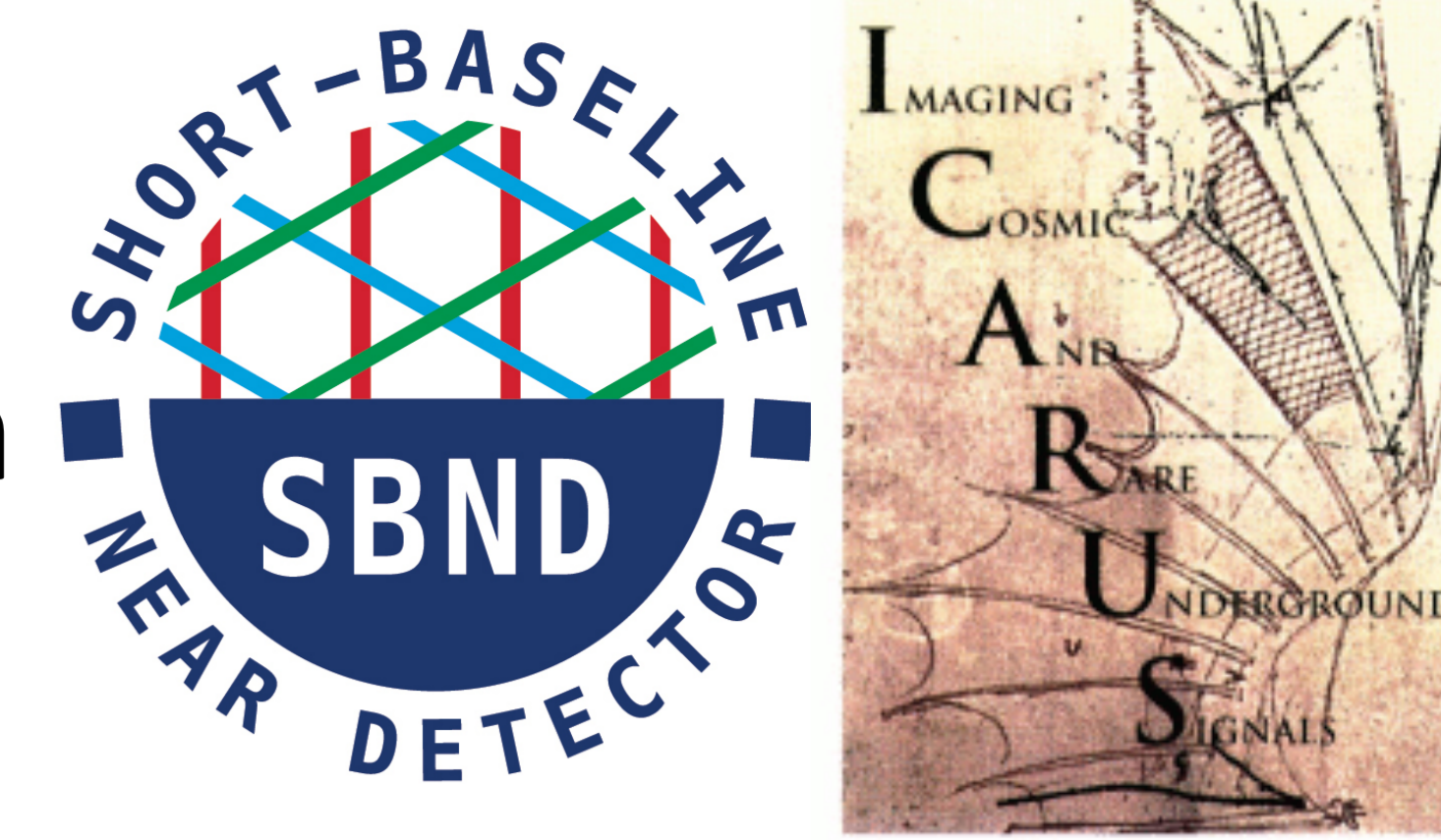


# Sensitivity to eV scale sterile neutrino oscillations at SBN

Jacob Larkin (Brookhaven National Lab & Stony Brook University) for the SBN collaboration



## 3+1 Oscillation Model

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

$$P_{\alpha \rightarrow \beta} = \left| \delta_{\alpha\beta} - \sin^2 2\theta_{\alpha\beta} \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right) \right|^2$$

$$\sin^2 2\theta_{\mu e} = 4 |U_{\mu 4}|^2 |U_{e 4}|^2$$

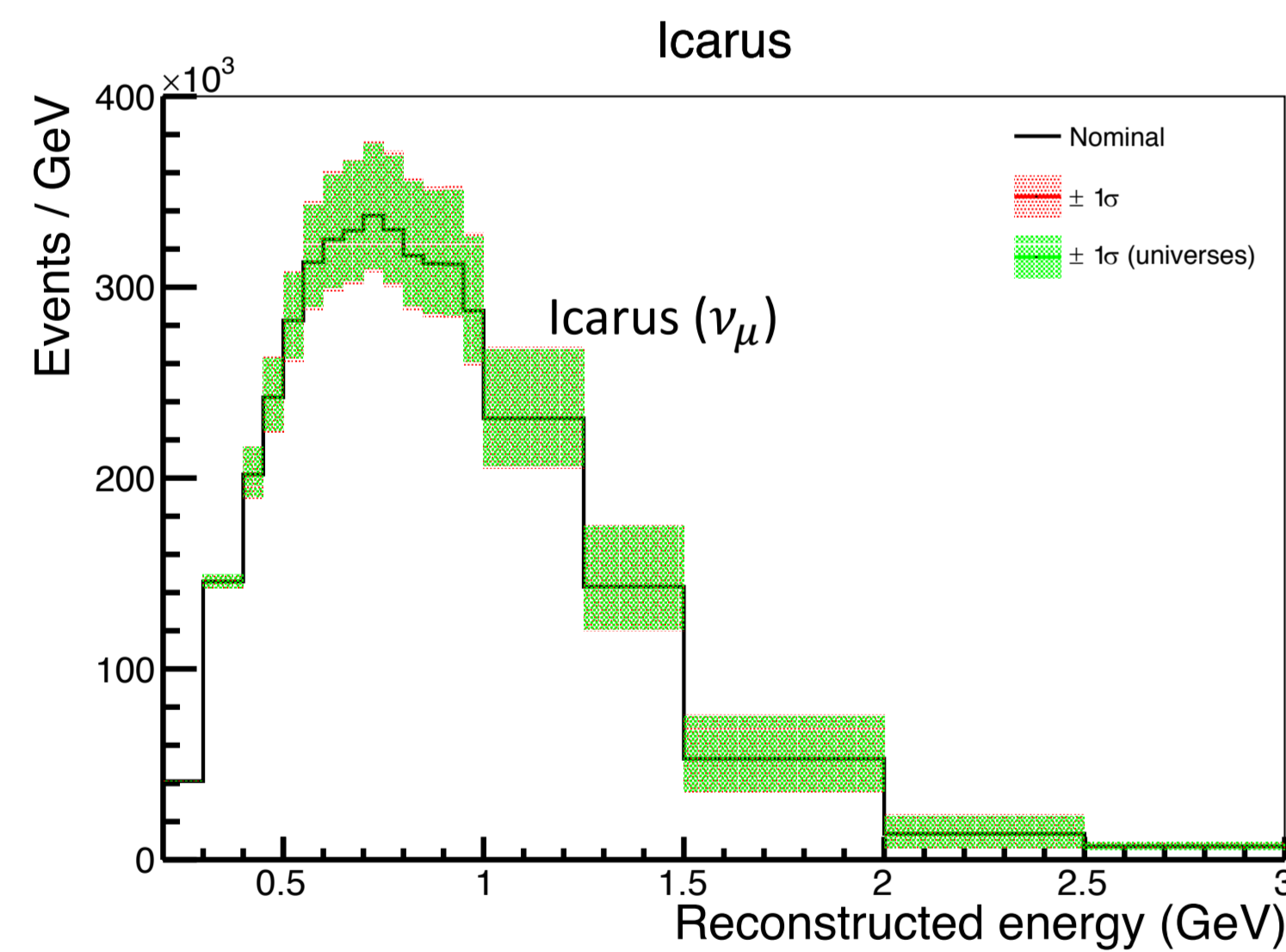
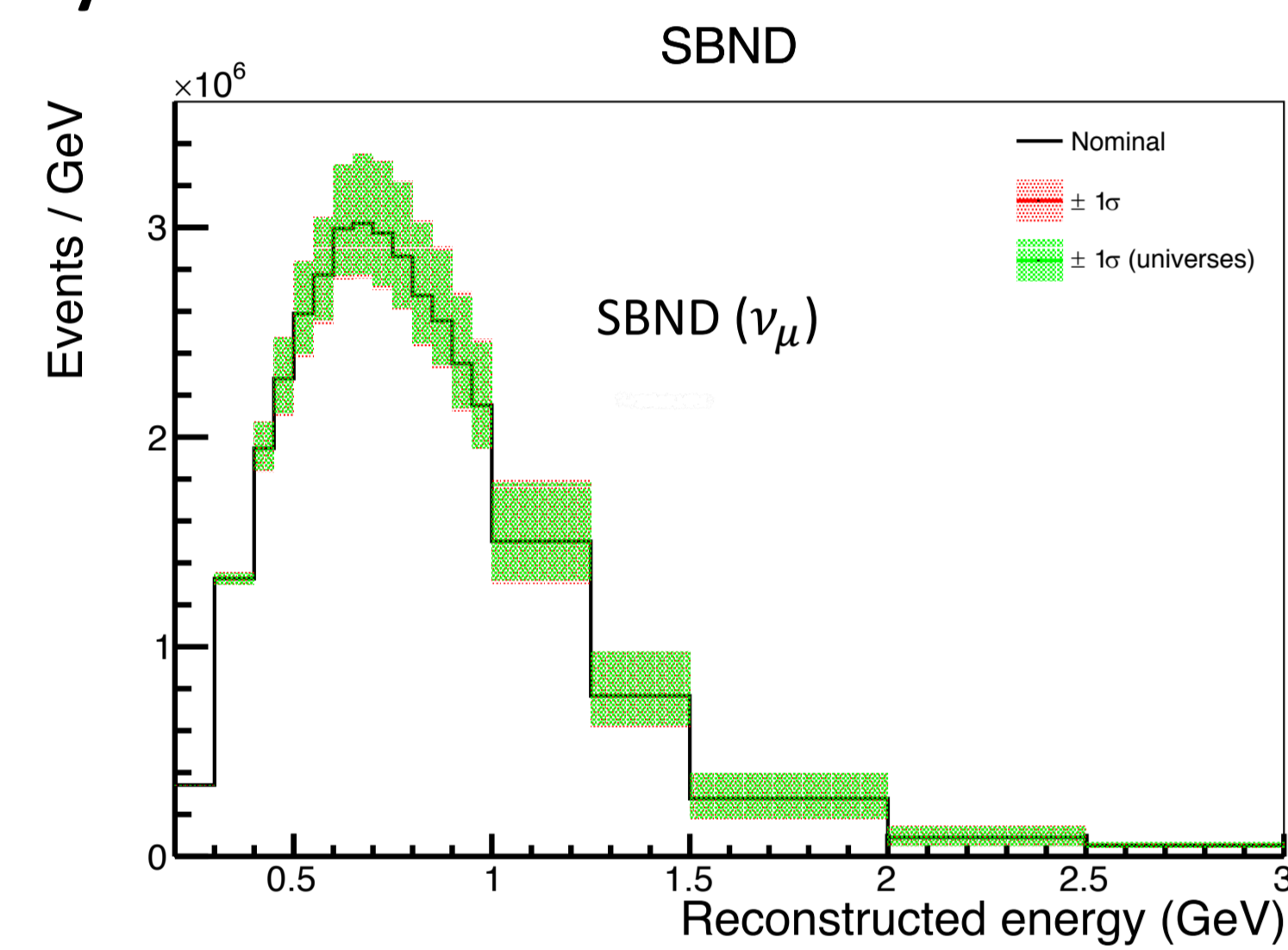
$$\sin^2 2\theta_{\mu\mu} = 4 |U_{\mu 4}|^2 (1 - |U_{\mu 4}|^2)$$

## CAFAna

- Developed originally for NOvA, and used in DUNE LBL analysis
- Analysis and fitting framework facilitates exploration of sensitivity and impact of systematics
- Using binned Poisson distributed data ( $M = \text{expected}$ ,  $D = \text{observed}$ ,  $P(\theta) = \text{penalty term}$ ):

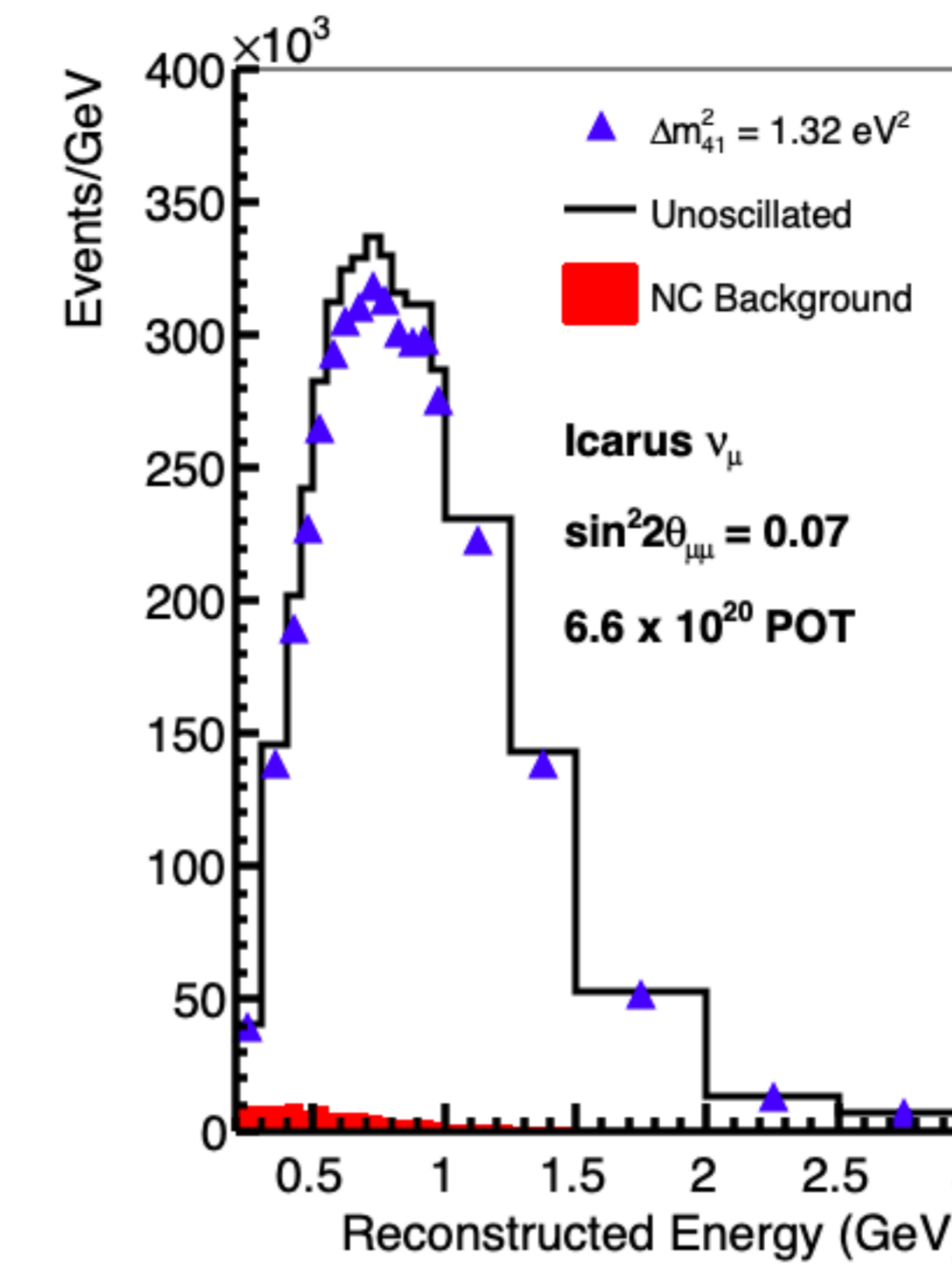
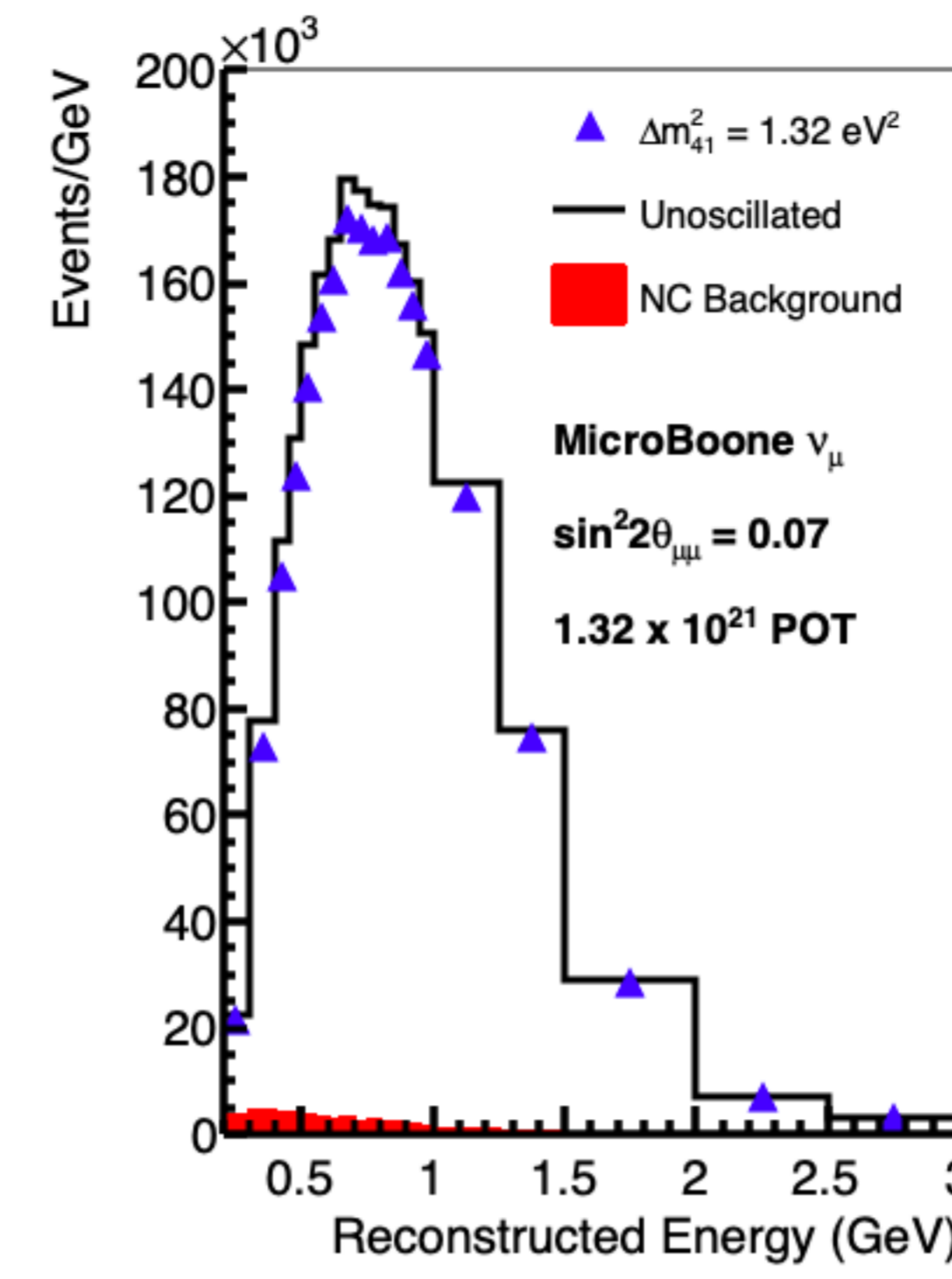
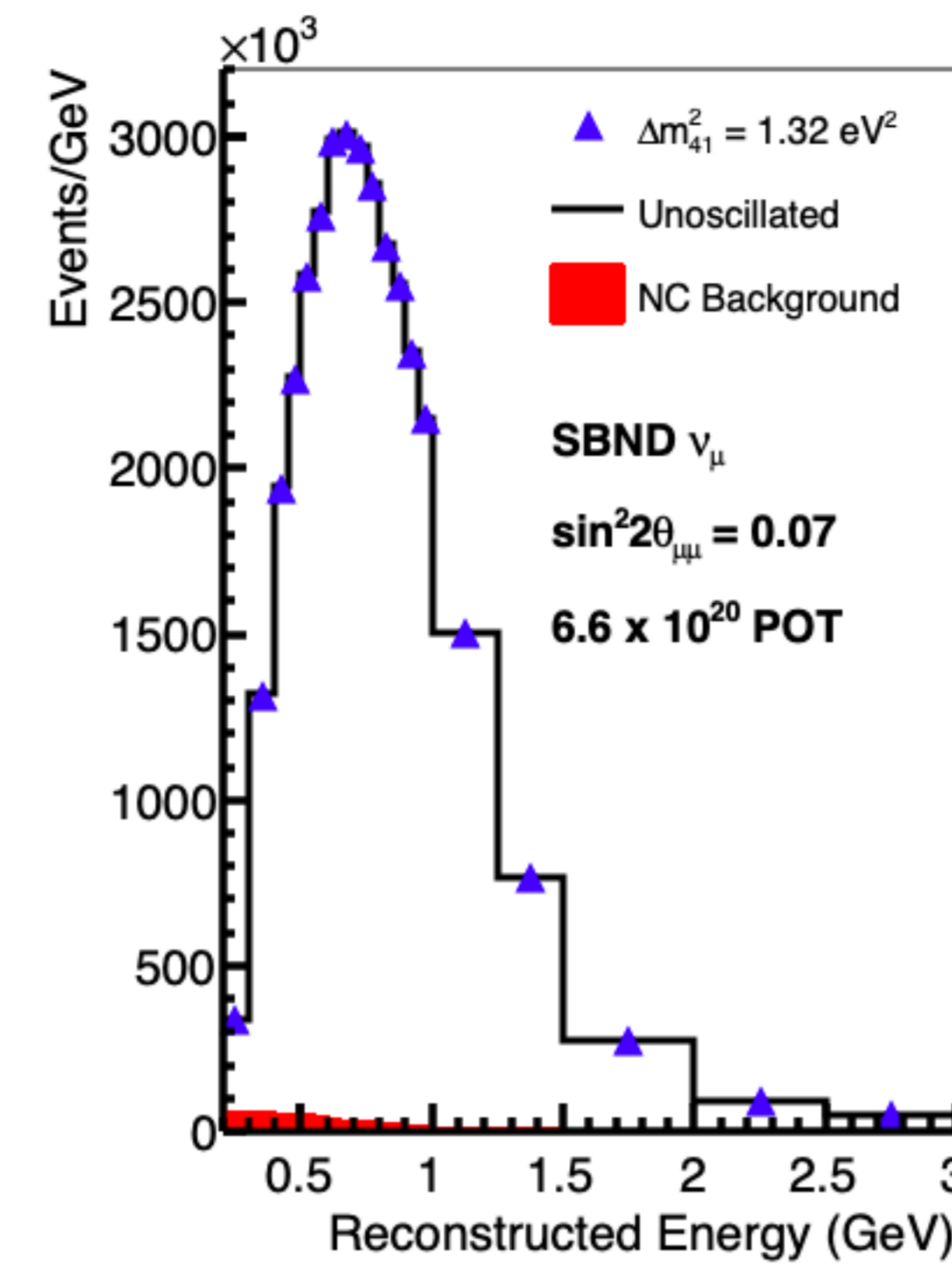
$$\chi^2 = 2 \sum_i^{N_{bins}} \left[ M_i(\theta) - D_i + D_i \log \left( \frac{D_i}{M_i(\theta)} \right) \right] + P(\theta)$$

## Systematics



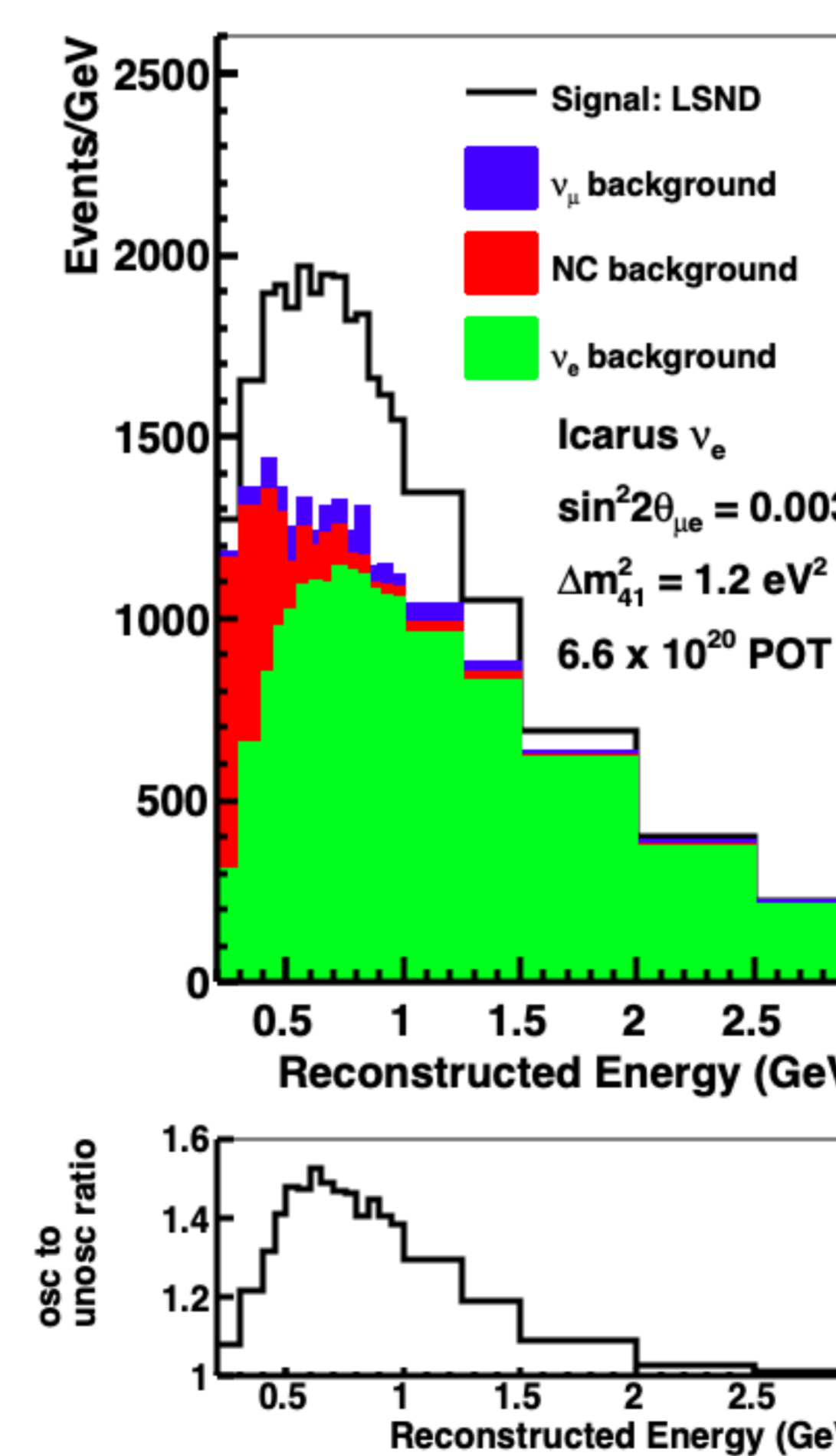
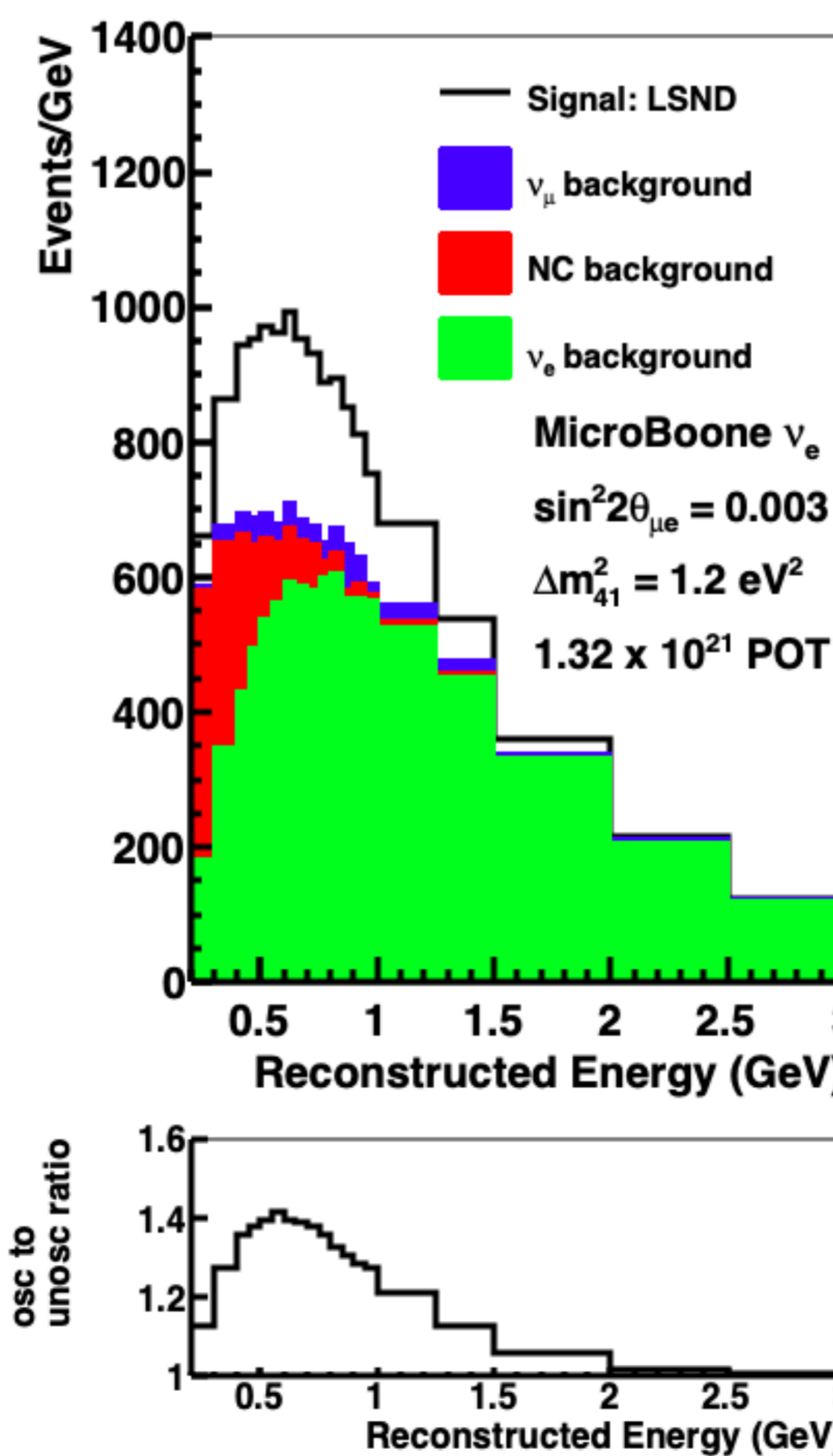
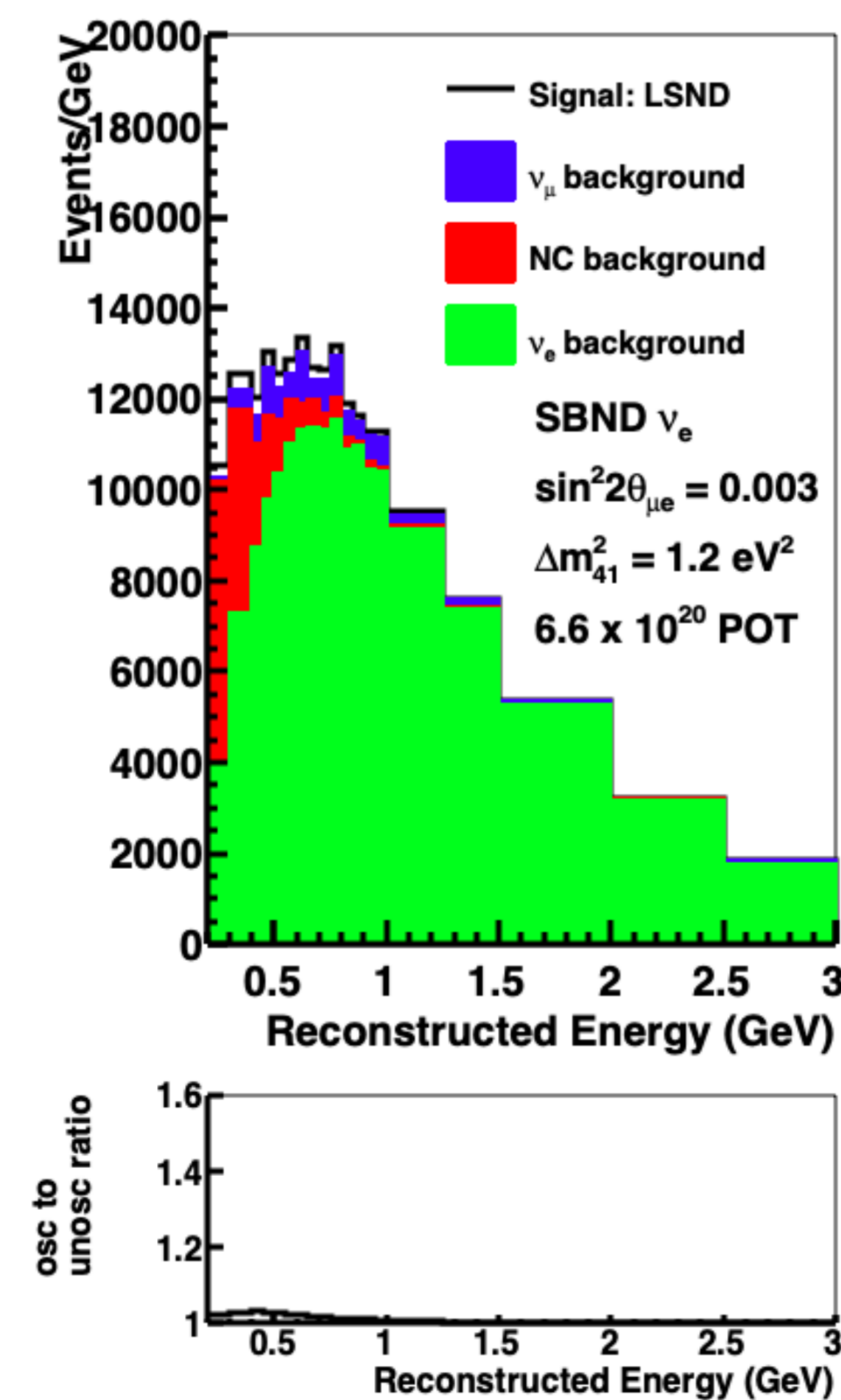
- Red: 68% error band from CAFAna systematic parameterization
- Green: 68% error band from 1000 generated universes
- CAFAna covers variations in the universes well
- Preliminary estimate of flux and interaction systematics included
- Detector effects not yet included

## Muon Neutrino Disappearance



- Oscillated spectra (blue triangles) show disappearance signal when compared to unoscillated spectra (solid line)
- Bottom plots show oscillated over unoscillated ratio in each bin, also showing disappearance near the peak of the MicroBoone and Icarus spectra

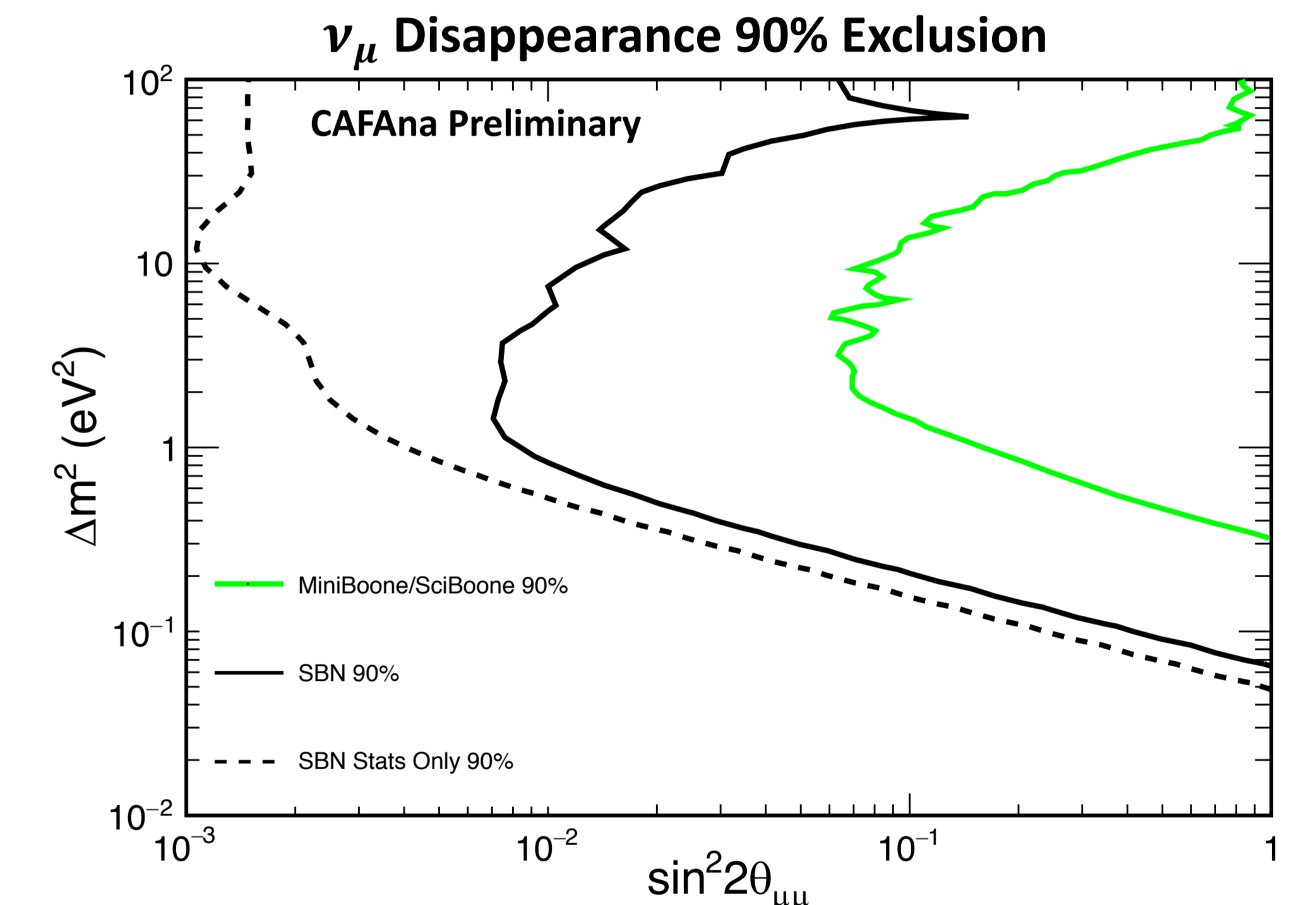
## Electron Neutrino Appearance



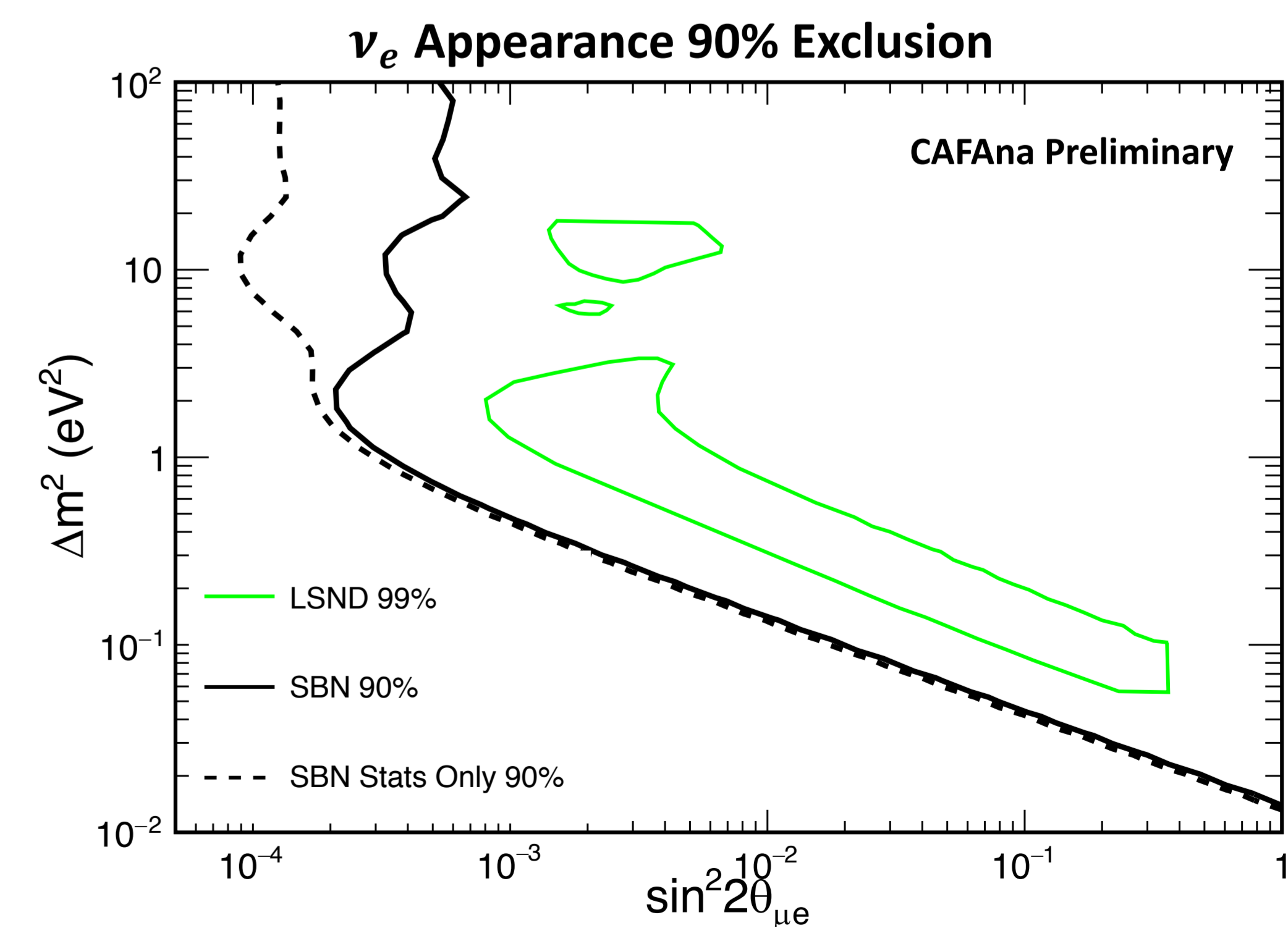
- Oscillated spectra (black line) show appearance signal when compared to background
- Bottom plots show oscillated over unoscillated ratio in each bin, also showing appearance near the peak of the MicroBoone and Icarus spectra

## Sensitivities

- Multi-detector experiment reduces impact of systematic uncertainty
- The oscillation sensitivities shown here, employ the same systematic error assignments used in the SBN proposal [arXiv:1503.01520] with known caveats. Work is currently in progress to improve the SBN oscillation sensitivity calculations by a) migrating to a better-motivated interaction model and modern uncertainty assignments from GENIE 3, b) incorporating a first evaluation of uncertainties in the detector response, and c) evaluating model dependencies and biases in the extrapolation from SBND to ICARUS (a-c not included here).



- Significant increase in sensitivity over MiniBoone/SciBoone (green contour) [arXiv:1106.5685]



- 3 detector sensitivity (black) covers the entire 99% allowed region from LSND (green) [arXiv:1803.10661v1]