

# Surrogate neutron capture reaction prospects for r-process nuclei

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FRIB and the GW170817 Kilonova

Facility for Rare Isotope Beams, 23-27 July 2018

# Surrogate neutron capture reaction prospects for r-process nuclei

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Potel<sup>(4)</sup>, Steve Pain<sup>(3)</sup>, David Walter<sup>(1)</sup>

*(1) Rutgers University*

*(2) Lawrence Livermore National Laboratory*

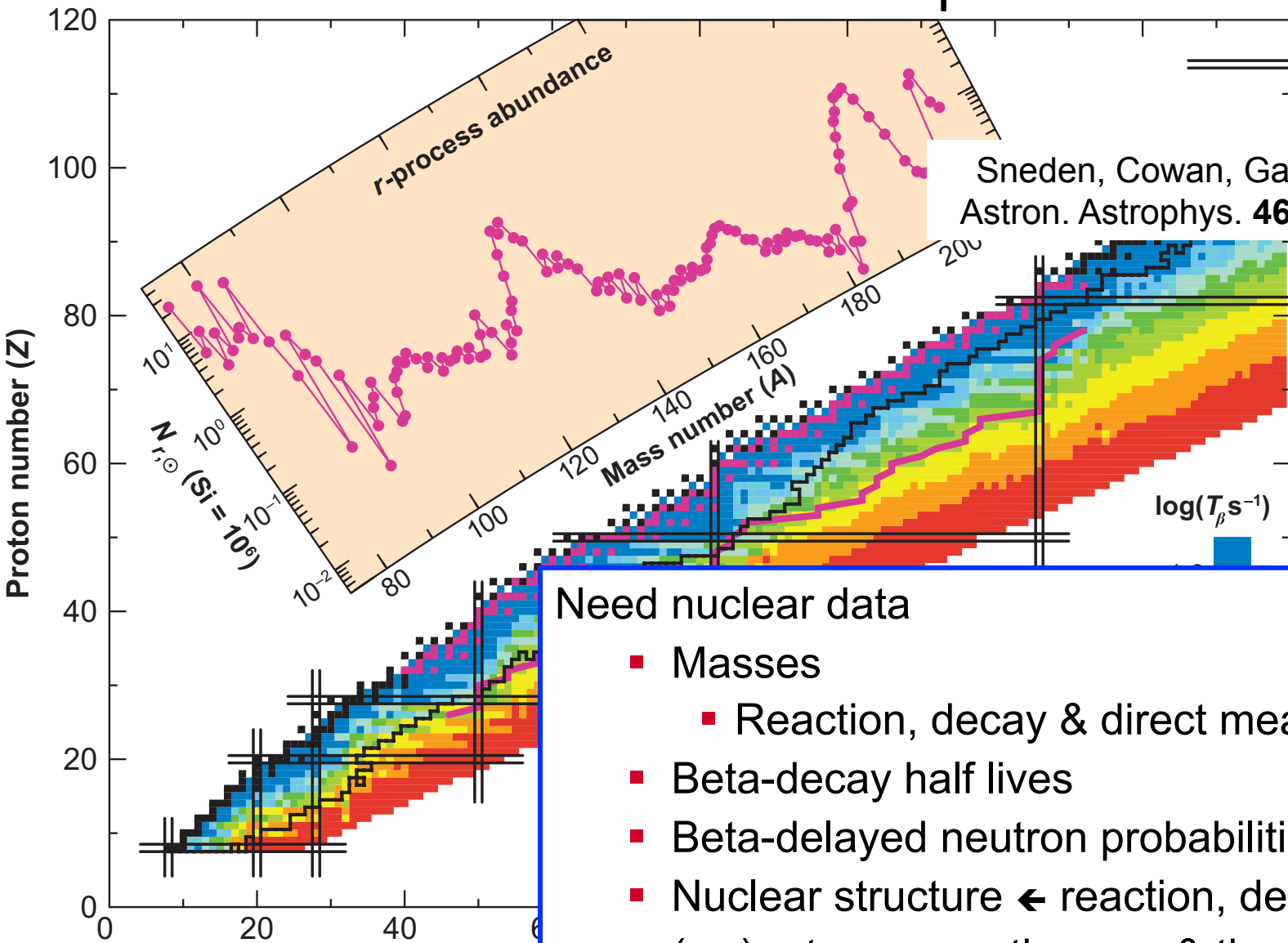
*(3) Oak Ridge National Laboratory*

*(4) Michigan State University & FRIB*

and the **ORRUBA, STAR-LiTeR and GODDESS**  
**collaborations**

**Funded in part by the** U.S. Department of Energy National  
Nuclear Security Administration & Office of Nuclear Physics and  
the National Science Foundation

depends on nuclear data

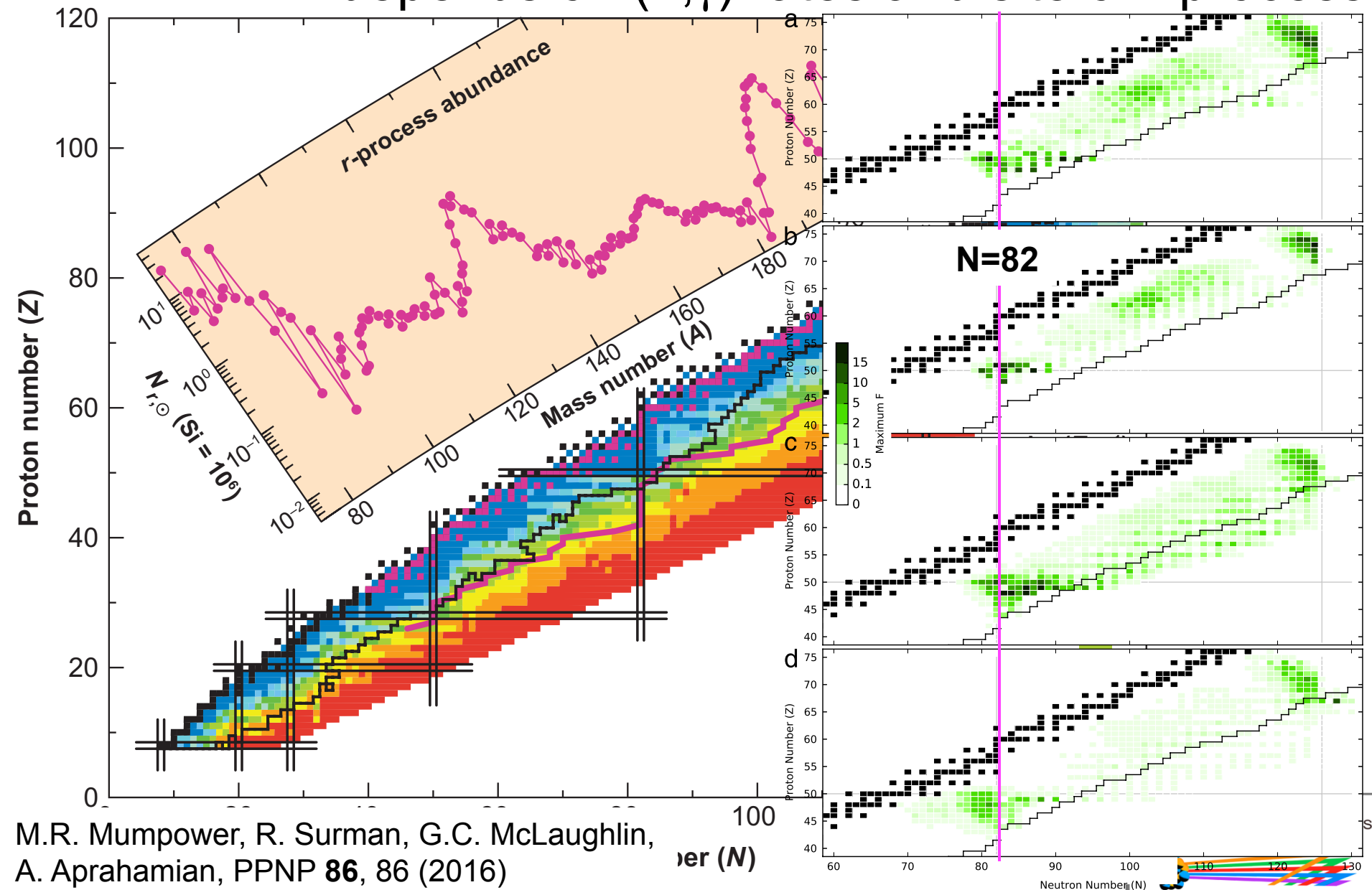


Snedden, Cowan, Gallino, *Annu. Rev. Astron. Astrophys.* **46**:241-288 (2008)

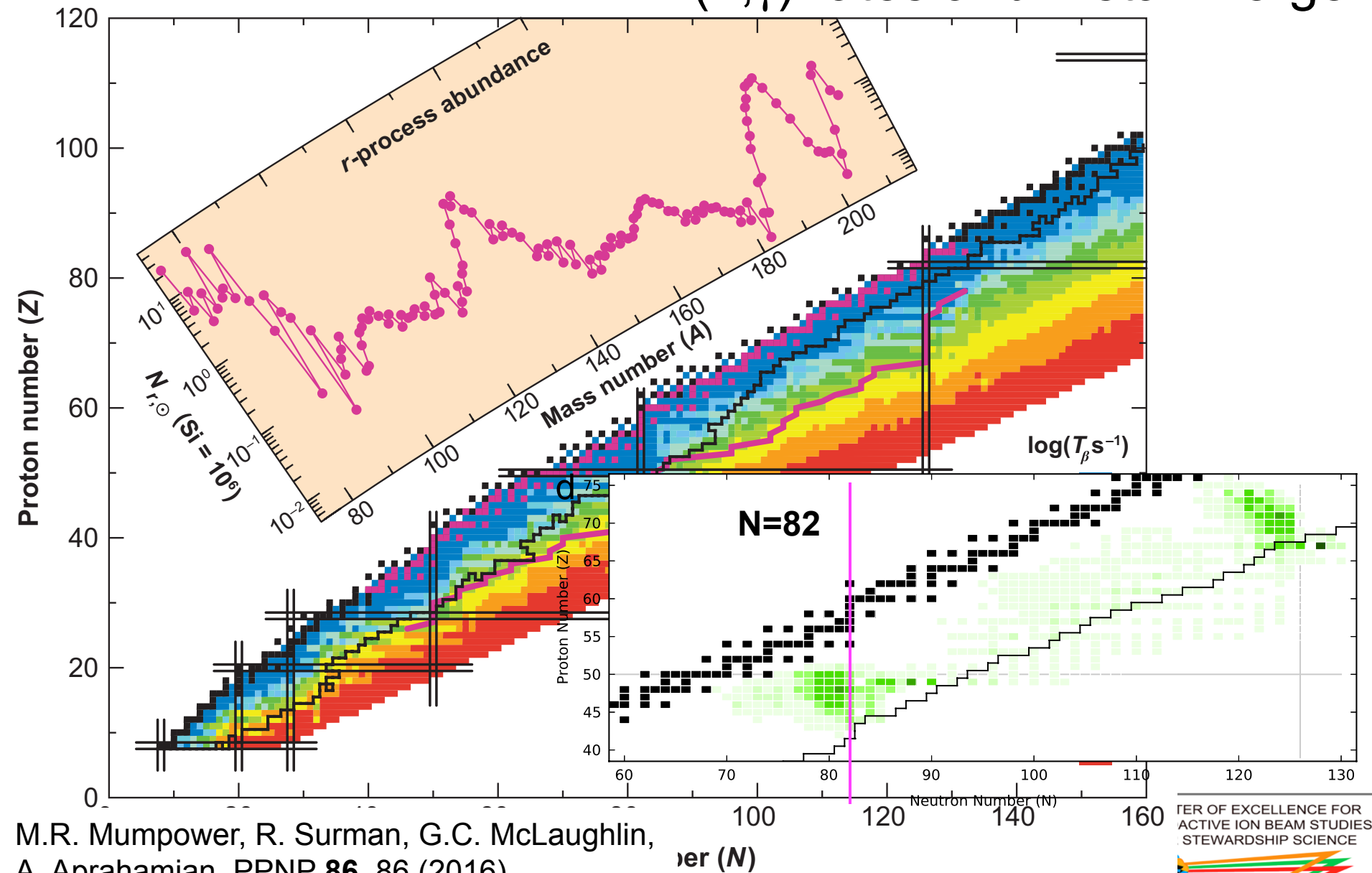
### Need nuclear data

- Masses
  - Reaction, decay & direct measurements
- Beta-decay half lives
- Beta-delayed neutron probabilities
- Nuclear structure ← reaction, decay & theory
- $(n,\gamma)$  rates ← reaction exp & theory studies

depends on  $(n, \gamma)$  rates and site of r process

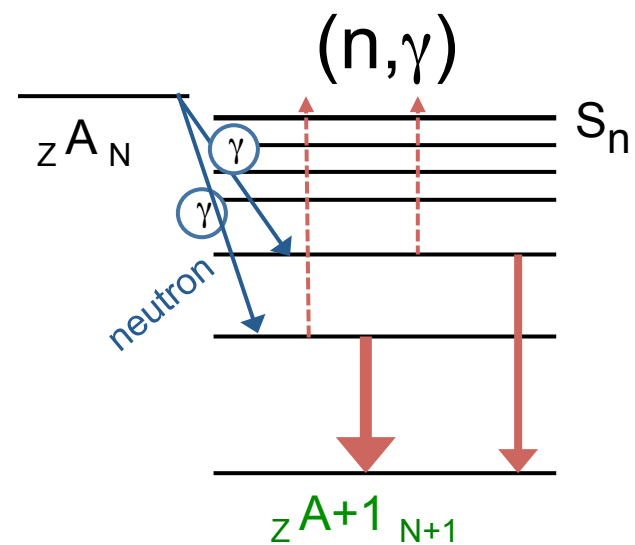
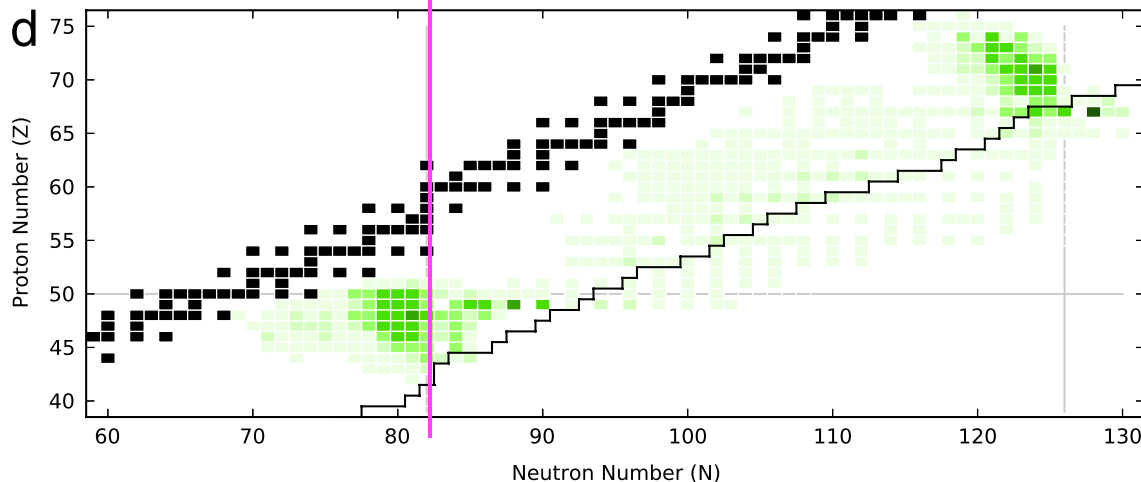


M.R. Mumpower, R. Surman, G.C. McLaughlin, A. Aprahamian, PPNP **86**, 86 (2016)



M.R. Mumpower, R. Surman, G.C. McLaughlin, A. Aprahamian, PPNP **86**, 86 (2016)

neutron capture dominated by direct capture



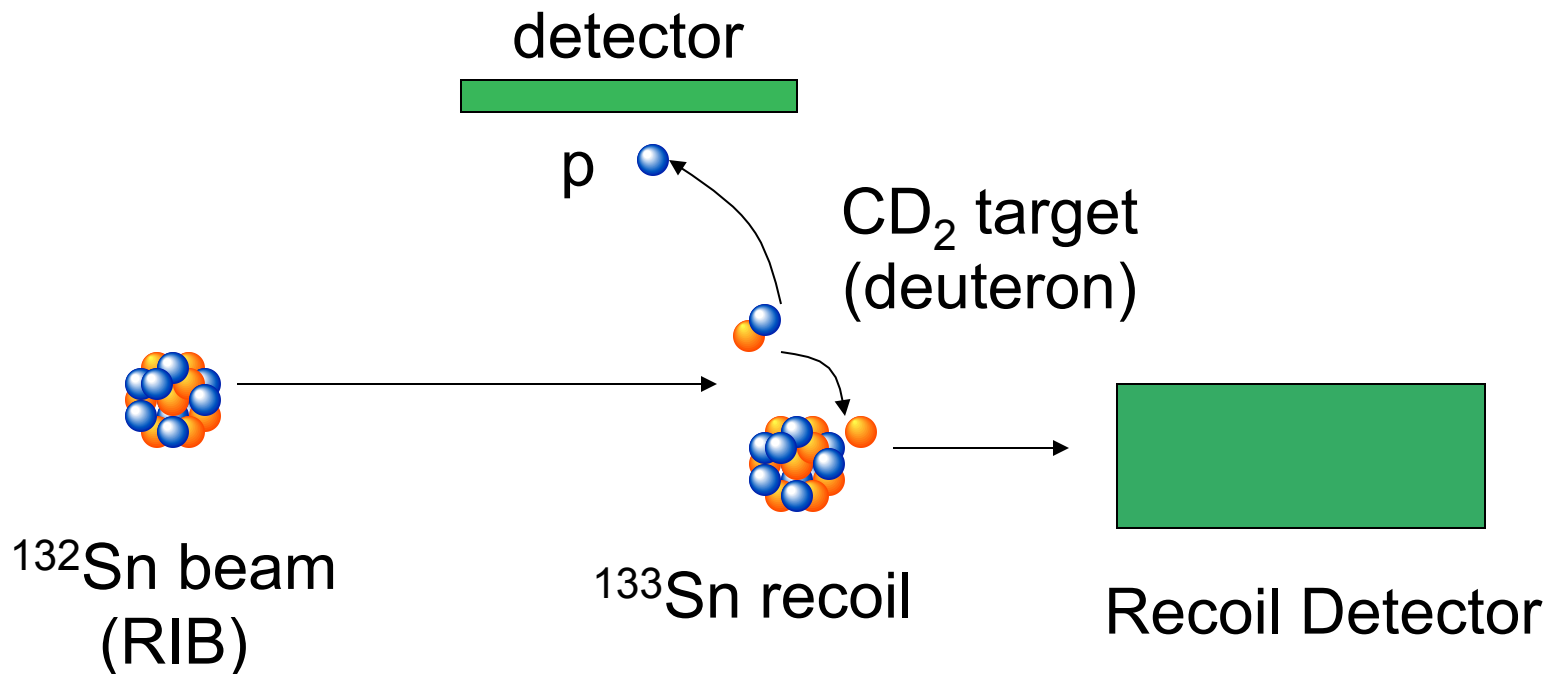
### Tin Z=50 isotopes n-star merger

Mumpower, et al. PPNP 2016

77	127	1.77
78	128	1.21
79	129	3.55
80	130	4.47
81	131	3.28
82	132	1.92

Inform by measuring neutron transfer  
e.g., (d,p) with n-rich RIBs



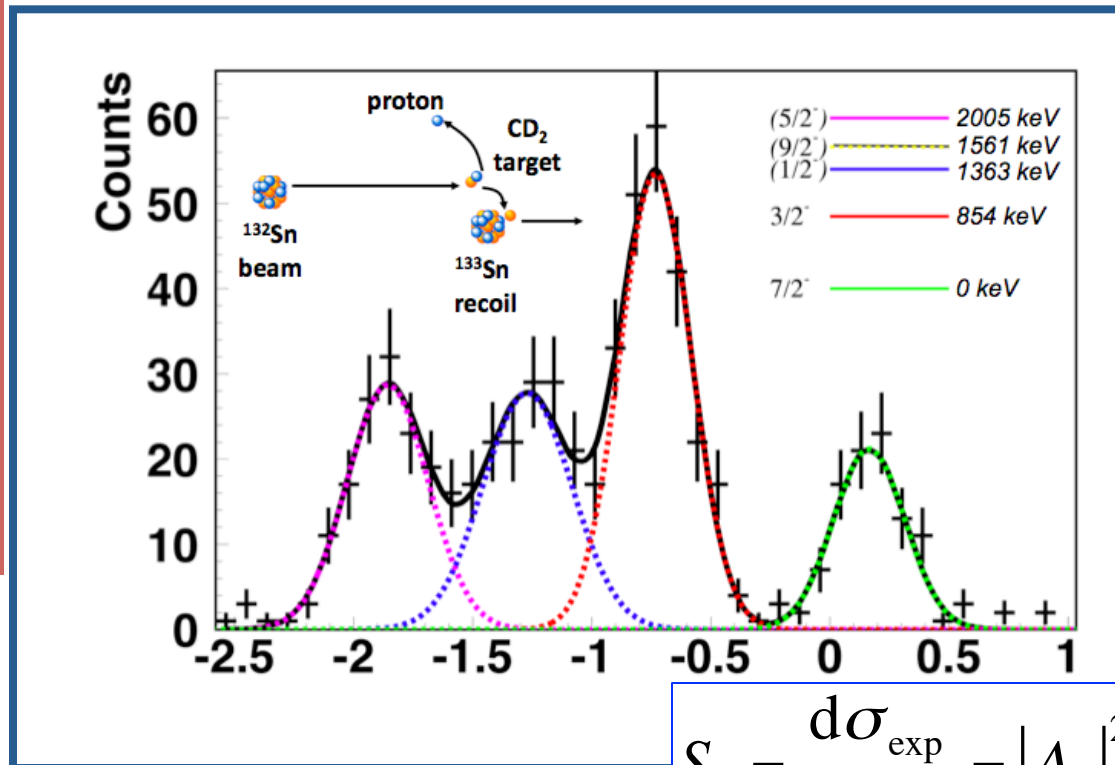


- Unfavorable kinematics → Reduced Q-value Resolution
- Rare Ion Beams (RIBs) are difficult and expensive to produce

Applicable to all isotopes which can be made into a beam

Identified  $2f_{7/2}$ ,  
 $3p_{3/2}$ ,  $(3p_{1/2})$ ,  $2f_{5/2}$   
 neutron strength in  
 $^{133}\text{Sn}$

K.L. Jones et al.  
 Nature, **465**,454 (2010)  
 Phys. Rev. C **84**, 034601 (2011)



$$S_{lj} = \frac{d\sigma_{\text{exp}}}{d\sigma_{\text{DW}}} = |A_{lj}|^2$$

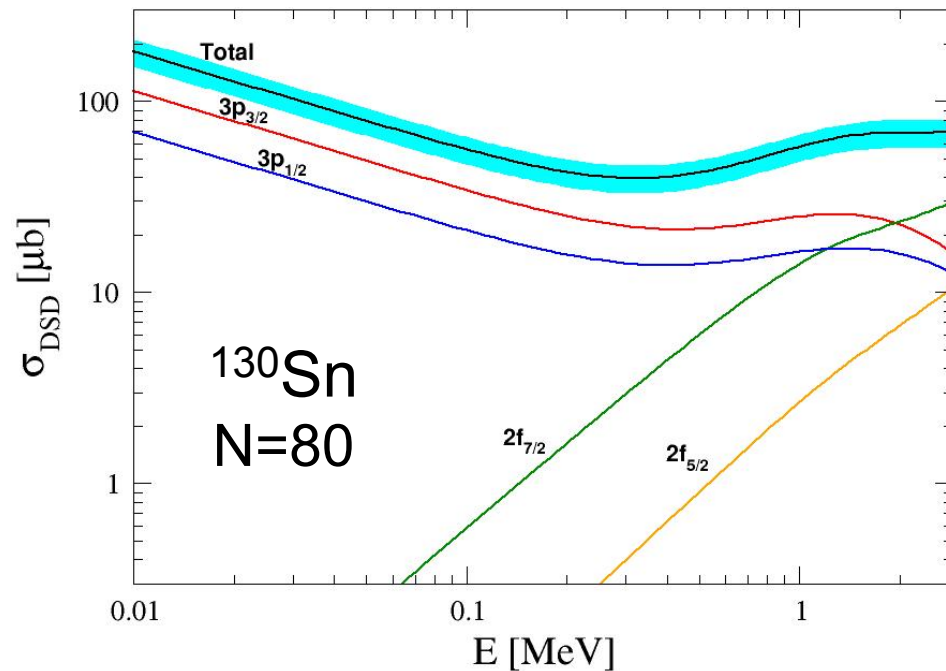
$E_x(\text{keV})$	$J^\pi$	Config	SF (DWBA)	SF (FR-ADWA)
0	$7/2^-$	$2f_{7/2}$	0.86(14)	1.00(8)
854	$3/2^-$	$3p_{3/2}$	0.92(14)	0.92(7)
1363(31)	$(1/2^-)$	$3p_{1/2}$	1.1(3)	1.2(2)
2005	$(5/2^-)$	$2f_{5/2}$	1.1(2)	1.2(3)

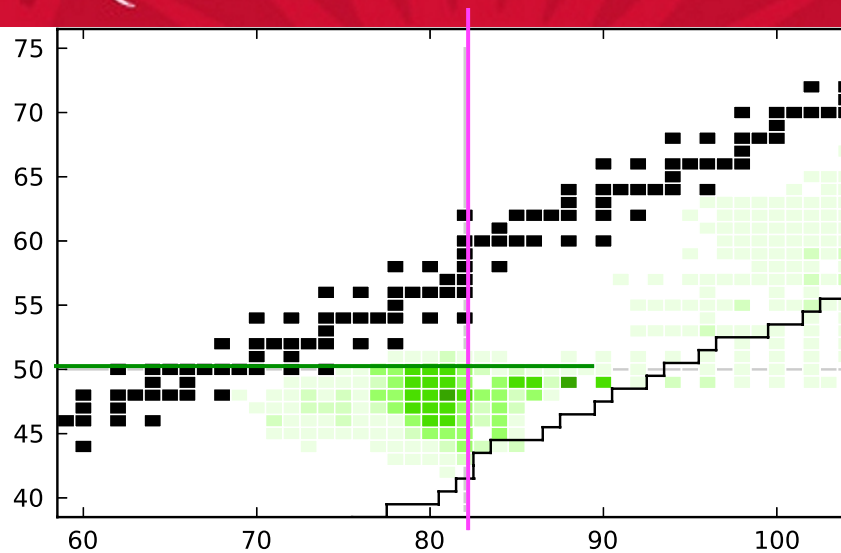


## Direct semi-direct direct capture with CUPIDO

- Incident n channel: Koning Delaroche potential
- Bound state: Bear Hodgson potential
- Semi-direct capture via GDR
  - Add GDR to s.p. EM operator
- Used measured SF and  $E_x$  to constrain
- Uncertainties  $\approx 20\%$

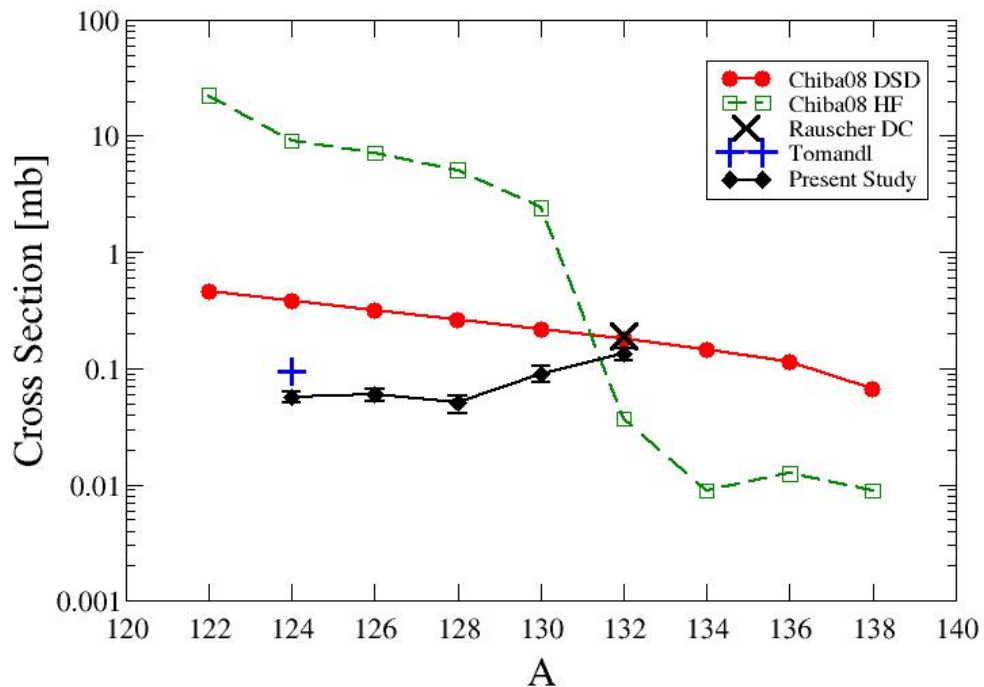
	30 keV $\sigma(n,\gamma)$ ( $\mu\text{b}$ )
$^{132}\text{Sn}(d,p)$	134(17)
$^{130}\text{Sn}(d,p)$	90(15)
$^{128}\text{Sn}(d,p)$	51(8)
$^{126}\text{Sn}(d,p)$	59(7)
$^{124}\text{Sn}(d,p)$	56(6)





Neutron star merger

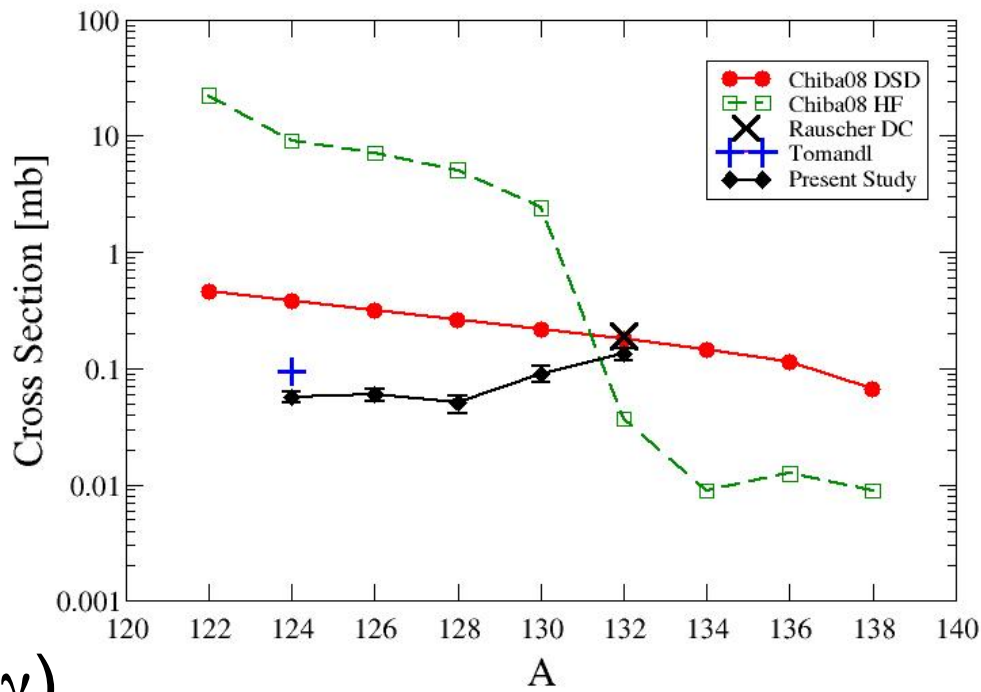
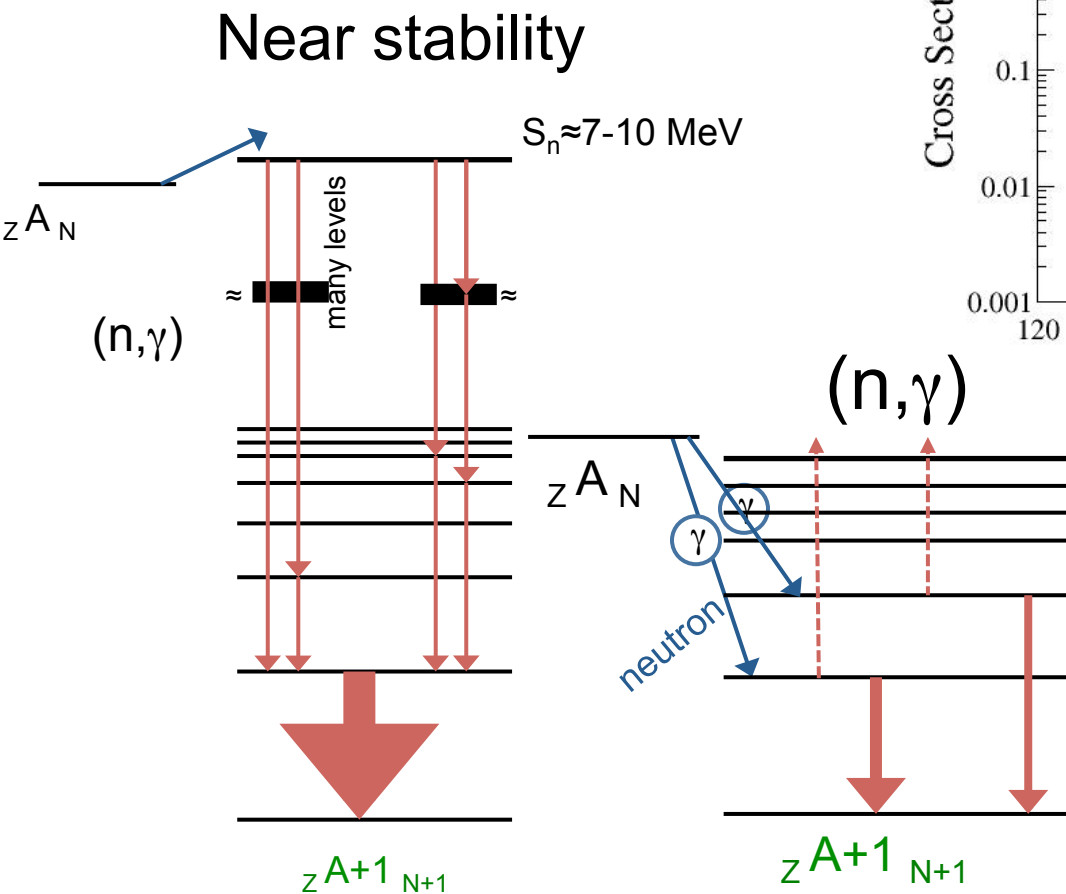
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**Sn(n,γ) vs A**  
 Theory: Chiba, et al. PRC 77, 015809 (2008)  
 DSD from exp: G. Arbanas, B. Manning



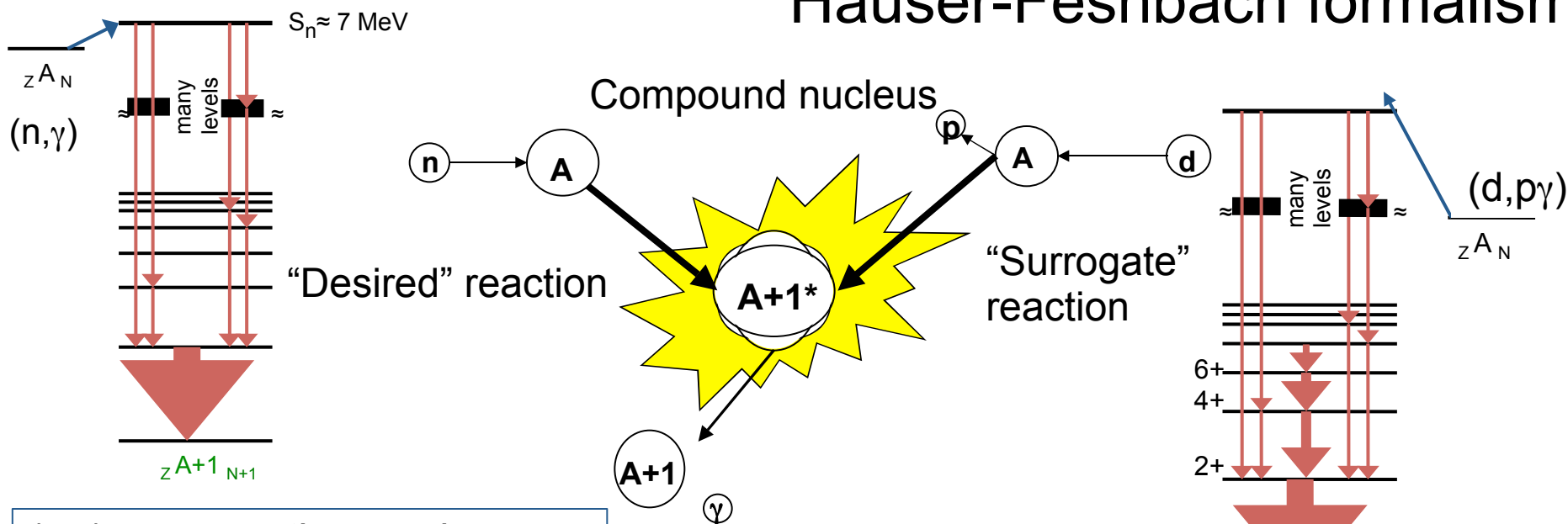
Statistical ( $n,\gamma$ ) dominates  $\sigma$  when  $N < N_{\text{magic}}$



Near waiting points

**Sn( $n,\gamma$ ) vs A**  
 Theory: Chiba, et al. PRC 77, 015809 (2008)  
 DSD from exp: G. Arbanas, B. Manning

# Surrogate reaction concept & Hauser-Feshbach formalism



$(n, \gamma)$  cross section can be written as product of compound nucleus formation and decay for every spin and parity:

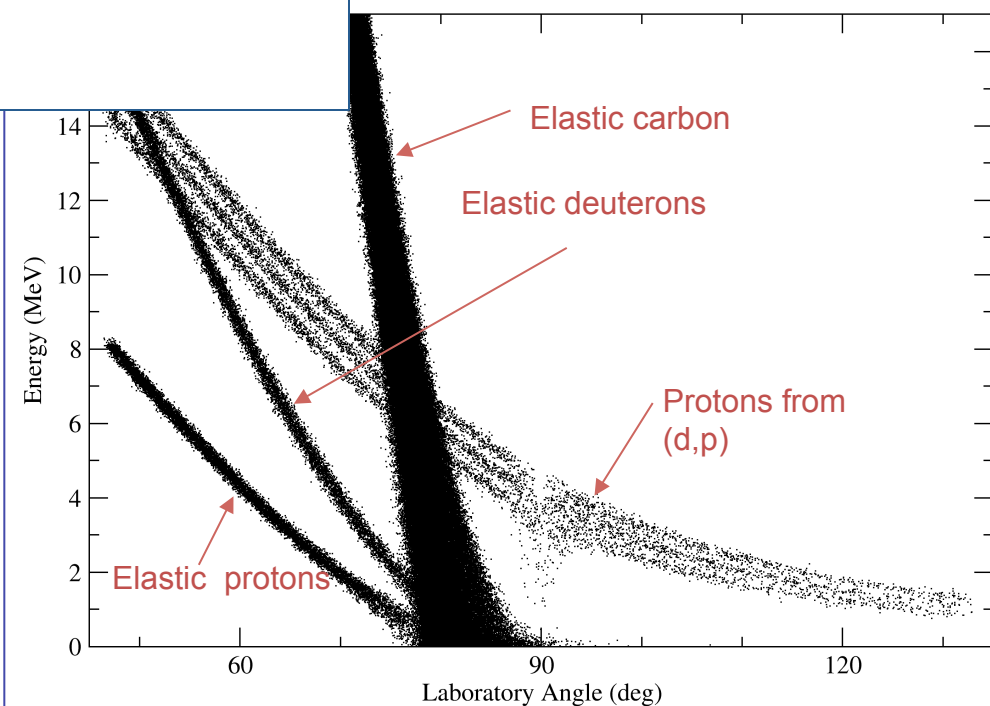
$$\sigma_{n\gamma}(E_n) = \sum_{J, \pi} \sigma_n^{CN}(E_x, J, \pi) G_\gamma^{CN}(E_x, J, \pi)$$

Surrogate particle-gamma coincidence can be written as product of compound nucleus formation and decay for every spin and parity:

$$P_{p\gamma}(E_x, \theta) = \sum_{J, \pi} F_{dp}^{CN}(E_x, J, \pi, \theta) G_\gamma^{CN}(E_x, J, \pi)$$

# Good candidate for (n, $\gamma$ ) surrogate with beams

- Relatively good match with spin distribution in (n, $\gamma$ ) which is dominated by  $\ell=0$
- Reaction predominantly one-step transfer of  $j=\ell\pm 1/2$  neutron
- “Easy” to produce CD<sub>2</sub> targets
- “Lower” beam energies (than heavier targets) to get above neutron separation energy
- Kinematics favors cleaner reaction



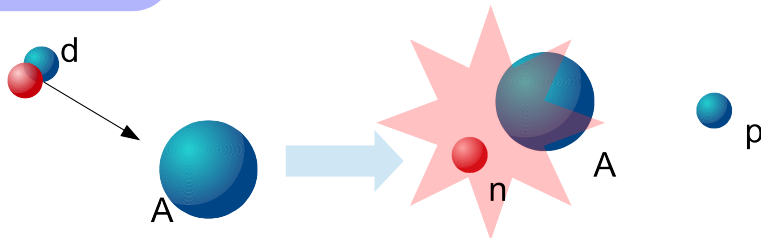
$$P_{p\gamma}(E_x, \theta) = \sum_{J, \pi} F_{dp}^{CN}(E_x, J, \pi, \theta) G_{\gamma}^{CN}(E_x, J, \pi)$$



# Neutron transfer (d,p) to unbound states, non-elastic breakup and surrogate for (n, $\gamma$ )

**step 1**

separation of the proton



## Two-step process

- d breakup; B.E. = 2.2 MeV

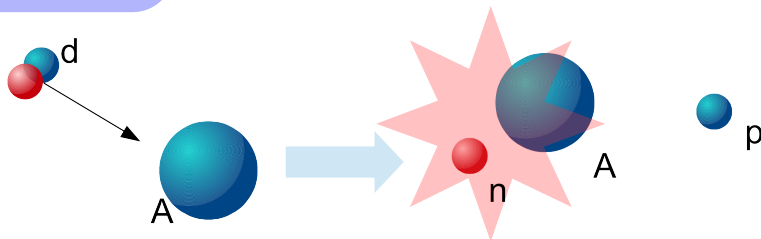
Gregory Potel et al. PRC 92, 034611(2015)  $\Rightarrow$  path to CN formation



# Neutron transfer (d,p) to unbound states, non-elastic breakup and surrogate for (n, $\gamma$ )

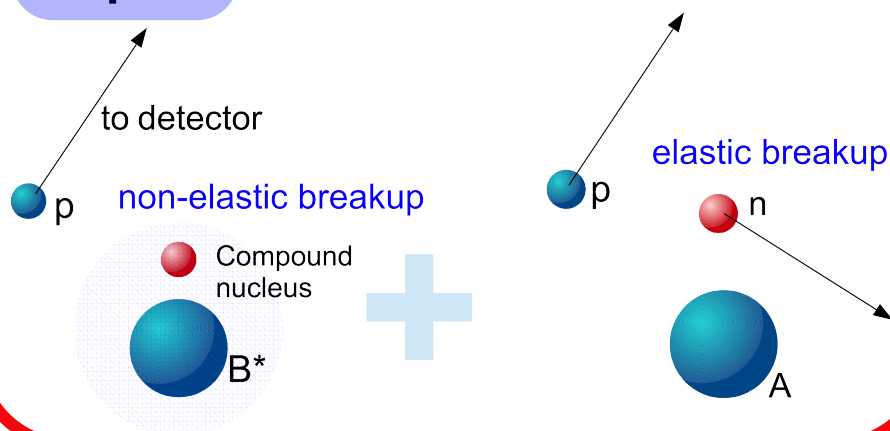
## step 1

separation of the proton



## step 2

propagation of  $n$  in the field of  $B^*$



## Two-step process

- d breakup; B.E. = 2.2 MeV
- n propagation
  - Elastic breakup
  - Non-elastic breakup  $\Rightarrow$  CN and surrogate (n, $\gamma$ )
  - Predicts  $J^\pi$  transfer

Gregory Potel et al. PRC 92, 034611(2015)  $\Rightarrow$  path to CN formation





$(d,p)$  reaction forms compound nucleus

Need to measure  $P(d,p\gamma)$

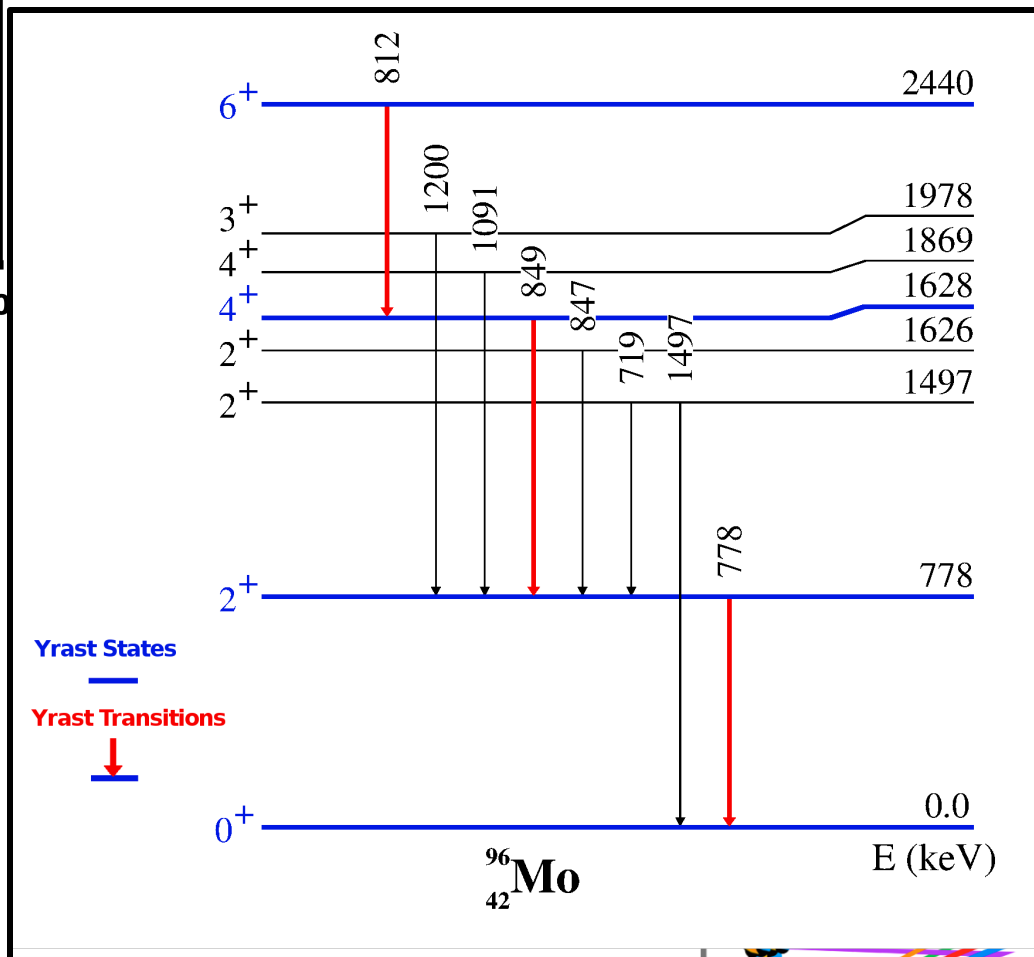
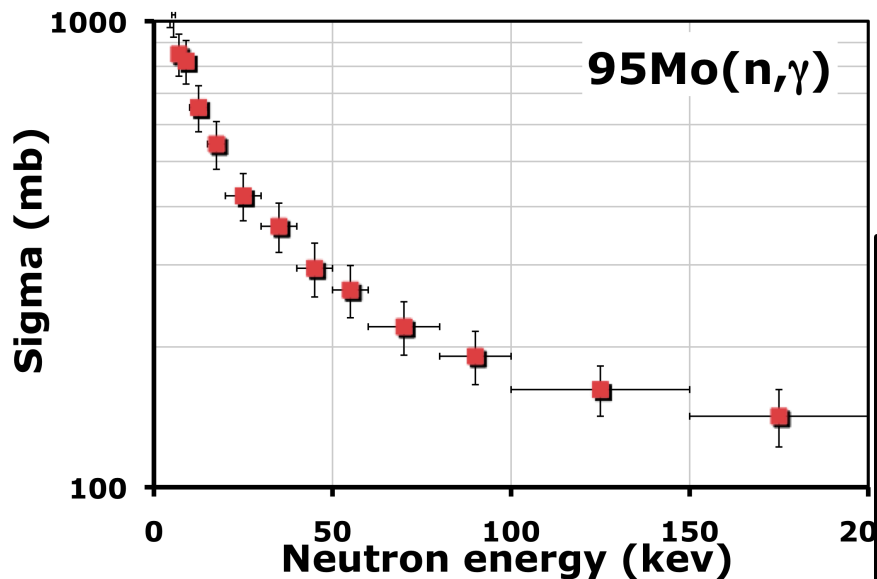
Need to deduce  $G^{CN}$  by fit to  $P(d,p\gamma)$  accounting for  $F^{CN}$

$$P_{p\gamma}(E_x, \theta) = \sum_{J, \pi} F_{dp}^{CN}(E_x, J, \pi, \theta) G_{\gamma}^{CN}(E_x, J, \pi)$$

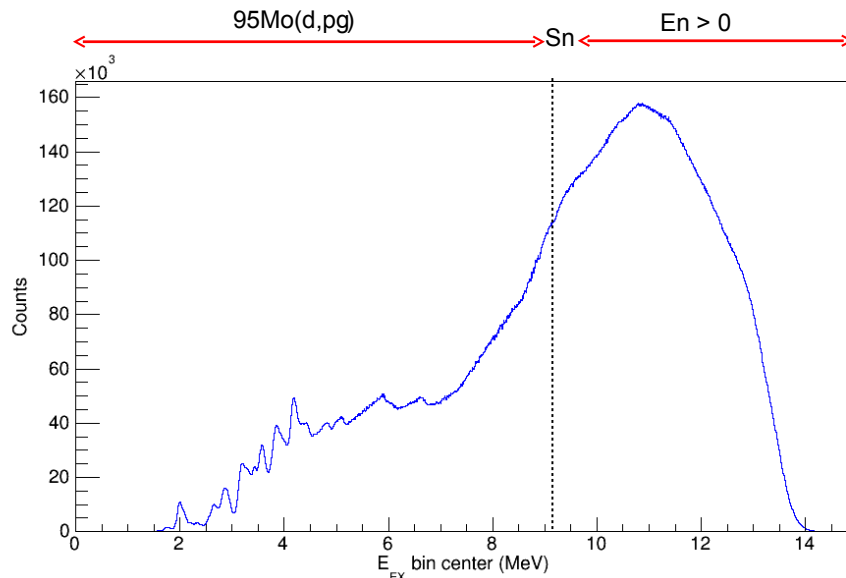
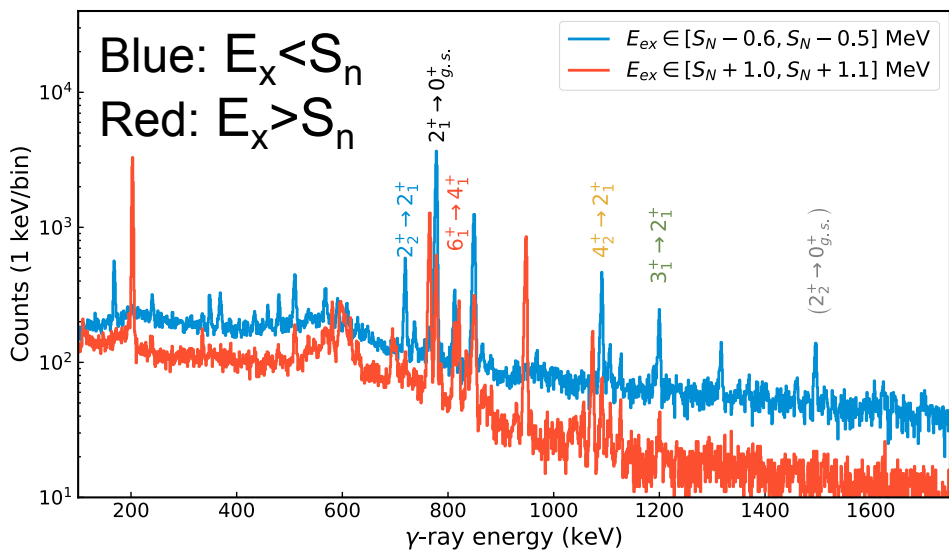
Validate with  $^{95}\text{Mo}(d,p\gamma)^{96}\text{Mo}$  reaction

$\sigma(n,\gamma)$  was measured and evaluated

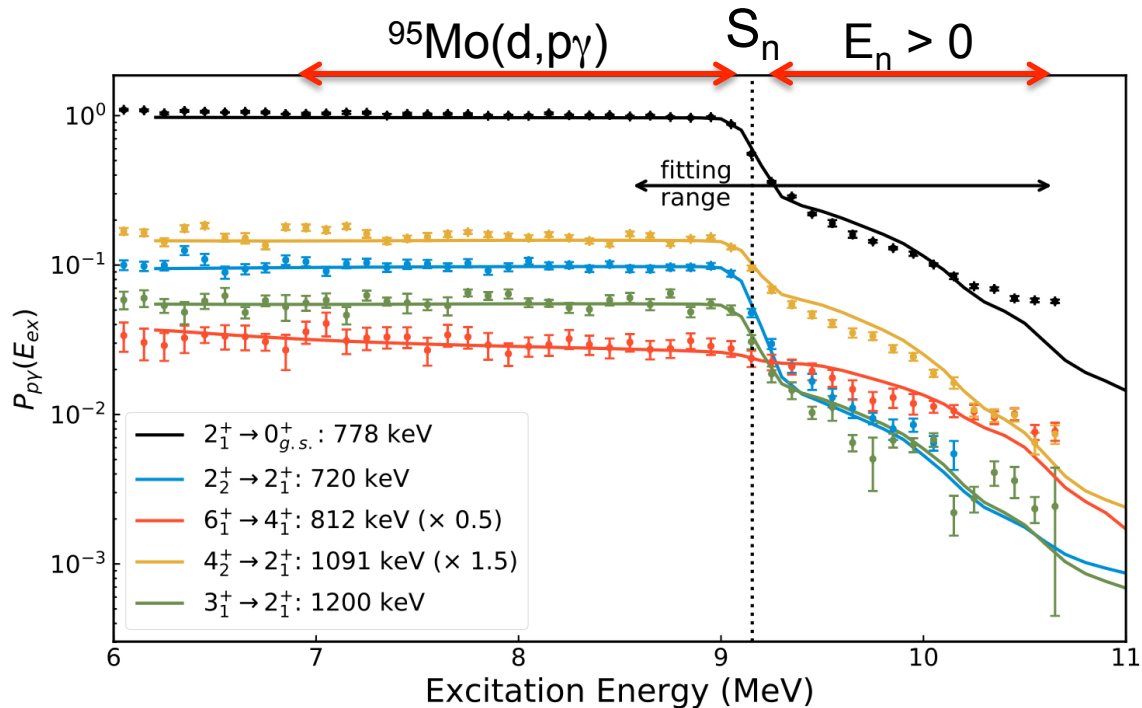
de L. Musgrove, et al., NPA **270**, 109 (1976)

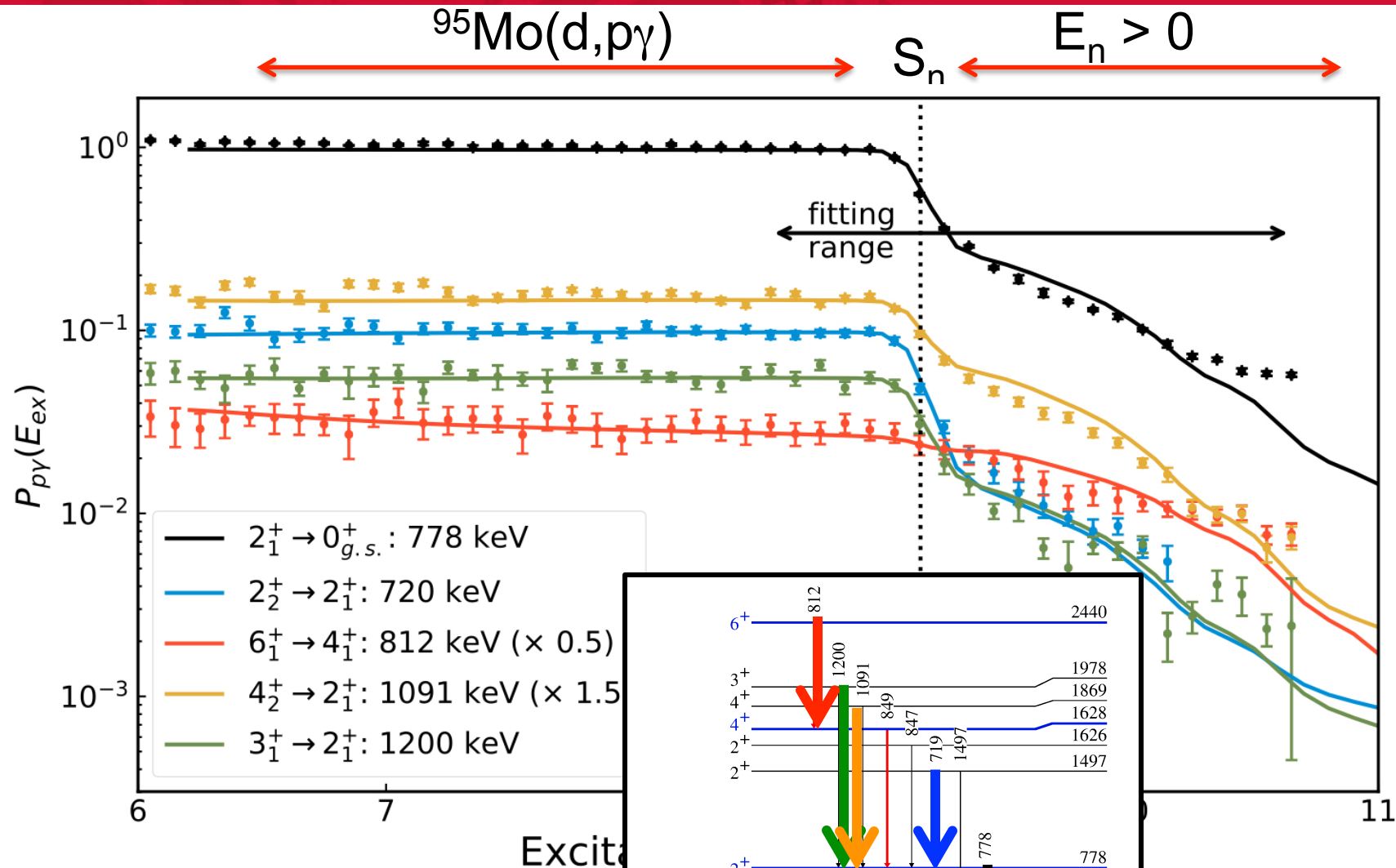


$^{96}\text{Mo}$  level scheme  
778 keV collecting transition

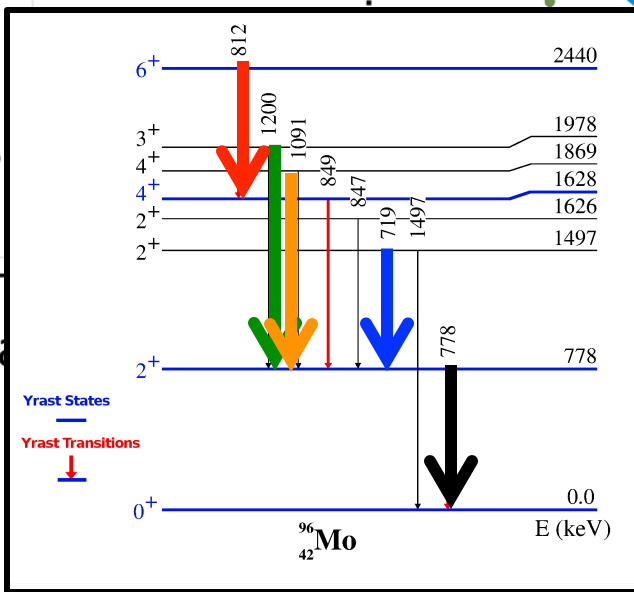


$$P(d, p\gamma) = \frac{N_{p-\gamma}}{N_p \epsilon}$$





- $2_1^+ \rightarrow 0_{g.s.}^+$ : 778 keV
- $2_2^+ \rightarrow 2_1^+$ : 720 keV
- $6_1^+ \rightarrow 4_1^+$ : 812 keV ( $\times 0.5$ )
- $4_2^+ \rightarrow 2_1^+$ : 1091 keV ( $\times 1.5$ )
- $3_1^+ \rightarrow 2_1^+$ : 1200 keV



Measured  $P(d,p\gamma)$

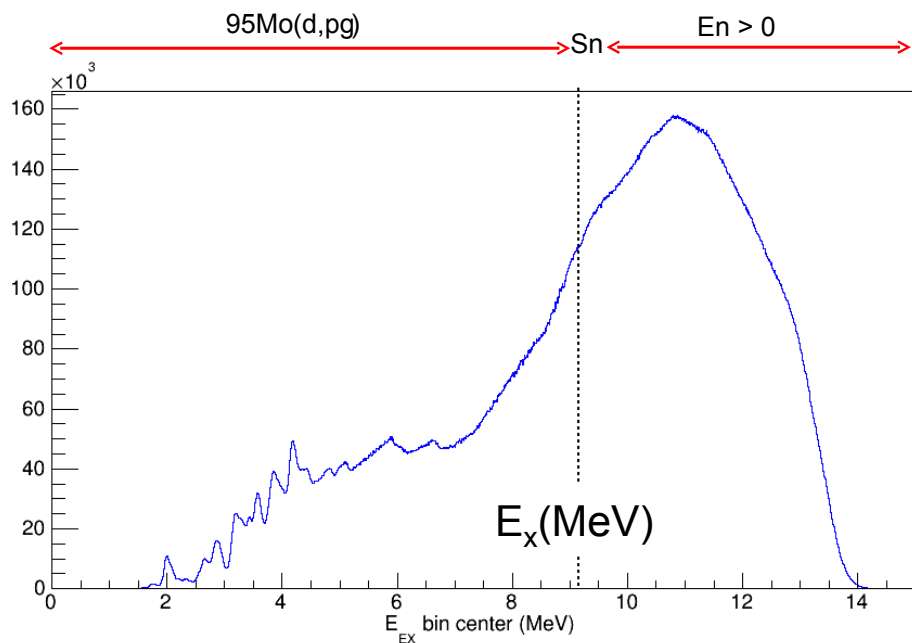
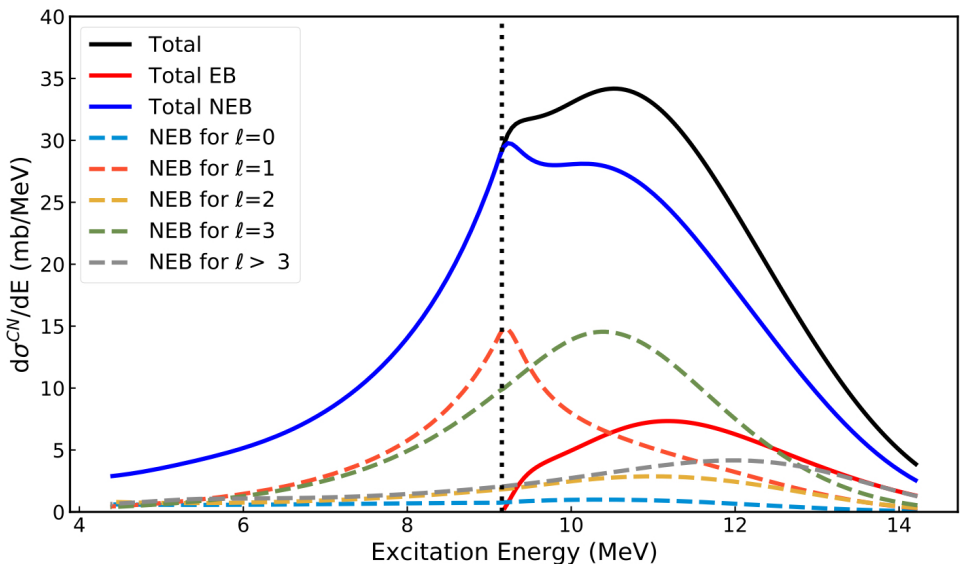
Calculate how  $(d,p)$  forms compound nucleus  $(E_x, J, \pi)$

➤ Deduce  $G^{\text{CN}}$  by fit to  $P(d,p\gamma)$  accounting for  $F^{\text{CN}}$

$$P_{p\gamma}(E_x, \theta) = \sum_{J, \pi} F_{dp}^{\text{CN}}(E_x, J, \pi, \theta) G_{\gamma}^{\text{CN}}(E_x, J, \pi)$$



## Potel d breakup calculations

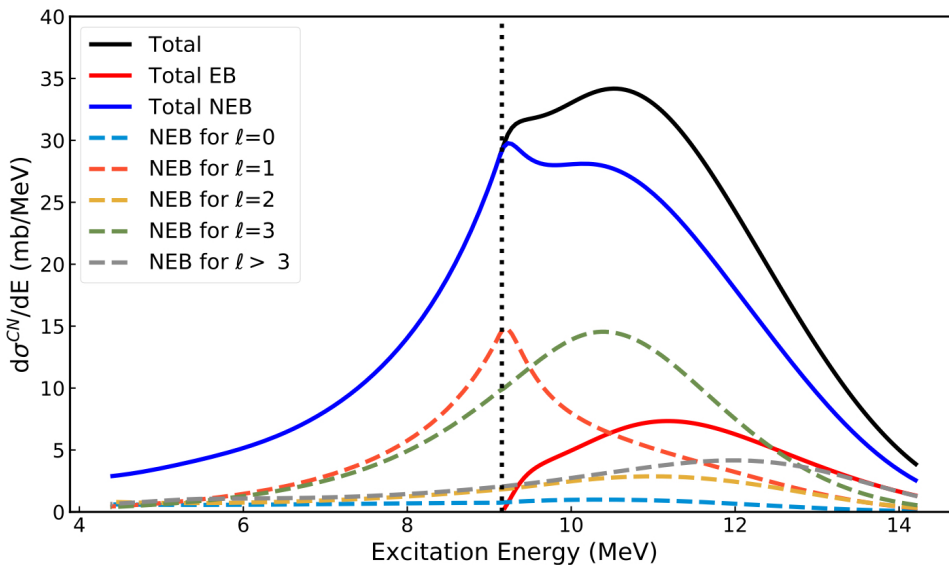


$^{95}\text{Mo}(\text{g.s.})=5/2^+$   
 $\ell=0 \Rightarrow 2^+, 3^+$  entry states



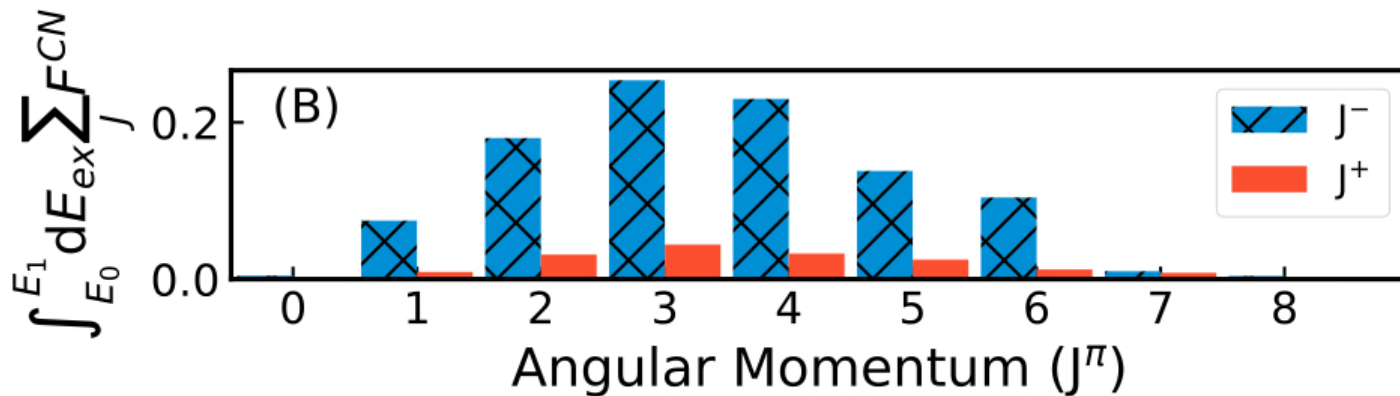
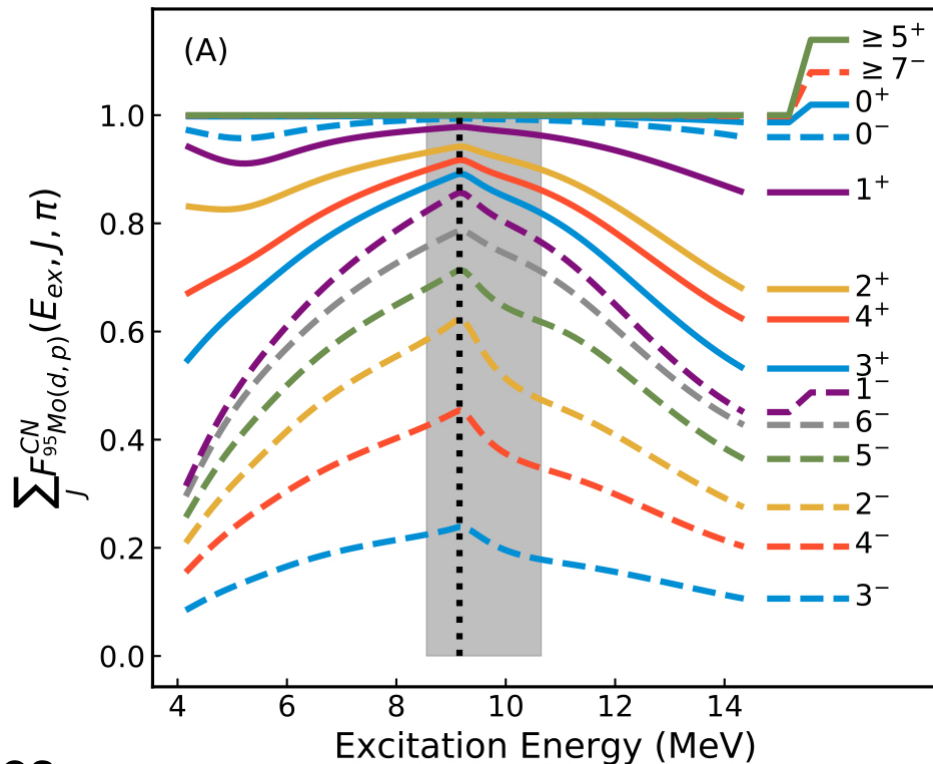
# RUTGERS Potel model for d breakup and $J^\pi$ distributions

## Potel d breakup calculations

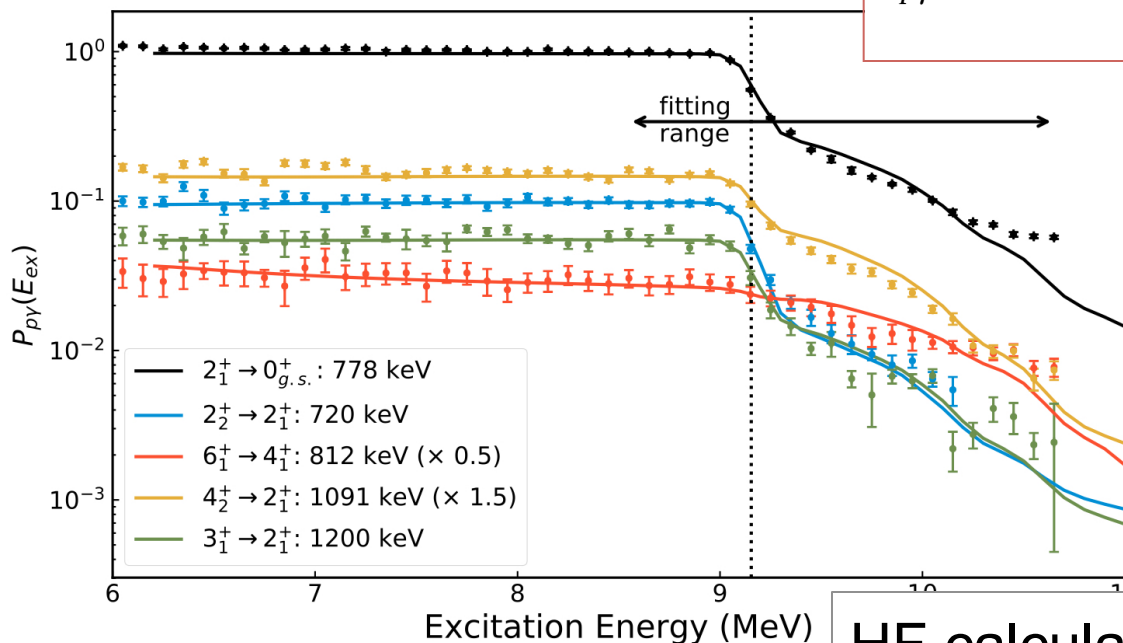


$^{95}\text{Mo}(\text{g.s.})=5/2^+$

$l=0 \Rightarrow 2^+, 3^+$  entry states



$$P_{p\gamma}(E_x, \theta) = \sum_{J, \pi} F_{dp}^{\text{CN}}(E_x, J, \pi, \theta) G_{\gamma}^{\text{CN}}(E_x, J, \pi)$$



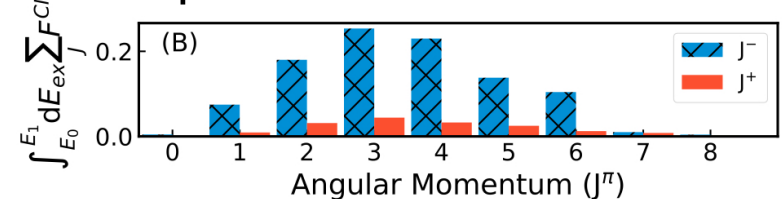
Surrogate (d,p $\gamma$ ) data

HF calculations (Jutta Escher)

- FCN from Potel
- Bayesian fit to observed  $P(d,p\gamma)$ 
  - Simple level density: Gilbert & Cameron
    - No norm to n resonance spacings
  - Simple Lorentzian  $\gamma$  strength function
    - No  $\langle \Gamma(\gamma) \rangle$

➤  $G^{\text{CN}}(E_x, J, \pi)$

$^{96}\text{Mo}$  spin distribution from Potel



G. Potel et al, PRC 92, 034611(2015)



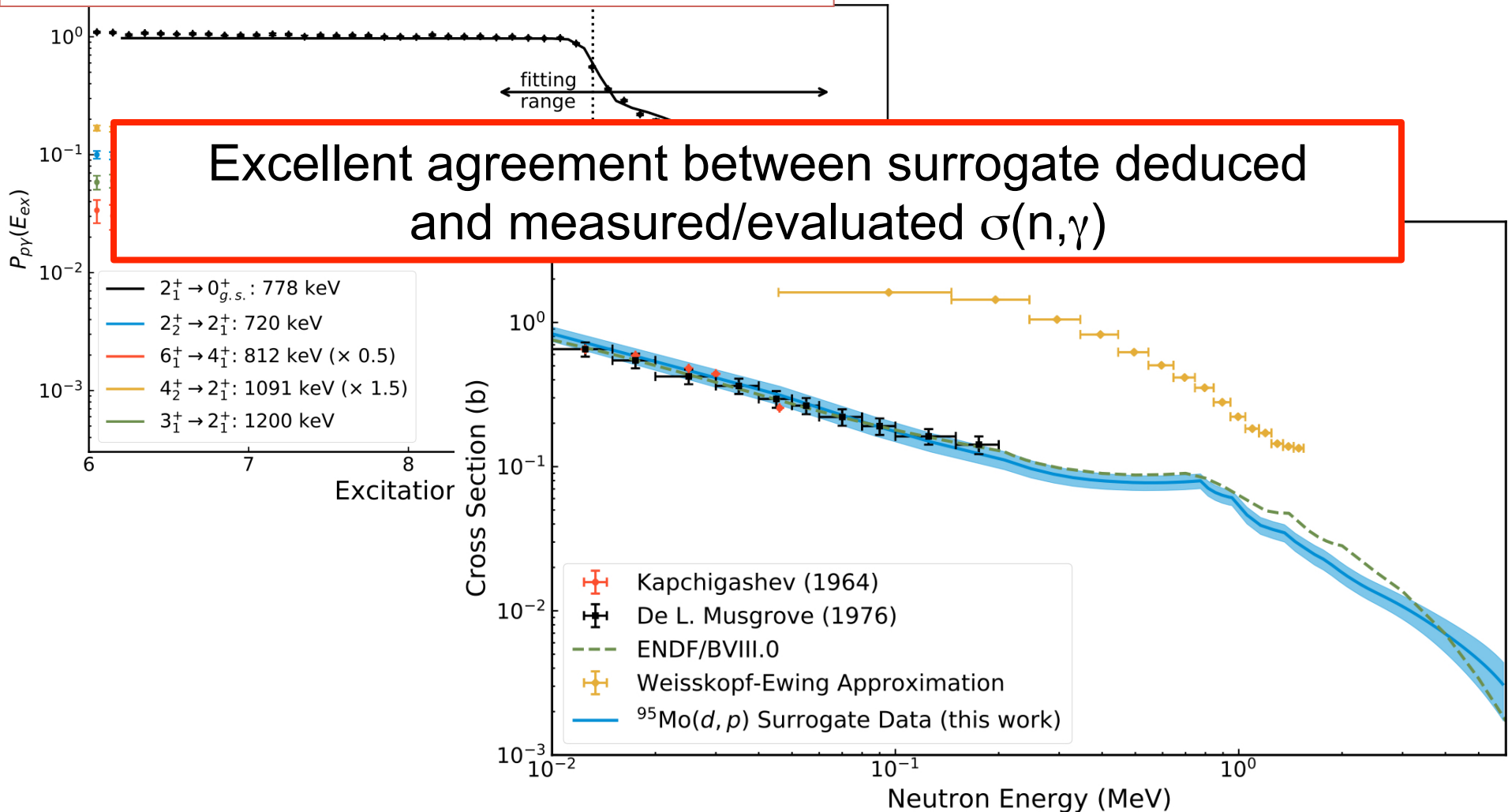
$$P_{p\gamma}(E_x, \theta) = \sum_{J, \pi} F_{dp}^{CN}(E_x, J, \pi, \theta) G_{\gamma}^{CN}(E_x, J, \pi)$$

$$\sigma_{n\gamma}(E_n) = \sum_{J, \pi} \sigma_n^{CN}(E_x, J, \pi) G_{\gamma}^{CN}(E_x, J, \pi)$$

- Deduce  $G^{CN}(E_x, J, \pi)$  from fit to data
- Calculate  $\sigma^{CN}$  w/ Koning-Delaroche optical potentials
- Deduce  $\sigma(n,\gamma)$  vs  $E_x$

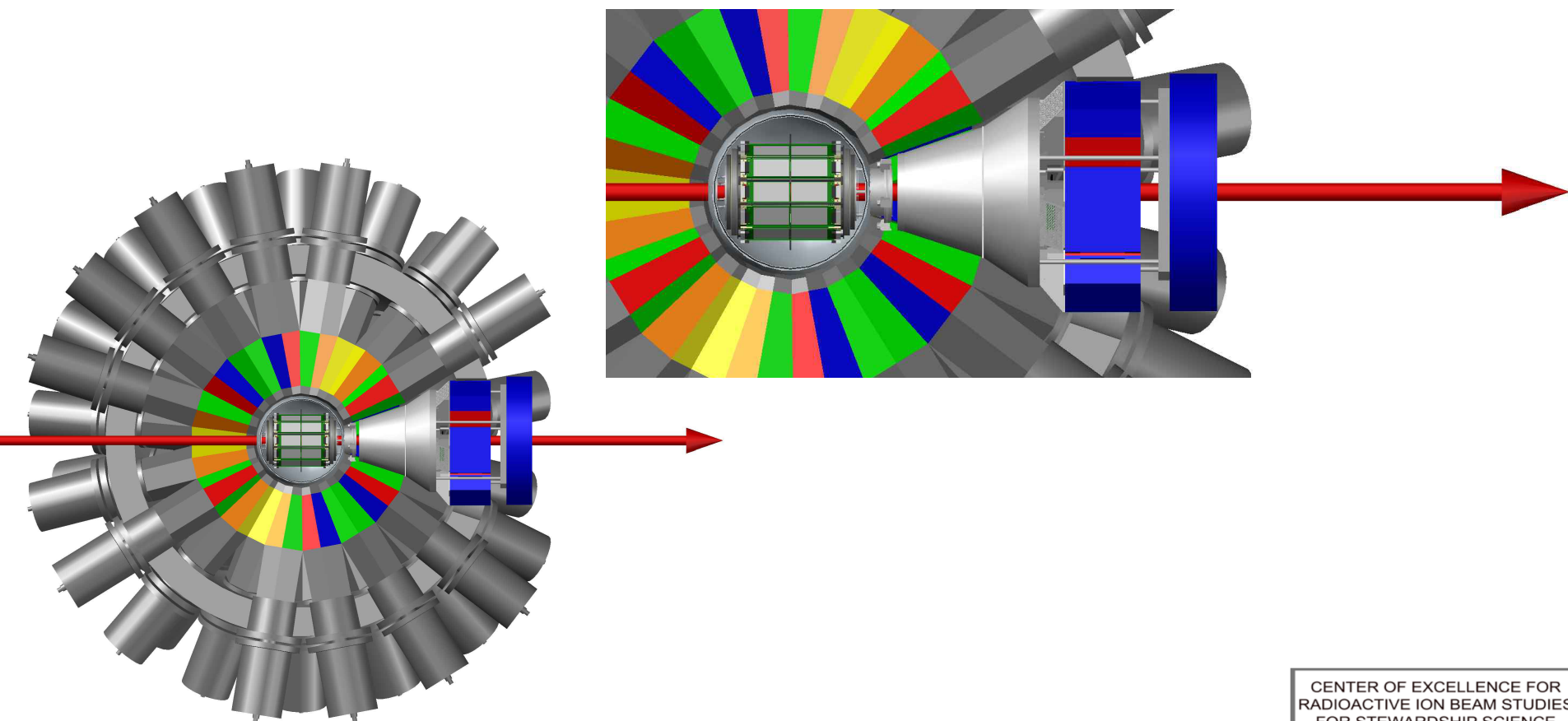


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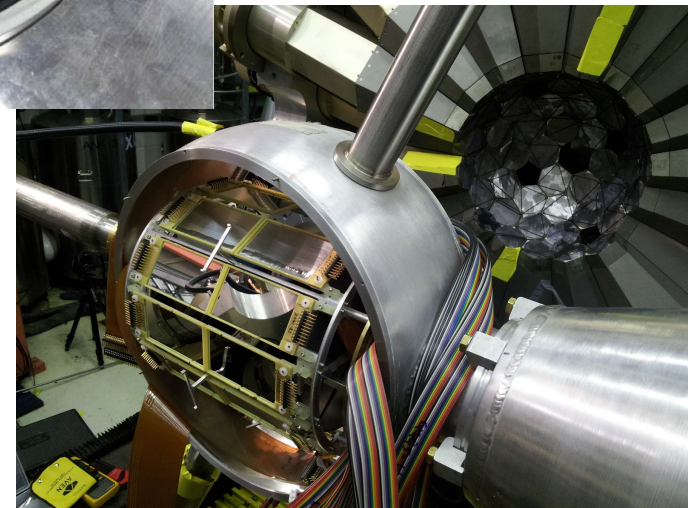
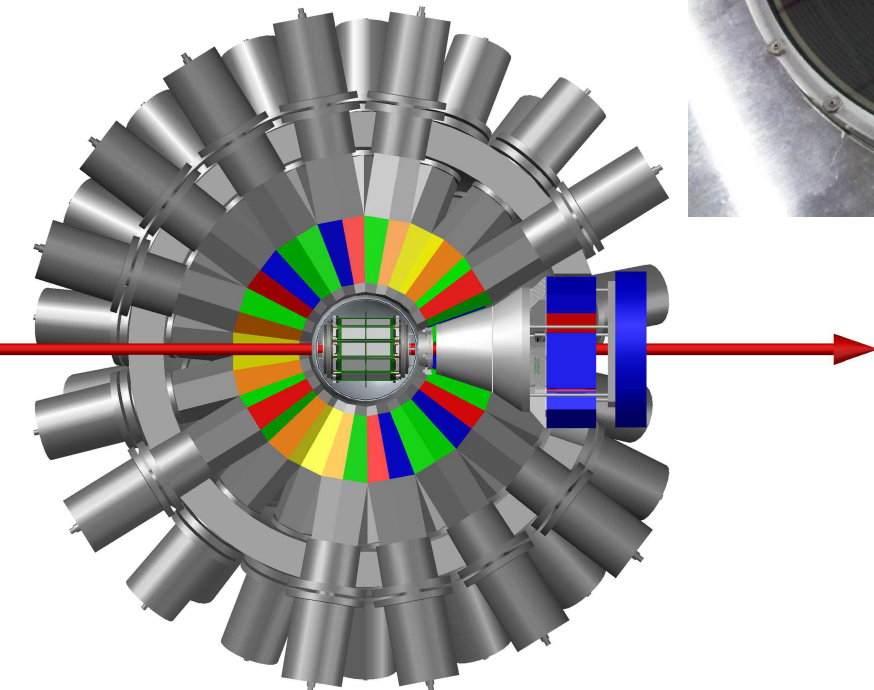
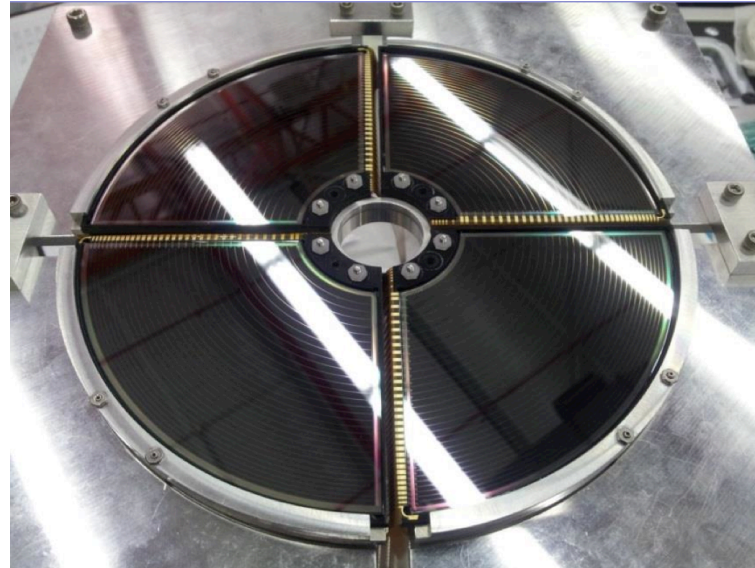


$$\sigma_{n\gamma}(E_n) = \sum_{J, \pi} \sigma_n^{CN}(E_x, J, \pi) G_{\gamma}^{CN}(E_x, J, \pi)$$

## Coupling charged particle &amp; gamma detector arrays

GODDESS: Gammasphere ORRUBA  
Dual Detectors for Experimental Structure Studies

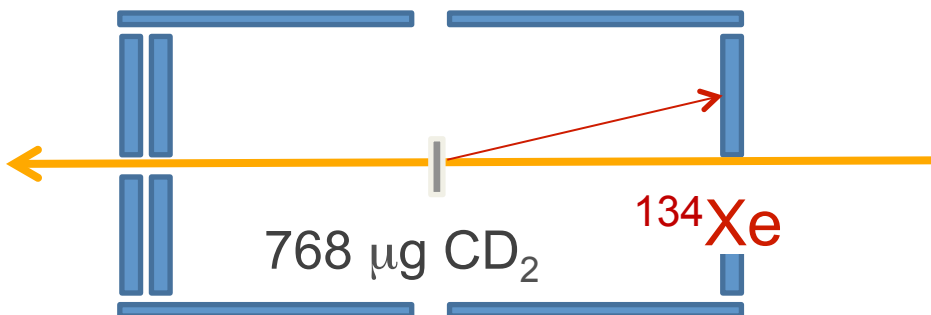
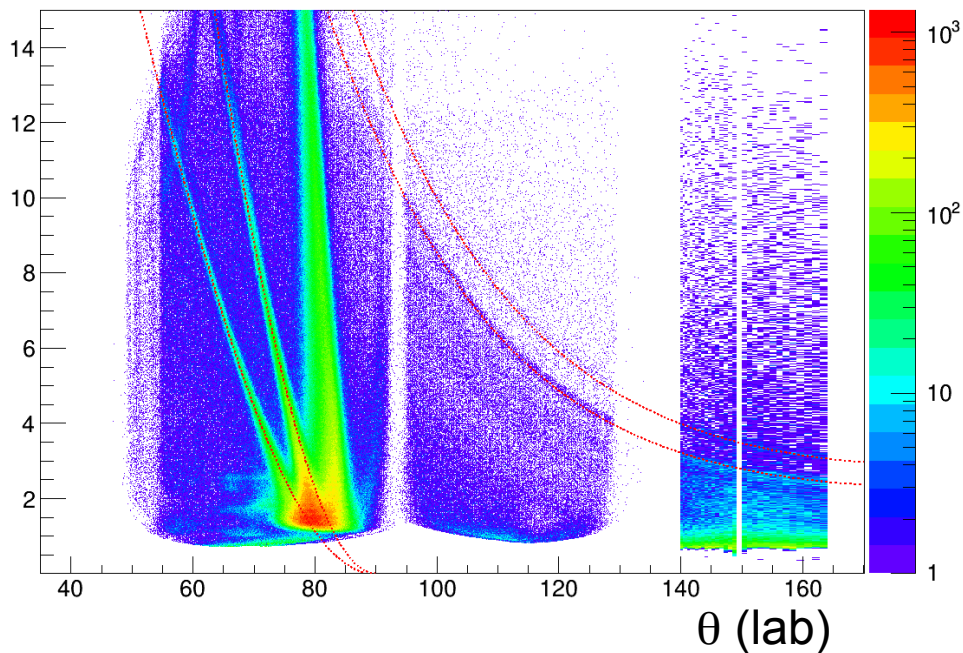
# GODDESS: Gammasphere ORRUBA Dual Detectors for Experimental Structure Studies



Oak Ridge Rutgers University Barrel Array + endcaps

N=80 isotone

Energy vs. Angle complete range



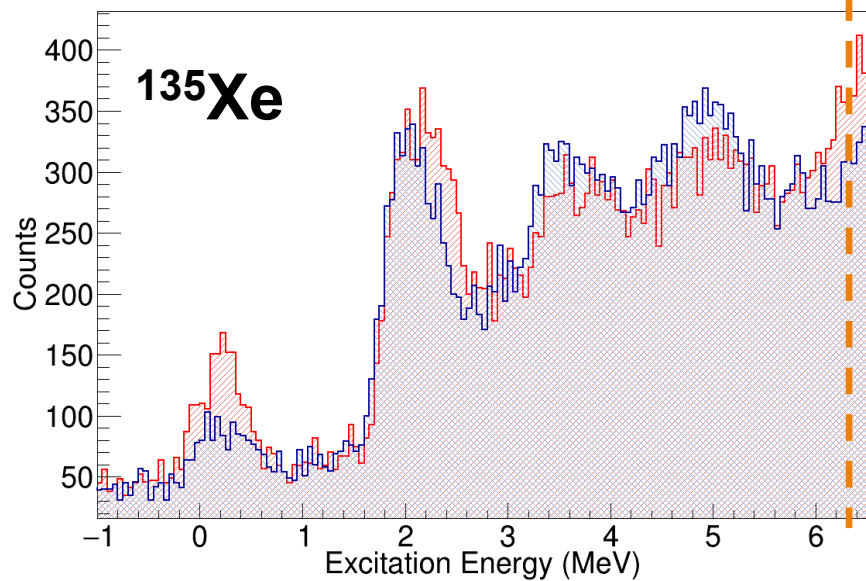
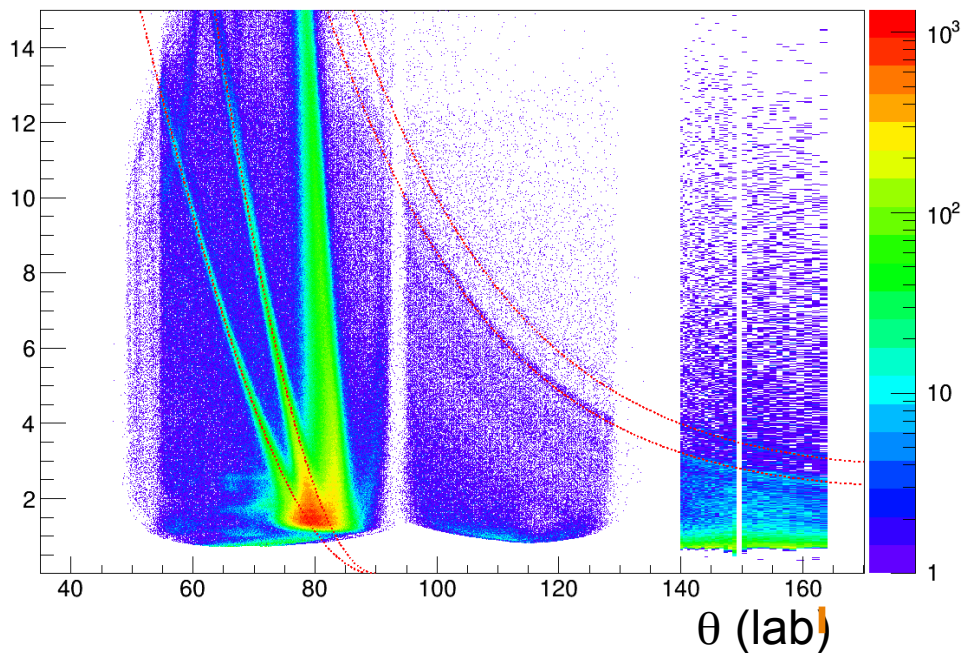
A. Lepailleur, private communication

FRIB TA July 2018

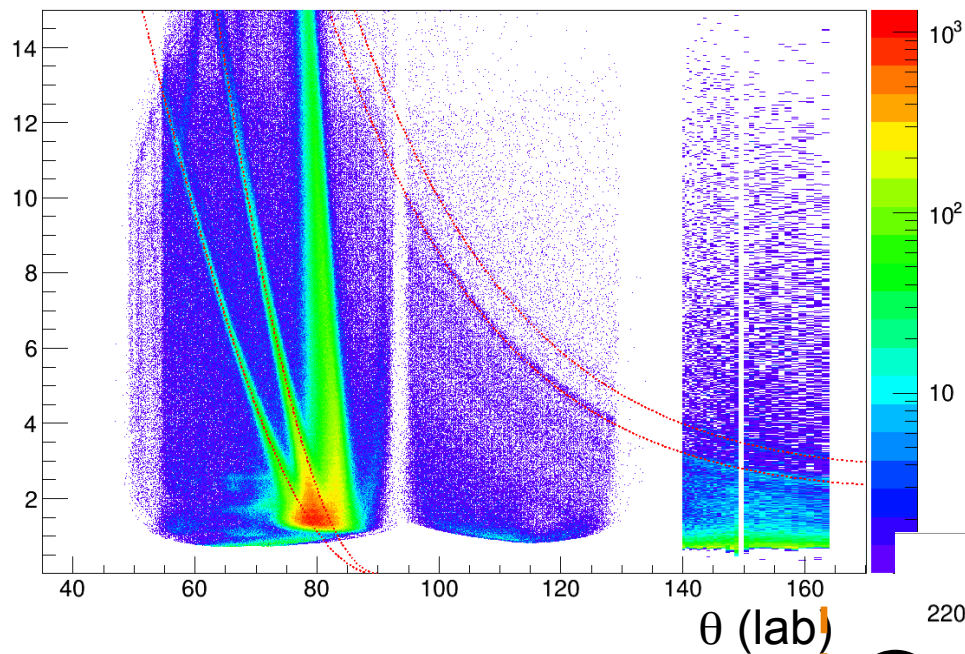


N=80 isotone

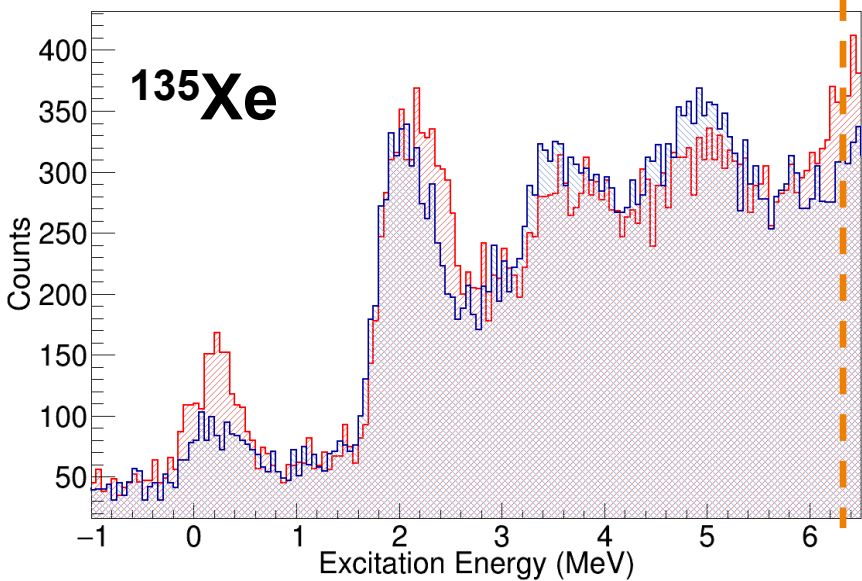
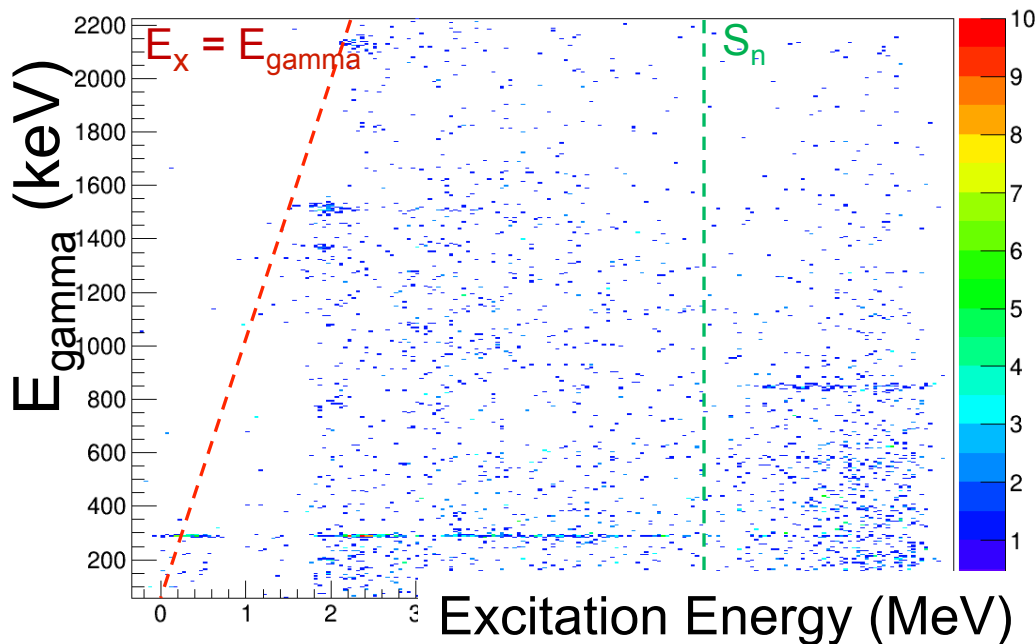
Energy vs. Angle complete range

 $^{135}\text{Xe}$   $E_x$  spectrumRed: QQQ5 (large  $\theta$ )Low- $\ell$  transfer important for DSDBlue: SX3 ( $90^\circ < \theta < 135^\circ$ )

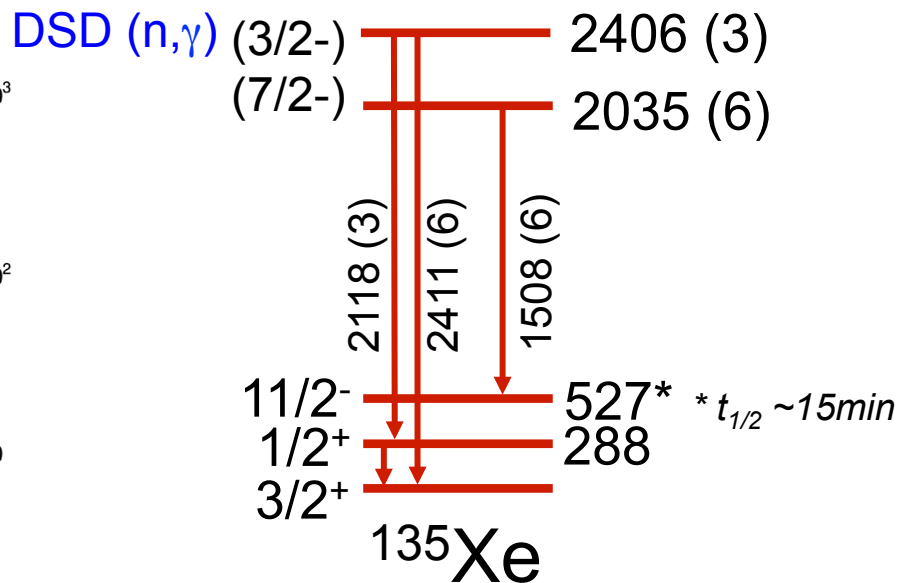
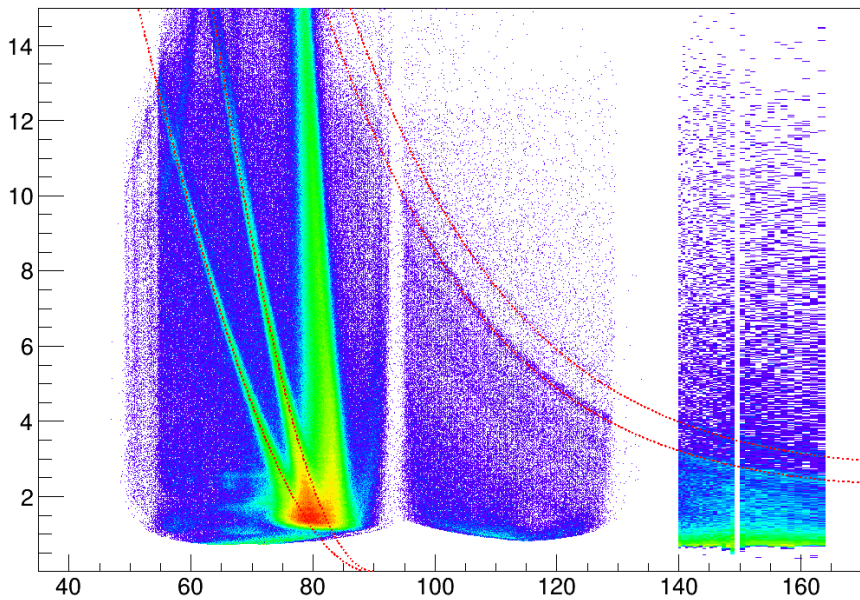
Energy vs. Angle complete range



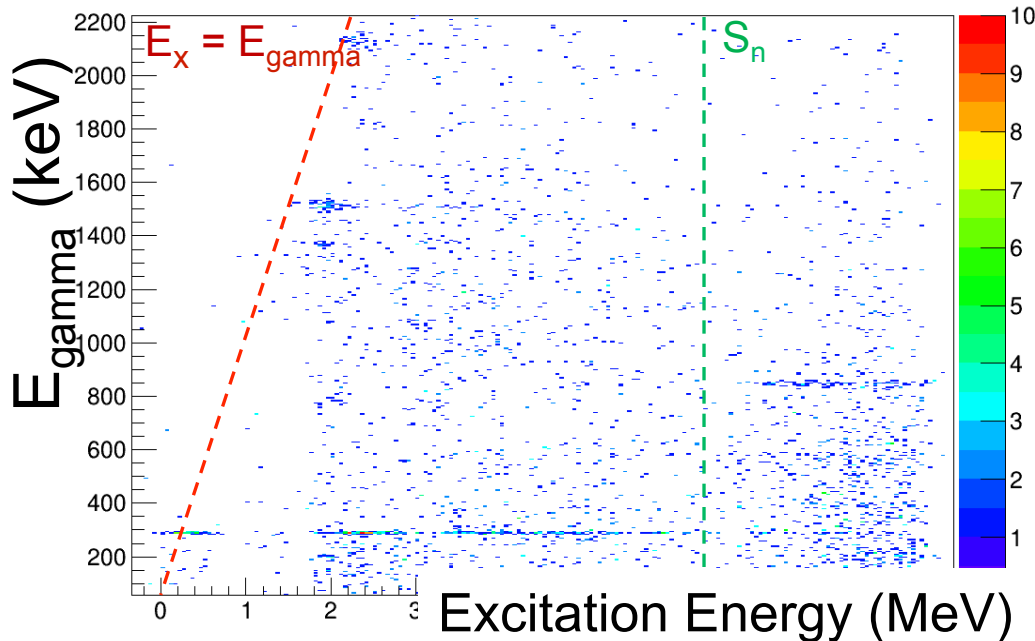
Gamma Energy BGO Veto vs. Excitation Energy QQQ5 U0



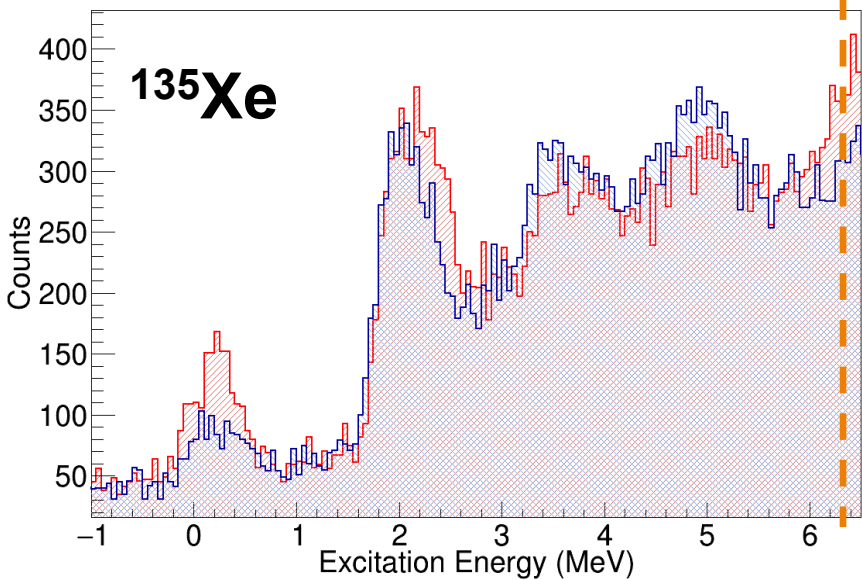
Energy vs. Angle complete range



Gamma Energy BGO Veto vs. Excitation Energy QQQ5 U0

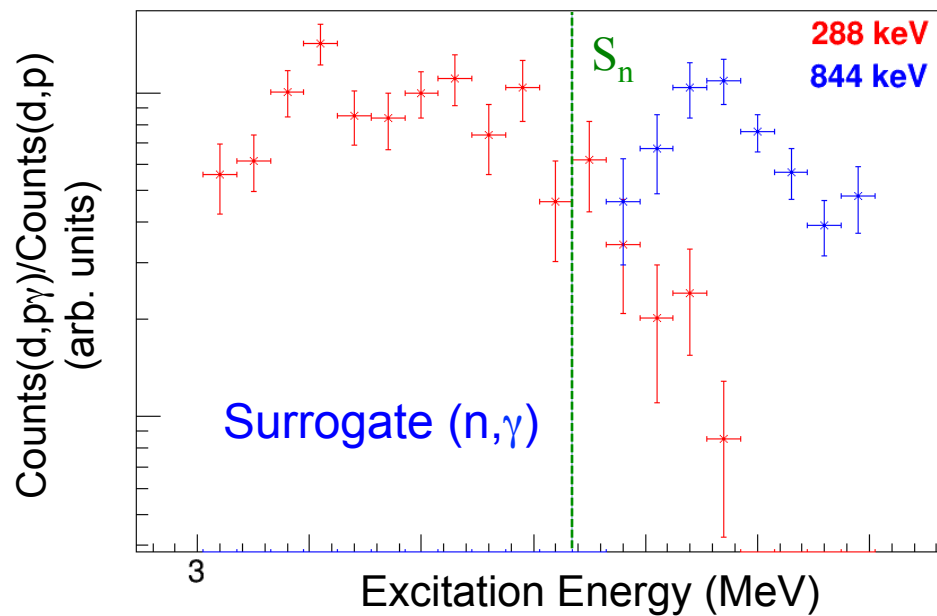
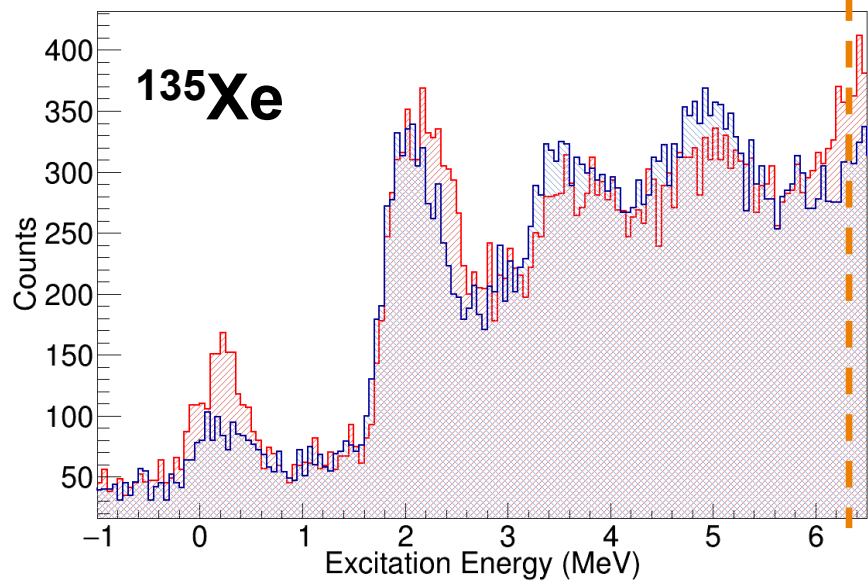
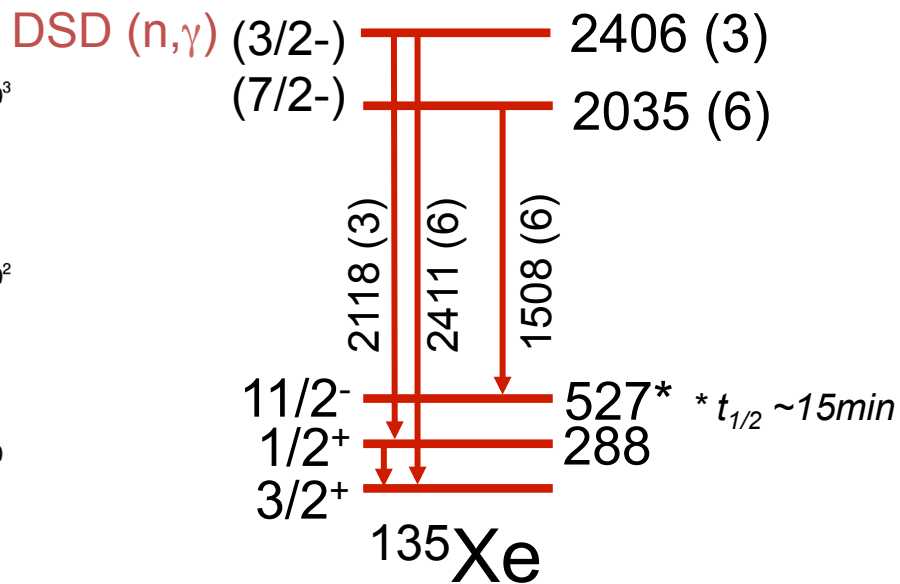
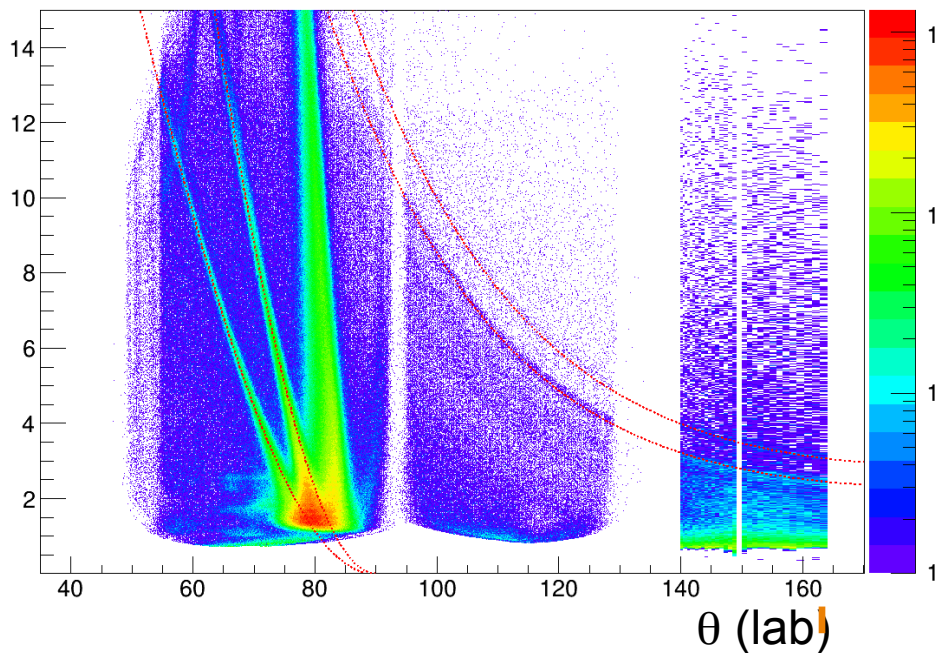


$\theta$  (lab)





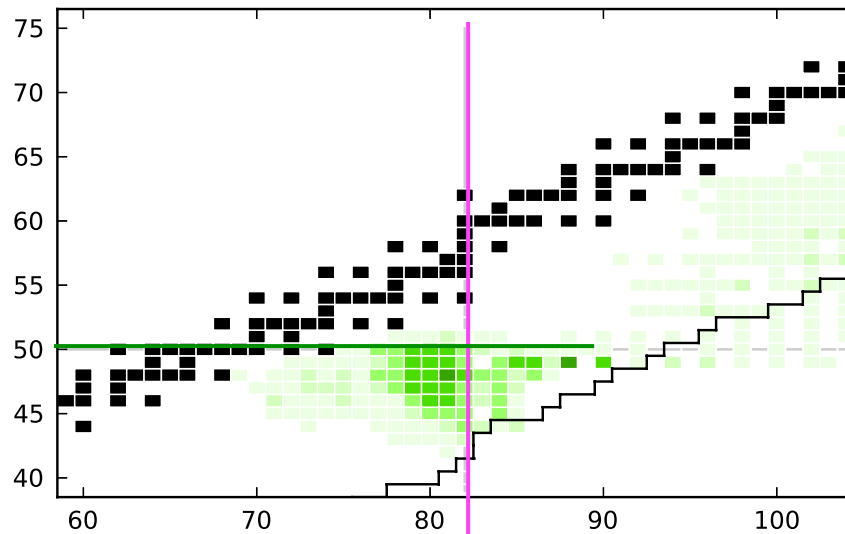
Energy vs. Angle complete range



Goal:  $\approx^{132}\text{Sn}$  isotopes important for n-star mergers  
 Will have to wait for FRIB

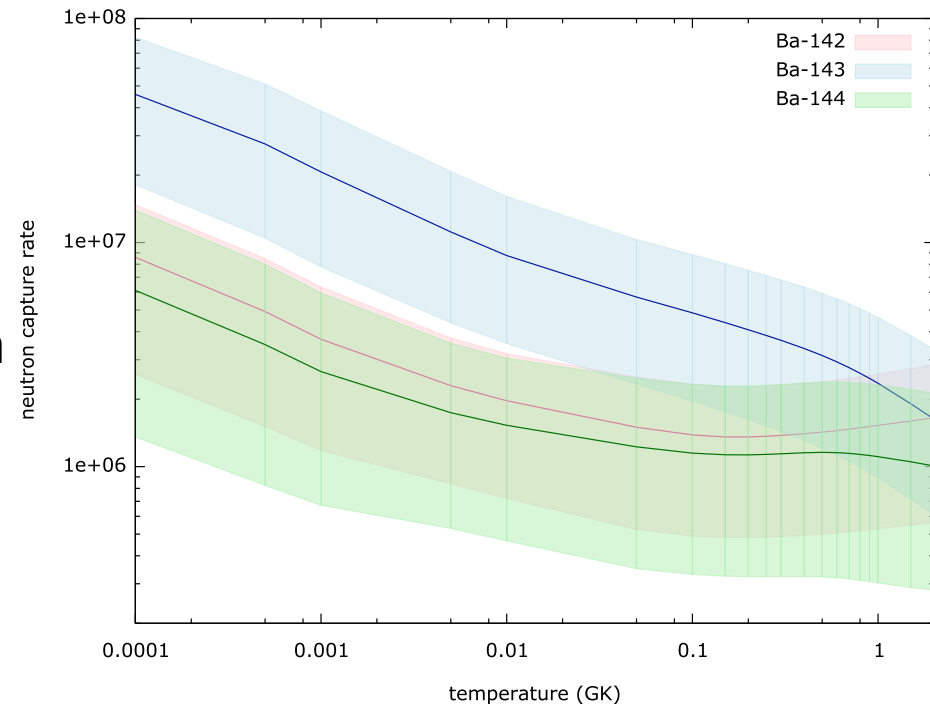
What can we do “now”?

- CARIBU  $^{252}\text{Cf}$  fragment beams
- Approved to measure  $^{143}\text{Ba}(d,p\gamma)$  w/ GODDESS



Ba( $n,\gamma$ ) rates from [Mum16]

Would be first surrogate  $(n,\gamma)$  on fission fragment to constrain  $(n,\gamma)$  in this region

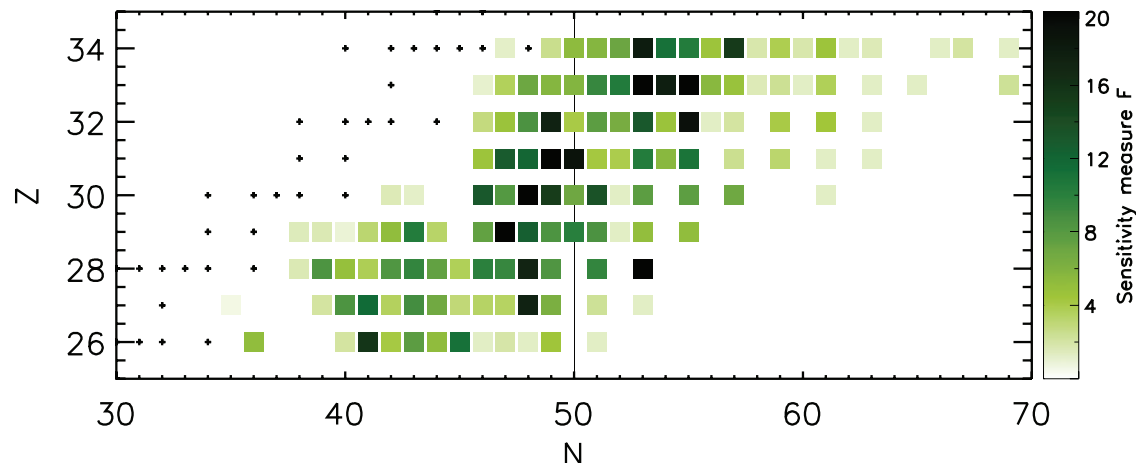
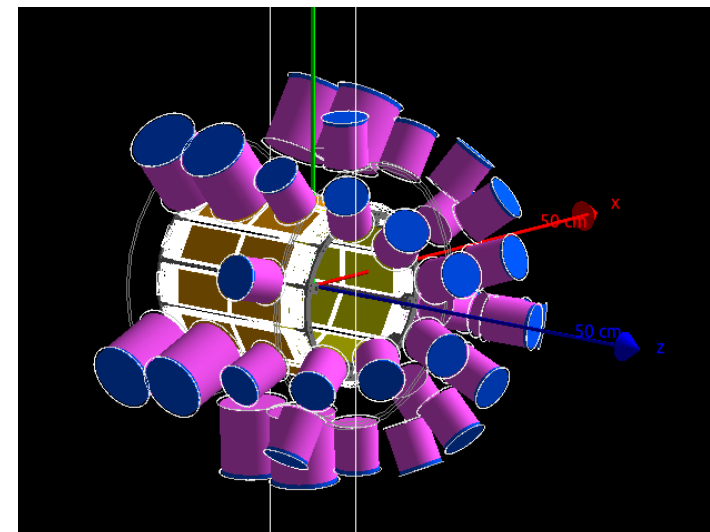
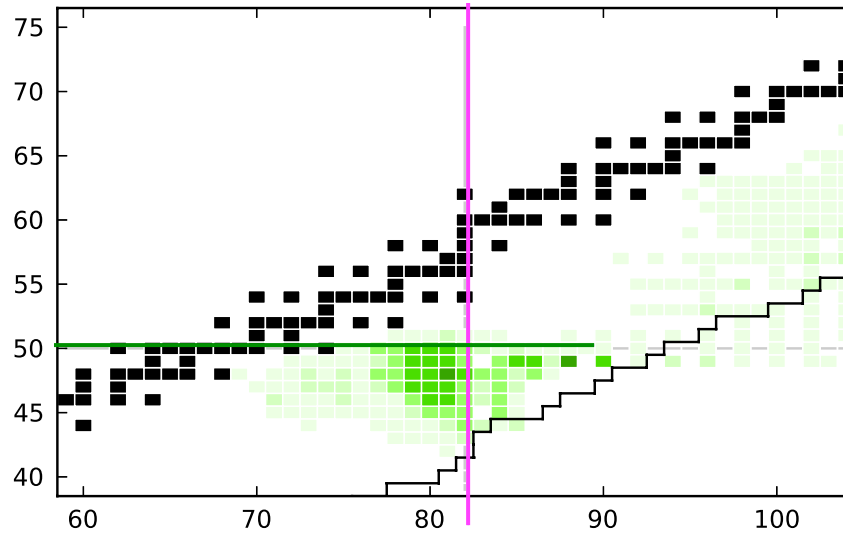


Goal:  $\approx^{132}\text{Sn}$  isotopes important for n-star mergers

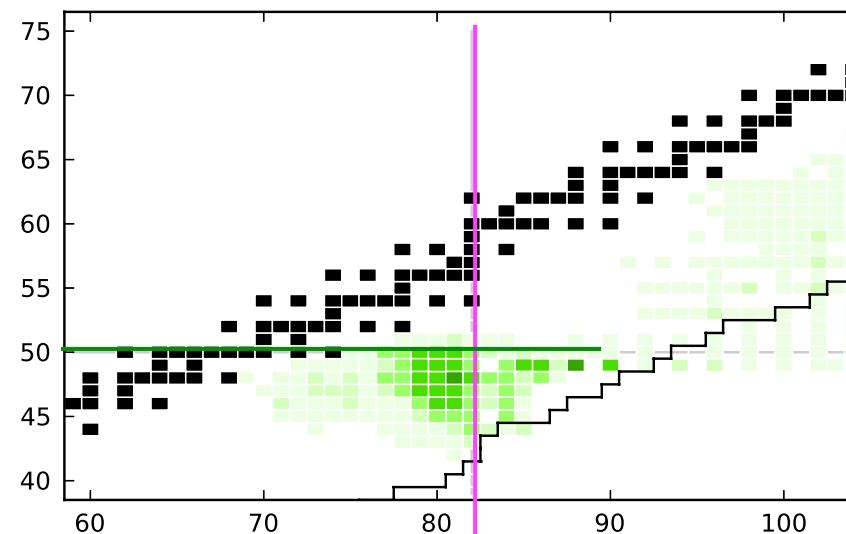
Will have to wait for FRIB

What can we do “now”?

- CARIBU  $^{252}\text{Cf}$  fragment beams
- Approved to measure  $^{143}\text{Ba}(d,p\gamma)$  w/ GODDESS
- NSCL  $\approx 80$  fast beams
- Approved to measure  $^{80}\text{Ge}(d,p\gamma)$  w/ ORRUBA+HAGRiD



- Understanding abundances from NSM r process is sensitive to  $(n,\gamma)$  rates, especially near shell closures, e.g.,  $^{130}\text{Sn}$ , and weakly bound nuclei with low level density
  - Need neutron transfer  $(d,p)$  to inform direct-semi-direct capture
- Unknown competition between DSD and CN  $(n,\gamma)$ 
  - Need validated surrogate for  $(n,\gamma)$
- Demonstrated that  $(d,p\gamma)$  is valid surrogate for  $(n,\gamma)$
- Demonstrated ability to measure  $(d,p)$  protons in coincidence with gamma rays
- Near term
  - $^{143}\text{Ba}(d,p\gamma)$
  - $^{80}\text{Ge}(d,p\gamma)$
- Goal: FRIB  $(d,p\gamma)$   
e.g., with  $^{130}\text{Sn}$  beams



# EXTRA SLIDES



## Hauser Feshbach code STAPRE (modified)

- Level Density
  - Use known discrete levels
  - Gilbert-Cameron level density
    - Fermi Gas model matched to Constant Temperature
- Gamma-ray strength function “usual” forms
  - E1, M1, E2, M2
  - E1 strength: modified, energy- and temperature-dependent Lorentzian
- Bayes’ method to obtain parameter constraints
- Monte-Carlo sampling of parameters to calculate cross sections

