

# Charge and energy calibration of the ProtoDUNE-SP detector using cosmic ray muons

Ajib Paudel, On behalf of the DUNE collaboration

Neutrino 2020 Poster ID:53

## INTRODUCTION

biggest LArTPC to date

ProtoDUNE-SP

Liquid Argon Time Projection Chamber (LArTPC) built at CERN neutrino platform.

The active volume is 6 m high, 7 m wide and 7.2 m along the drift direction.

Consists of 2 drift volumes of 3.6m drift length each.

Exposed to H4 beamline at CERN with known particle species of kaons(K<sup>+</sup>), positrons(e<sup>+</sup>), pions(π<sup>+</sup>) and protons(p).

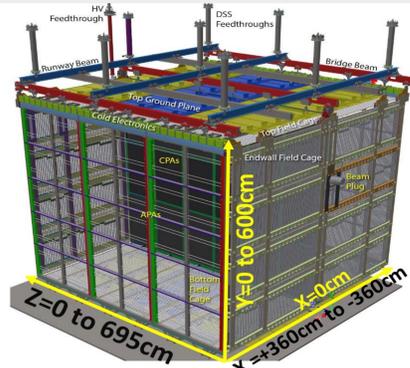


Fig 1: Major components of ProtoDUNE-SP

### LArTPC Calibration

Motivation for calibration:

- Hadronic cross-section on argon
- Detector response

Why Cosmics?

- ~10,000 muons/(minute.m<sup>2</sup>) reaches earth's surface.
- Cosmic ray muons cover the entire TPC in a short time (a few hours).
- dE/dx vs Kinetic energy well known

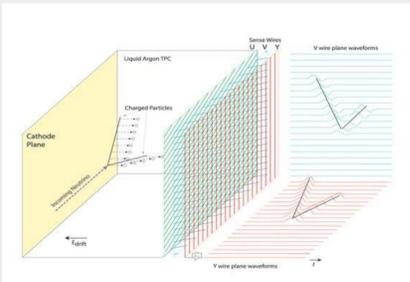


Fig 2: LArTPC and time vs wire readout

### dQ/dx non-uniformity

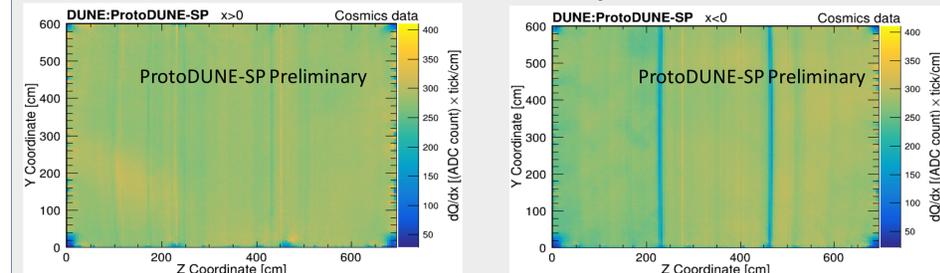


Fig 3: dQ/dx distribution for YZ plane for x>0 drift volume, left plot and x<0 drift volume, right plot

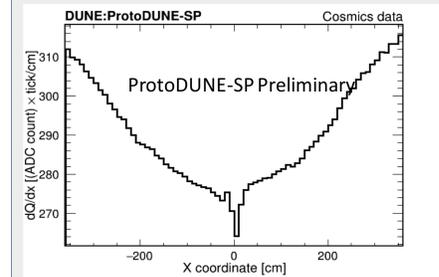


Fig 4: dQ/dx distribution vs drift coordinate

The charge deposition per unit length (dQ/dx) is affected by several factors including:

- Space-Charge Effect (SCE)
- Recombination Effect
- Electron attenuation
- Diffusion
- Electronics gain variation

The low dQ/dx stripes in YZ distribution for x<0 region is due to the presence of electron diverters at APA boundaries for x<0 (which moves charge away from the boundaries), while electron diverters are not present in x>0 region.

SCE and electronics gain variation are calibrated out using dedicated calibrations.

Detector response for the energetic through-going cosmic muons used as a data-driven correction, to remove remaining non-uniformities.

We are using a method like the MicroBooNE detector [1], which is discussed in next steps.

## METHODOLOGY

Charge calibration(dQ/dx calibration)

Using energetic through-going cosmic ray muons, charge calibration is carried out in two steps:

Step 1: YZ calibration

Divide YZ plane into 5 × 5 cm<sup>2</sup> bins

The YZ correction factor is:

$$C(y, z) = \frac{\text{median } dQ/dx \text{ for each bin}}{\text{median } dQ/dx \text{ for the entire TPC}}$$

Final calibrated dQ/dx is:

$$(dQ/dx)_{\text{calibrated}} = N(Q) \cdot C(x) \cdot C(y, z) \cdot (dQ/dx)_{\text{reconstructed}}$$

Step 2: X calibration

Divide X coordinate into 5cm bins, apply YZ correction to each hit

The X correction factor is:

$$C(x) = \frac{\text{median } dQ/dx \text{ for each bin}}{\text{median } dQ/dx \text{ for the entire TPC}}$$

Charge normalization constant,

$$N(Q) = \frac{\text{median } dQ/dx \text{ at the anode}}{\text{median } dQ/dx \text{ for the entire TPC}}$$

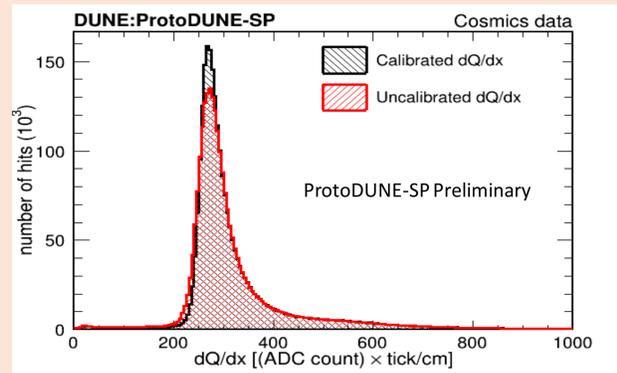


Fig5: dQ/dx distribution before and after charge calibration

Energy Scale Calibration (dE/dx calibration)

Stopping cosmic ray muons are used for energy scale calibration.

Landau-Vavilov [2] theory predicts stopping muon dE/dx vs kinetic energy(or residual range).

Calibrated dQ/dx  $\xrightarrow{\text{Modified Box Model } C_{\text{cal}}}$  dE/dx, C<sub>cal</sub> is a free parameter in the χ<sup>2</sup> minimization.

Modified Box Model [3]

$$\left(\frac{dE}{dx}\right)_{\text{calibrated}} = \exp\left(\frac{\left(\frac{dQ}{dx}\right)_{\text{calibrated}} \beta' W_{\text{ion}}}{C_{\text{cal}} \rho \ell}\right) - \alpha \left(\frac{\rho \ell}{\beta'}\right)$$

where

C<sub>cal</sub> = Calibration constant used to convert ADC values to number of electrons,

W<sub>ion</sub> = 23.6 × 10<sup>-6</sup> MeV/electron (the work function of argon),

ℓ = E field based on the measured space charge map,

ρ = 1.38 g/cm<sup>3</sup> (liquid argon density at a pressure of 124.106 kPa),

α = 0.93, and

β' = 0.212 (kV/cm)/(g/cm<sup>2</sup>)/MeV.

The last two parameters were measured by the ArgoNeuT experiment at an operational electric field strength of 0.481 kV/cm.

Calibration factor for ProtoDUNE-SP is measured to be C<sub>cal</sub>=(5.3955±0.0035)e-3 ADC/electrons.

References:

[1] <https://arxiv.org/abs/1907.11736>

[2] Particle Data Group

[3] ArgoNeuT collaboration, R. Acciarri et al., A Study of Electron Recombination Using Highly Ionizing Particles in the ArgoNeuT Liquid Argon TPC, JINST 8 (2013) P08005, [1306.1712].

## RESULTS

dE/dx vs residual range for Stopping cosmic muons

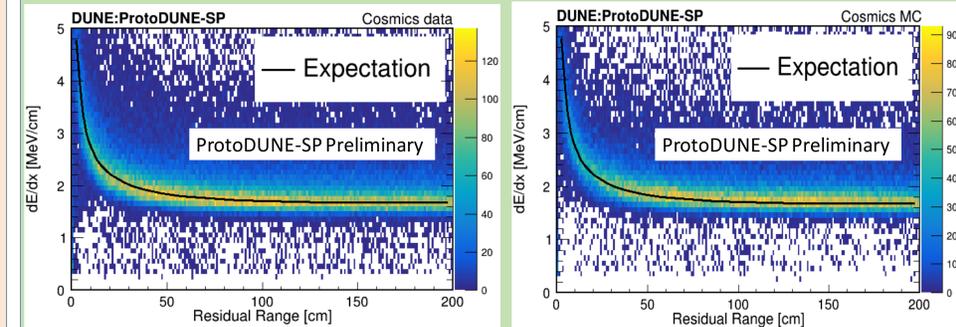


Fig 7: dE/dx vs residual range, ProtoDUNE-SP data, left and ProtoDUNE-SP simulation, right (Expectation is the Modified Box Model)

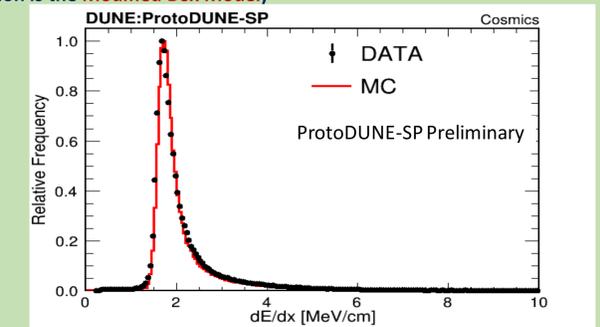


Fig 8: calibrated dE/dx for stopping muons

Cosmic Muon Based calibration constants applied to beam protons and muons:

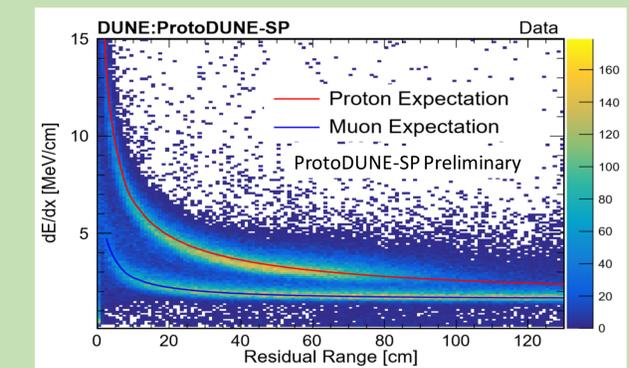


Fig 9: Proton-muon separation

Cosmic muon-based calibration factors has been used for test beam particles including pions, protons, muons, kaons and positrons. The agreement with theoretical prediction looked good.

Fig 8 shows good agreement between Monte-Carlo simulation dE/dx and data dE/dx for stopping muons.

Fig 9 shows, how well we can separate 1GeV/c beam protons and muons in ProtoDUNE-SP.

## SUMMARY:

- Cosmic muon calibration shows good results for test-beam particles as well.
- Shows excellent proton-muon separation.
- The result demonstrates high quality of the ProtoDUNE-SP detector.