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Motivations (3d)

- Mirror symmetry
- ▶ Dimensionful coupling, yet CFTs
- ► Holographic cosmology (Skenderis et al.)
 - ▶ Has exactly the quartic adjoint scalar potential terms that are needed
 - ightharpoonup Can do a better job than $\Lambda ext{CDM}$ at modeling small angle $\ell < 30$ CMB
 - ▶ Lattice simulations can potentially predict large angle statistics and CMB anomalies
- ightharpoonup Holography and ${
 m AdS_4} imes {
 m X_6}$



Parent theory

Dimensional reduction of N=1 6d SYM (all in adjoint).

$$\mathcal{L}=rac{1}{2g^2}\mathrm{Tr}F_{\mu
u}F_{\mu
u}+rac{i}{g}\mathrm{Tr}\Psi^TC\Gamma_{\mu}D_{\mu}\Psi$$
 $A_{\mu}
ightarrow A_i,\quad \phi_{lpha},\quad i=0,1,2;\quad lpha=1,2,3$ $\Psi_p,\quad p=1,\ldots,8
ightarrow \psi_a^I,\quad a=1,2;\quad I=1,2,3,4$

► Hence "N=4."



Continuum (twisted)

Spacetime group

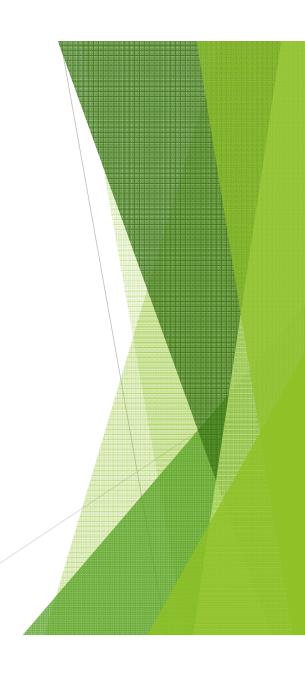
$$SO(4) \simeq SU(2)_l \times SU(2)_r$$

R-symmetry group

$$SU(2)_R \times U(1)_R$$

▶ Twisted rotation group

$$SU(2)' = \operatorname{diag}(SU(2)_r \times SU(2)_R)$$



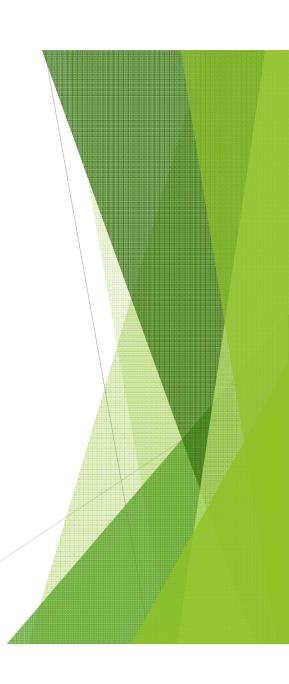
Continuum (twisted, cont.)

Donaldson-Witten twist

$$egin{array}{ll} g^2 \mathcal{L}_{4d}^{\mathcal{N}=2} &=& \mathrm{Tr}igg(rac{1}{4}\mathcal{F}_{\mu
u}\mathcal{F}^{\mu
u} + rac{1}{2}\mathcal{D}_{\mu}ar{\phi}\mathcal{D}^{\mu}\phi - lpha\left[\phi,ar{\phi}
ight]^2 \ && -rac{i}{2}\eta\mathcal{D}_{\mu}\psi^{\mu} + ilpha\phi\left\{\eta,\eta
ight\} - rac{i}{2}ar{\phi}\left\{\psi_{\mu},\psi^{\mu}
ight\} + \mathcal{L}_{\chi}igg), \end{array}$$

$$\mathcal{L}_{\chi} = ext{Tr}\left(rac{i}{8}\phi\left\{\chi_{\mu
u},\chi^{\mu
u}
ight\} - i\chi^{\mu
u}\mathcal{D}_{\mu}\psi_{
u}
ight)$$

ightharpoonup Dimensionally reduce to 3d. $\mathcal{D}_2 o [\phi_3,\cdot]$



Previous work

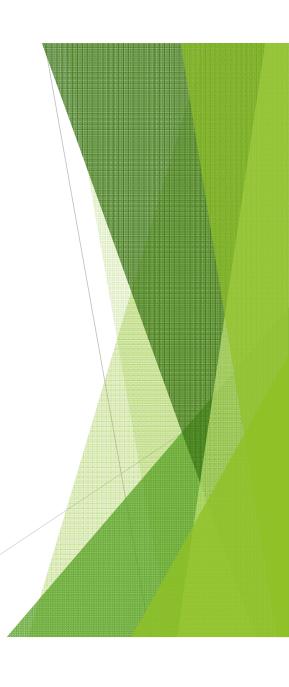
- Anosh Joseph [1307.3281]
- ▶ Blau-Thompson twist [hep-th/9612143]

$$SU(2)' = \operatorname{diag}(SU(2)_E \times SU(2)_N)$$

From the dimensional reduction

$$SO(6) \rightarrow SO(3) \times SO(3) \simeq SU(2)_E \times SU(2)_N$$

► Earlier Q=8 formulation by orbifold method [Cohen, Kaplan, Katz & Unsal 2003]



Our twisted lattice

- Must complexify everything.
- Dynamical lattice spacing, as usual in these twisted/orbifold lattices.
- ▶ Lift additional fields with generic mass terms.

$$\begin{split} \mathcal{L} &= \mathrm{Tr} \left(\frac{1}{4} \bar{\mathcal{F}}_{\mu\nu}(n) \mathcal{F}_{\mu\nu}(n) + \frac{1}{2} \bar{\mathcal{D}}_{\mu}^{+} \bar{\phi}(n) \mathcal{D}_{\mu}^{+} \phi(n) - \alpha \left[\phi(n), \bar{\phi}(n) \right]^{2} \right. \\ &+ \frac{i}{2} \bar{\mathcal{D}}_{\mu}^{+} \eta(n) \psi_{\mu}(n) + i \alpha \phi(n) \left\{ \eta(n), \eta(n) \right\} - \frac{i}{2} \bar{\phi}(n) \left(\psi_{\mu}(n) \bar{\psi}_{\mu}(n) + \bar{\psi}_{\mu} \left(n - e_{\mu} \right) \psi_{\mu} \left(n - e_{\mu} \right) \right) \right) + \mathcal{L}_{\chi}, \\ \mathcal{L}_{\chi} &= \mathrm{tr} \left[\frac{i}{8} \left(\phi(n) \bar{\chi}_{\mu\nu}(n) \chi_{\mu\nu}(n) + \phi \left(n + e_{\mu} + e_{\nu} \right) \chi_{\mu\nu}(n) \bar{\chi}_{\mu\nu}(n) \right) \right. \\ &- \frac{i}{2} \left(\bar{\chi}_{\mu\nu}(n) \bar{\mathcal{D}}_{\mu}^{+} \bar{\psi}_{\nu}(n) + \chi_{\mu\nu}(n) \mathcal{D}_{\mu}^{+} \psi_{\nu}(n) \right) \right]. \end{split}$$

► Lattice gauge invariance and Q invariance of

$$\chi_{\mu
u}(n) = rac{1}{2} \epsilon_{\mu
u
ho\lambda} ar{\chi}_{
ho\lambda} \left(n + e_{\mu} + e_{
u}
ight).$$

Implies

$$\sum_{\mu=1}^4 e_\mu = 0.$$

So, the theory must be 3d.



After using EOM,

$$\mathcal{L} = Q \operatorname{Tr} \left(rac{1}{4} \chi_{\mu
u}(n) \mathcal{F}_{\mu
u}(n) + rac{1}{2} ar{\mathcal{D}}_{\mu}^{+} ar{\phi}(n) \psi_{\mu}(n) + lpha \eta(n) \left[\phi(n), ar{\phi}(n)
ight]
ight)
onumber \ - rac{1}{8} \epsilon_{\mu
u
ho\lambda} \operatorname{tr} \left(\mathcal{F}_{\mu
u}(n) \mathcal{F}_{
ho\lambda} \left(n + e_{\mu} + e_{
u}
ight)
ight),$$

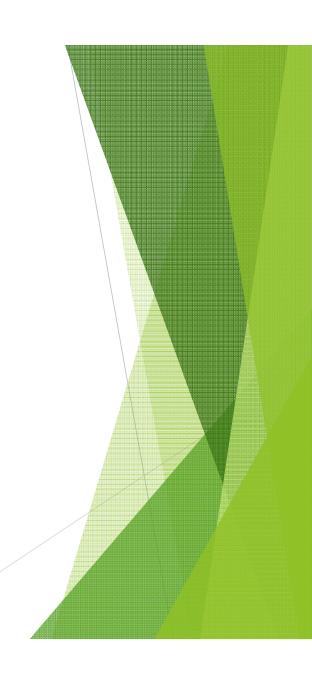
▶ Last term is Q invariant using lattice Bianchi identity.

▶ Where Q, which is nilpotent, acts as:

$$egin{aligned} Q\,\phi(n) &= 0,\;\; Q\,ar\phi(n) = i\eta(n),\ Q\,\eta(n) &= \left[ar\phi(n),\phi(n)
ight],\ Q\,\mathcal{U}_\mu(n) &= i\psi_\mu(n),\;\; Q\,ar{\mathcal{U}}_\mu(n) = -iar\psi_\mu(n)\ Q\,\psi_\mu(n) &= \mathcal{D}^+_\mu\phi(n),\;\; Q\,ar\psi_\mu(n) = ar{\mathcal{D}}^+_\mu\phi(n)\ Q\,\chi_{\mu
u}(n) &= ar{\mathcal{F}}_{\mu
u}(n) + rac{1}{2}\epsilon_{\mu
u
ho\lambda}\mathcal{F}_{
ho\lambda}\left(n+e_\mu+e_
u
ight). \end{aligned}$$

Counterterms

- Fine-tuning is calculable because the theory is super-renormalizable.
- As will see in arguments below, all counterterms are one-loop.



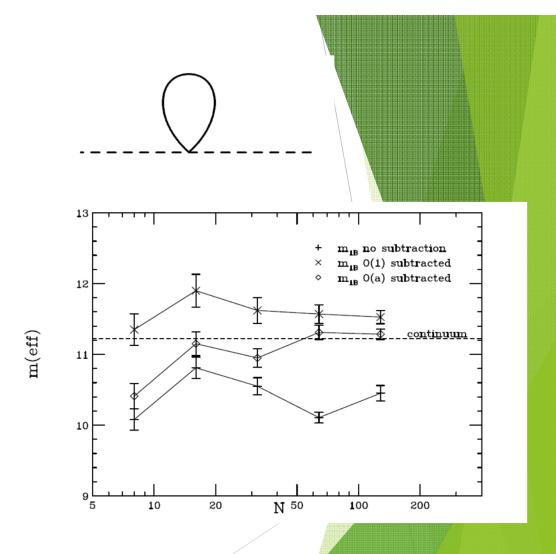
Power of super-renormalizability

- ▶ So, finite number of counterterms that can all be calculated at one loop in lattice perturbation theory.
- ▶ Includes the effects of the mass terms and scalar Q breaking.



Example from SUSY QM

- Finite CT in SUSY QM with naïve discretization [JG et al., 2004]
- Doublers appearing in this diagram (cancelled by scalar loop in the Q-exact case)
- Also shows power of Symanzik improvement



Prospects

- Currently coding up the Anosh Joseph discretization b/c it is much cleaner.
- Seems to have only one CT due to exact Q, point group, lattice gauge invariance.

$$\mathcal{L}(n) = rac{1}{g^2} Q ext{Tr} \left(\chi_{ab}(n) \mathcal{D}_a^+ \mathcal{U}_b(n) + \eta(n) ar{\mathcal{D}}_a^- \mathcal{U}_a(n) + rac{1}{2} \eta(n) d(n) + B_{abc}(n) ar{\mathcal{D}}_a^+ \chi_{bc}(n)
ight)$$

► For dimension counting it is best to have canonical normalization (otherwise everything is very confusing)

$$\Phi o g\Phi, \quad \Psi o g\Psi, \quad \mathcal{U}_m = rac{1}{ag}e^{ag\mathcal{A}_m}$$

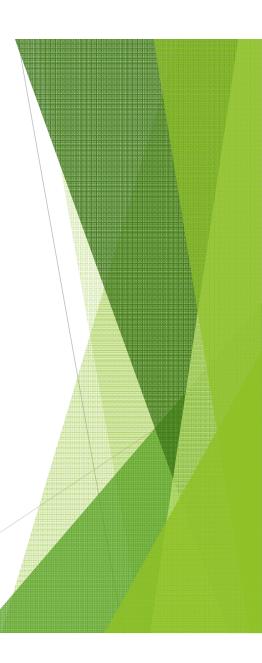
Renormalization in the Blau-Thompson twist

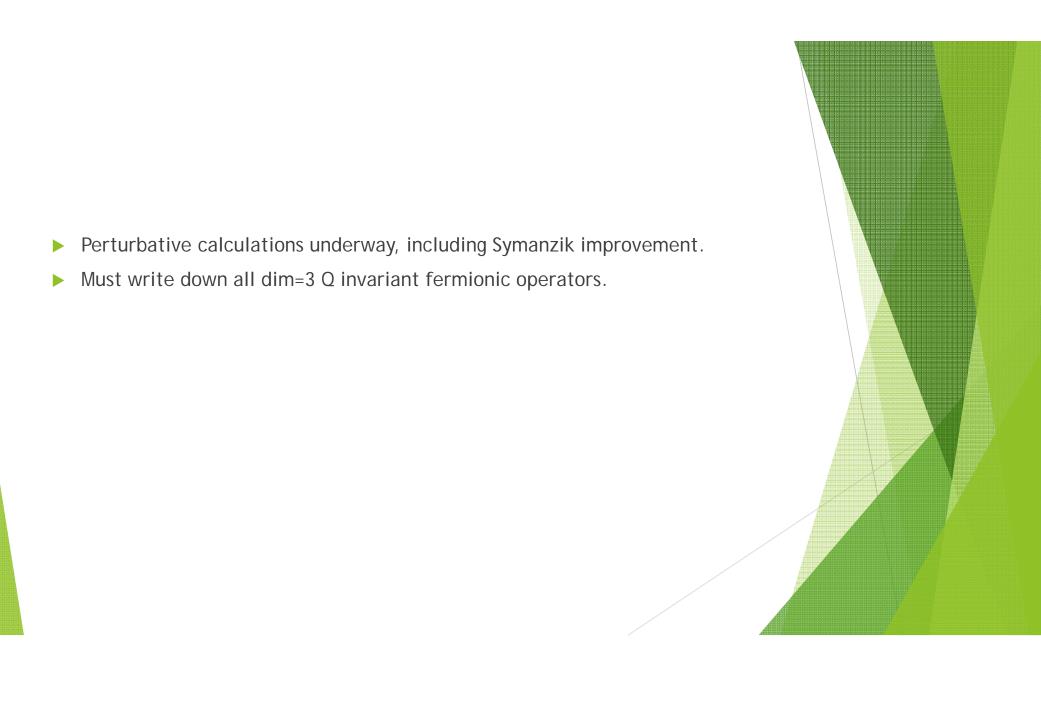
- \blacktriangleright [boson] = $\frac{1}{2}$, [fermion] = 1, [d]= $\frac{3}{2}$, [F]= $\frac{3}{2}$, [Q] = $\frac{1}{2}$
- So since loop is multiplied by $[g^2] = 1$
- ▶ Must have no more than dim=2 to be unsuppressed by lattice spacing.
- ▶ But also be fermionic if under Q.
- So restricting to Q-exact operators, must be of form

$$Q \text{Tr} \Psi$$
, $Q \text{Tr} \Psi \Phi$, $Q (\text{Tr} \Psi \text{Tr} \Phi)$

▶ Point group, lattice gauge invariance, limit to [shifts d EOM by constant]

$$Q{
m Tr}\eta={
m Tr}d$$





Conclusions

- ► The holographic cosmology doesn't really require SUSY, but it will be interesting to see how it impacts large angle predictions.
- ► Concrete realization of Symanzik improvement vs. SUSY in 3d should be very enlightening.
- In the Blau-Thompson twist we only have one 1-loop CT to determine in order to get full N=4 SUSY. Symanzik improvement will require more.