

# Searching for neutrino decoherence from quantum gravitational space-time fluctuations with IceCube

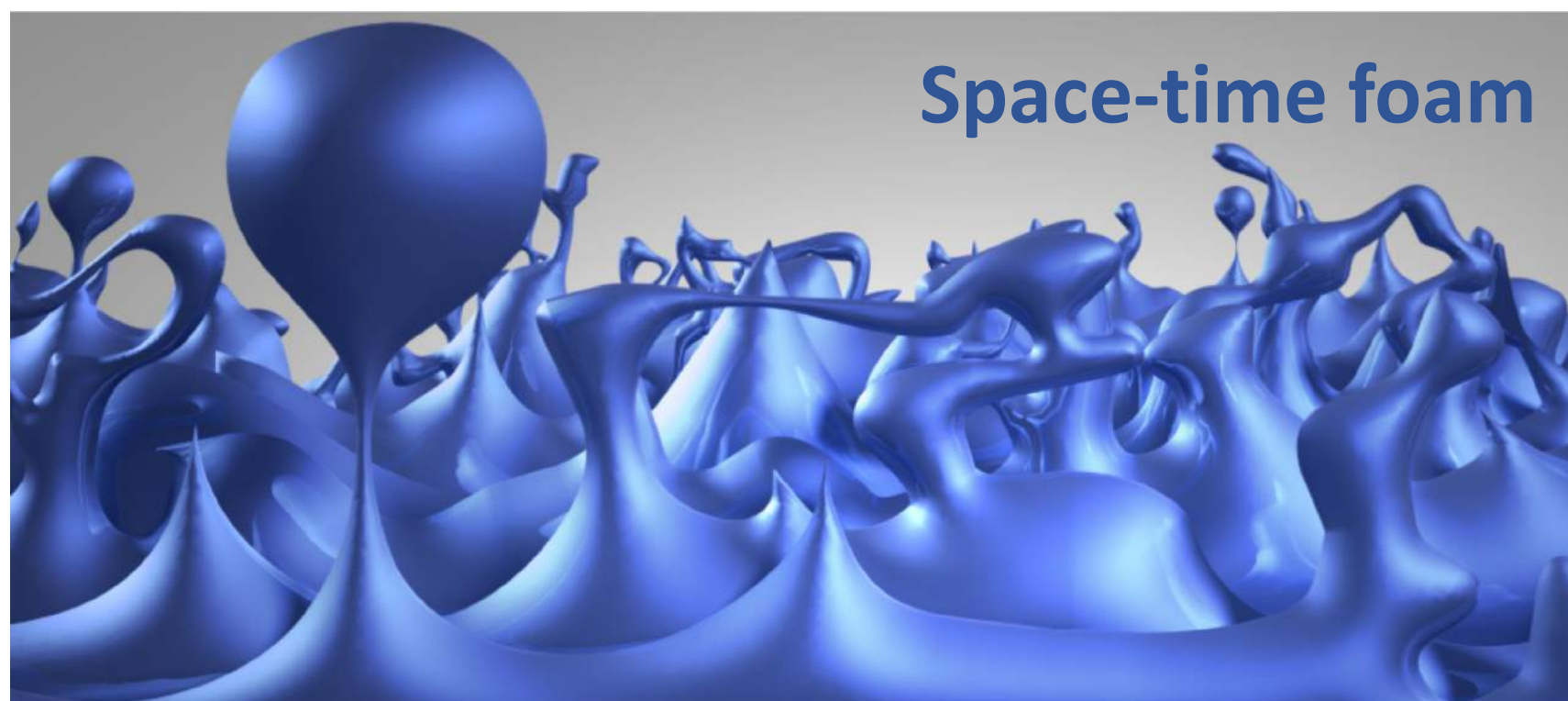
Tom Stuttard Niels Bohr Institute for the IceCube collaboration

## 1. Neutrino decoherence from quantum gravity

If gravity is a quantum force, the curvature of space-time may fluctuate at extremely small distance scales → **space-time foam**.

Sufficiently strong fluctuations conjectured to produce short-lived **virtual black holes** (VBH).

Stochastic interactions of neutrinos with VBHs can produce **loss of coherence** and other observable effects.

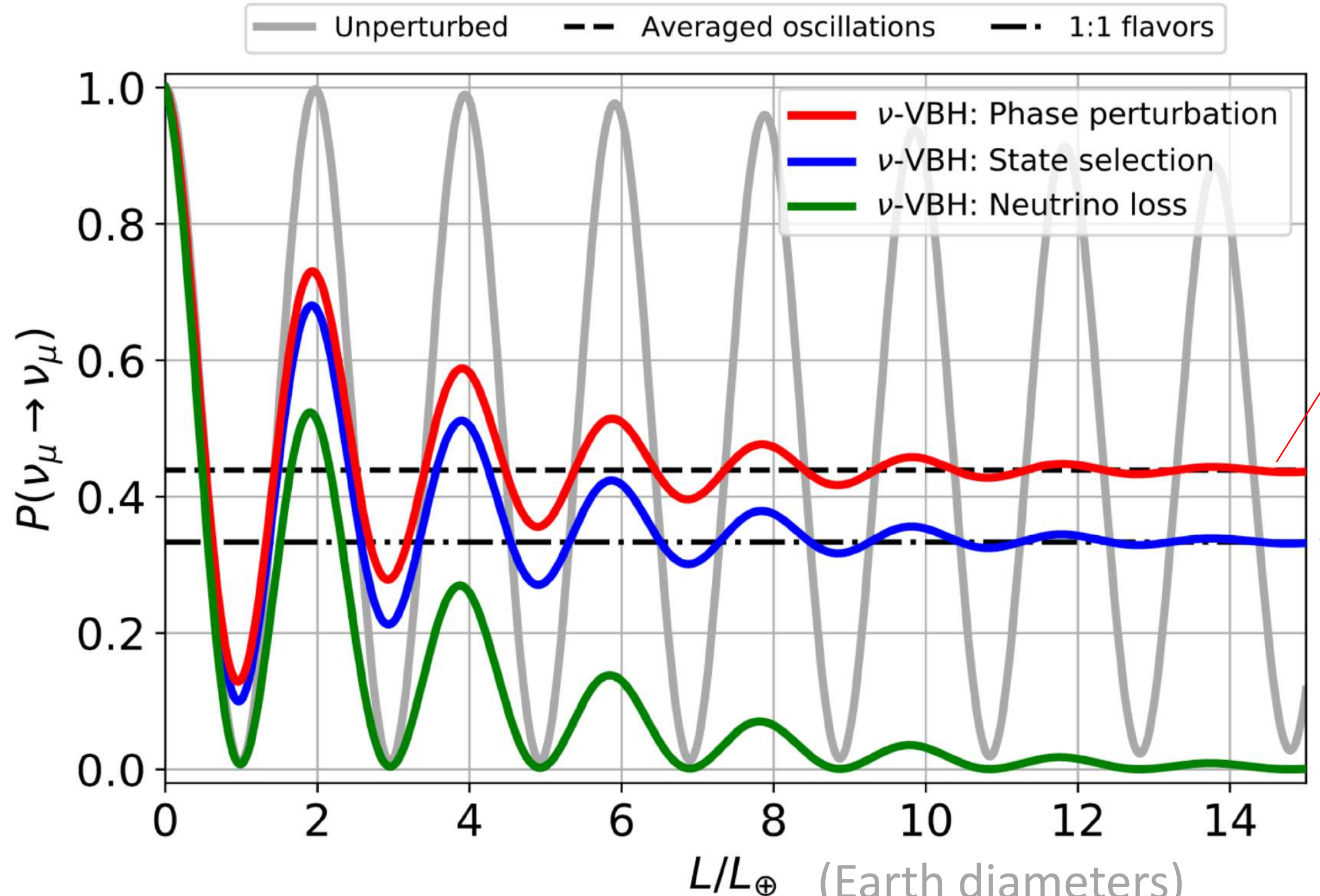


Credit: Chandra

## 2. Modelling ν-VBH interactions

Model by stochastically injecting heuristic ν-VBH interaction scenarios into neutrino propagation software model.

Effect on neutrino ensemble is **damping of oscillations**.



**3 scenarios tested:**

- Extreme phase perturbation:  $\nu \rightarrow \text{BH} \rightarrow \nu$  with  $\Delta\phi$
- Neutrino flavor/mass state democratically selected:  $\nu \rightarrow \text{BH} \rightarrow \nu_{e,\mu,\tau,1,2,3}$
- Neutrino lost:  $\nu \rightarrow \text{BH}$  with wavy lines indicating loss

Can reproduce signal analytically with open quantum system formalism:

$$\dot{\rho} = -i[H, \rho] - \mathcal{D}[\rho] \Rightarrow \mathcal{D}[\rho] = (D_{\mu\nu} \rho^\nu) b^\mu$$

$D_{\text{phase perturbation}} = \text{diag}(0 \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma)$

$D_{\text{state selected}} = \text{diag}(0 \ \Gamma \ \Gamma \ 0 \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ 0)$

$D_{\text{neutrino lost}} = \text{diag}(\Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma \ \Gamma)$

Allows interpretation of model limits in terms of underlying ν-VBH interaction parameters:

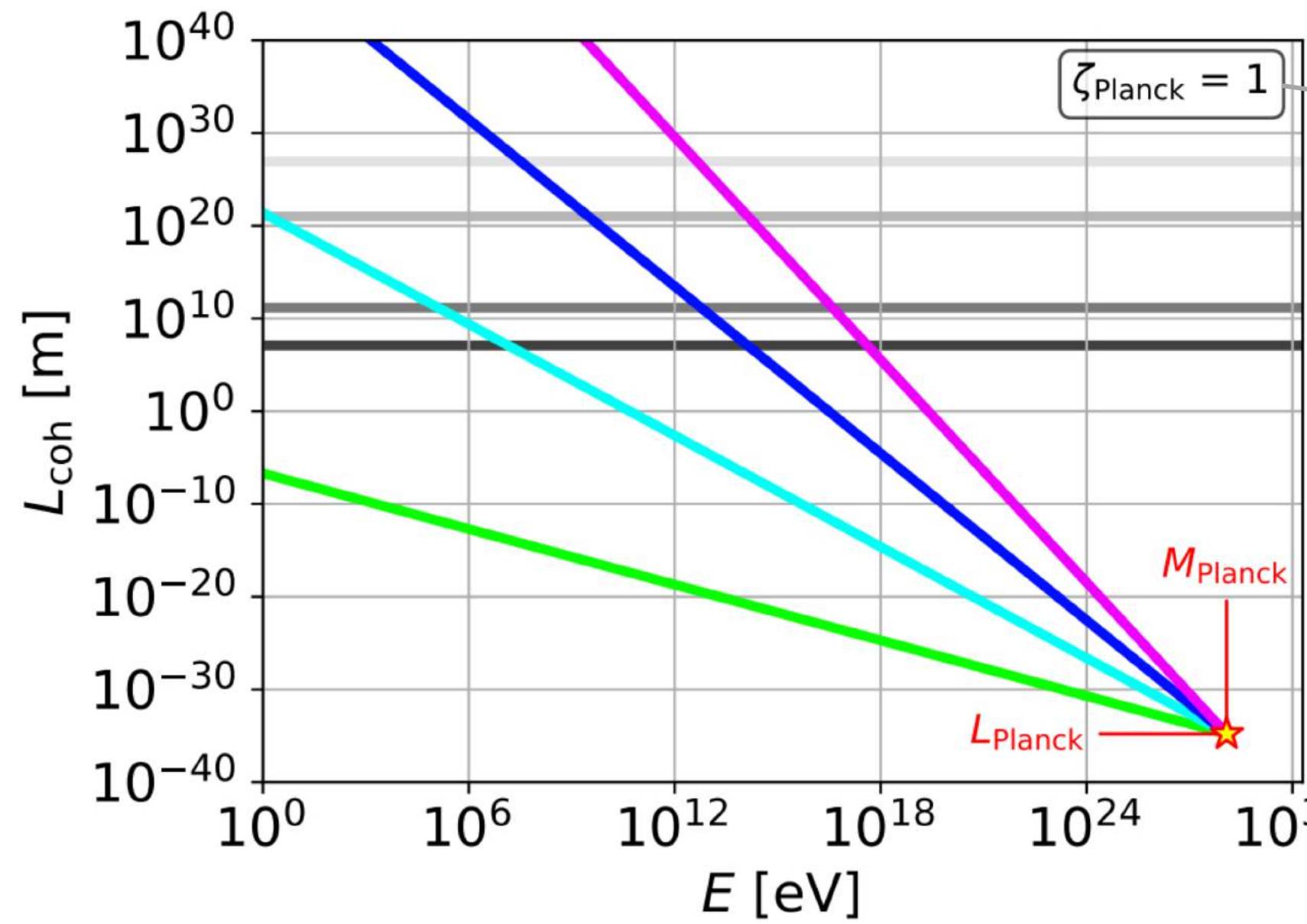
→  $1/\Gamma = \text{interaction mean free path}$ .

Coherence length:  $L_{\text{coh}} \equiv 1/\Gamma$  (= Earth diameter in example to left)

## 3. Energy dependence

No accepted quantum gravity theory → instead test a range of phenomenological ν-VBH interaction energy-dependence cases.

Express relative to Planck scale physics:  $\Gamma(E) = \zeta_{\text{Planck}} \frac{E^n}{M_{\text{Planck}}^{n-1}}$



$L_{\text{coh}}(E) = \frac{L_{\text{Planck}}}{\zeta_{\text{Planck}}} \left( \frac{M_{\text{Planck}}}{E} \right)^n$

‘Natural’ case:  $\zeta_{\text{Planck}} = 1$

- $n=1$ : Earth diameter
- $n=2$ : Earth-Sun distance
- $n=3$ : Milky Way diameter
- $n=4$ : Observable Universe

Sensitivity to ‘natural’ Planck scale physics for  $n < 3$  with atmospheric ν

## 4. A search with DeepCore

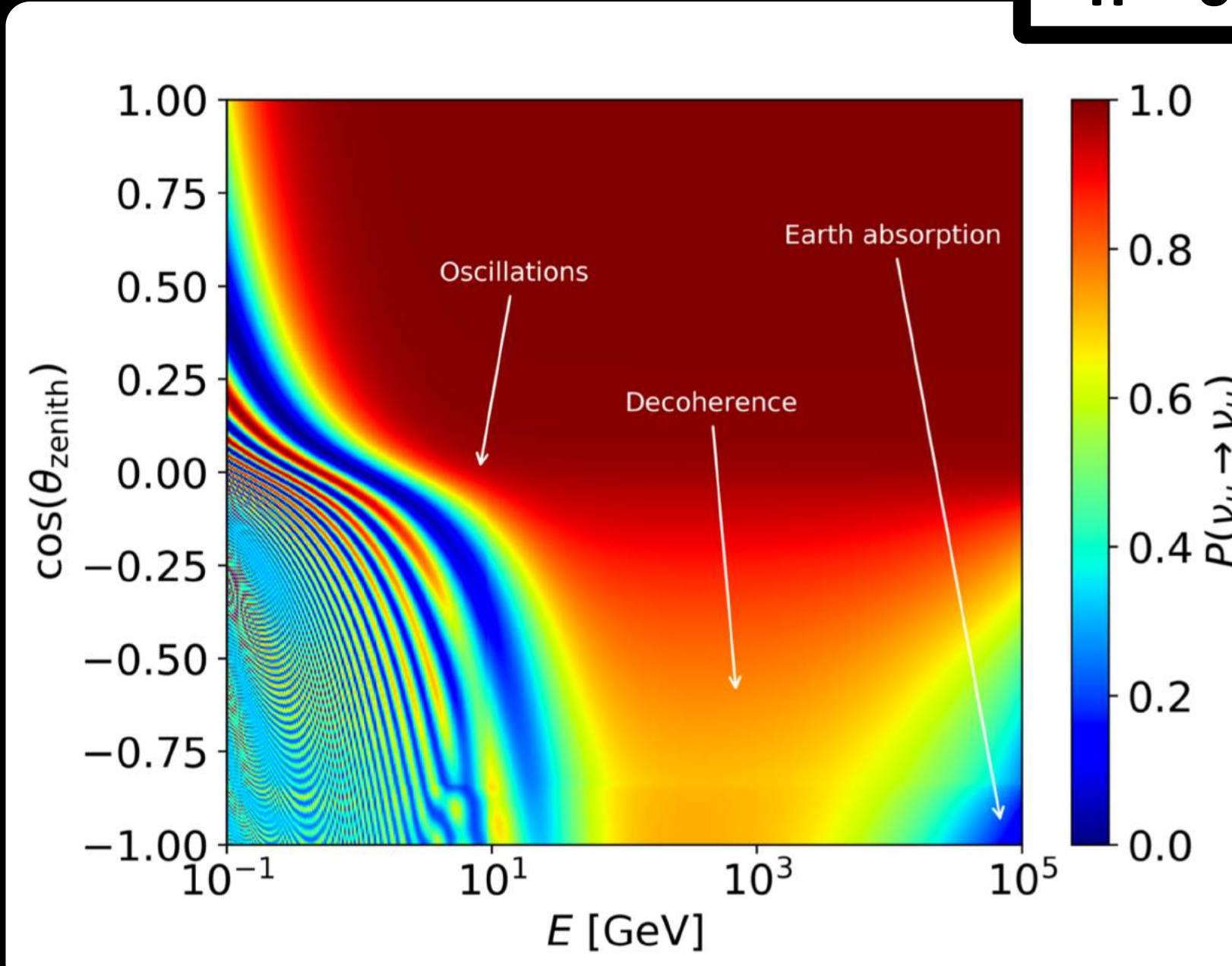
DeepCore is the dense core of the IceCube neutrino observatory.

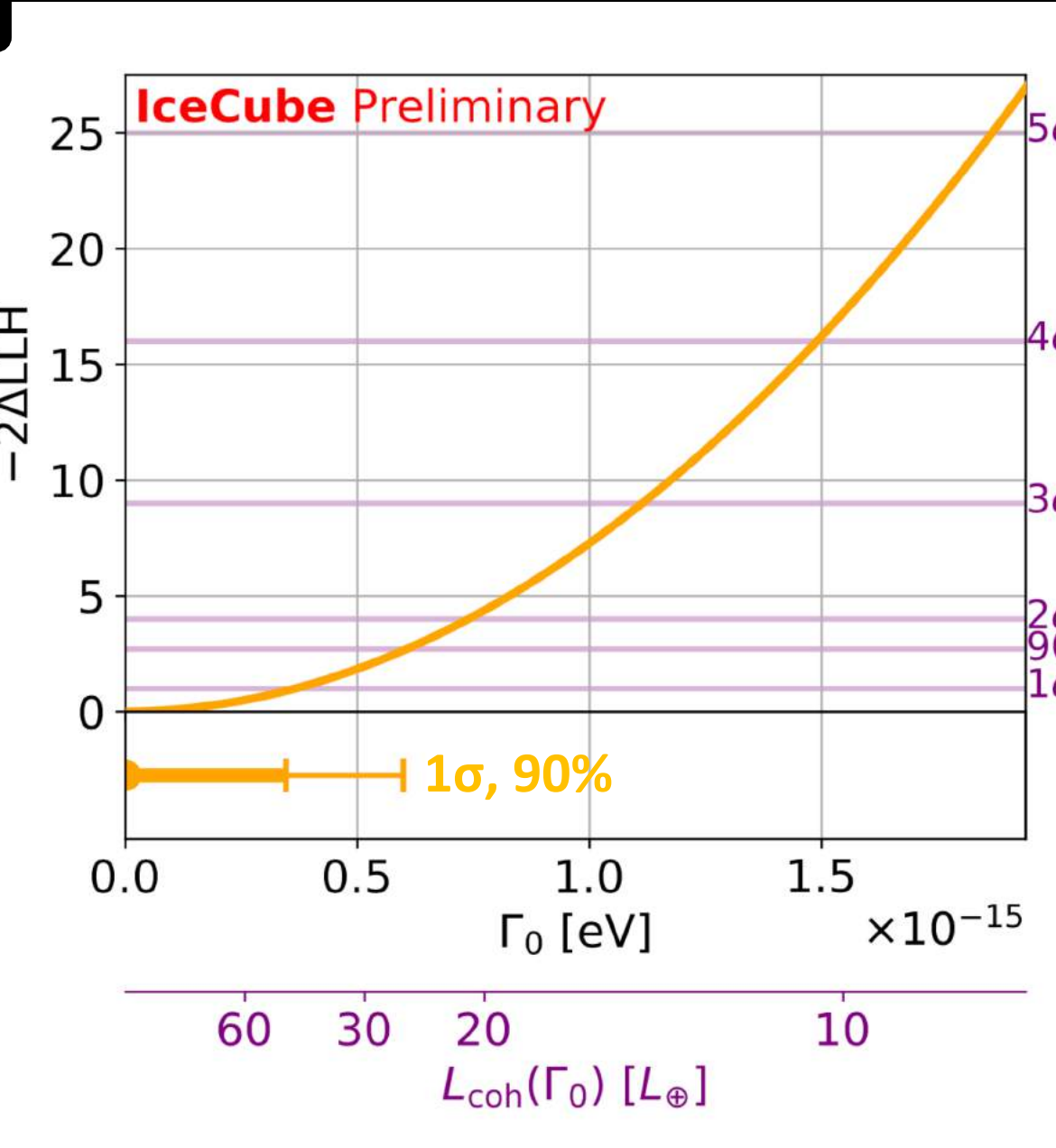
Performing an experimental search for modified atmospheric ν oscillations consistent with ν-VBH interactions.

Using new 8 year 5-300 GeV data sample (>300,000 ν) → results soon!

Example sensitivities:

**n = 0**





### References

- [1] G. Lindblad, Comm. Math. Phys. 48, 119 (1976)
- [2] E. Lisi, A. Marrone, and D. Montanino, Phys. Rev. Lett. 85, 1166 (2000)
- [3] F. Benatti and R. Floreanini, Journal of High Energy Physics, 032 (2000).
- [4] A. M. Gago, E. M. Santos, W. J. C. Teves, and R. Z.Funchal, arXiv:hep-ph/0208166 (2002)

### Coming soon...

- 1) Phenomenology paper (ν-VBH interactions, lightcone fluctuations)
- 2) DeepCore (GeV) and IceCube (TeV) experimental results

**n = 2**

