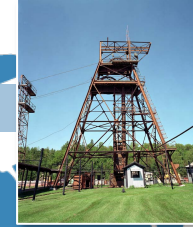


# NOVA STATUS

Patricia Vahle, On Behalf of the NOvA Collaboration  
William and Mary

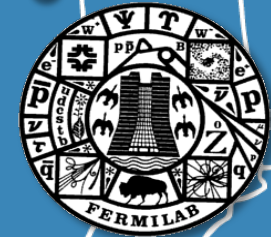
# NOvA

2



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- NOvA addresses compelling questions from P5 Science driver: Investigation of the Physics of Neutrino Mass
  - What is the Neutrino Mass Hierarchy?
  - Is there CP symmetry violation in neutrinos?
  - What is the pattern of mixing?
  - Is there more to it than  $3 \times 3$  PMNS?
  
- Long-baseline neutrino oscillation experiment
  - Study muon neutrino disappearance, electron neutrino appearance, and flavor-independent (NC) disappearance with both neutrinos and antineutrinos
  - 2 Detectors separated by 810 km
  - High power and high purity neutrino and antineutrino beams



# NOvA Collaboration



238 collaborators at 49 institutions across 7 countries



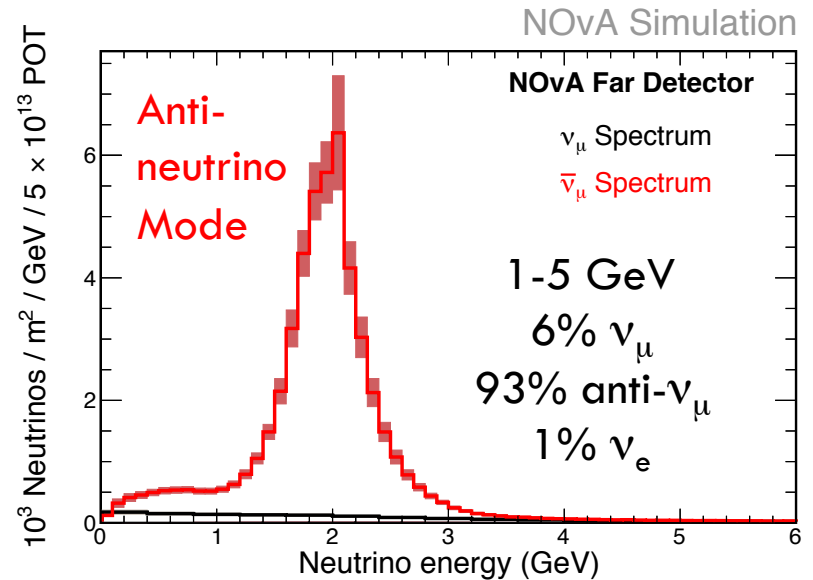
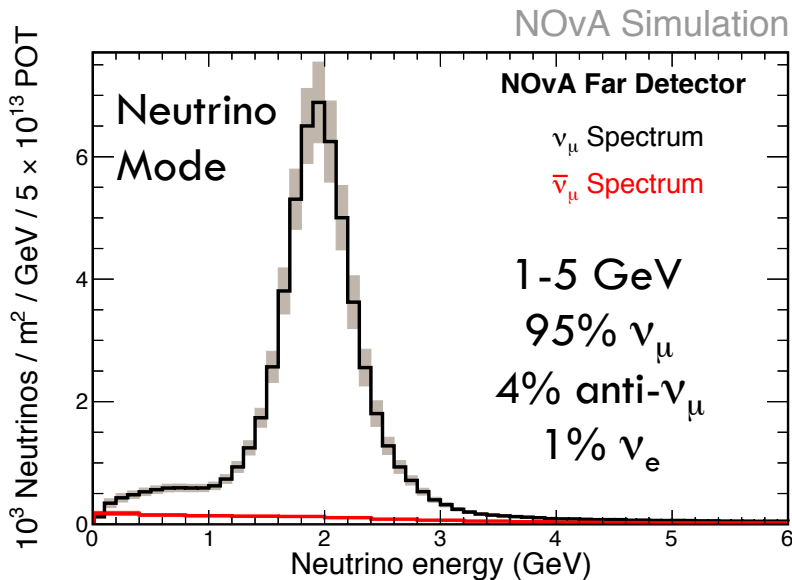
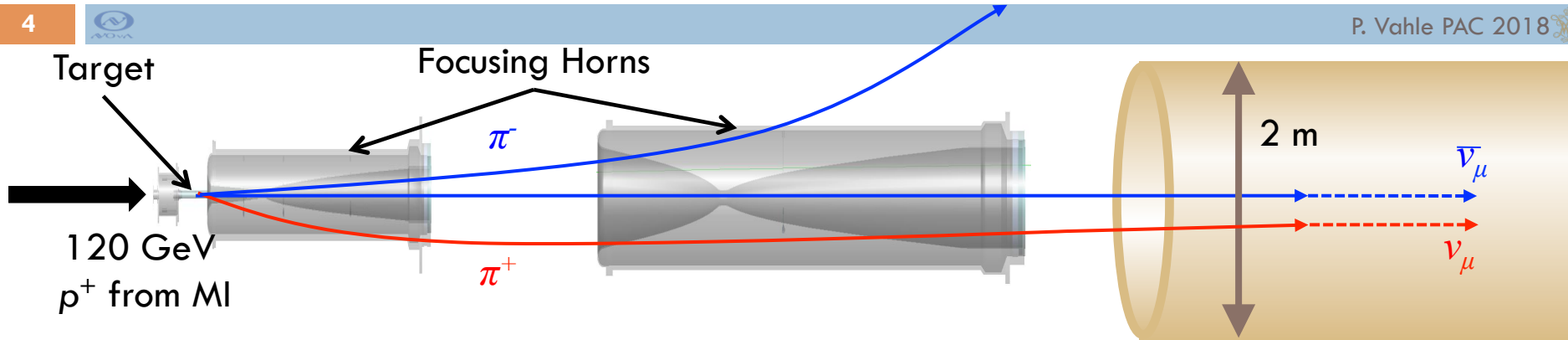
19 Remote Control Rooms across the globe

# Making an off-axis neutrino beam

4



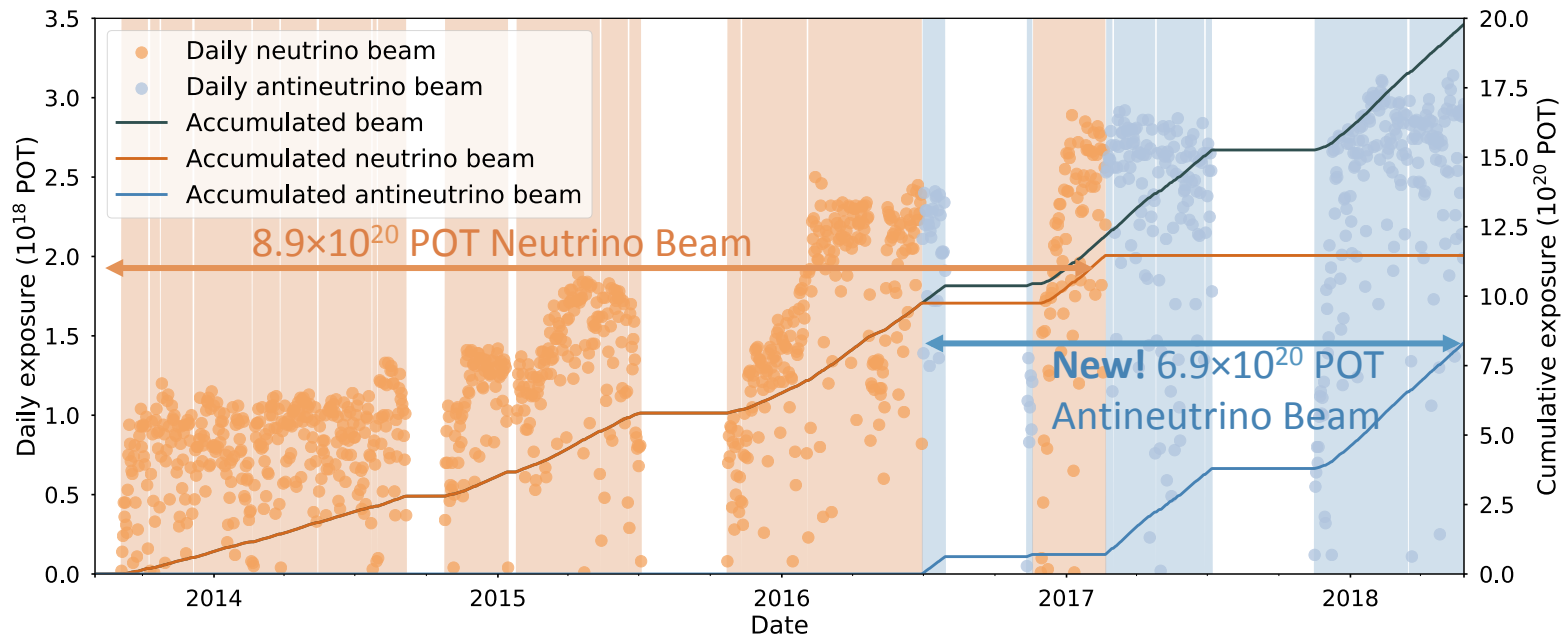
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# Beam Performance



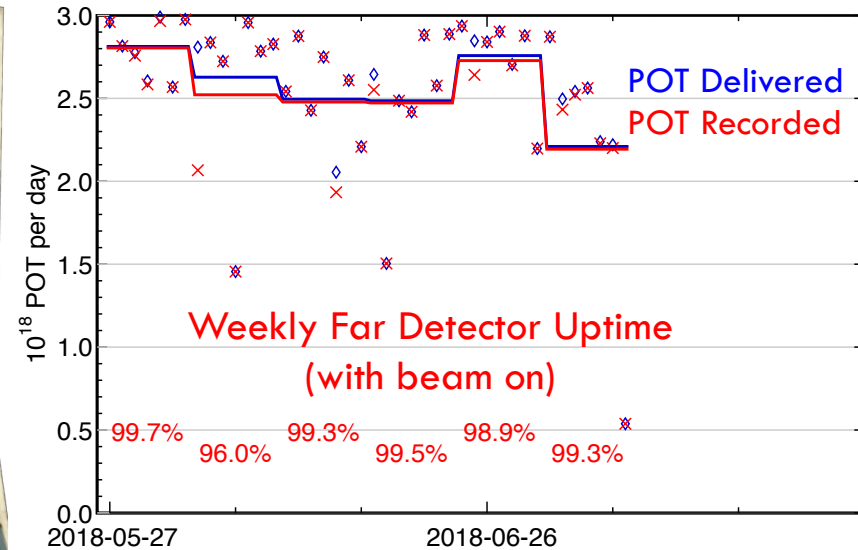
- Run at 700 kW design goal since January 2017
  - ▣ 5.64e20 POT delivered in FY18, exceeding goal, best year so far by 10%
  - ▣ 34 weeks of running at 98.7% uptime



# NOvA Detectors

6

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Far Detector  
14-kton  
896 planes  
344,064 channels

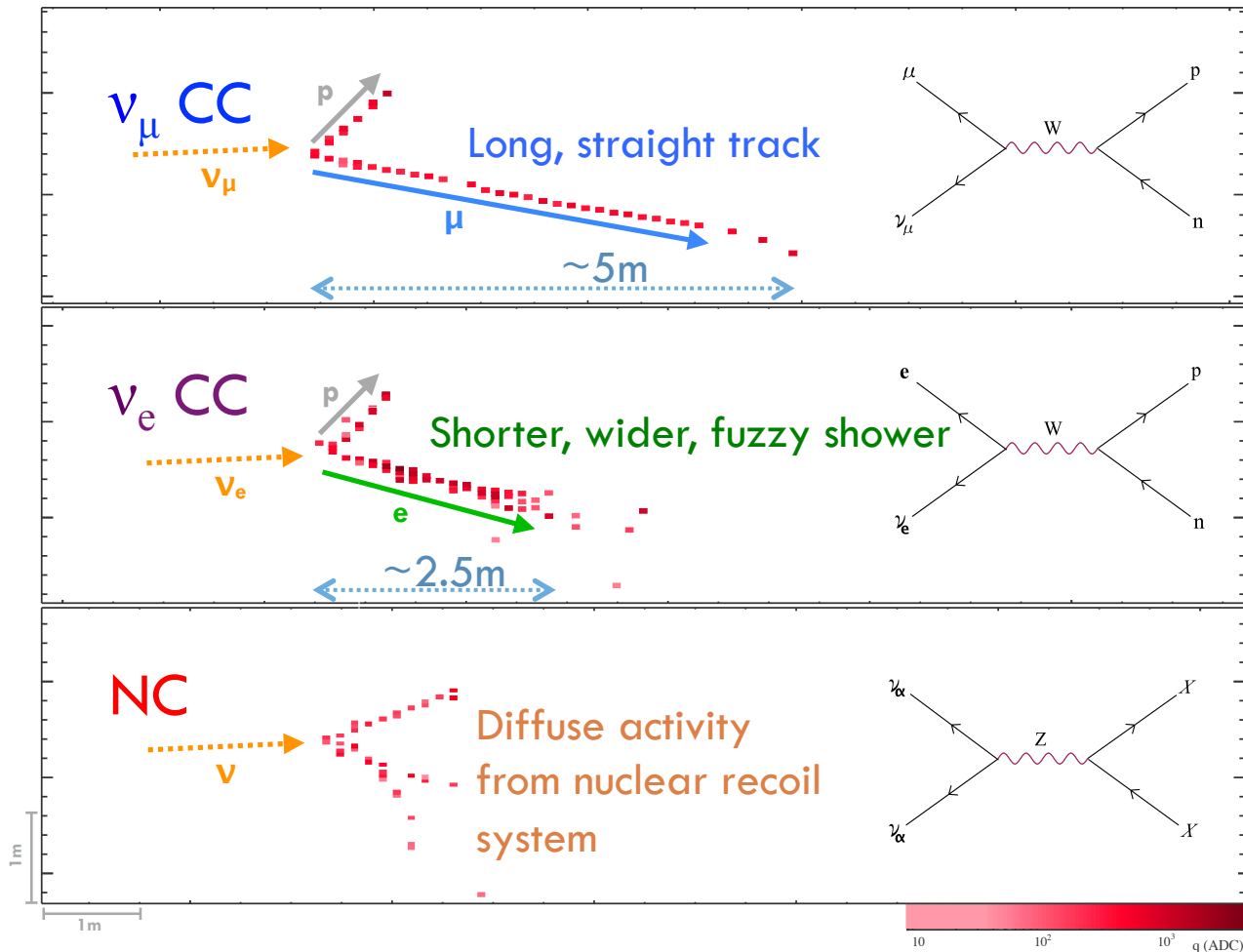
Near Detector  
290-ton

- Functionally identical detectors designed for electron ID
  - Low Z materials (PVC+Liquid Scint.)
  - 65% active
- ND: underground at FNAL
- FD: on the surface in Ash River, MN

More than 99.9% of 300k+ FD channels are operational!

# Event Selection

## Candidate Events from ND data



NOvA has pioneered the use of computer vision and deep learning techniques for event selection

- Calibrated hit maps are inputs to Convolutional Visual Network (CVN)
- Series of image processing transformations applied to extract abstract features
- Extracted features used as inputs to a conventional neural network to classify the event

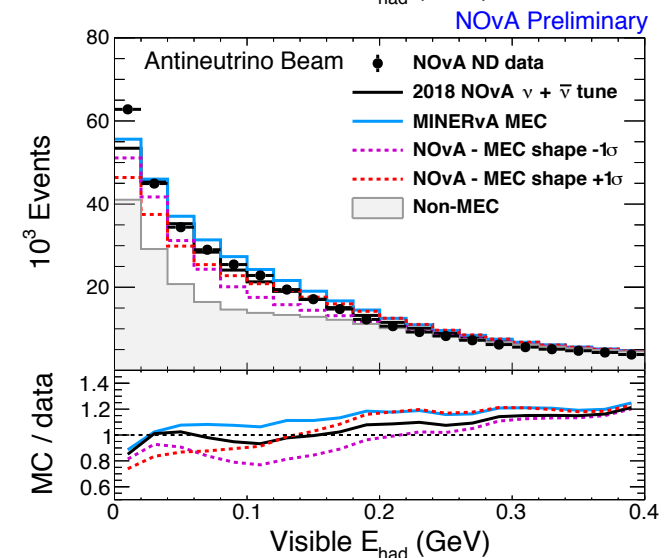
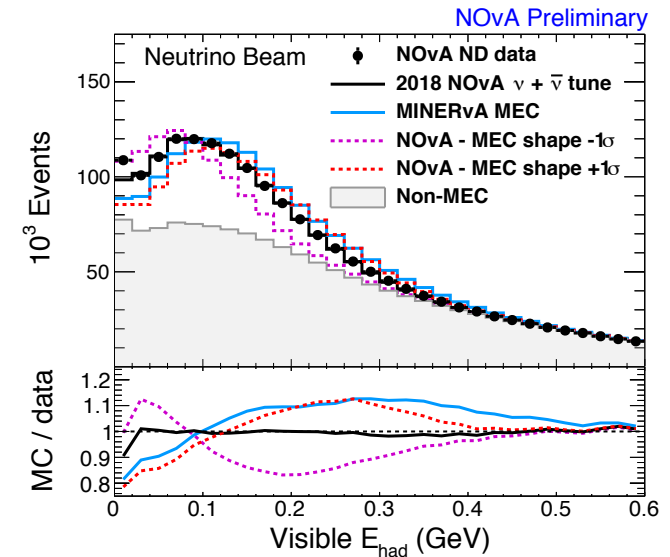
# NOvA Simulation

8



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- All NOvA oscillation analyses use ND data to predict FD
  - 2 detector technique naturally mitigates uncertainties on flux and cross sections
  - Simulation used for acceptance and resolution corrections, helps characterize impact of residual systematic uncertainties
- Perfect ND data/mc not required, tuning makes for a more robust prediction
  - Flux from PPFX—package developed on MINERvA based on hadron production data (Phys. Rev. D 94, 092005. 2016)
  - neutrino interaction tuning guided by MINERvA experience (arXiv:1705.02932 and arXiv:1601.01888)
  - Custom tuning of GENIE Empirical MEC model
- Workshop with MINERvA planned for this fall to share GENIE tuning techniques

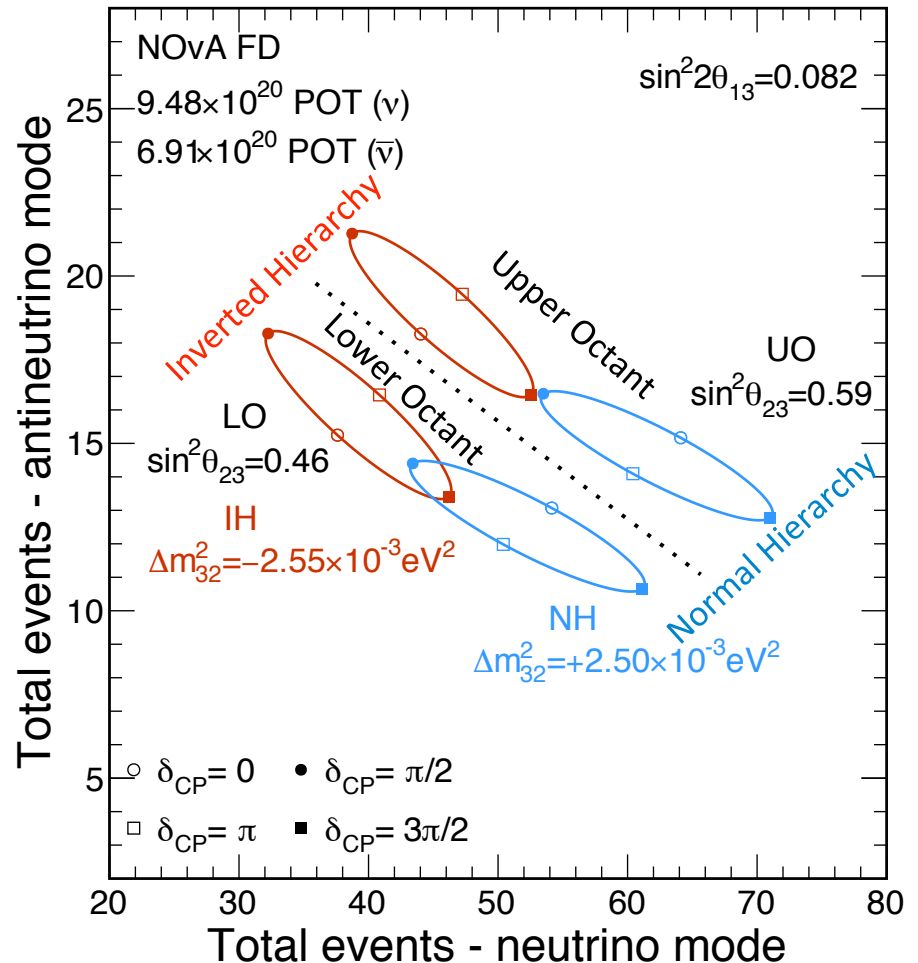




# Analysis Basics



- Compare electron-neutrino appearance probability in neutrinos to antineutrinos
  - ▣ in vacuum with no CP violation, the two should be the same
  - ▣ CP violation enhances oscillation probability for neutrinos while suppressing it for antineutrinos, or vice-versa
  - ▣ matter effects also introduce mass hierarchy dependent neutrino vs. antineutrino differences
  - ▣ upper octant enhances both neutrino and antineutrino oscillation probability, while lower octant suppresses both



# Analysis in more detail

Data

Area-normalized MC

Shape-only systematics

Wrong-sign

10



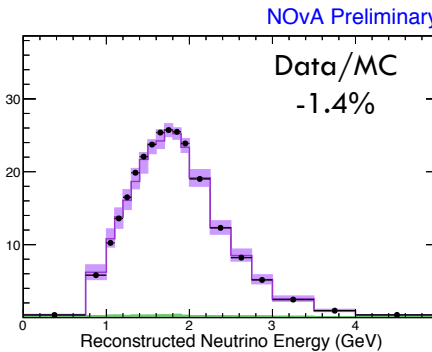
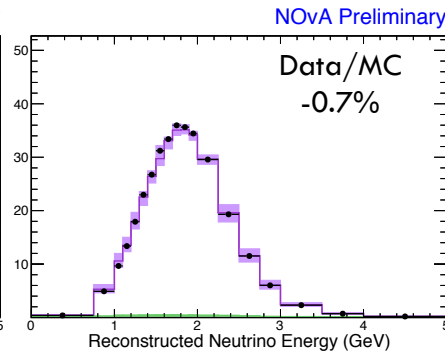
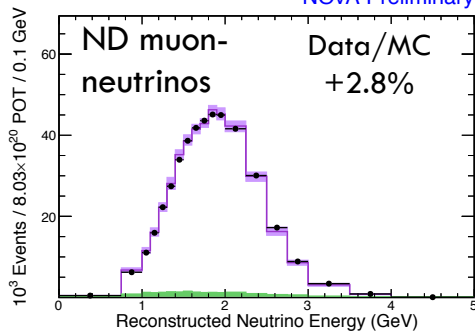
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- Full power comes from joint fit to energy dependence of both disappearance and appearance in neutrinos and antineutrinos
  - Muon neutrino spectra further separated by energy resolution
  - Electron neutrino spectra further separated by event sample purity

Quantile 1

Best Resolution  $\sim 6\%$

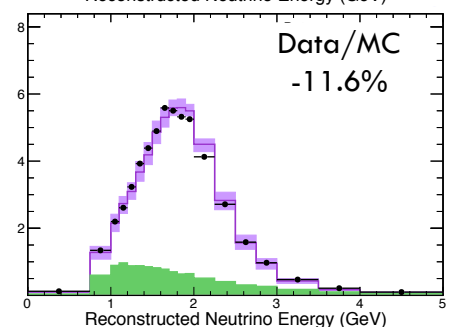
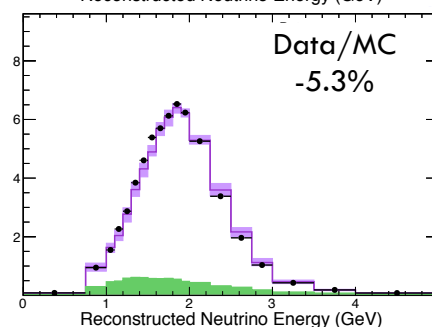
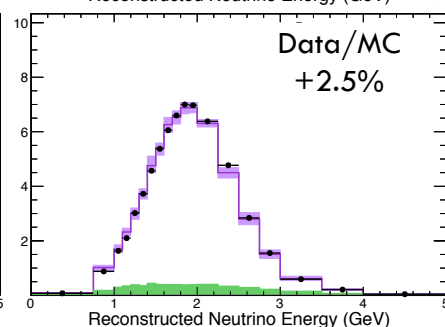
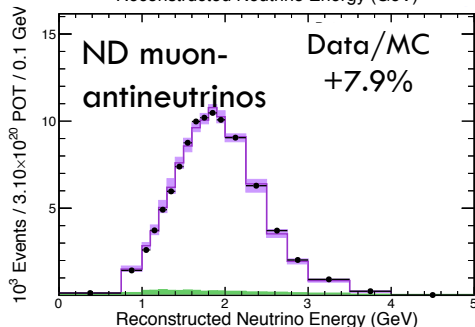
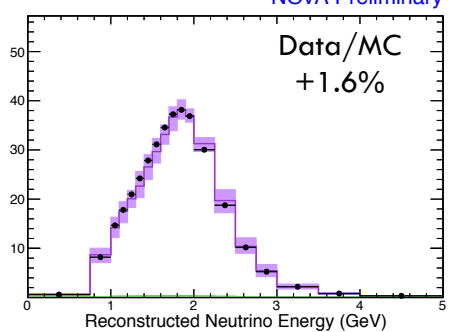
NOvA Preliminary



Quantile 4

Worst Resolution  $\sim 12\%$

NOvA Preliminary



# Appearance Results

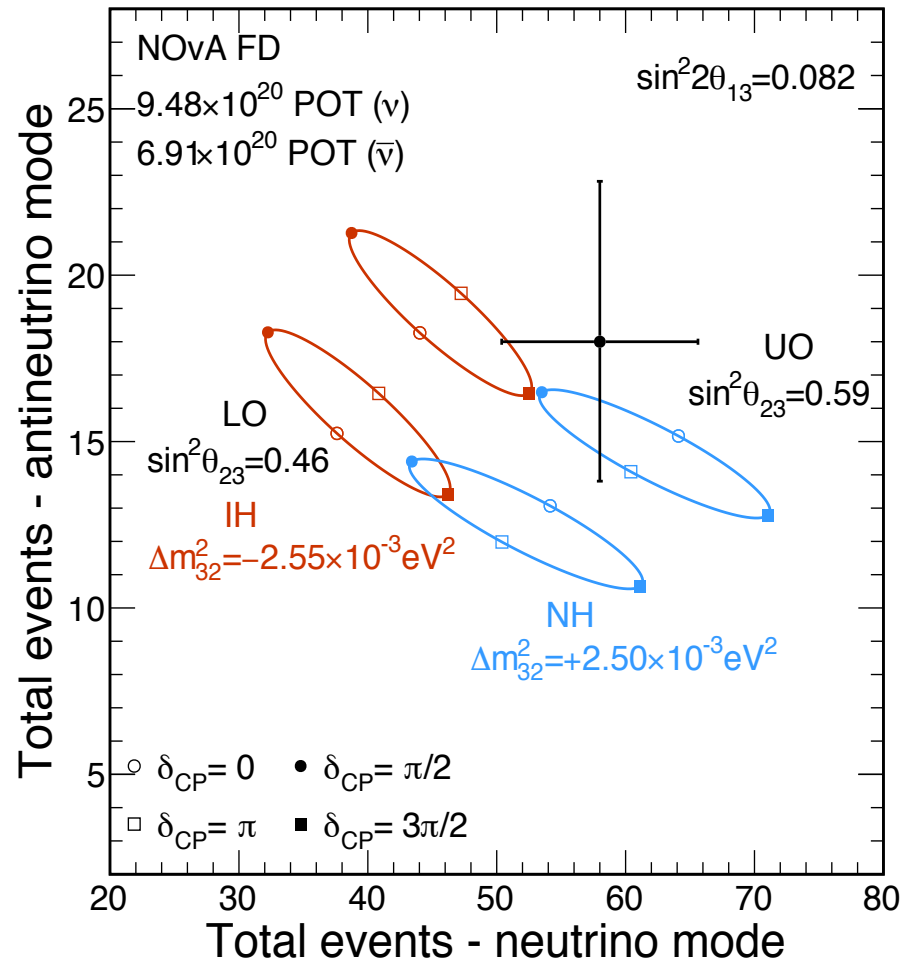
11



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- Neutrino mode
    - 58 events observed
    - 15 bkg expected
  - Antineutrino mode
    - 18 events observed
    - 5 bkg expected
- (including 1 wrong sign)

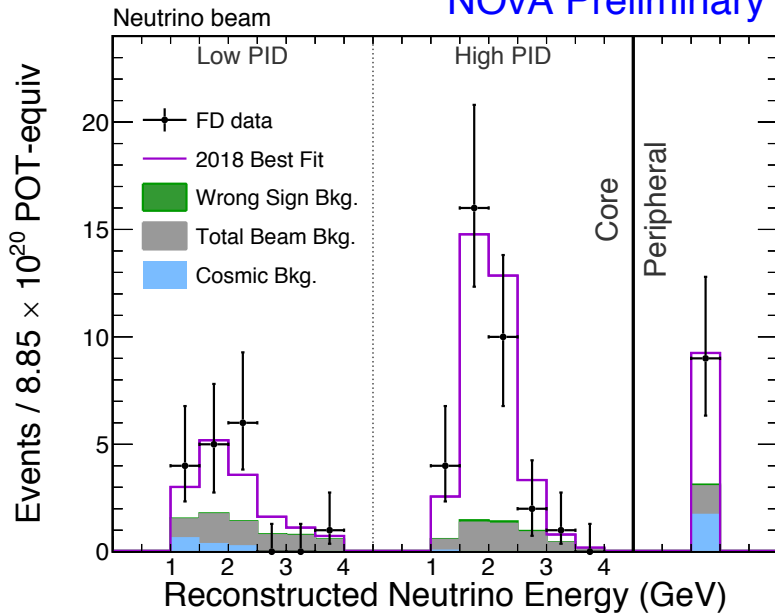
Strong ( $>4$  sigma) evidence of electron antineutrino appearance



# Appearance Spectra



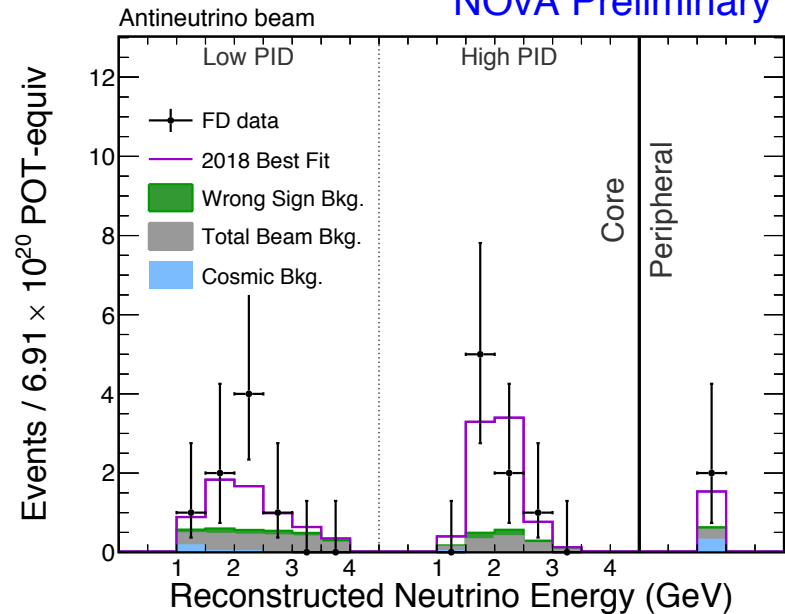
**NOvA Preliminary**



**Neutrino Mode**

Total Observed	58	Range
Total Prediction	59.0	30-75
Wrong-sign	0.7	0.3-1.0
Beam Bkgd.	11.1	
Cosmic Bkgd.	3.3	
Total Bkgd.	15.1	14.7-15.4

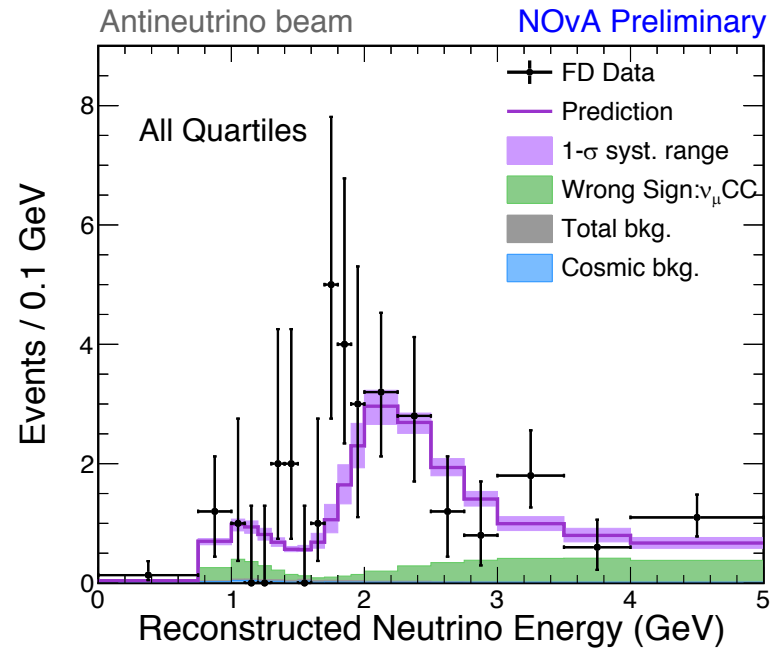
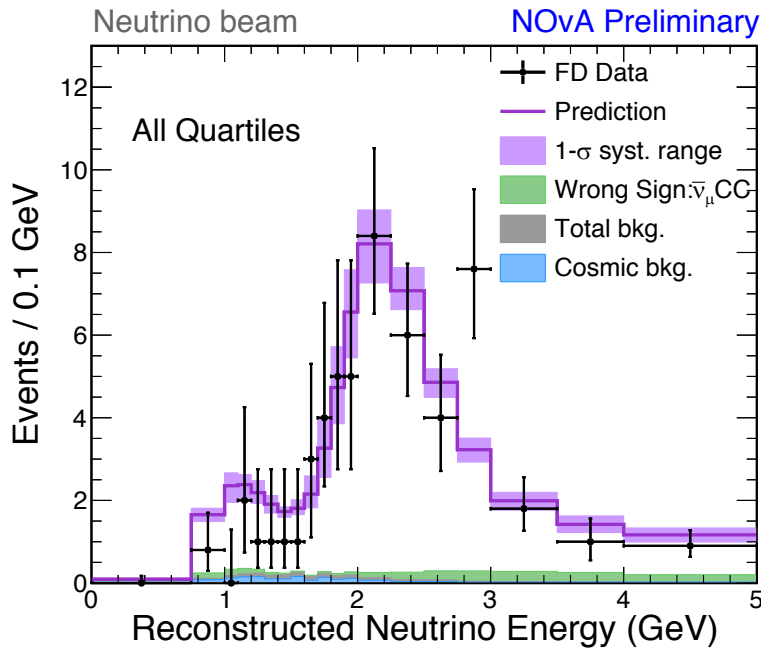
**NOvA Preliminary**



**Antineutrino Mode**

Total Observed	18	Range
Total Prediction	15.9	10-22
Wrong-sign	1.1	0.5-1.5
Beam Bkgd.	3.5	
Cosmic Bkgd.	0.7	
Total Bkgd.	5.3	4.7-5.7

# Disappearance Spectra



<b>Total Observed</b>	<b>113</b>
Best fit prediction	121
Cosmic Bkgd.	2.1
Beam Bkgd.	1.2
Unoscillated	730

Some tension in disappearance of neutrino vs. antineutrino; Antineutrinos show less disappearance. Results are compatible at better than 4% level.

<b>Total Observed</b>	<b>65</b>
Best fit prediction	50
Cosmic Bkgd.	0.5
Beam Bkgd.	0.6
Unoscillated	266

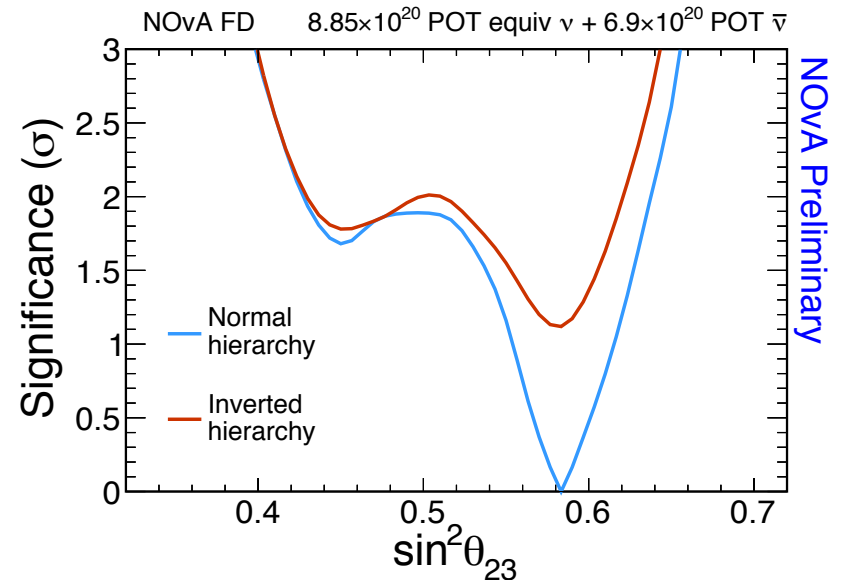
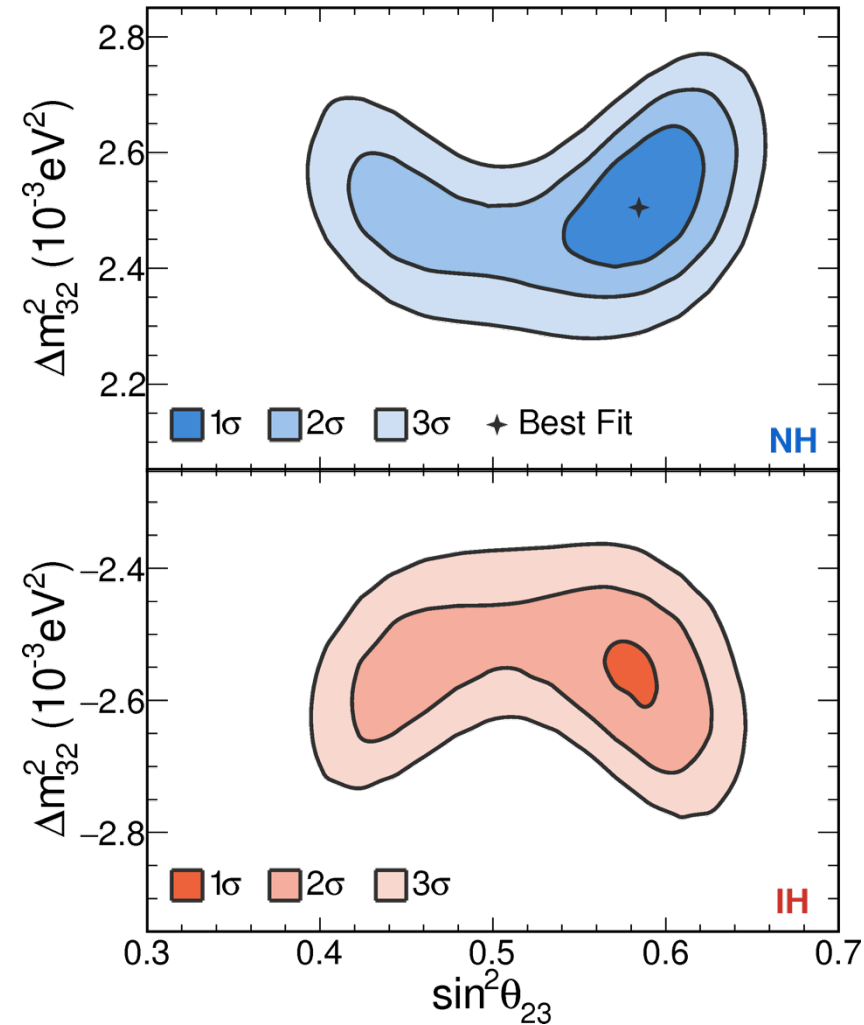
# Appearance and Disappearance Fit Results

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NOvA Preliminary



Best Fit

Normal Hierarchy

$$\Delta m_{32}^2 = 2.51_{-0.08}^{+0.12} \times 10^{-3} \text{eV}^2$$

$$\sin^2 \theta_{23} = 0.58 \pm 0.03$$

(Upper Octant)

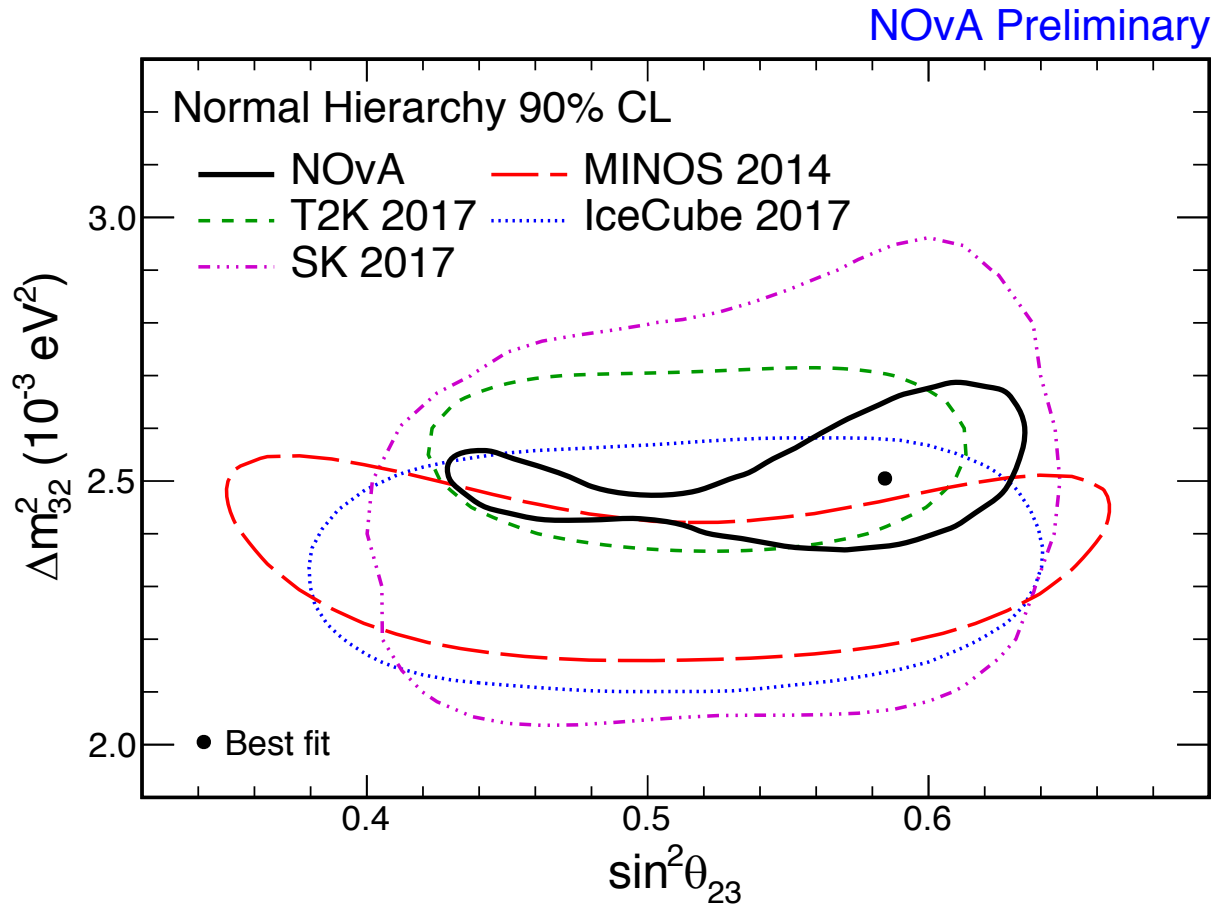
$$\delta_{CP} = 0.17\pi$$

# Appearance and Disappearance Fit Results

15

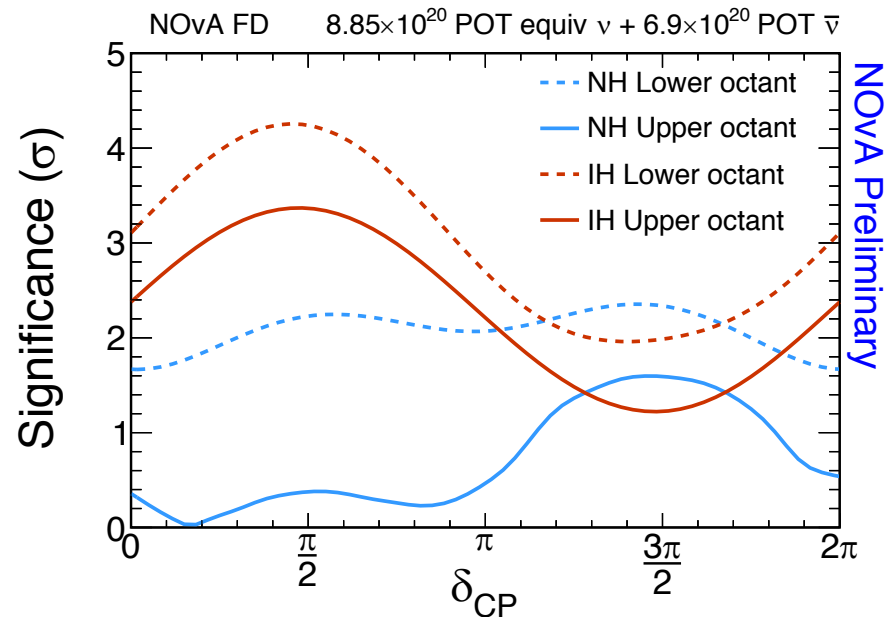
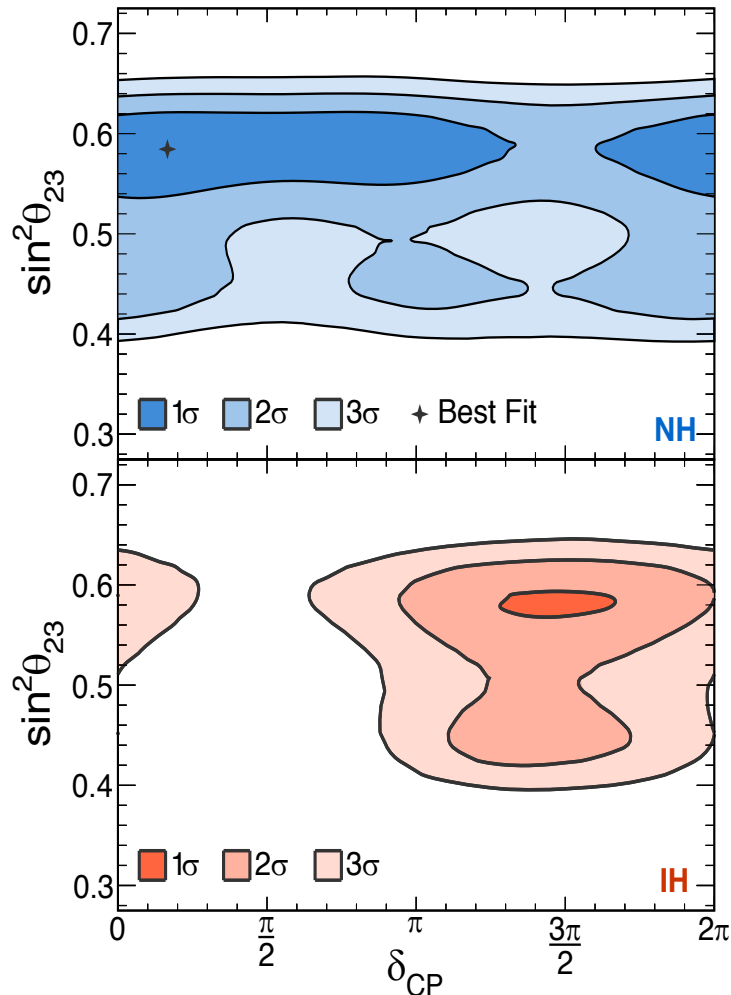


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# Appearance and Disappearance Fit Results

NOvA Preliminary



NOvA Preliminary

Normal Hierarchy preferred at 1.8 sigma  
 Exclude  $\pi/2$  in the IH at  $> 3$  sigma

Significances determined using Feldman-Cousins approach.  
 Leveraged high performance computing at NERSC



# Publications and PhDs

17



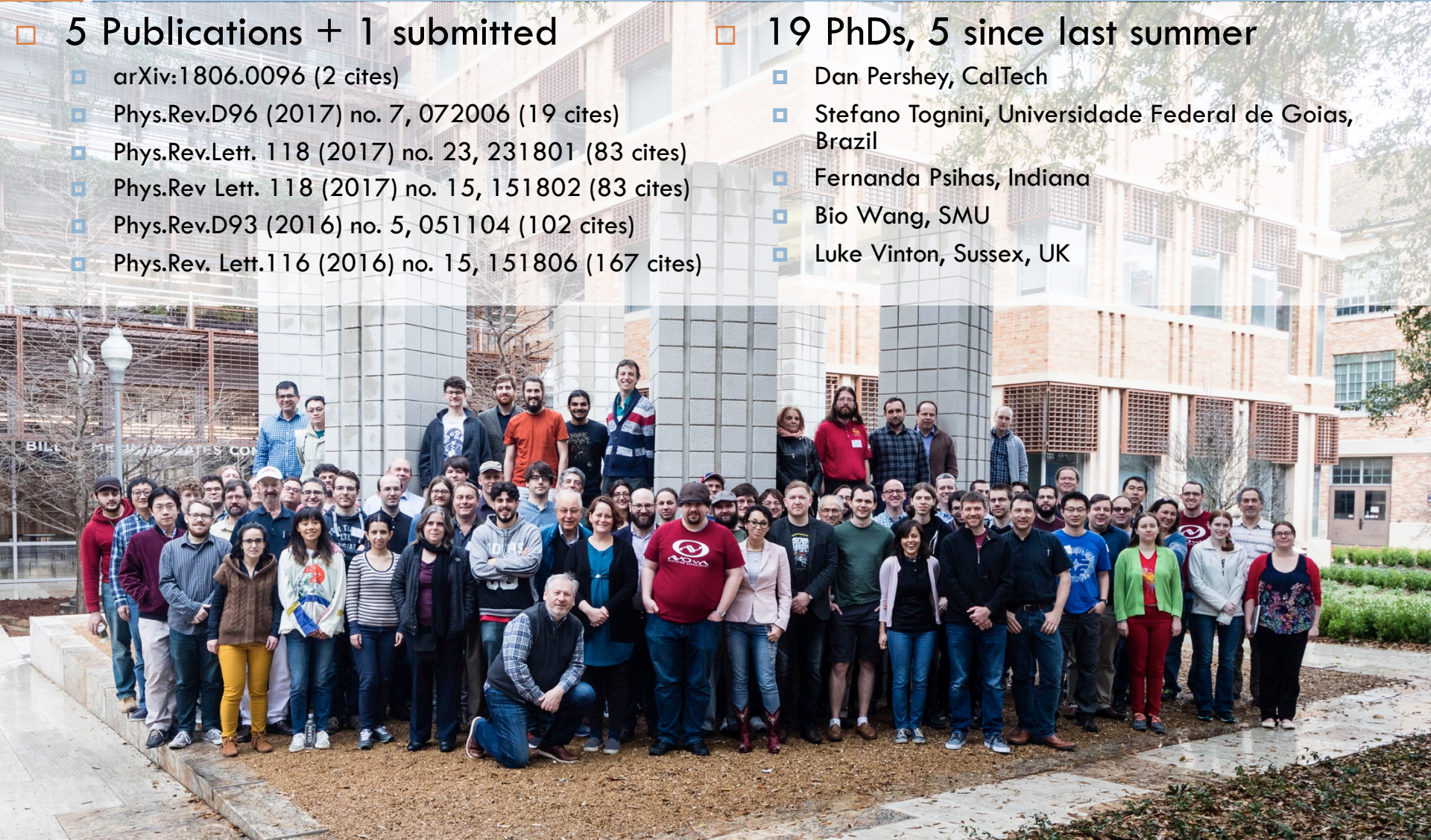
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## 5 Publications + 1 submitted

- arXiv:1806.0096 (2 cites)
- Phys.Rev.D96 (2017) no. 7, 072006 (19 cites)
- Phys.Rev.Lett. 118 (2017) no. 23, 231801 (83 cites)
- Phys.Rev Lett. 118 (2017) no. 15, 151802 (83 cites)
- Phys.Rev.D93 (2016) no. 5, 051104 (102 cites)
- Phys.Rev. Lett.116 (2016) no. 15, 151806 (167 cites)

## 19 PhDs, 5 since last summer

- Dan Pershey, CalTech
- Stefano Tognini, Universidade Federal de Goias, Brazil
- Fernanda Psihas, Indiana
- Bio Wang, SMU
- Luke Vinton, Sussex, UK



# Future Oscillation Analysis Strategy

18



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- Program planning advises us to expect to run to 2024
- Run plan:
  - Antineutrinos until  $1.2 \times 10^{20}$  POT (fulfills MINERvA request)
  - Approx. 50%-50% neutrino-antineutrino beyond
- Analysis plans
  - 2019: top up on antineutrinos
  - 2020: with  $\sim$ double neutrino exposure
  - Beyond: Increasing exposure, Joint analysis with T2K
- Analysis improvements in the works
  - More sophisticated cross section tuning
  - Better understanding of neutron response/simulation
  - More sophisticated ND decomposition for  $\nu_{ue}$
  - More sophisticated treatment of wrong sign
  - Testbeam



so far, 3 joint T2K meetings, with another planned this fall. Gaining mutual understanding of cross section uncertainties, with an eye to developing treatment of correlated systematics in joint fit

# Testbeam

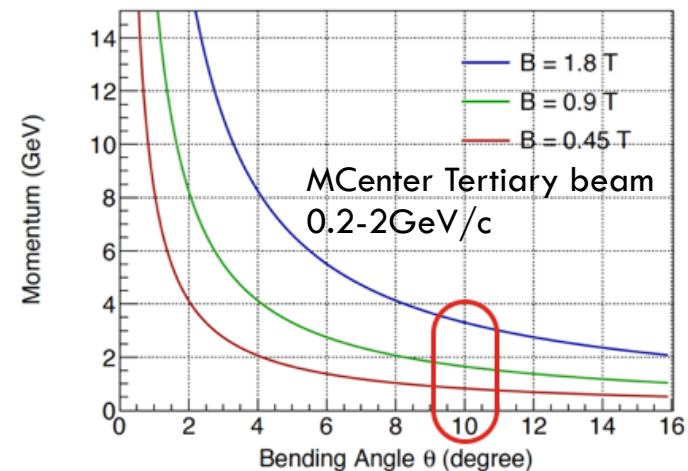
19



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- Testbeam program addresses dominant NOvA systematic errors
  - ▣ energy scale for muons, hadrons, EM showers
  - ▣ relative calibration
  - ▣ light/scintillator response model
- Provides real (tagged) data for sophisticated machine learning particle ID algorithms
- Installation begins this summer
  - ▣ Data collection starts end of the calendar year



# Proposed

## Accelerator Improvement Projects

20



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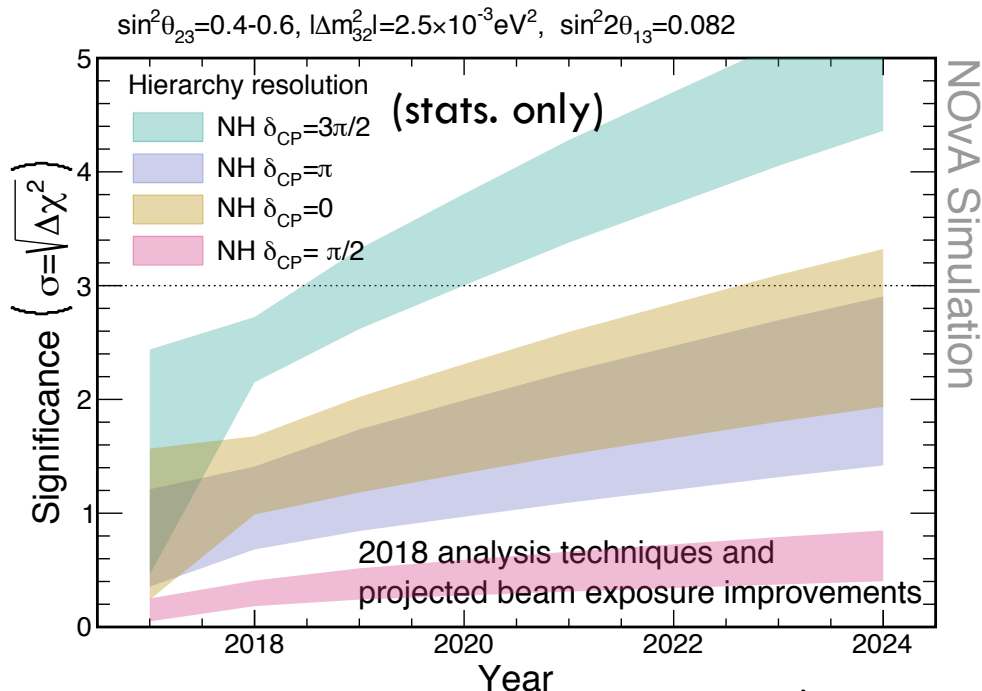
- 2 accelerator improvement projects have the potential to boost NuMI power beyond 700kW (up to 900kW-1MW)
  - Target System
    - New target design rated for 1MW
    - Improved horn stripline cooling
    - Radioactive water system upgrade
    - Target chase chiller and air handling upgrade
  - Intensity
    - Assorted projects to lower Booster losses
- Projects not only enable higher power, but improve reliability, mitigate risk
  - Lifetime extension to 2024
  - Support plan to run 40+ weeks/year (recent experience is 34-40)
- In our projections, we assume:
  - Power at 800 kW in FY19, 900kW in FY20-21, 1MW beyond
  - 40 weeks of running a year with uptime comparable to current running

# Future Sensitivity

21



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- Sensitivity dependent on true values in nature
- For favorable parameters consistent with results, we can achieve 3 sigma mass hierarchy sensitivity by 2020
  - ▣ 3 sigma sensitivity for 30-50% of delta CP range by 2024
  - ▣ >5 sigma in favorable cases by 2024, possible only with POT boosts from AIP work

□ Compare to other experiments\*:

- ▣ T2K reports Bayes factor NH/IH=7.9 (NH preferred at ~89%)
- ▣ SK prefers NH at at 80.6% to 96.7% (depending on  $\theta_{23}$ )
- ▣ Juno 3-4 sigma sensitivity 6 years after start in 2021 (depending on error on  $\Delta m_{\mu\mu}^2$ )
- ▣ KM3Net/Orca 5 sigma in 2024/2025 (depending on  $\theta_{23}$ )

\*Taken from Neutrino2018 talks by M. Wascko, Y. Hayuto, B. Wonsak, U. Katz

# Beyond 3 flavor oscillations



- Sterile Neutrinos
  - NC disappearance/anomalous muon neutrino (antineutrino) disappearance between ND and FD
  - SBL electron neutrino appearance in ND
  - Tau appearance in ND
- Neutrino interaction physics
  - Muon and electron neutrino charged current inclusive cross section
  - COH  $\text{Pi}^0$
  - CC  $\text{Pi}^0$  production
  - charged pion production
- Non-beam physics
  - Monopoles
  - Multi-muon cosmic ray events
  - Upward Muons
  - Supernova watch

# Summary

- First antineutrino results released this summer
  - ▣  $>4$  sigma evidence of electron-antineutrino appearance
  - ▣ 1.8 sigma preference for Normal Hierarchy
  - ▣ 1.8 sigma preference for non-maximal mixing
- Continue to take data in antineutrino mode until  $1.2 \times 10^{20}$  POT, then switch back to neutrinos
- Can achieve 3 sigma mass hierarchy sensitivity for favorable parameters by 2020
- With beam improvements can reach 3 sigma sensitivity to 30-50% of delta CP range by 2024,  $>5$  sigma for favorable parameters

24

# Backup

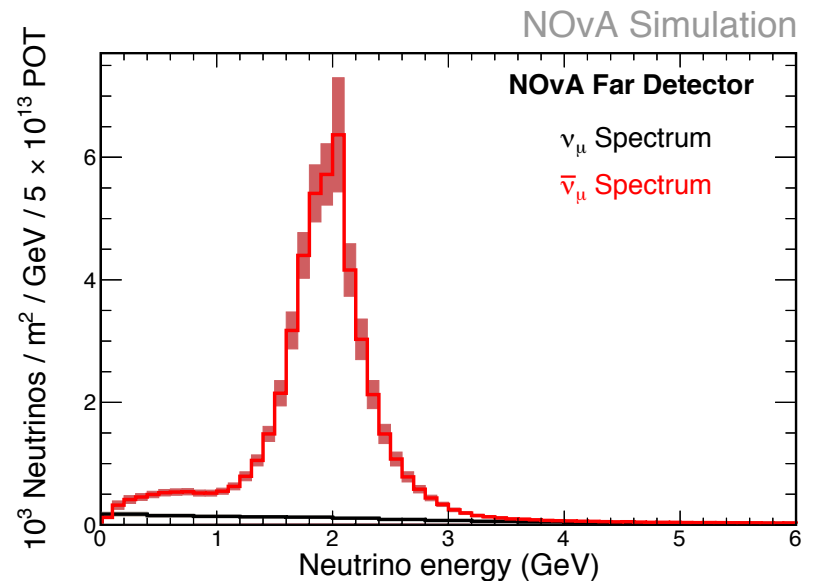
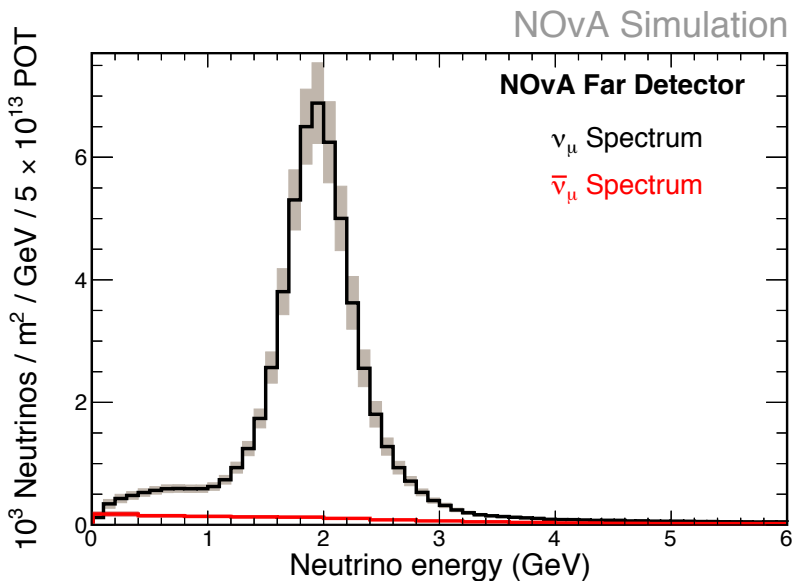
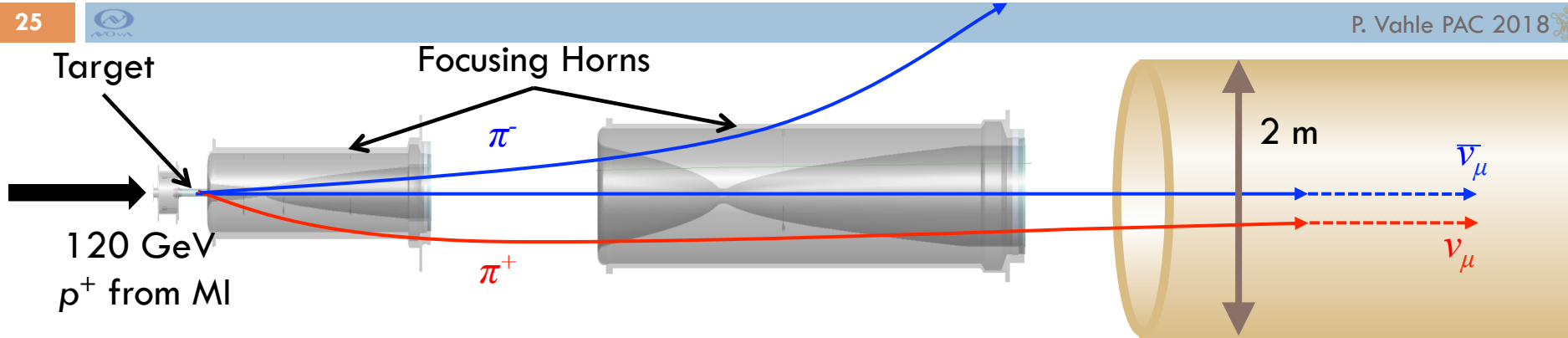


# Making an off-axis neutrino beam

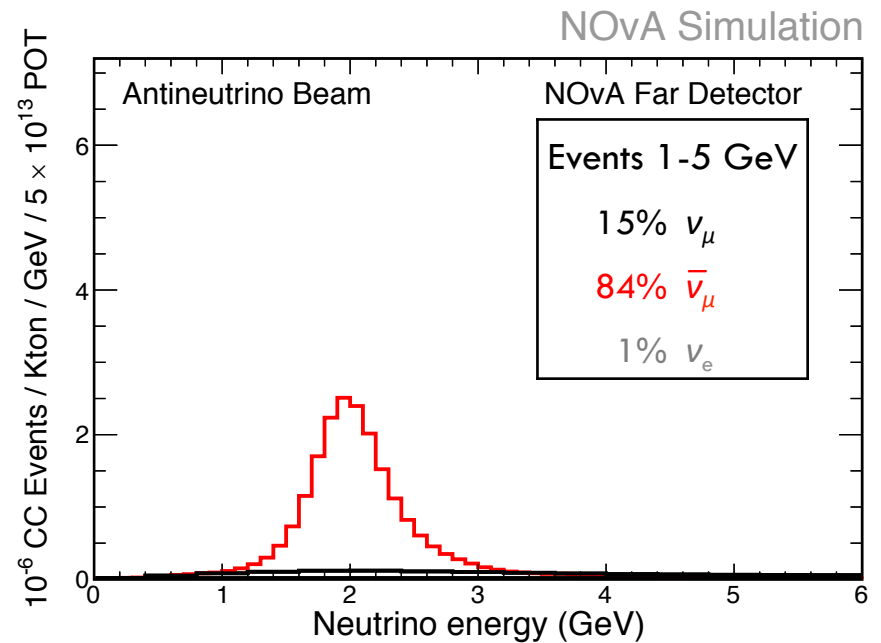
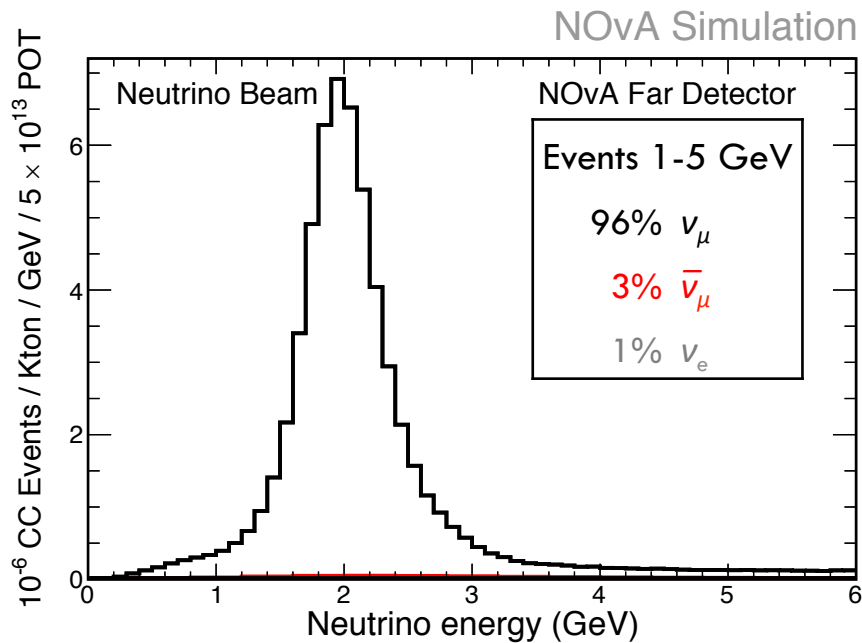
25



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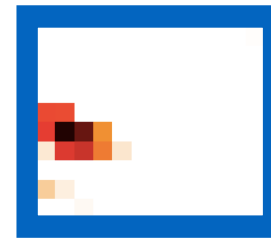
# Beam Spectra



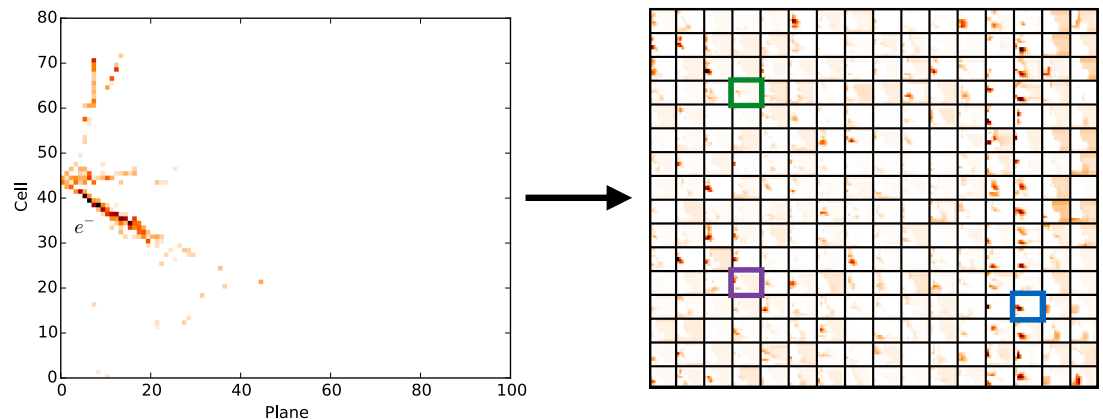
# Improved Event Selection

- This analysis features a new event selection technique based on ideas from computer vision and deep learning

- Calibrated hit maps are inputs to Convolutional Visual Network (CVN)



- Series of image processing transformations applied to extract abstract features
- Extracted features used as inputs to a conventional neural network to classify the event

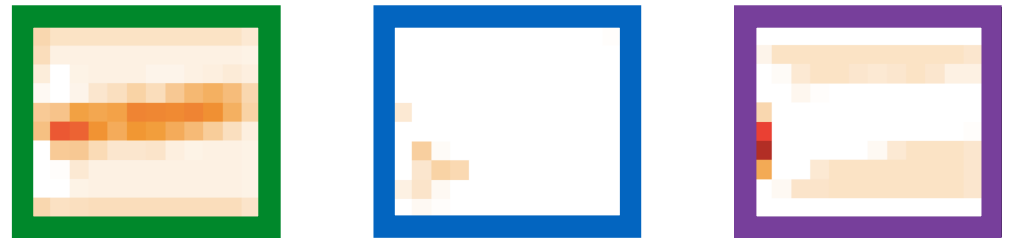


A. Aurisano et al., arXiv:1604.01444  
Posters P1.028 by A. Radovic, P1.032 by  
F. Psihas and A. Himmel for more detail

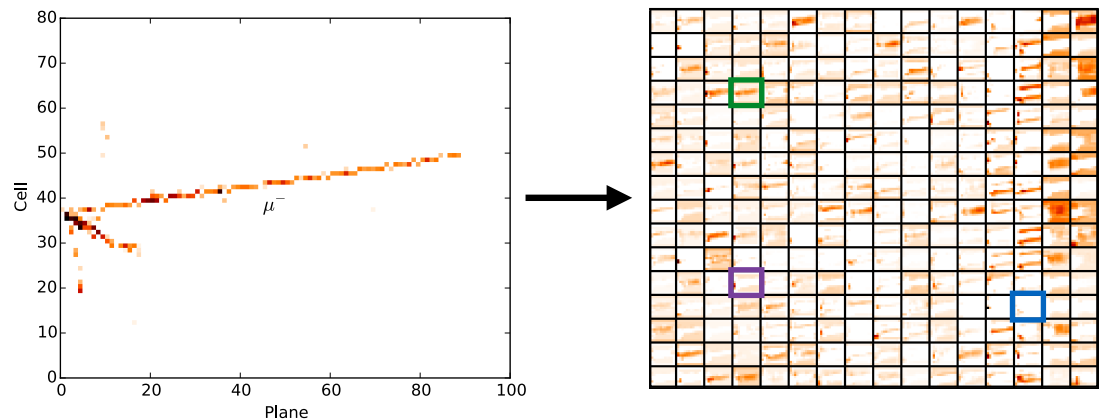
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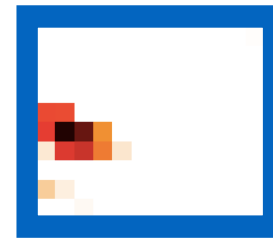


A. Aurisano et al., arXiv:1604.01444  
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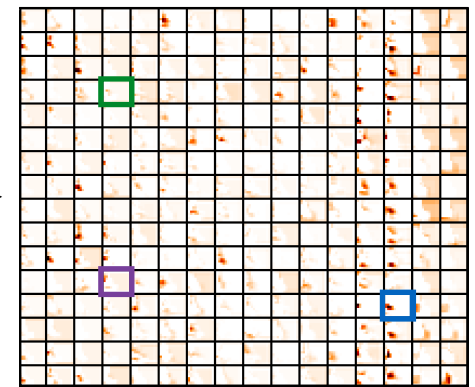
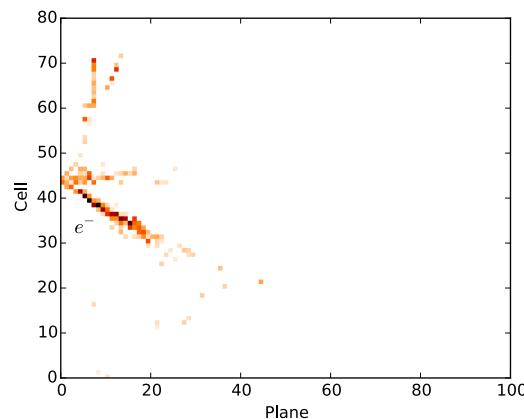
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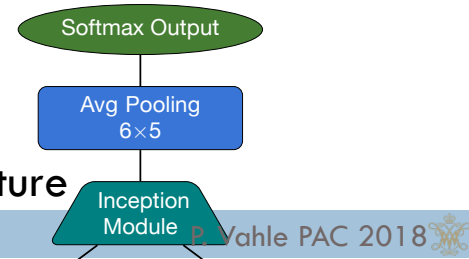


Improvement in sensitivity from CVN  
equivalent to 30% more exposure

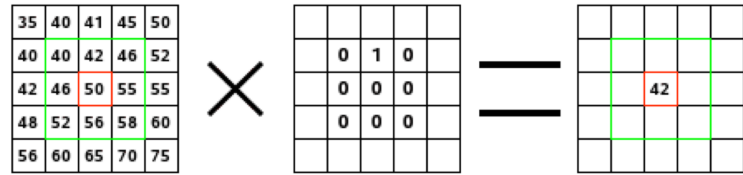
# CVN Architecture



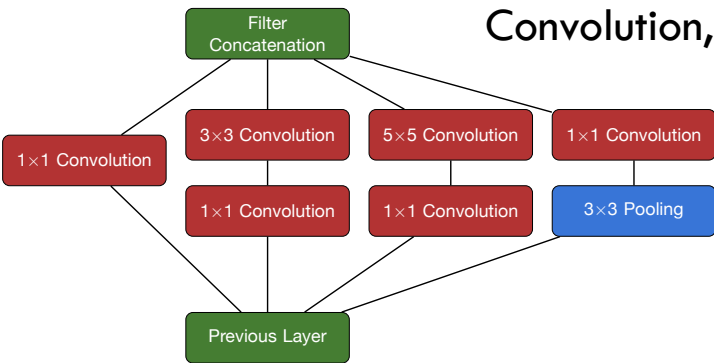
## CVN Architecture



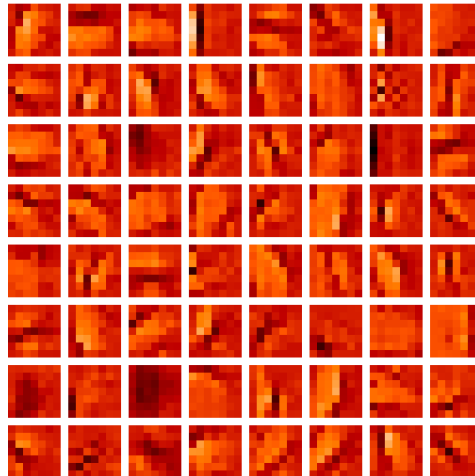
GoogLeNet  
Inception Module  
C. Szegedy et al.,  
arXiv:1409.4842



Example image processing transformation  
Convolution, or kernel map

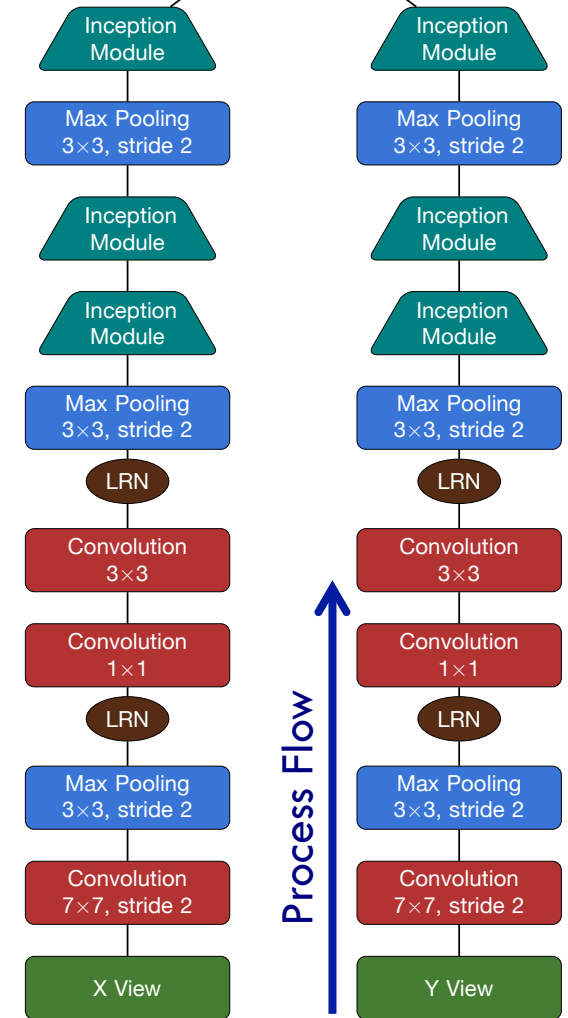


Example Convolutional Filter Layer



Network implemented and trained  
in the Caffe Framework  
(Y. Jia et al., arXiv:1408.5093)

Trained over 4.7M simulated events,  
Trained on FNAL GPU farm



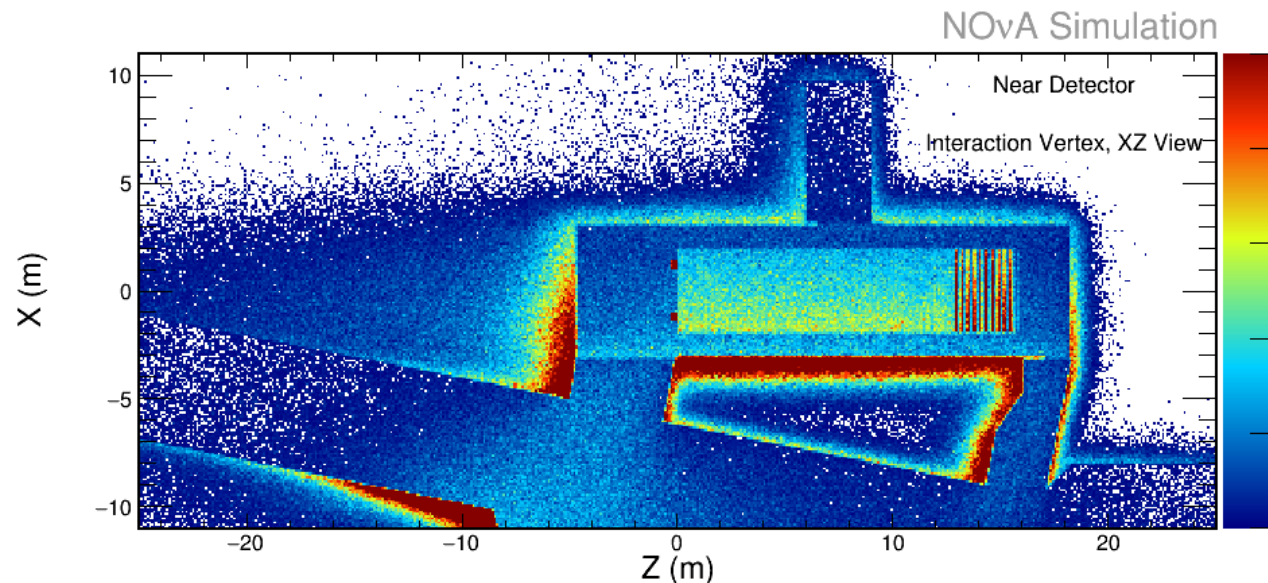
# Simulation

31



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- Beam line production, propagation and neutrino flux: FLUKA/Flugg
- Cosmic Ray flux: CRY
- Neutrino interaction and FSI: GENIE
- Detector: Simulation: Geant4
- Detector response: Custom simulation Routines



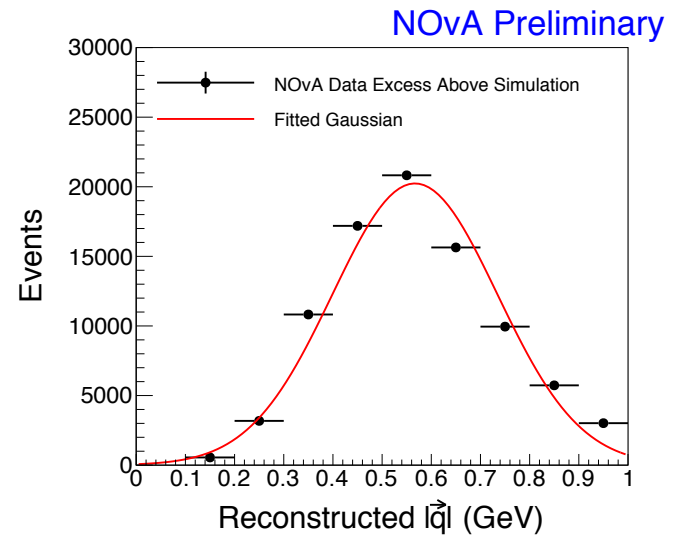
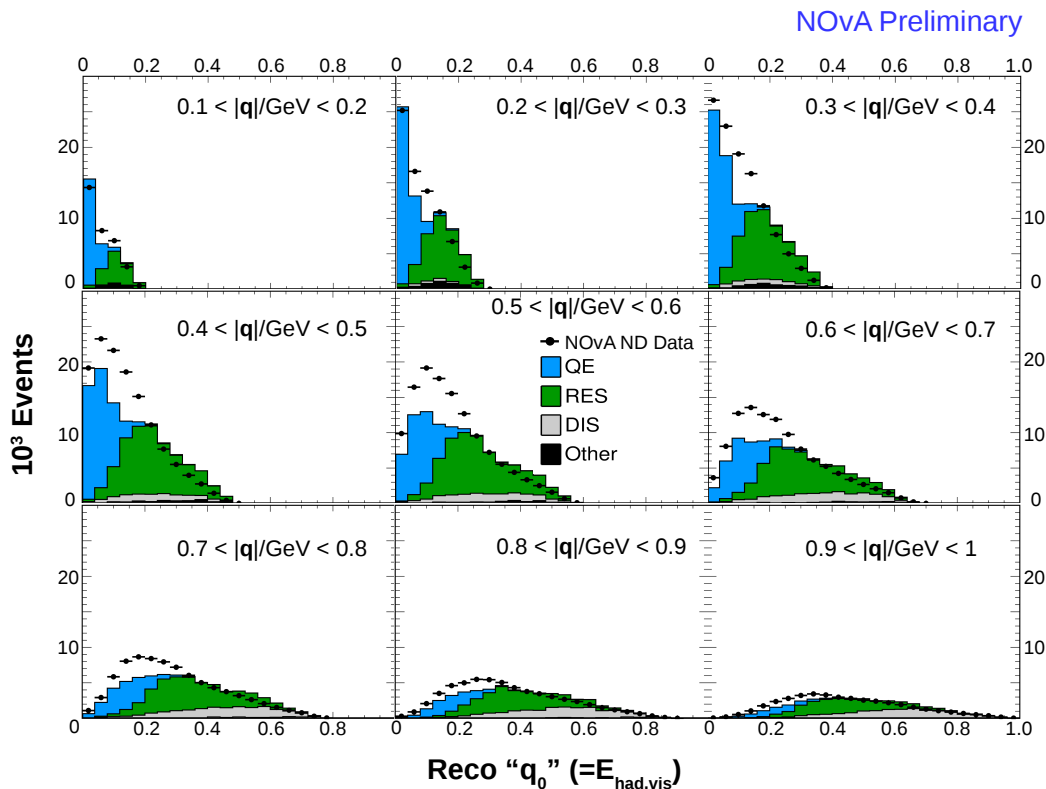
# Scattering in a Nuclear Environment

32



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- Near detector hadronic energy distribution suggests unsimulated process between quasi-elastic and delta production



Similar conclusions from MINERvA data reported in P.A. Rodrigues et al., PRL 116 (2016) 071802



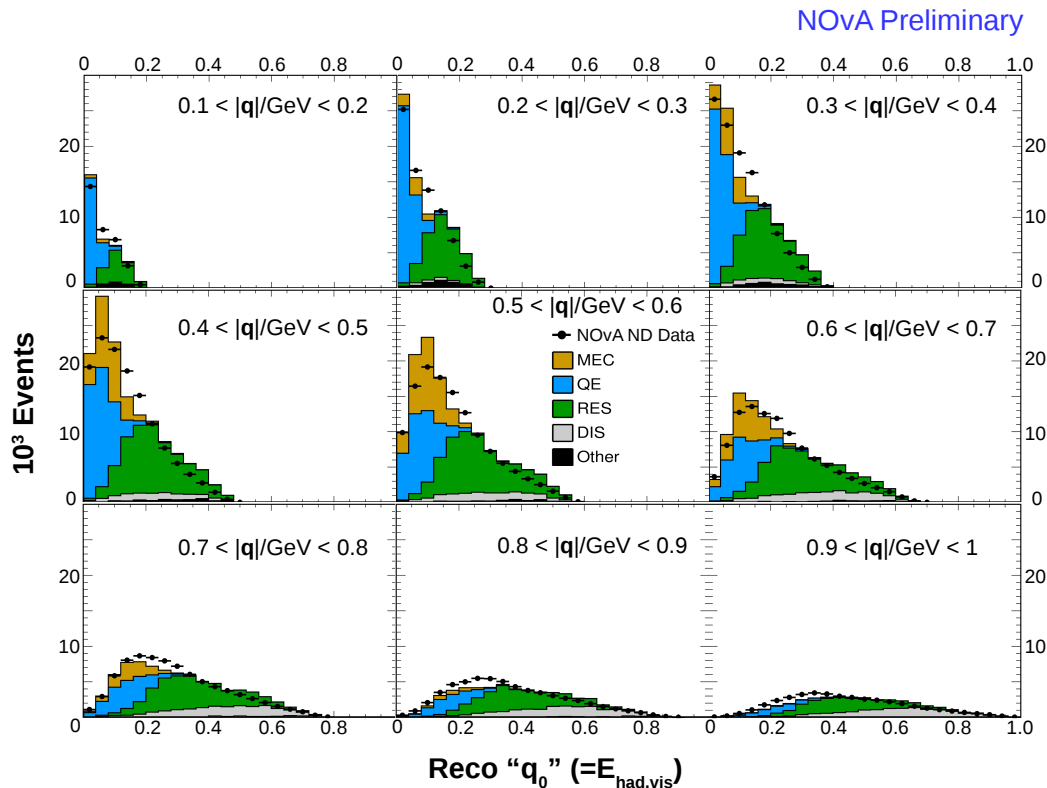
# Scattering in a Nuclear Environment

33



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- Enable GENIE empirical Meson Exchange Current Model
- Reweight model to match NOvA excess as a function of 3-momentum transfer



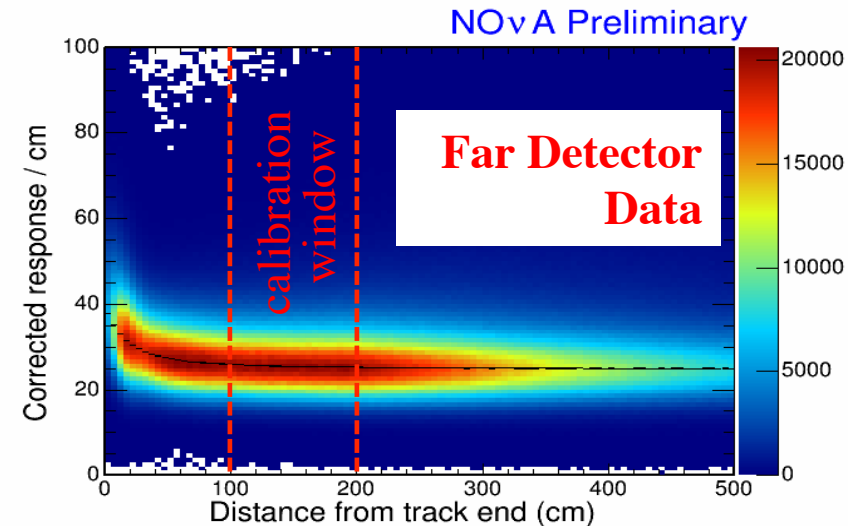
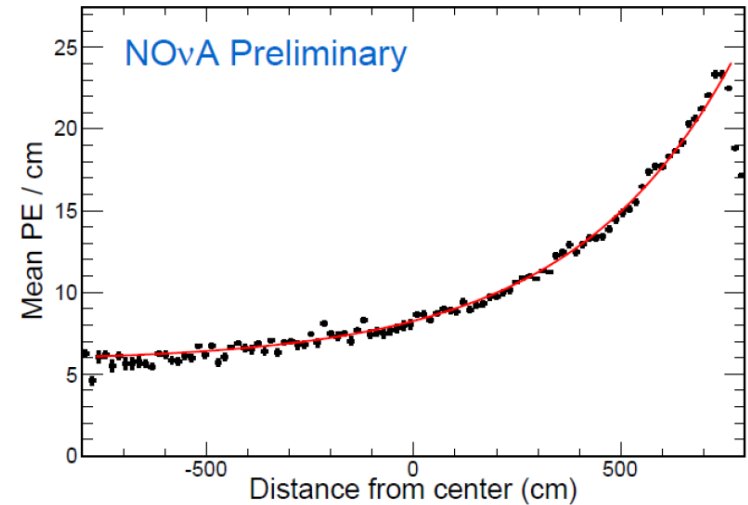
- Eliminate falloff of  $x_{sec}$  with energy
- Fix final state nucleon composition
- Match MEC  $q_0$  distribution to quasi-elastic distribution
- scale normalization in bins of 3-momentum transfer to match data/mc disagreement
- Include 50% systematic uncertainty on size of MEC component
- Additionally, reduce single non-resonant pion production by 50%

(P.A. Rodrigues et al, arXiv:1601.01888.)

# Calibration

- Calibration achieved using cosmic rays
- Light levels drop by a factor of 8 across a FD cell
- Stopping muons provide a standard candle

FD cosmic data - plane 84 (horizontal), cell 12



# Energy Scale

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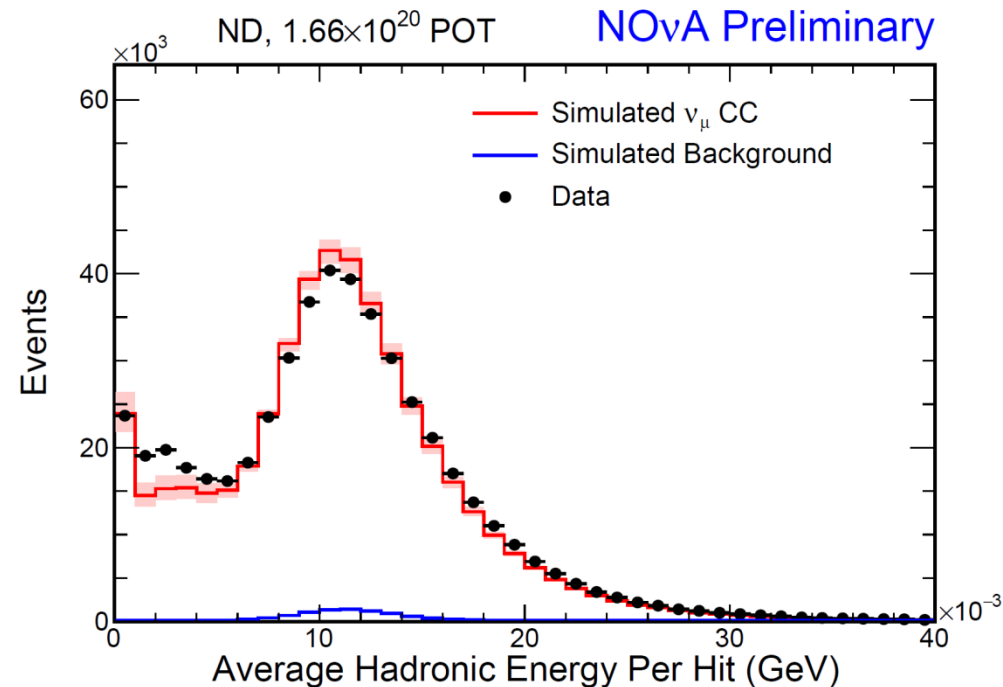
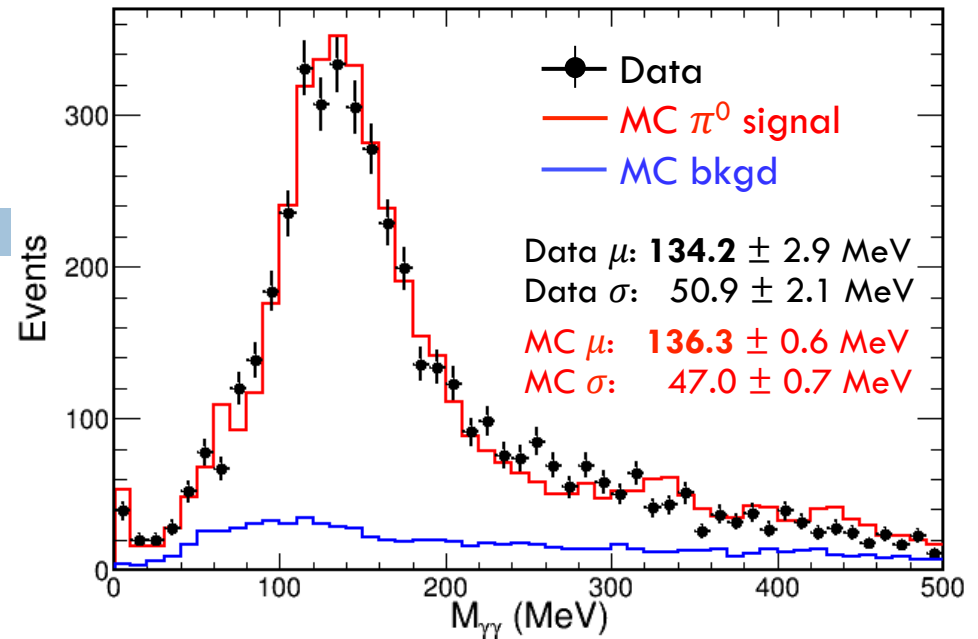
## Near Detector

- ▣ cosmic  $\mu$  dE/dx [ $\sim$ vertical]
- ▣ beam  $\mu$  dE/dx [ $\sim$ horizontal]
- ▣ Michel  $e^-$  spectrum
- ▣  $\pi^0$  mass
- ▣ hadronic shower  $E$ -per-hit

## Far Detector

- ▣ cosmic  $\mu$  dE/dx [ $\sim$ vertical]
- ▣ beam  $\mu$  dE/dx [ $\sim$ horizontal]
- ▣ Michel  $e^-$  spectrum

▣ All agree to 5%



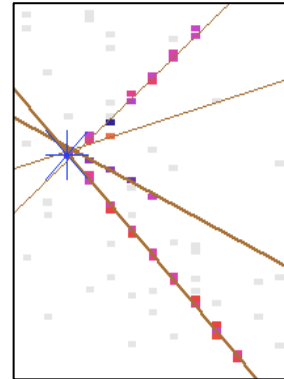
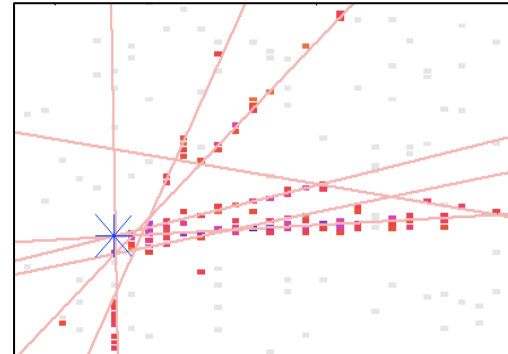
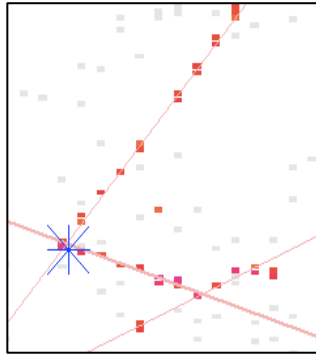
# Reconstruction

36

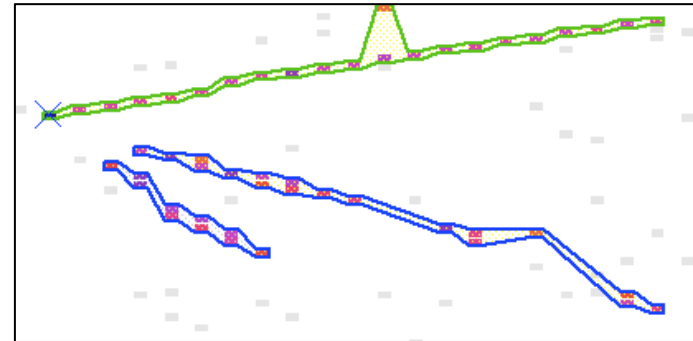
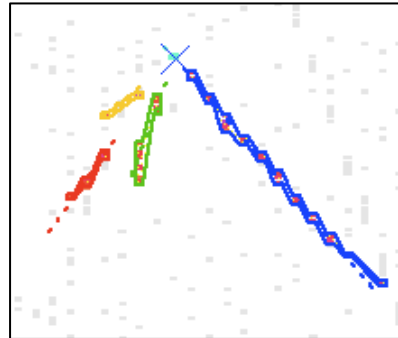


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**Vertexing:** Find lines of energy depositions w/ Hough transform CC events: 11 cm resolution



**Clustering:** Find clusters in angular space around vertex. Merge views via topology and prong  $dE/dx$



**Tracking:** Trace particle trajectories with **Kalman filter** tracker.

Also, **cosmic ray tracker:** lightweight, fast, and for large calibration samples, online monitoring.



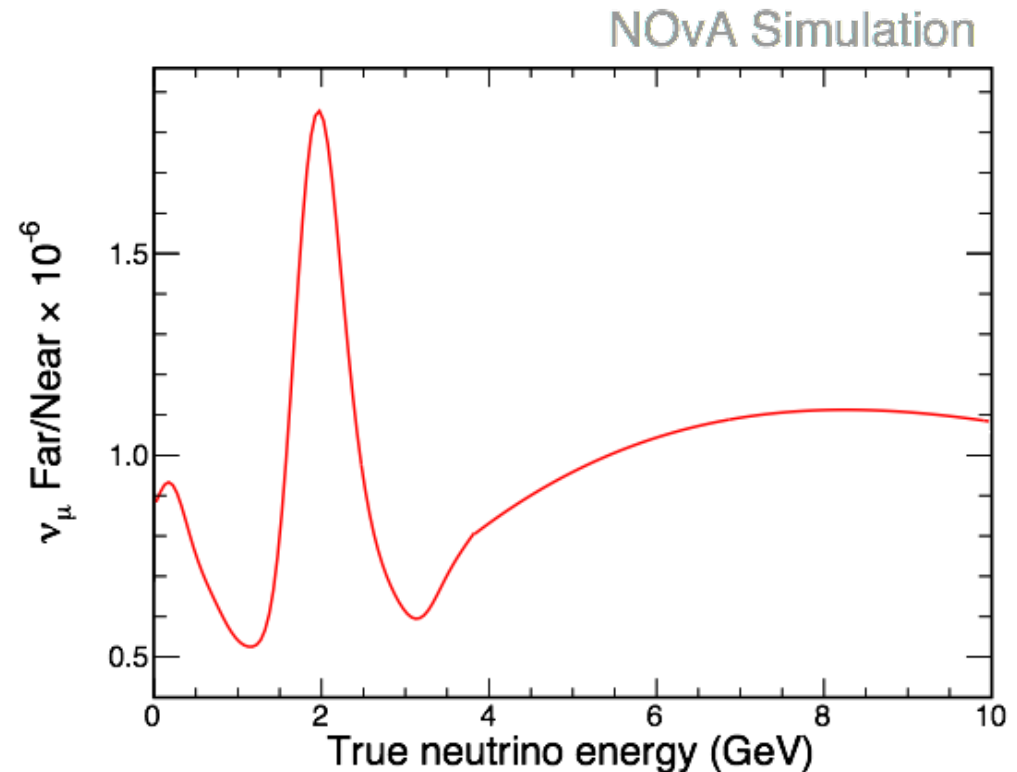
# Extrapolation

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- All NOvA analyses use ND data to predict the FD spectrum
- Muon Neutrino analyses: unfold reco energy, use true F/N ratio for FD prediction of track events
- Electron Neutrino signal: use Muon-neutrino FD prediction
- NC and Electron Neutrino background analyses: Far to Near spectrum in reconstructed energy ratio for FD prediction of shower events

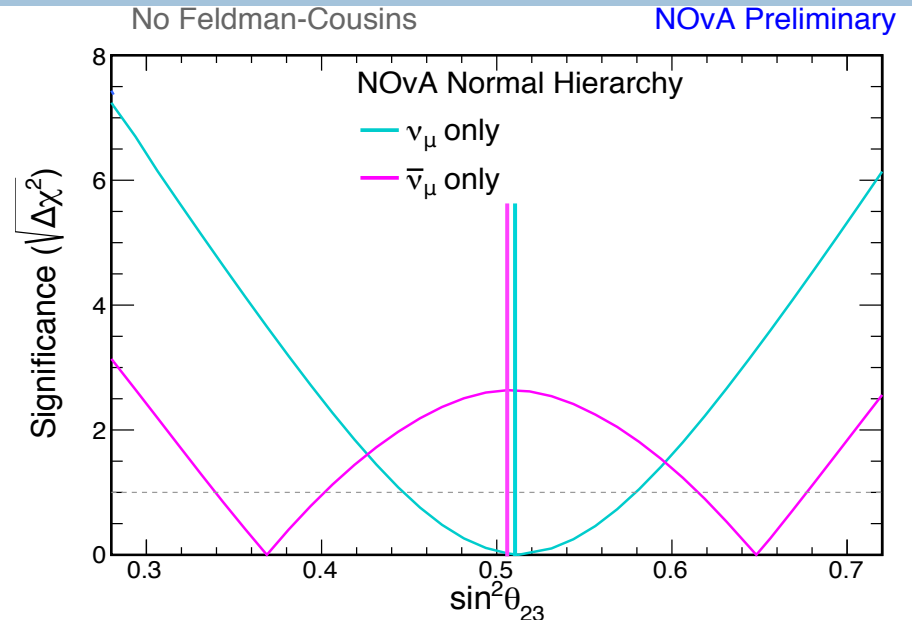
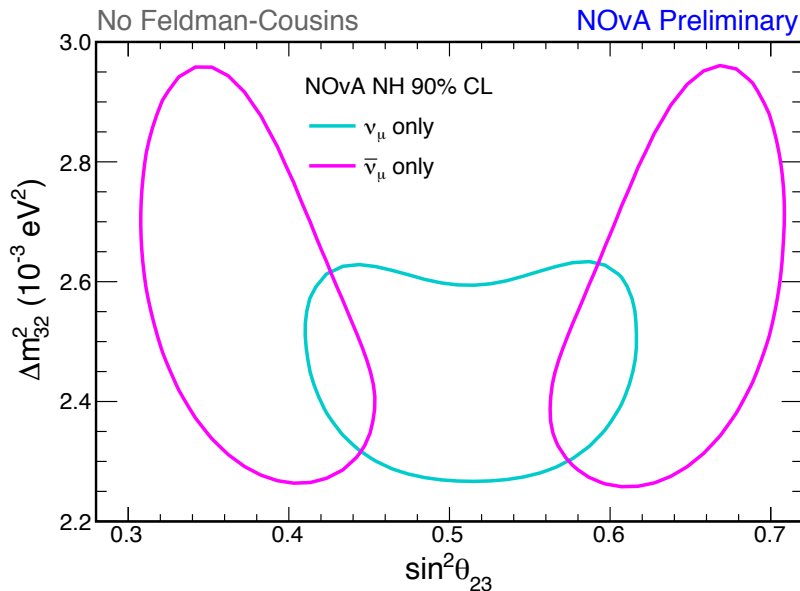


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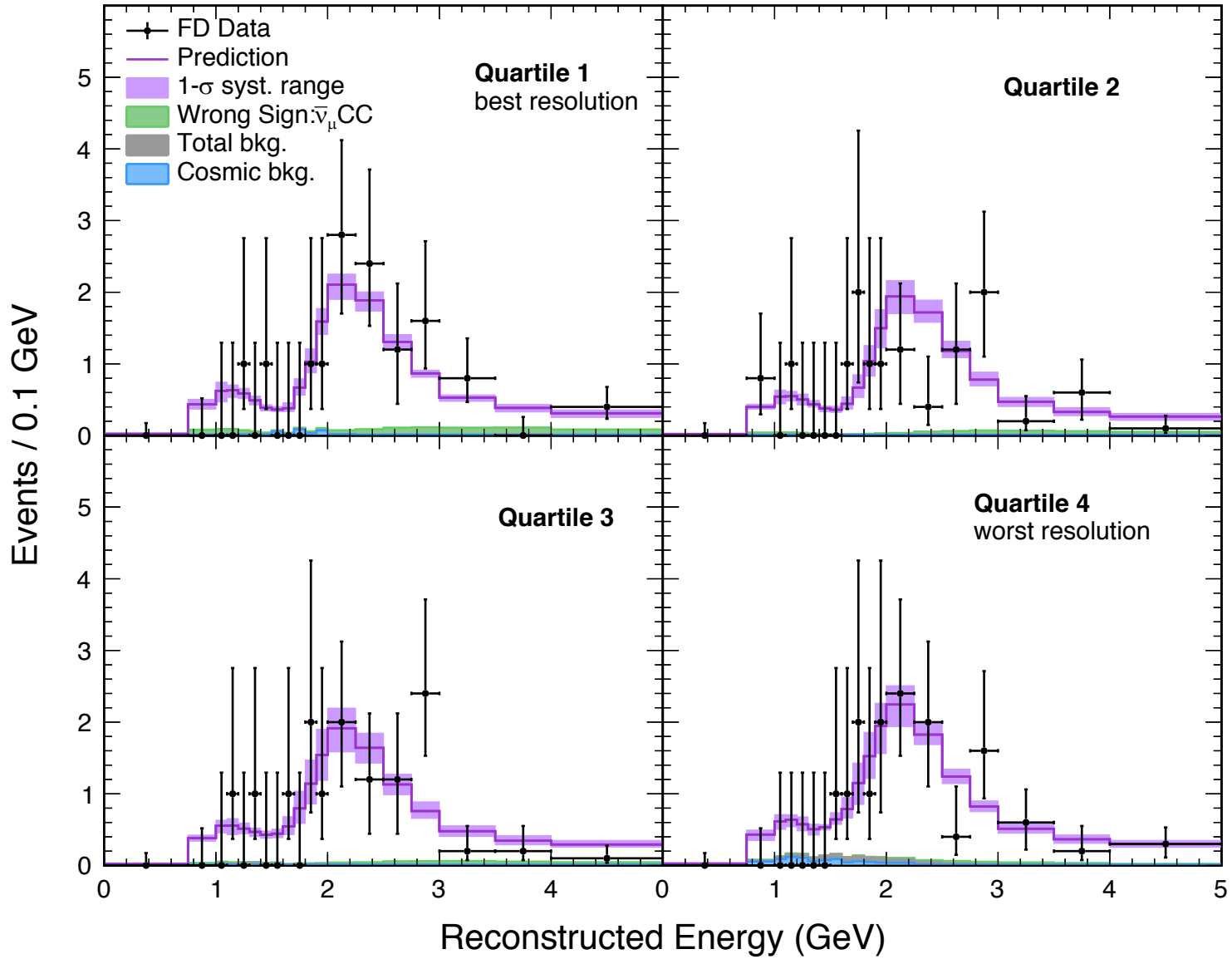
# Backup—Disappearance Results

# Disappearance

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- If fit separately, the  $\bar{\nu}_\mu$  data prefers non-maximal while  $\nu_\mu$  prefers maximal.
  - Consistent with joint oscillation parameters to  $>4\%$ .
- Matter effects introduce a small asymmetry in the point of maximal disappearance.
- Gives a  $1.3\sigma$  preference for the Upper Octant from *just* the  $\nu_\mu + \bar{\nu}_\mu$  fit in NH.
  - The asymmetry is flipped in the Inverted Hierarchy, so there is a similar preference for the lower octant there.

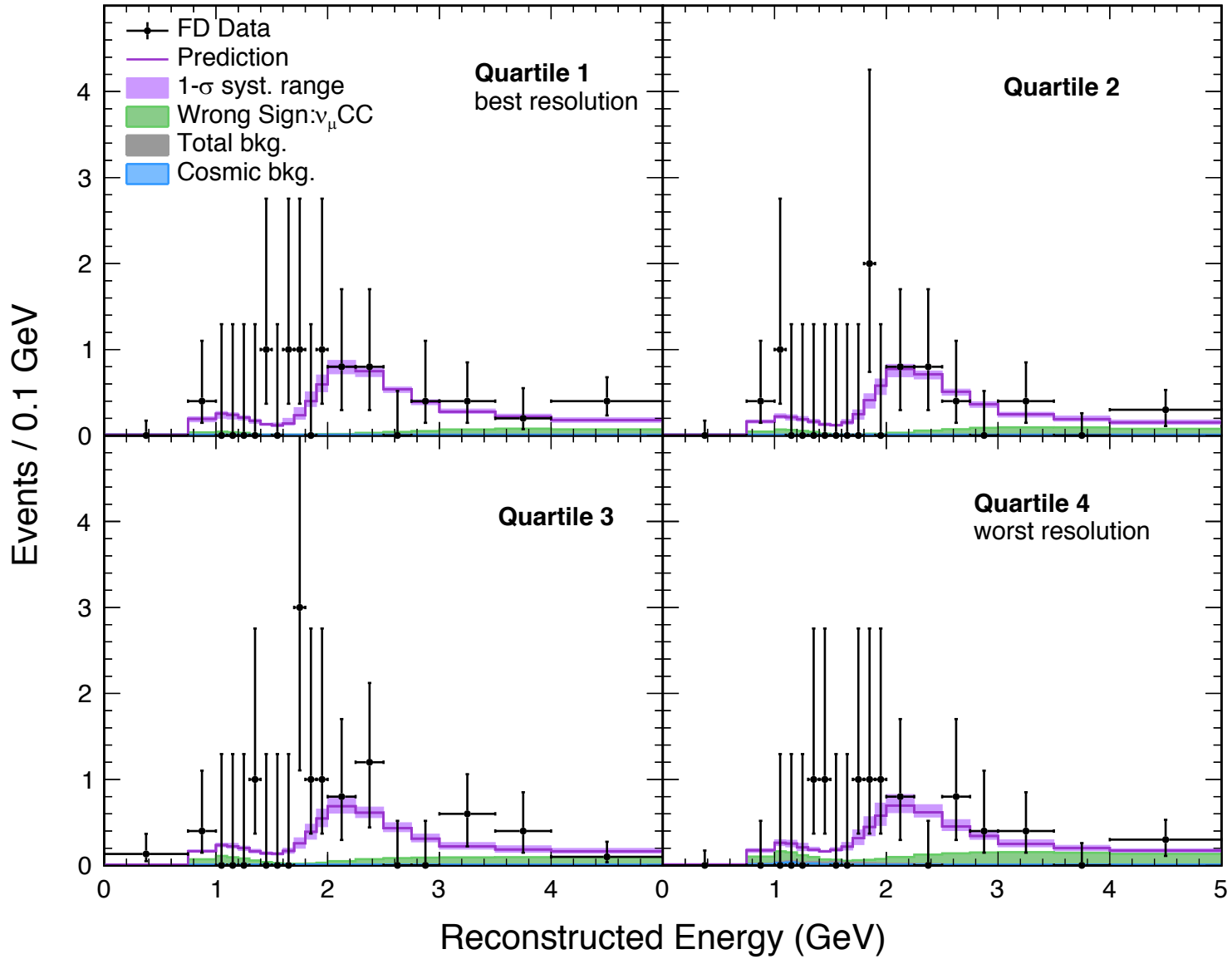




# Antineutrino beam

# NOvA Preliminary

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# Backup—NC Results

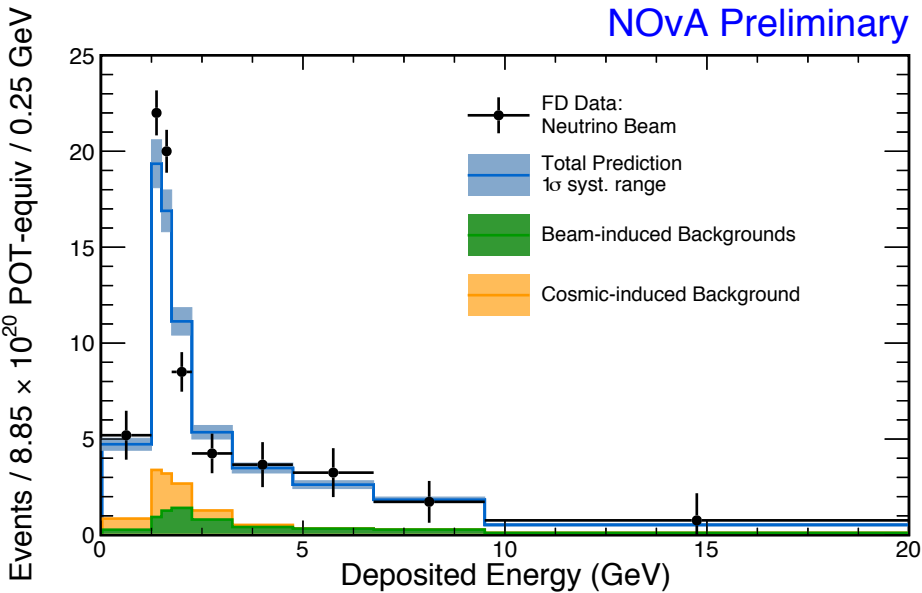
# NC Results

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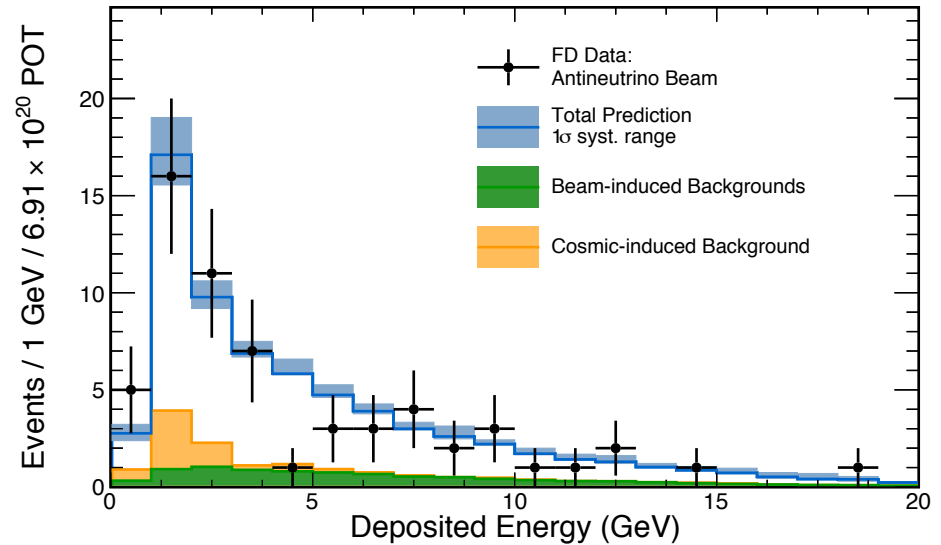
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NOvA Preliminary



Neutrino mode:  
expect  $188 \pm 13$  (syst.) with 38 bkg.  
Observe 201 events

NOvA Preliminary



Antineutrino mode:  
expect  $69 \pm 8$  (syst.) with 16 bkg.  
Observe 61 events

**No significant suppression of Neutral Current interactions in either neutrino or antineutrino mode**

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# Backup—Appearance Results

# Event Migration

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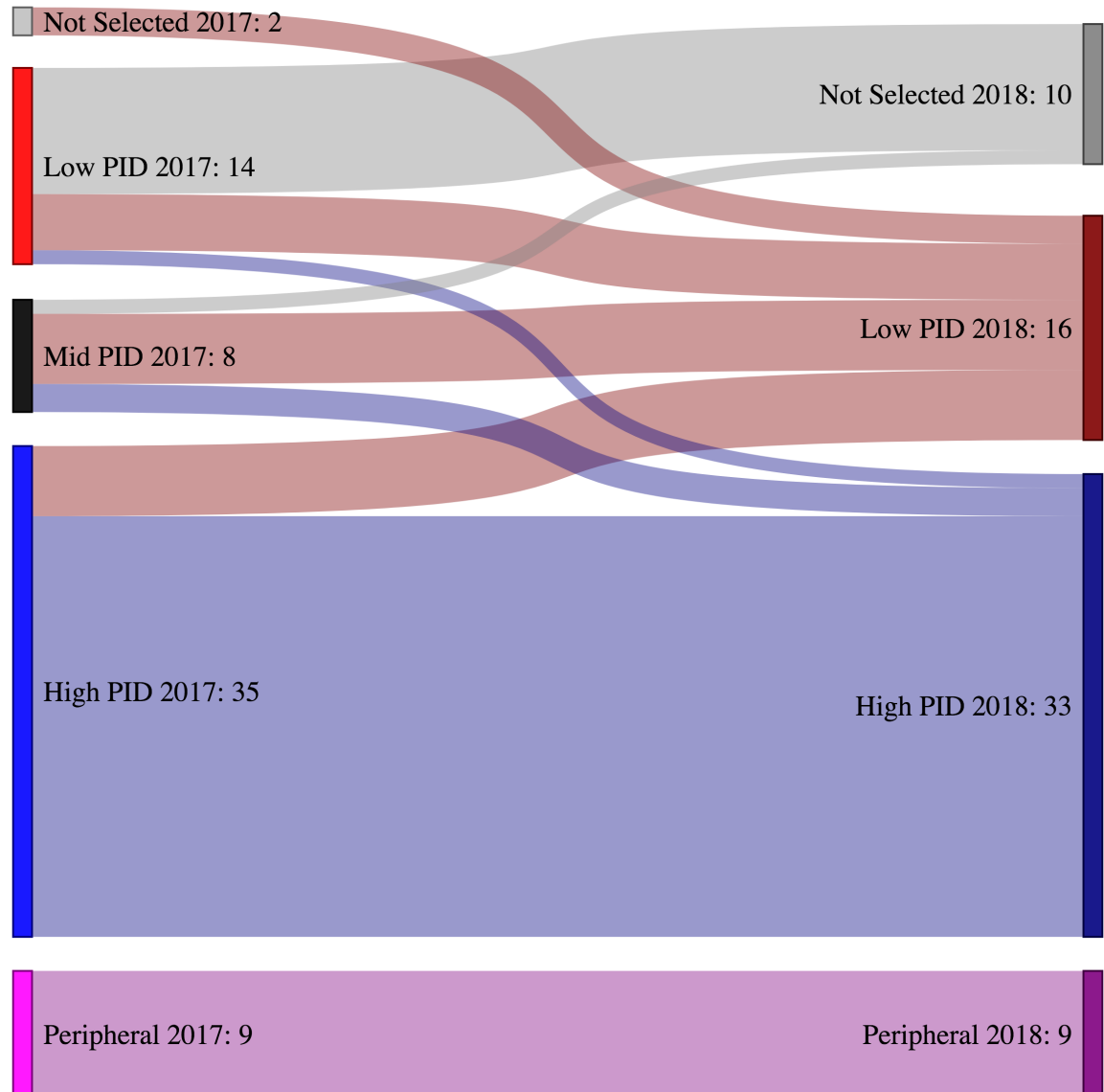


## Core:

- Expected to lose 7 and gain 3.5 vs. 2017.
- Actually lost 10 and gained 2.
- No events lost or gained in the high PID bin, but some did move from/to lower PID bins.

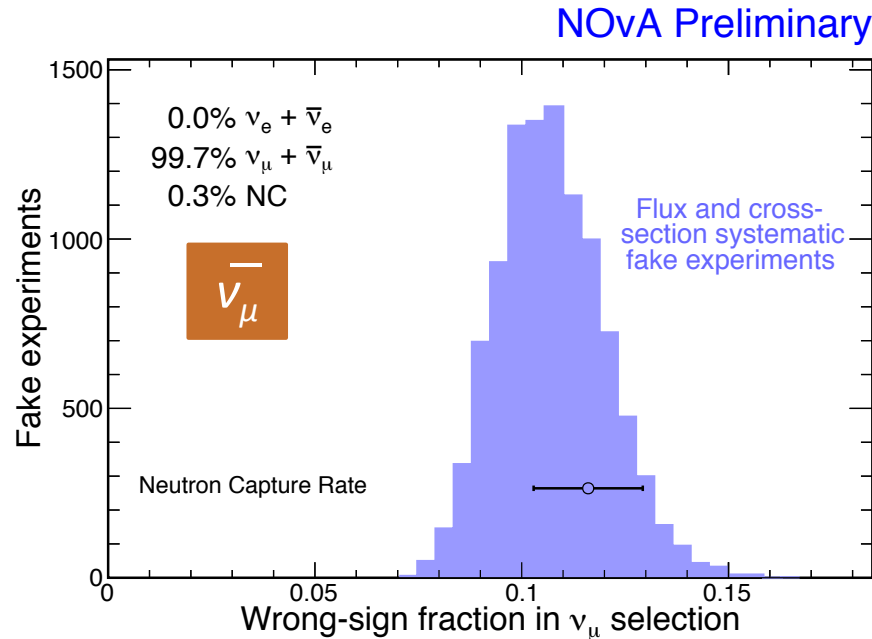
## Peripheral:

- Expected to gain and lose 2.5 events.
- Actually no events lost or gained.



# Wrong-sign Background

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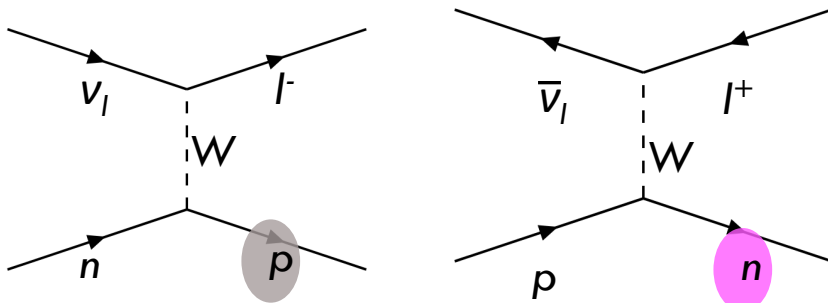


- The 11% wrong-sign fraction of the  $\bar{\nu}_\mu$  events is important since it becomes the WS background in the  $\bar{\nu}_e$  appearance analysis.
- $\sim 10\%$  systematic uncertainty on wrong-sign from flux and cross section
  - Does not include uncertainties from detector effects.
- Confirmed using data-driven cross-check of the wrong-sign contamination
  - 11% wrong-sign in the  $\bar{\nu}_\mu$  sample checked using neutron captures in the neutrino and antineutrino beams.

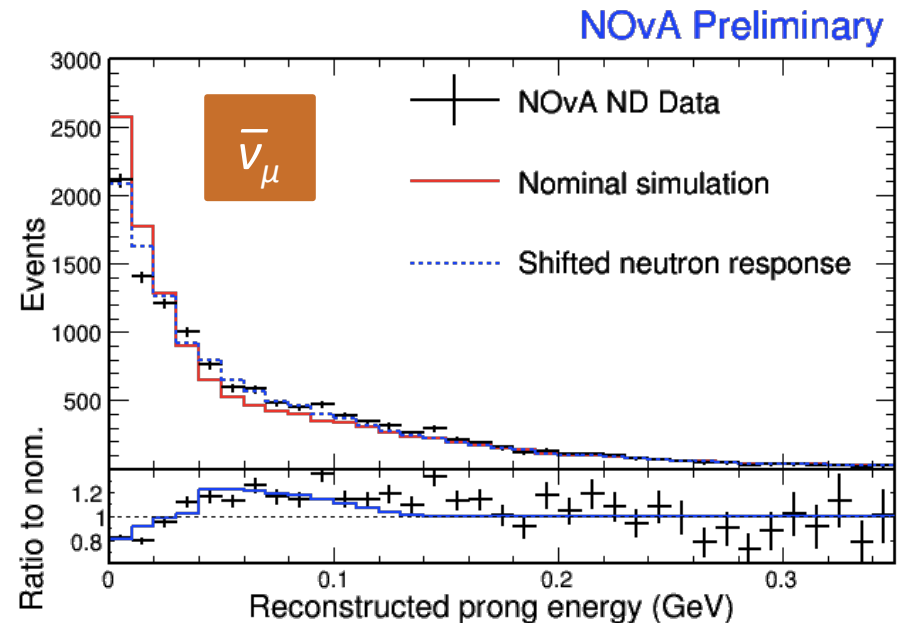
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# Backup—TestBeam

# New neutron response systematic



- $\bar{\nu}$ 's have neutrons where  $\nu$ 's have protons.
  - ▣ Often several hundred MeV of energy.
  - ▣ Modeling these fast neutrons is known to be challenging.
- See some discrepancies in an enriched sample of neutron-like prongs.
- New systematic introduced:
  - ▣ Scales the amount of deposited energy of some neutrons to cover the low-energy discrepancy.
- Shifts the mean  $\nu_\mu$  energy by 1% in the antineutrino beam and 0.5% in the neutrino beam.
  - ▣ Negligible impact was seen on selection efficiencies.





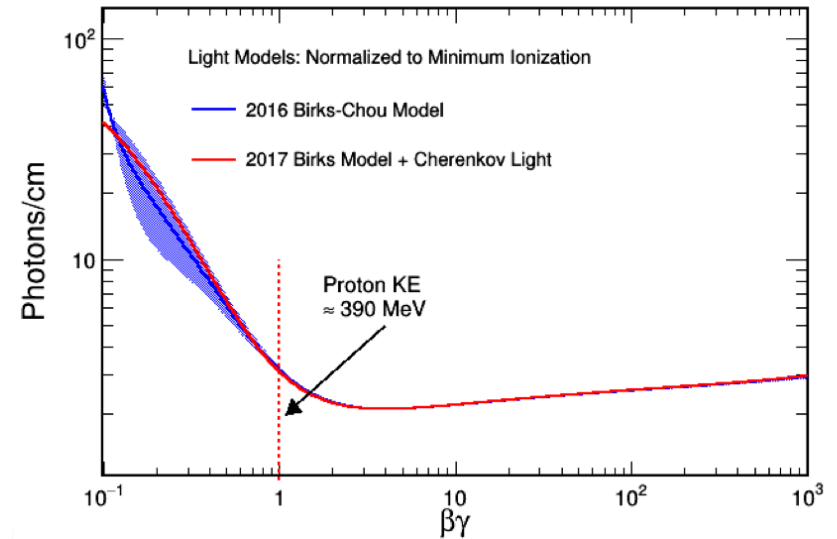
# Scintillator Model

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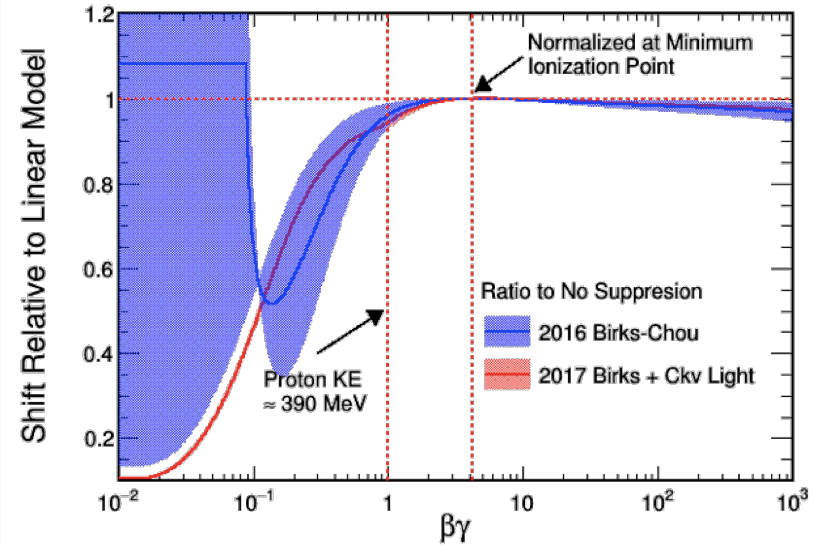


- Absorbed and re-emitted **Cherenkov light** is a small but important component of our scintillator response.
  - Particularly for low-energy protons in hadronic showers.
- Was one of our largest uncertainties, now reduced by an order of magnitude.
  - Previously accounted for with **second order terms in our scintillator model**.
  - Those terms were unusual, so we placed large systematics.
- Expected energy resolution for  $\nu_{\mu}$  CC events increased from 7% to 9%.

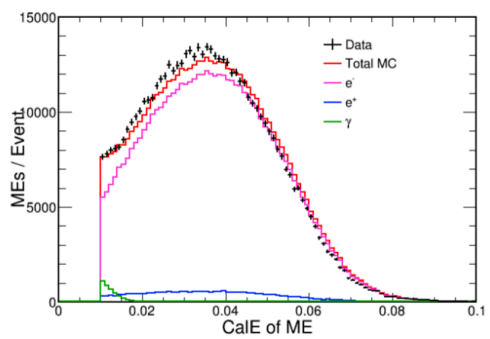
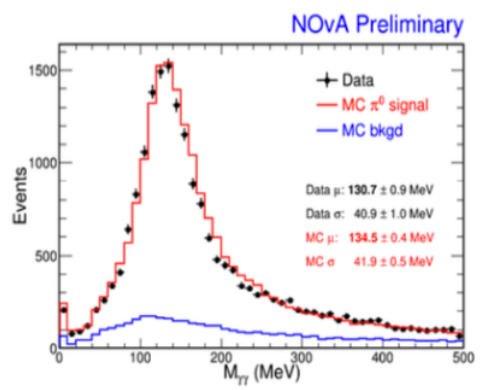
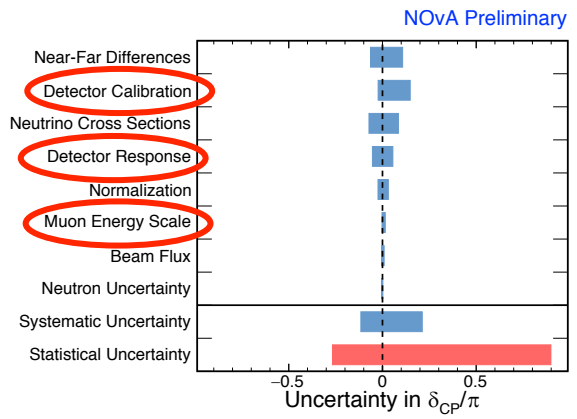
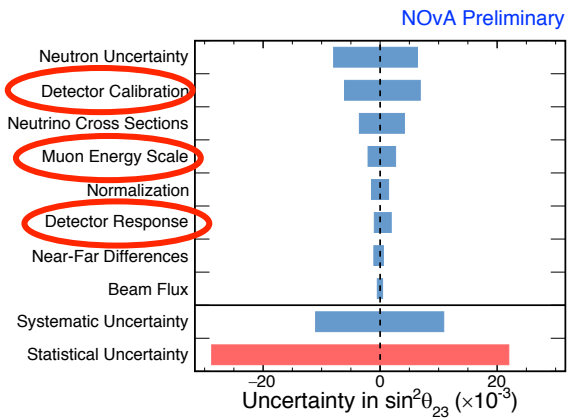
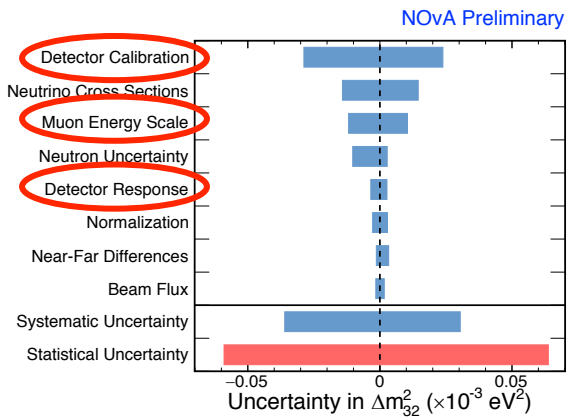
NOvA Simulation



NOvA Simulation



# Systematics and Testbeam Motivation



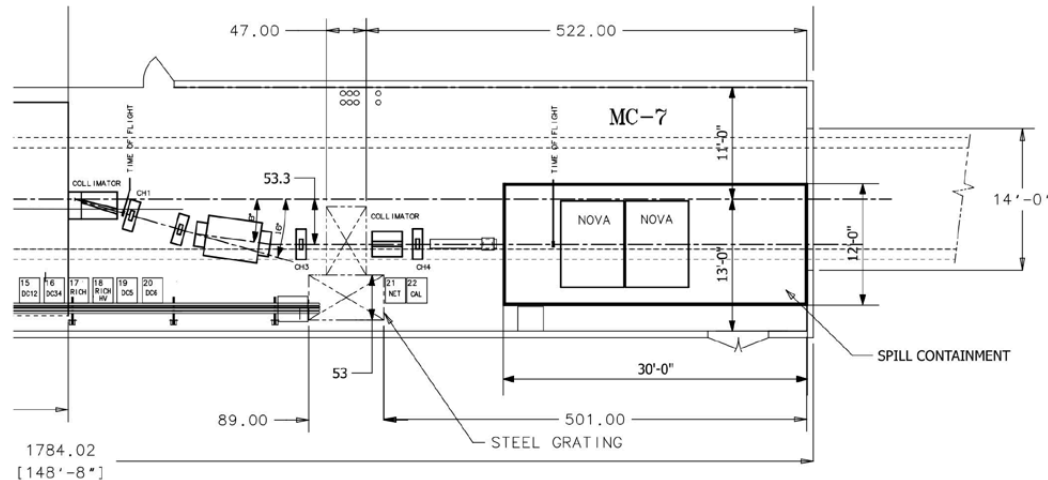
- Testbeam effort affords detailed understanding of detectors energy scale (muon, electromagnetic, and hadronic)
- Allows study of light model, scintillator quenching
- Provides real data for study of particle identification techniques

- Estimate of energy/calibration uncertainty comes from  $\sim 5\%$  difference in data vs. MC when comparing  $\pi^0$  mass and Michel e distributions
- Testbeam data decouples detector response from hadronic final state, neutrino cross section, flux uncertainties

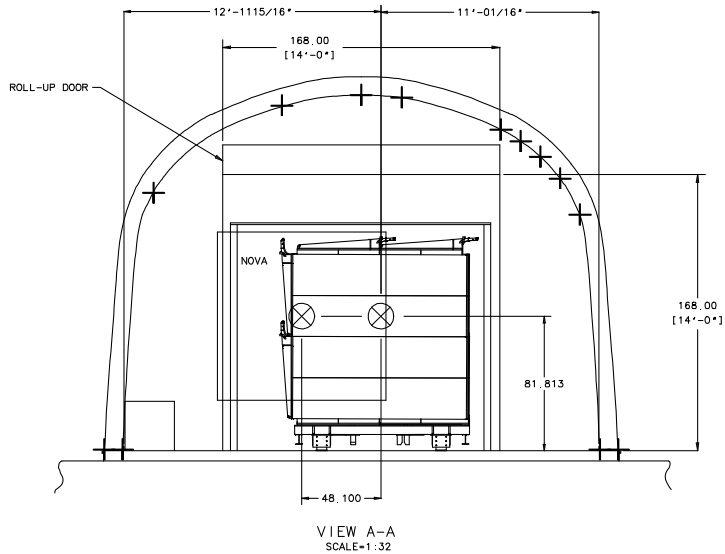
# Testbeam



- Installed in the MCenter Beamline
  - MC7
  - Space formerly used by MIPP
  - Downstream of Lariat
- Tertiary beam
  - e, mu, pi, p, K
  - momentum 0.2-2GeV/C



# Testbeam

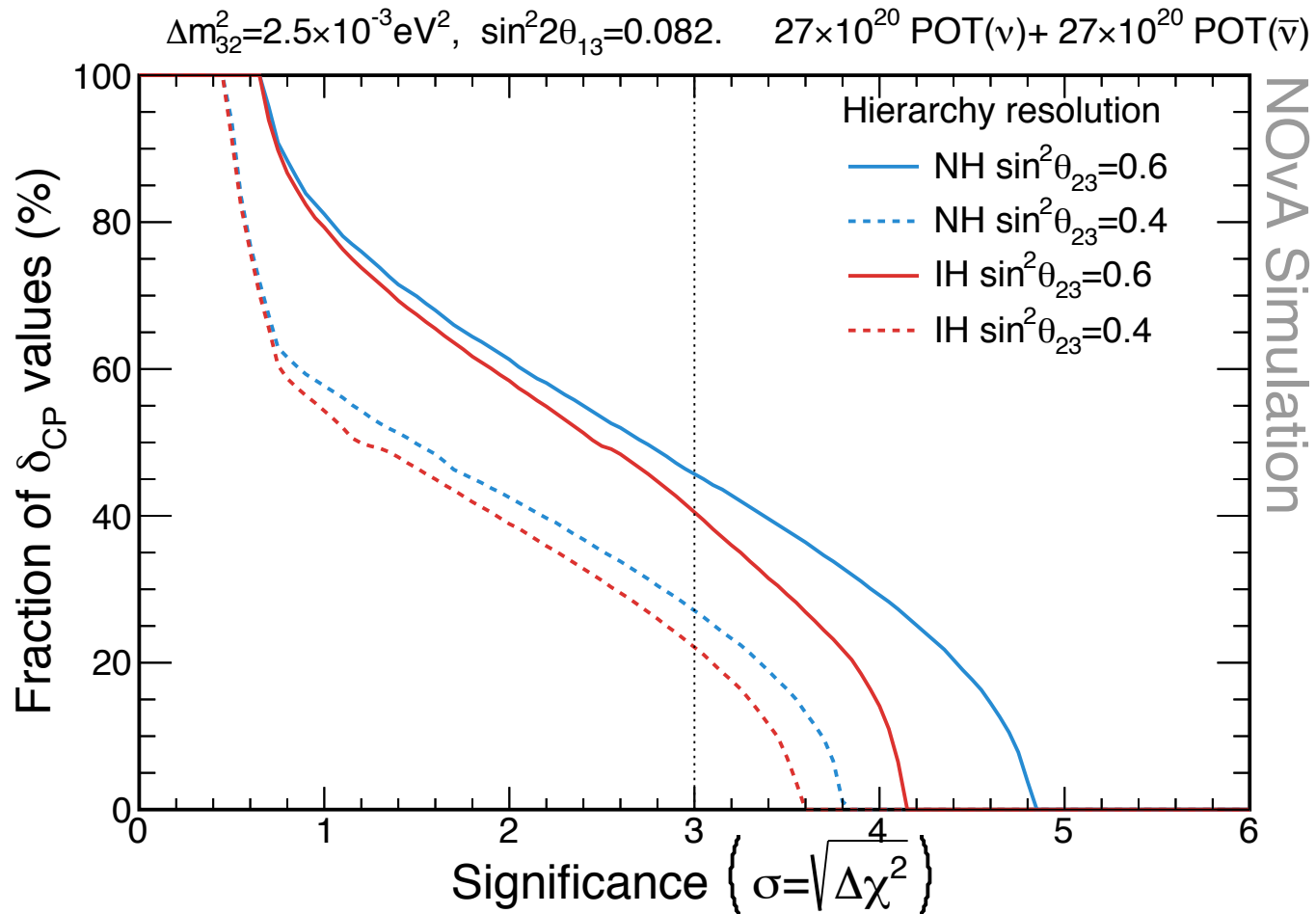


Detector Configuration	Two 2x2 31-plane blocks + One horiz. plane
Dimensions W(m) x H(m) x L(m)	2.6x2.6x4.1
Scintillator/plane (gallons)	86
Total Scintillator Volume (gallons)	5418
PVC Mass/plane (kg)	171.2
Adhesive Mass/plane (kg)	3.1
Total Empty Mass (kg)	10981
Total Scintillator Mass (kg)	17608
Total Detector Mass (kg)	28582
#APDs = #FEBs	126 (118 FEB 4.1 + 8 FEB 5.2)
#DCMs	3

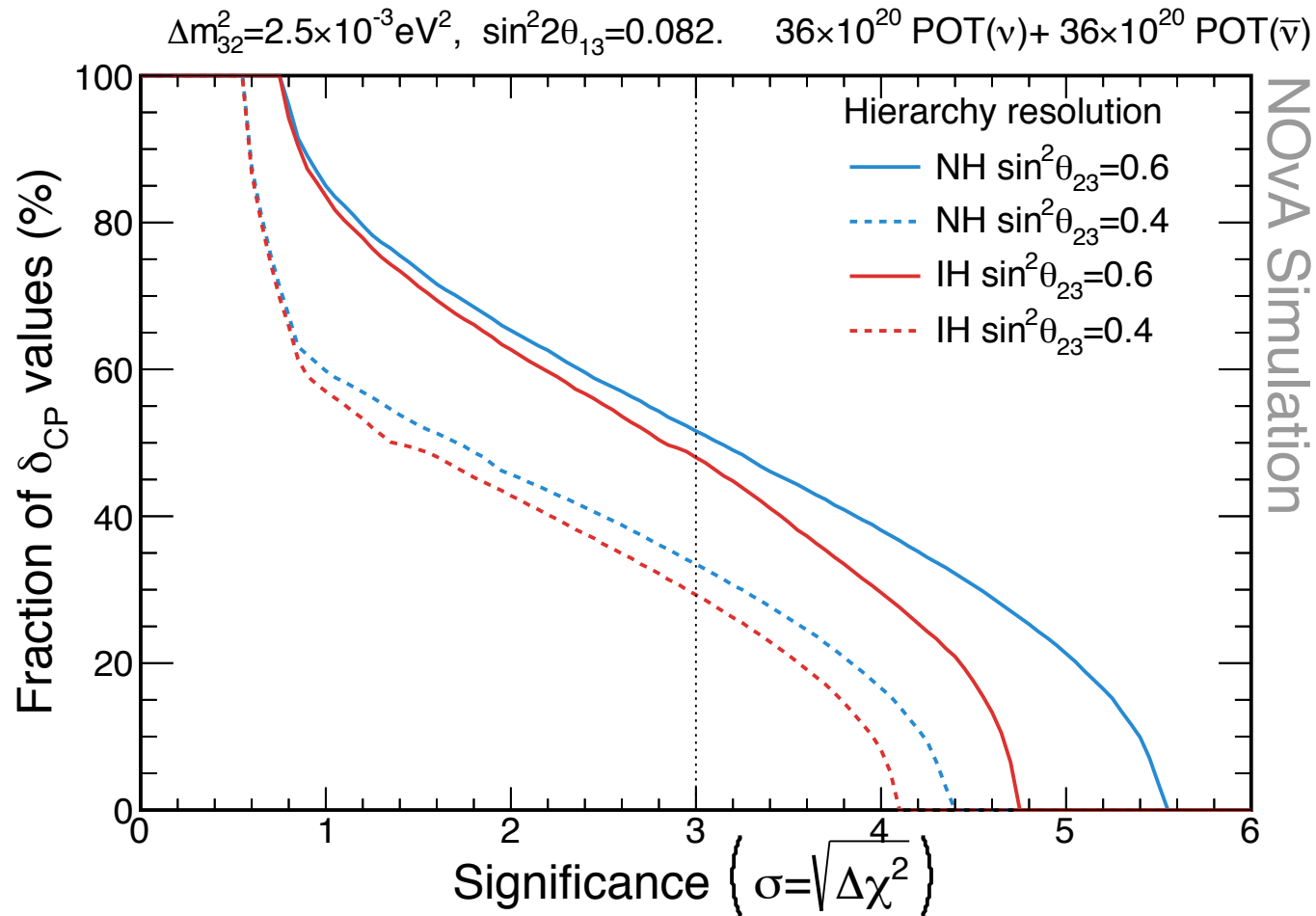
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# Backup—Sensitivity

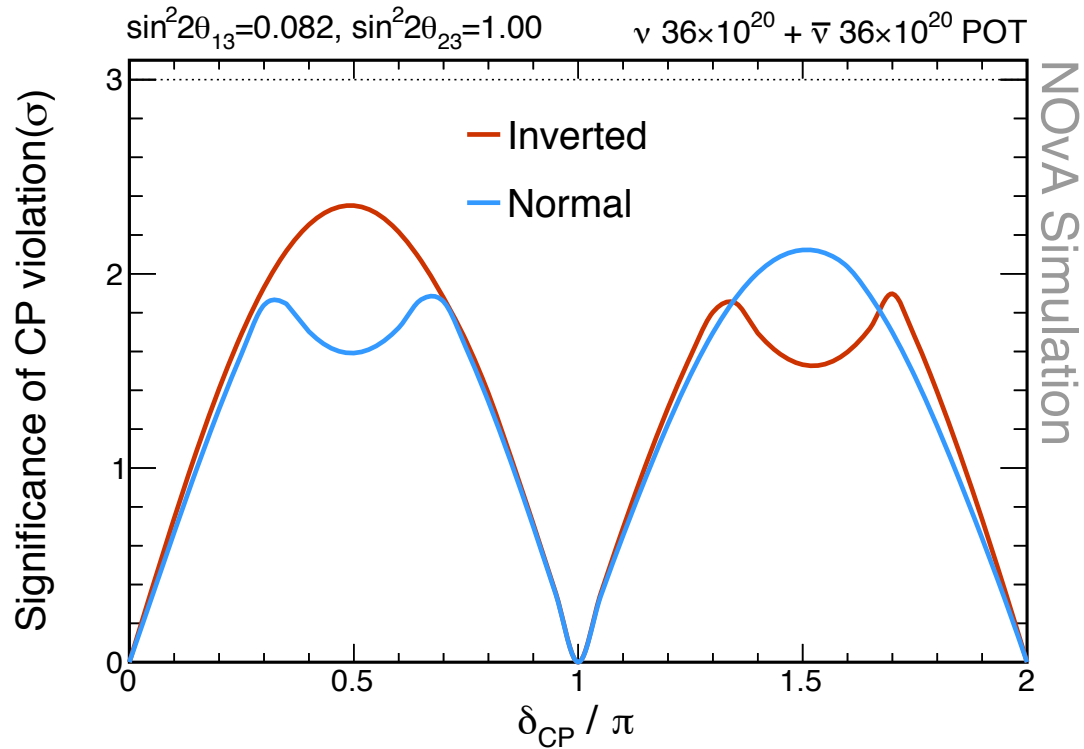
# AIP Sensitivity Gain



# AIP Sensitivity Gain



# CP Violation Sensitivity



$2+ \sigma$  sensitivity for CP violation in both hierarchies at  $\delta_{CP}=3\pi/2$  or  $\delta_{CP}=\pi/2$  (assuming unknown hierarchy) by 2024

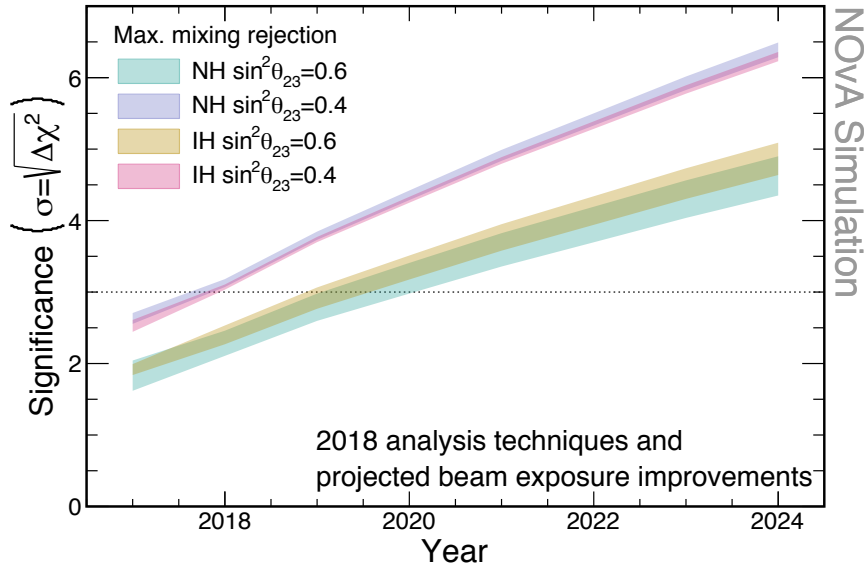


# Future Sensitivity



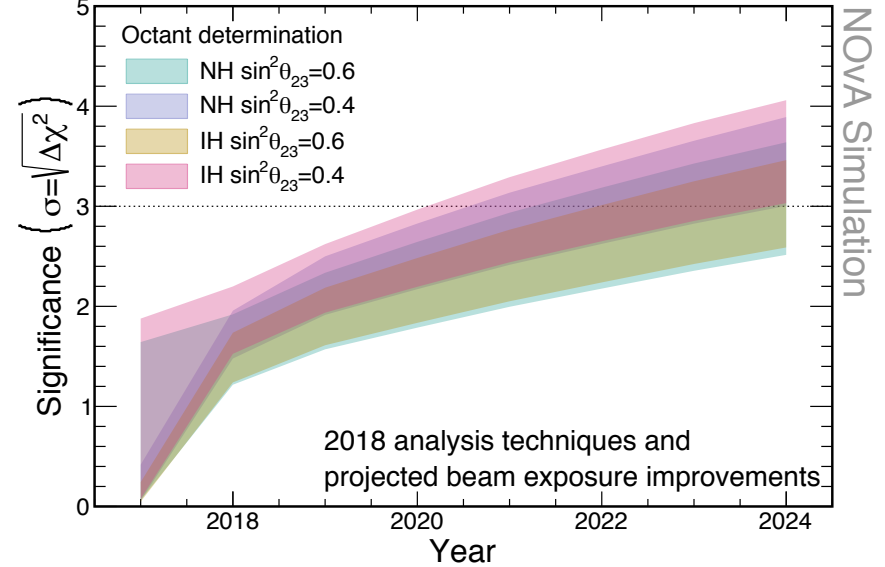
## Max mixing rejection

$\delta_{CP} \in [0, 2\pi], |\Delta m_{32}^2| = 2.5 \times 10^{-3} \text{eV}^2, \sin^2 2\theta_{13} = 0.082$



## Octant

$\delta_{CP} \in [0, 2\pi], |\Delta m_{32}^2| = 2.5 \times 10^{-3} \text{eV}^2, \sin^2 2\theta_{13} = 0.082$

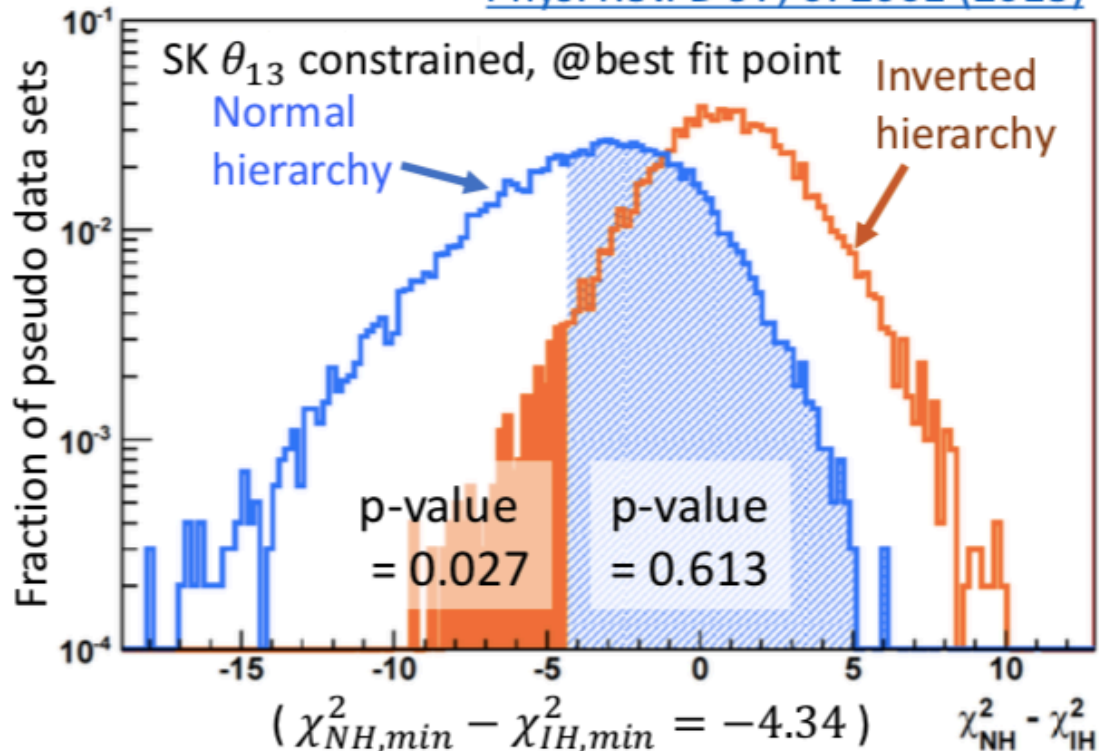
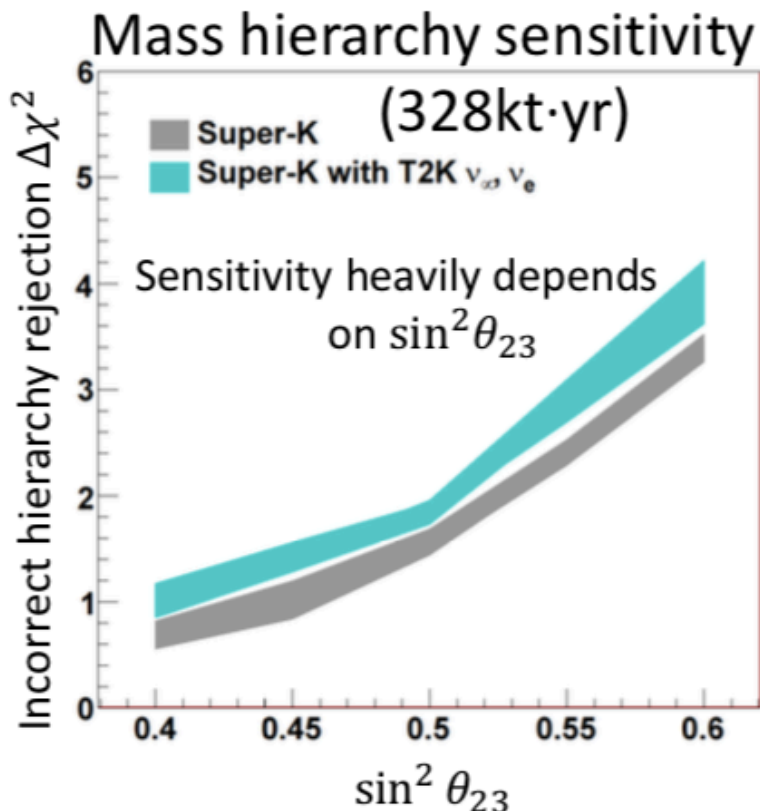


# Other Experiments



- T2K
  - M. Wascko—Nu2018
  - Bayes factor for NH/IH=7.9->NH preferred at 89%
- SuperK
  - Y. Hayuto—Nu2018
  - NH favored at 80.6 to 96.7% depending on  $\theta_{23}$
  - Expanding the fiducial volume
- Juno
  - Bjorn Wonsak—Nu2018
  - Start datataking in 2021
  - after 6 years, 3sigma or 4sigma MH resolution depending on constraint on  $\delta m_{\mu\mu}$
- KM3Net/Orca
  - U. Katz—Nu2018
  - 5sigma by 2024 at nova favored  $\theta_{23}$

[Phys. Rev. D 97, 072001 \(2018\)](#)



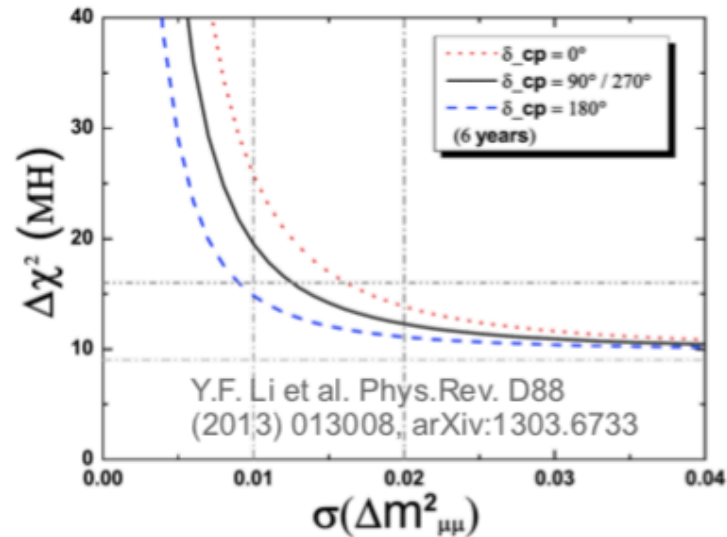
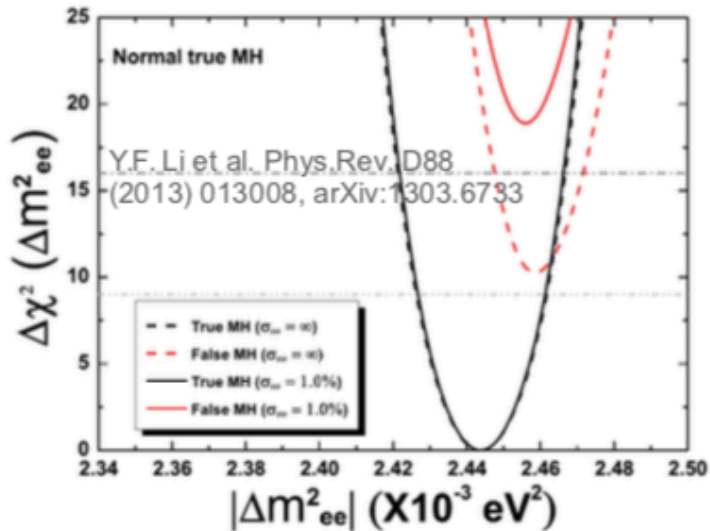
Estimate p-values using pseudo-data

for the smallest and largest  $\sin^2 \theta_{23}$ .

Hypothesis test  $\sim$  CL<sub>s</sub> method :  $CL_s(\text{IH rejection}) \equiv \frac{p_0(\text{IH})}{1-p_0(\text{NH})}$

Normal hierarchy is favored <span style="font-size: 2em; color: blue;">➔</span>	SK only	80.6 ~ 96.7%
	SK + T2Kmodel	91.5 ~ 94.5%

- **Measurement with or without constraint on  $\Delta m^2_{\mu\mu}$**



- **Sensitivity with 100k events (~6 yrs):**

- No constraint:  $\overline{\Delta\chi^2} > 9$
- With 1% constraint:  $\overline{\Delta\chi^2} > 16$

- **Reason for synergy:**

$$|\Delta m^2_{ee}| - |\Delta m^2_{\mu\mu}| = \pm \Delta m^2_{21} \cdot (\cos(2\theta_{12}) - \sin(2\theta_{12}) \sin(\theta_{13}) \tan(\theta_{23}) \cos(\delta))$$

Sign defined by MH

See H. Nunokawa et al, Phys.Rev. D72 (2005) 013009

# NMO measurement



- Primary signature:  
Energy-zenith distribution
- Inverse signatures for  $\nu$  and  $\bar{\nu}$ ,  
but signal measurable since  $\sigma(\nu) \approx 2 \sigma(\bar{\nu})$  and  $\Phi(\nu) > \Phi(\bar{\nu})$
- Measurement requires
  - best possible resolution in energy and zenith
  - separation  $\nu_e/\nu_\mu$
  - detailed understanding of systematics
- In-depth studies by KM3NeT and IceCube, extensive cooperation
- Results very similar

**P[2/161] S. Bouret**

Asimov and LLR median sensitivity after 3 years,  $\delta_{CP} = 0$

