First Results from ICEBERG PD

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first light from cosmics!



ARAPUCA and X ARAPUCA data were not taken simultaneously, don't make direct comparisons yet

0.8 µs pipeline delay, 3 µs readout window



some more examples



Collected using the ARAPUCA.

noise features



Two noise features that stand out immediately.

short term mu2e readout plans

- Identify noise sources: Trigger off of a single cosmic paddle, O(5k) events, with maximum readout window and Vb a few volts below breakdown (~44 V). Look at FFT to identify possible noise sources and develop methods for mitigation.
- Look for single photoelectrons: Do the same as above with Vb = 48 V and look for SPEs near pulse time that we've identified.
- **Identify through-going cosmic population:** Switch trigger to cosmic paddle coincidence and repeat again.

current data available

Coincident cosmic trigger data (3/11/19)

~300 triggers, Vb = 47 V (ARAPUCA and X-ARAPUCA read out consecutively)

Single cosmic paddle trigger data (3/14/19)

- \sim 5k triggers with Vb = 40 V (ARAPUCA only)
- \sim 1k triggers with Vb = 47 V (ARAPUCA only)
- ~ 2.5 k triggers with Vb = 48 V (ARAPUCA only)

On 3/14 I was simultaneously testing how quickly I can read things out. In the next round of data-taking I will be able to accumulate data a little faster.

data-taking plans for tomorrow

- If the schedule current schedule stays on track, the photon detectors will be on all day tomorrow and I plan to take lots of data.
- Accumulated more stats w/ the ARAPUCA at Vb = 48 V.
- Collect X-ARAPUCA data w/ Vb = 40 V and Vb = 48 V to compare.
- Do we want some data taken in 1 V steps b/w 44 V and 48 V, time permitting?
- Start identifying possible noise source and work on offline noise filtering options.
- When ICEBERG is emptied I plan to figure out why the second channel isn't being read out so we're ready for simultaneous readout during the next fill.

