

Novel constraints to recent MiniBooNE explanations: Laying siege to new physics

Carlos A. Argüelles

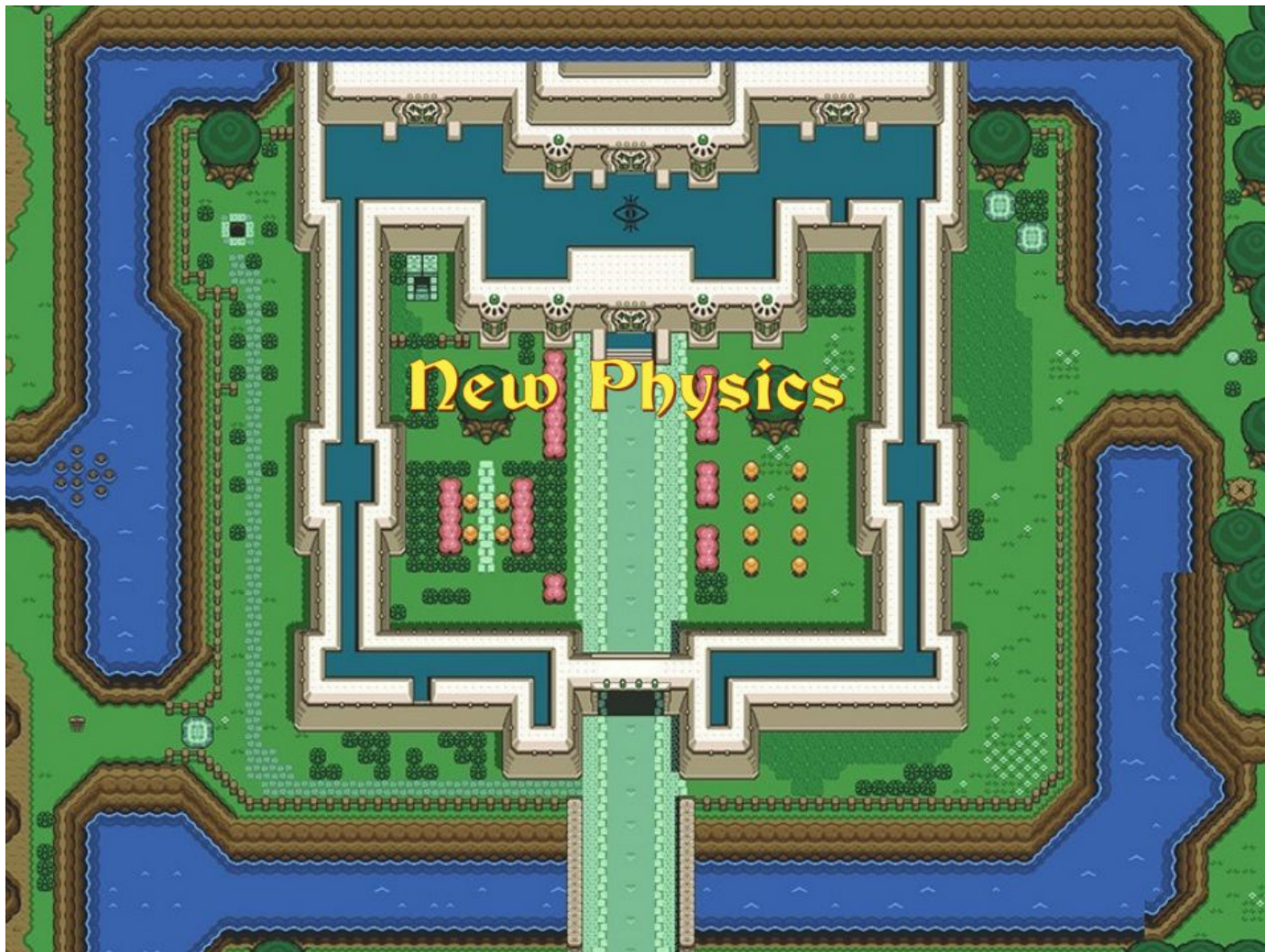
also starring

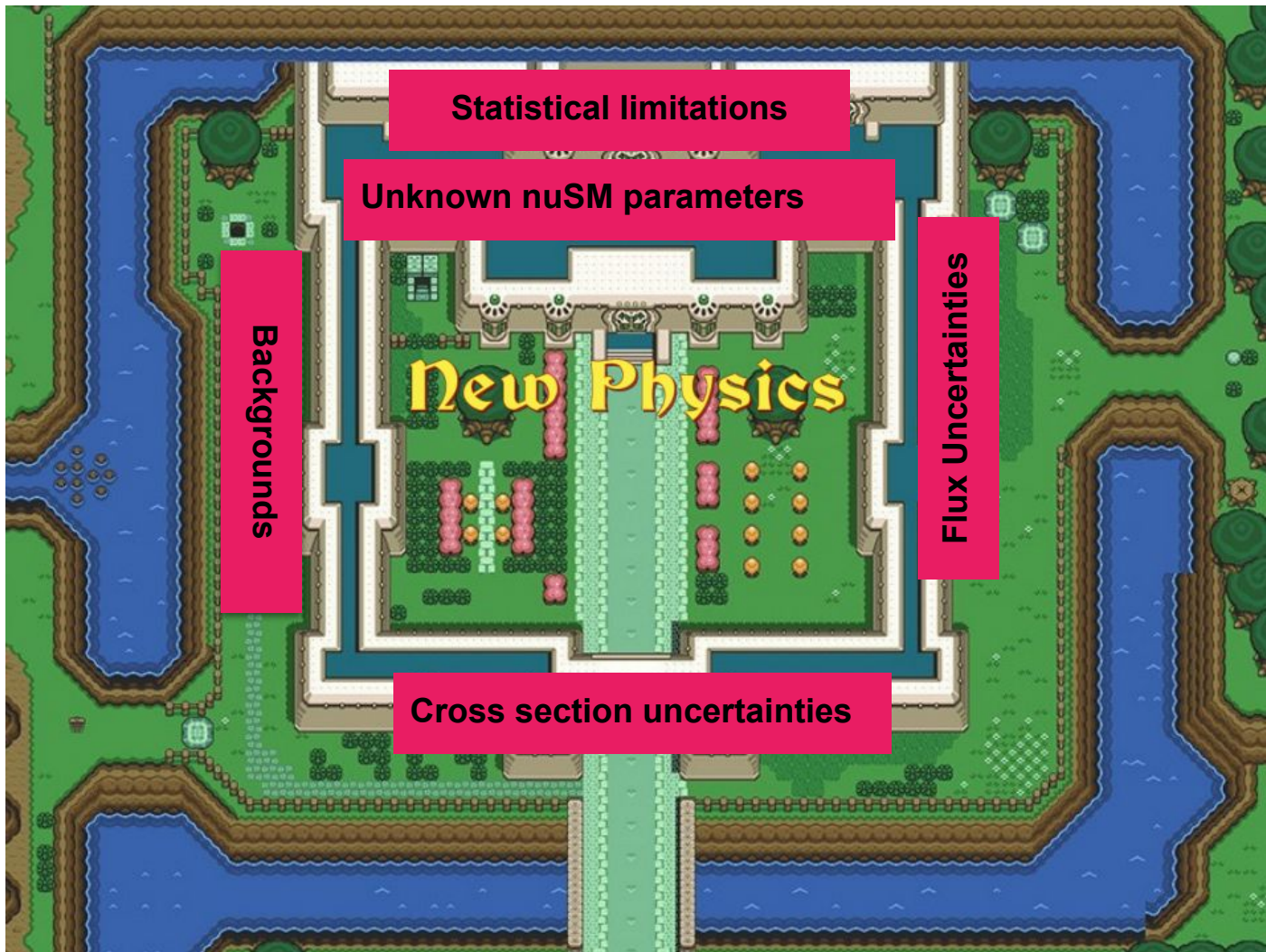
Matheus Hostert (Durham) and Yu-Dai Tsai (FNAL)

Siege

noun

a military operation in which enemy forces surround a town or building, cutting off essential supplies, with the aim of compelling the surrender of those inside.





Statistical limitations

Unknown nuSM parameters

Backgrounds

Flux Uncertainties

Cross section uncertainties

New Physics

The MiniBooNE excess

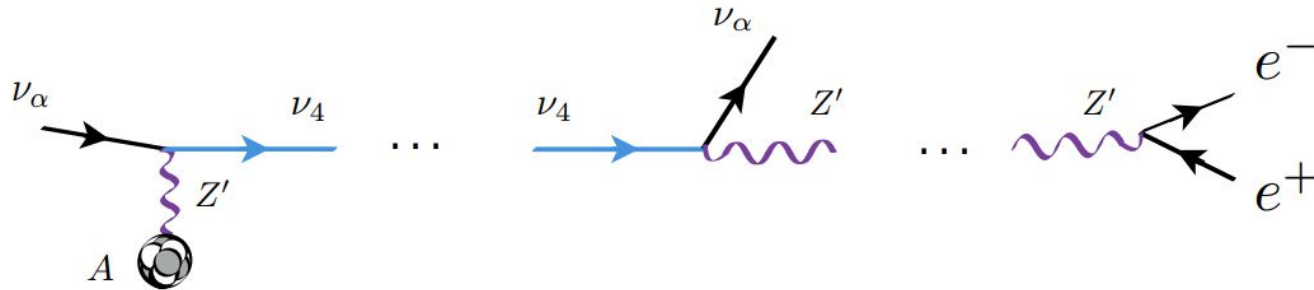
The MiniBooNE excess is an electron-neutrino-like observation above SM backgrounds. Explanation ideas:

- More electron-neutrino: $P(\nu_\mu \rightarrow \nu_e) > 3\nu_{SM}$ prediction. Sterile neutrino*?
- Electron-like background: more gammas?
- Recent proposal: new neutrino interaction that produces a pair of electrons.

* see talk tomorrow by C. Ternes@PONDD for a status on this; for tension and recent status with sterile neutrino explanations see Dentler et al. [1803.10661](#); also check out Moss et al. [1711.05921](#), Liao et al. [1810.01000](#), and Denton et al. [1811.01310](#) for recent ideas on how to get around these problems. See also Esmaili et al. [1810.11940](#) for complications with some of these ways out.

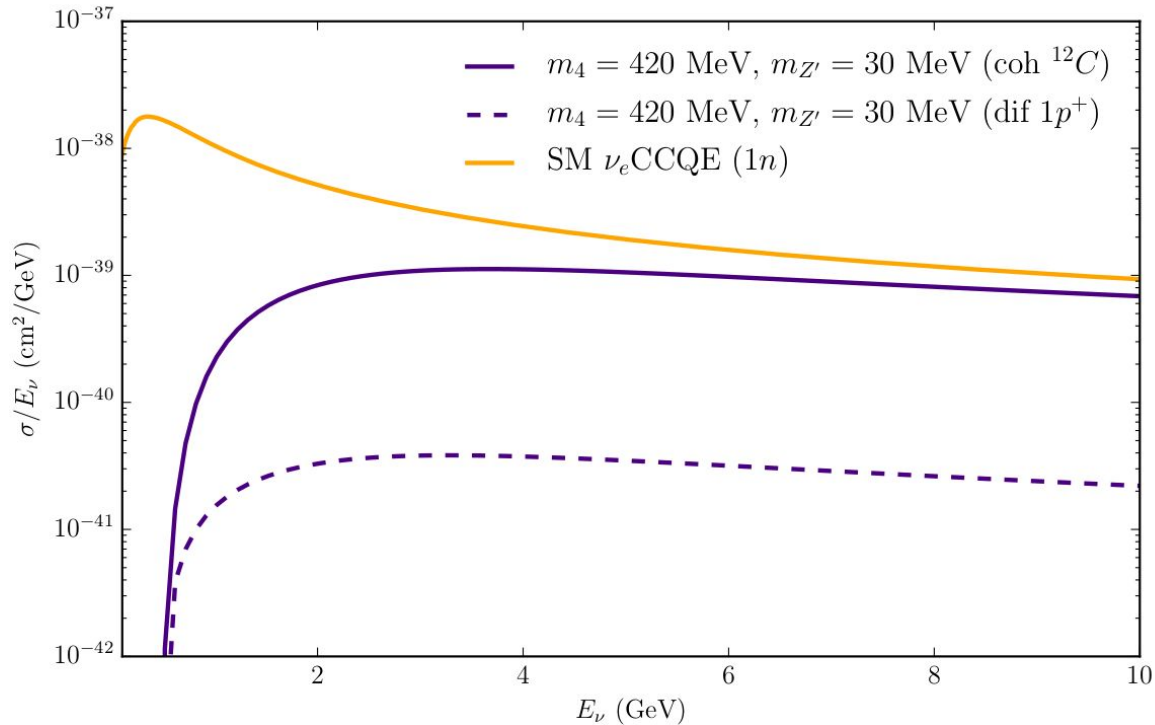
New proposal key features

- Little/no hadronic activity [depends on ratio of diffractive to coherent contributions]
- Small angular separation wrt to the booster beam. (see arXiv:1303.4587 and arXiv:1705.00353 for problems with angular distributions)
- Energy distribution must fit the excess observation.



See Bertuzzo et al. 1807.09877 and similar model by Ballett et al. 1808.02915.

How big of a cross section are we talking about?



Note that
CCQE is per
nucleon;

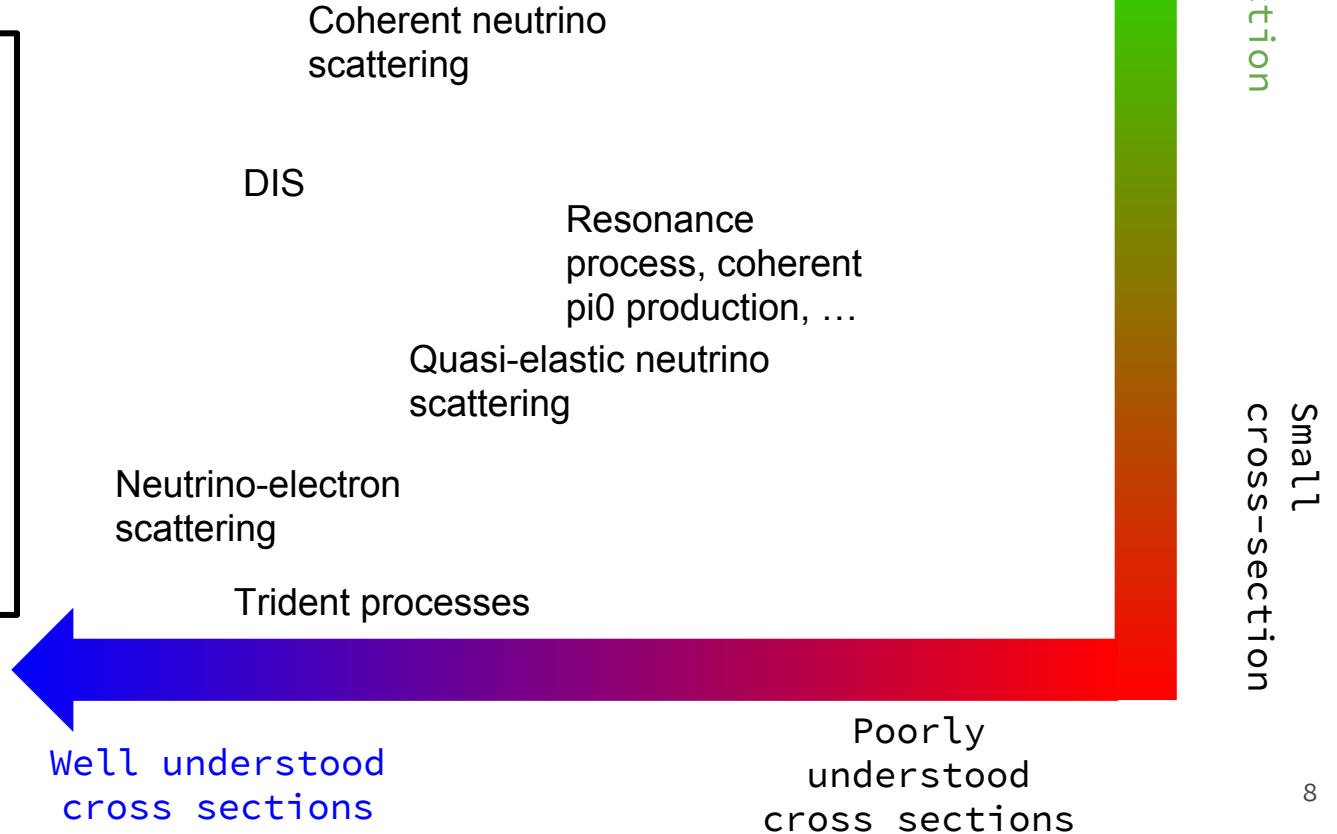
BSM-coh is
per atom.

Here we use benchmark (BP) point parameters reported by Bertuzzo et al. 1807.09877.

Strategy

NOTE that I am NOT putting scales in this diagram!

Looking for novel neutrino interactions requires understanding of our SM neutrino interactions. This is tough; see e.g. talks on Monday.



Strategy

NOTE that I am NOT putting scales in this diagram!

Looking for novel neutrino interactions requires understanding of our SM neutrino interactions. This is tough; see e.g. talks on Monday.

Coherent neutrino scattering

DIS

Resonance process, coherent π^0 production, ...

Quasi-elastic neutrino scattering

Neutrino-electron scattering

We are going to work here in this talk

Trident processes

Large cross-section

Small cross-section

Well understood cross sections

Poorly understood cross sections

Neutrino-electron scattering measurements

We have measured neutrino-electron scattering @:

- **LSND**
- **TEXONO**
- **Borexino**
- **SuperK**
- **MINERvA Low-Energy**
- **CHARM-II**

**Too low energies for BSM
case of interest**

**We will focus on these
experiments.**

Will measure it very soon @ **MINERvA Medium-Energy, NOvA,**
and later @ **DUNE.**

Strategy

Electron-neutrino-like scattering search



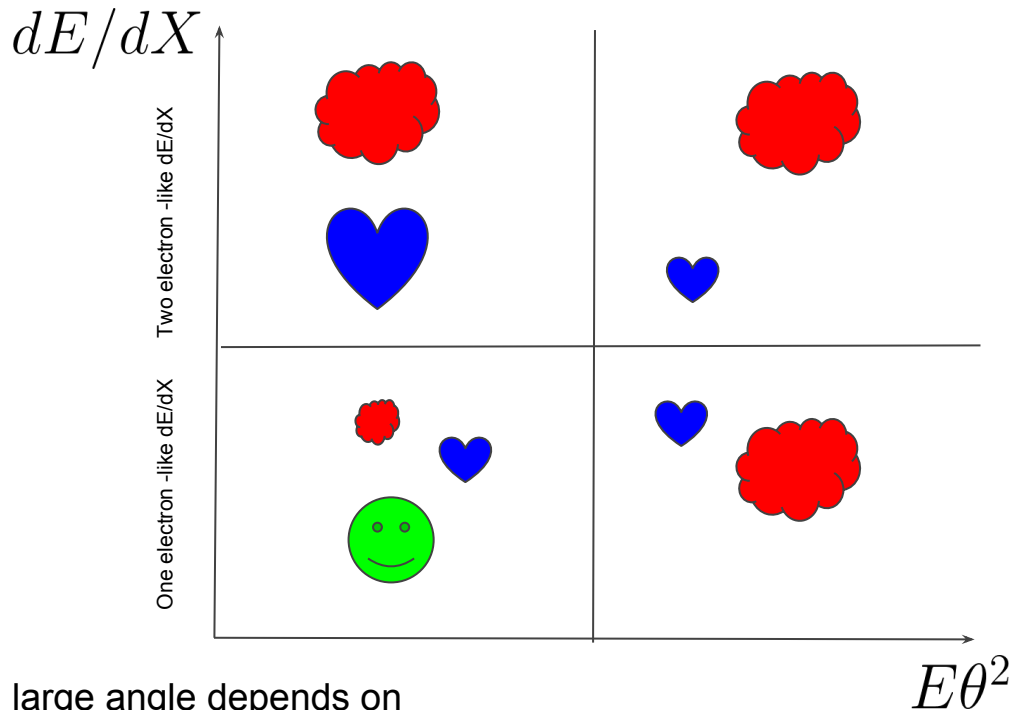
Background,
e.g. NCpi0



Neutrino-
electron
scat.



Recent BSM-MB
explanations

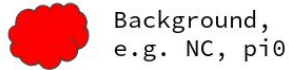


*how much *leakage* onto the large angle depends on model parameter and neutrino energy.

Working around limited information!

By design at final cut level CHARM-II and Minerva measurements have small backgrounds: also means small amount of BSM-signal leaking in. We cannot use the final event samples to constrain the new models :(!

Would be great if we had access to the reconstructed electron energy and angular distributions at different cut levels.



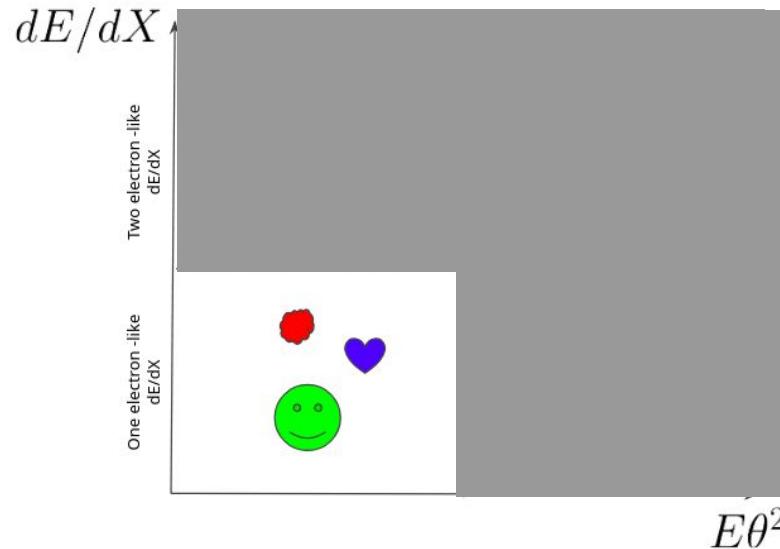
Background,
e.g. NC, π^0



Neutrino-
electron
scat.



Recent BSM-MB
explanations

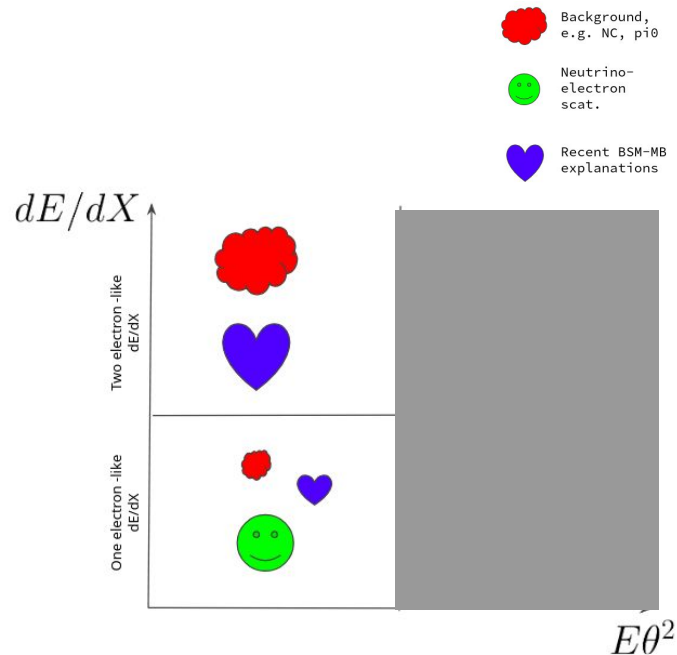
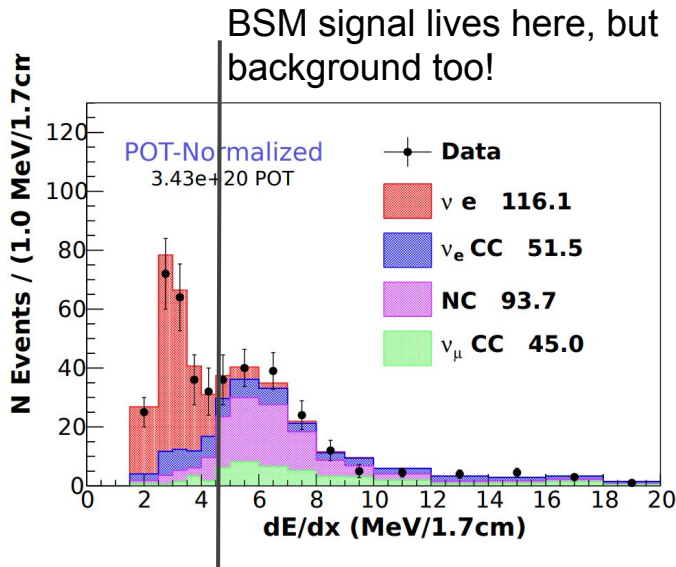


MINERvA analysis strategy

For MINERvA we are going to use the dE/dX distribution of candidate electron-neutrino scattering events.

All MINERvA cuts applied, except for the final dE/dX cut!

Note that backgrounds have been tuned here!



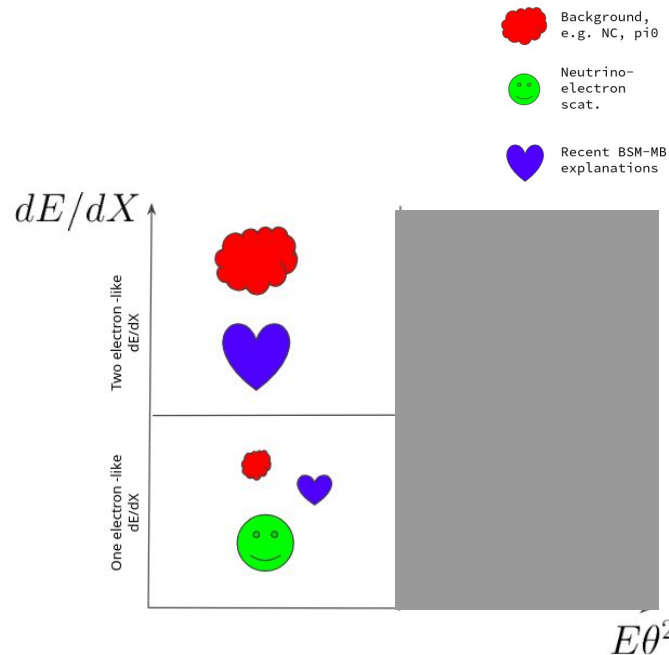
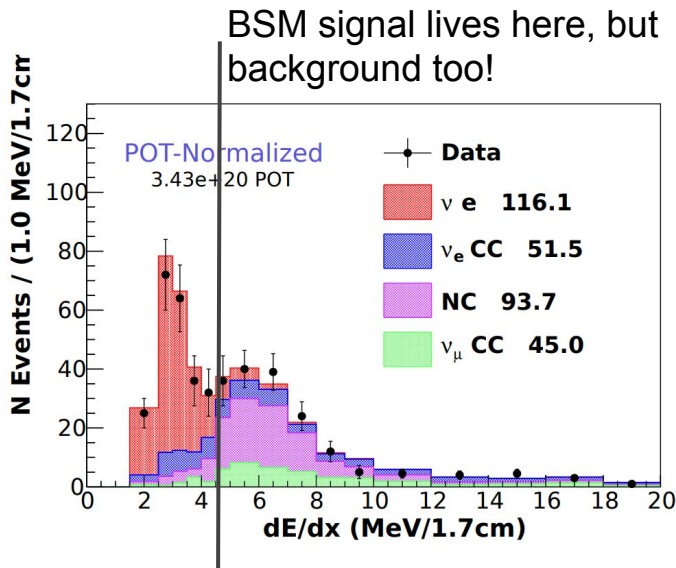
MINERvA analysis strategy

Parameter	Tuned value
ν_e	0.76 ± 0.03
ν_μ NC	0.64 ± 0.03
ν_μ CC	1.00 ± 0.02

For MINERvA we are going to use the dE/dX distribution of candidate electron-neutrino scattering events.

All MINERvA cuts applied, except for the final dE/dX cut!

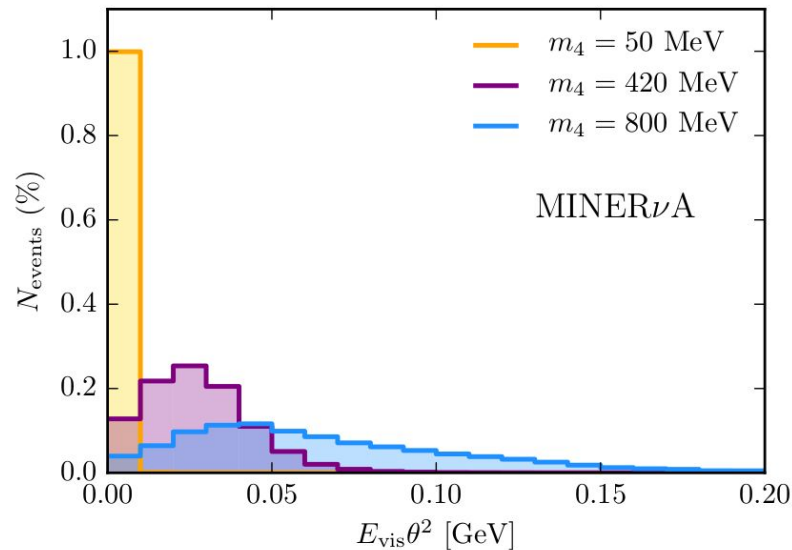
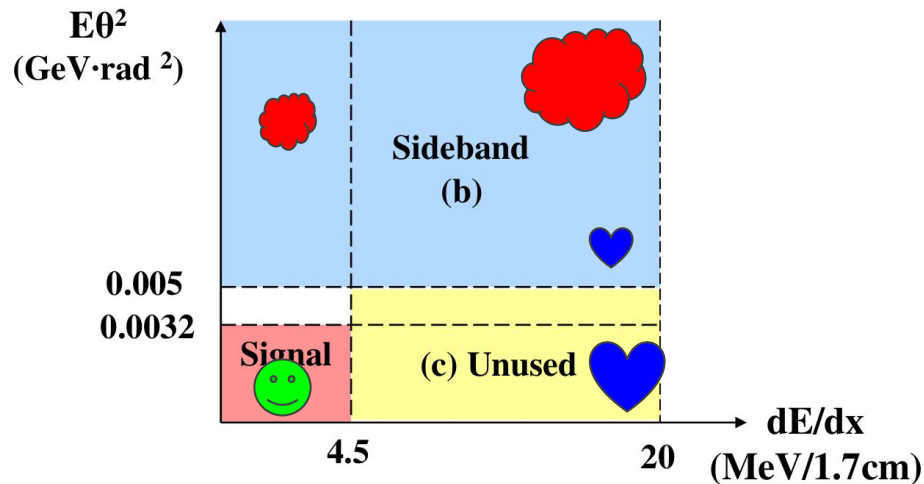
Note that backgrounds have been tuned here!



Sidebands used for tuning background on Minerva

Tuning parameters diagram from J. Park thesis:

<http://lss.fnal.gov/archive/thesis/2000/fermilab-thesis-2013-36.pdf>



For large heavy neutrino masses the BSM contribution leaks the sideband used to constrain the background on the neutrino electron scattering region.

Minerva: Our Analysis setup

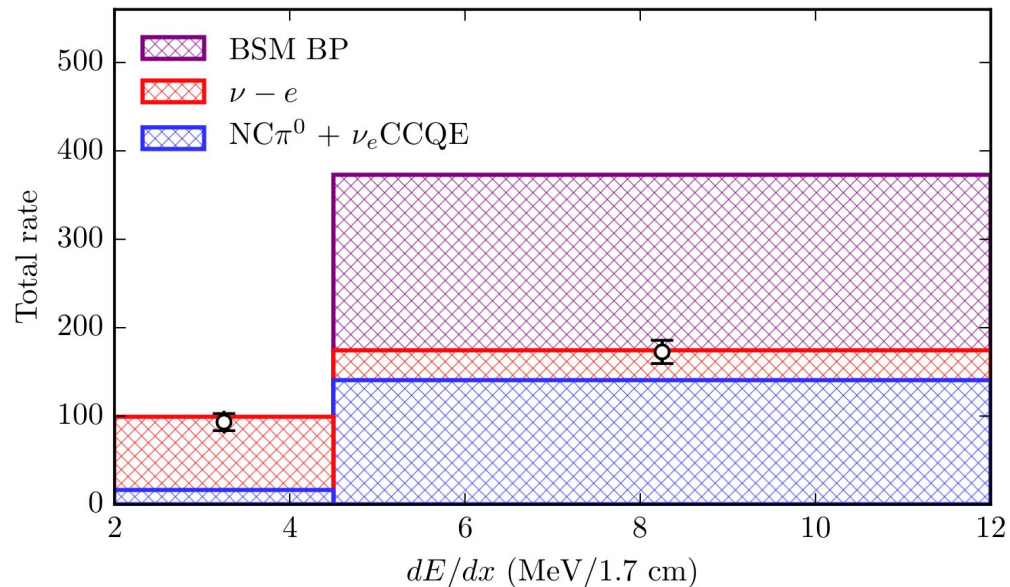
We use the following χ^2 definition:

$$\chi_{\alpha\beta}^2 = \frac{(N_{\text{data}} - (1 + \alpha + \beta)\mu_{\text{MC}}^{\text{BKG}} - (1 + \alpha)\mu_{\text{MC}}^{nu-e} - (1 + \alpha)\mu_{\text{BSM}})^2}{N_{\text{data}}} + \left(\frac{\alpha}{\sigma_{\alpha}}\right)^2 + \left(\frac{\beta}{\sigma_{\beta}}\right)^2$$

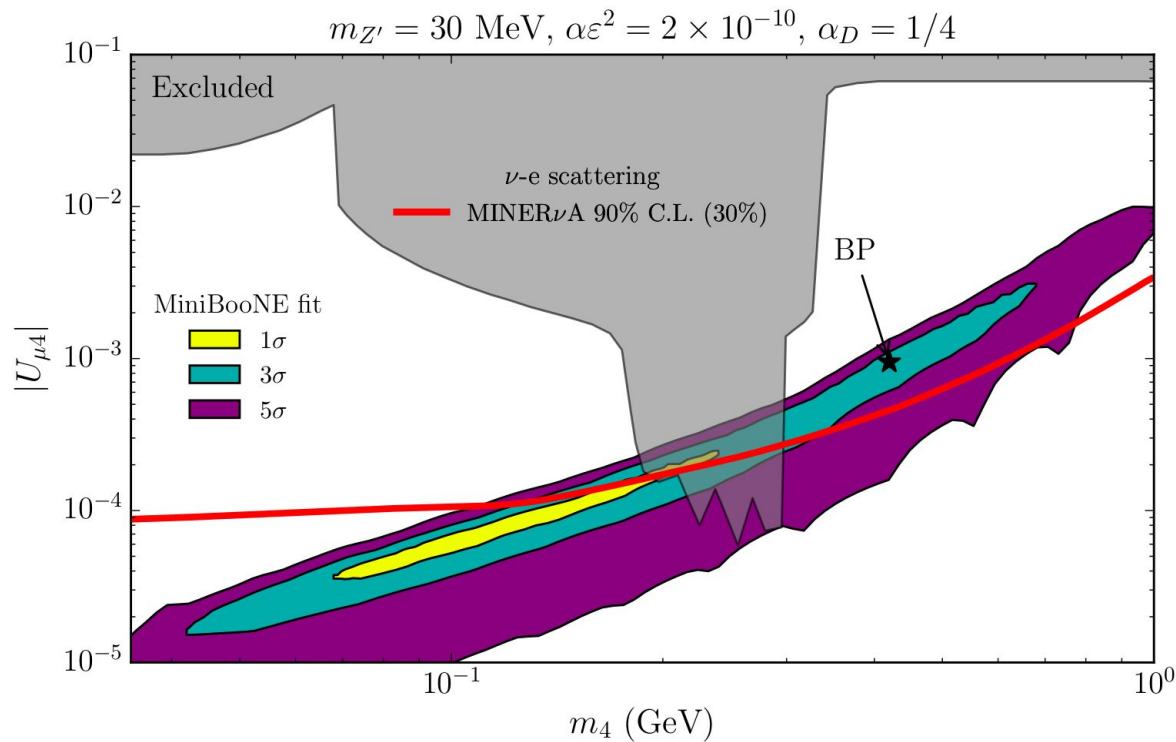
- We set $\sigma_{\alpha} = 10\%$ account for beam uncertainties.
- We set $\sigma_{\beta} = 30\%$ motivated by the amount of tuning; conservative with respect to tune normalization uncertainty.
- We include only coherent contribution to the BSM signal to avoid hadronic activity cuts.

MINERvA: Our Analysis setup

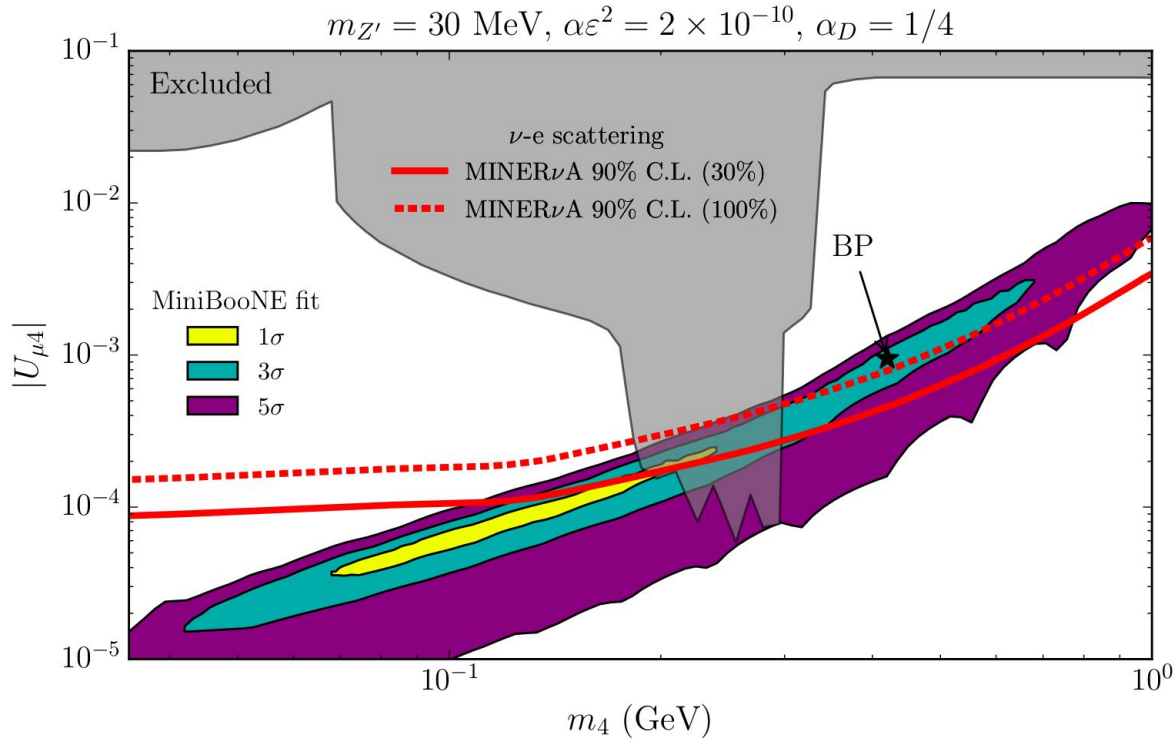
- We do a rate-only analysis on the single bin with $dE/dX > 4.5\text{MeV}/(1.7\text{cm})$
- We use 3.43×10^{20} POTs, Assume fiducial mass of 6.10 tons.



MINERvA result

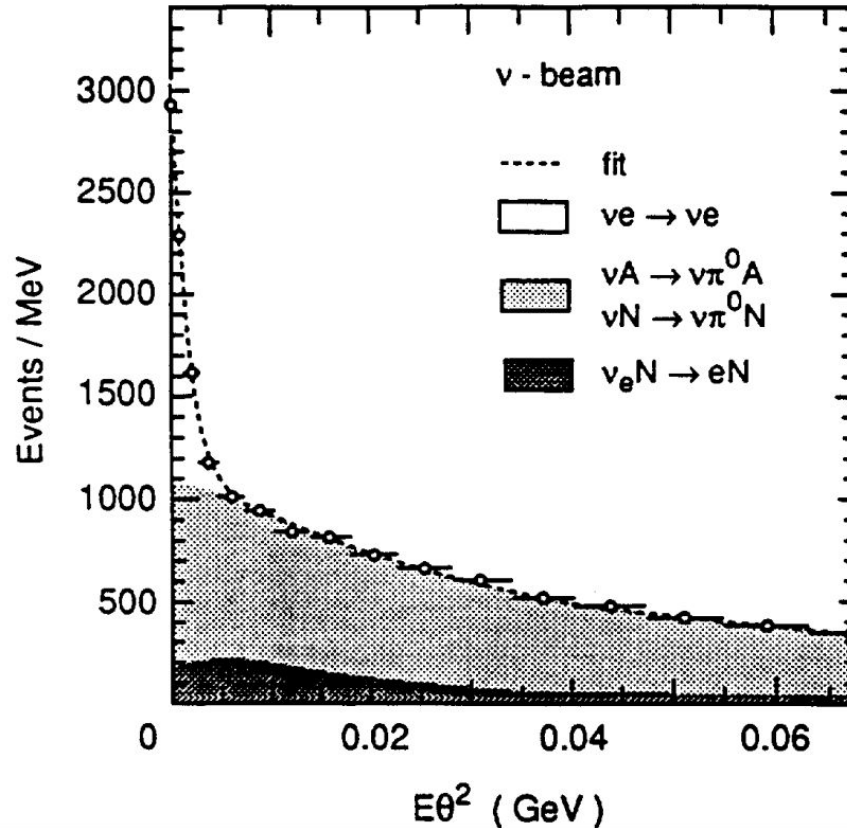


MINERvA result



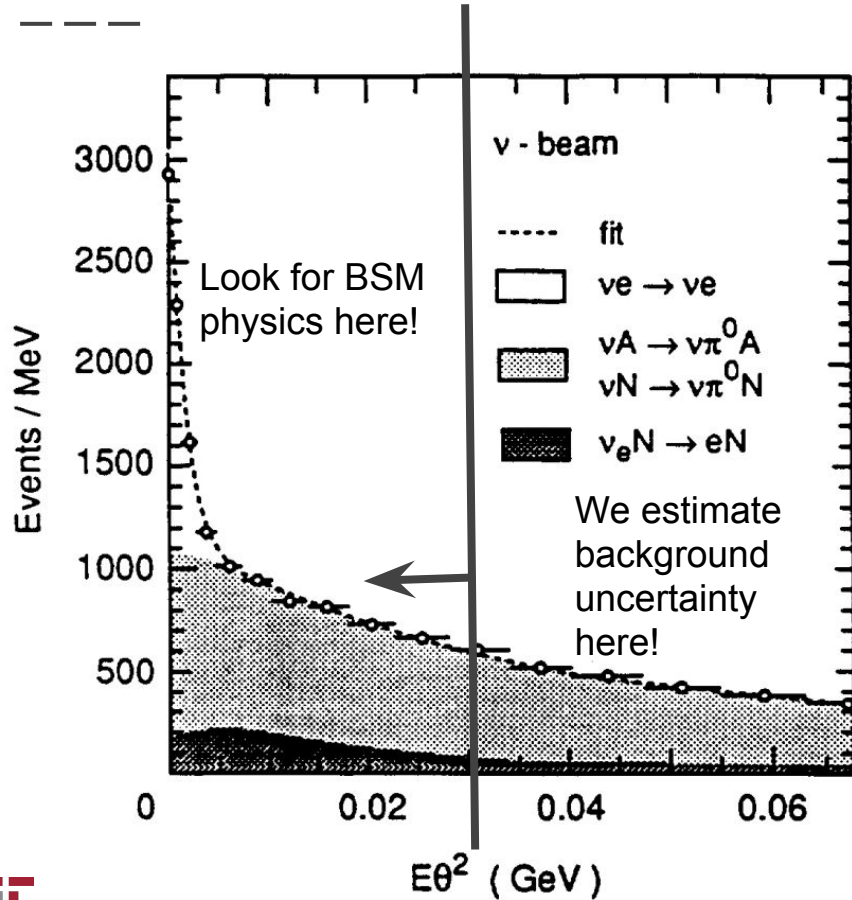
We checked that changing the background uncertainty from 30% to 100% changes the result by no more than a factor of two. The constraint power is coming from the BSM signal overshooting the data.

CHARM-II: complementary measurement

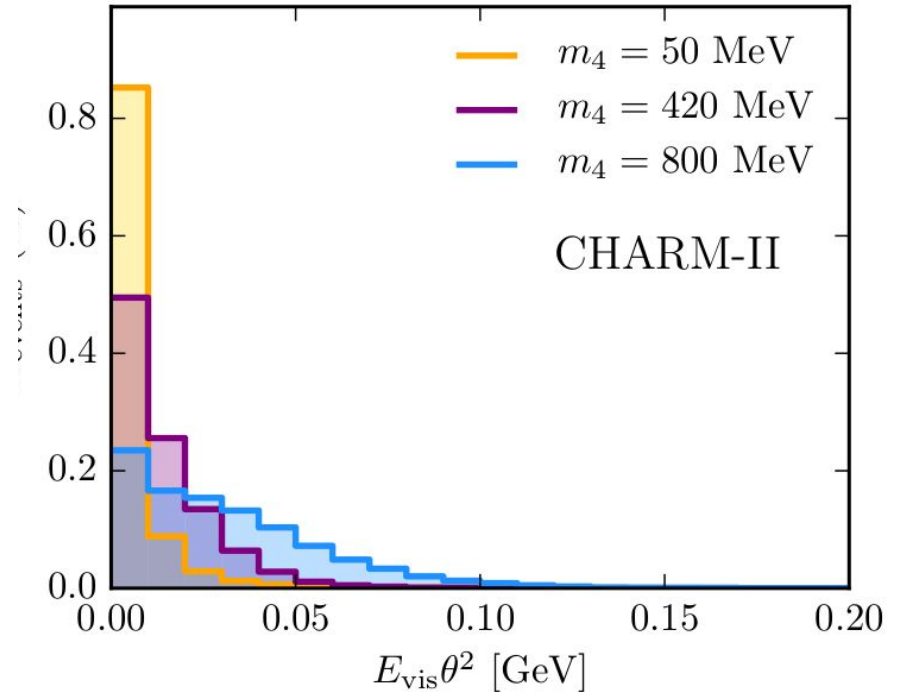


For CHARM-II we are going to use the $E\theta^2$ distribution before the final dE/dX cut is applied.

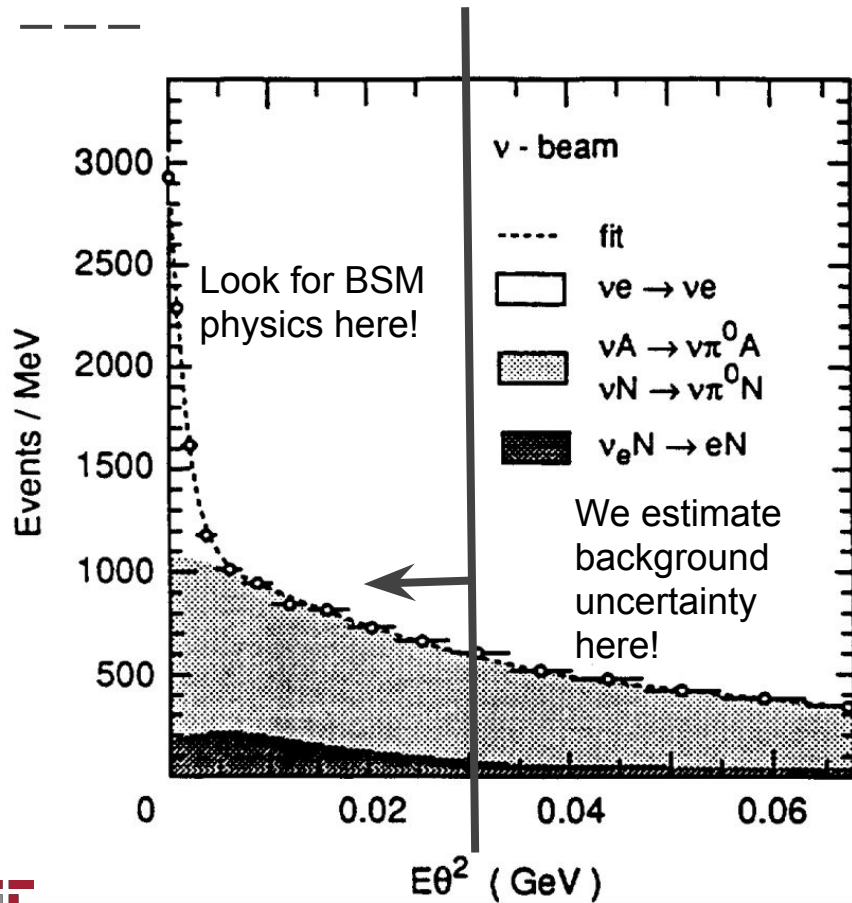
Finding “BSM-safe” sideband to measure background



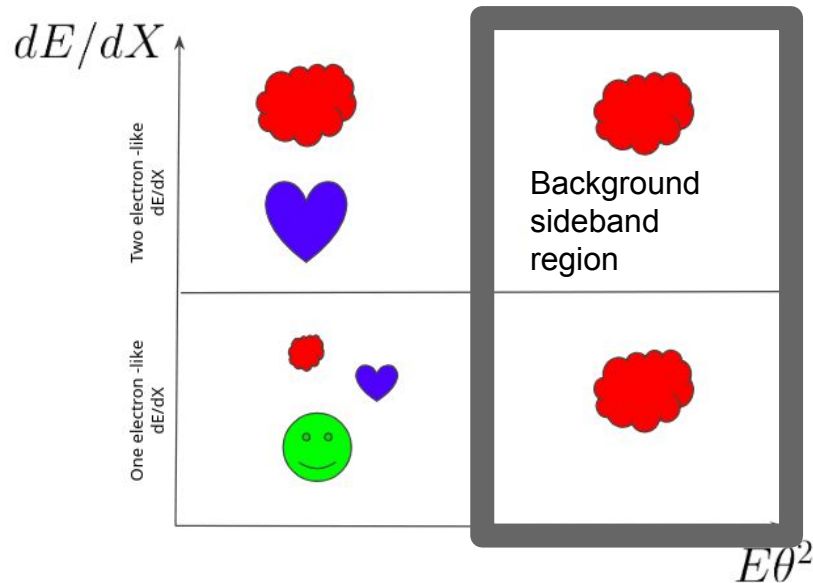
Angular distribution of BSM-signal



CHARM-II: complementary measurement



For CHARM-II we use the distribution before the angular cut and dE/dX were applied



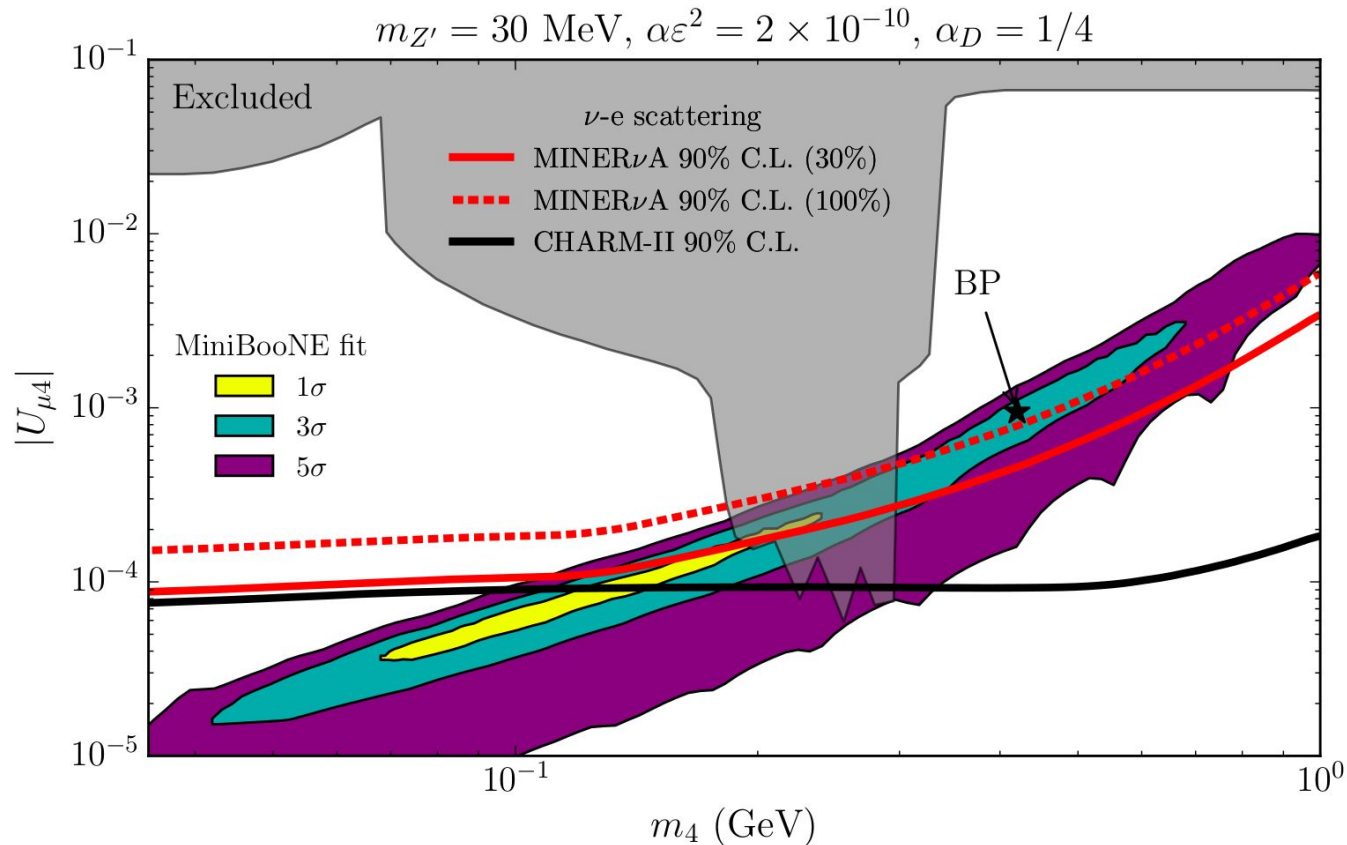
Use the region with $E\theta^2 > 0.03$ to obtain the background uncertainty.

Allow for rate/slope to change; with this we estimate its rate to be constrained to be $\sim 3\%$.

Our CHARM-II analysis setup details

- Rate-only analysis on a single bin with $E_{\theta}^2 < 0.03 \text{ GeV}$.
- Same χ^2 definition as in MINERvA, but updated uncertainties.
- Background norm. from sideband $\sim 3\%$; flux uncertainty $\sim 4\%$.
- We assume a fiducial mass of 547 tons, $\langle A \rangle \sim 20.7$, and $2.5e19 \text{ POT}$.

Putting it all together: the money plot



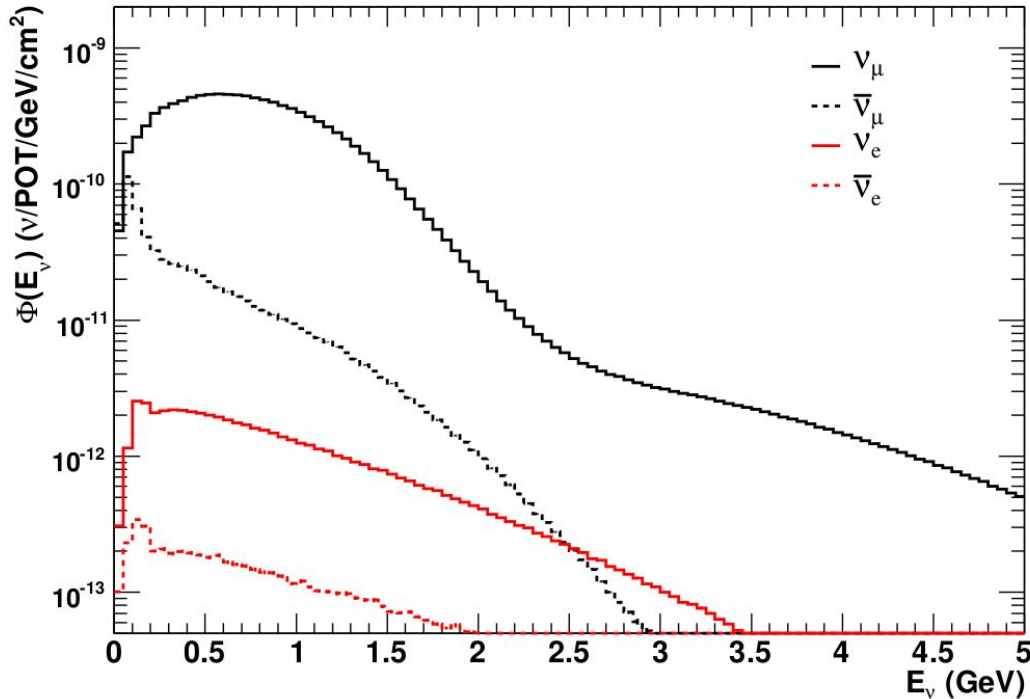
Take home message: lessons learned

- We are excited to see upcoming analyses by Minerva-ME and NOvA!
- We have used two different experiments to constrain recent MiniBooNE explanations. Tensions are large.
- Signals with small hadronic activity can be studied in conjunction with measurements of neutrino-electron scattering. See also M. Hostert (@PONDD-Tuesday), for similar claim for trident process searches.
- BSM-safe sidebands are critical if we are going to use tuned background predictions.

**Keep your eyes on the arXiv:
paper coming this week**

Bonus
slides!

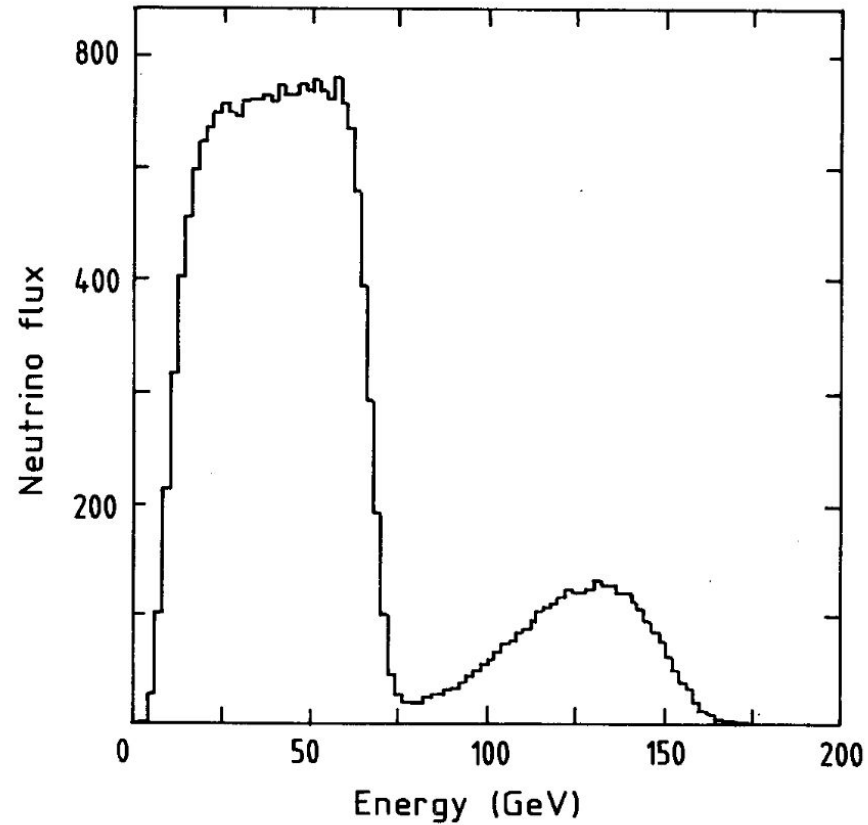
Booster beam primer



- Maxima < 1 GeV neutrino energy
- Production of heavy states via neutrino interaction: hard.
- Heavier BSM physics look like “effective”-interactions; then angular distribution of excess wont work.

Recent proposals:
Light new physics $\sim < \text{GeV}$.

CHARM-ii flux



CHARM-ii flux uncertainties

Systematic uncertainty of total cross-sections

Uncertainty	$\bar{\nu}$ (%)	ν (%)	$\bar{\nu}/\nu$ (%)
Absolute flux calibration (BCT, SSD combined)	2.0	2.0	-
Relative $\bar{\nu}/\nu$ flux	-	1.7	1.7
Dead-time	0.7	0.7	0.5
Target density	1.0	1.0	-
Event number systematic errors	1.7	0.2	1.7
Total systematic error	2.9	2.9	2.5