

Astrophysical neutrinos under the effect of active-sterile secret interactions

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Model

- Extension of the Standard Model: sterile neutrinos interacting with the active ones through secret interactions

$$\mathcal{L} = - \sum_{\alpha} \lambda_{\alpha} \bar{\nu}_{\alpha} \gamma_5 \nu_s \phi$$

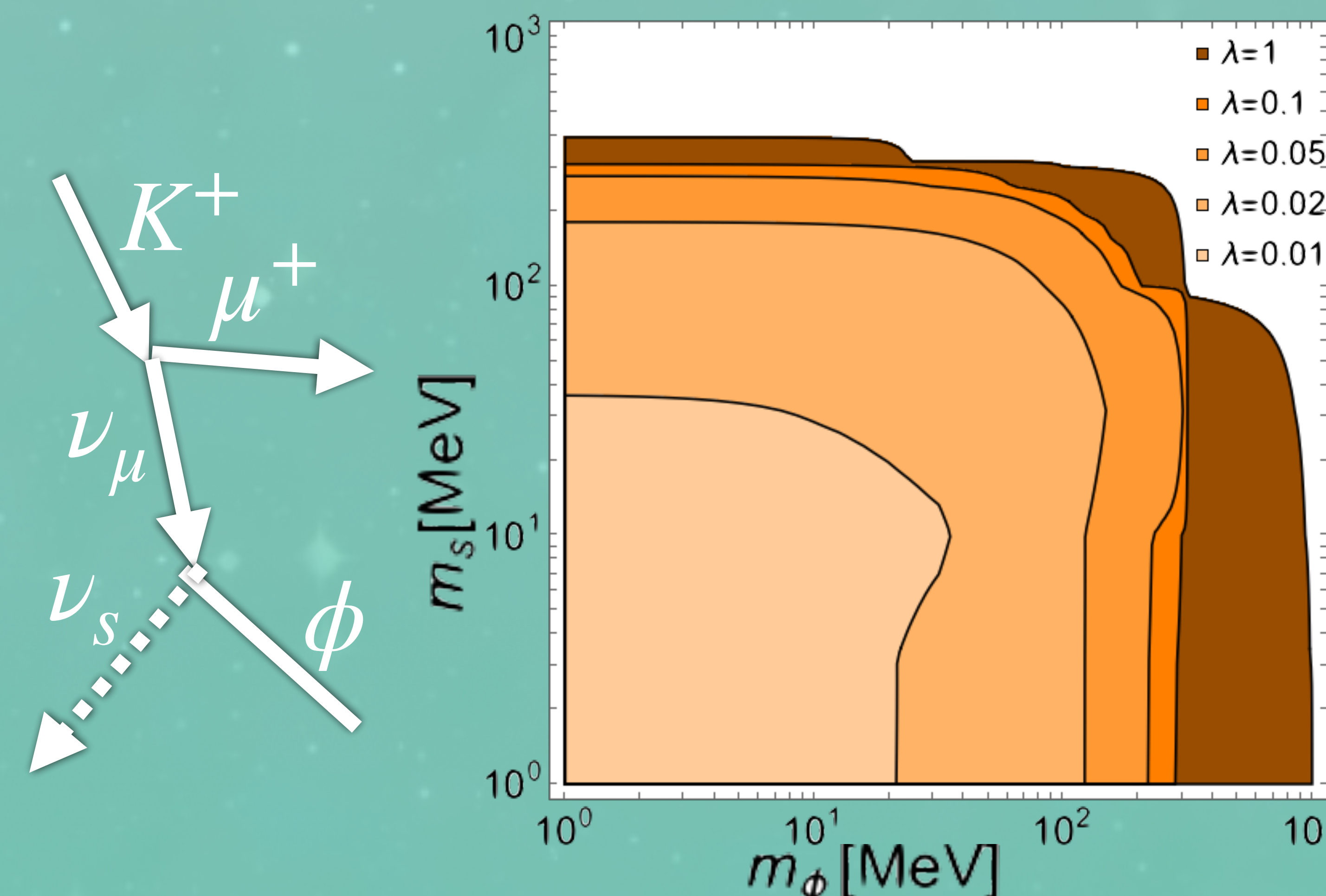
- Astrophysical active neutrinos collide with neutrinos from the Cosmic Neutrino Background (CNB) disappearing into sterile neutrinos

Cosmological constraints

- Sterile neutrinos and scalar particles must have large masses (above 10 MeV)

Laboratory constraints

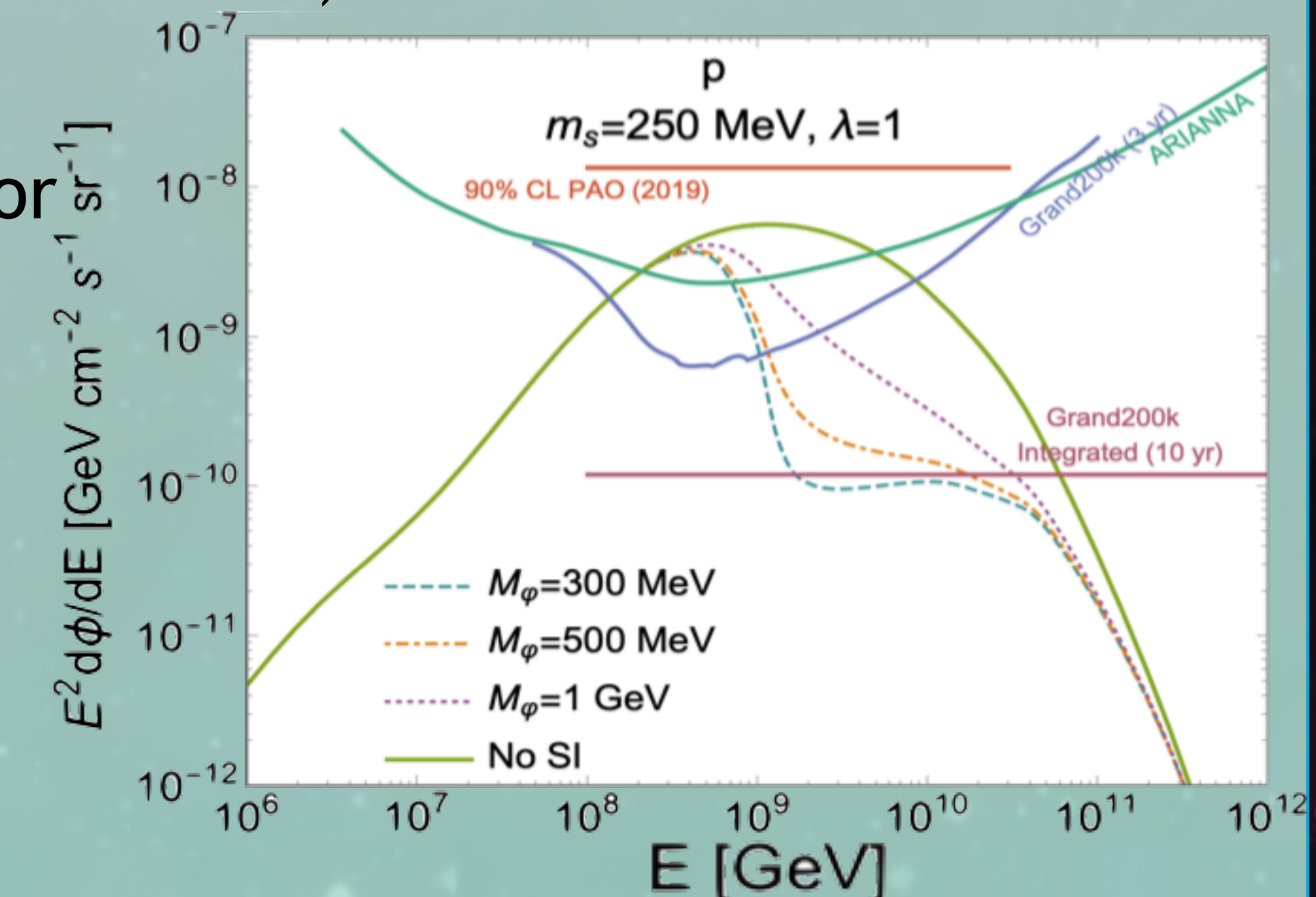
- Secret interactions open new decay channels for the mesons, so the coupling is constrained by kaon decay



Results - Equal couplings

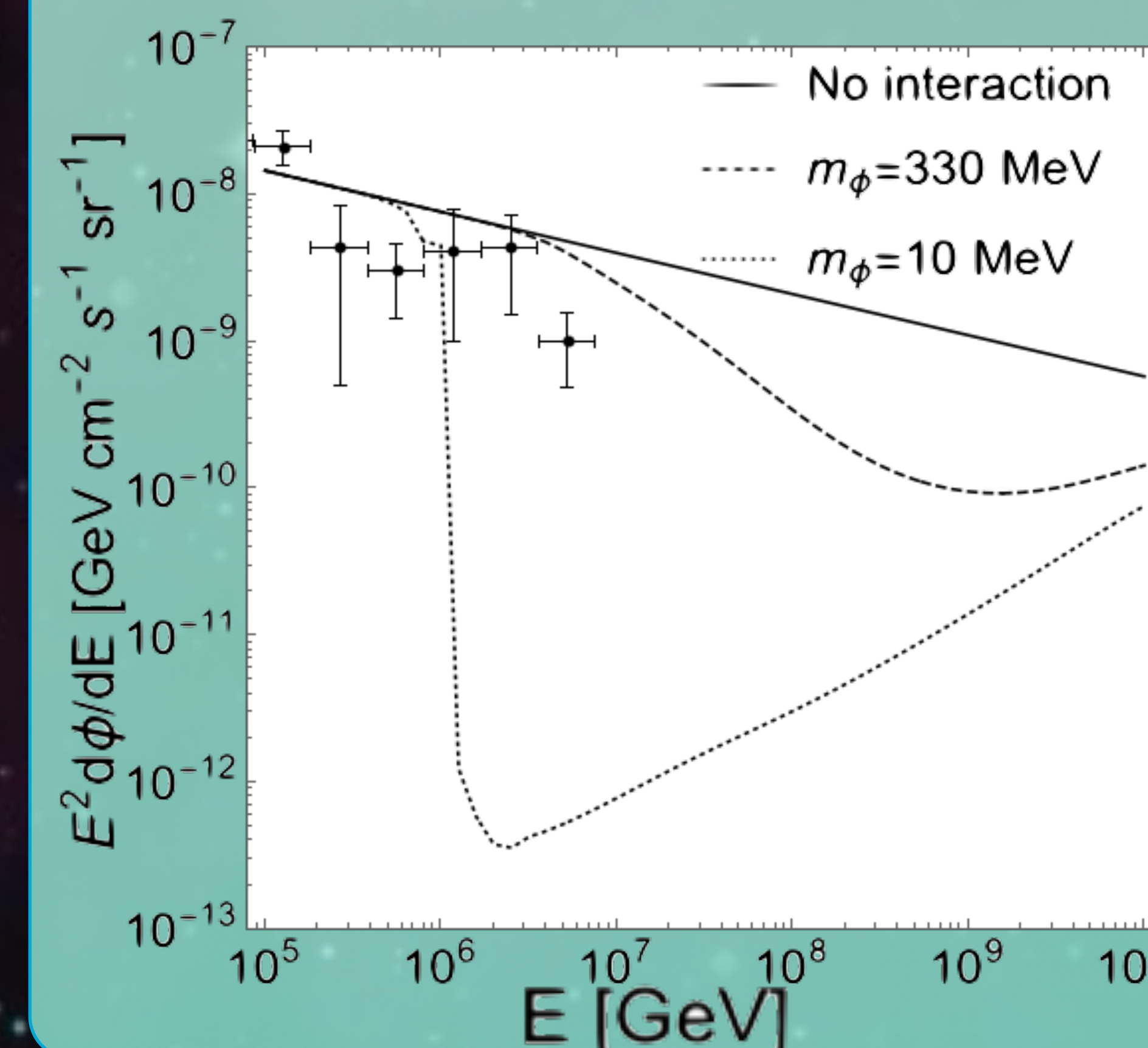
- If all neutrinos couple equally, kaon decay constrains the interaction: the effects are only important for high masses ($m_{\phi}, m_s \geq 250$ MeV)

- Absorption important for ultrahigh energy neutrinos, like cosmogenic neutrinos



Results - Tau coupling

- Since kaon cannot decay into tau leptons, the tau coupling is practically unconstrained, so we can choose smaller masses ($m_{\phi}, m_s \simeq 10$ MeV)



- Absorption important for PeV neutrinos from astrophysical sources
- We analyzed the effect on a benchmark power law

References

- D.F.G. Fiorillo, G. Miele, S. Morisi, N. Saviano, *Phys.Rev.D* 101 (2020) 8, 083024

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