

#### "First Measurement of the Total Neutron Cross Section on Argon Between 100 and 800 MeV"



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# Physics Challenges for DUNE

- DUNE will measure the oscillation probability of  $\nu_{\mu} \rightarrow \nu_{e}$  and  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$  as a function of neutrino energy
- Oscillation phenomena (like  $\delta_{CP}$ ,  $\theta$ 's,  $\Delta m^2$ 's, etc) depend on the neutrino energy
- A successful DUNE physics program will rely on a detailed understanding of the correlation between the true neutrino energy (from the beam) and the reconstructed neutrino energy (in the detector) – this is especially true for any differences between v and v



# Physics Challenges for DUNE

- First oscillation maximum is around 2.4 GeV (near the "resonance production region") - neutrino cross-sections are highly uncertain, especially in this region
- Many of the neutrino interactions in DUNE will involve baryon resonance production
- Final state interactions of the outgoing hadronic system are also uncertain – neutrons are often a part of this system emerging from the nucleus
- Typical neutron kinetic energies are on the order of hundreds of MeV – these need to be accounted for when reconstructing the incoming neutrino energy



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## LBNF Beam Example

- Neutrons carry away considerable amount of energy that escapes detection
- Different amounts of energy carried away between neutrinos and anti-neutrinos
- No neutron-argon cross-section data above 50 MeV (R.R. Winters et al., PRC 43, 492, 1991) need neutron data at high energies!







Plots from J. Chavez

GENIE used for a simple DUNE LAr volume

# Problems Entering into DUNE Physics

- Authors simulated the effect of missing energy in a calorimetric neutrino energy reconstruction for a "DUNE" detector
- Point represents true oscillation parameters with a green shaded region showing the contour around the correct result
- Fitted allowed region shifts relative to true value as we underestimate the missing energy – this is how wrong we could be (unknowingly)



#### Mini-CAPTAIN

- 400 kg instrumented mass inside a hexagonal TPC with 32 cm drift, 50 cm apothem
- 1000 wire channels, 3mm wire pitch
- 24 x 6 cm<sup>2</sup> PMT light detection system (Hamamatsu R8520-506 MOD)
- Same cold electronics as MicroBooNE









# Neutron Beam at LANL (LANSCE)

- Los Alamos Neutron Science Center WNR facility provides a high flux neutron beam with a broad energy spectrum up to 800 MeV
- Mini-CAPTAIN utilized this facility to collect high-energy neutron events in August 2017

Building 1302 Target 4

Flight Path 5: Nuclear Physics and Neutron Radiography

HIPPO: High-Pressure-Preferred Orientation Diffractometer

SMARTS: Spectrometer for Materials Research at Temperature and Stress Target

4FP60R

SPIDER

CHI-NU

ICE House



#### Mini-CAPTAIN Goals

- Produce neutron-argon cross-section measurements (total, exclusive channels) as a function of neutron kinetic energy
- Develop strategies for incorporating neutron ID and reconstruction in DUNE

4FP60R

SPIDER

CHI-NU

ICE House

Target 4

Flight Path 5: Nuclear Physics and Neutron Radiography

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Building 1302



DANCE: Detector for Advanced

Neutron Capture Experiments

#### Neutron Beam Structure

- Nominal time structure of the beam
  - sub-nanosecond pulses 1.8 microseconds apart within a 625 μs long macro pulse
  - Repetition rate: 100 Hz



- Facility is designed to deliver high flux we want <1 neutron per TPC drift length (200  $\mu$ s)
- Broad energy spectrum we would like our measurements to be as a function of neutron kinetic energy



- To achieve low neutron flux:
  - Constrain the shutters
  - Special low-intensity running nominal bunch spacing changed from 1.8  $\mu$ s to 200  $\mu$ s yielding 3 bunches per macropulse
- We ran in two modes:
  - 1. highly constrained shutters doesn't impact other users
  - 2. moderately constrained shutters with special bunch spacing annoys every other user in the facility



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## Neutron flux monitors

#### Fission chamber in 4FP15R

- For some measurements, an external understanding of the neutron spectrum may be important
- Fission chamber useful for high neutron fluxes (~10<sup>-5</sup> interaction rate) – standard facility equipment
- Scintillator detector (stilbene) useful for low neutron fluxes (~10<sup>-2</sup> interaction rate) – deployed by CAPTAIN
- Cross-calibrate scintillator detector with fission chamber at moderately high flux



### **CAPTAIN** Collaboration

- Alabama: Ion Stancu
- LBL: Craig Tull
- Boston University: Christopher Grant
- BNL: Hucheng Chen, Veljko Radeka, Craig Thorn
- UC Davis: Daine Danielson, Steven Gardiner, Emilja Pantic, Robert Svoboda
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- Stony Brook: Neha Dokania, Clark McGrew, Sergey Martynenko, Chiaki Yanagisawa

Spokesperson: Christopher Mauger; Deputy Spokesperson: Clark McGrew



# Triggering Mini-CAPTAIN

- Beam facility provides an RF signal for every micro pulse that is then distributed to TPC and photon detection system
- TPC takes data for 4.75 ms when the first RF from a 625  $\mu s$  macro pulse is received
  - 1.85 ms of buffered pre-trigger data
  - ~ 2.3 ms of post-trigger data; cosmics collected during this time (for calibration)
- All 3 micro pulses of a macro pulse fall within the same TPC acquisition window
- Photon detection system receives the RF signal independently and needs to be synced with TPC offline
  - Also allowed to self-trigger, independent of the RF trigger, if enough light is collected
  - Provided an accurate measure of the beginning of the drift time

#### Proton Candidate with a Cosmic Candidate



- Several hundred thousand beam triggers were collected
- Currently analyzed our special lowintensity run sample ("golden sample") - 3 bunches per macropulse
- Good LAr purity ( $e^{-1}$  lifetime > 72  $\mu$ s)
- Low electronics noise
- 1<sup>st</sup> measurement is an absolute total cross-section – also looking for exclusive channels

# **Basic Analysis Strategy**

- Find tracks in the time-projection chamber (TPC) in time with the beam and in the beam spot
- Match the tracks in the TPC to hits in the photon detection system (PDS)
- Use the timing from the PDS to determine the neutron energy kinetic energy bin
- For tracks (track starting position) in each kinetic energy bin, fit an exponential decay function



# Finding Tracks

- Raw signals from TPC wires were filtered for electronic noise – hits were formed from waveform peaks
- Custom-made algorithm found 2D clusters in a single plane and build 3D tracks from one track in collection plane and one track in an induction plane
- Beam is not precisely in the x-direction detector was slightly rotated relative to the beam direction
- Track cannot be more than  $\pm 27$  mm away from the beam path through the TPC

Plenty of possible cosmics and secondary interactions can be seen in addition to beam events



# Hit-finding efficiency

- Exposure of 4.3 hours at WNR yielded 115,880 reconstructed tracks
  - Require the following:
  - at least 10 MeV track
  - within 32  $\mu$ s of an RF signal
  - no more than ±27mm from beam line
    - $\rightarrow$  These cuts yield 9,911 tracks
- Lower x (downstream) has higher efficiency than upstream

   data analyses only uses downstream tracks (starting at or below wire 180) to avoid this inefficiency problem

 $\rightarrow$  Left with 2,631 tracks after this selection criteria



Gaps in the plot are wires that have been masked due to low efficiency

#### Cross section measurement

 Tracks are binned based on their kinetic energy – distributions of starting positions (x) fit in the direction of the beam with an exponential function

$$N(x) = N_0 e^{-n\sigma_T x} \qquad n = \rho_{Ar} \times \frac{N_A}{m_{Ar}} = 2.11 \times 10^{22} \text{ cm}^{-3} \qquad \begin{array}{l} \rho_{Ar} = 1.3973 \text{ g/cm}^3 \\ N_A = 6.022 \times 10^{22} \text{ n/mol} \\ m_{Ar} = 39.948 \text{ g/mol} \end{array}$$

Cross section extracted in each kinetic energy bin (shown below)

TABLE I. Neutron cross section in bins of kinetic energy. The  $\chi^2$  per degrees of freedom is presented, as well as the total number of tracks used for the fit in each bin. The exact functional form used for the fits is  $f(x) = c_1 e^{-c_2 x}$ . The cross section is extracted from the fitted parameter according to Equation 1.

Energy range [MeV]	Cross Section [barns]	$\chi^2/\mathrm{ndof}$	Number of tracks
100-199	$0.49{\pm}0.34$	1.48/3	264
199-296	$0.88{\pm}0.16$	11.81/7	536
296-369	$0.89{\pm}0.26$	4.739/5	329
369-481	$0.94{\pm}0.20$	8.262/6	413
481-674	$1.20{\pm}0.18$	5.713/6	624
674-900	$0.83 \pm 0.32$	0.1323/4	252

#### Cross section Measurement

- Current cross section measurement is dominated by statistical uncertainty
  - Uncertainty on position reconstruction of tracks was less than 4% and is negligible compared to statistical uncertainty
  - Reconstructed angle of track has less than a few percent uncertainty
  - Multiple track events could introduce as much as 10% variation in our cross-section – still not as large as statistical uncertainty
- Measured cross section is more or less consistent with models used in GEANT4 (QGSP\_BERT) and FLUKA. However, some fine-tuning of models could provide better agreement with the Mini-CAPTAIN data.



Note: noticeable deviation of data from models around 600 MeV

#### Conclusions

• Measured total neutron cross section on argon by Mini-CAPTAIN consistent with an energy averaged cross section of:

#### $0.91 \pm 0.10$ (stat.) $\pm 0.09$ (sys.) barns

- Measurements will help constrain uncertainties in models of neutron transport and help improve reconstruction performance in DUNE
- Data set used for this analysis corresponded to lowest intensity beam configuration – in future we will include additional data sets in other beam configurations