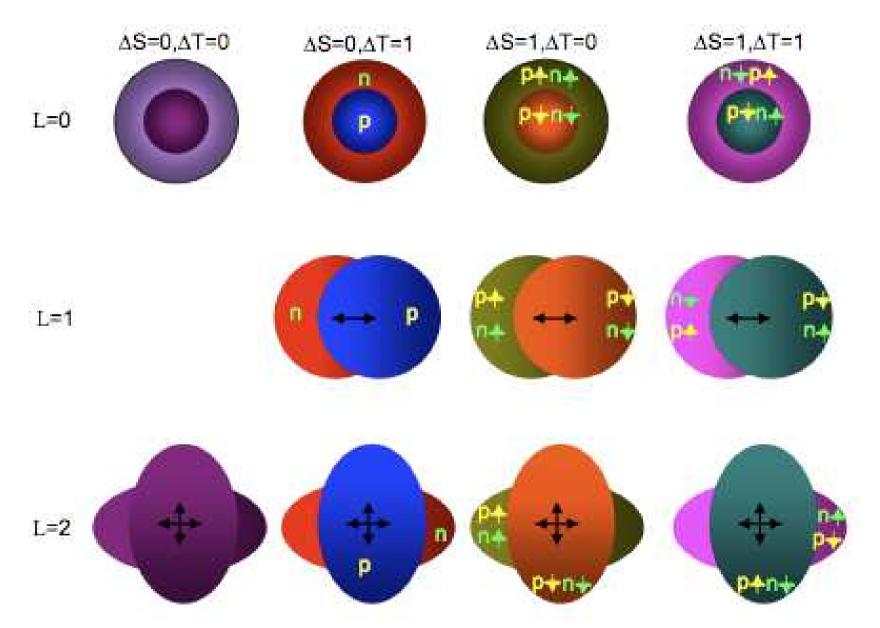
- Long-range correlations are correlations over the whole size of the nucleus
- They can redistribute the incoming energy transfer to the nucleus over all the nuclear constituents.
- They manifest themselves in collective excitations such as giant resonances

## The Random Phase approximation

Long-range correlations = low energy phenomena



### Long-range correlations = probing collective effects



https://cyclotron.tamu.edu/research/nuclear-structure/

#### RPA ... what's in a name

David Bohm & David Pines '52 – Condensed matter physics

Quantum mechanical interactions interaction between electrons

The Hamiltonian corresponding to a system of individual electrons is reexpressed such that the long-range part of the Coulomb interactions between electrons is described in terms of collective fields

This leads to the description of organized behavior of electrons brought along by long-range Coulomb interactions that couple together the motion of many electrons simultaneously = plasma oscillations

Neglecting the coupling between plasma vibrations of different momenta

### Propagators (in a nutshell)

Time evolution of a single-particle system characterized by quantum numbers  $\alpha$  at time  $t_0$ .

$$|\alpha, t_0; t\rangle = e^{-\frac{i}{\hbar}H(t-t_0)} |\alpha; t_0\rangle$$

$$i\hbar \frac{\partial}{\partial t} |\alpha, t_0; t\rangle = H |\alpha, t_0; t\rangle$$

Time-dependent Schrödinger equation

Ref.: Many-body Theory exposed! WH Dickhoff, D. Van Neck

$$\Psi(\vec{r},t) = \int d\vec{r}' \langle \vec{r} | e^{-\frac{i}{\hbar}H(t-t_0)} | \vec{r}' \rangle \langle \vec{r}' | \alpha, t_0 \rangle$$

$$= i\hbar \int d\vec{r}' G(\vec{r}, \vec{r}'; t, t_0) \Psi(\vec{r}', t_0)$$

Propagator or Green's function :  $G(\vec{r}, \vec{r}'; t, t_0) = -\frac{i}{\hbar} \langle \vec{r} | e^{-\frac{i}{\hbar}H(t-t_0)} | \vec{r}' \rangle$ 

$$G(\vec{r}, \vec{r}'; t, t_0) = -\frac{i}{\hbar} \langle 0 | \hat{a}_r e^{-\frac{i}{\hbar} H(t - t_0)} \hat{a}_{r'}^{\dagger} | 0 \rangle$$

$$= -\frac{i}{\hbar} \sum_{n} \langle 0 | \hat{a}_r | n \rangle \langle n | \hat{a}_{r'}^{\dagger} | 0 \rangle e^{-\frac{i}{\hbar} \varepsilon_n (t - t_0)}$$

$$= -\frac{i}{\hbar} \sum_{n} u(\vec{r}) u_n^* (\vec{r}') e^{-\frac{i}{\hbar} \varepsilon_n (t - t_0)}$$

#### Fourier transform:

# Impose causality

$$\theta(t - t_0) = -\frac{1}{2\pi i} \int dE' \frac{e^{-\frac{i}{\hbar}E'(t - t_0)}}{E' + i\eta}$$

$$G(\vec{r}, \vec{r}'; E) = -\frac{i}{\hbar} \int_{-\infty}^{+\infty} d(t - t_0) e^{\frac{i}{\hbar} E(t - t_0)} \theta(t - t_0) \sum_{n} u(\vec{r}) u_n^*(\vec{r}') e^{-\frac{i}{\hbar} \varepsilon_n (t - t_0)}$$

$$= \sum_{n} \int_{-\infty}^{\infty} dE' \delta(E - E' - \varepsilon_n) \frac{1}{-u_n(\vec{r})} u_n^*(\vec{r}') e^{-\frac{i}{\hbar} \varepsilon_n (t - t_0)}$$

$$= \sum_{n} \int dE' \delta(E - E' - \varepsilon_{n}) \frac{1}{E' + i\eta} u_{n}(\vec{r}) u_{n}^{*}(\vec{r}')$$

$$\sum_{n} u_{n}(\vec{r}) u_{n}^{*}(\vec{r}')$$

$$= \sum_{n} \frac{u_n(\vec{r})u_n^*(\vec{r}')}{E - \varepsilon_n + i\eta}$$

$$= \langle \vec{r} | \frac{1}{E - H + i\eta} | \vec{r}' \rangle$$

$$= \sum_{n} \frac{u_{n}(\vec{r})u_{n}^{*}(\vec{r}')}{E - \varepsilon_{n} + i\eta}$$

$$= \langle \vec{r} | \frac{1}{E - H + i\eta} | \vec{r}' \rangle$$

$$\delta(x - a) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} e^{i(x - a)t} dt$$

Or more general:

$$G(\alpha, \beta; E) = \langle 0 | a_{\alpha} \frac{1}{E - H + i\eta} a_{\beta}^{\dagger} | 0 \rangle$$



#### Any Hamiltonian

$$G(\alpha, \beta; E) = \langle 0 | a_{\alpha} \frac{1}{E - H + i\eta} a_{\beta}^{\dagger} | 0 \rangle$$

$$\sum_{\alpha} \frac{1}{\varepsilon_{\alpha} a_{\alpha}^{\dagger} a_{\alpha}} \frac{1}{\delta_{\alpha, \beta}}$$

$$\frac{1}{\delta_{\alpha, \beta}} \frac{1}{\delta_{\alpha, \beta$$

$$G^{(0)}(\alpha, \beta; E) = \langle 0 | a_{\alpha} \frac{1}{E - H_0 + i\eta} a_{\beta}^{\dagger} | 0 \rangle = \frac{\delta_{\alpha, \beta}^{\alpha}}{E - \varepsilon_{\alpha} + i\eta}$$

$$G = \frac{1}{E - H + i\eta}$$

Plug in : 
$$\frac{1}{A-B} = \frac{1}{A} + \frac{1}{A}B\frac{1}{A-B}$$
 Us 
$$G = G^{(0)} + G^{(0)}VG$$
 
$$G = G^{(0)} + G^{(0)}VG^{(0)} + G^{(0)}VG^{(0)} + \dots$$

$$G = G^{(0)} + G^{(0)}VG$$

$$G = G^{(0)} + G^{(0)}VG^{(0)} + G^{(0)}VG^{(0)}VG^{(0)} + \dots$$

$$\langle \alpha | \frac{1}{E - H + i\eta} | \beta \rangle = \langle \alpha | \frac{1}{E - H_0 + i\eta} | \beta \rangle$$

$$+ \sum_{\alpha} \langle \alpha | \frac{1}{E - H_0 + i\eta} | \gamma \rangle \langle \gamma | V | \delta \rangle \langle \delta | \frac{1}{E - H + i\eta} | \beta \rangle$$

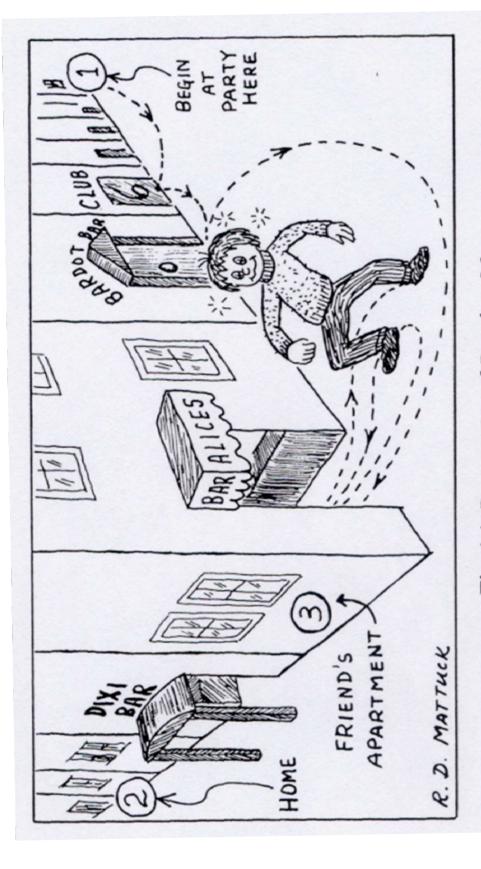
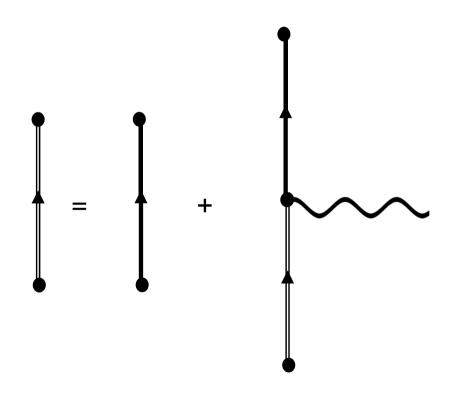


Fig. 1.1 Propagation of Drunken Man

(Reproduced with the kind permission of The Encyclopedia of Physics)

$$G = G^{(0)} + G^{(0)}VG$$

$$G = G^{(0)} + G^{(0)}VG^{(0)} + G^{(0)}VG^{(0)}VG^{(0)} + \dots$$



# N-body systems

N-body state can be written as

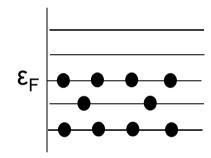
$$|\Phi_n\rangle = |\lambda_1 \lambda_2 .... \lambda_N\rangle = a_{\lambda_1}^{\dagger} a_{\lambda_2}^{\dagger} .... a_{\lambda_N}^{\dagger} |0\rangle$$

With energy

$$E_n = \sum_{i=1}^N \varepsilon_{\lambda_i}$$

The Fermi sea is written as

$$|\Phi_0\rangle = \prod_{\lambda_i \le F} a_{\lambda_i}^{\dagger} |0\rangle$$



$$\widehat{H}_0 a_{\alpha}^{\dagger} |\Phi_0\rangle = (E_0 + \varepsilon_{\alpha}) a_{\alpha}^{\dagger} |\Phi_0\rangle \quad \alpha > F$$

$$\widehat{H}_0 a_{\alpha} |\Phi_0\rangle = (E_0 - \varepsilon_{\alpha}) a_{\alpha}^{\dagger} |\Phi_0\rangle \quad \alpha < F$$

So now we also need hole propagation!

# N-body systems

In general, an N-body state can be written as

$$|\Phi_n\rangle = |\lambda_1 \lambda_2 .... \lambda_N\rangle = a_{\lambda_1}^{\dagger} a_{\lambda_2}^{\dagger} .... a_{\lambda_N}^{\dagger} |0\rangle$$

With energy

$$E_n = \sum_{i=1}^N \varepsilon_{\lambda_i}$$

The Fermi sea is written as

$$\widehat{H}_{0}a_{\alpha}^{\dagger} |\Phi_{0}\rangle = (E_{0} + \varepsilon_{\alpha}) a_{\alpha}^{\dagger} |\Phi_{0}\rangle \qquad \alpha > F 
\widehat{H}_{0}a_{\alpha} |\Phi_{0}\rangle = (E_{0} - \varepsilon_{\alpha}) a_{\alpha}^{\dagger} |\Phi_{0}\rangle \qquad \alpha < F 
\stackrel{\varepsilon_{F}}{\longleftarrow} \qquad \stackrel{\varepsilon_{F}}{\longleftarrow} \qquad$$

So now we also need hole propagation!

# N-body systems

In general, an N-body state can be written as

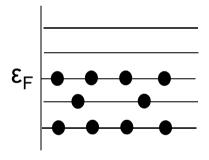
$$|\Phi_n\rangle = |\lambda_1 \lambda_2 .... \lambda_N\rangle = a_{\lambda_1}^{\dagger} a_{\lambda_2}^{\dagger} .... a_{\lambda_N}^{\dagger} |0\rangle$$

With energy

$$E_n = \sum_{i=1}^N \varepsilon_{\lambda_i}$$

The Fermi sea is written as

$$|\Phi_0\rangle = \prod_{\lambda_i \le F} a_{\lambda_i}^{\dagger} |0\rangle$$



$$\widehat{H}_{0}a_{\alpha}^{\dagger} |\Phi_{0}\rangle = (E_{0} + \varepsilon_{\alpha}) a_{\alpha}^{\dagger} |\Phi_{0}\rangle \qquad \alpha > F$$

$$\widehat{H}_{0}a_{\alpha} |\Phi_{0}\rangle = (E_{0} - \varepsilon_{\alpha}) a_{\alpha}^{\dagger} |\Phi_{0}\rangle \qquad \alpha < F \quad \varepsilon_{F}$$

So now we also need hole propagation!

Interacting system :  $H_0 \rightarrow H$ 

$$G(\alpha, \beta; E) = \langle 0 | a_{\alpha} \frac{1}{E - H + i\eta} a_{\beta}^{\dagger} | 0 \rangle$$
 Single-particle

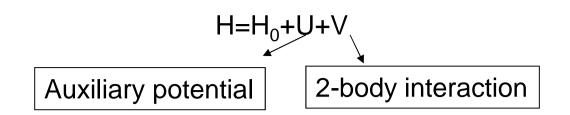
Particle states above Fermi

leve 
$$G(\alpha,\beta;E) = \sum_{m} \frac{\left\langle \Phi_{0}^{N} \left| \left| a_{\alpha} \left| \Phi_{m}^{N+1} \right\rangle \left\langle \Phi_{m}^{N+1} \right| \left| a_{\beta}^{\dagger} \left| \Phi_{0}^{N} \right\rangle \right. \right|}{E - \left( E_{m}^{N+1} - E_{0}^{N} \right) + i\eta} + \sum_{n} \frac{\left\langle \Phi_{0}^{N} \left| \left| a_{\beta}^{\dagger} \left| \Phi_{n}^{N-1} \right\rangle \left\langle \Phi_{n}^{N-1} \left| \left| a_{\alpha} \left| \Phi_{0}^{N} \right\rangle \right. \right|}{E - \left( E_{0}^{N} - E_{n}^{N-1} \right) - i\eta}$$

Hole states below Fermi level

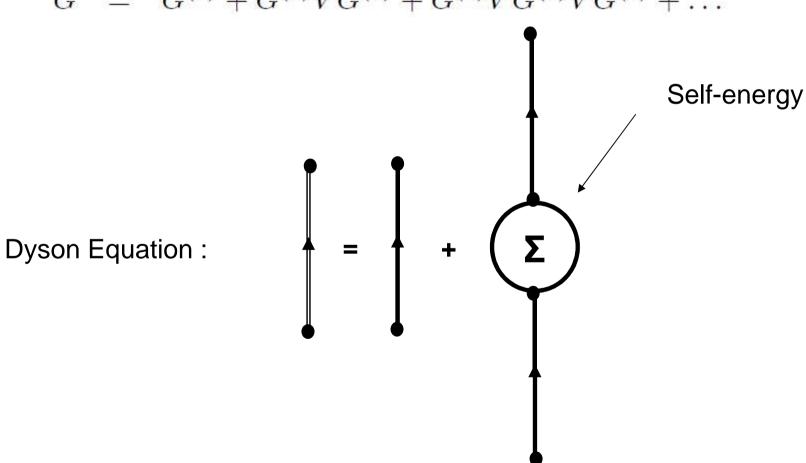
One particle (hole) in an N-body system

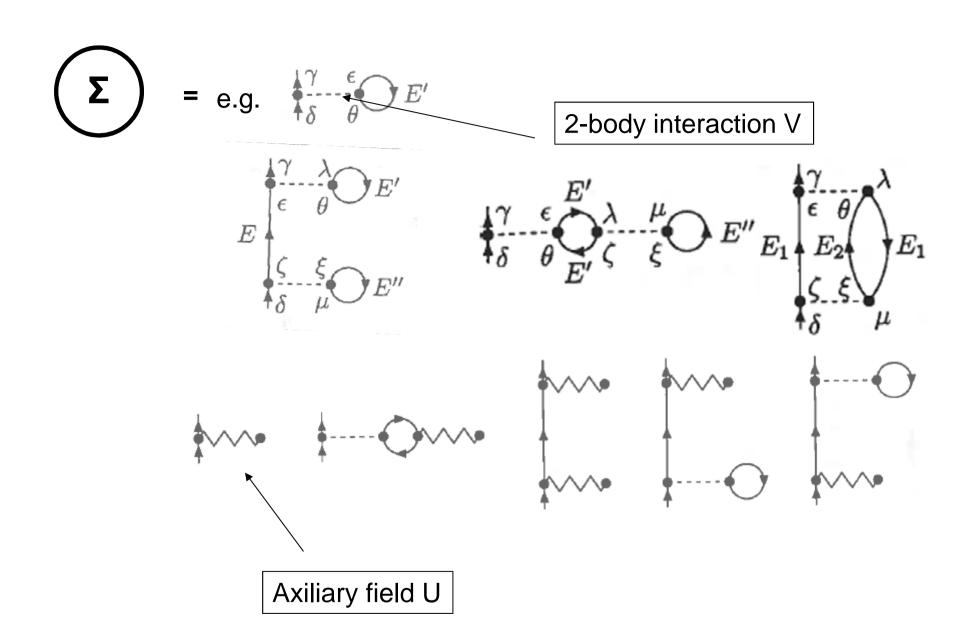
Lehmann representation of the single-particle progagator



$$G = G^{(0)} + G^{(0)}VG$$
  

$$G = G^{(0)} + G^{(0)}VG^{(0)} + G^{(0)}VG^{(0)}VG^{(0)} + \dots$$

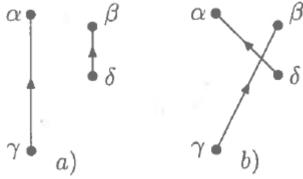




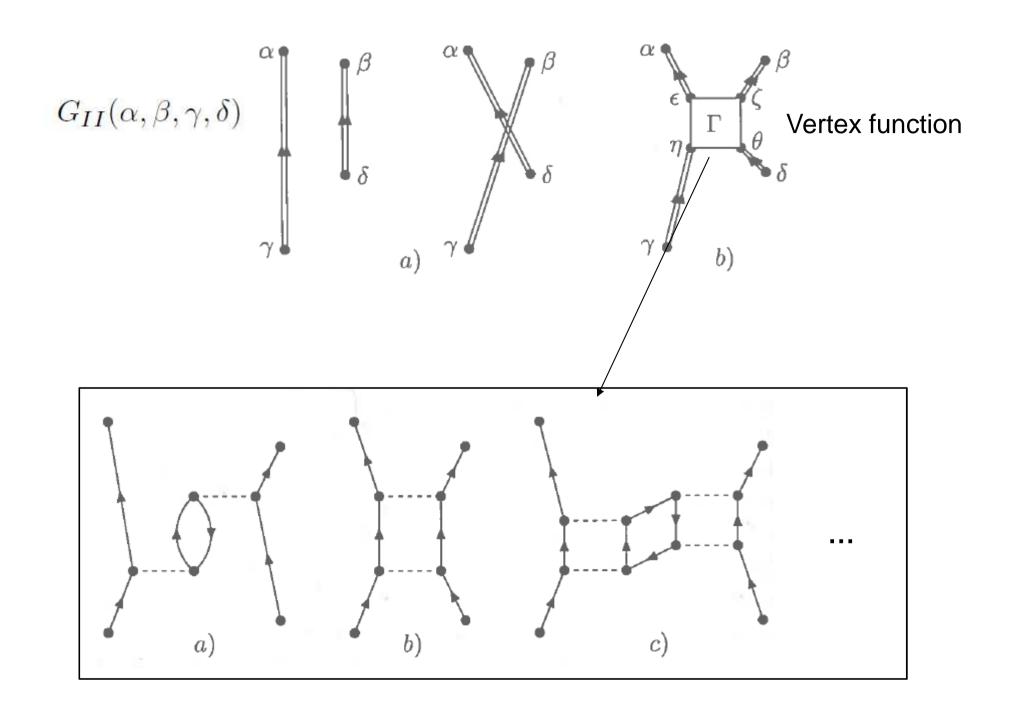
. . .

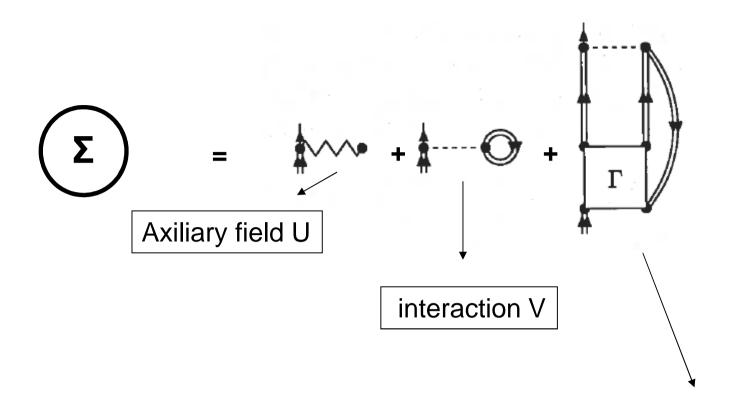
# Two-particle propagator

Non-interacting  $G_{II}^{(0)}(\alpha,\beta,\gamma,\delta)$ 



$$G_{II}^{(1)}(\alpha,\beta,\gamma,\delta)$$





Higher order contributions



Reformulation of the self-energy of a dressed particle

# Hartree-Fock propagator

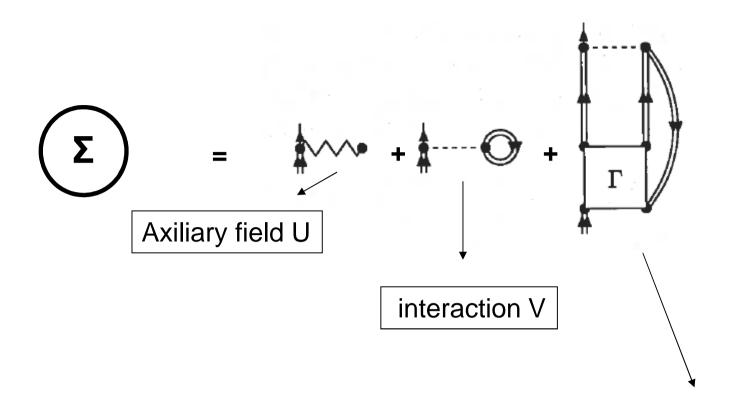
$$\widehat{H} = \widehat{T} + \widehat{V} = (\widehat{T} + \widehat{U}) + (\widehat{V} - \widehat{U})$$

$$\widehat{H}_0 = \widehat{T} + \widehat{U} = \sum_{\alpha} \varepsilon_{\alpha} a_{\alpha}^{\dagger} a_{\alpha}$$

$$G^{(0)}(\alpha, \beta; E) = \delta_{\alpha, \beta} \left[ \frac{\theta(\alpha - F)}{E - \varepsilon_{\alpha} + i\eta} + \frac{\theta(F - \alpha)}{E - \varepsilon_{\alpha} - i\eta} \right]$$

 $G^{HF}(\alpha, \beta; E) = G^{(0)}(\alpha, \beta; E) + \sum_{\gamma, \delta} G^{(0)}(\alpha, \gamma; E) \Sigma^{HF}(\gamma, \delta) G^{HF}(\delta, \beta; E)$ 

$$\Sigma^{HF}(\gamma, \delta; E) = -\langle \gamma | U | \delta \rangle - i \int \frac{dE'}{2\pi} \sum_{\mu\nu} \langle \gamma \mu | V | \delta \nu \rangle G^{HF}(\nu\mu; E')$$



Higher order contributions



Reformulation of the self-energy of a dressed particle

$$G^{HF}(\alpha, \beta; E) = G^{(0)}(\alpha, \beta; E) + \sum_{\gamma, \delta} G^{(0)}(\alpha, \gamma; E) \Sigma^{HF}(\gamma, \delta) G^{HF}(\delta, \beta; E)$$

$$\Sigma^{HF}(\gamma, \delta; E) = -\langle \gamma | U | \delta \rangle - i \int \frac{dE'}{2\pi} \sum_{\mu\nu} \langle \gamma \mu | V | \delta \nu \rangle G^{HF}(\nu\mu; E')$$



# Excited states: Particle-hole or polarization propagator

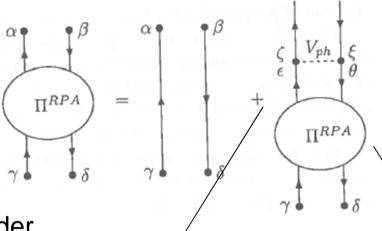
$$\Pi(\alpha, \beta^{-1}; \gamma, \delta^{-1}; E) = \sum_{n \neq 0} \frac{\langle \Phi_0^N | a_\beta^\dagger a_\alpha | \Phi_n^N \rangle \langle \Phi_n^N | a_\gamma^\dagger a_\delta | \Phi_0^N \rangle}{E - (E_n^N - E_0^N) + i\eta}$$

$$\epsilon_{\mathsf{F}} = \sum_{n \neq 0} \frac{\langle \Phi_0^N | a_\beta^\dagger a_\alpha | \Phi_n^N \rangle \langle \Phi_n^N | a_\beta^\dagger a_\alpha | \Phi_0^N \rangle}{E + (E_n^N - E_0^N) - i\eta}$$

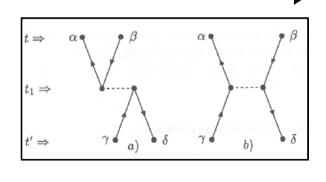
$$\Pi^{(0)}(\alpha, \beta^{-1}; \gamma, \delta^{-1}; E) = \int \frac{dE'}{2\pi} G^{(0)}(\alpha, \gamma; E + E') G^{(0)}(\delta^{-1}, \beta^{-1}; E')$$

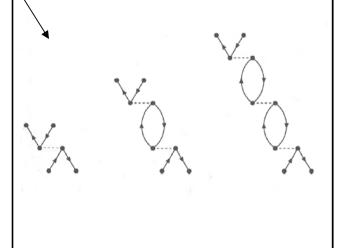
$$=$$
 $\begin{pmatrix} \alpha & & & \beta \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & &$ 

$$\Pi^{RPA}(\alpha, \beta^{-1}; \gamma, \delta^{-1}; E) = \Pi^{(0)}(\alpha, \beta^{-1}; \gamma, \delta^{-1}; E) 
+ \sum_{\varepsilon\theta\zeta\xi} \Pi^{(0)}(\alpha, \beta^{-1}; \zeta, \xi^{-1}; E) \langle \zeta\xi^{-1} | V_{ph} | \varepsilon\theta^{-1} \rangle \Pi^{RPA}(\varepsilon, \theta^{-1}; \gamma, \delta^{-1}; E)$$



First order  $\gamma$   $\downarrow$   $\delta$  Higher order RPA diagrams





# 2-body interaction V?

#### Landau-Migdal

$$V = c_0 \{ f_0(\rho) + f_0'(\rho)\vec{\tau}_1\vec{\tau}_2 + g_0(\rho)\vec{\sigma}_1\vec{\sigma}_2 + g_0'(\rho)\vec{\sigma}_1\vec{\sigma}_2\vec{\tau}_1\vec{\tau}_2 \}$$

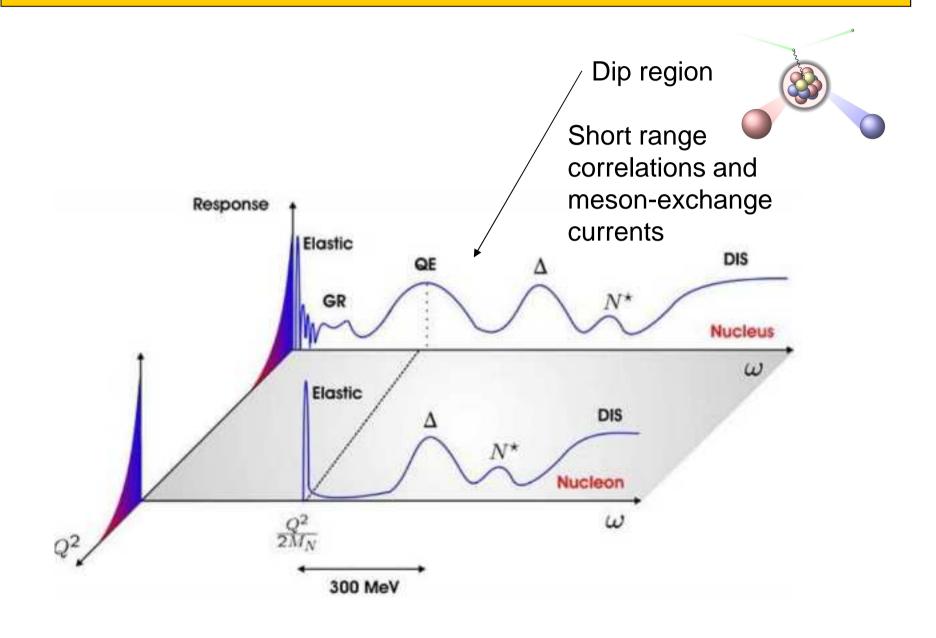
$$f(
ho(r)) = (1 - 
ho(r)) \ f^{ext} \ + \ 
ho(r) \ f^{int}$$
 
$$ho(r) \ = \ \frac{1}{1 + e^{rac{r - R}{a}}}$$

# 2-body interaction V?

### Skyrme

$$\begin{split} V(\vec{r_1},\vec{r_2}) &= t_0 \; (1+x_0 \hat{P}_\sigma) \; \delta(\vec{r_1}-\vec{r_2}) \\ &- \frac{1}{8} t_1 \left[ (\overleftarrow{\nabla}_1 - \overleftarrow{\nabla}_2)^2 \; \delta(\vec{r_1}-\vec{r_2}) + \delta(\vec{r_1}-\vec{r_2}) \; (\overrightarrow{\nabla}_1 - \overrightarrow{\nabla}_2)^2 \right] \\ &+ \frac{1}{4} t_2 (\overleftarrow{\nabla}_1 - \overleftarrow{\nabla}_2) \; \delta(\vec{r_1}-\vec{r_2}) \; (\overrightarrow{\nabla}_1 - \overrightarrow{\nabla}_2) \\ &+ i W_0 \; (\overrightarrow{\sigma}_1 + \overrightarrow{\sigma}_2) \cdot (\overleftarrow{\nabla}_1 - \overleftarrow{\nabla}_2) \times \delta(\vec{r_1}-\vec{r_2}) \; (\overrightarrow{\nabla}_1 - \overrightarrow{\nabla}_2) \\ &+ \frac{1}{6} t_3 \; (1-x_3) \; (1+\hat{P}_\sigma) \; \rho \; \frac{(\vec{r_1}+\vec{r_2})}{2} \; \delta(\vec{r_1}-\vec{r_2}) \\ &+ \frac{e^2}{|\vec{r_1}-\vec{r_2}|} \; + \; x_3 t_3 \; \delta(\vec{r_1}-\vec{r_2}) \; \delta(\vec{r_1}-\vec{r_3}) \\ &- \frac{1}{24} t_4 \left\{ \left[ (\overleftarrow{\nabla}_1 - \overleftarrow{\nabla}_2)^2 + (\overleftarrow{\nabla}_2 - \overleftarrow{\nabla}_3)^2 + (\overleftarrow{\nabla}_3 - \overleftarrow{\nabla}_1)^2 \right] \right\} \\ &\delta(\vec{r_1}-\vec{r_2}) \; \delta(\vec{r_1}-\vec{r_3}) + \; \delta(\vec{r_1}-\vec{r_2}) \; \delta(\vec{r_1}-\vec{r_3}) \\ &\left\{ \left[ (\overrightarrow{\nabla}_1 - \overrightarrow{\nabla}_2)^2 + (\overrightarrow{\nabla}_2 - \overrightarrow{\nabla}_3)^2 + (\overrightarrow{\nabla}_3 - \overleftarrow{\nabla}_1)^2 \right] \right\}. \end{split}$$

### Other correlations



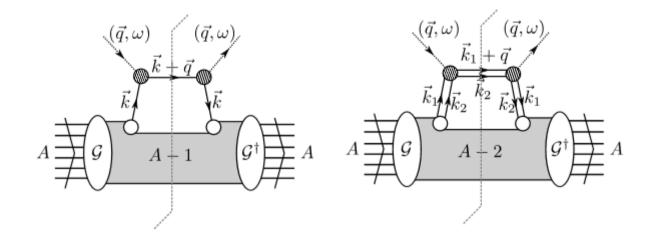
# Short-range correlations

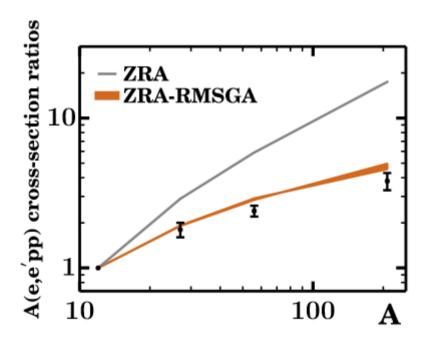


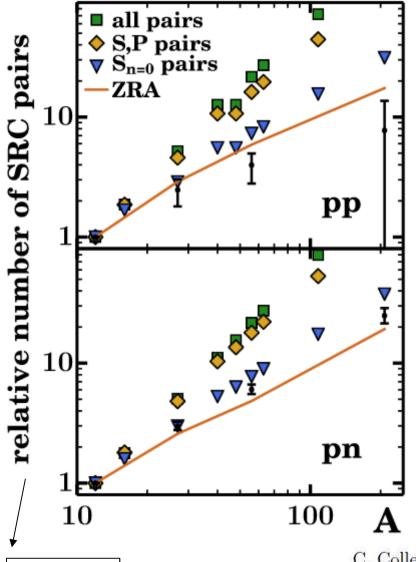
- The short-range repulsive character of this force, which correlates with the Pauli exclusion principle, results in a large mean free path of the nucleons with respect to the size of the nucleus
- In an independent particle model nucleons move independently from each other in a mean field
- This approach fails to capture short-range features of nucleon-nucleon correlations
- SRC: short-range repulsive, tensor component of the nuclear force
- Individual nucleons receive large momenta compared to the Fermi momentum

IPM single-particle orbitals are depleted by SRC\* and higher energy levels are populated

- Short-range correlations affect 1nucleon and 2nucleon knockout processes
- The SRC-prone nucleon pairs are predominantly in a back-to-back configuration with a small center-of-mass and high relative momentum
- Final-state interactions for the outgoing nucleons affect the experimental observations







To <sup>12</sup>C

- mass dependence ~ A<sup>1.12</sup>:soft!
- Predominantly pn, s-pairs
- Universal over mass tabel
- Tensor force dominates at short distances

C. Colle, W. Cosyn, and J. Ryckebusch, Phys. Rev. C 93, 034608 (2016)
 C. Colle, W. Cosyn, J. Ryckebusch, and M. Vanhalst, Phys. Rev. C 89, 024603 (2014)

024604 (2015).

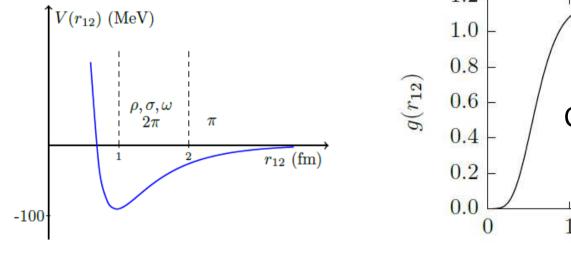
C. Colle, O. Hen, W. Cosyn, I. Korover, E. Piasetzky, J. Ryckebusch, and L. B. Weinstein, Phys. Rev. C 92,

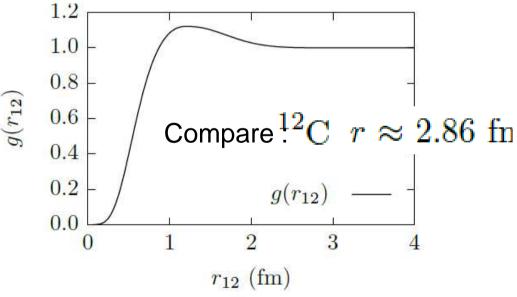
Data: CLAS A(e,e',pN) data

## Two-body density:

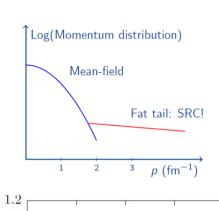
$$ho^{[2]}(m{r}_1,m{r}_2)= 
ho^{[1]}(m{r}_1)
ho^{[1]}(m{r}_2)g(r_{12})$$
 Independent particle model

#### Correlation function





# Short-range correlations



spin-isospin (x5)

0.8

0.4

0.0

 $f(r_{12})$ 

$$|\Psi\rangle = \frac{1}{\sqrt{\mathcal{N}}}\widehat{\mathcal{G}}|\Phi\rangle$$

with

$$\widehat{\mathcal{G}} \approx \widehat{\mathcal{S}} \left( \prod_{i < j}^{A} \left[ 1 + \widehat{l}(i, j) \right] \right)$$

$$\widehat{l}(i,j) = -g_c(r_{ij}) + f_{\sigma\tau}(r_{ij}) \left( \vec{\sigma}_i \cdot \vec{\sigma}_j \right) \left( \vec{\tau}_i \cdot \vec{\tau}_j \right)$$

Shifting the complexity induced by correlations from the wave functions to the operators

$$+ f_{t\tau}(r_{ij}) \widehat{S}_{ij} \left( \vec{\tau}_i \cdot \vec{\tau}_j \right),\,$$

$$\langle \Psi_{\rm f} | \widehat{J}_{\mu}^{\rm nucl} | \Psi_{\rm i} \rangle = \frac{1}{\sqrt{\mathcal{N}_{\rm i} \mathcal{N}_{\rm f}}} \langle \Phi_{\rm f} | \widehat{J}_{\mu}^{\rm eff} | \Phi_{\rm i} \rangle$$

$$r_{12}$$
 (fm)
$$10^{-2}$$

$$10^{-3}$$

$$10^{-4}$$

$$10^{-6}$$

$$0$$

$$200$$

$$400$$

$$600$$

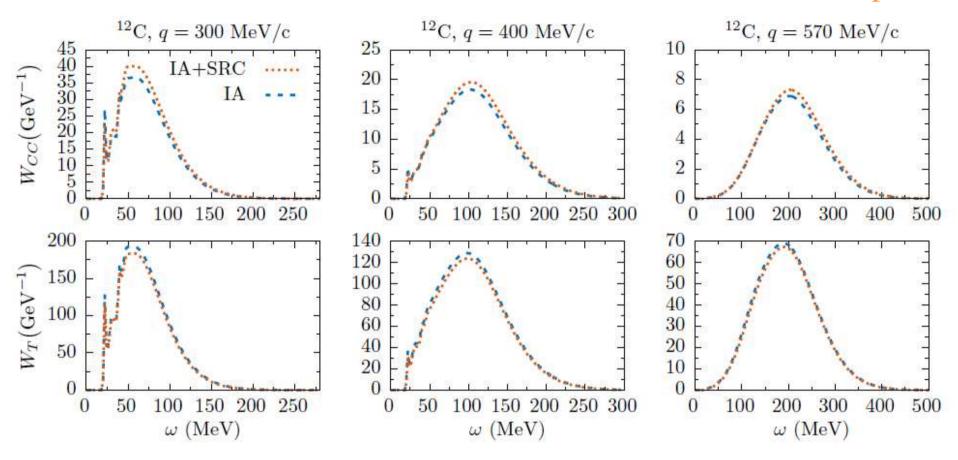
$$800$$

$$k_{12}$$
 (MeV/c)

$$\widehat{J}_{\mu}^{\text{eff}} \approx \sum_{i=1}^{A} \widehat{J}_{\mu}^{[1]}(i) + \sum_{i < j}^{A} \widehat{J}_{\mu}^{[1], \text{in}}(i, j) + \left[ \sum_{i < j}^{A} \widehat{J}_{\mu}^{[1], \text{in}}(i, j) \right]^{\dagger}$$

$$\widehat{J}_{\mu}^{[1],\text{in}}(i,j) = \left[\widehat{J}_{\mu}^{[1]}(i) + \widehat{J}_{\mu}^{[1]}(j)\right] \widehat{l}(i,j)$$

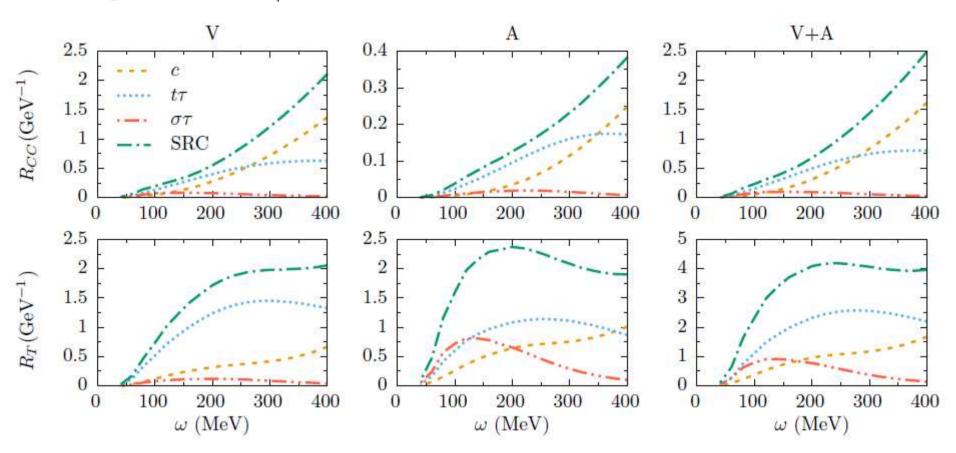
# SRC neutrinos 1p1h



- Reduction of transverse response
- Enhancement of Coulomb-longitudinal

## SRC neutrinos 2p2h

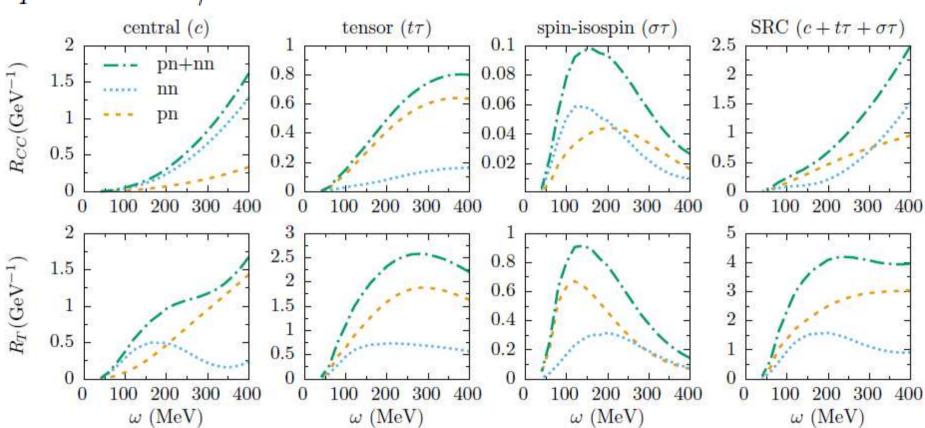
$$q = 400 \text{ MeV/c}$$



- Vector and axial contributions have comparable strength
- Tensor often dominates, but not for all kinematics

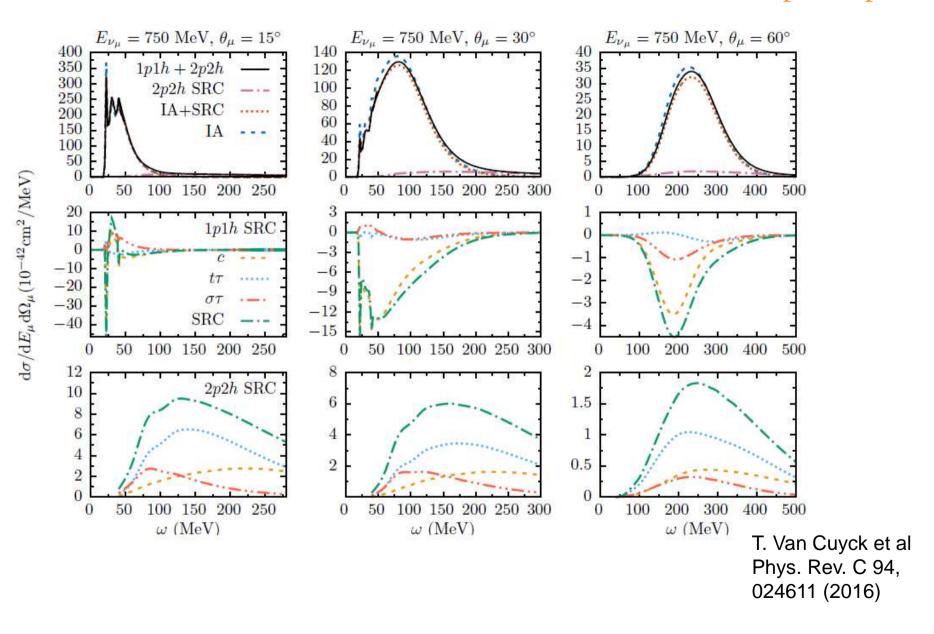
# SRC neutrinos 2p2h

q = 400 MeV/c



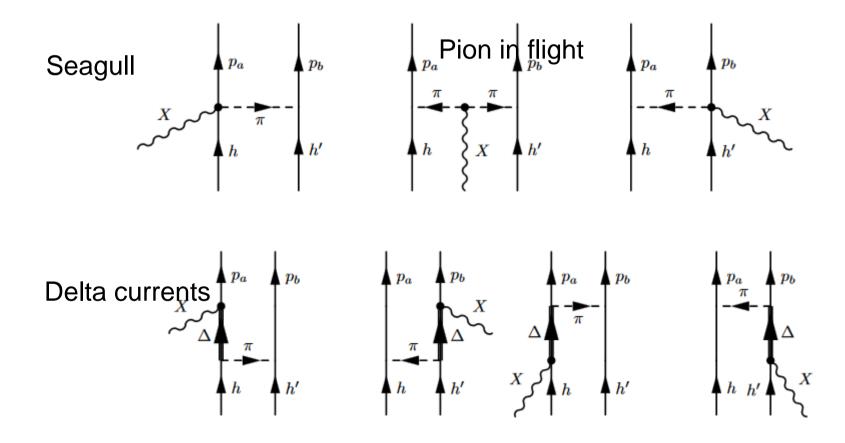
- Vector and axial contributions have comparable strength
- Tensor often dominates, but not for all kinematics
- pn pairs dominate

# SRC neutrinos 1p1h+2p2h

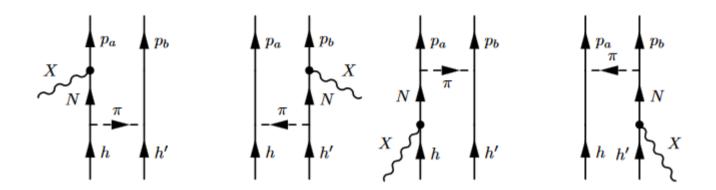


# Meson-exchange currents

When an electroweak boson interacts with a pair of nucleons which are correlated through the exchange of a meson, this will cause the knockout of one or both of the particles from the nucleus. The boson was interacting with a current consisting of two nucleons, a two-body current, called a MEC



#### Correlation currents

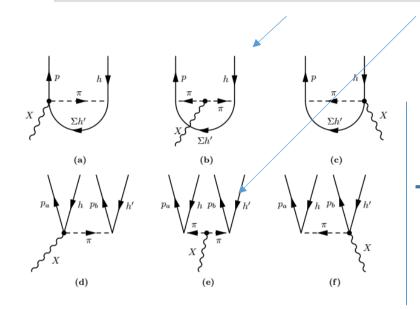


Already included in mean field models!

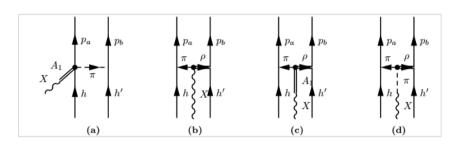
#### Contributions of heavier mesons:

m<sub>π</sub>≈135MeV, m<sub>ρ</sub>≈775MeV, m<sub>ω</sub>≈782MeV

# MEC in 1p1h and 2p2h



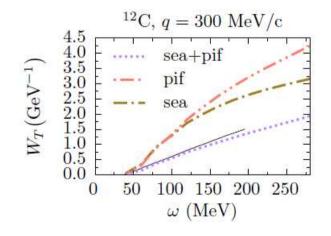
#### Axial contributions:

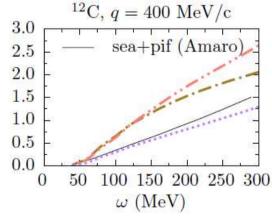


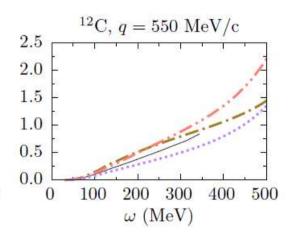
$$\widehat{\rho}_{A}^{[2],\text{axi}}(\boldsymbol{q}) = \frac{i}{g_{A}} \left( \frac{f_{\pi NN}}{m_{\pi}} \right)^{2} (\boldsymbol{I}_{V}) \left( F_{\pi}(\boldsymbol{q}_{2}^{2}) \Gamma_{\pi}^{2}(\boldsymbol{q}_{2}^{2}) \frac{\boldsymbol{\sigma}_{2} \cdot \boldsymbol{q}_{2}}{\boldsymbol{q}_{2}^{2} + m_{\pi}^{2}} - F_{\pi}(\boldsymbol{q}_{1}^{2}) \Gamma_{\pi}^{2}(\boldsymbol{q}_{1}^{2}) \frac{\boldsymbol{\sigma}_{1} \cdot \boldsymbol{q}_{1}}{\boldsymbol{q}_{1}^{2} + m_{\pi}^{2}} \right)$$

I. Towner, Nucl. Phys.A542, 631 (1992)

<sup>12</sup>C(e, e')





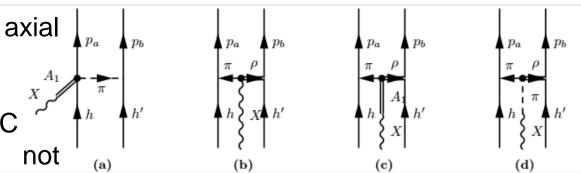


Only seagull have counterpart

timelike

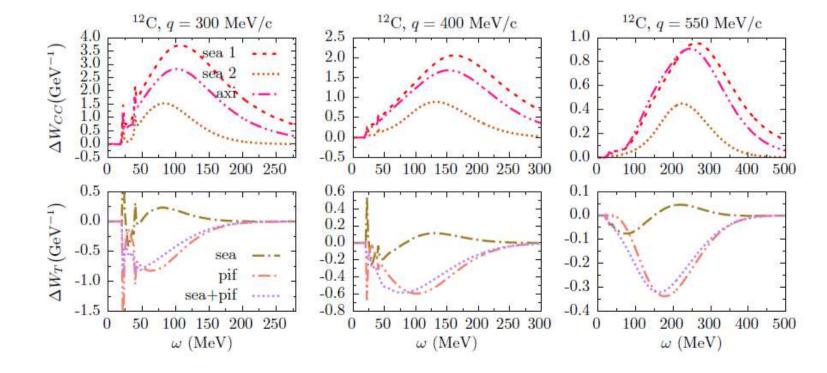
Partially constrained by PCAC

 Non-relativistic reduction unambiguous

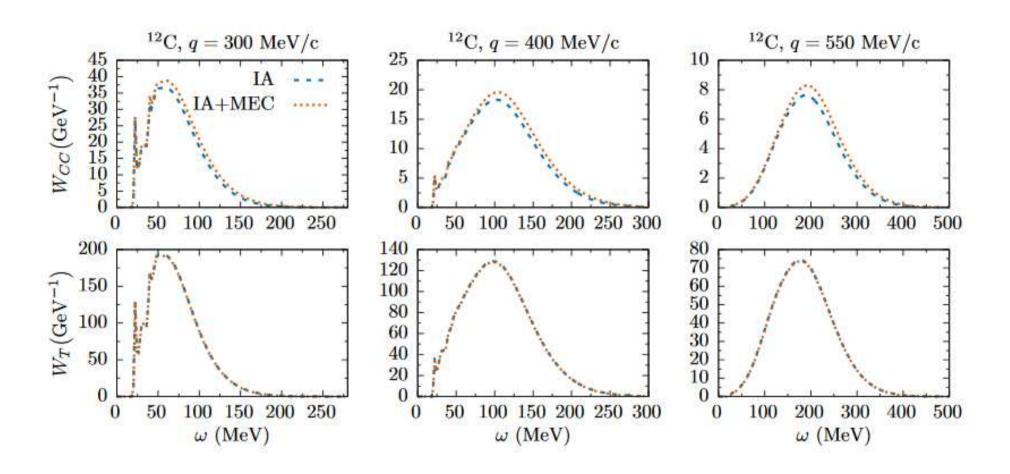


I. Towner, Nucl. Phys.A542, 631 (1992)

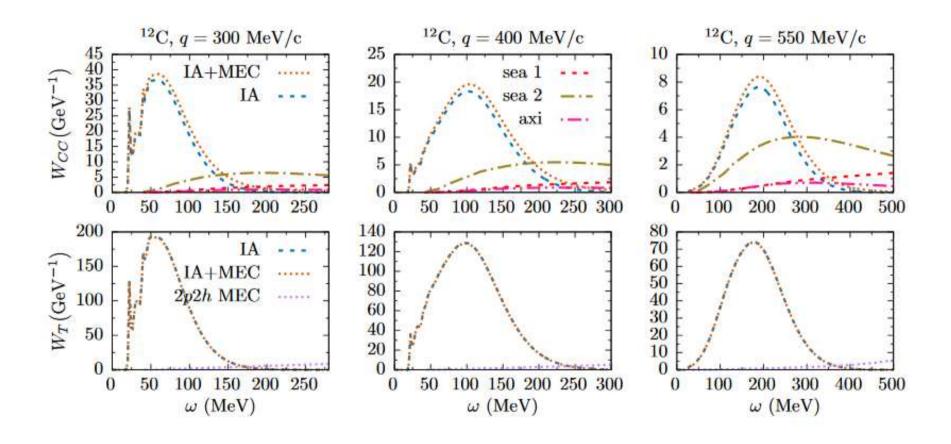
$$\widehat{\rho}_{A}^{[2],\text{axi}}(\boldsymbol{q}) = \frac{i}{g_{A}} \left( \frac{f_{\pi NN}}{m_{\pi}} \right)^{2} (\boldsymbol{I}_{V}) \left( F_{\pi}(\boldsymbol{q}_{2}^{2}) \Gamma_{\pi}^{2}(\boldsymbol{q}_{2}^{2}) \frac{\boldsymbol{\sigma}_{2} \cdot \boldsymbol{q}_{2}}{\boldsymbol{q}_{2}^{2} + m_{\pi}^{2}} - F_{\pi}(\boldsymbol{q}_{1}^{2}) \Gamma_{\pi}^{2}(\boldsymbol{q}_{1}^{2}) \frac{\boldsymbol{\sigma}_{1} \cdot \boldsymbol{q}_{1}}{\boldsymbol{q}_{1}^{2} + m_{\pi}^{2}} \right)$$



# Seagull and pif in1p1h



# Seagull and pif in 2p2h



T. Van Cuyck, N. Jachowicz, R. González-Jiménez, et al. PRC95, 054611 (2017)

#### 2-nucleon knockout cross sections

#### 2-nucleon knockout:

$$\begin{split} \frac{\mathrm{d}\sigma^X}{\mathrm{d}E_f\mathrm{d}\Omega_f\mathrm{d}T_b\mathrm{d}\Omega_b\mathrm{d}\Omega_a} &= \frac{p_ap_bE_aE_b}{(2\pi)^6}g_{rec}^{-1}\sigma^X\zeta \\ &\qquad \times \left[v_{CC}W_{CC} + v_{CL}W_{CL} + v_{LL}W_{LL} + v_TW_T + v_{TT}W_{TT} + v_{TC}W_{TC} + v_{TL}W_{TL} + h(v_{T'}W_{T'} + v_{TC'}W_{TC'} + v_{TL'}W_{TL'})\right], \end{split} \tag{4}. \end{split}$$
 with: 
$$\mathcal{J}_{\lambda} = \langle \Phi_f^{(A-2)}(E^{exc}, J_RM_R); \; \boldsymbol{p}_a m_{s_a}; \; \boldsymbol{p}_b m_{s_b} \, | \, \hat{J}_{\lambda}(\boldsymbol{q}) \, | \, \Phi_{gs} \, \rangle \\ &\qquad | \, \Phi^{2p2h} \, \rangle = | \, \Phi_f^{(A-2)}(E^{exc}, J_RM_R); \; \boldsymbol{p}_a m_{s_a}; \; \boldsymbol{p}_b m_{s_b} \, \rangle_{as} \end{split}$$

#### Some issues:

- 7-dimensional integrals
- Need to fold over incoming neutrino flux!
- Huge numbers of diagrams
- Ambiguities in non-relativistic reduction of contributions
- Little constraints for axial parameters

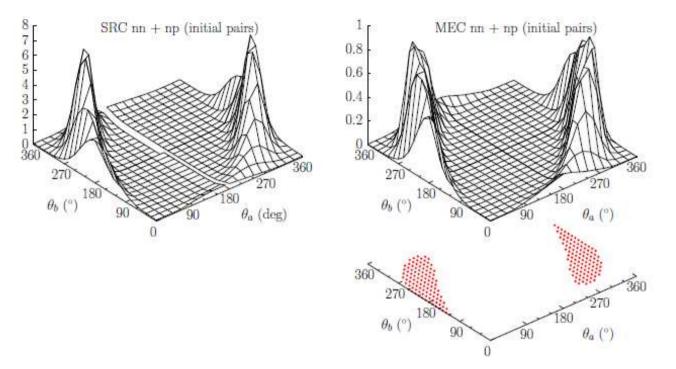


Figure 4.5: The  $^{12}\text{C}(\nu_{\mu}, \mu^- N_a N_b)$  cross section  $(N_a = \text{p}, N_b = \text{p'}, \text{n})$  at  $\epsilon_{\nu_{\mu}} = 750$  MeV,  $\epsilon_{\mu} = 550$  MeV,  $\theta_{\mu} = 15^{\circ}$  and  $T_{\text{p}} = 50$  MeV for in-plane kinematics. Left with SRCs, right with MECs, the bottom plot shows the  $(\theta_a, \theta_b)$  regions with  $P_{12} < 300$  MeV/c.

- Strength residing in restricted part of phase space
- $ullet p_bpprox p_b^{ave}$
- Quasi-deuteron kinematics

#### Some issues:

- 7-dimensional integrals
- Need to fold over incoming neutrino flux!
- Huge numbers of diagrams
- divergences
- Ambiguities in non-relativistic reduction of contributions
- Little constraints for axial parameters

#### Approximation schemes:

- Reduce number or range of integrations :
  - Frozen approximation (Ruiz-Simo, Amaro et.a l ArXiv1703.1018)
  - Modified convolution approximation (Ruiz-Simo, Amaro et.al, ArXiv1706.06377)
- Choice of subset of diagrams and terms

#### What about correlations on the neutrino market?

 Neutrino-Nucleus Cross Sections for Oscillation Experiments

Teppei Katori and Marco Martini

arXiv:1611.07770

 NuSTEC White Paper: Status and Challenges of Neutrino-Nucleus Scattering

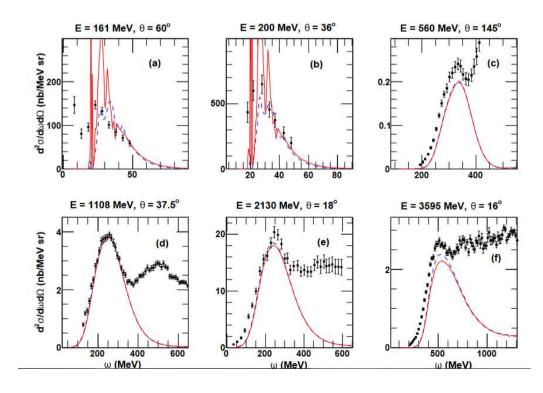
arXiv:1706.03621

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# What about correlations on the neutrino market?

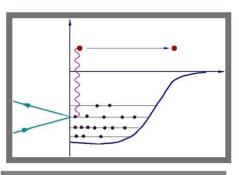
# **Ghent model**

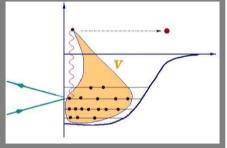
- Hartree-Fock mean field
- Continuum RPA
- Residual interaction : Skyrme



#### **CRPA**

- •Green's function approach
- •Skyrme SkE2 residual interaction
- •self-consistent calculations





$$= \left| \begin{array}{c} \\ \\ \\ \end{array} \right| + \left| \begin{array}{c} \\ \\ \end{array} \right| + \left| \left| \begin{array}{c} \\ \\ \end{array} \right| + \left| \left| \begin{array}{c} \\ \\ \end{array} \right| + \left| \left| \begin{array}{c} \\ \\ \end{array} \right| + \left| \left| \left| \begin{array}{c} \\ \\ \end{array} \right| + \left| \left| \left| \left|$$

$$\Pi^{(RPA)}(x_1, x_2; \omega) = \Pi^{(0)}(x_1, x_2; \omega) + \frac{1}{\hbar} \int dx \int dx' \ \Pi^{(0)}(x_1, x; \omega) \ \widetilde{V}(x, x') \ \Pi^{(RPA)}(x', x_2; \omega)$$

Solving the RPA equations in coordinate space

$$|\Psi_{C}(E)\rangle = \left|ph^{-1}(E)\right\rangle + \int dx_{1} \int dx_{2} \ \tilde{V}(x_{1}, x_{2})$$

$$\sum_{c'} \mathcal{P} \int d\varepsilon_{p'} \left[ \frac{\psi_{h'}(x_{1})\psi_{p'}^{\dagger}(x_{1}, \varepsilon_{p'})}{E - \varepsilon_{p'h'}} \left| p'h'^{-1}(\varepsilon_{p'h'}) \right\rangle \right]$$

$$- \frac{\psi_{h'}^{\dagger}(x_{1})\psi_{p'}(x_{1}, \varepsilon_{p'})}{E + \varepsilon_{p'h'}} \left| h'p'^{-1}(-\varepsilon_{p'h'}) \right\rangle \right] \left\langle \Psi_{0} \left| \hat{\psi}^{\dagger}(x_{2})\hat{\psi}(x_{2}) \right| \Psi_{C}(E) \right\rangle$$

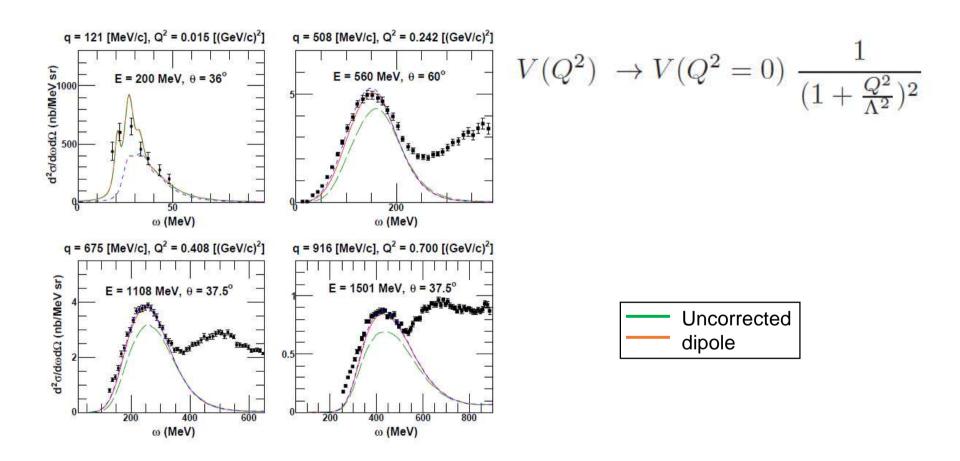
What we really need is transition densities:

$$\begin{split} \langle \Psi_{0} || X_{\eta J} || \Psi_{C}(J; E) \rangle_{r} &= - \langle h || X_{\eta J} || p(\varepsilon_{ph}) \rangle_{r} \\ &+ \sum_{\mu, \nu} \int dr_{1} \int dr_{2} \ U_{\mu\nu}^{J}(r_{1}, r_{2}) \ \mathcal{R} \left( R_{\eta\mu; J}^{(0)}(r, r_{1}; E) \right) \ \langle \Psi_{0} || X_{\nu J} || \Psi_{C}(J; E) \rangle_{r_{2}} \end{split}$$

So in the end we have to solve a set of coupled equations, that after discretizing on a mesh in coordinate space, translates into a matrix inversion:

$$\rho_C^{RPA} = -\frac{1}{1 - R U} \rho_C^{HF}$$

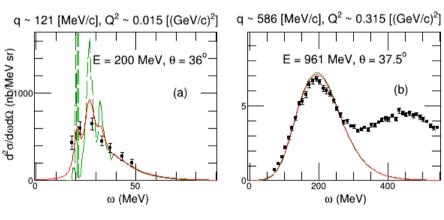
•Regularization of the residual interaction:



#### • Final state interactions:

-taken into account through the calculations of the wave function of the outgoing nucleon in the (real) nuclear potential generated using the Skyrme force

-influence of the spreading width of the particle states is implemented through a folding procedure



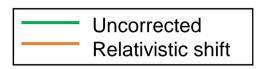
Bare RPA

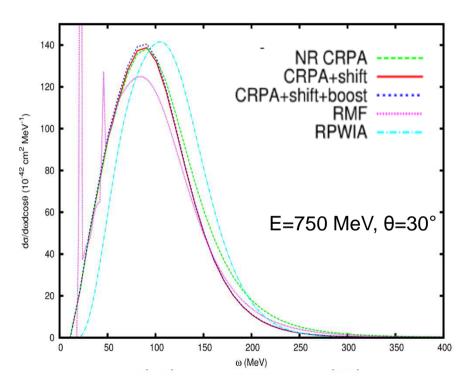
folding

$$R'(q,\omega') = \int_{-\infty}^{\infty} d\omega \ R(q,\omega) \ L(\omega,\omega'),$$

$$L(\omega, \omega') = \frac{1}{2\pi} \left[ \frac{\Gamma}{(\omega - \omega')^2 + (\Gamma/2)^2} \right].$$

•Relativistic corrections at higher energies (S. Jeschonnek and T. Donnelly, PRC57, 2438 (1998)):





#### Shift:

$$\lambda \to \lambda(\lambda + 1)$$
  $\lambda = \omega/2M_N$ 

 The outgoing nucleon obtains the correct relativistic momentum

$$p = \sqrt{T^2 + 2MT}$$

 Shifts the QE peak to the correct relativistic position

#### Boost:

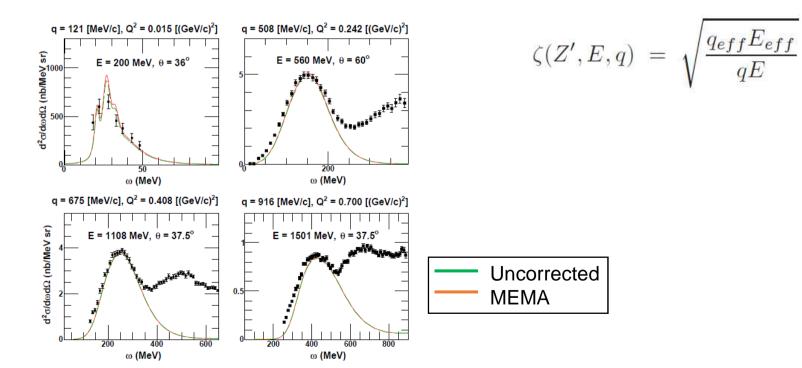
$$egin{aligned} R_{ ext{CC}}^{ ext{V}}(q,\omega) &
ightarrow rac{q^2}{q^2-\omega^2} \, R_{ ext{CC}}^{ ext{V}}(q,\omega) \,, \ R_{ ext{LL}}^{ ext{A}}(q,\omega) &
ightarrow \left(1+rac{q^2-\omega^2}{4m^2}
ight) \, R_{ ext{LL}}^{ ext{A}}(q,\omega) \,, \ R_{ ext{T}}^{ ext{V}}(q,\omega) &
ightarrow rac{q^2-\omega^2}{q^2} \, R_{ ext{T}}^{ ext{V}}(q,\omega) \,, \ R_{ ext{T}}^{ ext{A}}(q,\omega) &
ightarrow \left(1+rac{q^2-\omega^2}{4m^2}
ight) \, R_{ ext{T}}^{ ext{A}}(q,\omega) \,, \ R_{ ext{T}'}^{ ext{VA}}(q,\omega) &
ightarrow \sqrt{rac{q^2-\omega^2}{q^2}} \, \sqrt{1+rac{q^2-\omega^2}{4m^2}} \, R_{ ext{T}'}^{ ext{VA}}(q,\omega) \,. \end{aligned}$$

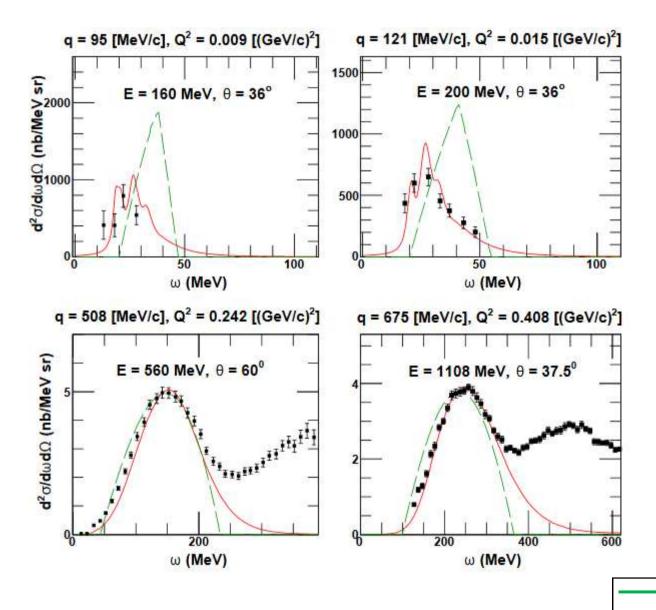
- •Coulomb correction for the outgoing lepton in charged-current interactions :
  - ✓ Low energies : Fermi function

$$F(Z', E) = \frac{2\pi\eta}{1 - e^{-2\pi\eta}} \qquad \eta \sim \mp Z'\alpha$$

✓ High energies : modified effective momentum approximation (J. Engel, PRC57.2004 (1998))

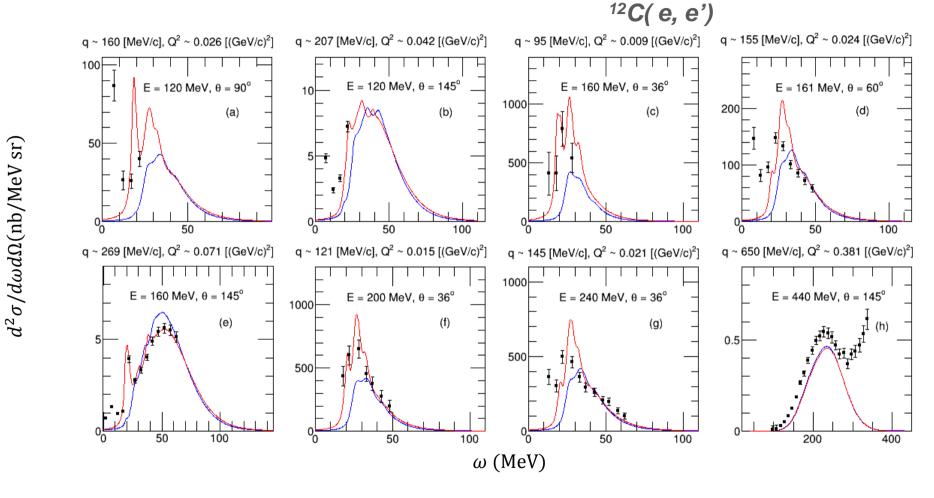
$$q_{eff} = q + 1.5 \left(\frac{Z'\alpha\hbar c}{R}\right), \qquad \Psi_l^{eff} = \zeta(Z', E, q) \Psi_l,$$



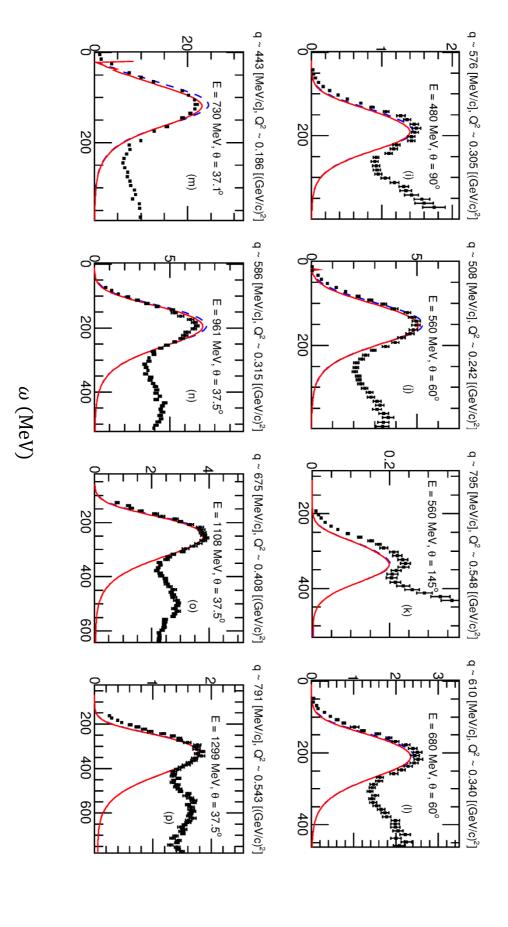


Fermi gas

Hartree-Fock



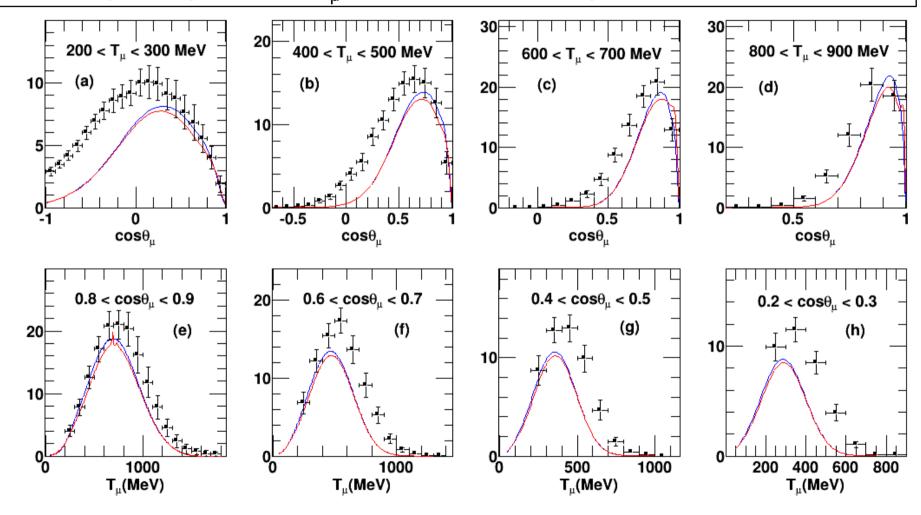
P. Barreau et al., Nucl. Phys. A402, 515 (1983), J. S. O'Connell et al., Phys. Rev. C35, 1063 (1987), R. M. Sealock et al., Phys. Rev. Lett.62, 1350 (1989).,D. S. Bagdasaryan et al., YERPHI-1077-40-88 (1988),D. B. Day et al., Phys. Rev. C 48, 1849 (1993).,D. Zeller, DESY-F23-73-2 (1973).



V. Pandey, N. Jachowicz, M. Martini, et al. Phys. Rev. C94, 054609 (2016)

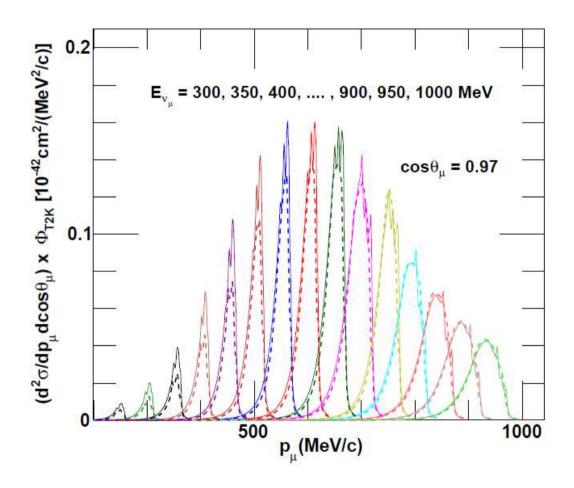
# MiniBooNe $v_{\mu}$

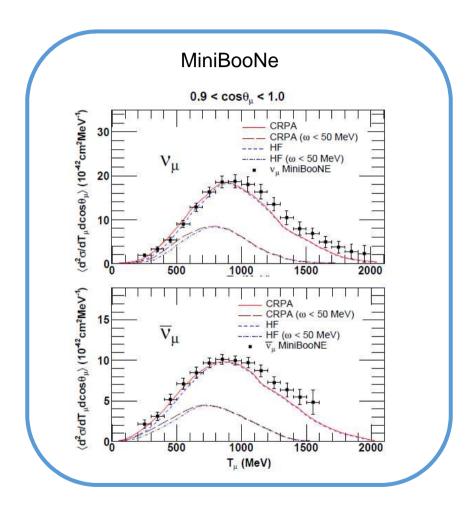
- Satisfactory general agreement
- Good agreement for forward scattering
- Missing strength for low T<sub>µ</sub>, backward scattering

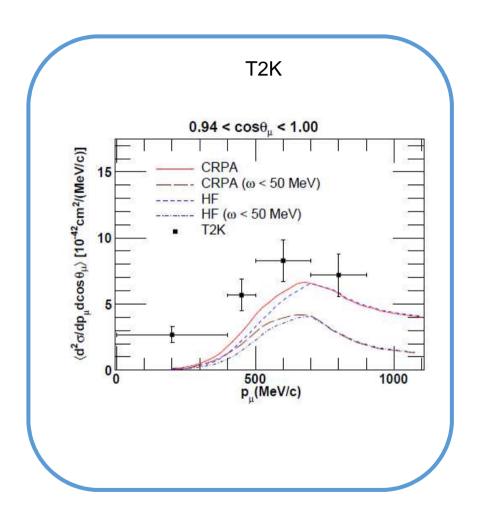


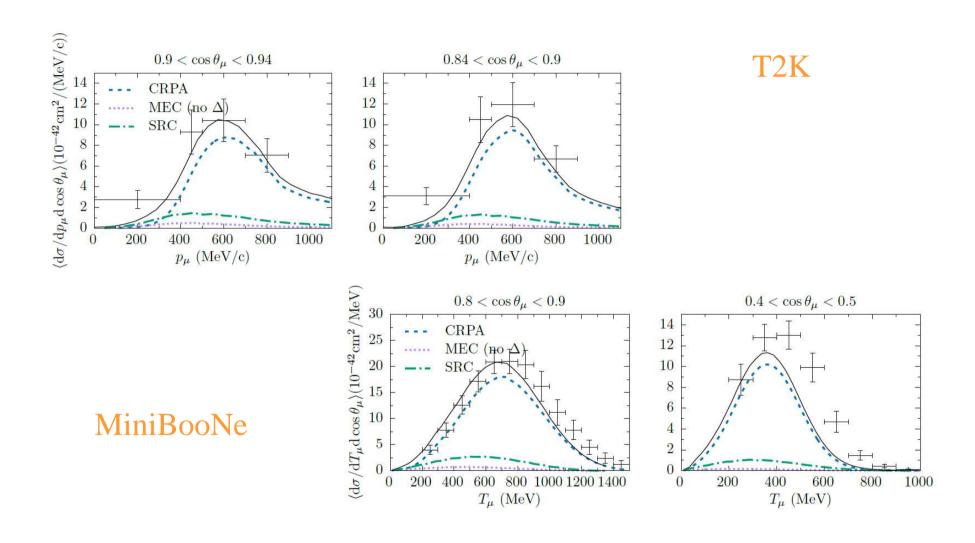
V. Pandey et al, Phys. Rev. C 92, 024606 (2015)

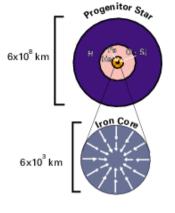
# Forward scattering

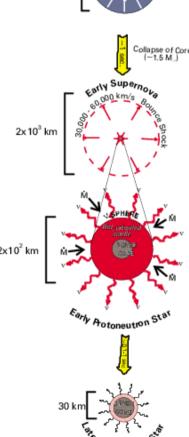






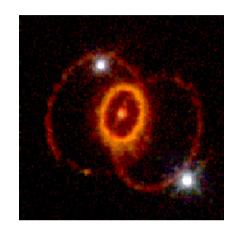


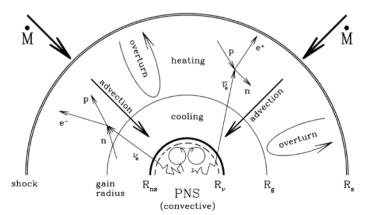




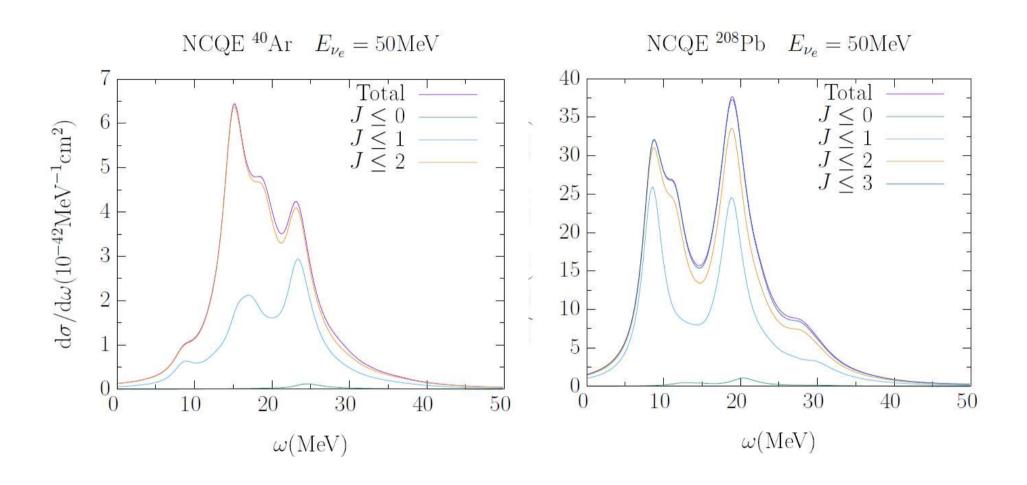
# Neutrinos in a core-collapse supernova

- weak interactions are important neutrinos are produced in the neutronization processes characterizing the gravitational collapse
- neutrinos are responsible for the cooling of the proto-neutron star
- neutrinonucleosynthesis
- energy deposition by neutrinos might reheat the stalled shock wave and cause a delayed explosion
- terrestrial detection of supernova neutrinos



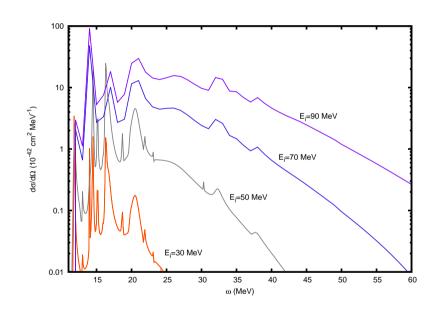


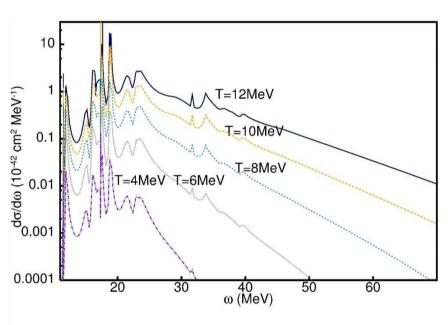
# Neutrino scattering results at low energies:



# Folded cross sections supernova neutrino spectra:

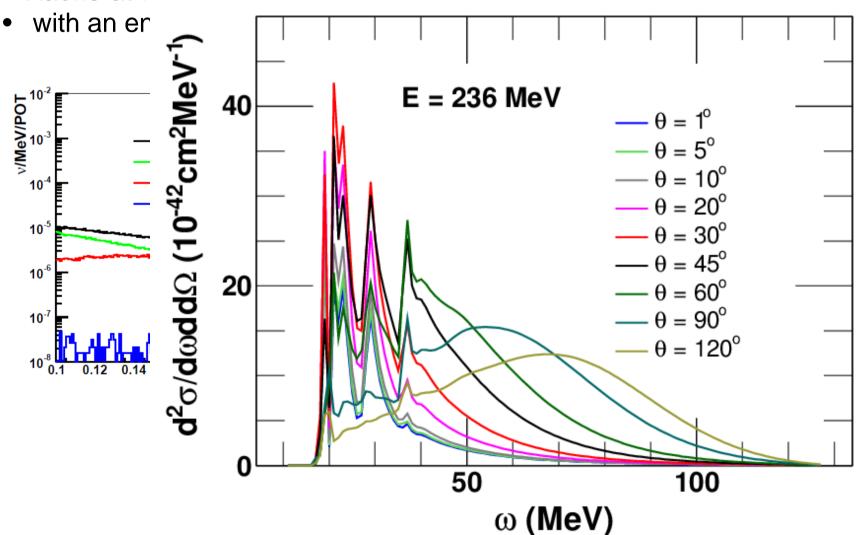
<sup>16</sup>O(v, v') <sup>16</sup>O\*





# 236 MeV neutrinos

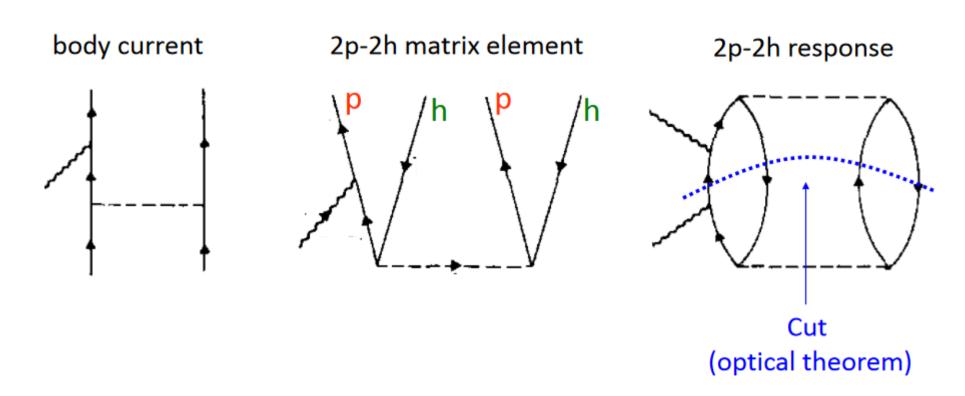
- Protons on Carbon generate Kaons
- Kaons-at-rest- decay ... primarily in v...



Different diagrammatic languages are used ...

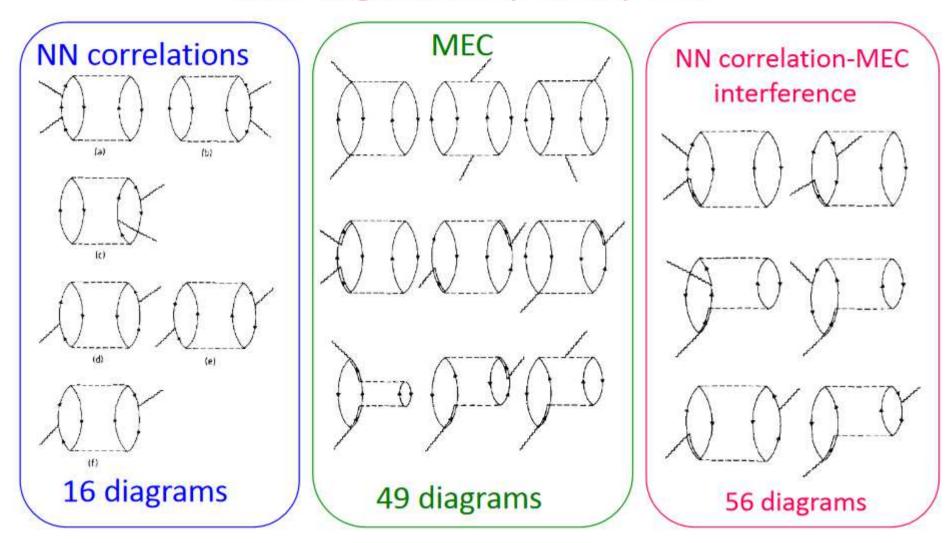
# Two particle-two hole sector (2p-2h)

### Three equivalent representations of the same process



Final state: two particles-two holes

# Some diagrams for 2p-2h responses



(Some already included in a mean field)

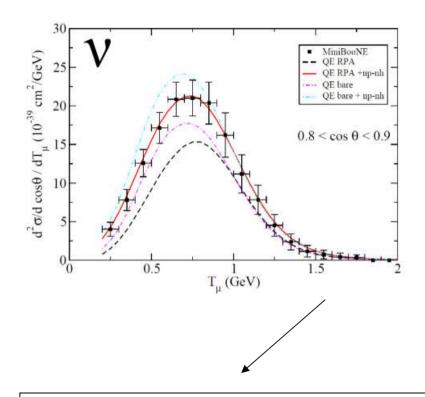
#### What about correlations in models on the neutrino market?

### Martini, Ericson model

- Starting point :Local Fermi gas, relativized
- RPA equations are solved in momentum space : Lindhard equation
- Correlations are introduced with a Landau-Migdal interaction : effective parametrization of pion- and exchange, contact force, non-relativistic reduction for MEC
- Includes delta degrees of freedom in RPA ( Ericson-Ericson Lorentz-Lorentz effect)

References (non-exhaustive list): PRC 80, 065501 (2009); PRC 81, 045502 (2010); PRC 84, 055502 (2011); PRD 85, 093012 (2012); PRD 87, 013009 (2013); PRC 87, 065501 (2013); PRC 90, 025501 (2014); PRC 91, 035501 (2015); PRC 94, 015501 (2016), etc.

# Martini, Ericson model



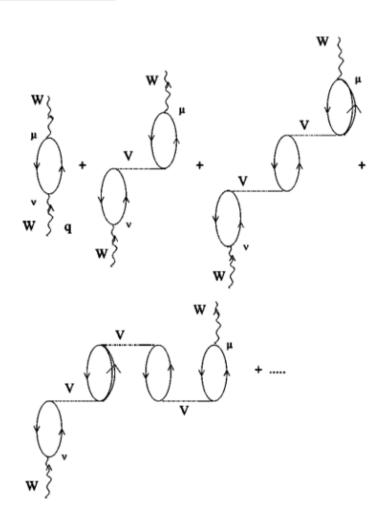
First to explain MiniBooNE QE data including np-nh effects, M\_A~ 1



#### What about RPAs on the neutrino market?

# Valencia model (Nieves et al.)

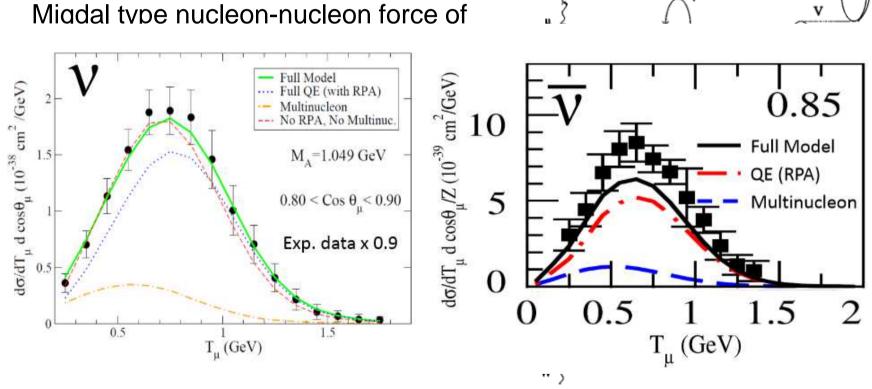
- Starting point :Local Fermi gas
- MEC, correlation with an effective Landau-Migdal type nucleon-nucleon force of pion-, and rho meson correlation currents
- relativistic
- Includes delta degrees of freedom
- Dressed nucleon propagators in the nuclear medium~ spectral function
- Includes SRC effects and 2p-2h degrees of freedom



#### What about RPAs on the neutrino market?

# Valencia model (Nieves et al.)

- Starting point :Local Fermi gas, relativized
- Correlations with an effective Landau-Middal type nucleon-nucleon force of

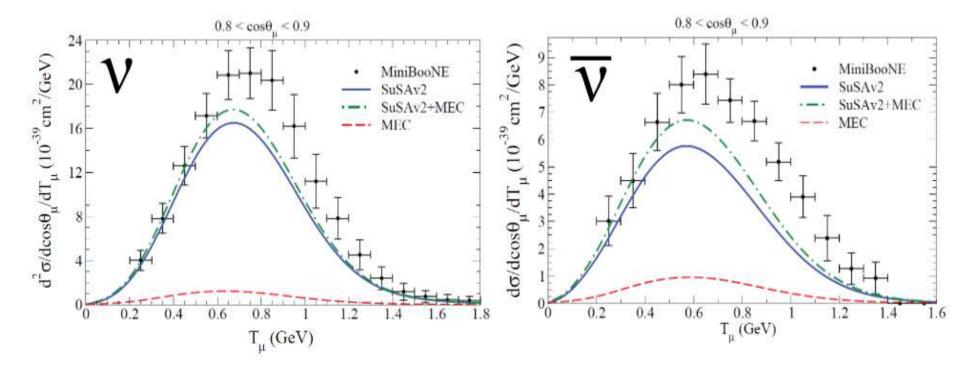


References (non-exhaustive): PRC 83, 045501 (2011); PLB 707, 72 (2012); PRD 85, 113008 (2012); PLB 721, 90 (2013); 88, 113007 (2013), etc.

#### What about correlations in models on the neutrino market?

### Susa model

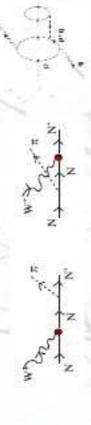
'Effective' scaling approach for genuine QE + MEC correlations

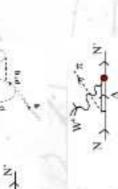


<u>References (non-exhaustive)</u>: Phys.Lett.B696,151 (2011); Phys.Rev.Lett. 108, 152501 (2012); .Phys. Rev.C90, 035501 (2014); Phys. Rev. D91, 073004 (2015) etc.

 "A priori" very similar models (microscopic) give very different results.

Nieves ∆ Martini ∆





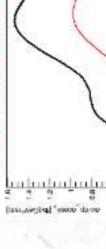




Models have parameters

Prediction : we will see more models

with parameters to adjust to the data



# Models have limits!

Prediction: collaborations th-exp will help to understand model validity.

F.Sanchez, Outlook to 2020. Nuint 2017. Toronto.

# \*Selected\* Open issues (NuSTEC white paper) http://lanl.arxiv.org/pdf/1706.03621

After the acceptance of two-body currents as relevant contribution to the CCQE cross section, several issues still remain. The most urgent one is that of agreement between different models, and between models and experiments. Theoretical results need to be compared in a systematic way to all available data, and validated against electron-scattering data. The various assumptions and differences in models that lead to discrepancies need to be understood. This would be of great help in assessing the range of validity of each approach and facilitate the incorporation of more detailed models in generators

- From a purely theoretical view, the modeling of outgoing hadrons and hadronic final-state interactions is an issue that needs increased efforts
- Interferences between various nuclear effects and a meticulous study of double counting hazards. It is important to identify model-dependences and basis-dependent separations between different approaches.
- The development of consistent models able to cover all experimental needs from 200 MeV to 10s of GeVs is an open issue. None of the theories currently in use cover this vast energy region, models to match and fill the gaps between different predictions need to be developed