# **DUNE FD DAQ**

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## **DUNE FD DAQ**

- The DUNE FD DAQ system design has progressed rapidly in recent months
- This talk summarizes:
  - DAQ Baseline Design
    - DUNE FD SP DAQ baseline design has passed internal CDR review in December 2018
  - DAQ Data Selection System
  - **Design Justification** (ProtoDUNE and simulations)

### **DUNE FD DAQ Design**



- DAQ is split between 4850ft level and surface: Bulk of processing/buffering underground, minimizing data traffic to surface.
- Strong power, cooling, space constraints in CUC (600kW for all 4 modules, ~50 racks)

### **DUNE FD DAQ Design (SP+DP)**



# **DUNE FD DAQ Design (SP+DP)**

- Scope:
  - Begins at fibers from the detector (electrically decoupled from cryostats)
  - Ends at WAN network interface (fibers to FNAL via ESnet)
  - Provides common computing and network services for other systems
  - All slow control and safety functions are outside DAQ
- Functionality:
  - Provides basic timing and synchronization for sub-detectors
  - Receives, synchronizes, compresses, buffers streaming data from sub-detectors
  - Extracts trigger primitives from data, summarizing local activity in detector; makes local, module, and cross-module trigger decisions
  - Builds "event records" from selected space-time volumes, buffers, and relays those to permanent storage
  - Carries out local data reduction and filtering as required

# **DUNE FD DAQ Design (SP+DP)**

#### Design principles:

- A single, scalable system design for all detector modules
- Ability to self-trigger with zero dead-time and with high efficiency for targeted physics signals and to record triggered data covering the full detector
- Evolutionary design: begin with very conservative design for first module
- Preserve possibility of additional capacity as required

#### Key challenges:

- Supernova Burst (SNB) requirements: large buffering upstream, low fake trigger rates
- Long ("permanent") commissioning state: partition-able system
- Difficult-to-access location: reliability and remote operation
- Power, cooling and space considerations

### **DUNE FD DAQ: SP-focused in next slides**



For 1 SP Module:



#### For 1 SP Module:

1 (+1) Data Selection (MLT) server +networking +CCM Surface control room One of these in DUNE FD SNEWS, LBNF etc LBNF etc LBNF etc To / from other MLTs Filter

Common to all four modules: 1 (+1) Data Selection (ETM) server Timing System Interface + networking

+CCM

Also: Trigger Primitive Local Storage System (~1PB)

#### For 1 SP Module:

30 Event Builder server +networking +CCM

5 Data Selection (HLT) servers +networking +CCM

Storage buffer (~1PB)





# Substantiating Design: 1. ProtoDUNE-SP

#### **Demonstrated Hardware Components:**

#### • Front-end Readout:

- 2 out of 6 APAs are now read out by FELIX-based Front-end system (~DUNE FD system without Co-processor(s) and only 1 APA/FELIX board)
  - $\rightarrow$  interface to front-end electronics, scalability, data flow
  - $\rightarrow$  host server requirements and specifications
  - $\rightarrow$  platform for further system development (co-processor, trigger primitives)
- Back-end DAQ: Event Builder farm, CCM machines, disk buffers
  - $\rightarrow$  system partitioning, scalability
  - → platform for further system development (CCM, IPC, Data Flow Orchestrator)
- Data Selection/Timing: (External Triggering)
  - $\rightarrow$  Timing Distribution System

### Substantiating Design: 1. ProtoDUNE-SP

• Key challenge for DUNE FD: (what is new with respect to ProtoDUNE, MicroBooNE, SBND, ICARUS...):

#### **Continuous self-triggering of detector**

- Self-triggering demonstration is being planned for ProtoDUNE (target: 2019, and ~2021).
- Other detectors (e.g. ICARUS, MicroBooNE) have demonstrated self-triggering in coincidence with external gates (data and trigger rates are capped).
   In DUNE-FD: Such ``throttling" limits physics sensitivity to off-beam physics.

→ Need intelligent trigger (Data Selection) design and validation on reliable simulation, plus accurate noise and background predictions.

(This has been the Data Selection WG focus thus far).

### **Data Selection sub-system**

### **Baseline Data Selection Strategy and Hierarchy: TPC-based**

Lowest Level

Trigger Primitives: "hit on a wire/channel"

Trigger Candidates: "clusters of hits"

Trigger Command:

(prompts "event record")

"localized (HE) interaction" "extended (SNB) interactions"

Filter:

down-selection of event records

**Highest Level** 

### **Data Selection sub-system**



Block diagram of DUNE DAQ Data Selection sub-system, illustrating hierarchical structure of sub-system design.

### **Expected Trigger Rates:**



( $\rightarrow$  100s of losslessly compressed data from entire module to event builder; rate ~1/month)

### TPC Trigger Primitive Simulations (SP)

Significant effort dedicated on trigger primitive simulations (see backup slides):

- Charge collection efficiency and fake rates due to noise and radiologicals have been studied as a function of hit threshold with MonteCarlo
- CPU trigger primitive generation speed on CPU (4 cores) was demonstrated to keep up with expected raw data rates, offline
- Trigger primitive rates measured at ProtoDUNE-SP in situ. Effort on understanding and removing contribution from cosmics/cosmogenics and (known) noisy channels is ongoing.
- Full-stream, single-APA, online trigger primitive generation on CPU (10 cores) demonstrated at ProtoDUNE SP

### TPC Trigger Candidate Simulations (SP)



Trigger Candidates formed by clustering Trigger Primitives in Channel (collection) vs. Time

Efficiency for triggering on highest Charge APA• Win for electrons

Trigger Candidates with sufficient energy can serve as Trigger Commands for **"localized (HE) interaction"** events: beam, atmospheric neutrinos, nucleon decay, neutronantineutron oscillation, cosmics

#### Simulation demonstrates >99% trigger efficiency for >100 MeV visible energy.





### TPC Trigger Candidate Simulations (SP)

Low-energy trigger candidates serve as input to SNB trigger:

120000 Entries 18.06 Mean \$td Dev 6.821 0.6 0.5 ┶╴ ┶╴ 0.4 0.3 0.6 0.2 0.5 Cluster configuration: 5 Hit ADC threshold: 8 0.1 0.4 0L 5 10 15 20 25 30 35 40 45 50 0.3 Ev [MeV] 0.2 Cluster configuration: 5 A. Booth, 0.1 Hit ADC threshold: 18 P. Lasorak 0L @ Sussex 50 Ev [MeV] 5 10 15 20 25 30 35 45 40

Simulation demonstrates ~20-30% trigger candidate efficiency for individual SN neutrino interactions.

### TPC SNB Trigger Simulation (SP)

**SNB trigger: Multiplicity-based:** low-energy trigger candidates over (up to) 10 seconds. (An energy-weighted multiplicity count scheme could be applied to increase efficiency/minimize background.)



SNB trigger studies demonstrate high (>90%) galactic coverage while keeping fake SNB trigger rate to 1/month.

Trigger Rates and ProtoDUNE-SP Event Builder experience
→ Extrapolation to DUNE FD SP Back-End DAQ

- Simulation studies support physics and rates requirements:
  - >99% trigger efficiency for "localized (HE) interactions," and
  - >90% galactic coverage for SNB's can be maintained while keeping:
  - "localized (HE) interaction" trigger rate to ~ 0.1 Hz (<1 Hz)
  - "extended (SN) interactions" (SNB) trigger rate to ~ 1/month
- Additionally, ProtoDUNE-SP Back-End DAQ has demonstrated our ability to keep up with 1/25<sup>th</sup> the size of a single DUNE FD SP module, for trigger rates up to 40 Hz and 3 ms readout window (x5 design goal)
- Allows confident scaling of ProtoDUNE-SP Back-End DAQ system to DUNE FD SP scale.

### TPC HLT (Filter) SP Simulations: E.g. CNN-based APA-frame image classification/filtering in GPUs

- TP's generated from a large number of wires per APA (over a few APAs)
- Multiple TC reported to MLT (one or more per APA over a few APAs)
- MLT is aware of detector partition and operating state of each APA and accelerator timing → decides it's a valid HE candidate event
- MLT sends trigger command (interpreted as 5.4ms readout over entire module) to Back-end System
- Data retrieved → Event Builder
- HLT can optionally remove APAs which are "empty"



#### **Post-Event-Builder:**

E.g. CNN trained to classify "empty" APA-frames can provide further radiological+noise reduction (e.g. empty frames can be dropped post event builder).

### TPC HLT (Filter) SP Simulations: E.g. CNN-based APA-frame image classification/filtering in GPUs

- TP's generated from a small number of wires per APA (over multiple APAs)
- Multiple TC reported to MLT (~one per APA over multiple APAs and over an extended period of time (longer than 1 drift))
- MLT is aware of detector partition and operating state of each APA and accelerator timing → decides there are valid LE candidate events → decides this is a valid SNB event
- MLT sends trigger command (interpreted as 100s readout over entire module) to Back-end System
- Data retrieved → Event Builder
- HLT can optionally remove APAs which are "empty"

#### **Post-Event-Builder:**

E.g. CNN trained to classify "empty" APA-frames can provide further radiological+noise reduction (e.g. empty frames can be dropped post event builder).



### TPC HLT (Filter) SP Simulations: E.g. CNN-based APA-frame image classification/filtering in GPUs

RAD	RAD	SN	n-nbar	atmo. nu	p-decay	cosmic	1	100 F		Rad. scores (true rad. samples)				
score cut	frame efficiency	frame efficiency	frame efficiency	frame efficiency	frame efficiency	frame efficiency	pa	80			RA RA	D_scores D_scores	_misident	lified
<0.1	0.73% (99.44% rejection)	89.18%	99.98%	92.24%	99.29%	92.57%	rmalize	60		events mis-identified as SN by Pierre's clustering algo				
<0.01	0.14% (99.82% rejection)	83.27%	99.98%	91.01%	99.18%	92.46%	Area no	40			Y. Jv	va, G.	Ge	
<0.001	0.033% (99.969% rejection)	77.11%	99.98%	89.76%	99.04%	92.24%		20		@ Columbia				
<0.0001	0.011% (99.989% rejection)	69.74%	99.97%	88.39%	98.74%	91.71%		Ū	0.0	0.2	0.4	0.6 d scores	0.8 for all r	1.( ad(+noise)

This scheme has been shown to correctly tag mis-identified events for low-energy clustering trigger approach.

CNN vgg16b network performance: **20-30ms classification time per APA-frame** on single GPU card.

# **Closing remarks**

#### **DUNE FD DAQ Conceptual Design (Baseline) is complete.**

Substantial efforts in recent months have gone into substantiating design and costing, through **simulations and ProtoDUNE-SP** experience.

What has been presented is being incorporated in the **TDR**.

Coming up:

- ProtoDUNE-DP
- **Parallel (simulation) efforts ongoing on PDS-based data selection**, for SNB trigger (see talks at January collaboration meeting)
  - SP: P. Lasorak @ Sussex, D. Pershey @ Duke
  - DP: J. Soto, A. Gallego Ros @ CIEMAT

...and groups already thinking about future evolutions of the DAQ design!

### **Backup Slides**



- Constraints:
  - Large data volume from Single Phase TPC
  - Relatively poor signal/noise ratio
  - Tight power limits for DAQ
  - Data from all readout (DP, SP / TPC, PDS) received into Felix PCIe board
- Baseline solution: Perform some functions on FPGA co-processor to SP TPC
  - reduce bandwidth to server PCs
  - Reduces total system power (still under investigation)
  - Approach successful at LHC experiments
    - Re-using firmware development infrastructure from CMS
- Functions on FPGA:
  - Hit finding
  - Lossless data compression
  - Data buffering for SNB triggering (O(10-seconds))
  - Storage of losslessly compressed data for SNB triggers (O(100-seconds))

- DUNE variant of Felix
  - Connection to coprocessor
    - Uses VITA standard
       FMC+ connector.
  - Uses same firmware / software as Atlas Felix
    - .... which is large
       majority of
       development effort
- For more details see <u>talk at DUNE</u> <u>Collaboration</u> <u>meeting, September</u> <u>2018</u>



- Initial studies have shown that hit finding in FPGA straightforward
  - current algorithms consider each wire separately
  - Low FPGA resource usage
- "Low cost" data compression in FPGA straightforward
  - Encode differences between samples with fixed table.
  - Achieves almost same performance as "best possible" compression.
  - Very low FPGA resource usage
  - "Best possible" compression, e.g. Huffman compression with dynamic tables, tried at ProtoDUNE but uses significant FPGA resources.
- Working towards "full chain" tests in Lab then ProtoDUNE

### **Co-processor Firmware**

- Time-multiplexing framework
  - Logic runs at high multiple of TDC sample rate. Reduces resource usage. Used extensively at LHC
  - Data divided up into packets and sent along communication channels between processing blocks
  - Each processing block cycles through data from 64 wires
  - For more details see talk at DUNE Collaboration meeting, Jan 19



### Summary Table—Uncompressed!

Source	Annual Data Volume/ 10 kt	Assumptions
Beam interactions	27 ТВ	800 vs+800 dirt $\mu s$ ; everything read out for 5.4 ms (no ZS); 10 MeV threshold in coincidence with beam time; include cosmics
Cosmics (+atmospherics) Scheme 1	10 PB	All wires in 5.4 ms window around HE event
Radiologicals	< 1 PB	Weighting scheme for SN bursts and other LE events; dump all 10 s for each burst trigger; tune fake rate to be < 100/year.
Front-End cals	200 TB	Worst case of measuring every single ADC bin with 100 measurements/point; four times/year
Radioactive source cals	100 TB	Source rate < 10 Hz; only one APA readout; PDS is negligible; full readout window per tag; no ZS
Laser cals	200 TB	1x10 <sup>6</sup> total laser pulses; tight ZS for both induction and collection; ½ of all wires in TPC illuminated
Random Triggers	60 TB	Same as cosmics scheme; rate is 45/day
Trigger Primitives	< 2 PB	Only collection wires; 12 b/primitive; 4 primitive types; <sup>39</sup> Ar dominates;

### Lossless compression could yield x4 reduction (cf. Viren) 5

### **Uncompressed Event Rates**

Localized HE triggers (1 SP module): 5.4ms x 150 x 2560 x 2 MHz x 12 bit = 6.22 GB

Extended SNB triggers (1 SP module): 100s x 150 x 2560 x 2 MHz x 12 bit = 115 TB

# **Trigger Primitive Definitions**

• Currently in TDR: Conservative TP Model:

• channel address	32bit
• time of hit	64bit
• time over threshold	16bit
• ADC sum	32bit
• error flag	16bit

corresponding to a total of 20 Bytes.

 Expected to be dominated by Ar39 radiological backgrounds (~100Hz per channel for 0 threshold)

### **TPC Trigger Primitive Simulations**



Figure 5: Efficiency and hit rate, default noise. The rate of hits due to noise (ie, not matched with any true energy deposition) is too low to show up on this plot.

Ongoing effort on trigger primitive generation: Efficiency and TP rates due to noise and radiologicals.

### **TPC Trigger Primitive Simulations**



Figure 6: Efficiency and hit rate, with default noise increased by 50%.

#### Effect of noise increase.

### **TPC Trigger Primitive Simulations**



Ongoing effort on filtering (simulation based) and trigger primitive generation speed. Figure 7: Time for Trigger Primitive generation in milliseconds of processing over APA milliseconds, as a function of the number of threads. The test was done on an existing Xeon Gold 6140 system. The boxes indicate the median and interquartile range of times for various iterations of the test, and the whiskers indicate the minimum and maximum after 1000 repeats of thet test. By 4 threads (cores), the algorithm can keep up with detector data rate.

### **ProtoDUNE-SP Trigger Primitives**

"Pure" 39Ar rate should be ~33kHz/ APA (half of DUNE, since only one side of APA is active).

Currently working on understanding additionally expected contribution from cosmics and (known) noisy channels in ProtoDUNE-SP.



Also, full stream, single APA trigger primitive generation on CPU (10 cores) demonstrated at ProtoDUNE-SP!

### **ProtoDUNE-SP Noise Studies**





- Coherent noise @ 10-40 kHz
- Remove by subtracting median ADC value across every 16 channels, independently for each tick
- ♦ ENC: ~100 e<sup>-</sup> effect

### **ProtoDUNE-SP Ar39 Rates**



- After the three fixes, spectrum looking much closer to that observed at MicroBooNE
  - Still broader due to longer wires, thus higher noise
  - Lower energy reach due to less recombination at 500 V/cm (MicroBooNE runs at 273 V/cm)
- Next: fit to MC templates and extract rate

## **DUNE FD SP DAQ Schedule & Milestones**



Current status:

- Done: Establish baseline design
- In Progress: Demonstrate baseline design in protodune
- Planned: Complete engineering design

### **Consortium Membership**

Country Institute		Primary interests	Links	
CERN	CERN	FELIX software		
France	Lyon	DP interface	DPE	
Japan	Iwate	DP interface	DPE	
Japan	КЕК	DP interface	DPE	
Japan	NIT Kure	DP interface	DPE	
Netherlands	NIKHEF	FELIX firmware		
UK	Birmingham	Co-processor		
UK	Bristol	Timing system; firmware		
UK	Edinburgh	Computing interface	Computing	
UK	Imperial	CCM; firmware		
UK	Liverpool	PD-SP coordination	APA	
UK	Oxford	Co-processor; data selection		
		CCM; firmware;		
UK	RAL	infrastructure / installation	тс	
UK	Sussex	Data selection		
UK	UCL	Firmware		
UK	Warwick	PD-SP PDS	PDS	

Country	Institute	Primary interests	Links
USA	BNL	FELIX hardware	CE
USA	Columbia	Data selection	
USA	Duke	Data selection	
USA	FNAL	Back-end	Many
USA	lowa	Data selection	
USA	Minnesota	Infrastructure / instalation	
USA	Notre Dame	Data selection	Calibration
USA	Penn	Data selection; timing system	
USA	PNNL	TBD	
USA	SDSMT	Data selection	
USA	SLAC	CE interface	CE
USA	UCD	DQM	
USA	Colorado	Data Selection	
USA	UCI	TBD	
Portugal	1 institution	Calibration interface	Calibration
Italy	1 institution	Data selection	
Canada	3 institutions	FELIX	
Czech republic	1 institution	CCM	
France	1 institution	Firmware	DPE

# **Consortium Organization**

- Consortium Leader:
- Technical Leader:

Dave Newbold (STFC) Georgia Karagiorgi (Columbia)

• TDR Editors: Brett Viren (BNL) & Georgia Karagiorgi (Columbia)

#### Working groups:

- WG1. Architecture
  - Lead: Giles Barr (Oxford) & Giovanna Lehman-Miotto (CERN)
- WG2. Hardware
  - Lead: David Cussans (Bristol) & Matthew Graham (SLAC)
- WG3. Data Selection
  - Lead: Josh Klein (U. Penn.)
- WG4. Back-end
  - Lead: Kurt Biery (FNAL)
- WG5. Integration/Infrastructure
  - Lead: Alec Habig (UMN Duluth)