

Detecting Supernova Neutrinos using the DUNE Photon Detection System



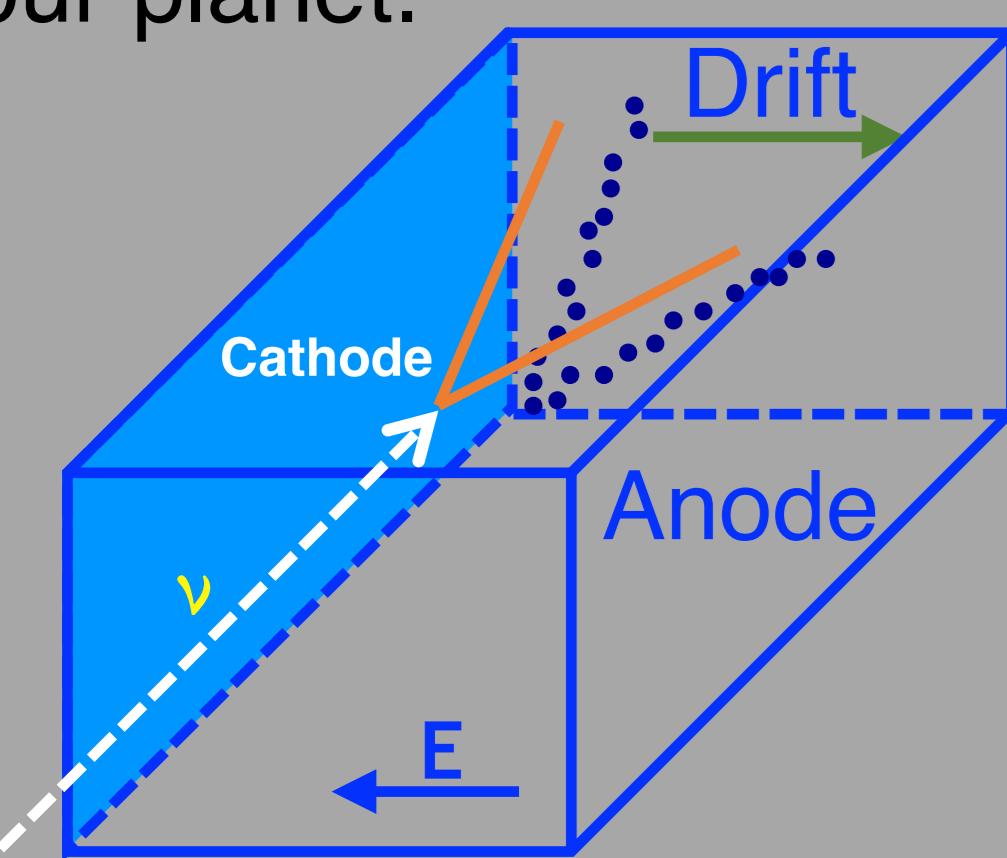
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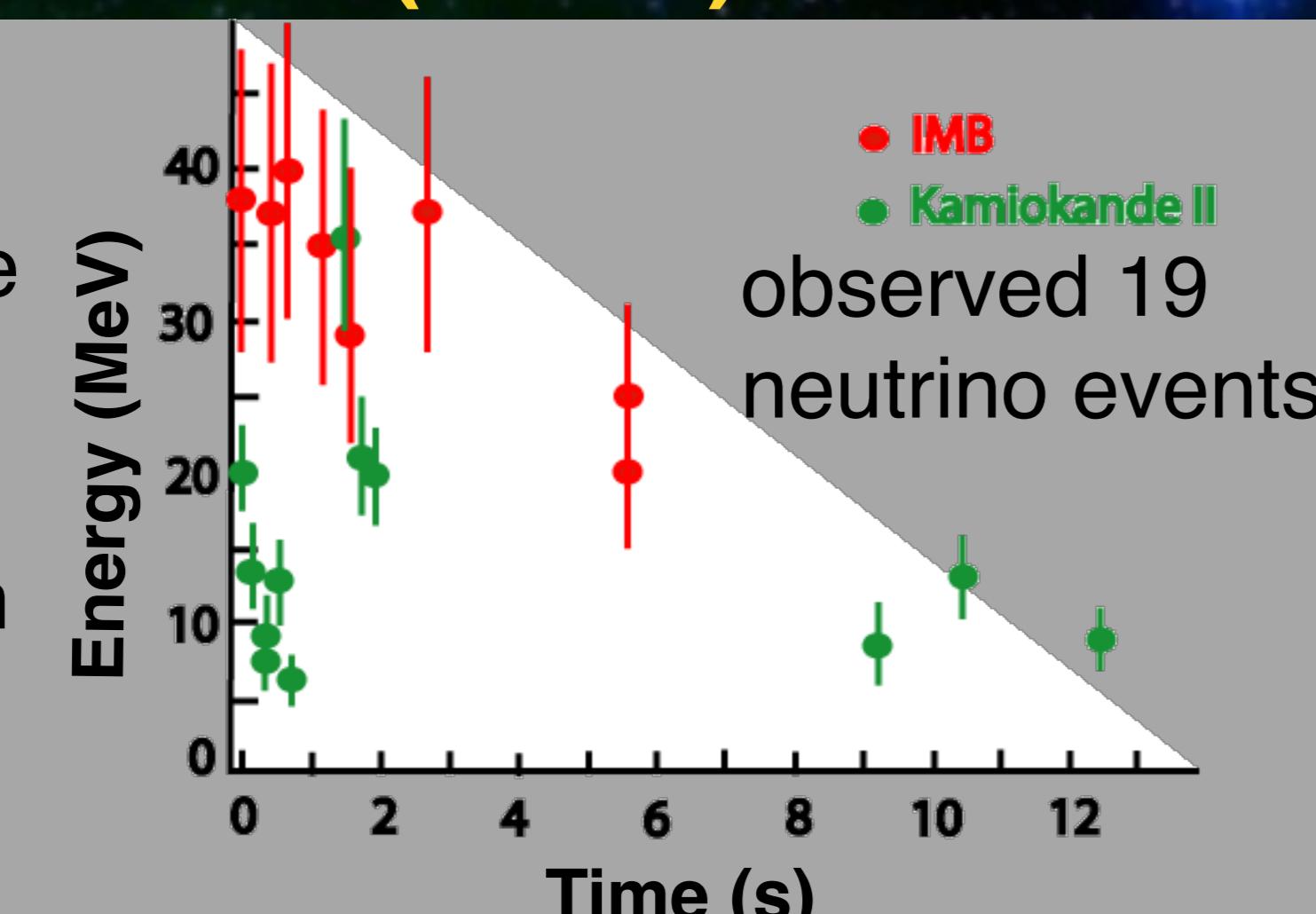
1. Supernova Neutrino Bursts (SNBs) and DUNE

- Neutrinos of all flavors, few tens of MeV energy were produced from a core-collapse Supernova (SN).

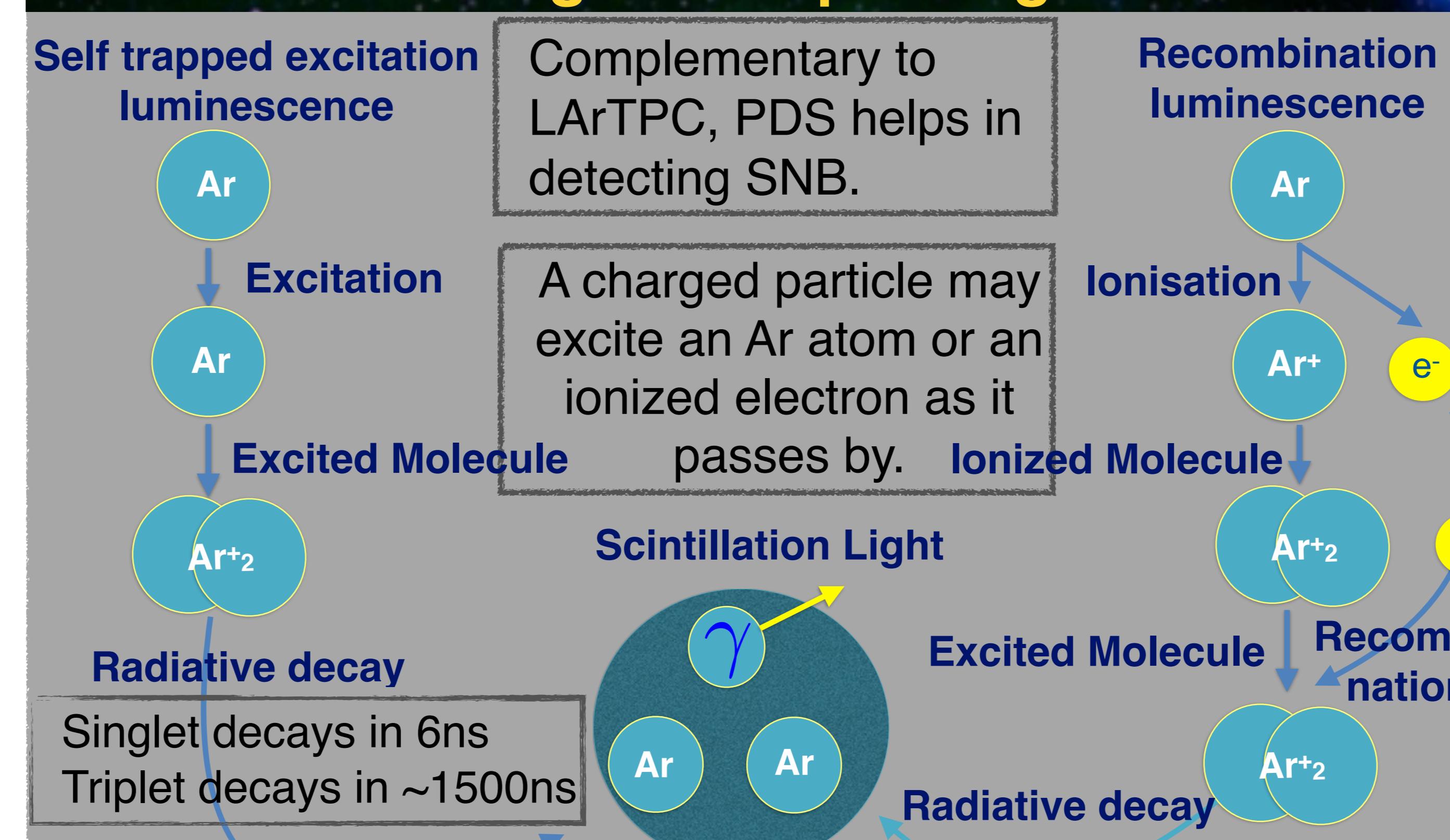
- Neutrinos were observed from SN1987A, 50 kpc away from our planet.



- DUNE's Liquid Argon Time Projection Chamber (LArTPC) detects low energies supernova neutrino events.
- DUNE uses ARAPUCA¹ photon detectors to trap photons of a certain wavelength.

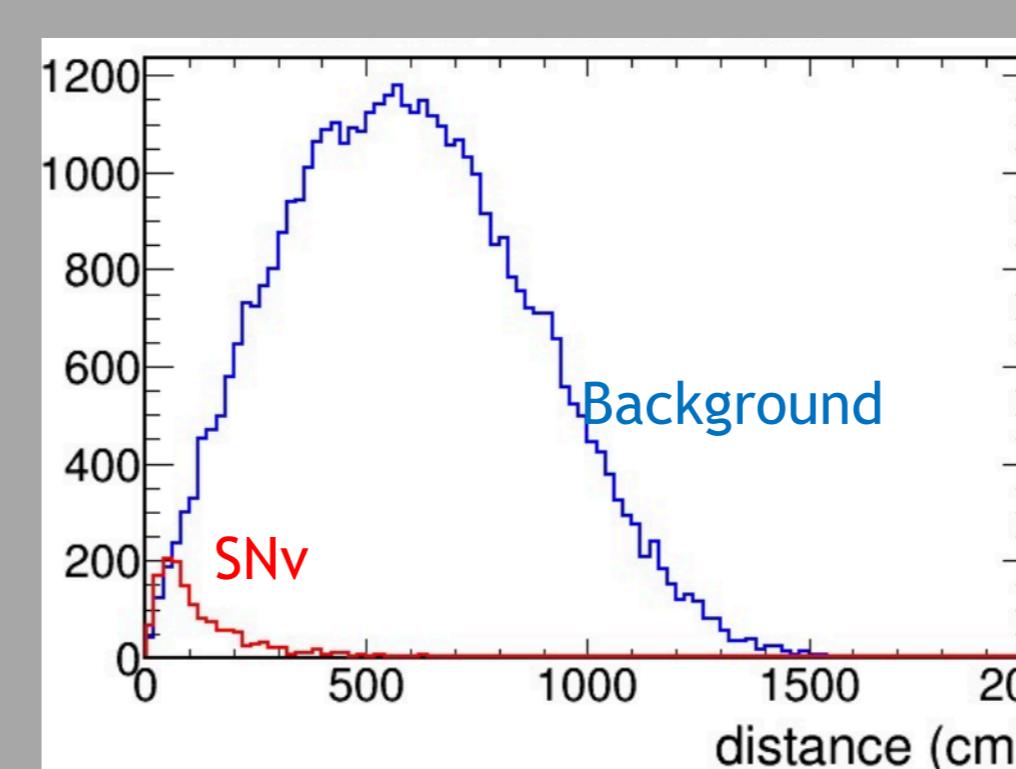
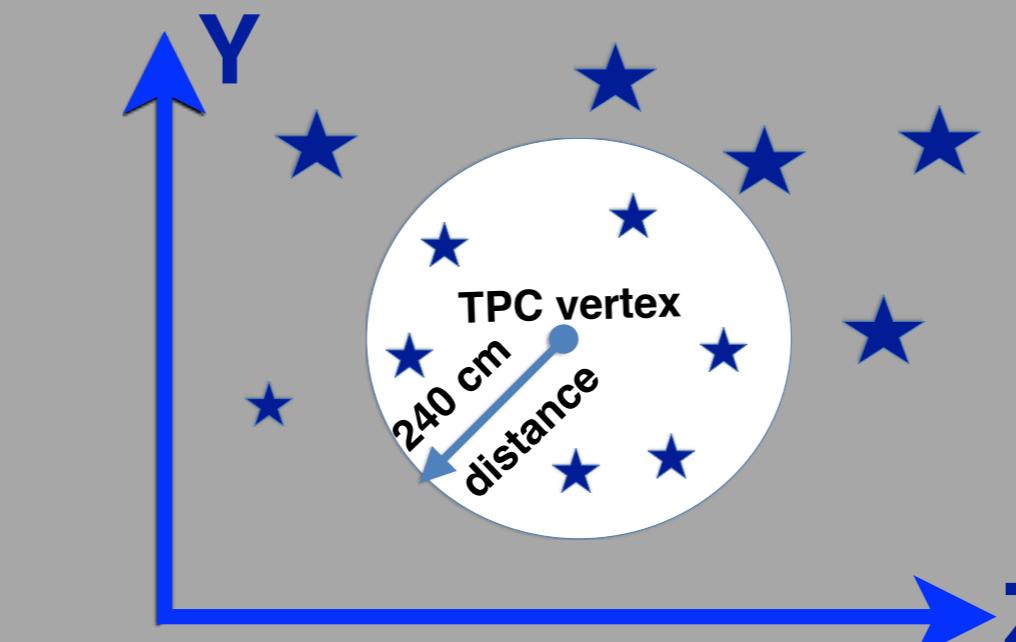


2. Scintillation Light in Liquid Argon



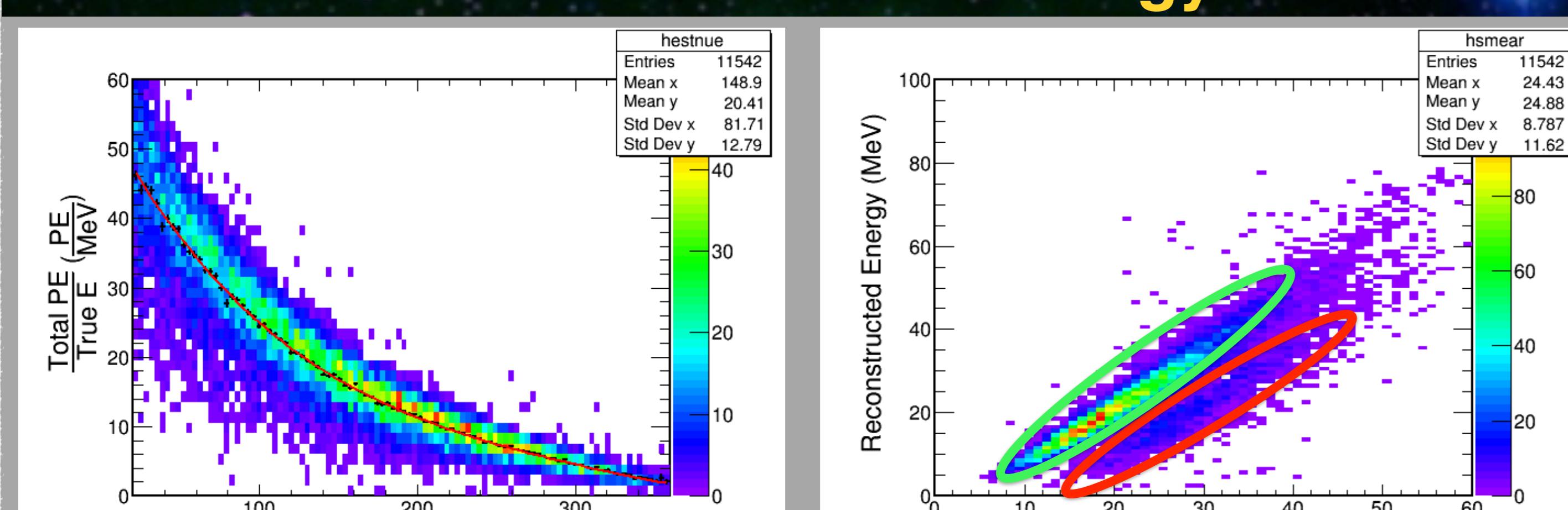
5. Matching PDS and TPC Activity

- **Flash Matching:** Coincidence between PDS flash and TPC positions will reduce the rate of uncorrelated background in the two systems.



- **Matching PDS and TPC events requires:**
 - The flash be within one drift time before the TPC time.
 - Vertex reconstruction is within 240 cm for PDS and TPC event in the Y-Z plane.
 - Consider the largest flash that remains as the match.

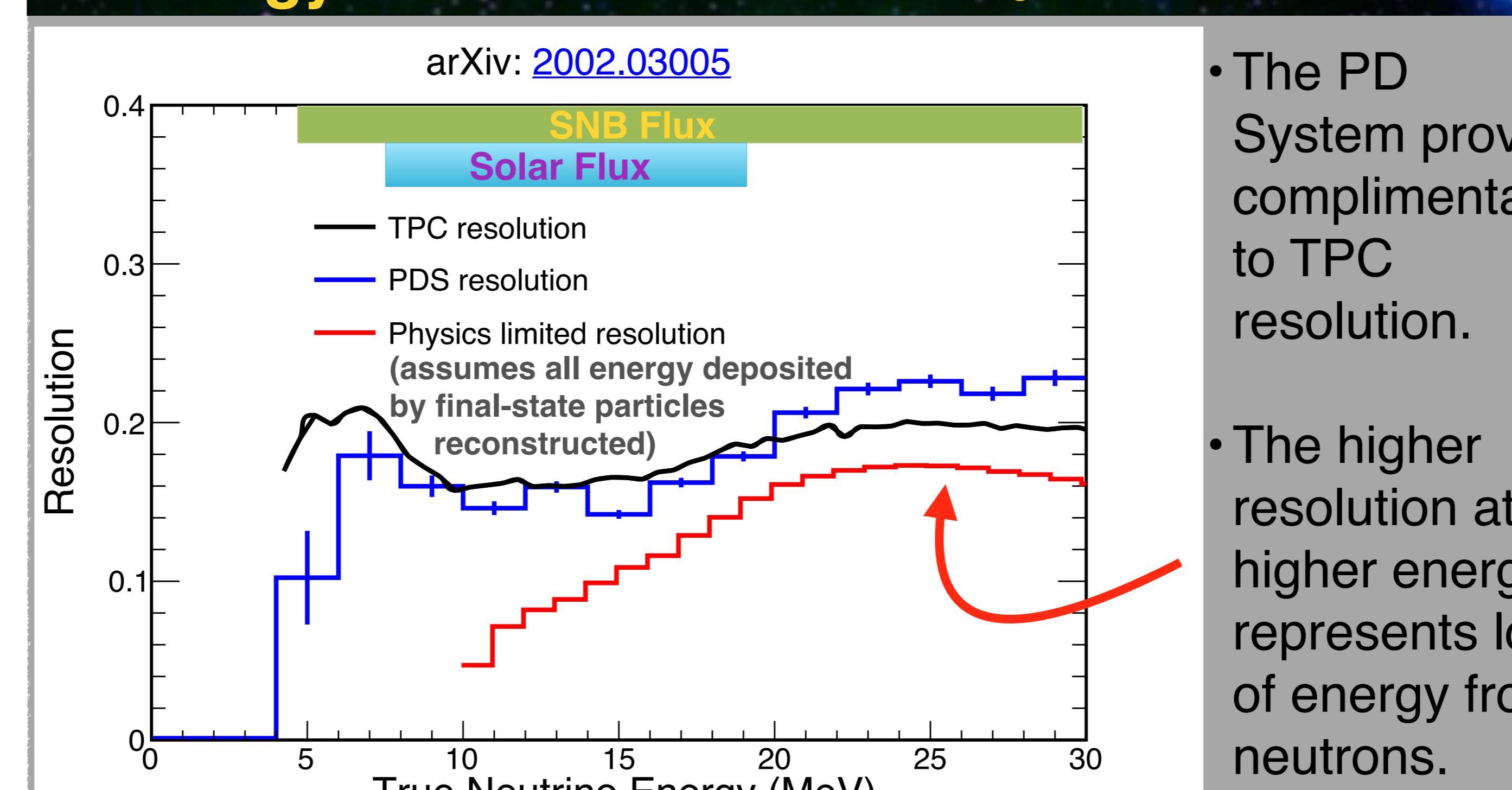
7. Reconstruction of Neutrino Energy



Rayleigh scattering attenuates the observed light yield as a function of the distance to the PDS detectors.

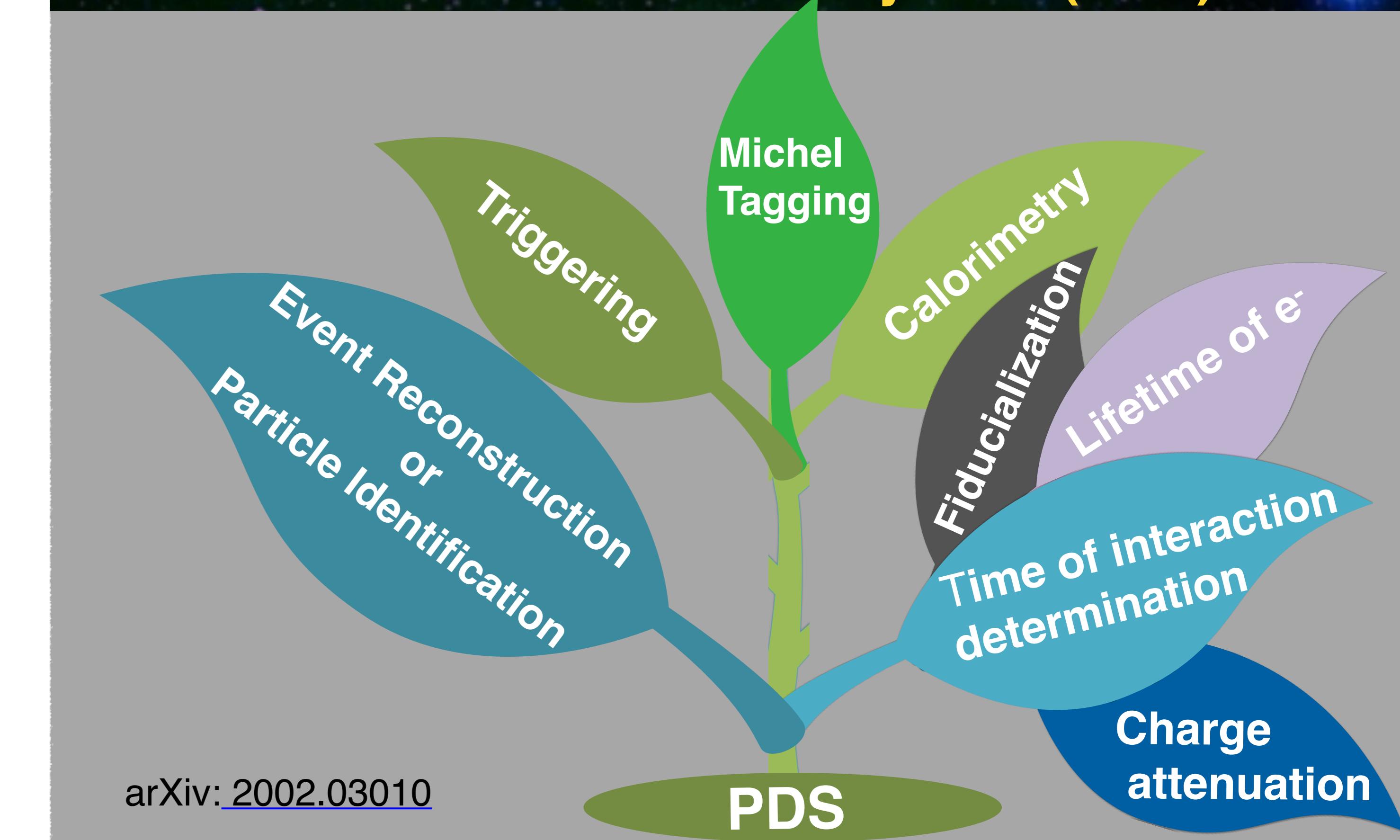
Two populations: events **with** and **without** neutron emission.

8. Energy Resolution for Low ν_e CC Interaction



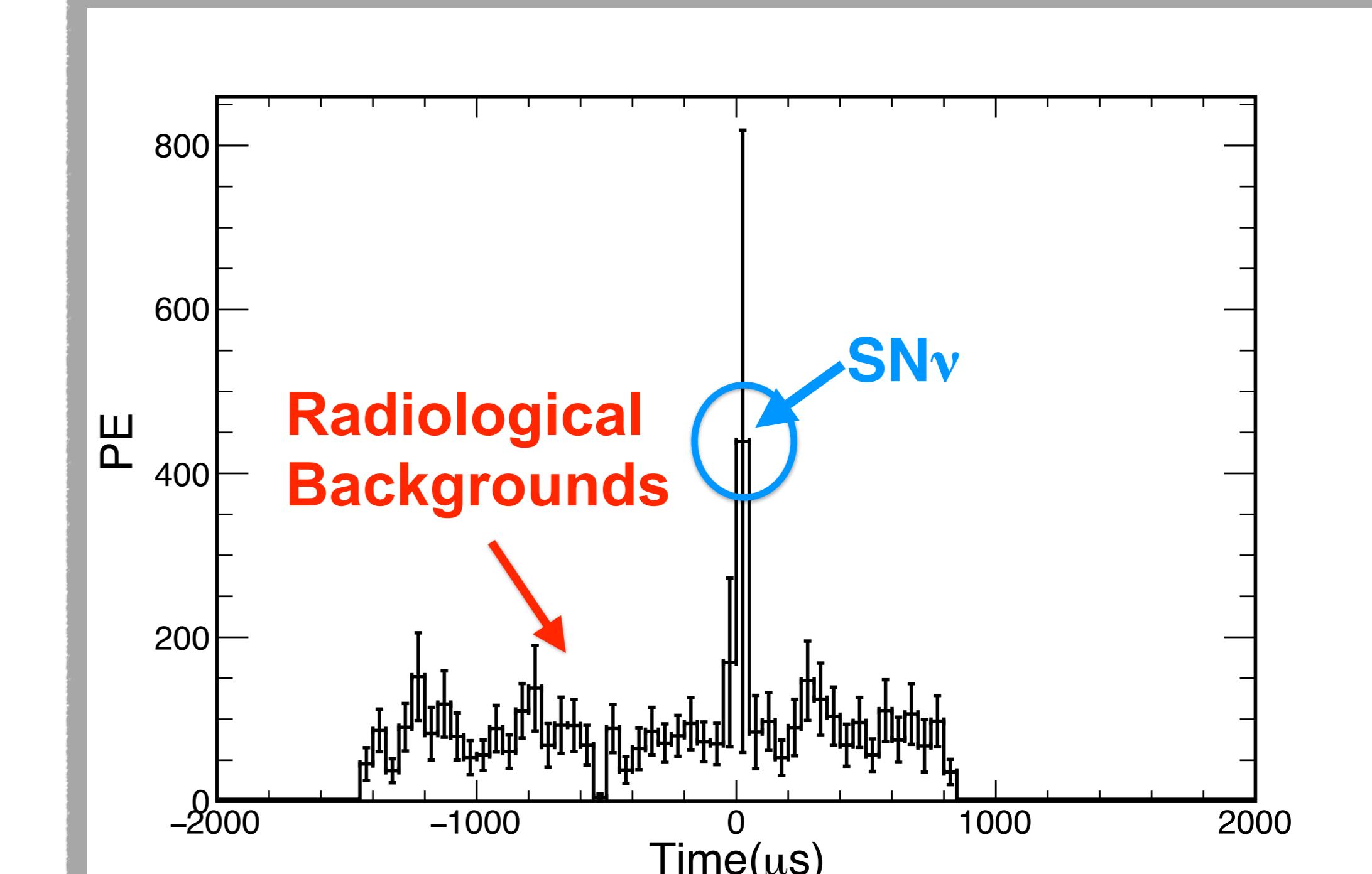
- The PD System provide complimentary to TPC resolution.
- The higher resolution at higher energy represents loss of energy from neutrons.

3. Goal of Photon Detection System (PDS)



arXiv: 2002.03010

6. Supernova Neutrino Calorimetry



SN neutrinos are embedded with lots of radiological backgrounds and signal peaked at time zero micro second.

9. Summary

