THE DUNE NEAR DETECTOR **REFERENCE DESIGN**

H. A. Tanaka









INTRODUCTION

- Agenda for the review
- Introduction to LBNF/DUNE
- The Overarching Requirements
- The PRISM Concept
- The DUNE Near Detector Reference design
- Subsystem overviews (LAr, MPD, 3DST-S)
- Towards the CDR and ND Requirements



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AGENDA

0800 - 0900	Executive Session		
0900 - 1000	DUNE ND Requirements and Reference Overview 40+ 20 (Tanaka)		
1000 - 1015	Break		
1015 - 1100	Constraining interaction model uncertainty	30 + 15 (Wilkinson)	
1100 - 1145	Long-baseline physics analysis status overview	30 + 15 (Worcester)	
1145 - 1230	20 DUNE-PRISM capability for determine energy response 30 + 15 (Wilking)		
1230	Committee gets lunch from Director's CR and returns		
1245 - 1330	Executive Session over lunch		
1330 - 1400	Direct flux measurements with neutrino-elastic scattering	20 + 10 (Marshall)	
	DUNE Near Detector subsystem details		
1400 - 1430	LAr	20 + 10 (Sinclair)	
1430 - 1445	Break		
1445 - 1515	MPD	20 + 10 (Bross)	
1515 - 1545	3DST-S	20 + 10 (Manly)	
1545 - 1645	Executive Session		
1645 - 1715	Close out		



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GUIDANCE FROM THE COMMITTEE

- DUNE should provide the ND review committee with an "existence proof" of a plausible and achievable ND design that will meet the requirements set in the physics TDR
 - This need not be an optimized design, but enough to permit sign-off of on the physics TDR
- The review committee will advise the LBNC on whether the ND concept is feasible and appropriate
- The review committee will provide early constructive feedback to DUNE, as DUNE prepares to complete the ND CDR
- From the DUNE collaboration:
 - We sincerely thank the committee for its in-depth review of the ND "Executive Summary" document and for the insightful and probing questions.
 - Our submitted document is our primary response but we will return to key points through the presentations

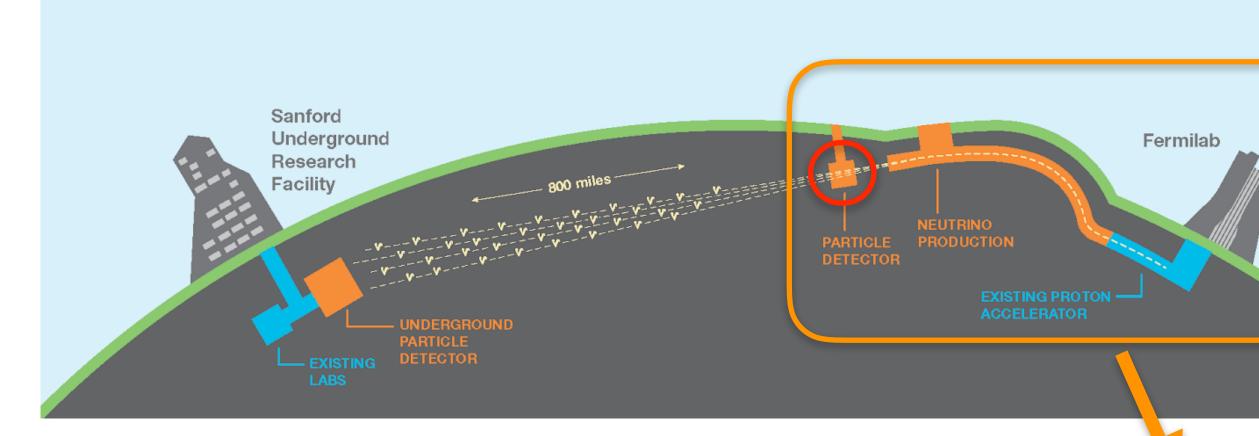




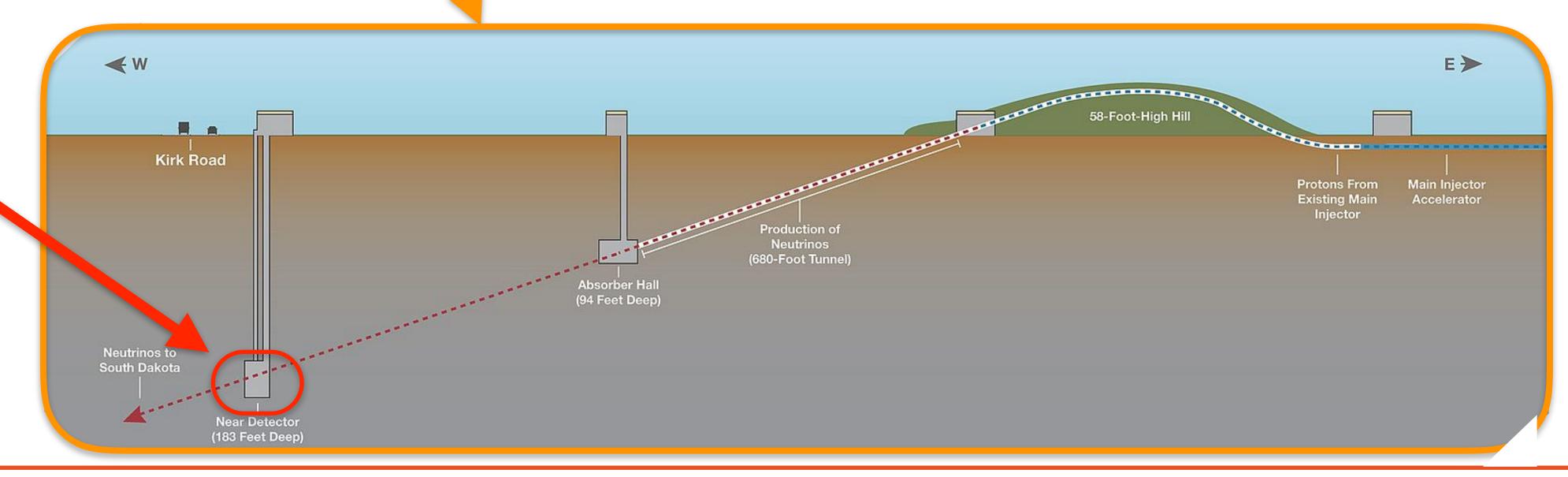




DUNE/LBNF AND THE NEAR DETECTOR



 At 1.2 MW, ~0.16 events/ton/spill at the near site





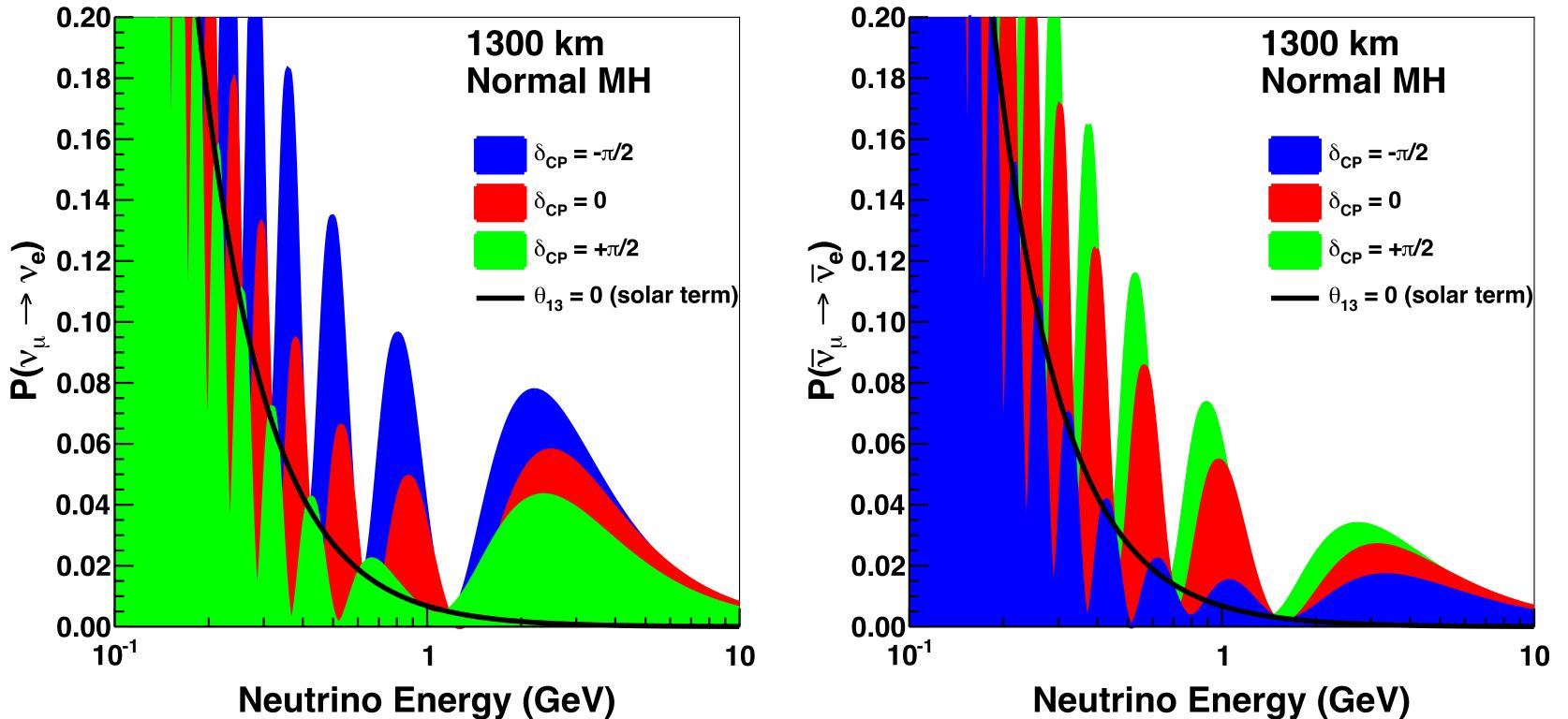


- LBNF sends an intense, broadband ν_μ = (ν_μ) beam 1280 km from FNAL to the DUNE far detector (FD) at SURF
- Here, we talk about measurements at the near site, 575 m from the production target





NEUTRINO OSCILLATIONS AT DUNE



- events observed at the far detector
- Fully exploit broad band LBNF neutrino beam to measure oscillation probability vs. E_{v}



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more in E. Worcester's talk

• Neutrino oscillation parameters are extracted via the rate and energy spectrum of $v_{\mu}(\bar{v}_{\mu})$, $v_e(\bar{v}_e)$





ROLE OF NEAR DETECTOR

- long baseline analysis.

$$N_{FD}(\nu_{\alpha} \to \nu_{\beta}, E_{rec}) = \int dE_{\nu} \quad \Phi_{\nu_{\alpha}}(E_{\nu}) \quad \sigma_{\nu_{\beta}}(E_{\nu}) \quad R_{\nu_{\beta}}(E_{\nu}, E_{rec}) \quad P(\nu_{\alpha} \to \nu_{\beta}, E_{\nu})$$

which involves

- Φ : the initial neutrino flux •
- σ : the neutrino interaction cross section
- **R**: the response, encapsulating how a neutrino with energy E_v is reconstructed in the detector (E_{rec})
 - incorporates both the modelling of the *v*-Ar interaction final state and the detector response
- oscillation parameters with sufficient precision.







• The primary role of the near detector (ND) is to serve to serve as the experiment's control for the

• It must provide enough information to sufficiently model each component of the following convolution

such that the number and spectrum of v_{μ} , \bar{v}_{μ} , v_{e} , \bar{v}_{e} at the FD can be predicted as a function of the





BEYOND THE MODELS:

- The neutrino oscillation analysis is built on models of:
 - The neutrino flux using Monte Carlo simulation tuned with external data (e.g. hadron production @ NA61) -
 - Neutrino interactions using a generator incorporating nuclear physics and tuned to historic data
 - The detectors using full Geant4 simulations tuned with calibrations and other data (e.g. ProtoDUNE)
- This foundation is known not to be solid:
 - We know that the neutrino interaction models are deficient, particularly with regard to the nuclear physics Current data cannot be fully explained by existing models

 - Recent "surprises" in qualitatively new effects (e.g. multinucleon interactions) not previously considered
 - etc. . . . there are bound to be more surprises
 - Neutrino beams are based on very well known physics, but are complex and operate in harsh environments careful monitoring and precise cross checks on the neutrino flux are critical
- DUNE ND must be designed to address these issues





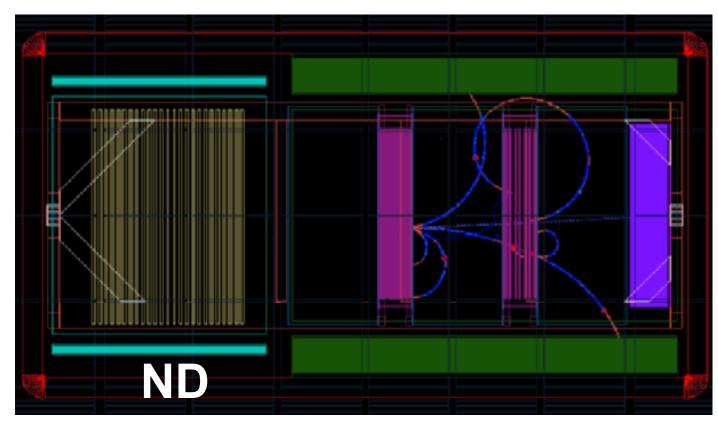


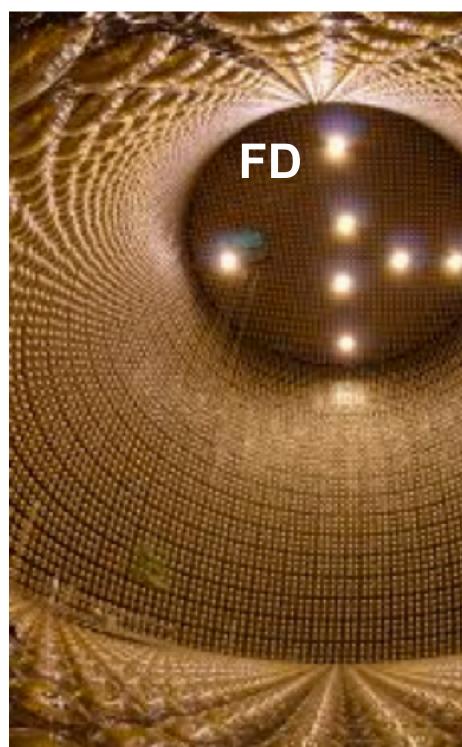


ONGOING EXPERIENCE

• T2K:

- Near detectors with capabilities beyond those of the far detectors can further inform flux, neutrino interaction modeling, etc.
- subsystems have specialized roles





• DUNE Approach

- -

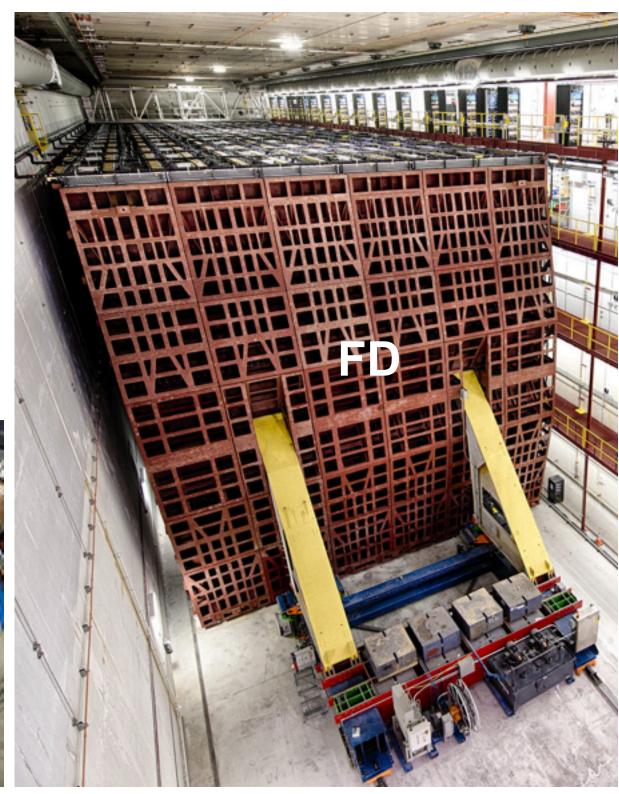


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- NOvA
 - Near detector that is functional identical to far detector can directly inform far detector distributions and cancel systematic errors





Near detectors with capabilities beyond the far detector are needed to sufficiently constrain uncertainties A detector functionally very similar to the FD is needed to inform how the interactions will actually appear there





DRAFT OVERARCHING REQUIREMENTS* (I)

O0 Predict the neutrino spectrum at the FD: The ND must provide a prediction for the energy spectrum of v_{μ} , \bar{v}_{μ} , v_e , and \bar{v}_e at the FD. The prediction must be provided as a function of the oscillation parameters and systematic uncertainties must be small enough to achieve the required CP coverage. This is the primary requirement of the DUNE ND.

00.1	Measure interactions on argon	Me nuc kine
O0.2	Measure the neutrino energy	Re bia acł be
O0.3	Constrain the cross section model	Me sec cau true ach



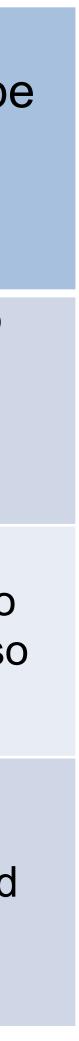


easure neutrino interactions on argon to reduce uncertainties due to clear modeling, determine the neutrino flavor, and measure the full nematic range of the interactions that will be seen at the FD.

econstruct the neutrino energy in CC events and control for any ases in energy scale or resolution, keeping them small enough to hieve the required CP coverage. These measurements must also transferable to the FD

easure neutrino cross-sections in order to constrain the crossction model used in the oscillation analysis. Mismodeling that uses incorrect FD predictions as a function of neutrino flavor and le or reconstructed energy must be constrained well enough to hieve the required CP coverage.

* Requires DUNE Executive Board sign off





DRAFT OVERARCHING REQUIRMENTS* (II)

O0.4	Measure neutrino flux
O0.5	Obtain data with different neutrino fluxes
O0.6	Monitor the neutrino beam





Measure neutrino fluxes as a function of flavor and neutrino energy. This allows for neutrino cross-section measurements to be made and constrains the beam model and the extrapolation of neutrino energy spectra from the ND to the FD.

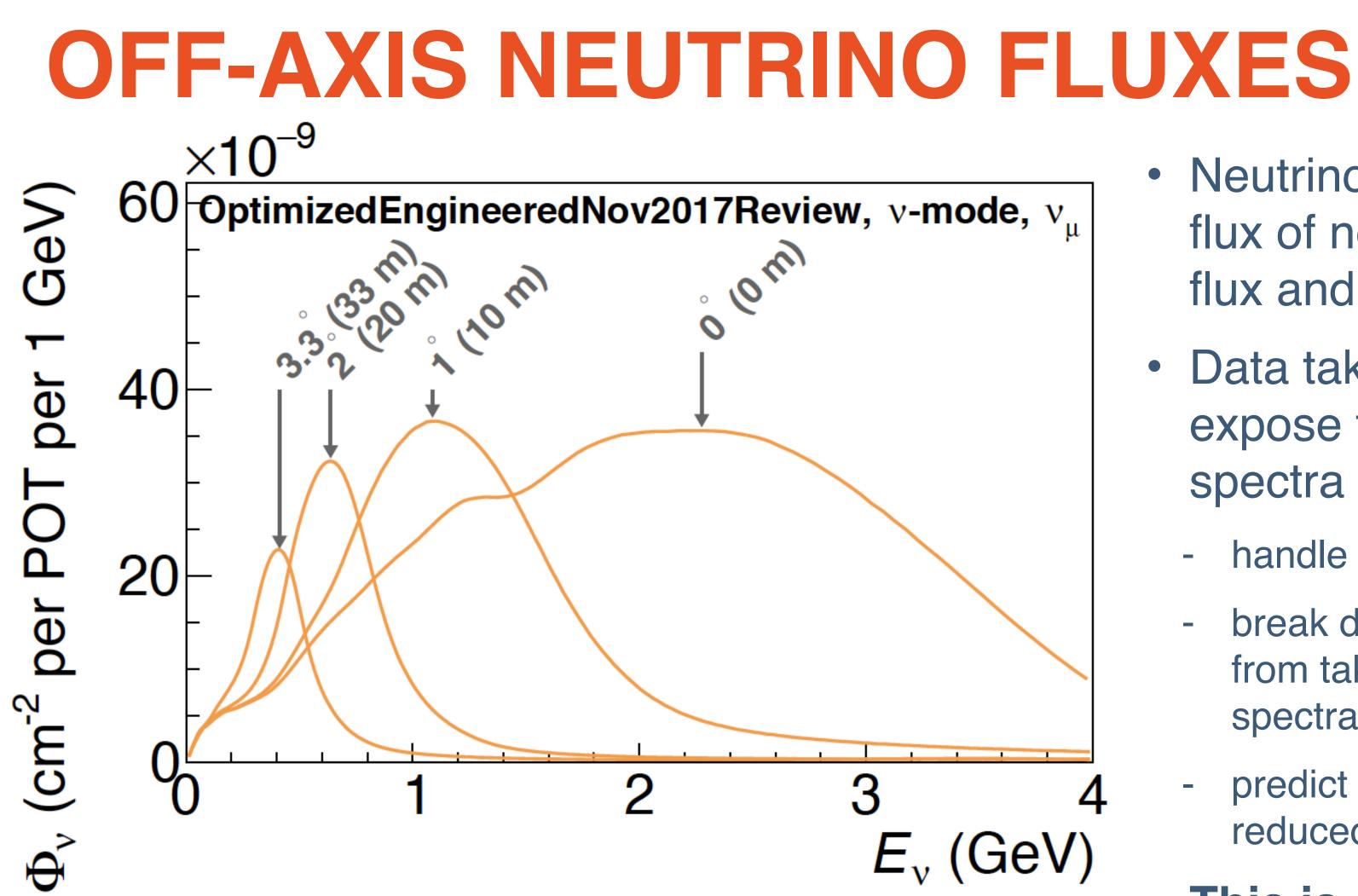
Measure neutrino interactions in different beam fluxes (especially ones with different mean energies) to disentangle flux and crosssections, verify the beam model, and guard against systematic uncertainties on the neutrino energy reconstruction.

Monitor the neutrino beam energy spectrum with sufficient statistics to be sensitive to intentional or accidental changes in the beam on short timescales. The precise requirement will be informed by the run plan as well as experience from previous experiments.

* Requires DUNE Executive Board sign off







more in M. Wilking's talk





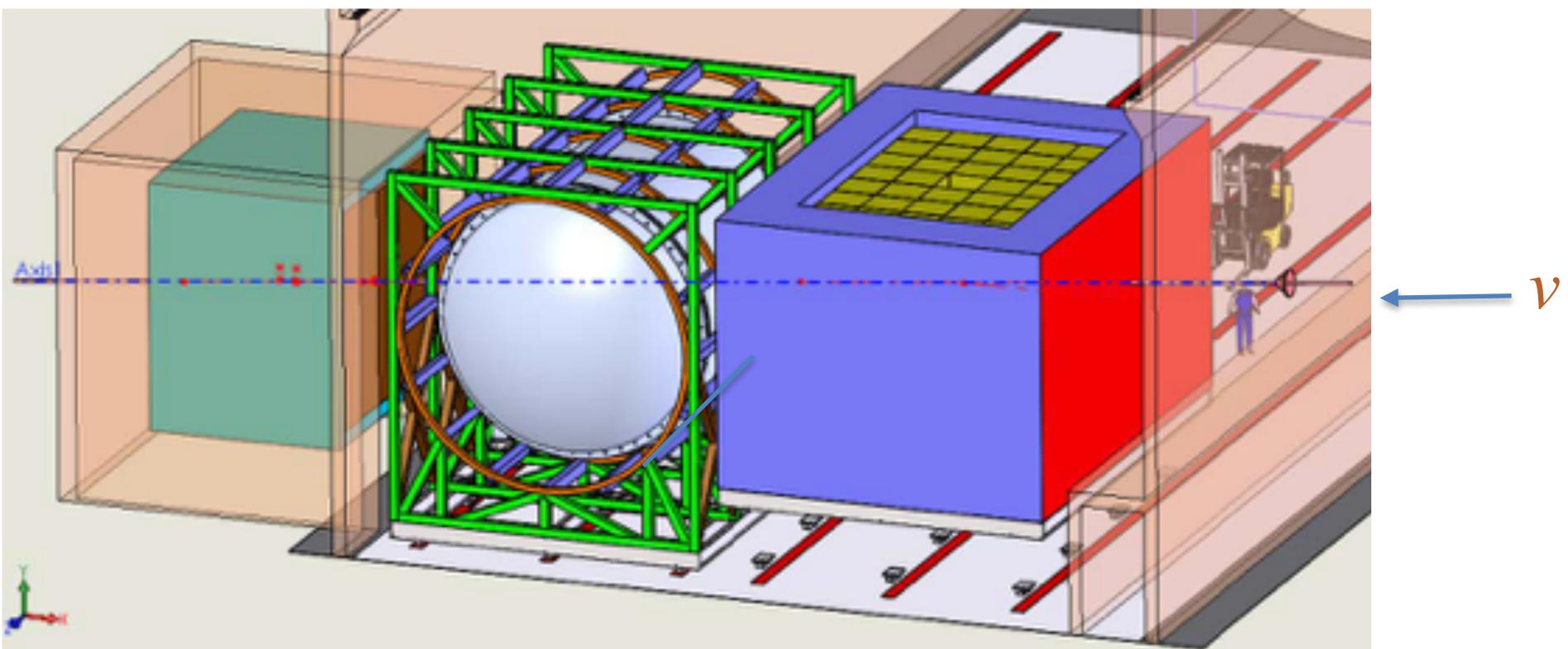
 Neutrino interactions observed in a fixed flux of neutrinos are a convolution of the flux and cross section

- Data taken at varying off-axis positions expose the detector to fluxes with different spectra ("PRISM")
 - handle to deconvolve the flux and cross section
 - break degeneracies in modelling that may result from taking data only on-axis (one off-axis position/ spectra)
 - predict oscillated neutrino event spectra at FD with reduced model dependence
- This is a qualitatively new element of ND design that we will introduce to DUNE



DUNE

DETECTOR OVERVIEW



- From left-to-right (upstream-to-downstream)
 - LAr: Liquid Argon Time Projection Chambers -
 - MPD: Multi-Purpose Detector
 - 3DST-S: Three Dimensional Scintillating Tracker-Spectrometer -







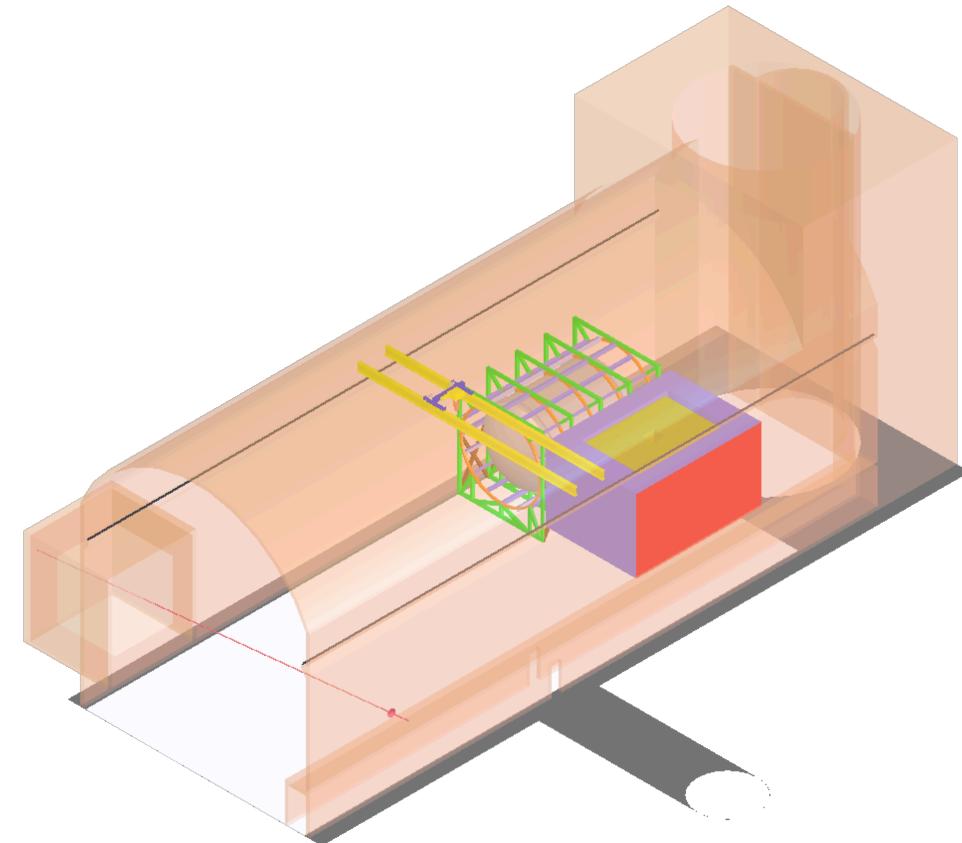


THE NEAR DETECTOR HALL

- 3DST-S remains on-axis as a beam monitor







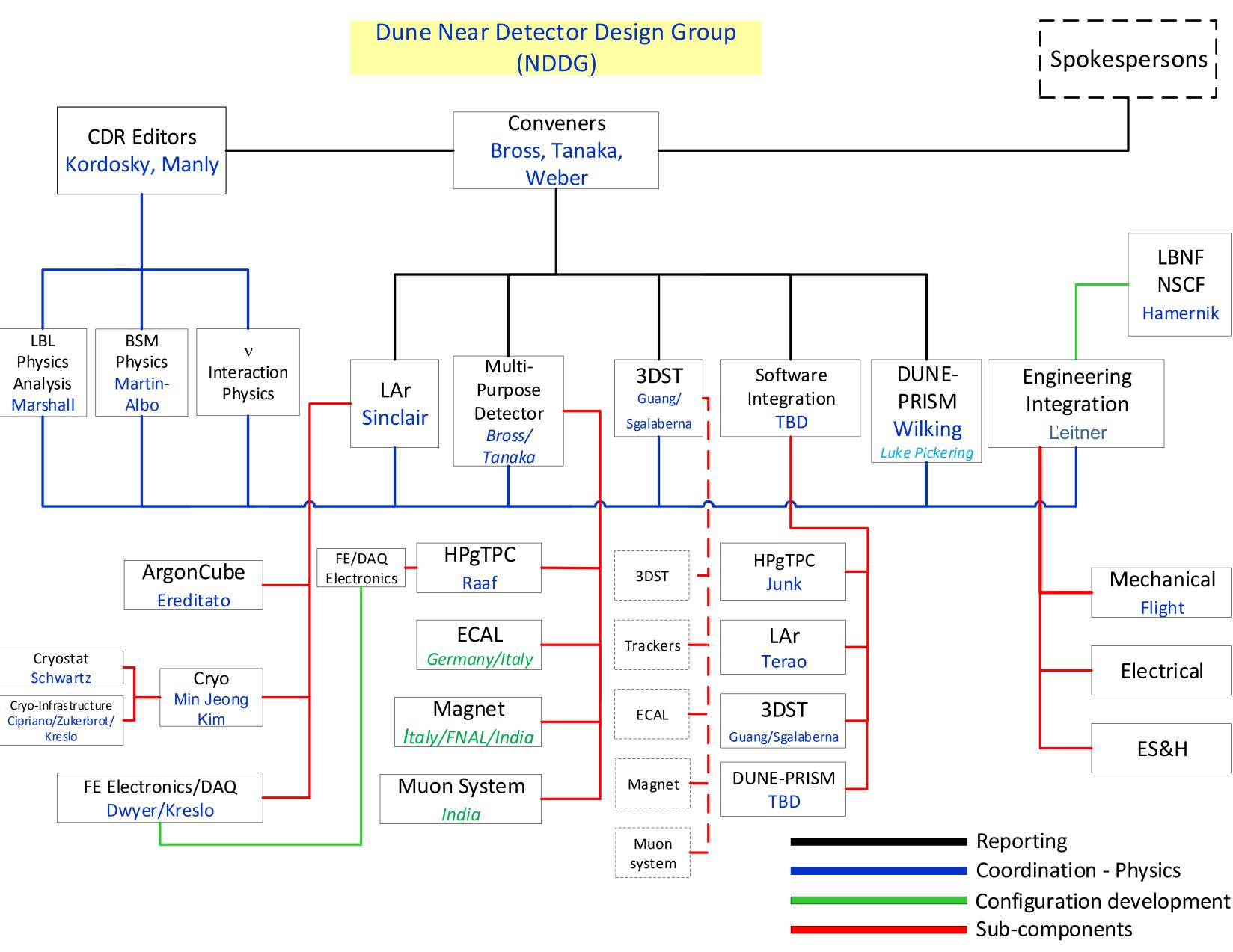
• LAr + MPD move transverse to the beam to effect DUNE-PRISM (*i.e.* exposure to off-axis fluxes)

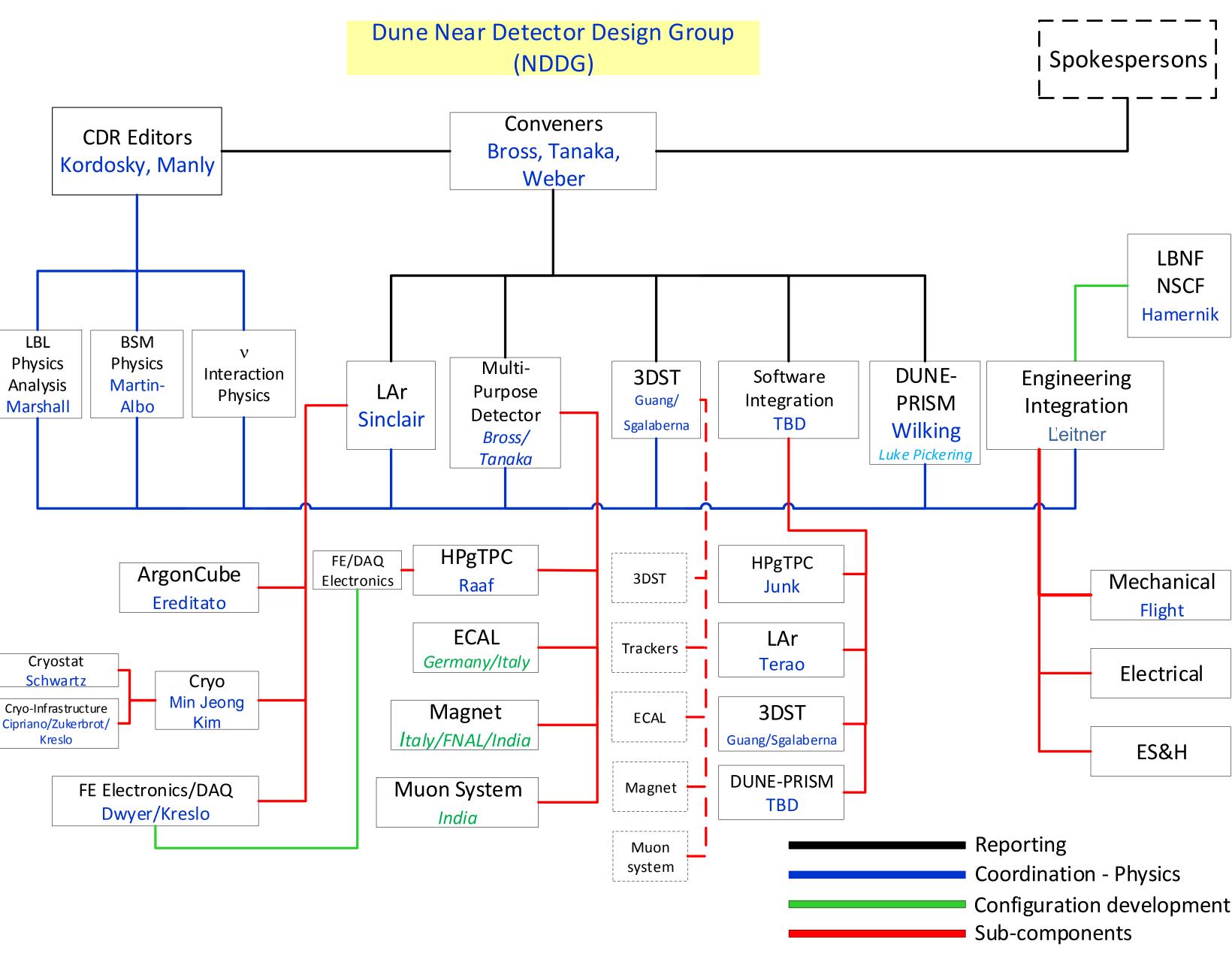




ORGANIZATION

- Near Detector Design Group formally launched in 9/2018
- Biweekly general meetings
- Individual subgroup meetings
- Regular engineering/facility interface and integration meetings









DUNE

MATRIX OF SUBCOMPONENT ROLES

Complementarity in:

- Event reconstruction:
 - MPD necessary to complete reconstruction of events in LAr detector
 - ECAL necessary to complete reconstruction of HPgTPC interactions (like collider detector)
- Viewing of *v*-Ar interactions
 - MPD provides look at "bare" *v*-Ar interaction with sign selection, very low thresholds, and minimal secondary interactions
 - LAr provides look at "full" v-Ar interaction as seen by FD
- Capability and Robustness:
 - 3DST-S provides look at *v*-CH interactions with novel neutron detection capabilities
 - 3DST-S provides detailed on-axis beam monitoring



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MF

	Essential Characteristics	Primary Function	Select Phys Aims
٨r	Mass	Experimental control for FD	$ u_{\mu}\left(\overline{\nu}_{\mu}\right)CC$
	Target nucleus (Ar)	Measure unoscillated Evspectra	v-e scattering
	Technology: FD-like	Flux determination	$v_e + \overline{v}_e \ CC$
			v interaction mo
PD	Magnetic field	Experimental control for LArTPCs	$v_{\mu}\left(\overline{v}_{\mu} ight)$ CC
	Target nucleus (Ar)	Momentum analyze µ from LAr	$v_e\left(\overline{v}_e\right)$ CC
	Low density		v interaction mo
	Hermeticity	with low threshold, minimal secondary interactions	
)ST-S	On axis	Neutrino beam monitor	On-axis flux sta
	Mass	Neutron detection	v interaction models dependence
	Magnetic field		<i>v</i> -e scattering
	CH target		Ŭ
JNE-	LAr+MPD move off-axis	Change neutrino flux spectrum	Deconvolute flu
RISM	011-0215		Energy response
			Provide FD-like spectrum at NE
			Uncover mismo



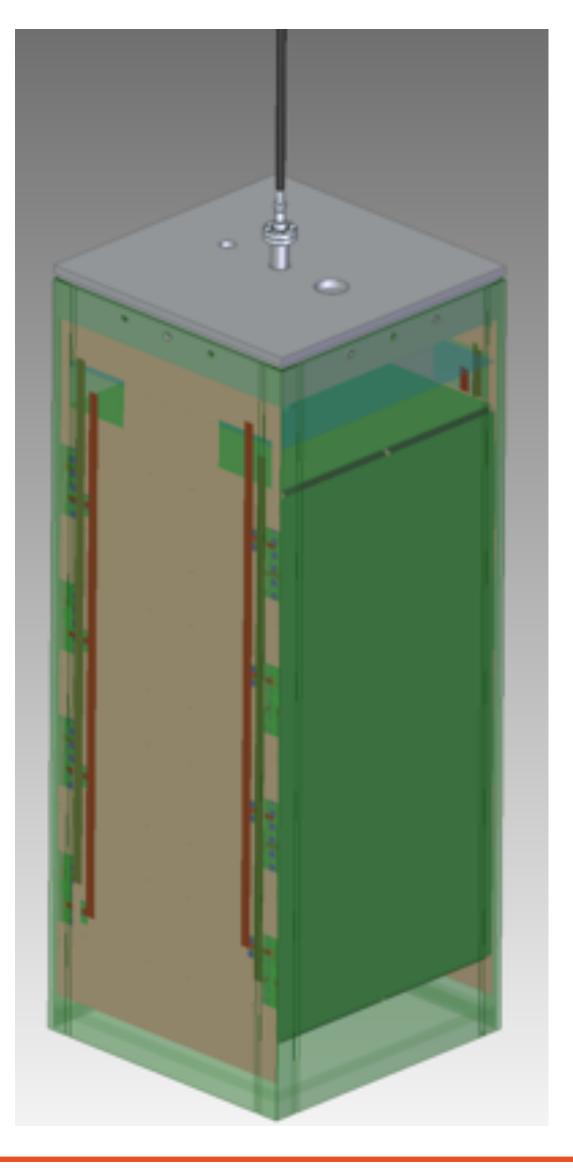
LAR DETECTOR

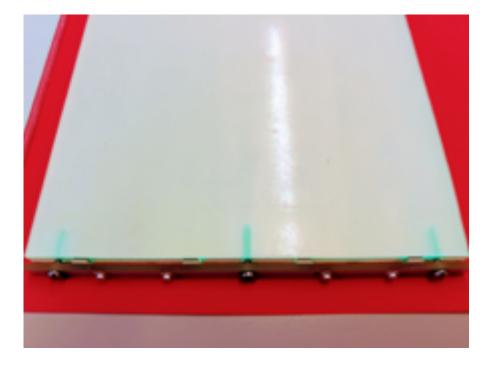
- Liquid Argon Time Projection Chamber (LArTPC) based on modularized design:
 - Same basic detection principles as DUNE Far Detector
- Detector built from 1 x 1 x 3 m³ self-contained modules
 - two sided-drift with pixel readout proves direct 3D representation of ionization activity
 - optical isolation provides a priori localization of scintillation signal
- Full detector is an array of 7 (width) x 5 (depth) modules with a total of ~150 tons of active LAr volume



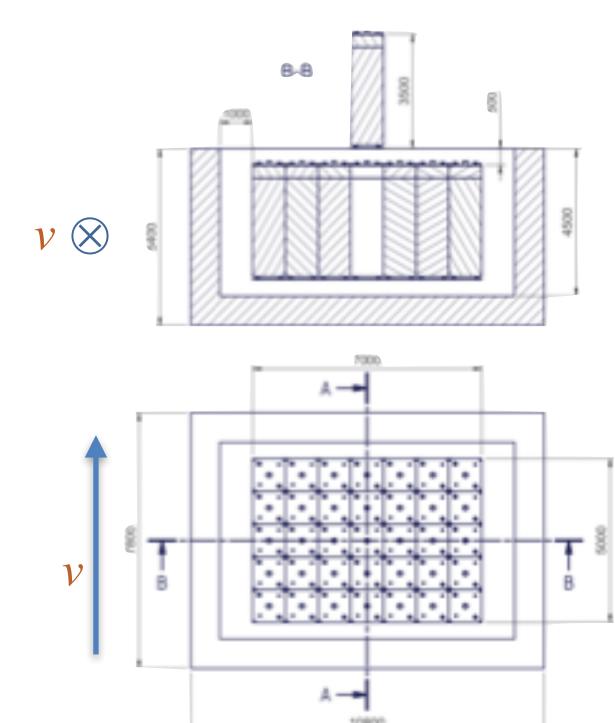


more in J. Sinclair's talk













LAR: BASIC CONSIDERATIONS

- High rates in the Near Detector Hall
 - Pixelization:
 - Overcome occupancy and ambiguity issues in wire readout detectors
 - Modularization
 - Faster drift and localization of scintillation signal to aid in associating to individual interactions
- Number/size of overall detector
 - Muon momentum reconstruction and hadron containment
- Primary and most direct view into how events look at FD
- Sufficient mass and capability for flux measurements via v-e elastic scattering, low-v



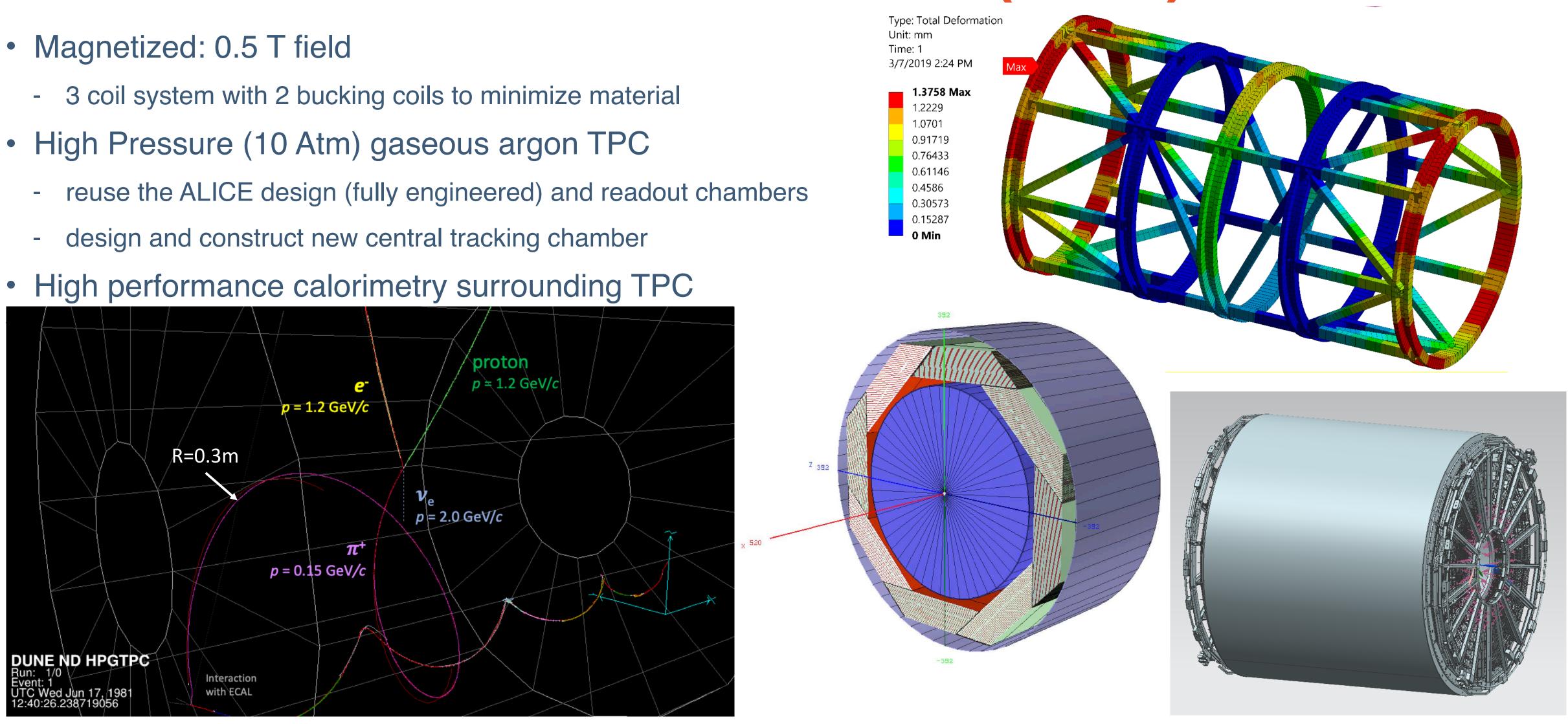






MULTI-PURPOSE DETECTOR (MPD)

- - 3 coil system with 2 bucking coils to minimize material





more in A. Bross's talk





MPD: BASIC CONSIDERATIONS

- Downstream muon spectrometer for LAr detector with HPgTPC
 - Acceptance (covers high energy muons coming from LAr)
 - Minimize material between LAr/HPgTPC (Air coil magnet, ribbed pressure vessel, optimize upstream ECAL)
- Observe *v*-Ar events with:
 - Sign-selection, magnetic spectrometry -
 - Very low tracking thresholds -
 - Minimal secondary interactions
- Flavor specific measurements by tagging lepton/sign
- Fully observe "bare" v-Ar interactions:
- disentangle: v-Ar from secondary interactions from detector response
- powerful tool in transferring ND measurements to FD



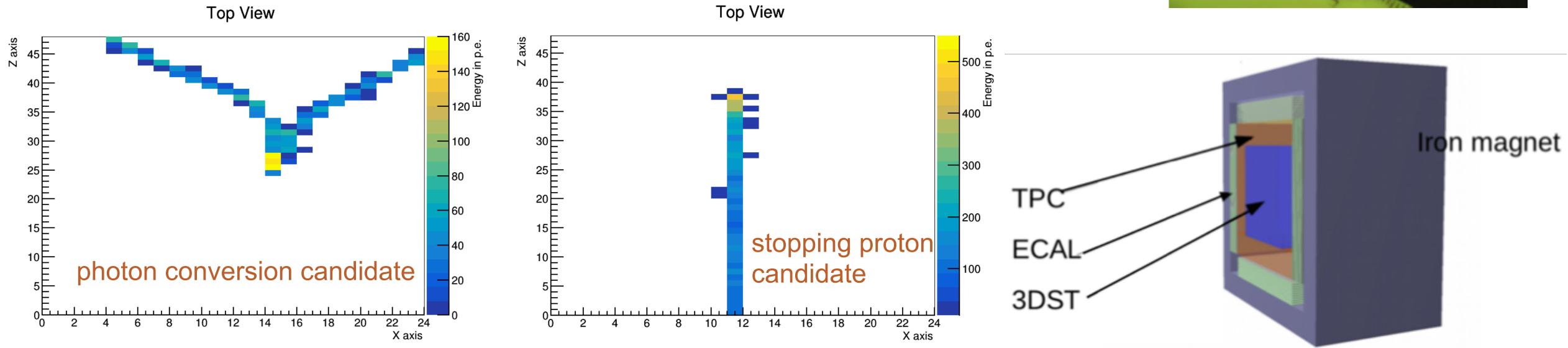






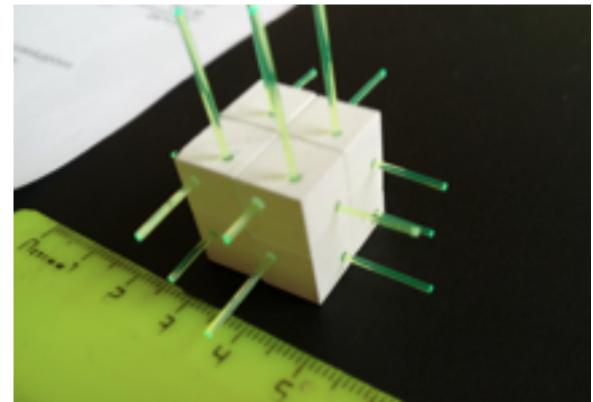
3DST-S

- Active scintillating target composed of 1x1x1xcm³ scintillator cubes
 - 2.4 x 2.4 x 2 m³ total volume -
 - fine-grained, isotropic tracking (proton tracking to ~300 MeV/c)
 - neutron tagging and spectrometry by time-of-flight
- Surrounded by TPC for tracking and ECAL in magnetic field





more in S. Manly's talk





3DST-S: BASIC CONSIDERATIONS

- On-axis flux monitor
 - sufficient rate, spectrometry capabilities, and transverse span
- Neutron detection
 - New capability in neutrino detectors
 - Nascent capabilities in MINERvA show potential
- *v*-CH sample
 - Cross check *v*-A modelling across A
 - Connect to "historic" data sets





Provides cross check on flux measurements with very different detector technology and capabilities



BACK TO OVERARCHING REQUIREMENTS

O0 Predict the neutrino spectrum at the FD: The ND must provide a prediction for the energy spectrum of v_{μ} , \bar{v}_{μ} , v_e , and \bar{v}_e at the FD. The prediction must be provided as a function of the oscillation parameters and systematic uncertainties must be small enough to achieve the required CP coverage. This is the primary requirement of the DUNE ND.

00.1	Measure interactions on argon	LAr prin MP cap
O0.2	Measure the neutrino energy	Bo inte 3D neu
O0.3	Constrain the cross section model	LA Cro O0 spe



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Ar: very high statistics of v-Ar interactions using the same detection nciples as the FD assisted by the MPD for μ spectrometry.

PD: a very detailed look at the same interactions with additional pabilities (very low thresholds, magnetic spectrometry, etc.).

oth LAr and MPD have ability to measure neutrino energy in v-Ar eractions, with complementary methods described in O0.1

OST-S has similar capabilities for *v*-CH interactions with additional utron detection capabilities

r + MPD samples as described above

oss checks from 3DST-S on CH with neutron detection

0.6 with LAr + MPD provide powerful variation in underlying energy ectrum to deconvolute the model from the flux.





BACK TO OVERARCHING REQUIREMENTS

00.4	Measure neutrino flux
00.5	Obtain data with different neutrino fluxes
00.6	Monitor the neutrino beam





Primarily LAr detector for $v_{\mu} / \overline{v}_{\mu} / (v_e + \overline{v}_e)$

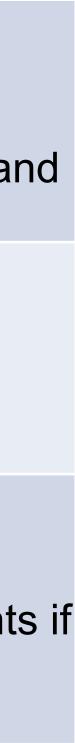
MPD will provide separation into v/\overline{v}

3DST-S will provide cross checks with different reconstruction and detection capabilities

DUNE-PRISM will be effected by moving LAr+MPD

3DST-S will provide dedicated measurements on-axis where potential variations are likely to be most visible

LAr+MPD can be moved to provide supplemental measurements if needed





TOWARDS A CDR REQUIREMENT TABLE

- Overarching (O):
 - Basic statements of the role/purpose of the near detector in the experiment
 - No reference to individual subsystems of particular implementations -
- **Capabilities (C):**
 - Measurements that the ND must make to deliver overarching requirements (e.g. v-e flux measurement) Refers to subsystem that provide the capability but not the particular implementation
- **Performance (P):**
 - Detector performance (efficiencies, resolution, etc.) needed to deliver the capabilities -Refers to specific subsystems and particular implementations as needed
- Technical (T):
 - Parameters/other properties (e.g. pixel pitch, magnetic field) specific to detectors to deliver the required performance



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• We proposes hierarchy of requirements working down from the overarching requirements:





CASE STUDY: BEAM MONITORING

Label	Name	Description	Driven by	Drives
00.6	Monitor the neutrino beam	Monitor the neutrino beam energy spectrum with sufficient statistics to be sensitive to intentional or accidental changes in the beam on short timescales. The precise requirement will be informed by the run plan as well as experience from previous experiments.	00	
C1	Monitor the neutrino event rate	The neutrino event rate must be measured to R% in D days	O0.6	P1
C2	Monitor the neutrino beam direction	The neutrino beam center in the coordinates transverse to the beam direction must be measured to C cm in D days	O0.6	P1, P2
C3	Monitor the neutrino beam spectrum	The neutrino event rate must be measured differentially to be sensitive to S% variations in the flux in B neutrino energy bins.	O0.6	P1, P3
P1	Neutrino event rate	The neutrino beam monitor must detect N neutrino interactions/day	C1, C2, C3	T1
P2	Vertex resolution	The neutrino beam monitor must resolve neutrino vertices to V cm	C2	T2
P3	Spectrometry	The neutrino beam monitor must measure muon energy to M% and hadron energy to H%	C3	Т3







CONCLUSIONS:

- The DUNE Near Detector Reference design involves three interconnected and complementary subsystems
 - LAr TPC observes v-Ar interactions based on the same basic detection principles as the Far Detector with enhancements to cope with the high rate environment
 - MPD provides muon spectrometry for the LAr detector and a precise view of the v-Ar with very low threshold, sign selection, and minimal secondary effects.
 - 3DST-S provides dedicated on-axis monitoring, v-CH interactions with powerful neutron detection capabilities that can cross check the neutrino model and provide cross checks on the neutrino flux measurement
- The DUNE Near Detector will employ the novel PRISM concept to expose the LAr+MPD system to varying neutrino spectra by moving the detectors off-axis
 - De-convolves the flux and cross section and provides powerful cross checks on the model
- While further study and optimization is needed, we believe this concept is the right one and is plausible in implementation and in meeting the needs of the DUNE physics goals



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