

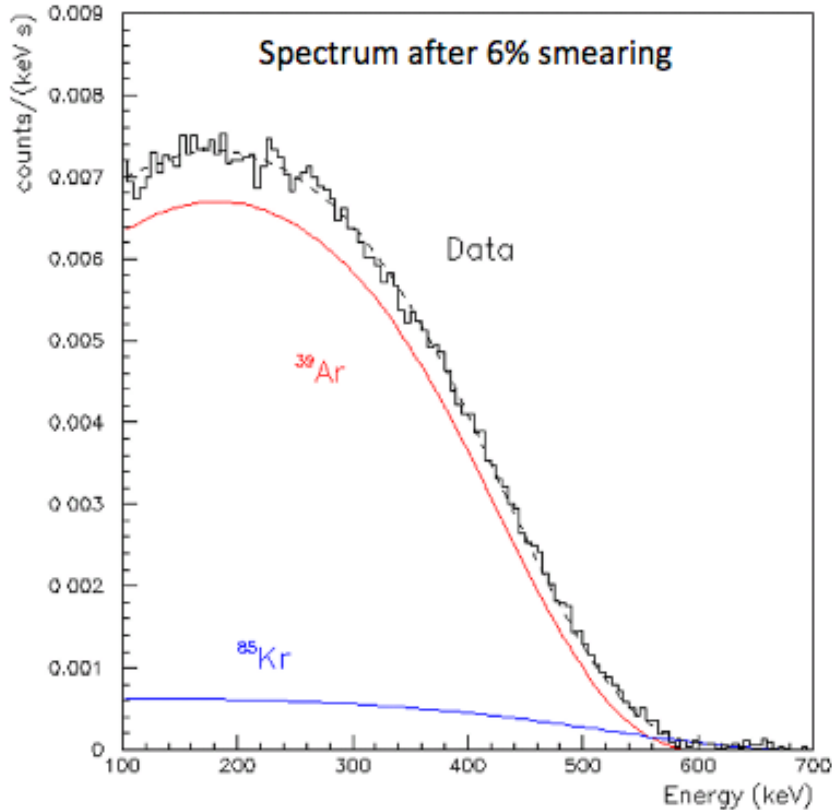
# Reconstruction of $^{39}\text{Ar}$ Beta Decays at ProtoDUNE-SP

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Colorado State University

*ProtoDUNE Sim/Reco Meeting*  
*February 13<sup>th</sup>, 2019*

- ◆ Many calibrations done at MicroBooNE utilize cosmic rays (e.g. electron lifetime)
  - MicroBooNE on surface → **4000 cosmics/second**
- ◆ Not a reliable option at DUNE FD due to being almost a mile underground
  - DUNE FD: **4000 cosmics/day** (and **20 Michels/day**)
  - ... and this is for an entire 10 kt module!
  - Corresponds to 5 cosmics/day/m<sup>3</sup>
- ◆ Cosmics can still help, but need alternative charge sources for calibrations
- ◆ Plenty of <sup>39</sup>Ar **beta decays** at DUNE FD (O(50000) per readout) – good option that should be explored for DUNE, studied first at MicroBooNE/ProtoDUNE

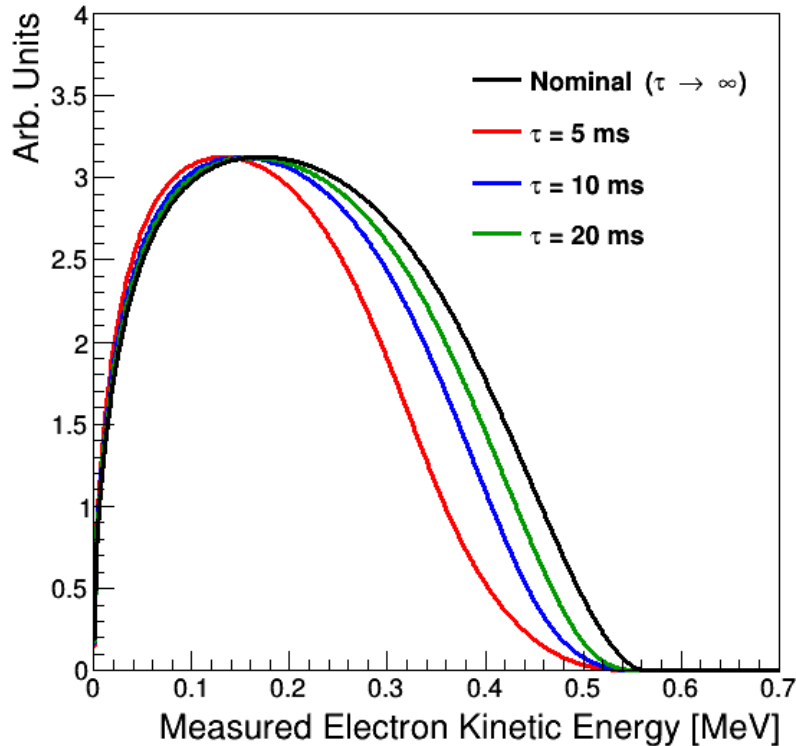
# $^{39}\text{Ar}$ Beta Decays



**Benetti et al., “Measurement of the specific activity of Ar-39 in natural argon” (2006).**

- ◆  $^{39}\text{Ar}$  beta decay cut-off energy is 565 keV
  - Roughly **half** of the energy deposited on a single wire by a MIP at DUNE
- ◆ Several things smear observed **charge spectrum**, e.g.:
  - Electronics noise
  - Recombination fluctuations
  - Unknown location of  $^{39}\text{Ar}$  decay in drift direction
- ◆ For last point: we know decays are **uniform in x**

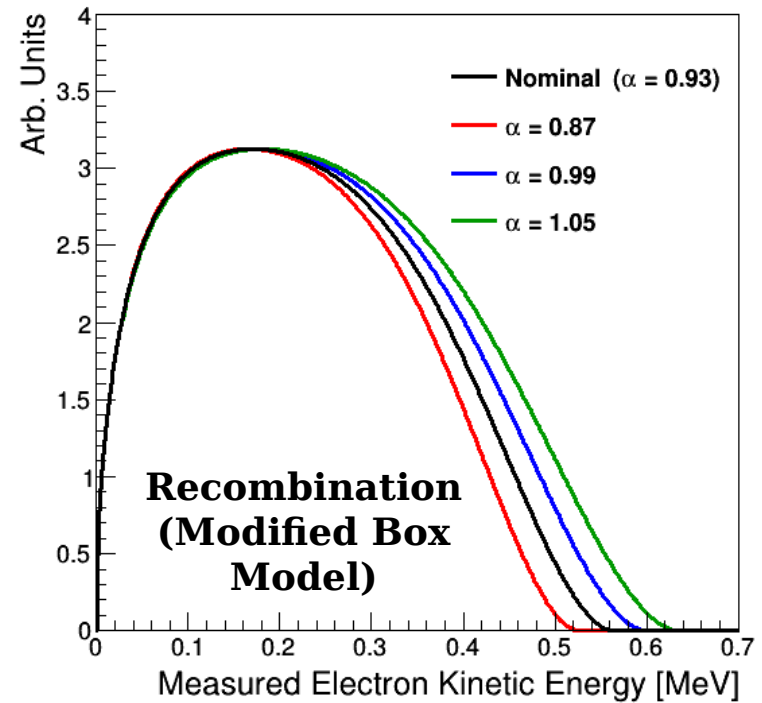
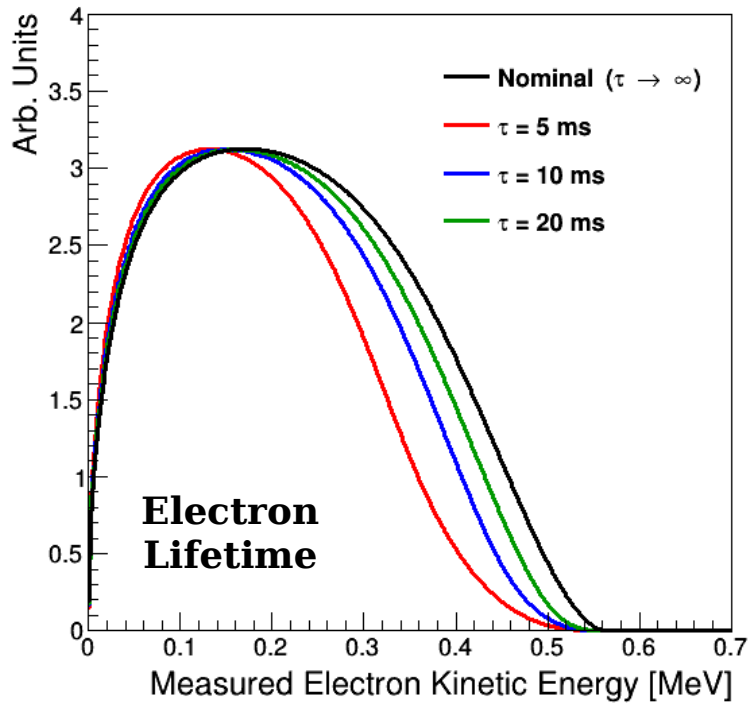
# $^{39}\text{Ar}$ Beta Decays



**Example Use Case:**  
**Fine-Grained Electron**  
**Lifetime Measurement**

- ◆  $^{39}\text{Ar}$  beta decay cut-off energy is 565 keV
  - Roughly **half** of the energy deposited on a single wire by a MIP at DUNE
- ◆ Several things smear observed **charge spectrum**, e.g.:
  - Electronics noise
  - Recombination fluctuations
  - Unknown location of  $^{39}\text{Ar}$  decay in drift direction
- ◆ For last point: we know decays are **uniform in x**

# e<sup>-</sup> Lifetime vs. Recomb.



- ◆ Electron lifetime and recombination both impact spectrum, but in different ways → largely separable
- ◆ Noise also leads to smearing, but this can be measured very precisely with noise data

# *Studies at MicroBooNE*

## Study of Reconstructed $^{39}\text{Ar}$ Beta Decays at the MicroBooNE Detector

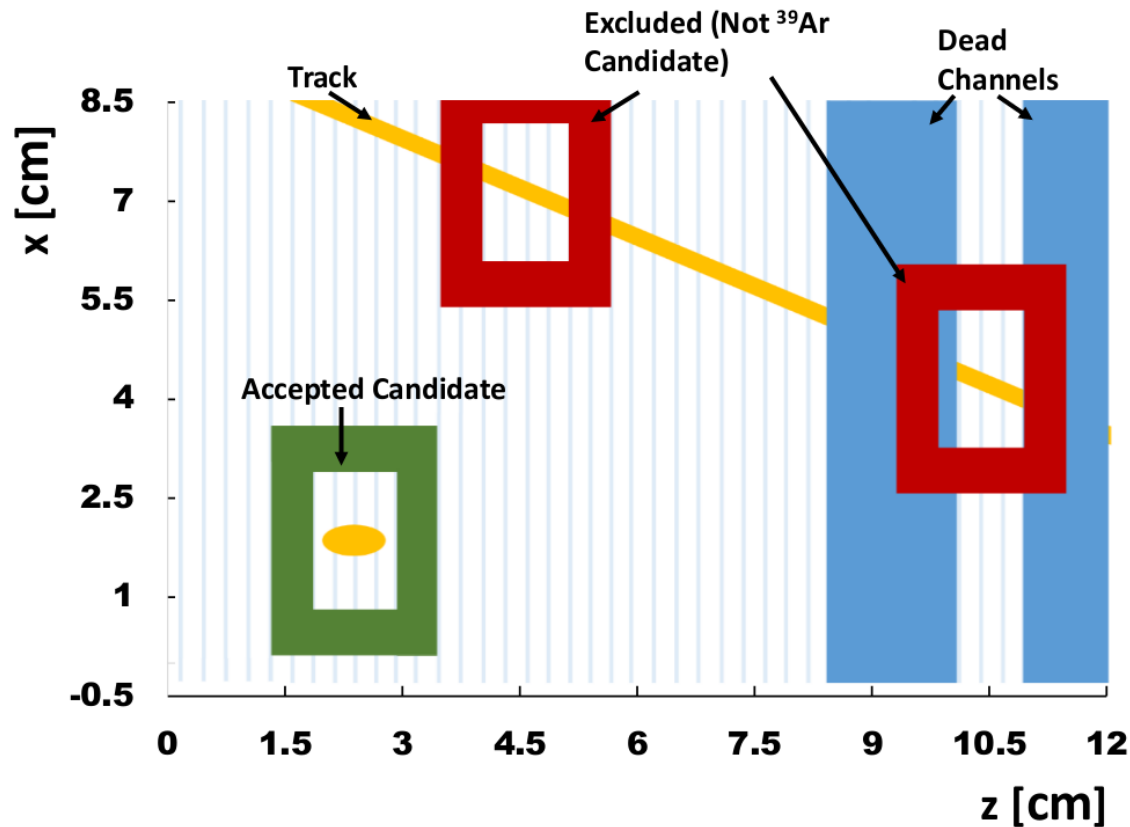
The MicroBooNE Collaboration\*

June 28, 2018

In atmospheric argon,  $^{39}\text{Ar}$  beta decays occur at a rate of roughly one Becquerel per kilogram; as a result, large liquid argon time projection chambers (LArTPCs) see plentiful amounts of these decays in each event readout window. These  $^{39}\text{Ar}$  beta decays can provide a variety of different uses in LArTPC experiments. They allow for the study of reconstructing point-like ionization charge deposition in the detector, which is relevant for the reconstruction of supernova neutrino interactions. The point-like topology and well-known energy spectrum of  $^{39}\text{Ar}$  beta decays also provides a unique handle for calibrations. Presented here is the first study of reconstructing  $^{39}\text{Ar}$  beta decays in the MicroBooNE LArTPC. The spectrum of reconstructed electron energies from  $^{39}\text{Ar}$  beta decays, measured using raw TPC waveforms, is found to have a similar end point as the predicted distribution. Additionally, the signal shapes of reconstructed  $^{39}\text{Ar}$  beta decay candidates are used to validate the collection plane wire field response.

- ◆ See MicroBooNE **public note** on  $^{39}\text{Ar}$  beta decays
- ◆ Will summarize main results in following slides

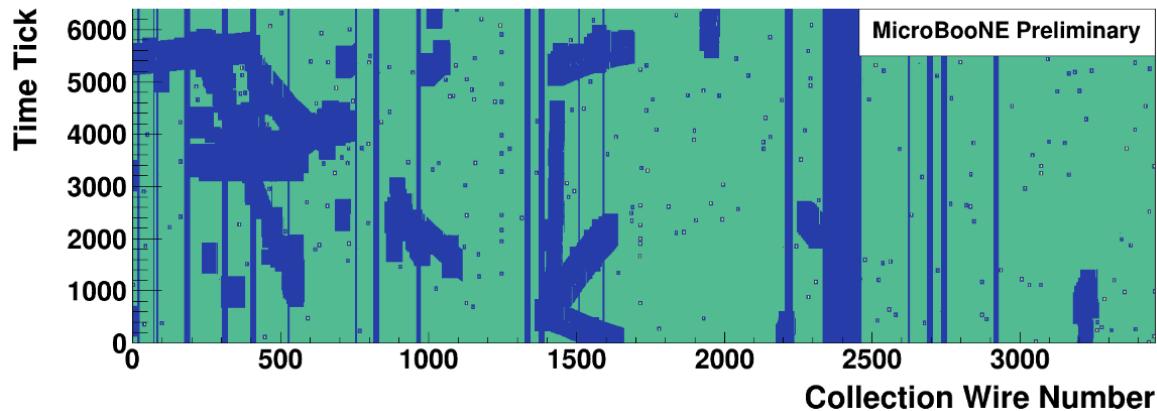
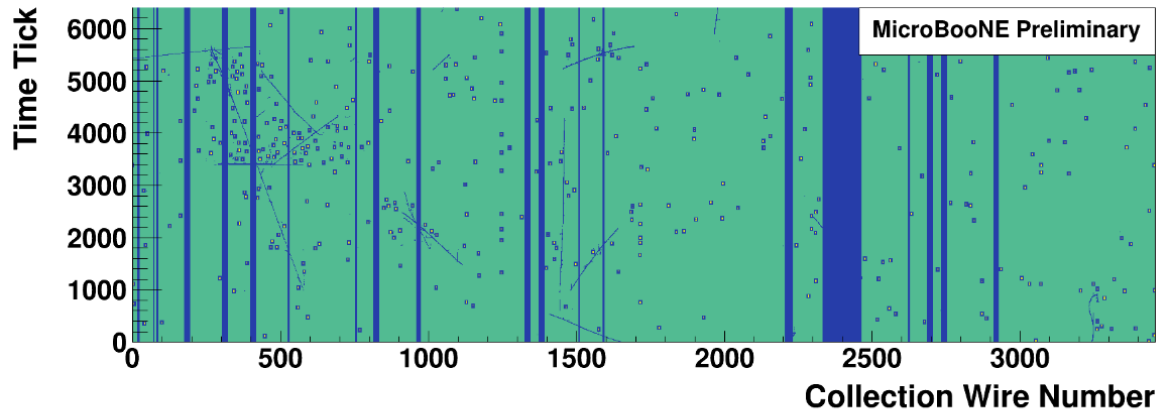
# Candidate Selection



- ◆ Select point-like topology; reject track-like topology
- ◆ Account for dead channels in selection
- ◆ Only looking at collection plane (for now)

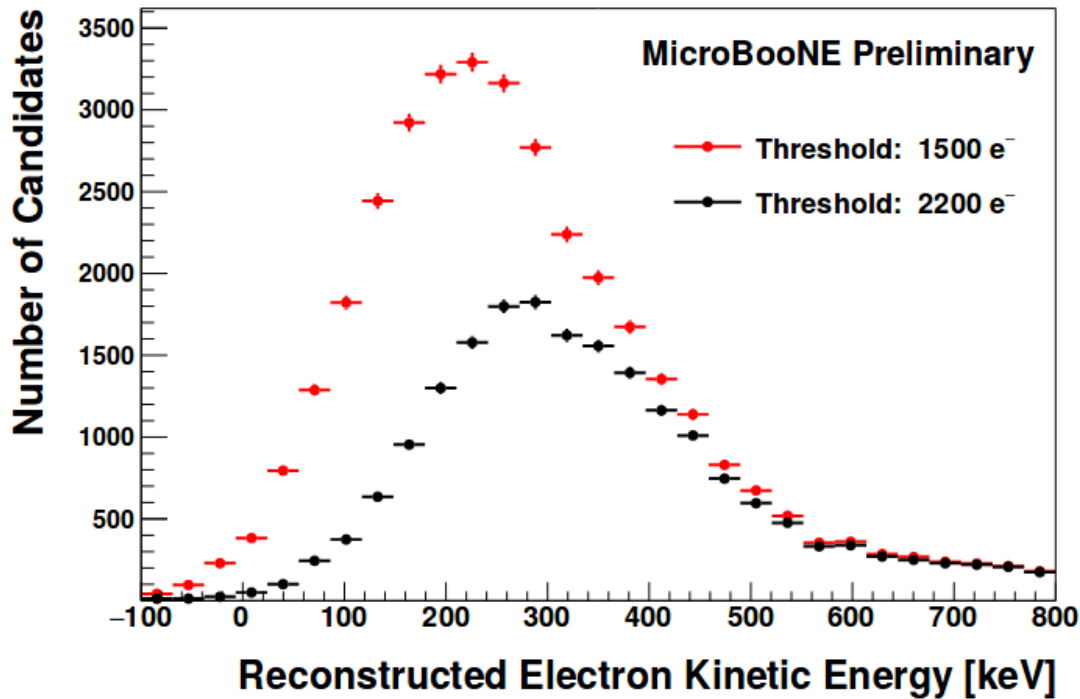


# Track Proximity Veto



- ◆ Ignore point-like activity near tracks that is more likely to be cosmogenic in nature (within  $\sim 30$  cm)

# Energy Spectrum



$$E = \frac{GI}{RK} \times Q$$

G = Gain [e<sup>-</sup>/ADC]

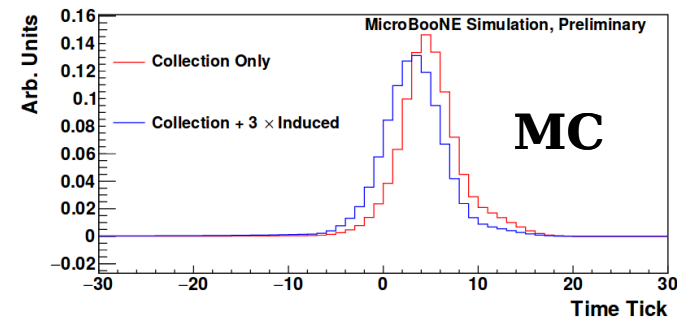
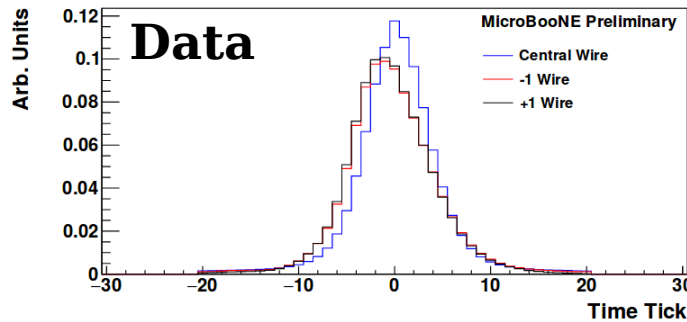
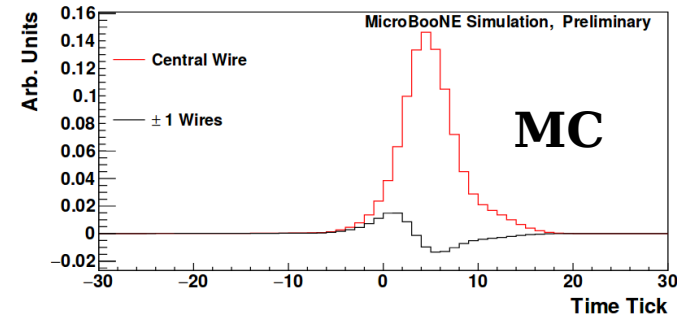
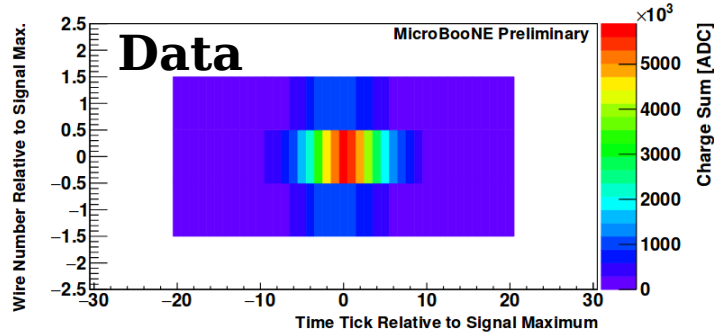
I = 0.0236 keV/e<sup>-</sup>

R = Recombination  
Factor

K = Elec. Response  
Area-to-Amp. Ratio

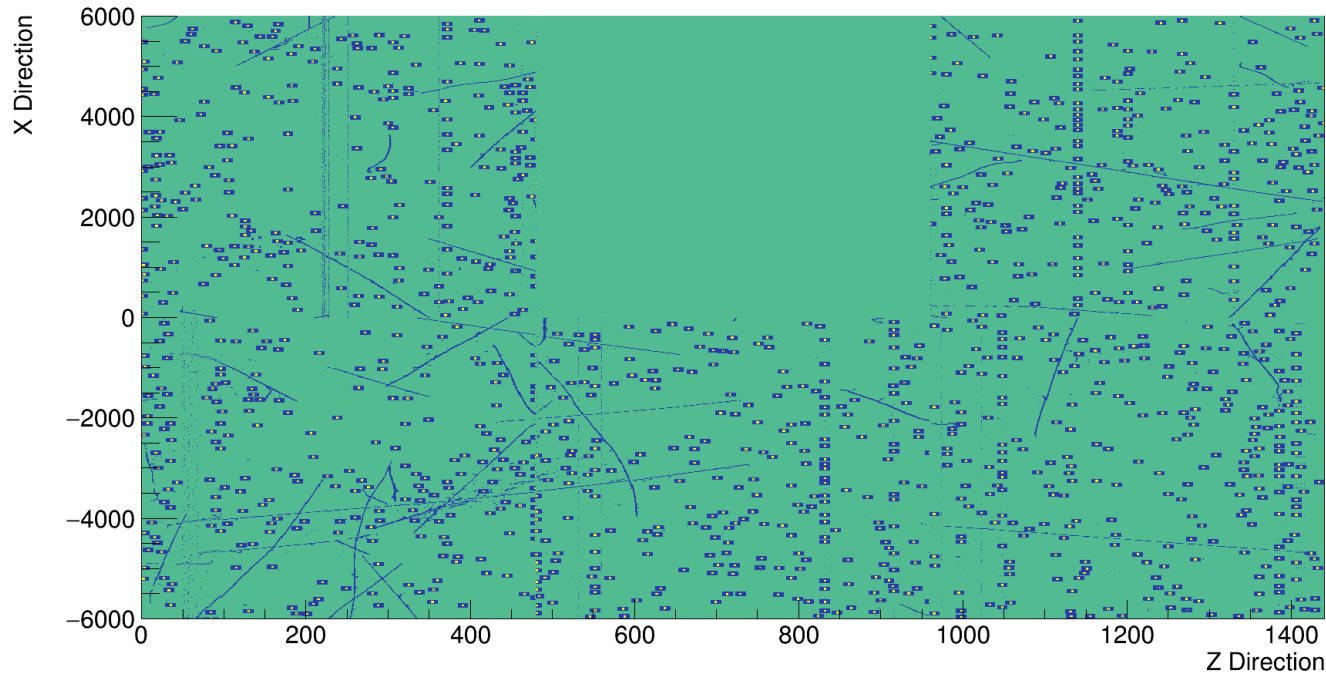
- ◆ Reconstruct energy spectrum in 3 wire by 40 time tick window around candidate peak
- ◆ Reconstructed end point in correct place (~565 keV)
- ◆ High-energy tail: cosmogenic background

# Signal Shape



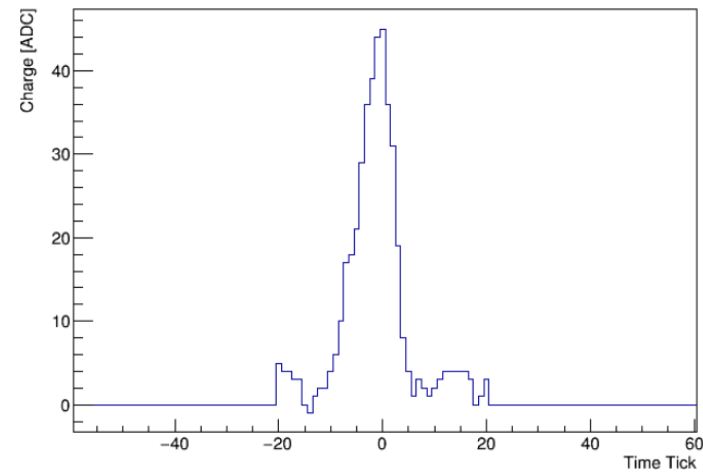
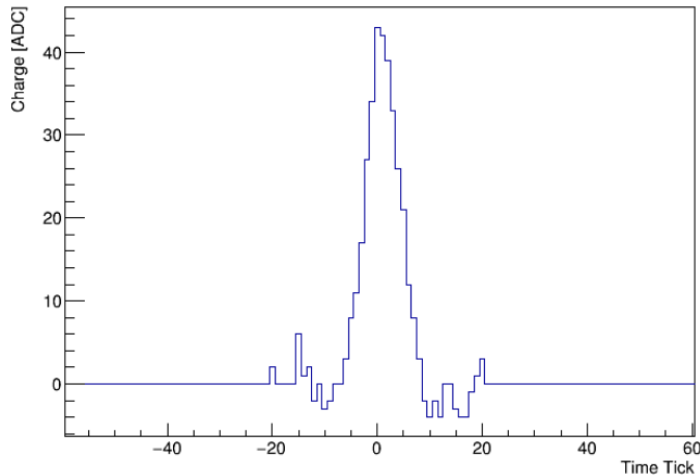
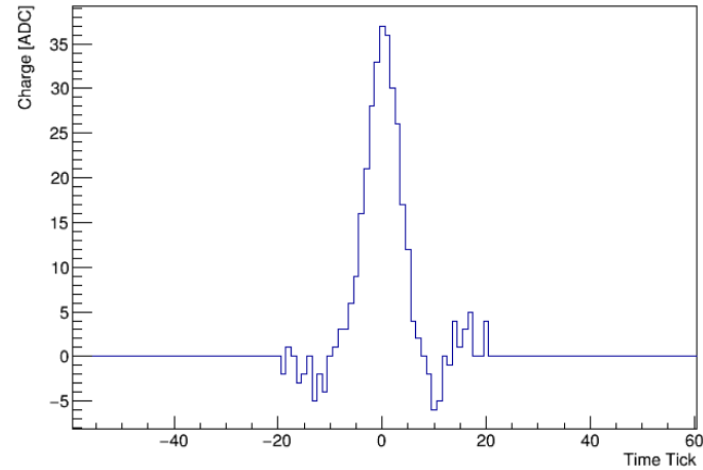
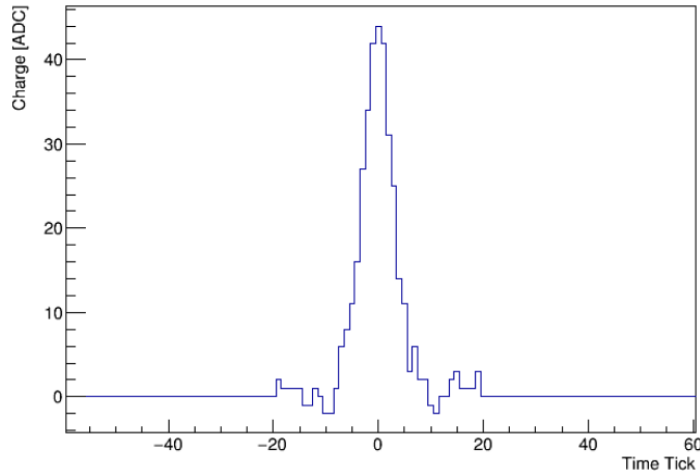
- ◆ Average signal shape reconstructed by lining up signals in time; include central **and** side wires
- ◆ Signs of induced charge effects on collection plane
- ◆ Use to study diffusion, wire field response variations

# *Studies at ProtoDUNE-SP*



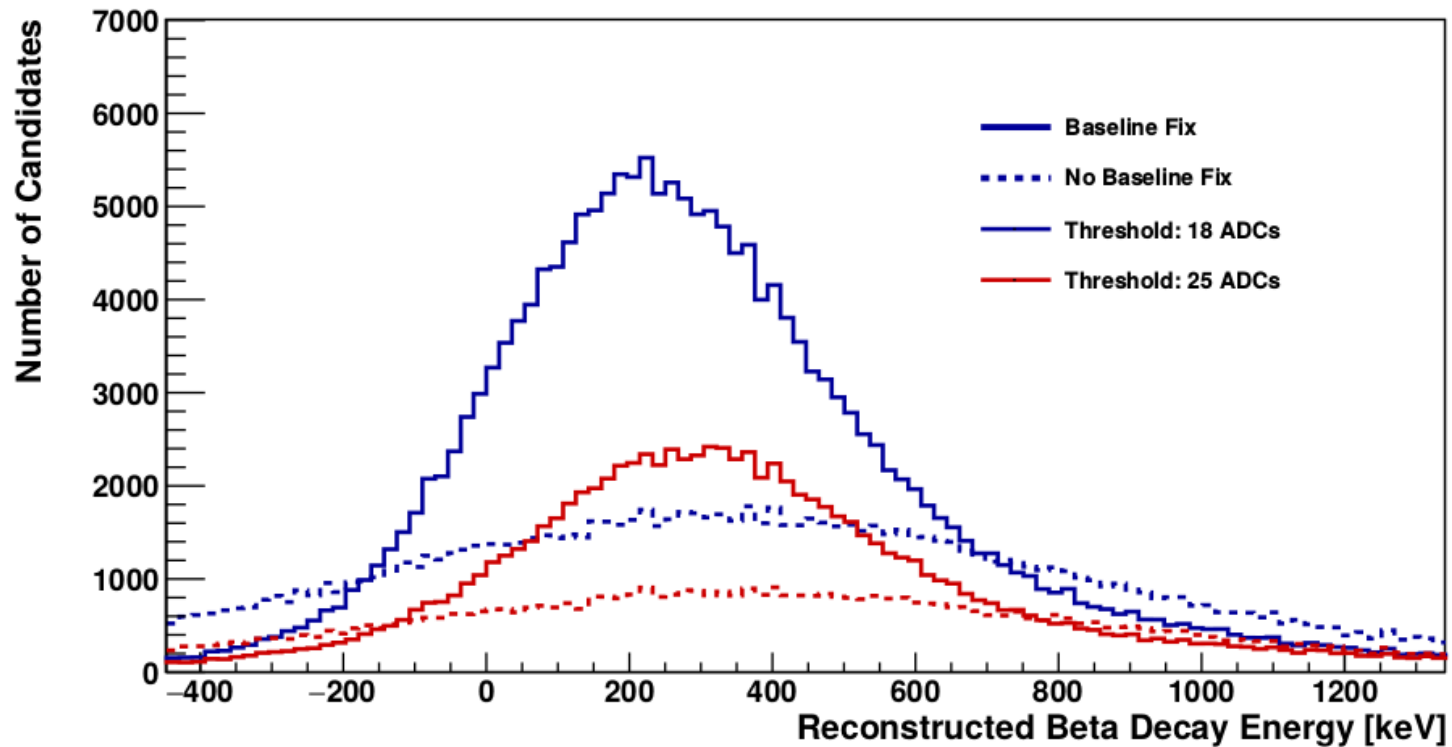
- ◆ Largely use same methodology as at MicroBooNE
- ◆ For now, not vetoing regions near tracks (next step)
- ◆ Run 5451 (missing one APA); study 150 events
- ◆ Signs of noisy wires contributing fakes (minor)

# Example Candidates



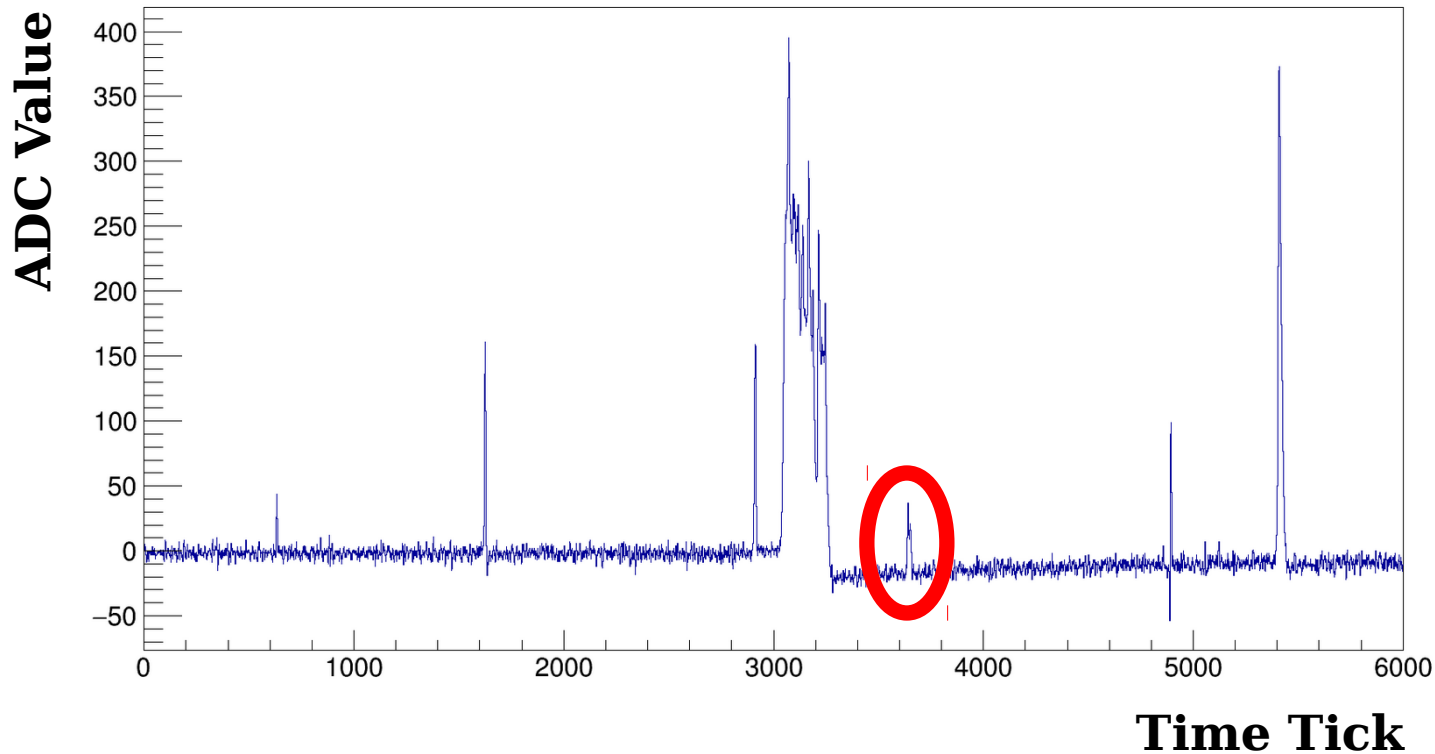
◆ Example  $^{39}\text{Ar}$  beta decay candidates (300-600 keV)

# First Spectrum



- ◆ First spectrum **looked really bad**
- ◆ Ended up being due to sagging waveforms after larger ionization signals (AC-coupled electronics)
- ◆ Moved to local baseline estimate w/ sidebands

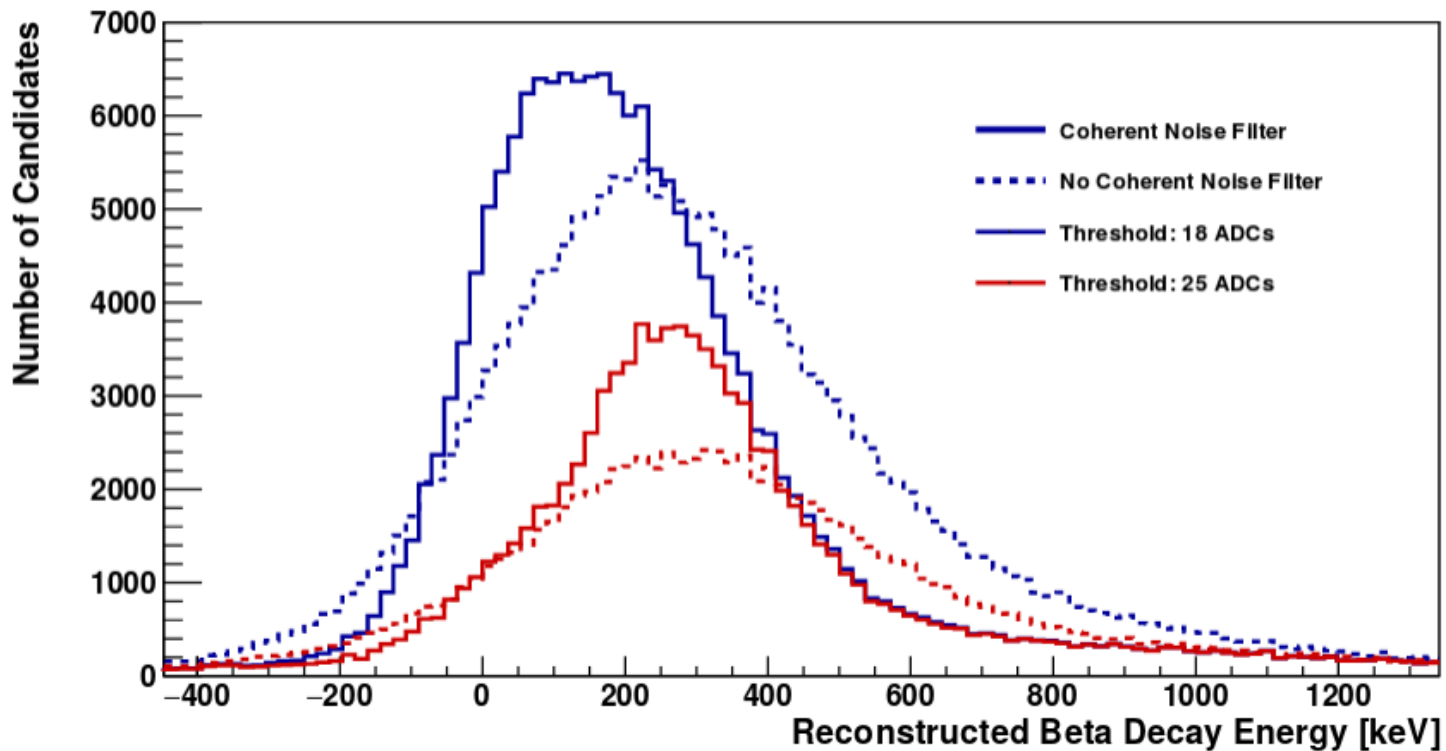
# Baseline Deviations



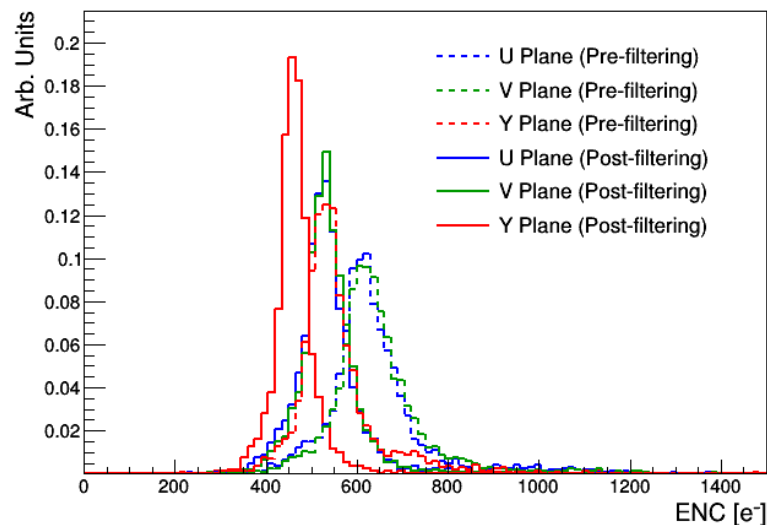
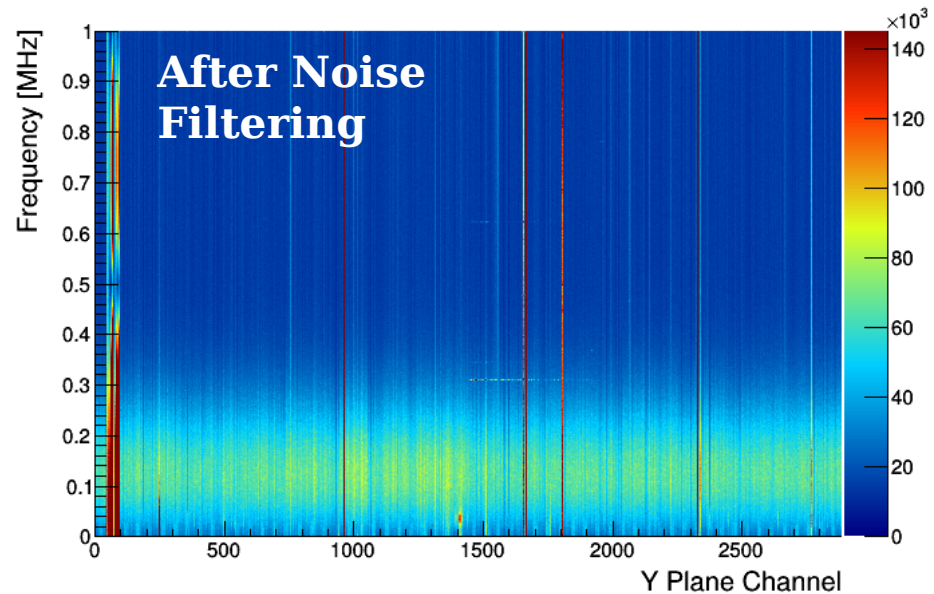
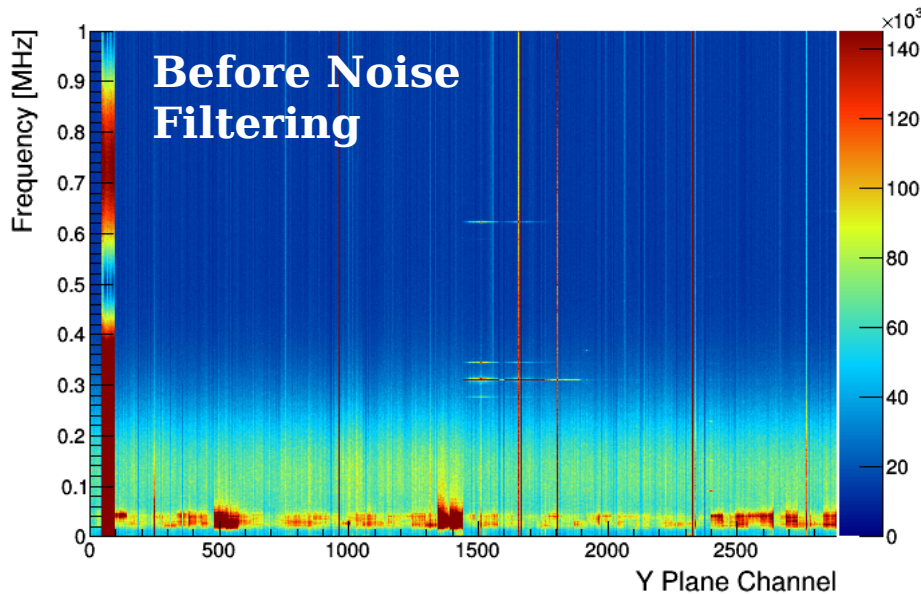
- ◆ Some  $^{39}\text{Ar}$  beta decay candidates hidden in sag of waveform - local baseline correction (sidebands of integration window) fixes energy estimate
- ◆ Really should correct **before** selecting candidates



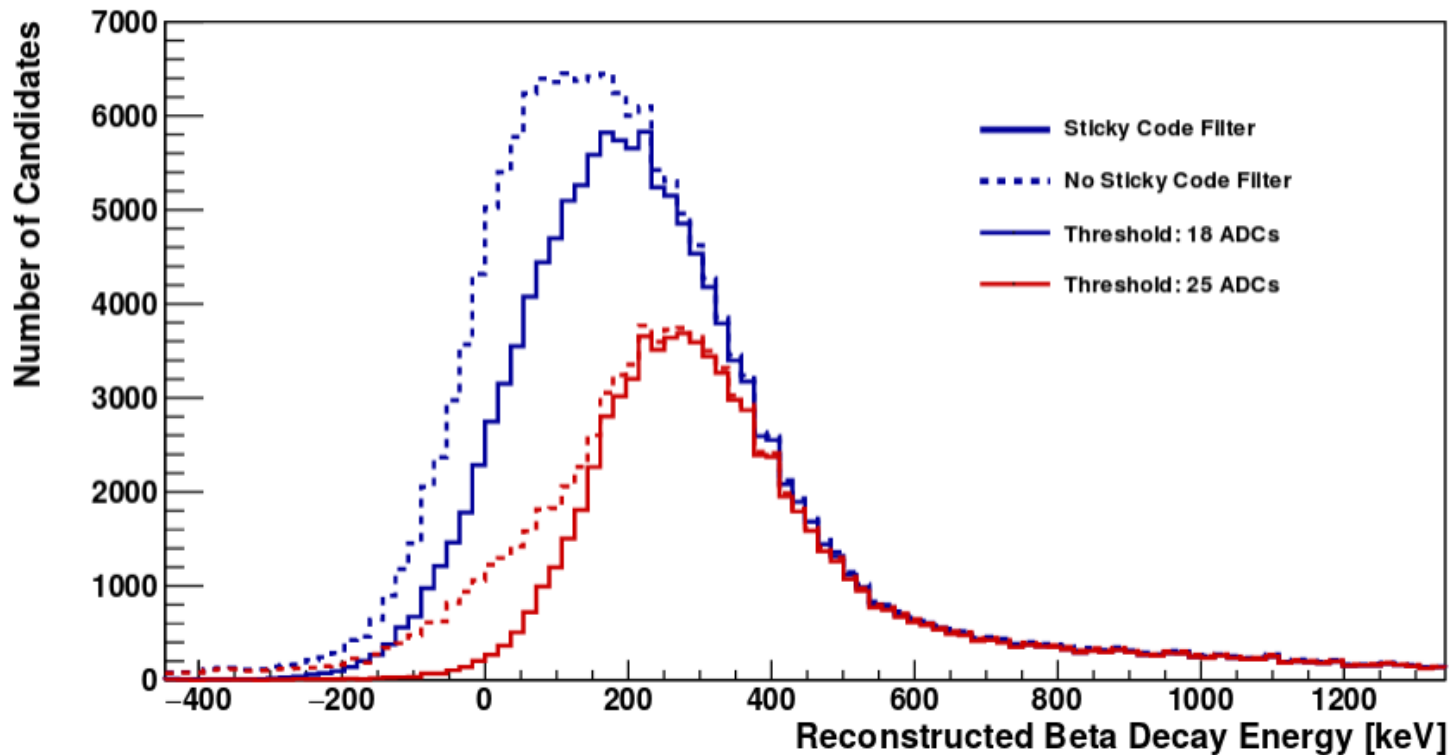
# Better Spectrum



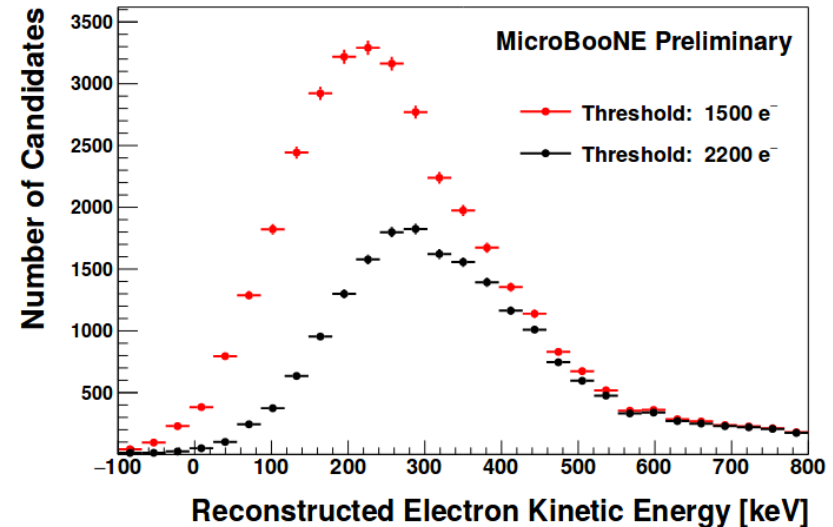
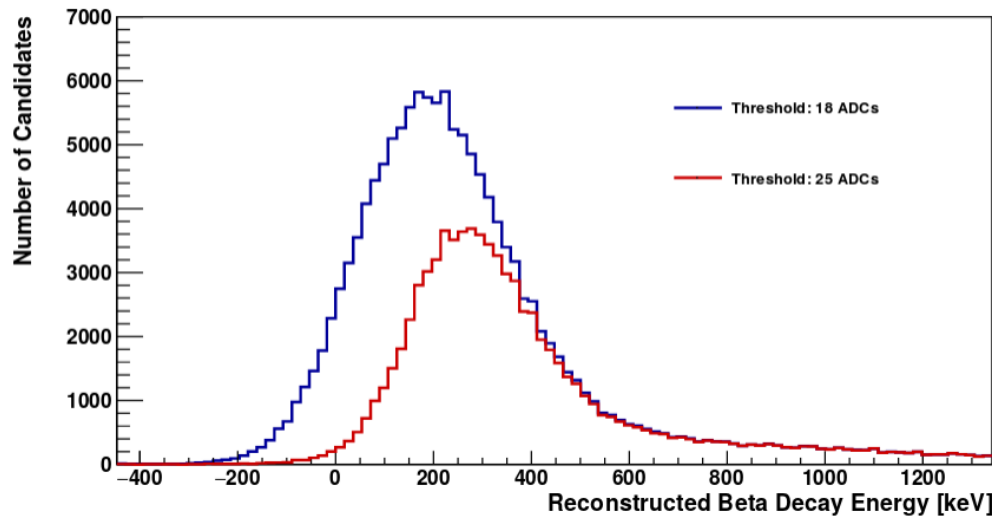
- ◆ Spectrum more narrow, but still much more broad than seen in MicroBooNE
- ◆ Can be improved further with coherent noise removal - see next slide



- ◆ Coherent noise @ 10-40 kHz
- ◆ Remove by subtracting median ADC value across every 16 channels, independently for each tick
- ◆ ENC:  $\sim 100 e^-$  effect

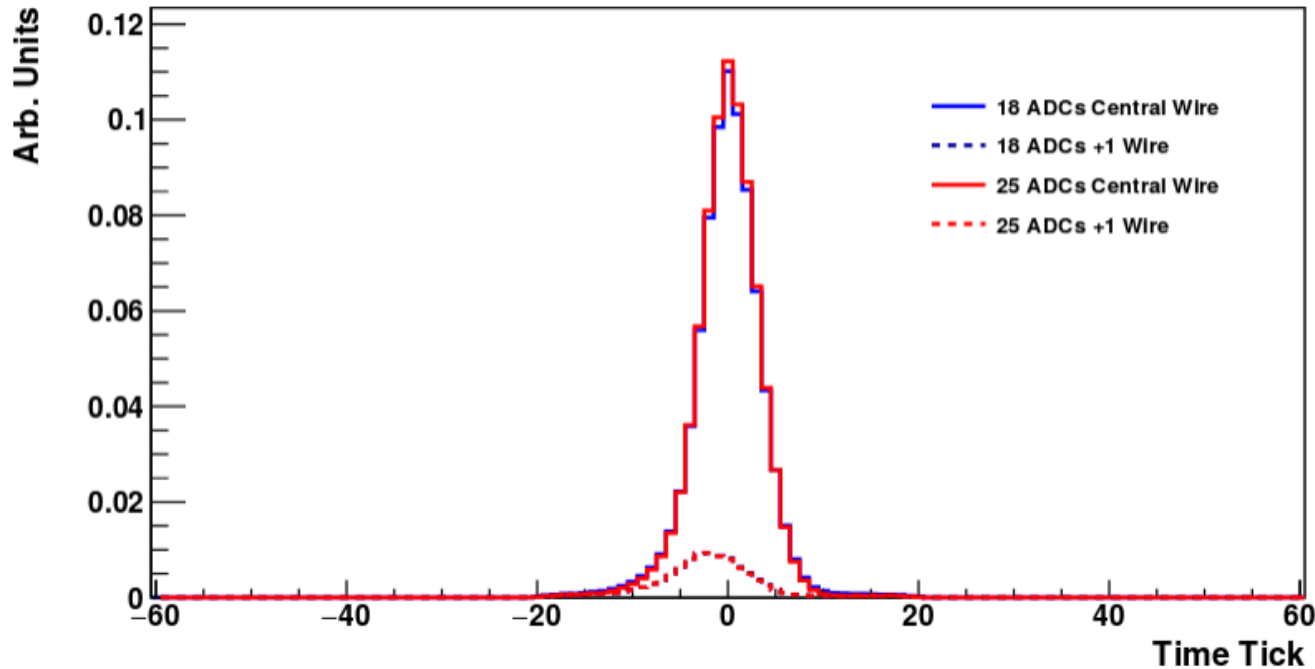


- ◆ Much better, but strange low-energy features
- ◆ Largely a result of **sticky codes** (ADC issues); can filter out with pulse shape discrimination
  - Reject candidates with abrupt jump in ADC across time



- ◆ After the three fixes, spectrum looking much closer to that observed at MicroBooNE
  - Still broader due to longer wires, thus higher noise
  - Lower energy reach due to less recombination at 500 V/cm (MicroBooNE runs at 273 V/cm)
- ◆ Next: fit to MC templates and extract **rate**

# Signal Shape

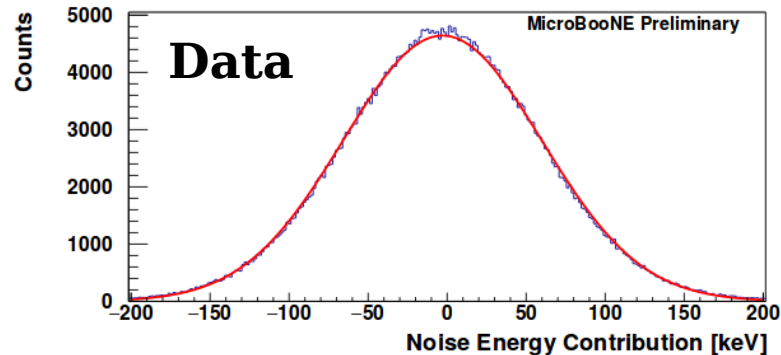
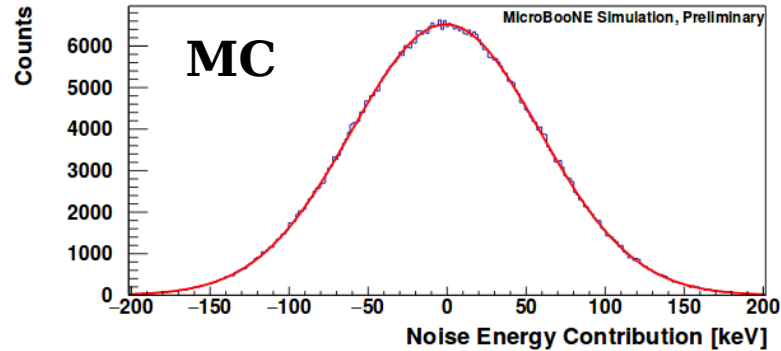


- ◆ Also looked at average signal shape
- ◆ Side wires see very little relative signal – diffusion is not too significant of an effect, on average
  - Helps that wire spacing is 5 mm (3 mm at MicroBooNE)
- ◆ Can see hints of induced charge here as well

- ◆  $^{39}\text{Ar}$  beta decay could be very useful for calibrations at DUNE FD (in lieu of cosmics)
- ◆ First studied at MicroBooNE; now at ProtoDUNE-SP
  - Reconstructed energy spectrum looks reasonable
  - Signal shape consistent with point-like ionization
- ◆ Need to compare to MC as next step
  - Use MC templates to fit to data, extract rate ( $\sim 1$  kg/Bq)
- ◆ DAQ will be a challenge with nominal approach to forming trigger primitives – rate very high!
  - Instead: consider doing prompt analysis on FPGA, and read out only shape/energy “histogram” (one per wire)
  - Use round-robin approach to decrease bandwidth

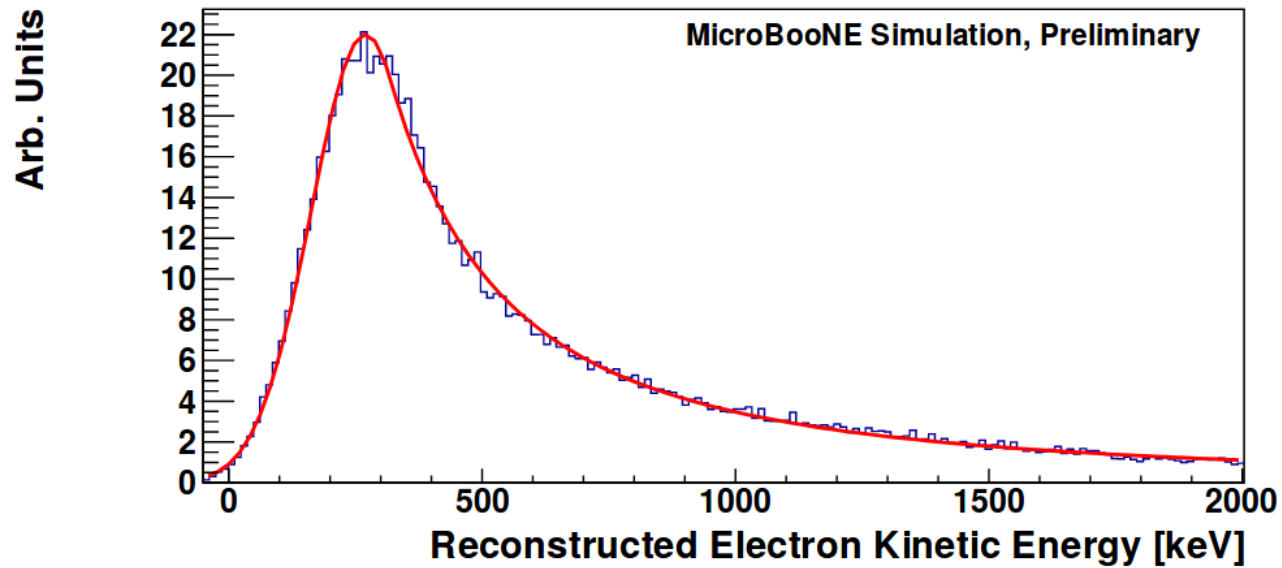
# Backup

# Noise Comparison

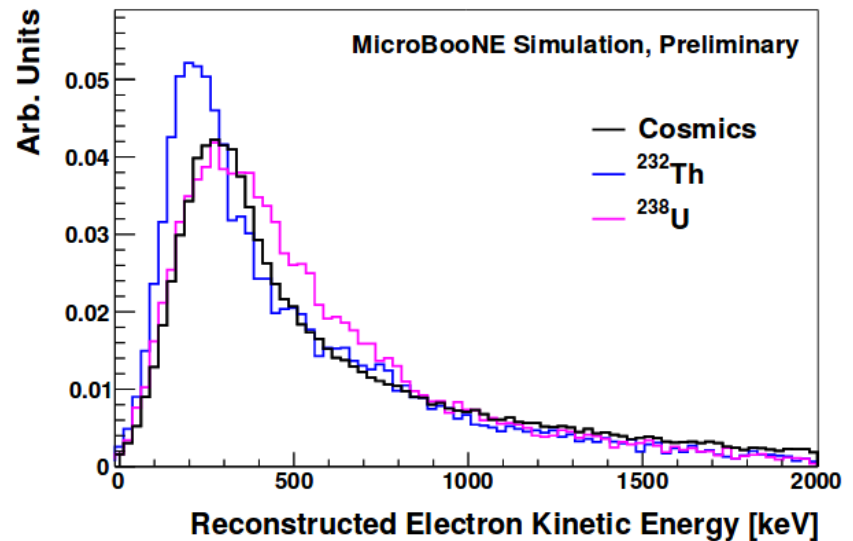
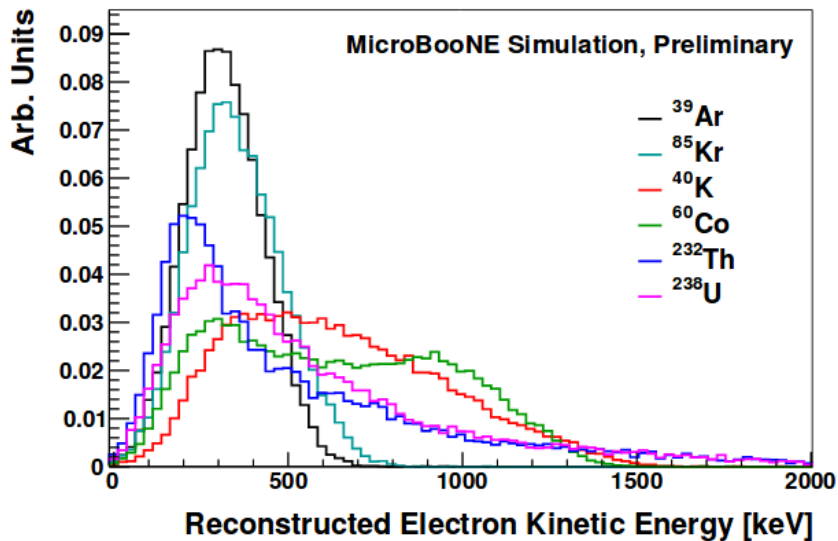


- ◆ From MicroBooNE public note
- ◆ Compare spectrum in data and MC, selecting windows randomly across readout window
  - Data and MC very similar ( $\sim 60$  keV RMS)





- ◆ From MicroBooNE public note
- ◆ Spectrum fits well to Crystal Ball function



- ◆ From MicroBooNE public note
- ◆ Similar shapes from cosmic background,  $^{232}\text{Th}$ , and  $^{238}\text{U}$  → simpler to separate underground