



Determination
of the LBNF
Neutrino Flux

Mary Bishai
(on behalf of
LBNF/DUNE)
Brookhaven
National Lab

The LBNF
Beamline

Flux Modeling
and
Uncertainties

Near
Detector(s)
Flux
Measurements

Muon
Monitors

Summary

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NuInt 2018, 15-19 October 2018, Gran Sasso Science Institute, Italy

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- 2 Flux Modeling and Uncertainties
- 3 Near Detector(s) Flux Measurements
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- 5 Summary

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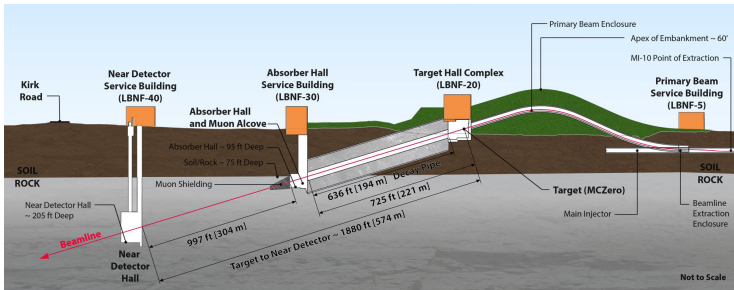
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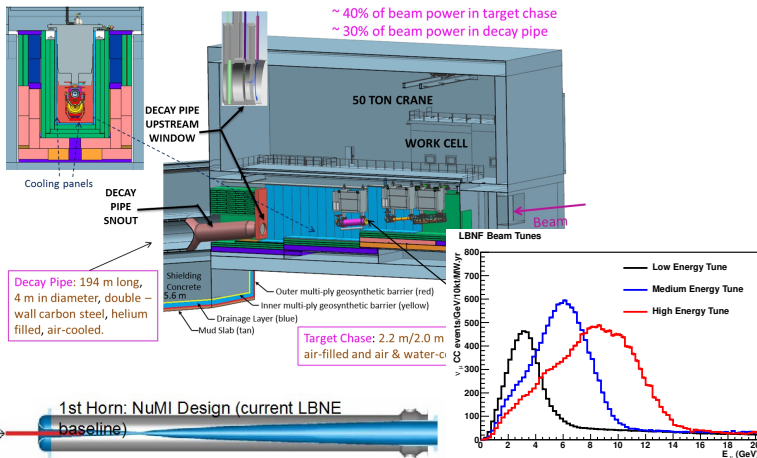
Summary



- Primary proton beam 60-120 GeV
- Initial 1.2 MW beam power, upgradable to 2.4 MW
- Embankment allows target complex to be at grade
- Wide-band beam (on-axis) based on NuMI design with tunable energy spectrum.
- Decay pipe: 194m x 4m diameter, He filled

ND default: 574 m from target, FD: 1300 km

Initial conceptual design was of a *tunable wide-band* NuMI-style focusing:



LBNF has switched to CPV optimized focusing design with 3 horns

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Optimization of flux for Physics: CP Violation

Laura Fields

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- The 2015 CD1R reference design for LBNF/DUNE is a NuMI-like movable target (segmented rectangular graphite fins with water cooling $\approx 1\text{m}$ long) and 2 modified NuMI horns 6.6m apart
- In Sep 2017 LBNF adopted a focusing design with 3 horns optimized using a *genetic algorithm* with the physics parameter to be measured (CPV sensitivity) used to gauge fitness.
- Target geometry is optimized at the same time, as well as proton beam energy with realistic Main Injector power profile (1.03 MW at 60 GeV to 1.2 MW at 120 GeV).
- Limits on horn current, diameter and length are imposed based on experience with T2K and NuMI horn manufacturing
- Limits on horn separation imposed based on size of target chase.

Optimized horn design with 297kA current :



Parameter	Value	Parameter	Value
Horn A Length (mm)	2218	Horn A F1 (% of length)	53
Horn A R1 (mm)	43	Horn A OC Radius (mm)	369
Horn A R2 (mm)	33		
Horn B Length (mm)	3932	Horn C Length (mm)	2184
Horn B R1 (mm)	159	Horn C R1 (mm)	284
Horn B R2 (mm)	81	Horn C R2 (mm)	131
Horn B R3 (mm)	225	Horn C R3 (mm)	362
Horn B F1 (% of length)	31	Horn C F1 (% of length)	20
Horn B F2 (% of length)	22	Horn C F2 (% of length)	9
Horn B F3 (% of length)	2	Horn C F3 (% of length)	7
Horn B F4 (% of length)	16	Horn C F4 (% of length)	35
Horn B OC Radius (mm)	634	Horn C OC Radius (mm)	634
Horn B Position (mm)	2956	Horn C Position (mm)	17806

Optimized target is 4λ (2m C) with $\sigma_{\text{beam}} = 2.7\text{mm}$, $E_p \sim 110\text{ GeV}$



Optimization of flux for Physics: CP Violation

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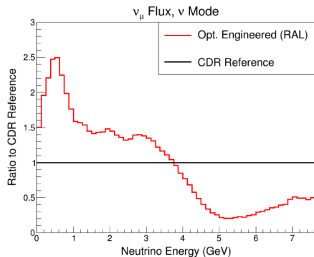
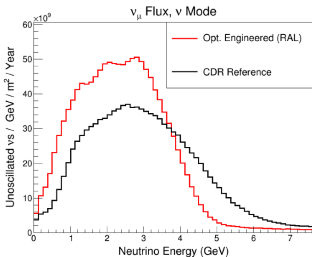
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Flux Modeling and Uncertainties

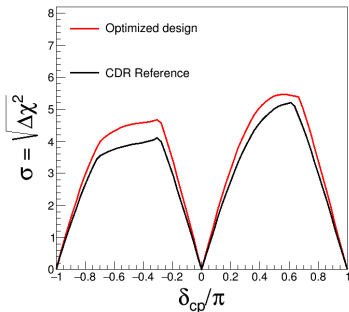
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CP violation sensitivity



Computationally advanced optimization techniques = significant gain in flux and CPV sensitivity from *many small changes*

Gain in sensitivity \equiv 70% increase in FD mass

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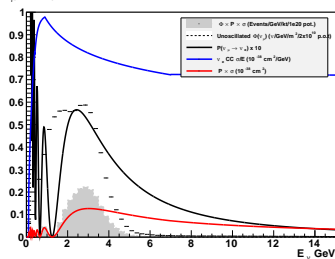
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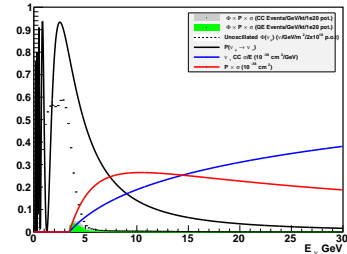
NuMI-like reference design could be tuned to higher energy to observe $\nu_\mu \rightarrow \nu_\tau$ with high statistics.

2015 two horn optimized design $E_p = 66$ GeV:

$\nu_\mu \rightarrow \nu_e$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_\tau$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_e$ 290 events

$\nu_\mu \rightarrow \nu_\tau$ 60 events

in 40 ktons, 1 year at 1.2 MW

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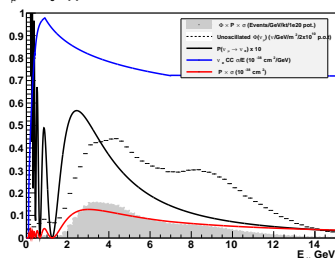
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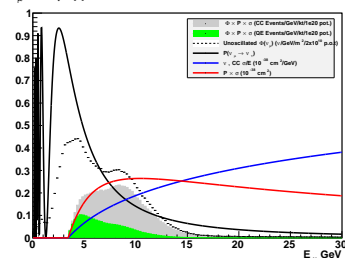
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LBNF target -2m from horn 1, NuMI focusing 230 kA, horns 17m apart

$\nu_\mu \rightarrow \nu_e$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_\tau$ Appearance at 1300 km



$\nu_\mu \rightarrow \nu_e$ 330 events $\nu_\mu \rightarrow \nu_\tau$ 700 events
in 40 ktons, 1 year at 1.2 MW

Increase ν_τ appearance 10x!!!

Increase high energy ν_e appearance - good for NSI/Sterile searches

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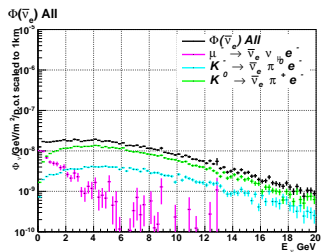
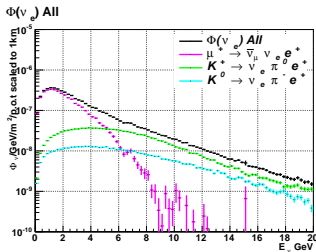
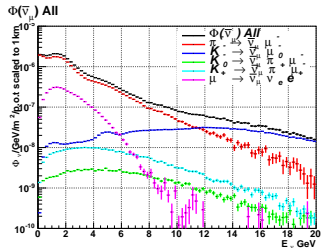
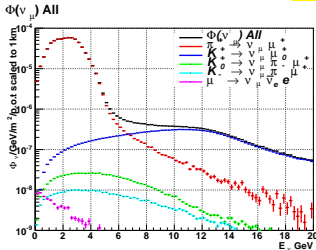
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ND 570



Baseline scaled to 1km from middle of decay channel

Flux components at near and far

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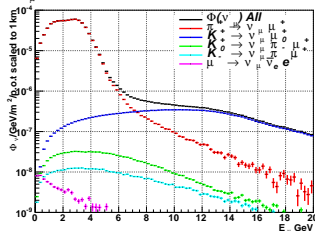
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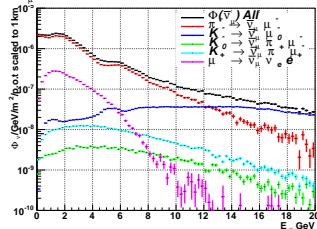
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FD 1300km

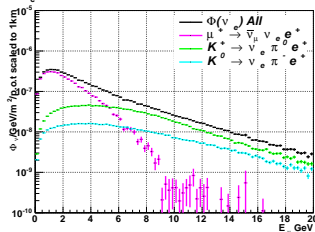
$\Phi(\nu_\mu)$ All



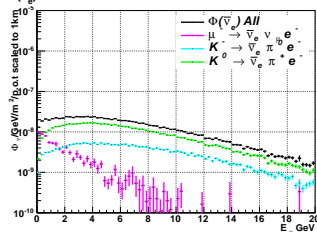
$\Phi(\bar{\nu}_\mu)$ All



$\Phi(\nu_e)$ All



$\Phi(\bar{\nu}_e)$ All



Baseline scaled to 1km from middle of decay channel

PPFX Package

- PPFX is a package that was **developed to correct the MINERvA flux simulation** to 'match' external hadron production data, e.g NA49:
- How the external data constraint **works in practice**:

- **Complete information** about cascades leading to a neutrino is recorded for each proton on target and stored in the flux tuples
 - Including interaction materials and ancestor kinematics
- In MINERvA analyses, **neutrino events are weighted** by:

$$w_{\text{HP}} = \frac{f_{\text{Data}}(x_F, p_T, E)}{f_{\text{MC}}(x_f, p_T, E)} \quad f = E \frac{d^3\sigma}{dp^3}$$

- Weights are applied for incident protons $12 < E_p < 120$ GeV/c, **scaled by Fluka** and checked by comparing to NA61 pC @ 31 GeV [Phys. Rev. C84 (2011)034604]
- Weights for events with **multiple interactions** in the ancestor chain are the product of the weight for each interaction

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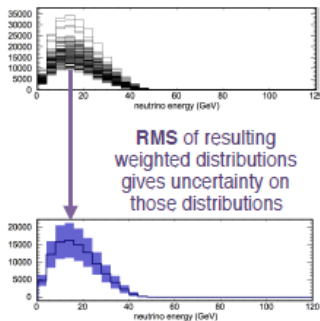
PPFX Package

- ✦ **Uncertainties** on the external data constraints are propagated to uncertainties on our flux and other simulated distributions using a “**Many-Universes**” method:
- ✦ For each event, in addition to the central value weights we have discussed:

$$w = e^{-L\rho(\sigma_{\text{DATA}} - \sigma_{\text{MC}})} \left(\prod_{\text{reweightable interactions}} \frac{f_{\text{Data}}(x_f, p_T, E)}{f_{\text{MC}}(x_f, p_T, E)} \right)$$

We also store many (~1000) weights constructed from **data cross sections varied according to their uncertainties** (taking into account correlations)

For interactions **uncovered by data, large (40%) are assumed**, correlated with neighboring bins



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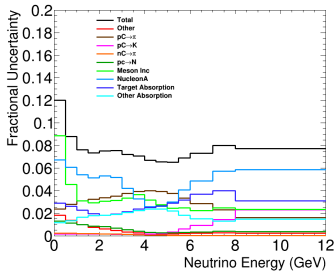
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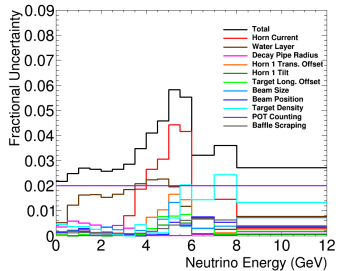


Hadron Prod. Uncertainties

NA49 dataset used to constrain



Pion production by neutrons from data (assuming isospin symmetry)
Nucleon incident interactions not covered by data



Focusing Uncertainties

Detailed focusing uncertainties based on the NuMI experience in MiNERνA.

Used CPV optimized design with simplified 2 horns.

Need to use engineered 3 horn design

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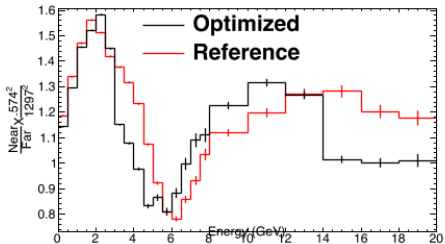
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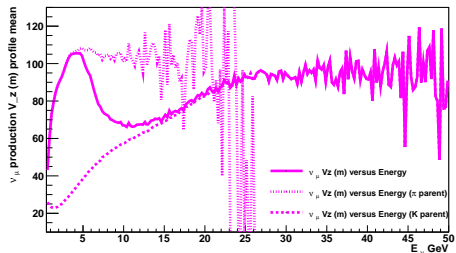
Simple ratio of far spectrum/near spectrum:

Neutrino parent decay location in decay pipe:

π/K decay kinematics and decay channel geometry are primary reason for strange shape of F/N ratio



ν_{μ} events at FD (1300km)



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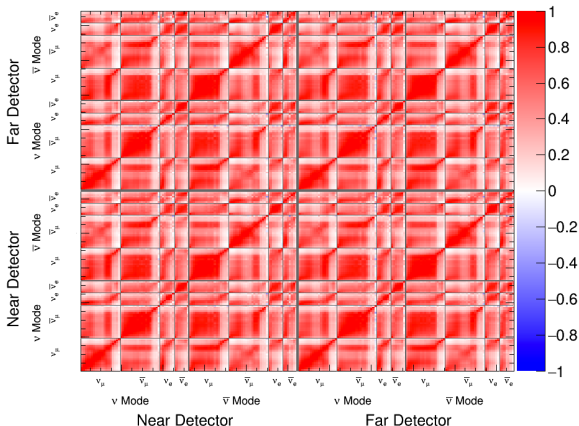
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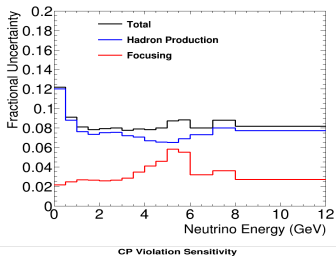
Summary

To correctly relate near to far fluxes - need to use a correlation matrix:

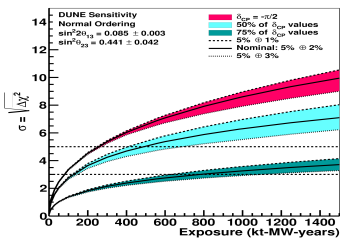
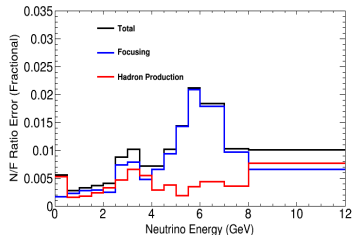


Flux correlation matrix comes from simulation and is highly correlated

Uncertainty on ND flux prediction



Residual uncertainty at FD



How well do we actually trust the simulation to correctly estimate the uncertainties on near \rightarrow far extrapolation?



Near Detector Flux Measurement Strategies

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Technique	Flavor	Absolute normalization	Relative flux $\Phi(E_\nu)$	Near Detector requirements
NC Scattering $\nu_\mu e^- \rightarrow \nu_\mu e^-$	ν_μ	2.5%	$\sim 5\%$	e^- ID θ_e Resolution e^-/e^+ Separation
Inverse muon decay $\nu_\mu e^- \rightarrow \mu^- \nu_e$	ν_μ	3%		μ^- ID θ_μ Resolution 2-Track ($\mu+X$) Resolution μ energy scale
CC QE $\nu_\mu n \rightarrow \mu^- p$ $Q^2 \rightarrow 0$	ν_μ	3 – 5%	5 – 10%	D target p Angular resolution p energy resolution Back-Subtraction
CC QE $\bar{\nu}_\mu p \rightarrow \mu^+ n$ $Q^2 \rightarrow 0$	$\bar{\nu}_\mu$	5%	10%	H target Back-Subtraction
Low-ν_0	ν_μ		2.0%	μ^- vs μ^+ E_μ -Scale Low- E_{Had} Resolution
Low-ν_0	$\bar{\nu}_\mu$		2.0%	μ^- vs μ^+ E_μ -Scale Low- E_{Had} Resolution
Low-ν_0	$\nu_e/\bar{\nu}_e$	1-3%	2.0%	e^-/e^+ Separation (K_L^0)
CC	ν_e/ν_μ	<1%	$\sim 2\%$	e^- ID & μ^- ID p_e/p_μ Resolution
CC	$\bar{\nu}_e/\bar{\nu}_\mu$	<1%	$\sim 2\%$	e^+ ID & μ^+ ID p_e/p_μ Resolution
Low-ν_0/CohPi	$\bar{\nu}_\mu/\nu_\mu$	$\sim 2\%$	$\sim 2\%$	μ^+ ID & μ^- ID p_μ Resolution E_{Had} Resolution

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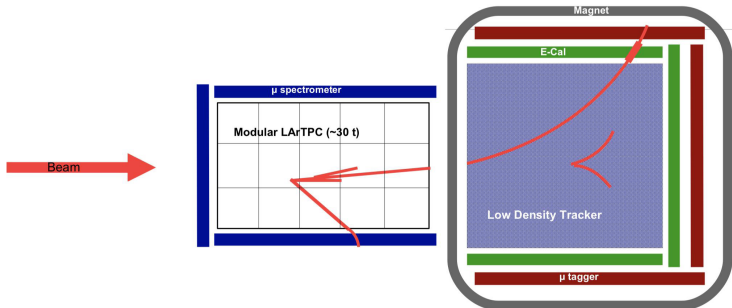
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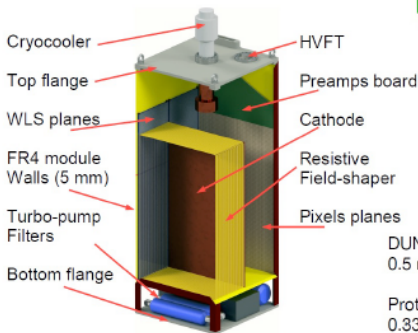
Summary

An unmagnetized LArTPC with a low density tracker embedded in a large magnet with EM sampling calorimeters surrounded by muon spectrometers

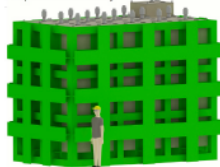


Expect 2.5×10^5 neutrino interactions/ton Ar/1e20 POT at 574m

ArgonCube

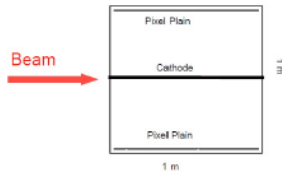


Proposed Geometry



CAD of the modular LArTPC for the DUNE ND complex

Module



DUNE ND modules: $1.0 \times 1.0 \times 2.0 \text{ m}^3$.
0.5 m drift length

Prototype modules $0.67 \times 0.67 \times 1.8 \text{ m}^3$.
0.33 m drift

Allows for high statistics measurements on Ar and ND/FD detector systematics constraints

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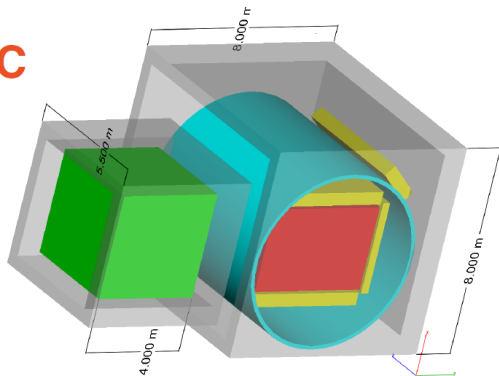
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HP GAr TPC

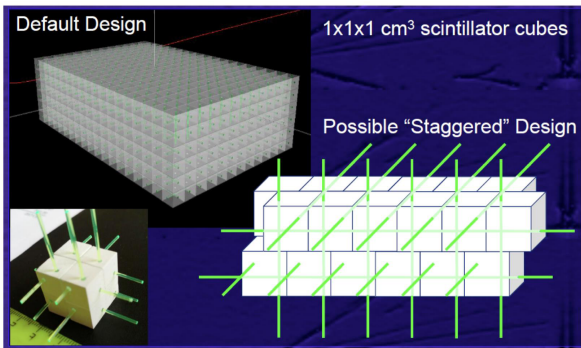
- Magnet (0.5T)
- LAr
- TPC_{hpg}
- EM Cal (20X₀)
- Steel (4λ₀)

LAr:
with 2 X 2 m FV
~ 7 X₀ annulus
~ 1.25 λ₀ annulus



GArTPC allows for precision measurement of ν -Ar vertex activity

3DST



Improved beam flux measurements (for example using low- ν method and CC QE) on a lighter target and when combined with GARTPC allows for measurement of $\pi^0\gamma\gamma$.

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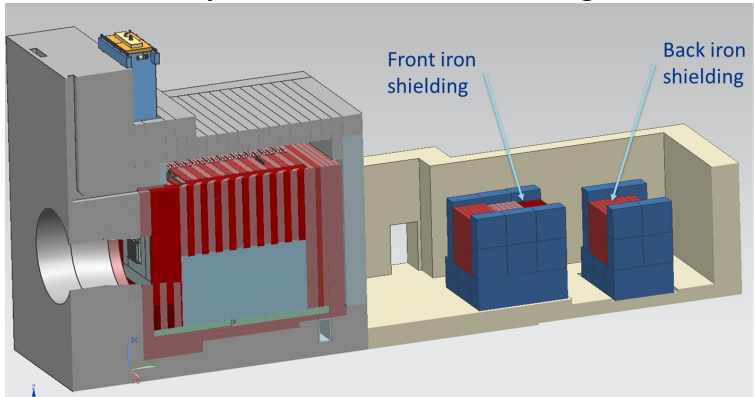
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Layout of Muon Alcove and Shielding



High intensity makes it difficult to measure μ spectrum accurately. With a 2.4 MW beam, the absorber thickness is too large to sample the lower energy muons. But these systems play an essential role in monitoring *flux stability*

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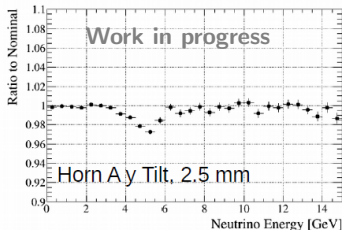
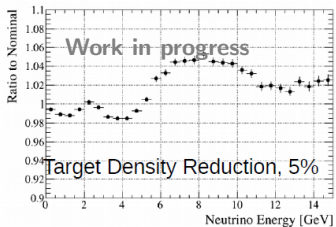
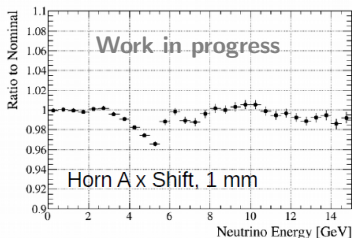
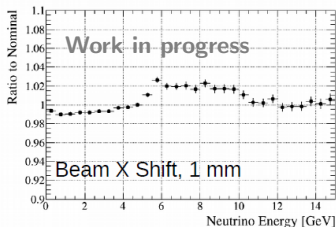
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ν Spectrum Changes



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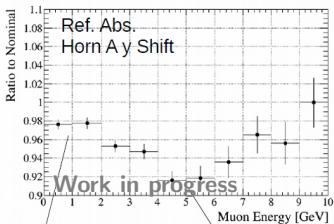
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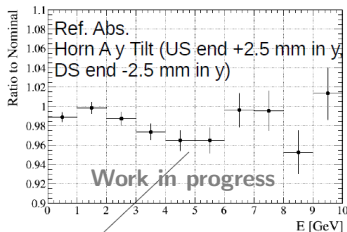
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μ Spectrum Changes



Reduction in total flux

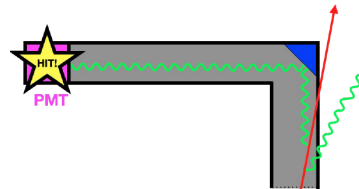
Shape changes at max near 5 GeV



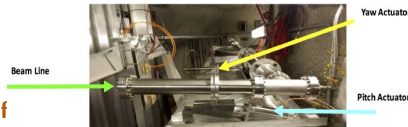
Changes are v. small - need novel detector concepts

- **Array of ionization detectors:** Measures muon beam center and intensity. Spill by spill monitoring of beam stability. Both diamond and silicon under study
- **Threshold gas Cherenkov detector (R&D):** Uses signal intensity at different gas pressure and angles to extract rough muon spectrum.
- **Stopped muon counters (R&D):** separate stations with steel shielding in between could measure muon flux at several energies. Better measurement of beam flux spectrum and composition.

Gas Cherenkov counter concept:



Prototype in NuMI beamline:



Currently only ionization detectors included in the beam design.

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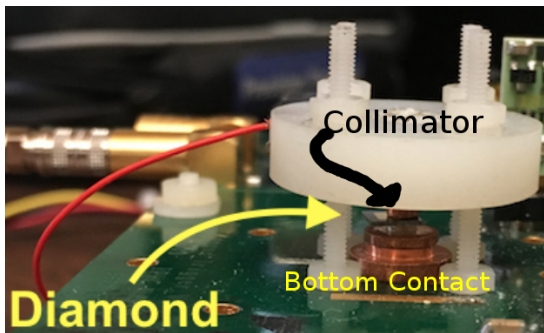
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Flux
Measurements

Muon
Monitors

Summary



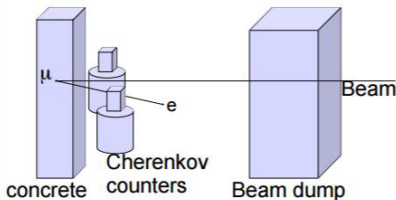
Use polycrystal chemical vapor deposition (pCVD) diamond - detects ionizing radiation when a large voltage potential (1V per μm of thickness) is applied across two sides of the diamond. Diamond is radiation hard.

pCVD detector prototype installed in NuMI during 2018 shutdown.

From K. Hiraide, *Muon monitor using the decay electrons*, NBI2003 Workshop



Strategy

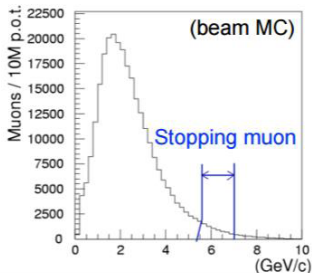


- Counting the decay electrons from muons stopping at the wall of μ -pit
- Measuring spatial and time distributions of events

- Energy loss of muons in the beam dump
- Range of electrons in the concrete



We can measure muons of
5.2~7.0 GeV/c
 by counting the decay electrons



Stopped Muon Prototype

Determination of the LBNF Neutrino Flux

Mary Bishai
(on behalf of LBNF/DUNE)
Brookhaven National Lab

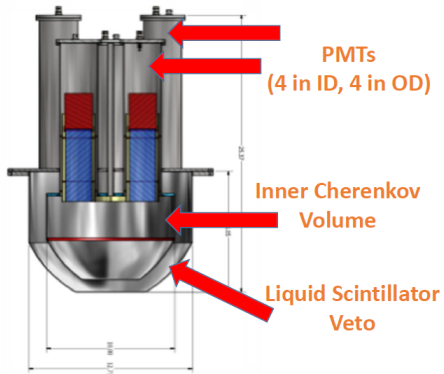
The LBNF Beamline

Flux Modeling and Uncertainties

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Prototypes being commissioned with cosmics.



Summary and Conclusions

Determination
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- The next generation of long-baseline experiments requires determination of the flux at the **1-2% level**.
- LBNF is a wide-band beam with tunable capabilities \Rightarrow requires beam flux/spectrum measurements over a large range of energies.
- Very high intensity beams (MW class) are needed \Rightarrow challenging near detector designs. Difficult to keep the same technology near and far. **For LBNF/DUNE the near detector could be a gas tracker using an Ar target to match the liquid Argon TPC far detector**
- Focusing uncertainties dominate the residual uncertainties in the near to far extrapolation at LBNF/DUNE \Rightarrow determination of the hadron production from the target *is necessary but not sufficient* for a-priori calculations of the neutrino flux. **Do we need a spectrometer following the horns?**
- **Measurements of the muon flux after the absorber is difficult but necessary to monitor the tertiary beam stability. R&D is ongoing on new technologies.**