

Tianlai



Peter Timbie
UW-Madison

Tremendous Arrays Workshop
Brookhaven National Lab
July 30, 2018

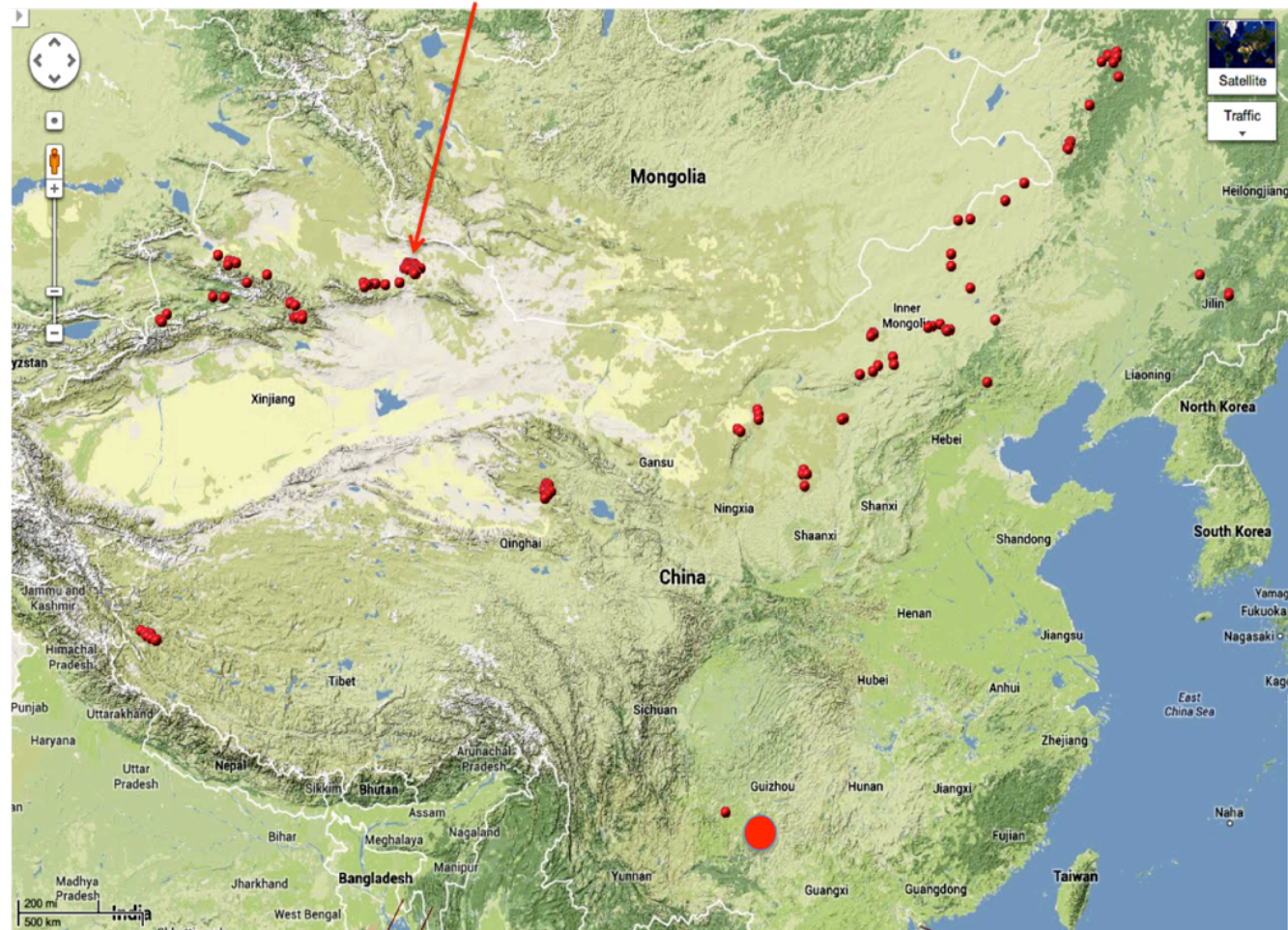


Site Selection

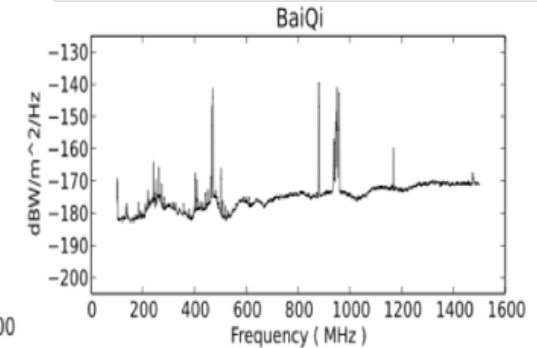
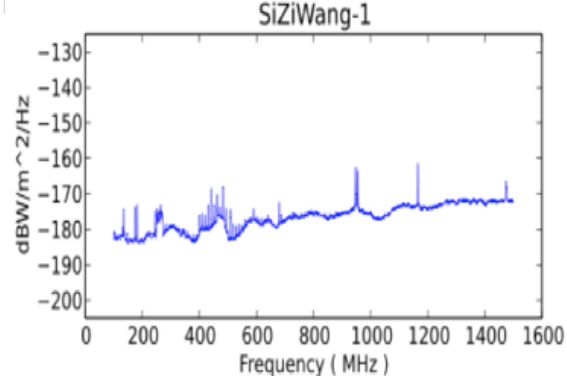
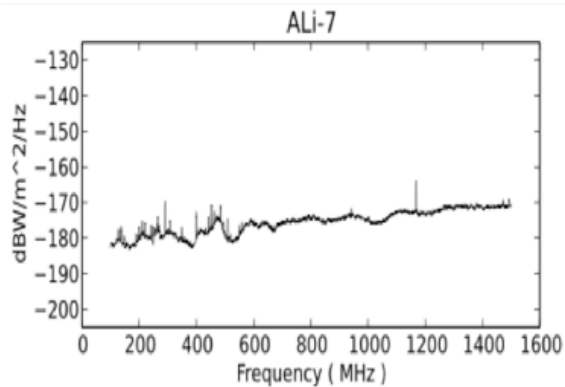
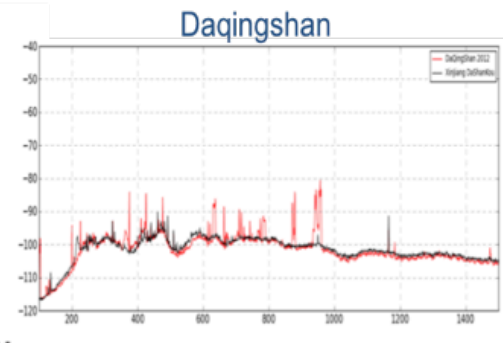
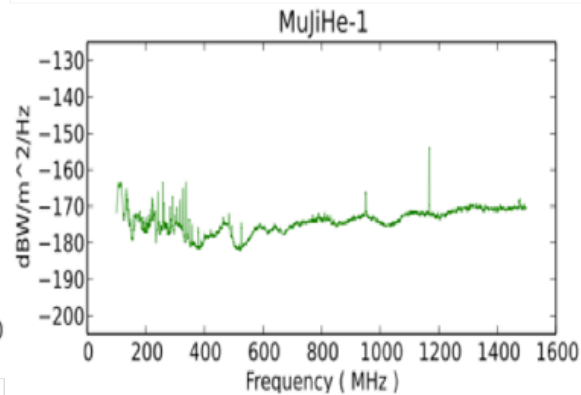
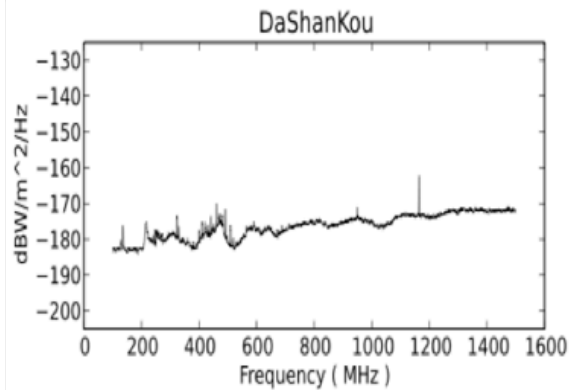
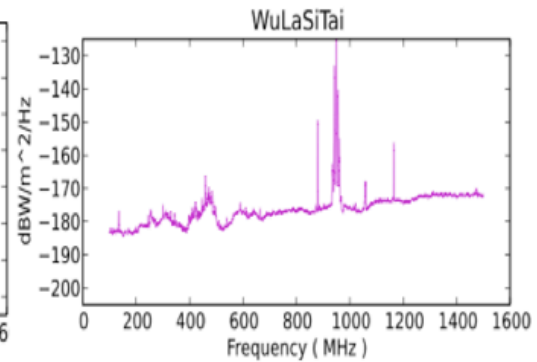
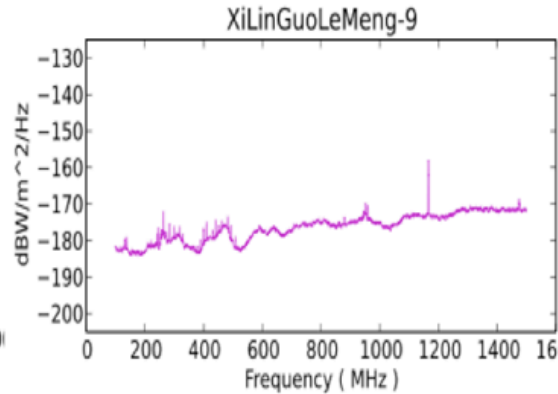
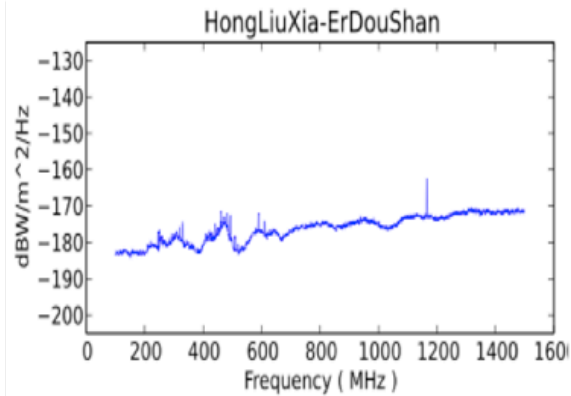
Site Requirements:

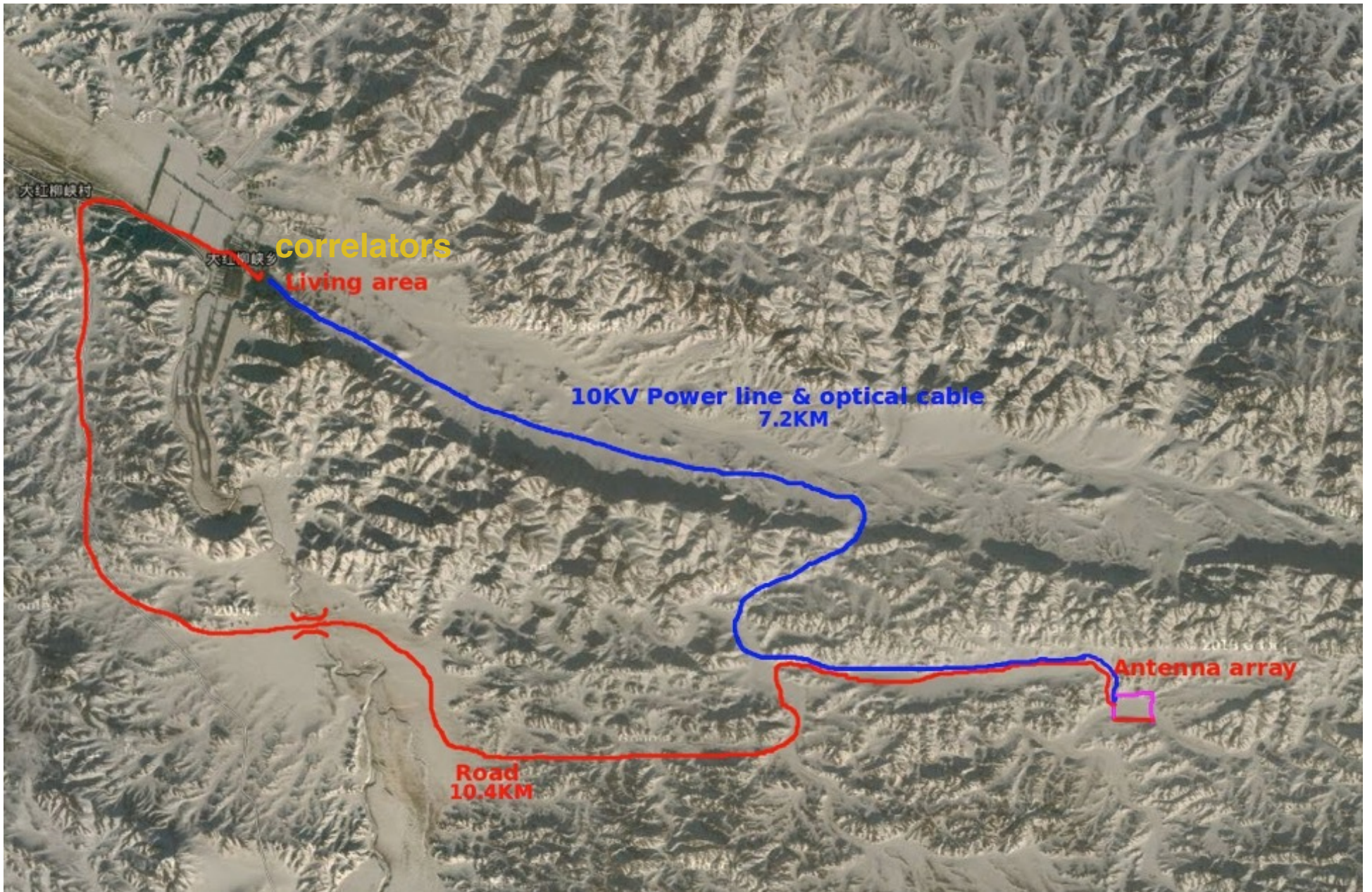
To minimize radio frequency interference (RFI), typically sparsely populated area, surrounded by mountains and hills

The Selected Site Hongliuxia



RFI spectra for some sites





Tianlai Facility



Receiver and Correlator



Tianlai Participants

China: Xuelei Chen (PI at NAOC),
CETC-54, Institute of Automation,
Hangzhou Dianzi U., XAO

US: J. Peterson (CMU)
P. Timbie, S. Das (Wisconsin)
A. Stebbins, J. Marriner (Fermilab),
G. Tucker (Brown)

France: R. Ansari, J.E Campagne,
M. Moniez (LAL/IN2P3)
J.-M. Martin, P. Colom (Obs. Paris)

Song (KASI), L. Wolz (Melbourne),
T-C. Chang (JPL), U-L. Pen (CITA)



The concept of “**tianlai**”-- the **heavenly sound** was coined by ancient Chinese philosopher Zhuang-Zi (Chuang-Tzu, 369BC-286BC)

Changjun.Gao



Fengquan Wu



Huli Shi



Juyong Zhang



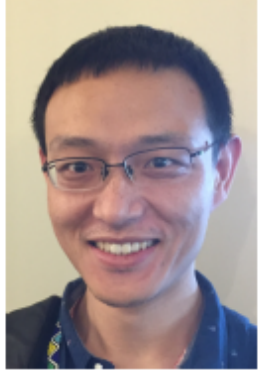
Xin Wang



Yan.Gong



Bin.Yue



Yichao Li



Yougang.Wang



Shifan Zuo



Yidong.Xu



Jialu.Zhu



Donghao Liu



Gong.Cheng



Rongli.Wang



Boqin Zhu



Wang Weixian



Tiantian Li



Jingyu Zhang



Jixia.Li



Zhiping Chen



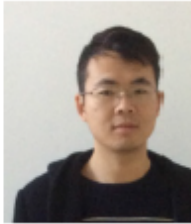
Qizhi Huang



Chenhui .Niu



Wenkai.Hu



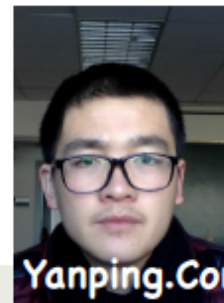
Kaifeng.Yu



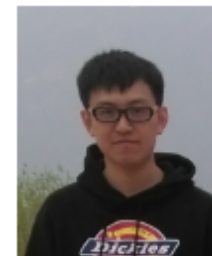
Wenxiao.Xu



Shijie.Sun



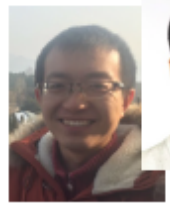
Yanping.Cong



Jiawei.Li



Qunxiong.Wang



Tianxian Mao



Ye.Cao

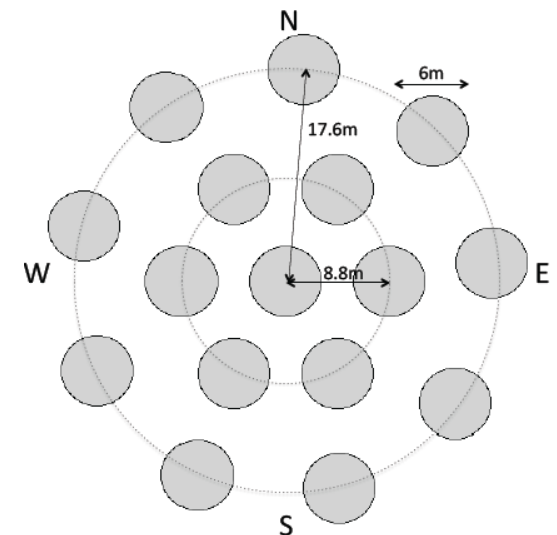
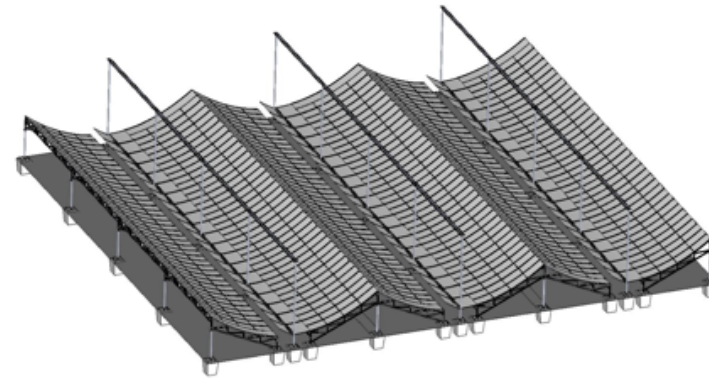
U.S. Participation in Tianlai

- In addition to large Chinese / French participation
 - Peter Timbie / Santanu Das / Trevor Oxholm & UG's - University of Wisconsin
 - Albert Stebbins / John Marriner - Fermilab
 - Jeff Peterson - Carnegie Mellon
 - Greg Tucker - Brown
 - open to new participants
- NSF grant: reduce/analyze 1st 3 years of data
 - Timbie (PI) / Stebbins received NSF-AAG Grant (2016 - 2019)
 - **Tianlai Analysis Center**
 - computing/storage @ FNAL & Open Science Grid
- Manpower limited

Tianlai Configuration(s)

- **Band** 700-800MHz ($0.58 < z < 1.05$)
1024 frequency channels
($\delta v = 100\text{kHz}$ $\delta z = 0.0002$)
tunable in 600-1420MHz
- **Cylinder Array** 3 x 15m x 40m cylinders
96 dual polarization feeds
4 sec sampling
- **Dish Array** 16 x 6m dishes
16 dual polarization feeds
1 sec sampling
- **Pathfinder+ Cylinder Array**
216 dual polarization feeds
4 sec sampling
- **Full Cylinder Array** 8 x 15m x 120m
2048 dual polarization feeds
400-1420MHz

Pathfinders to demonstrate
basic principle
and encounter issues rapidly



Pathfinder Highly Configurable Transit Telescope

- Tuneable:
 - $600 < \nu < 1420\text{MHz}$ ($0 < z < 1.36$) [fixed 100 MHz bandwidth]
- Cylinders:
 - equal/unequal spacing of feeds on cylinder
 - feed placement on cylinders may be same or different
 - redundancy vs broader u-v coverage (less mode mixing)
- Dishes:
 - pointable / tracking
 - arrangeable in any ground configuration

R&D: plan to play with all the knobs

Tianlai *Long Term Goals*

With **lessons learned** from current pathfinder arrays fill out and expand arrays ...

- make lower noise / higher resolution images of Northern sky.
- make HI maps, $P_{\text{HI}}[k]$ for cosmological parameters e.g. BAO/DE.

Table 2
Telescope Configurations Used in This Paper

Bull et al. ApJ 2015

Experiments	T_{inst} [K]	$N_d \times N_b$	D_{dish} [m]	D_{min} [m]	D_{max} [m]	ν_{crit} [MHz]	$\nu_{\text{max}}^{\text{IM}}$ [MHz]	$\nu_{\text{min}}^{\text{IM}}$ [MHz]	$\Delta\nu^{\text{IM}}$ [MHz]	z_{min}	z_{max}	S_{area} [deg ²]
Targeted IM												
• BAOBAB-128	40	128 × 1	1.6	1.6	26.0	...	900	600	300	0.58	1.37	1,000
BINGO	50	1 × 50	25.0	1260	960	300	0.13	0.48	5,000
◇ CHIME	50	1280 × 1	20.0	800	400	400	0.77	2.55	25,000
FAST	20	1 × 20	500.0	1000	400	600	0.42	2.55	2,000
• MFAA	50	3100 × 1	2.4	0.1	250.0	950	950	450	500	0.49	2.16	5,000
◇ Tianlai	50	2048 × 1	15.0	950	550	400	0.49	1.58	25,000

Table 4
1D Marginal Constraints (68% CL) on the Extended Λ CDM Model, Including the Planck Prior

Experiments	A / 10^{-2}	h / 10^{-3}	Ω_K / 10^{-4}	Ω_{DE} / 10^{-3}	n_s / 10^{-4}	σ_8 / 10^{-3}	γ / 10^{-2}	w_0 / 10^{-2}	w_a / 10^{-2}	FOM
BAOBAB-128	24.3	50.2	71.3	36.6	38.5	8.1	9.0	33.3	71.4	8.0
BINGO	25.8	30.8	90.0	16.1	38.5	8.2	2.8	44.1	172.5	7.8
CHIME	3.0	8.7	9.7	7.1	30.2	5.2	3.4	5.0	15.1	288.1
FAST	7.5	13.5	16.0	10.1	33.5	6.4	3.2	7.1	18.5	144.7
MFAA	5.7	11.9	14.1	9.1	32.2	6.0	3.1	6.3	17.2	165.7
Tianlai	3.6	8.0	11.9	6.3	28.7	4.9	2.4	4.0	12.0	383.3

Tianlai *Pathfinder* Goals

- Implement large interferometric **arrays** in a quiet site in western China to obtain a high fidelity 3D image of Northern Sky w/ 100MHz bandwidth. Accurate **beam calibration** is essential.
- Compare cylinder arrays with dish arrays (also cross correlate dish/cylinder)
- Experiment with cylinder feed arrangements / dish placements.
- Experiment with different calibration schemes: artificial calibration sources on ground and in air; natural calibrations sources: Sun and other bright sources, pulsars, holography; numerical modeling of beam.
- Deep imaging of North Polar Cap (NPC) with dish array.
- Optimize algorithms **and** telescope arrangement for best foreground removal.
- Construct redshift space HI emission maps especially in NPC.
- Vary frequency band up to 1.4GHz (lo-z 21cm) so as to make maps to compare with LSS with optical redshift surveys SDSS and NCCRS

Timeline

- **2014:**

- basic infrastructure: roads, buildings, power, optical fibers
- electronics design

- **2015:**

- scientific infrastructure: reflectors finished
- much of electronics installed
- first fringe
- engineering / debugging

- **2016:**

- engineering / debugging
- astronomical imaging of bright sources
- several runs with dish array : mostly NCC
- 1 run with with cylinders

- **2017:**

- fix hardware issues (some down time)
- develop test calibration methods (mostly offline analysis development while taking data)
- cylinder+dish runs

- **2018:**

- production runs

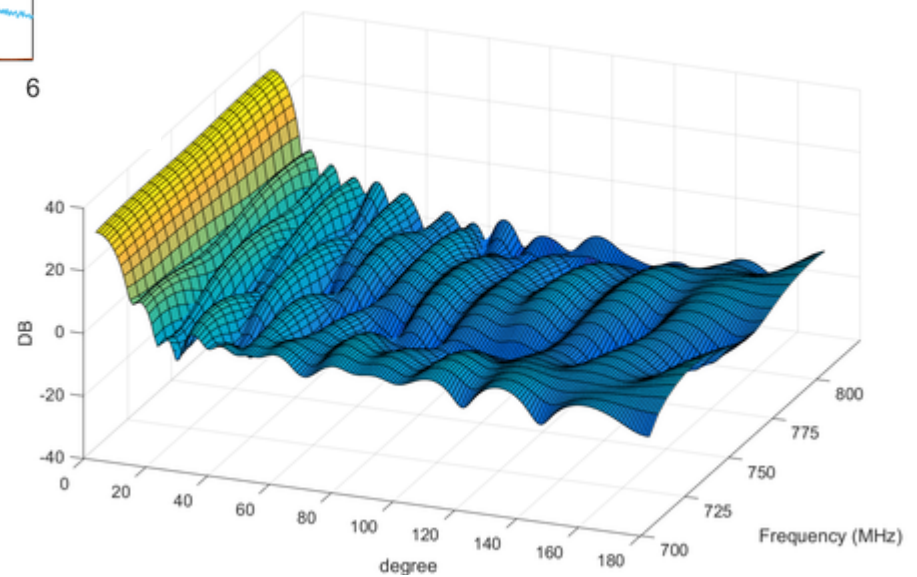
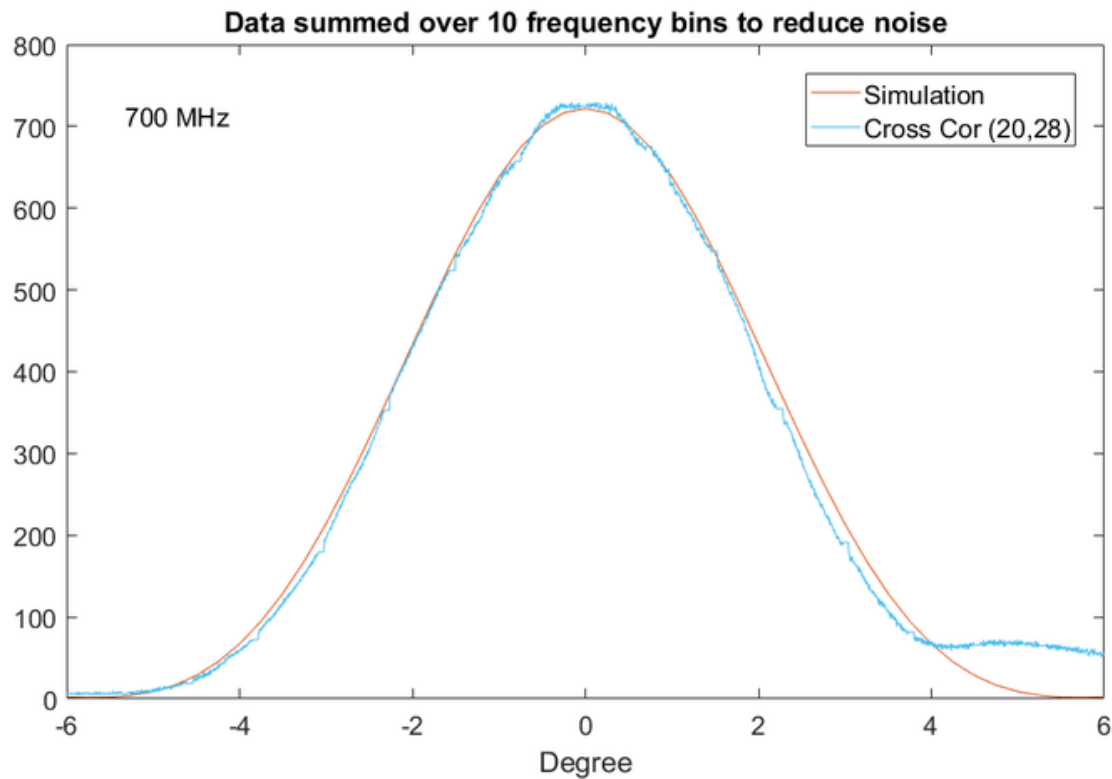
- **2019:**

- production runs
- tune dishes to lo-z
- transient backend / outriggers?

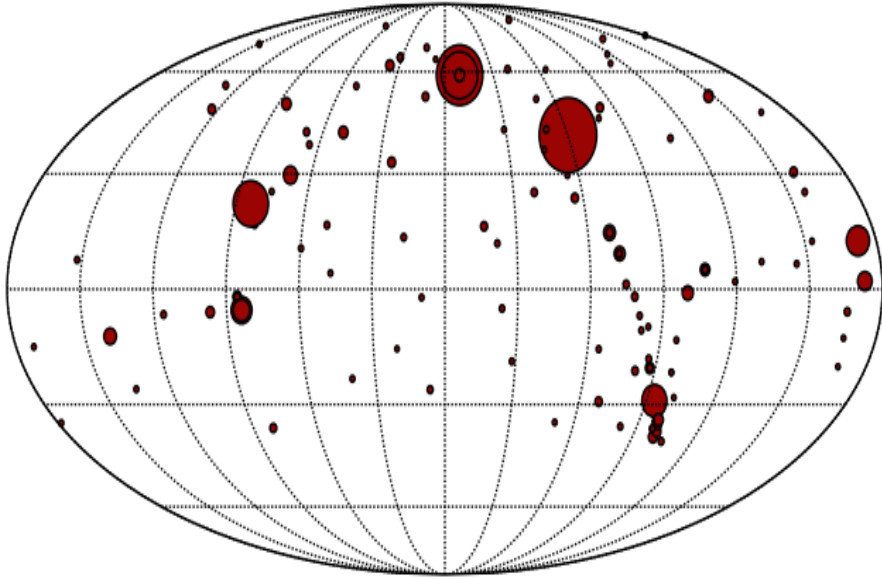
Tianlai Analysis Center

- software development
- computer resources / data storage
- data reduction
- [data visualization tools/repository](#)
- feedback for instrumentation / observational strategy
- EM simulations and modeling of beams
- beam calibration techniques
- [transient detections](#)

EM Simulations

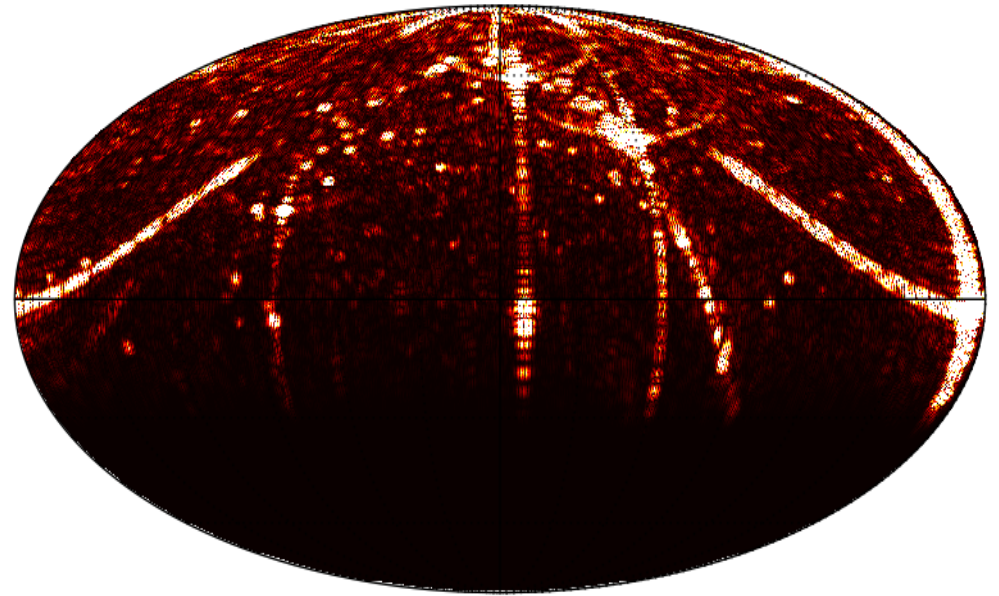


First Light: sky maps with cylinders



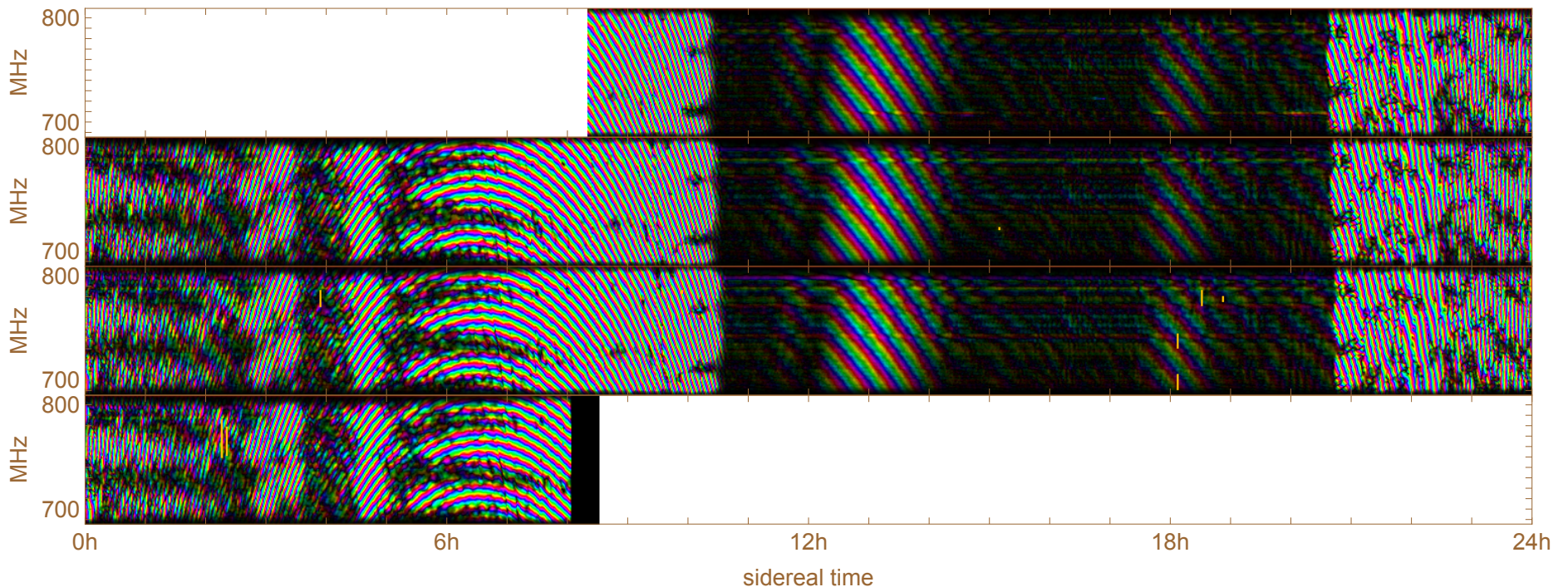
NVSS (1.4GHz) bright sources

S. Zuo et al., in preparation



First light image with 5 frequency channels (0.1 MHz each) from data taken during first light 2016.09.27-2016.09.30

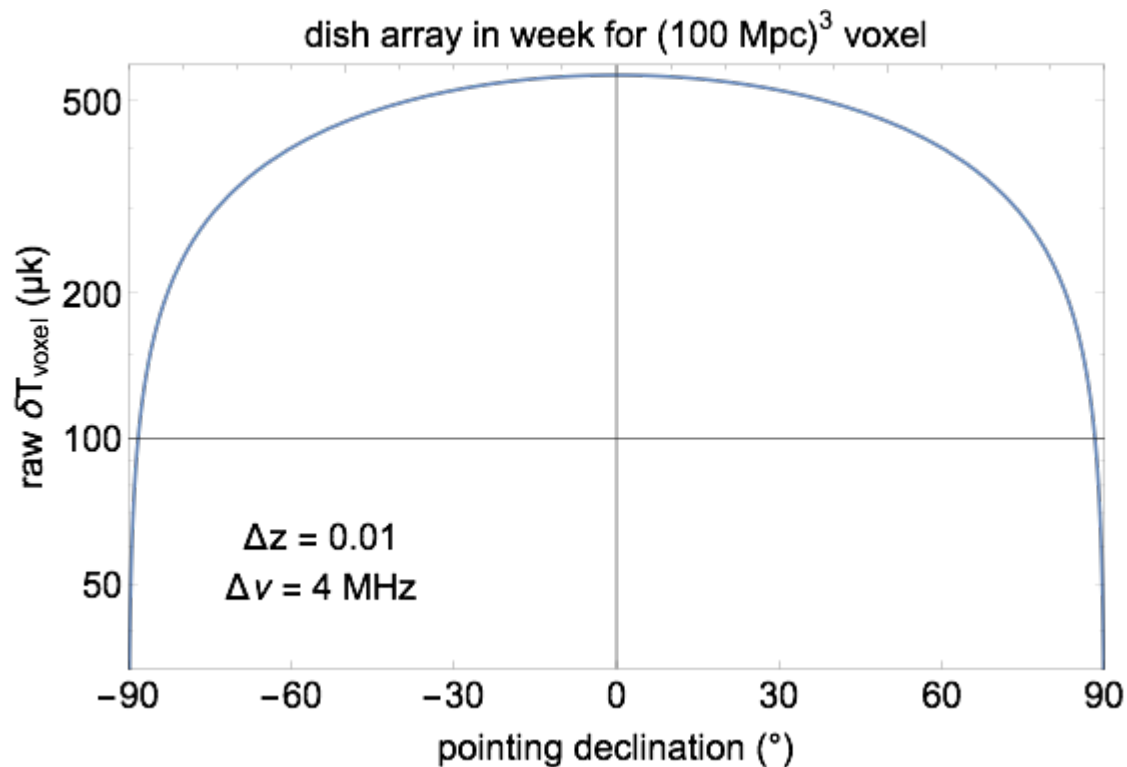
First Light: waterfall plots from dish array



A. Stebbins'
visibility visualizer

- Visibility for one of 528 baselines in dish array
- Pointed at NCP for 4 days, June 2016
- Data avg'd over 1 minute, 1 MHz
- Intensity is magnitude of visibility, color is phase

Planned North Polar Survey with Dish Array

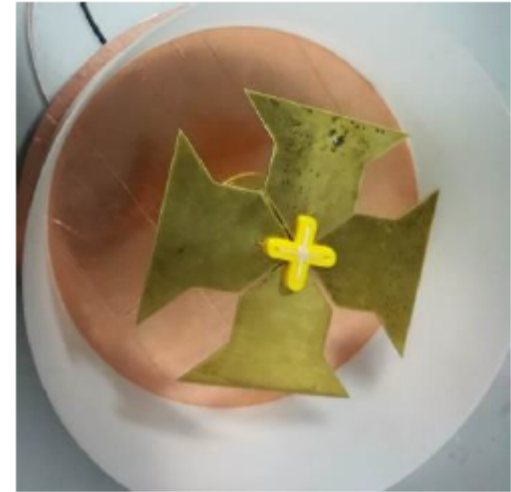


- By pointing dish array toward pole will integrate down to low map noise temperature rapidly.
- **WANTED:** NCC optical spectroscopic survey of existing photometric survey to compare to

Beam Mapping with Drone



1.6 m



- Matric-600 Pro
- 18 min. flight for 5.5 kg payload
- 2500 m max altitude (but 500 m regulations)
- Lightweight dish feed as transmitter
- First tests with broadband noise source

Tianlai Target Of Opportunity

- Tianlai Analysis Center NSF Grant ends in 2019!
- **but** Tianlai will continue and expand.
- continued/expanded US participation after 2019?
- most flexible developmental machine for HI intensity mapping
- excellent site wrt RFI.
- Tianlai buildout \gtrsim CHIME $t_{\text{Tianlai}} > t_{\text{CHIME}}$
- low entry costs relative to a new scope
- Chinese lead in funding

21 cm Cosmology workshop and Tianlai Collaboration Meeting

2018-09-17--2018-09-21 Pingtang

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Important Deadlines

- Deadline for the registration: **August 5, 2018**

