



### $v_{\tau}$ containment studies

Jeremy Hewes  $v_{\tau}$  and high-energy beam tune meeting 1st November 2018



## Introduction

- Looking at  $v_{\tau}$  interactions using a full simulation and reconstruction chain:
  - v events simulated using GENIE.
  - GEANT4 propagates MC particles through detector.
  - Detector simulation to convert these particles into signals on wires.
  - Reconstruction chain:
    - Wire processing.
    - Hit finding & clustering.
    - Object reconstruction (tracks, showers).
    - Neutrino energy reconstruction.



# Motivation

- Studies by Saul Monsalve, Leigh Whitehead, Alex Radovic et al. show great promise for CVN (convolutional visual network).
- Given its success at identifying  $v_e$  and  $v_\mu$  events, how effective is CVN at identifying  $v_\tau$  interactions?
- Their model performed poorly on  $v_{\tau}$  events due to a severe lack of MC statistics on which to train the network.
- Produced an enriched v<sub>τ</sub> MC sample to add to their dataset, but this introduced new problems.
  - Now at higher energies, v<sub>τ</sub> events overrepresented in training sample.
  - Network classifies any high-energy neutrino interaction as a  $v_{\tau}$ .







# Training samples

- Produced high-stats (~800k total) simulation samples for six neutrino types (v<sub>e</sub>, v<sub>µ</sub>, v<sub>τ</sub>, v̄<sub>e</sub>, v̄<sub>µ</sub>, v̄<sub>τ</sub>) with an event rate approximately flat in true neutrino energy between 3.5 and 80 GeV.
- Trained a ResNet network, but accuracy for  $v_{\tau}$  events significantly worse than other neutrino flavours.
- The small 1x2x2 workspace geometry used meant the majority of  $v_{\tau}$  events were uncontained.





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# Containment studies

- Examining how well-contained events are in each available geometry before beginning another round of high-stats MC production.
- Clearly workspace is too small, but 1x2x6 and 10kt are both candidates.
- Use new lower-energy warped flux (150 MeV - 15 GeV).
- Produced low-stats (~1000) samples for  $v_e$ ,  $v_\mu$ ,  $v_\tau$  in both detector geometries.
- Flux warping to approximate event rate flat in energy is less accurate for v<sub>τ</sub>s, likely because of lower cross-sections at lower energies.
  - Should produce separate warped flux for  $v_{\tau}$  events to correct for this.





# Containment definition

- Simulation provides "MC reconstructed" objects based on simulated ionization electrons.
  - Construct "MC showers" and "MC tracks" for particles which are above thresholds and visible in the detector.
  - A version of MC truth information which describes what we see in the detector.
- Require all MC reco objects have a start and end point inside the fiducial volume.
- Strict definition of containment every particle visible in the detector must be completely contained.





### ve containment





### $v_{\mu}$ containment





### $v_{\tau}$ containment





## Atmospheric containment

- We have an MC sample of beam neutrinos with an event rate ~flat in energy. What can this tell us about atmospheric neutrinos?
- Downward-going neutrinos only have ~12m of detector in which to stop.
- Define a 12x12x12 subset of the detector, and see what % of events with a vertex inside this volume are contained within it.



# Containment



- Containment in 12x12x12 volume generally very poor.
- $v_{\mu}$  containment in CDR is ~75%, but this only considers the muon containment, whereas we consider *everything*.

Table 4.3: Atmospheric neutrino event rates including oscillations in 350 kt  $\cdot$  year with a LArTPC, fully or partially contained in the detector fiducial volume.

Sample	Event Rate
fully contained electron-like sample	14,053
fully contained muon-like sample	20,853
partially contained muon-like sample	6,871

 Would be interesting to do this study for simulated atmospherics in full detector.



## Angular resolution

- First-pass look at angular resolution:
  - Calculate  $\cos \theta_v$  between true neutrino direction and beam direction.
  - Calculate  $\cos \theta_{vis}$  between directionality of visible system and beam direction.
    - Direction of visible system is calculated from the directions of all MC reco particles, weighted by momentum.
  - Take the difference between these two to calculate angular resolution.
  - Look at the mean and spread in these as a function of true neutrino energy.



## Angular resolution — all events





### Angular resolution — all $v_e$ events



![](_page_14_Picture_0.jpeg)

### Angular resolution — all $v_{\mu}$ events

![](_page_14_Figure_2.jpeg)

![](_page_15_Picture_0.jpeg)

### Angular resolution — all $v_{\tau}$ events

![](_page_15_Figure_2.jpeg)

![](_page_16_Picture_0.jpeg)

## Conclusions

- Containment is an issue even for the larger "reduced" detector 1x2x6 geometry.
  - Gives us motivation to produce MC in the full DUNE 10kt geometry for our next round.
- We are able to:
  - Produce migration matrices for true v energy vs calorimetric energy for contained events.
  - Make statements on what proportion of interactions we expect to be contained.
  - Any other useful inputs for sensitivity studies?