

# Mu2e Experimental Issues affected by Slow Extraction Variations

R. BernsteinSlow Extraction Workshop22 July 2019

**Basic Problem** More Less Useful Data! Data Calibration Drifts Vetoes Finding a balance Detector Backgrounds **Trips DAQ** Overload

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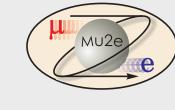
R. Bernstein

Slow Extraction 2019





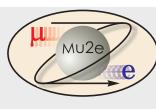
- Relative Importance varies across time scales and any given experiment, but for example:
  - tracker momentum reconstruction degrades from confusion: pile-up leads to tossing tracks
  - extra energy in calorimeter making cluster finding difficult and increases noise terms in resolution function; gain can fluctuate with intensity variations, which is very hard to undo in analysis
  - matching tracker to calorimeter (or any subsystem to any other subsystem)
  - particle ID
  - veto systems lead to unacceptable losses
  - combinatorial background
  - DAQ, front-end electronics, and trigger: bandwidths (or single-event upsets)



# Implying More Measurements

- Spill uniformity has to be monitored carefully
  - and there can be multiple time scales even within one detector system
    - for example, calorimeter has ps timing of start of shower; ns for matching to tracker; µs for trigger/DAQ

## What is µe Conversion?

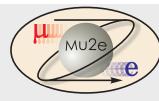


muon converts to electron in the field of a nucleus

$$\mu^- N \to e^- N \qquad \text{does not change}$$
 
$$R_{\mu e} = \frac{\Gamma(\mu^- + N(A,Z) \to e^- + N(A,Z))}{\Gamma(\mu^- + N(A,Z) \to \text{all muon captures})}$$

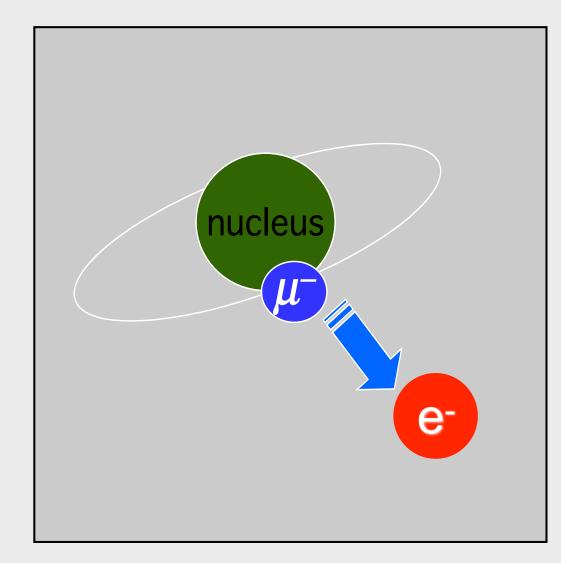
- Standard Model Background of 10-54
- Charged Lepton Flavor Violation (CLFV)
  - can measure a signal with SES of ~ 3 x 10<sup>-17</sup>
- Related Processes: $\mu$  or  $\tau \to e\gamma$ ,  $\tau \to 3l$ ,  $K_L \to \mu e$ , and more

# **Experimental Signal**

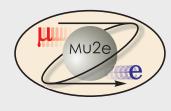


$$\mu^- N \rightarrow e^- N$$

- A Single Monoenergetic Electron
- If N = AI, E<sub>e</sub> = 105. MeV
  - electron energy depends on Z
- Nucleus coherently recoils off outgoing electron, no breakup



#### Overview Of Processes

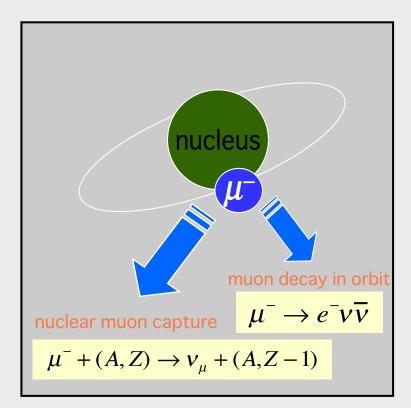


 $\mu$ - stops in thin Al foil

the Bohr radius is  $\sim 20 \text{ fm}$ , so the  $\mu$ - sees the nucleus



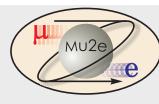
muon capture, muon "falls into" nucleus: normalization



60% capture 40% decay

Decay in Orbit: background

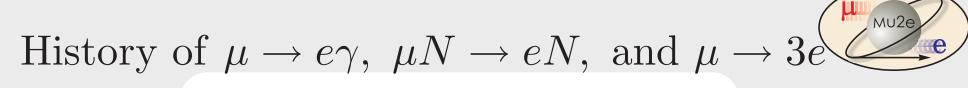
#### Mu2e In a Sound Bite





-Roni Harnik

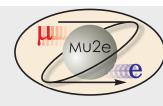
- Why are there flavors?
- Why is there more than one generation, and always three!





RHB and P.S. Cooper, Phys Rept C (1307.5787)

#### How Rare is That?



Pretty Rare: let us know if this happens to you!

Probability of	
rolling a 7 with two dice	1.67E-01
rolling a 12 with two dice	2.78E-02
getting 10 heads in a row flipping a coin	9.77E-04
drawing a royal flush (no wild cards)	1.54E-06
getting struck by lightning in one year in the US	2.00E-06
winning Pick-5	5.41E-08
winning MEGA-millions lottery (5 numbers+megaball)	3.86E-09
your house getting hit by a meteorite this year	2.28E-10
drawing two royal flushes in a row (fresh decks)	2.37E-12
your house getting hit by a meteorite today	6.24E-13
getting 53 heads in a row flipping a coin	1.11E-16
your house getting hit by a meteorite AND you being	
struck by lightning both within the next six months	1.14E-16
your house getting hit by a meteorite AND you being	
struck by lightning both within the next three months	2.85E-17

### Measuring 10<sup>-17</sup>

- The captured muon is in a 1s state and the wave function overlaps the nucleus (picture ~ to scale)
- We can turn this into an effective luminosity
- Luminosity = density x velocity

$$|\psi(0)|^2 \times \alpha Z = \frac{m_\mu^3 Z^4 \alpha^4}{\pi} = 8 \times 10^{43} \text{ cm}^{-2} \text{ sec}^{-1}$$

AI

- Times  $10^{10}$  muons/sec X 2  $\mu$ sec lifetime
- Effective Luminosity of 10<sup>48</sup> cm<sup>-2</sup>sec<sup>-1</sup>



# Mu2e Specific: How Does this work in one experiment

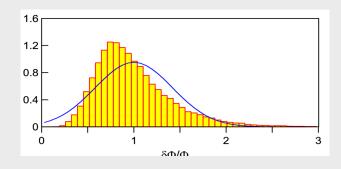
- We see all of these effects but I will focus on just a couple
- Mu2e is "primarily" concerned with backgrounds for this discussion:
  - we need to have our vetoes as close to 100% live as possible
  - we don't want "fake hits" or accidental activity causing track misreconstuction and large tails on resolution
  - and of course we want as high an efficiency as possible.
     Too-intense spills lead to loss of data.

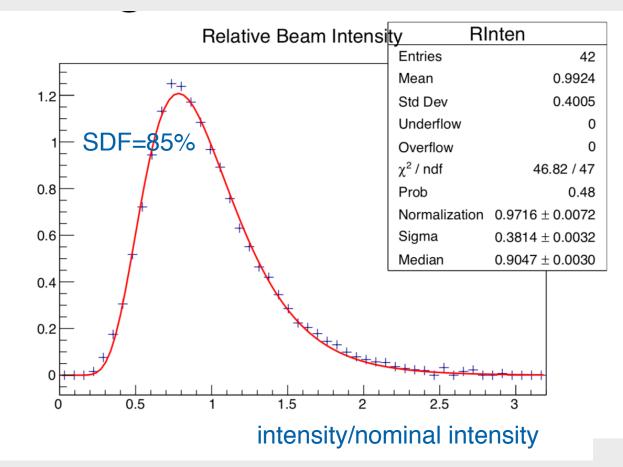
### Modeling Intensity Variations

We have the fitted distribution in our simulations.
 We did most of the simulations with a +- 50% variation. We recently went to the model below but can change it and test the result (in progress)

$$SDF = \frac{1}{1 + \left(\frac{\sigma_I}{I}\right)^2}$$

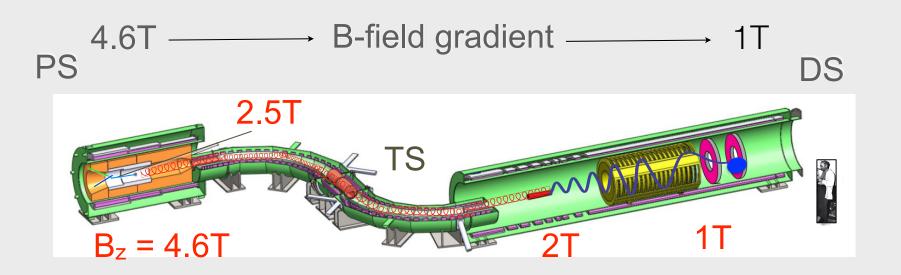
this is from Protvino
<a href="https://indico.gsi.de/event/4496/">https://indico.gsi.de/event/4496/</a>





we have no special reason to think we will do this well

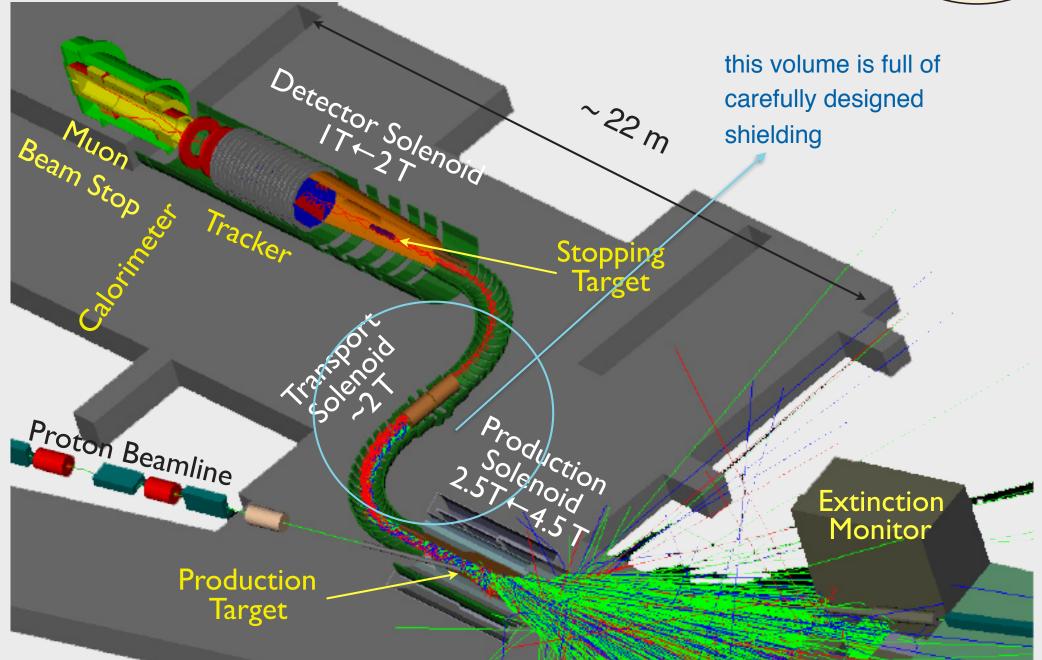
# Mu2e Muon Beam: Three Solenoids and Gradient



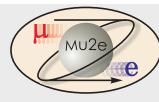
- Target protons at 8 GeV inside superconducting solenoid
- Capture muons and guide through S-shaped region to Al stopping target
- Gradient fields used to collect and transport muons

#### Beam's Eye View

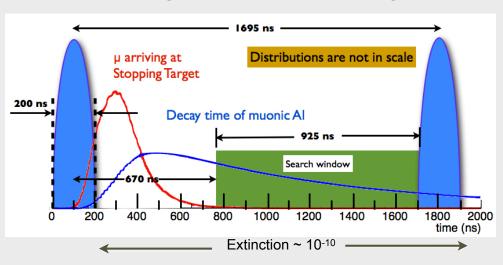




#### Pulsed Beam Structure

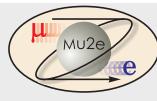


- Tied to prompt rate and machine: FNAL "perfect"
- Want pulse duration <<  $au_{\mu}^{\rm Al}$ , pulse separation  $pprox au_{\mu}^{\rm Al}$ 
  - FNAL Debuncher has circumference 1.7 $\mu$ sec , ~x2  $au_{\mu}^{
    m Al}$
- Extinction between pulses < 10<sup>-10</sup> needed
  - = # protons out of pulse/# protons in pulse

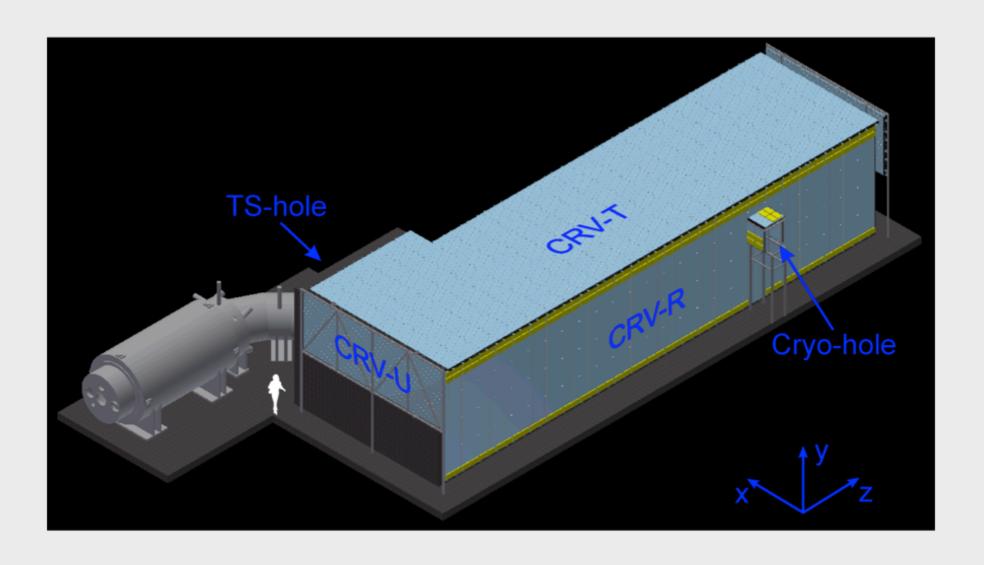


 10-10 based on simulation of prompt backgrounds and beamline

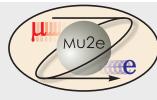
# Cosmic Ray Veto



**CRV** 

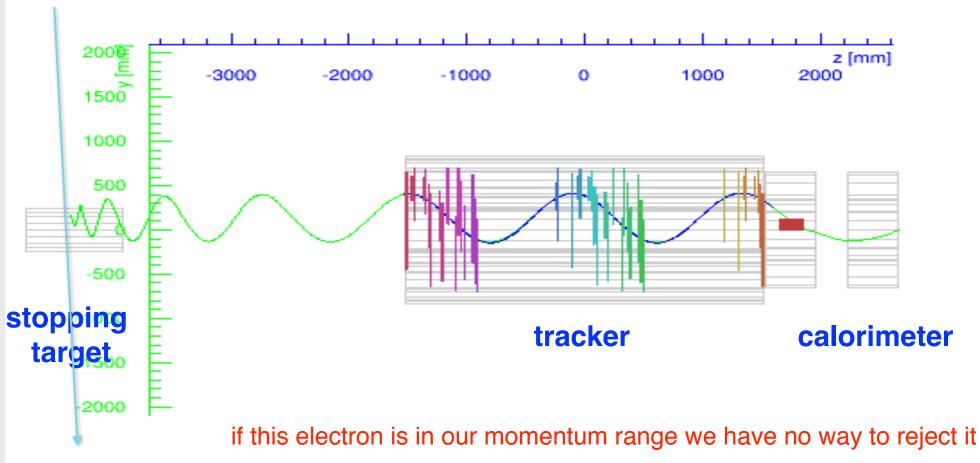


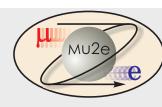
## Sources of CR Background



 Other sources of CR backgrounds come from CR entering upstream in Transport Solenoid, creating electron that is then transported to stopping target



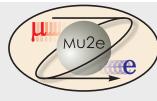




# Beam-induced sources of Dead Time

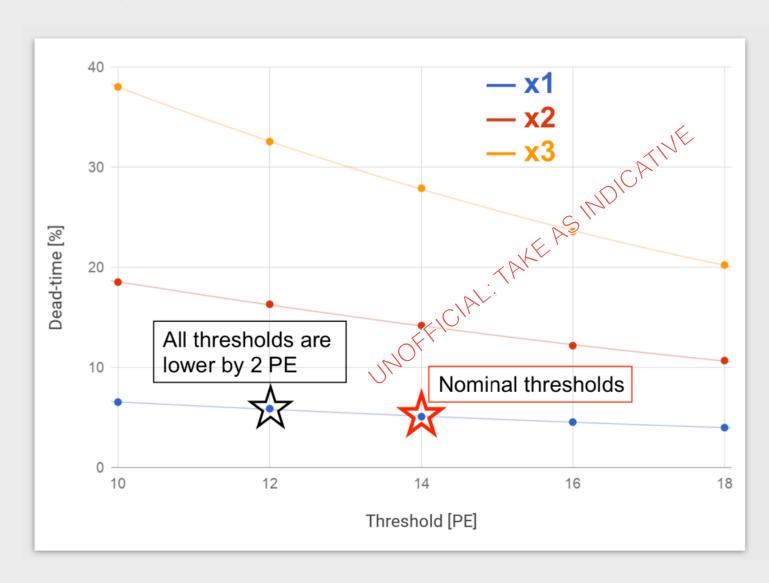
- Neutrons from Production Solenoid: neutrons get thermalized, captured, and produce delayed photons
- Fast neutrons from muon captures in beam line and the stopping target
  - fast neutron recoil off a proton deposits energy

#### **CRV Thresholds**

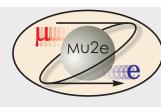


#### Y. Oksuzian

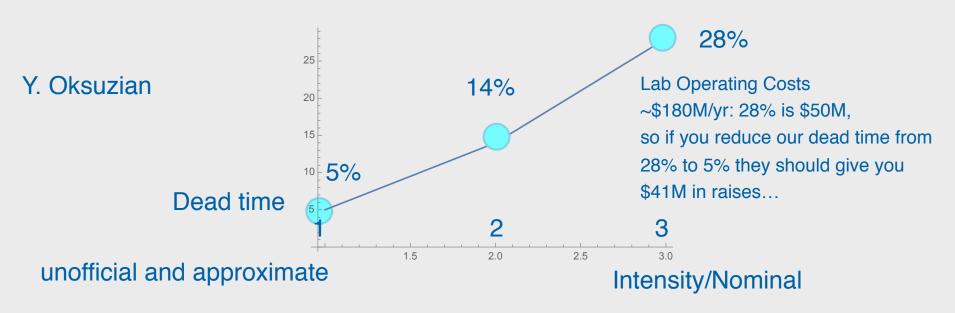
#### unofficial and approximate!!

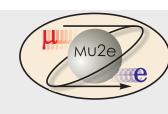


### Cosmic Ray Veto



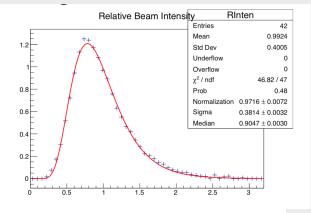
- 3/4 layer veto needs to be 99.99% efficient
- All of this not far from 8 kW, 8 GeV beam, leading to dead time from neutrons. Have to cut if CR veto is dead or we will be open to background
- Dead time:



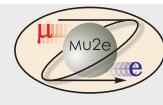


#### Work Out Some Numbers

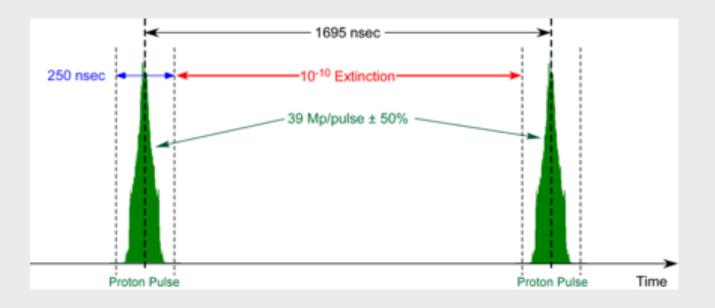
- This plot represents the 1695 ns period of 3.9e19
  - protons/micropulse
- Fit to a lognormal distribution
- Integrate against
  - this intensity function: dead time = 3.8%
  - flat intensity function from 0 to 3: dead time = 11%
- The difference is 3 months of shifts and beam to reach our goals



#### **Pulse Duration**

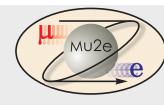


 Mu2e is not as concerned here, as long as the distribution doesn't lose beam outside the indicated limits

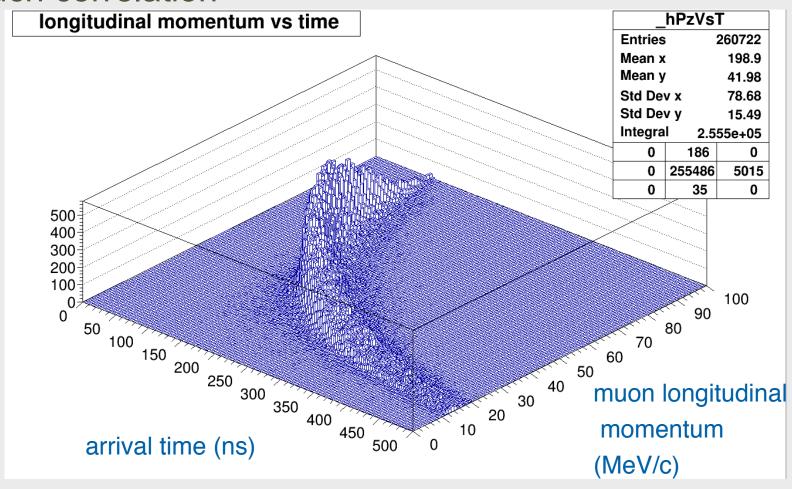


 This is because our magnet system uses solenoids. The particles travel in helices and the pitch angle (or p<sub>T</sub>) distribution of the muons smears out the arrival time, so there is no close correlation between arrival time and momentum

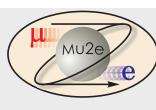
#### **Pulse Duration**



 Here is the correlation for a delta-fcn beam; smear this with 125 ns in time and you see there won't be much correlation

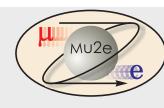


### Other Detector Systems



- Calorimeter, Tracker, extinction monitor, FEE, DAQ, not terribly sensitive
- Just don't give us superbuckets (old phrase meaning x10 or more nominal)
  - these can be real problems in rare decay experiments; we're fortunate that's not a source of background in Mu2e
  - we would have to cut these

### Summary



- Spill non-uniformity affects every experiment differently but in general:
  - background, dead time go up
  - resolution and stability get worse
- Mu2e's problem is primarily dead time