

# Hadron Measurement with Transition Radiation

*Intro*

*TR for hadrons in DUNE*

*Rates*

*Optics and photon detection*

*19 Apr 2019*

*Leo Bellantoni*

*Detector working group*

# Hadron Measurement with Transition Radiation

**Warning: This idea  
does not quite work yet**

*19 Apr 2019*

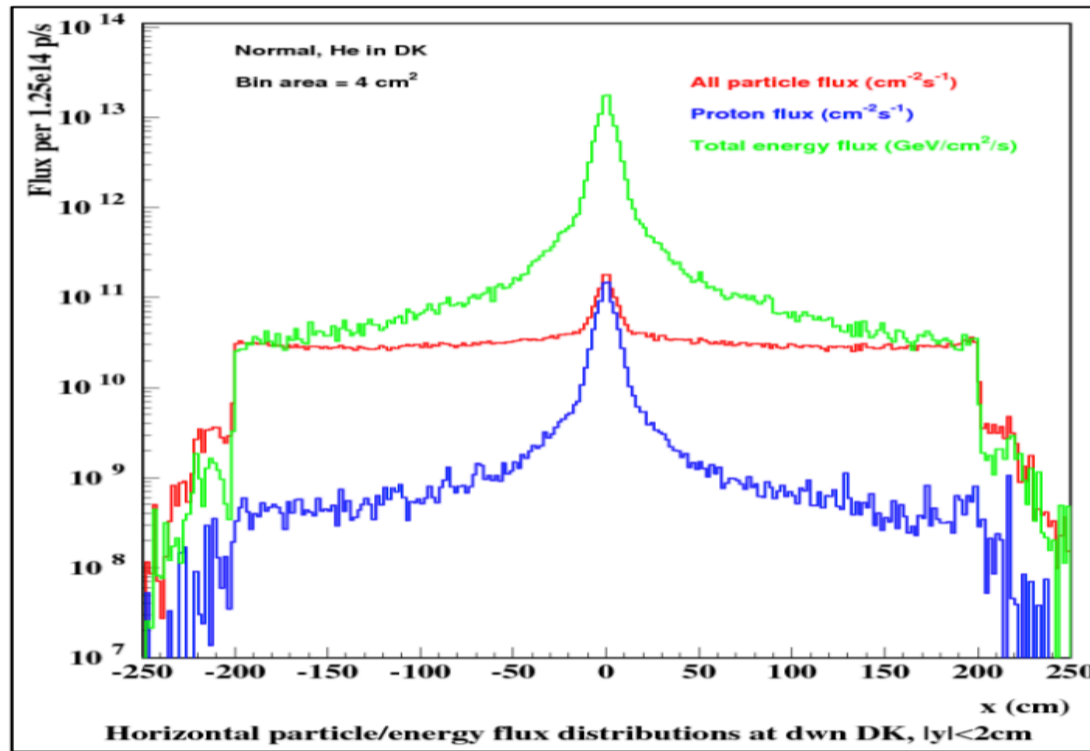
*Leo Bellantoni*

*ASR working group*

# Introduction

- We would *really* like to know about the hadrons in the decay pipe
- To this end, we had a fairly advanced proposal to put a spectrometer in the beamline
- I think the last experiment to try to instrument the decay pipe was K2K. They put a Cherenkov counter upstream of the hadron monitors; set the index of refraction,  $n$ , so that it was blind to protons that had not interacted in the target.
  - The analysis did go into their final result
  - They didn't try it again for T2K
  - Setting  $n$  like that makes them blind to most of the  $\pi^\pm$  and all of the  $K^\pm$  spectrum  $\Rightarrow$  Not *super* useful

# Introduction



CDR Annex 3  
Fig 2-91

2.4 MW on target

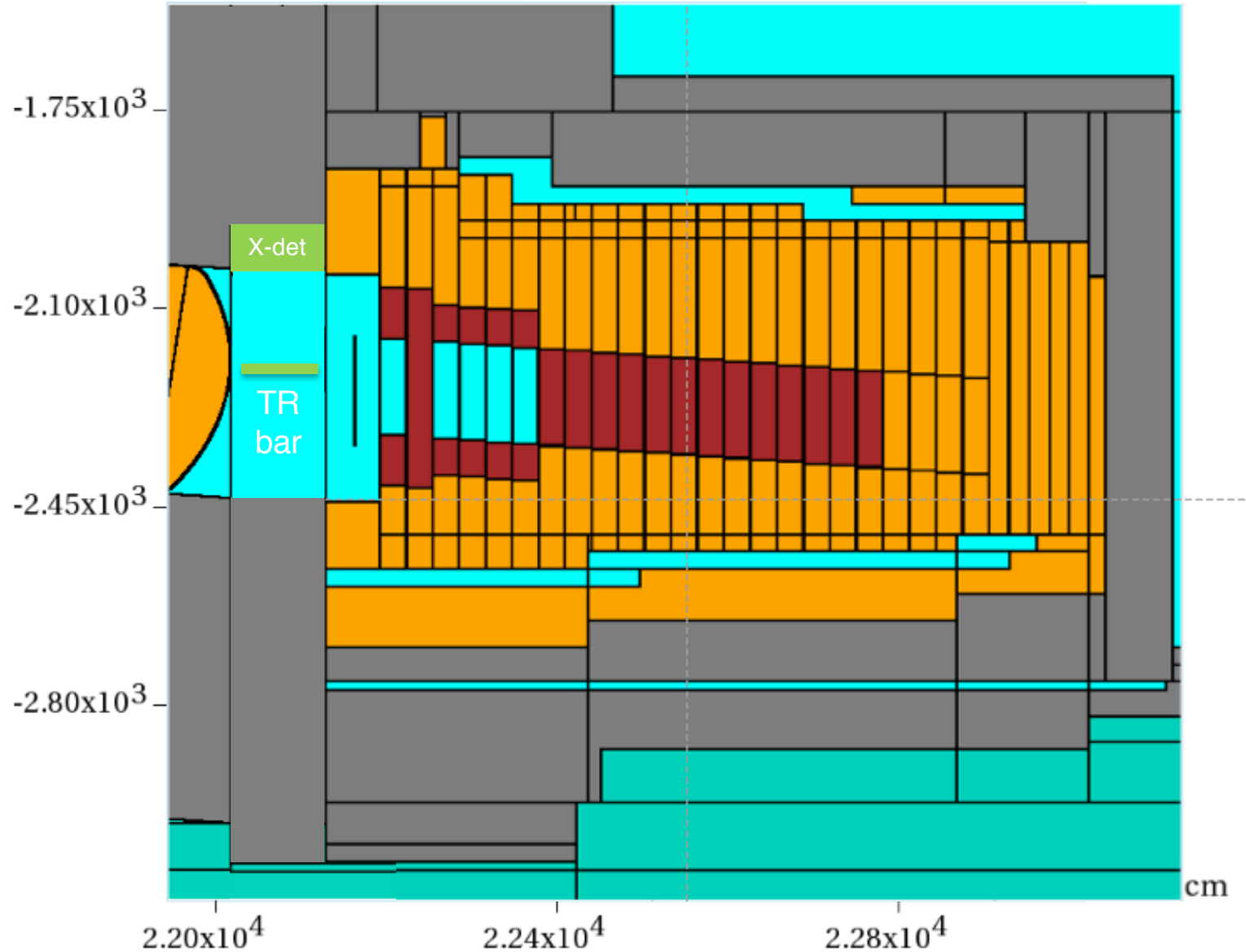
time-averaged  
rate

(Mokhov)

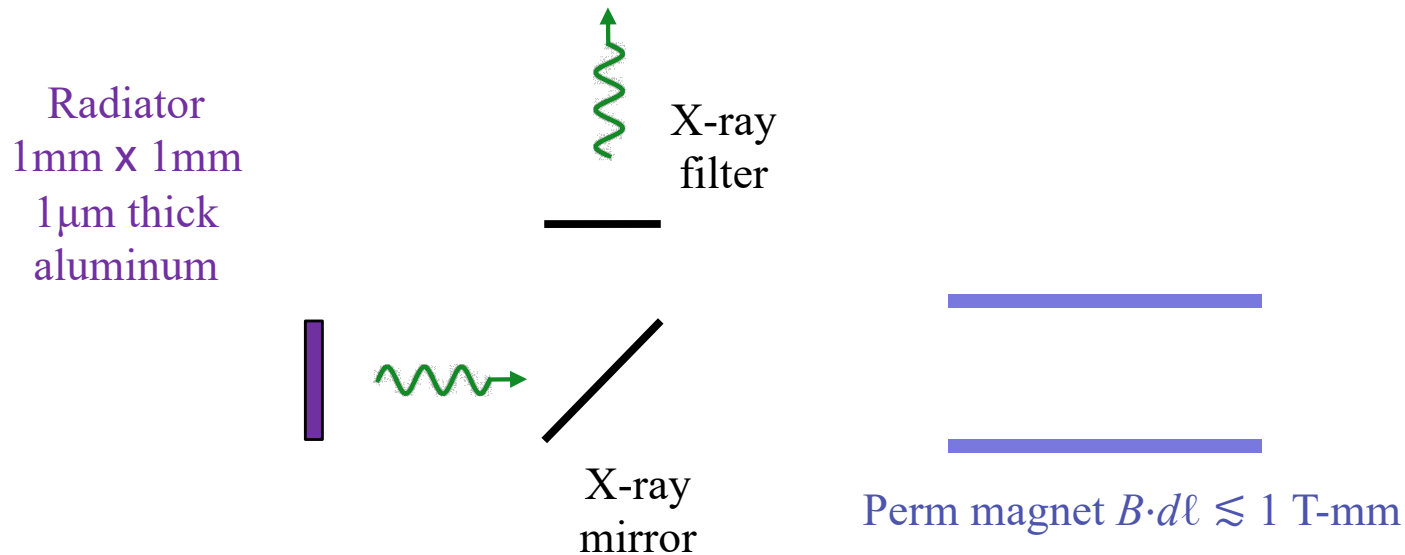
**Our protons are a large part of the flux only within 30cm of the beamline axis... could think of Cherenkov outside that region. But rates are very high... so try transition radiation which typically produces less light than Cherenkov radiation**

# TR for hadrons in DUNE

- Need vacuum between TR bar &  $\gamma$  det, but could be poor vacuum - pump & close
- Smaller intrusion than spectrometer
- No PID though - just spectrum of  $\gamma$  for charged particles at a single (or few) points, angles in the beam



## TR bar - upstream end

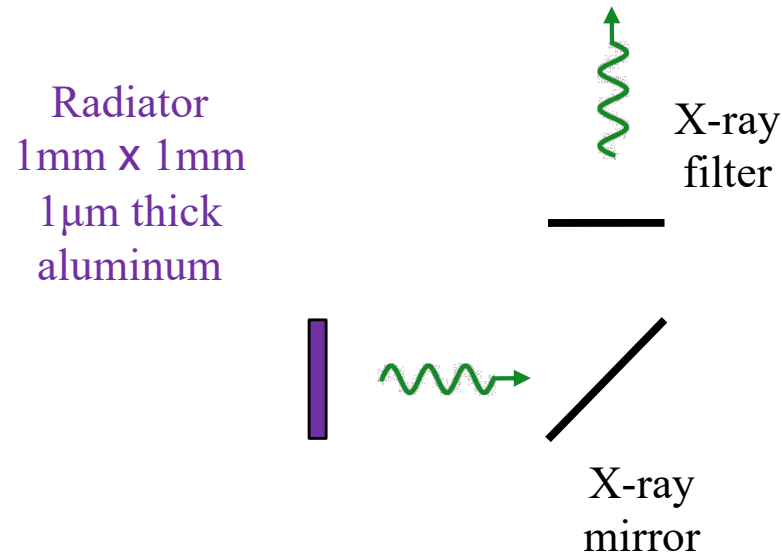


Al has high plasma frequency  $\omega_p = 32.9 \text{ eV}$ . 1µm thickness for absorption.  
Absorption limits to  $E(\gamma) < 70 \text{ eV}$  (even from the back)

X-ray mirrors use total external reflection (metals have  $\Re(n) < 1$  in X-ray)  
For Au a 70 eV X-ray reflects through  $\leq 2 \times 66^\circ$

There will be electrons of low energy from 2ndary emission upstream  
If they are below 100 MeV, and looking for a  $\sim 4 \text{ mrad}$  bend,  
need  $B \cdot \ell = 1.3 \text{ T mm}$ . That's 1mm of Neodymium magnet

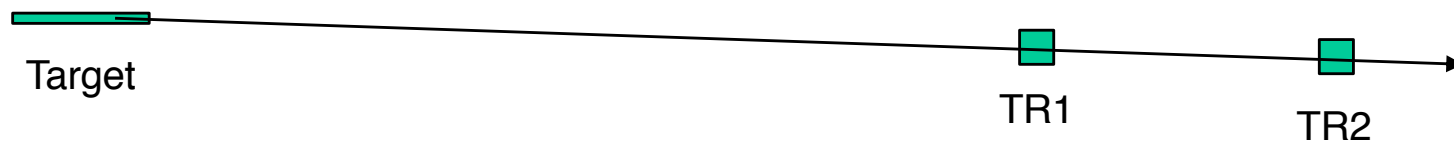
## TR bar – downstream end



The same thing, but no sweeper magnet.  
1 m downstream to get ~1mrad resolution

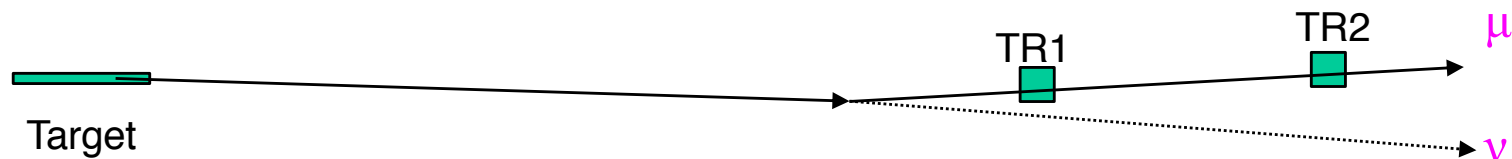
# Rate calculation

*Existing rate calculation is really very approximate. A  $10\mu\text{s}$  spill will have between 1550 and 15100 signal coincidences depending on  $\gamma$ . But this calculation assumes all the flux is in a straight line from target*



*Actually a lot of the beam is muons that have substantial angle re target  
Could tilt TR2 relative to TR1 and go off axis to see mostly muons,*

- 1) reducing the signal coincidence rates*
- 2) getting a more direct measure of the  $\pi \rightarrow \nu\mu$  decay*
- 3) and a much lower coincidence rate*





# Photon detection

Even behind the cement, there is still a large flux of  $\mu^\pm$

They are very different from the X-ray photons – they form tracks, have different energy deposits. There are plenty of discriminants between the two. But hits from tracks can kill signal acceptance

Want a purely photonic system rather than e.g. Si or wires

Tetraphenyl butadiene, p-terphenyl, and diphenyl stilbene have been used at 58.4 nm, not that far from that filter's bandpass

*McKinsey et.al., NIM 132 (1997) 351*

Princeton Instruments uses proprietary compound (which might be  $\text{GdO}_2\text{S:Tb}$ ) for this energy range

Some literature work needed here...

## Obviously a long way to go

Photon detection: will the fluor compound have a light component with long lifetime? That could get excited by  $\mu$  (even behind the cement) and saturate all the pixels

Singles rate still pretty high for full intensity; a very challenging hardware level trigger is most likely needed

What do we learn from measurement of the  $\gamma$  spectrum at a few points & directions at end of pipe?

Radiation hardness of permanent magnet?

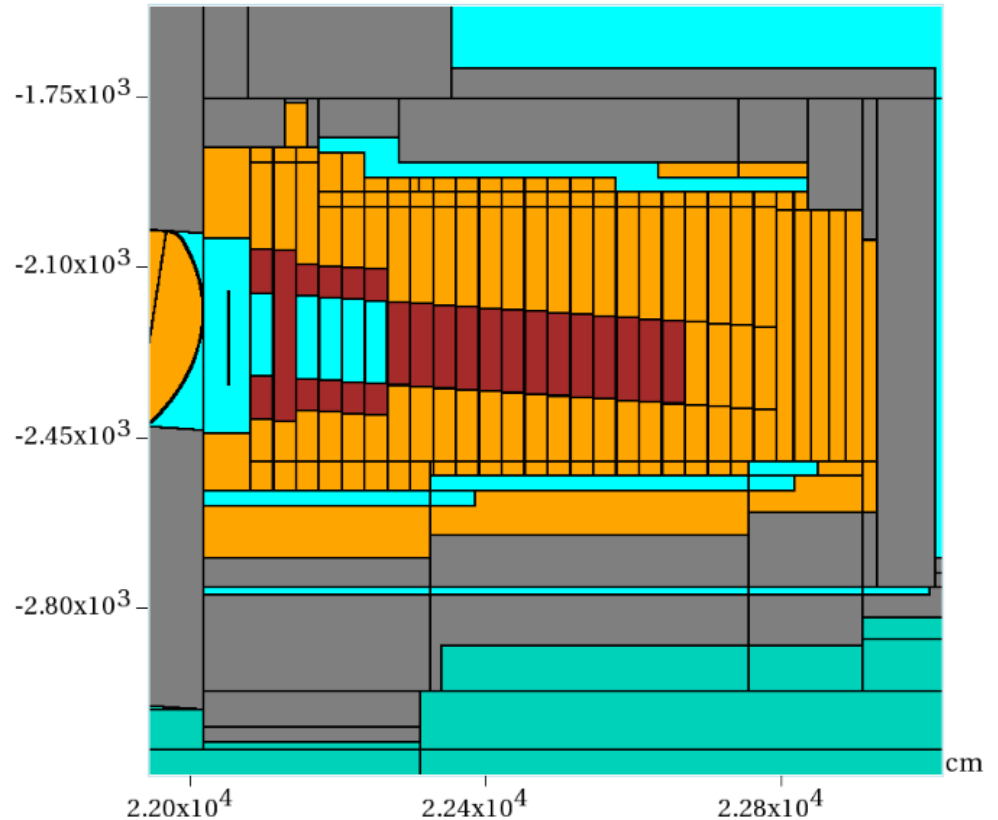
Calibration – can't get an absolute spectrum without it

Alignment & thermal shock

# In A Single Slide

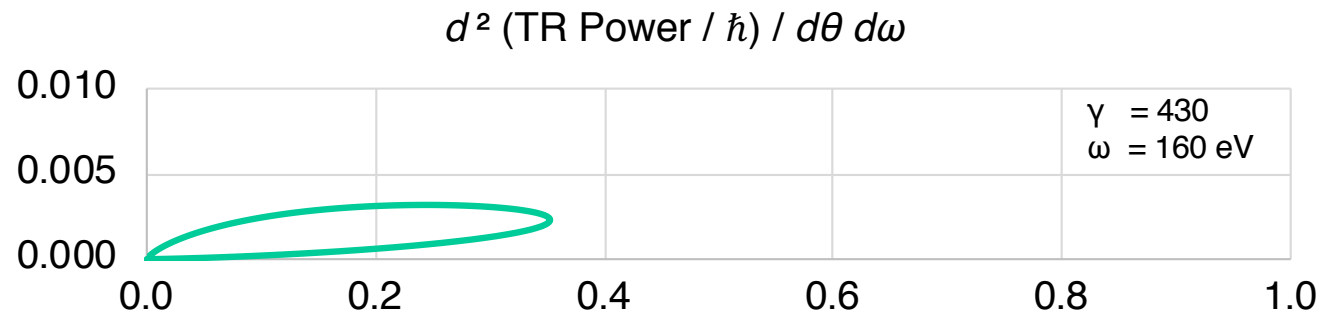
## TR for hadrons in DUNE

- Not a lot of space between decay pipe & absorber – about 0.6m & the hadron monitor has to fit there
- Assume that  $\approx 1\text{m}$  could be found between decay pipe and hadron monitor
- Detection of the TR has to be behind cement, above beamline, and photonic not electronic
- Two TR radiators,  $\sim 1\text{m}$  apart, make X-ray TR, followed by mirrors to bounce it up behind the cement



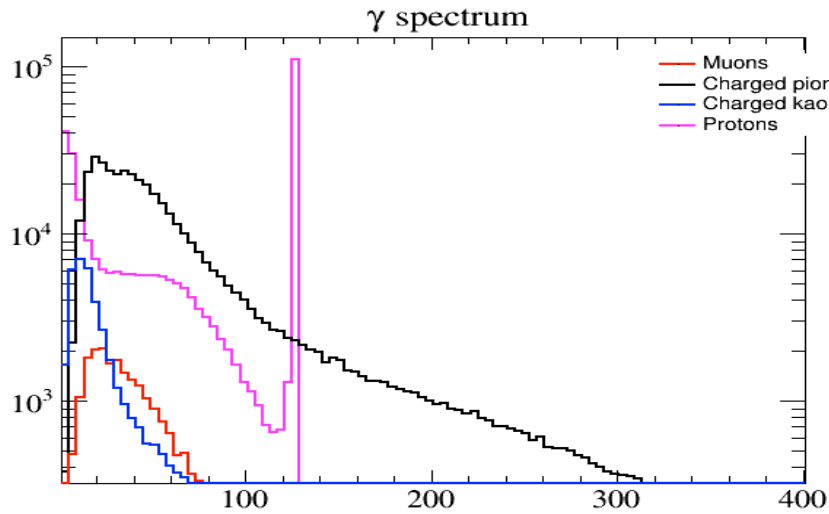
# A Transition Radiation Primer

- A moving charged particle in vacuum has some EM fields and does not radiate
- That same particle in a material (of plasma frequency  $\omega_p$ ) has different fields and does not radiate either
- The transition from one set of fields to the other accelerates the electrons in the material, thereby producing radiation
- Most of the radiation is forward, at angle  $\theta \approx 1/\gamma$



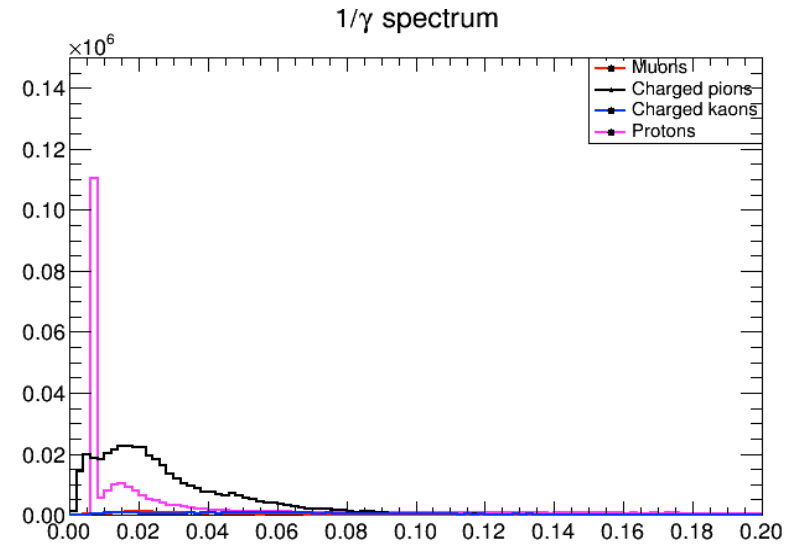
- Total radiated energy is proportional to  $\gamma$ :  $I = \alpha \hbar \omega_p \gamma / 3$
- Nearly all the radiation happens at frequencies below  $\gamma \omega_p$
- Interested in  $\gamma \approx 2$  (1 GeV  $K^\pm$ ) to  $\gamma \approx 250$  (35 GeV  $\pi^\pm$ )

# TR for hadrons in DUNE



Showing particles that enter g41bne decay pipe at angle so that they'd exit end of pipe if they don't decay

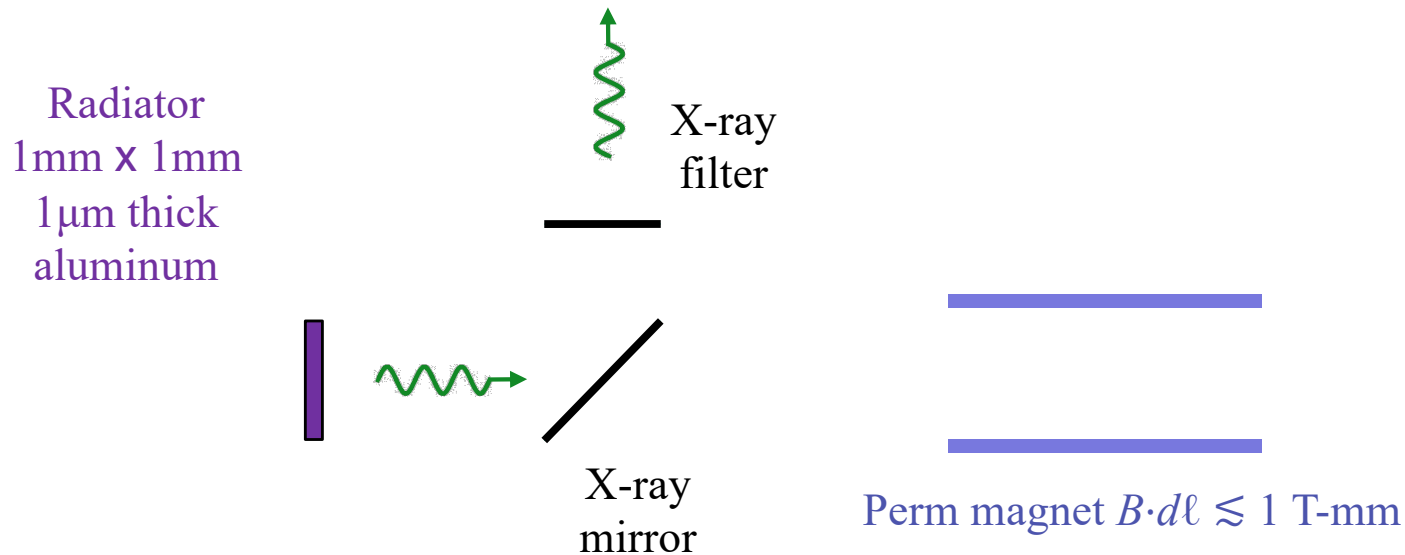
$e^\pm$  are bright sources of TR but we have them only from interactions in upstream material and so are at low momentum



$1/\gamma$  spectrum suggests we want angular resolution  $\sim 1\text{-}2$  mrad

Angle spanned by decay pipe endcap  $\approx 9$  mrad

## TR bar – upstream end



Al has high plasma frequency  $\omega_p = 32.9 \text{ eV}$ . 1µm thickness for absorption.

Lumpkin et al. use this for their optical TR radiators.

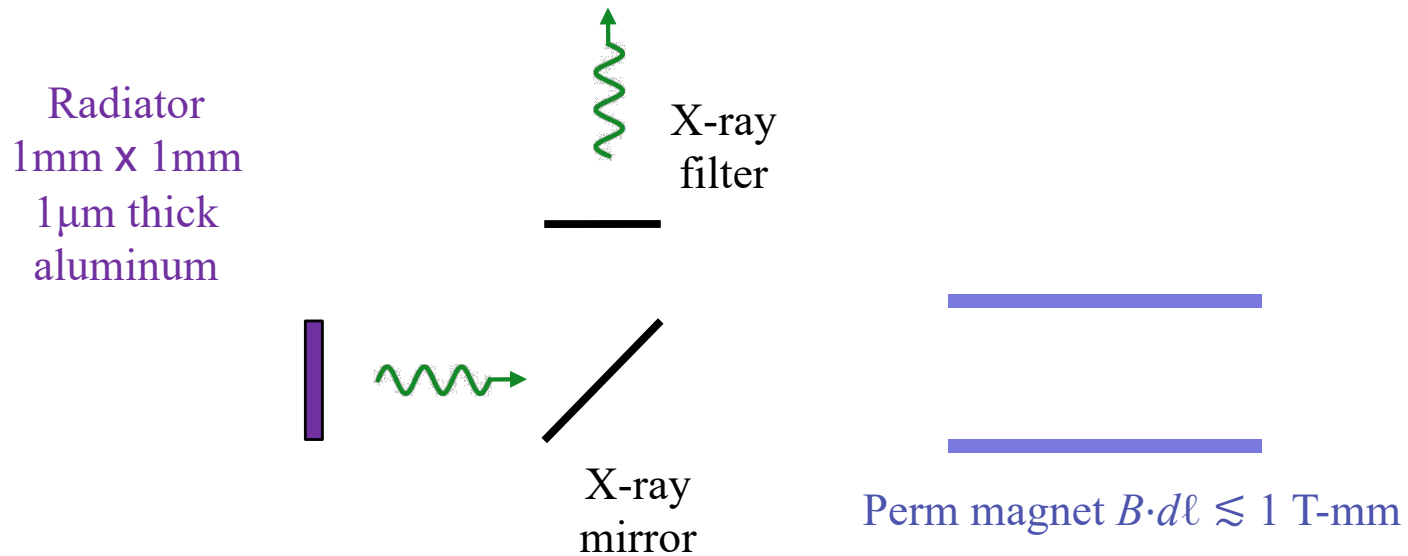
Absorption limits to  $E(\gamma) < 70 \text{ eV}$  (even from the back)

X-ray mirrors use total external reflection (metals have  $\Re(n) < 1$  in X-ray)

Only grazing angles reflect &  $(\theta_{\max})(E_\gamma [\text{eV}]) = \text{const.}$

For Au, const.  $\sim 80 \text{ mrad keV}$ , so a 70 eV X-ray reflects through  $\leq 2 \times 66^\circ$

## TR bar – upstream end



There will be electrons of low energy from 2ndary emission upstream  
If they are below 100 MeV, and looking for a ~4 mrad bend,  
need  $B \cdot \ell = 1.3$  T mm. That's 1mm of Neodymium magnet



# X-ray filter



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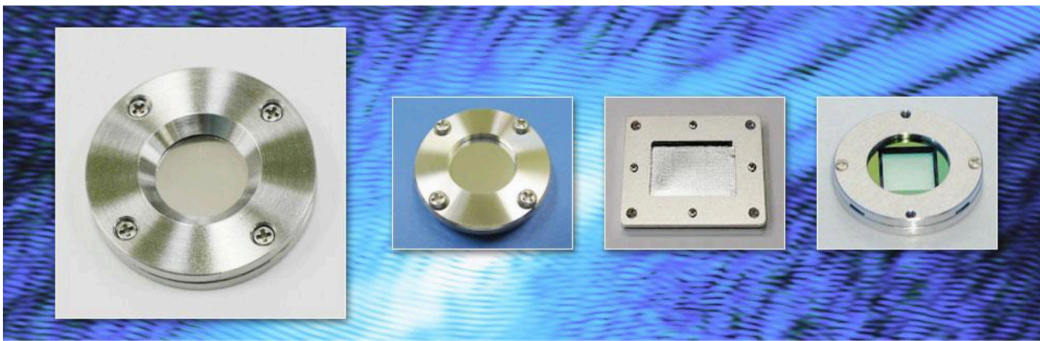
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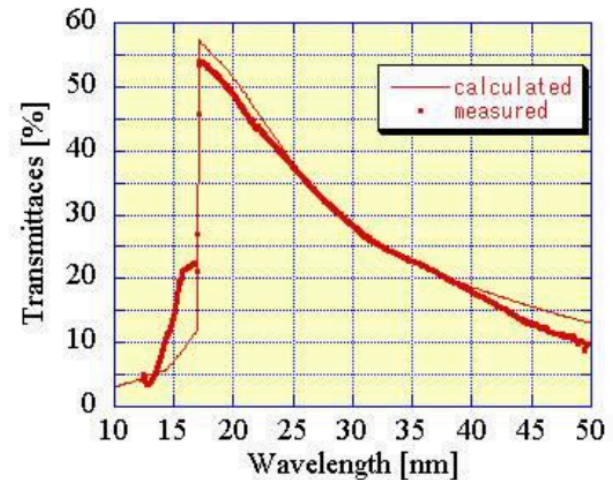
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## XUV filters



Premium XUV filters with high transmittance and long lifetime



X-ray filters are commercially available

They are ~100nm thin films 1cm across – pulse heating could be a mechanical issue

I'll use 18 – 45nm (27-70 eV) bandpass for flux calculations, 25% efficiency

# Rate calculation

*The PDG gives the  $N^0$  of photons with  $E > \hbar\omega_0$  for 1 transition*

$$N_\gamma(\hbar\omega > \hbar\omega_0) = \frac{\alpha}{\pi} \left[ \left( \ln \frac{\gamma \hbar\omega_p}{\hbar\omega_0} - 1 \right)^2 + \frac{\pi^2}{12} \right]$$

*not exact as the derivation assumes  $\gamma \gg 1$ , but we want to go down to  $\gamma = 2$*

*For range 18 – 45 nm,  $N_\gamma$  ranges from 0.0083 ( $\gamma = 2$ ) to 0.0259 ( $\gamma = 250$ )*

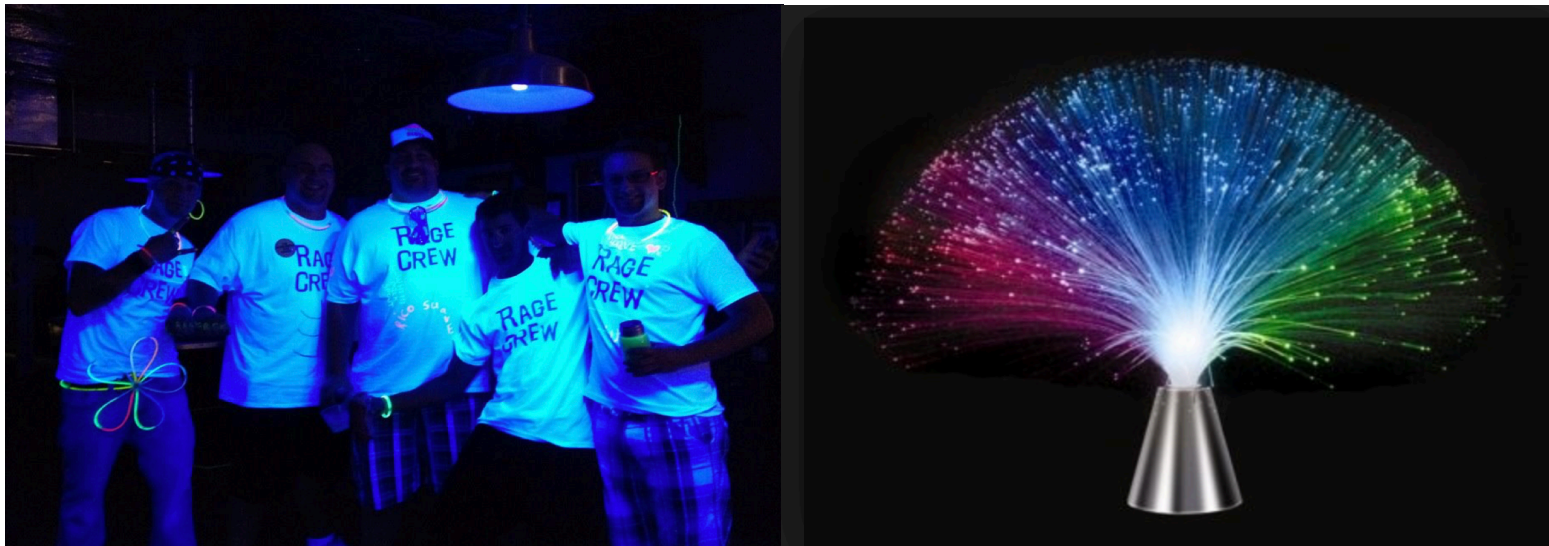
*Instantaneous flux is (1.2 s/10  $\mu$ s)( $\sim 3 \times 10^{10}$  Hz/cm<sup>2</sup>) =  $3.6 \times 10^{15}$  Hz/cm<sup>2</sup>;  
times 0.1 cm x 0.1 cm, that's 36 THz.*

*Times (0.0083  $\cdot$  0.25)<sup>2</sup>, 155 MHz; times (0.0259  $\cdot$  0.25)<sup>2</sup>, 1510 MHz  
assuming no losses in reflection and detection (neither is plausible)*

*For 10  $\mu$ s, between 1550 and 15100 signal coincidences per cycle*

# Photon detection

Maybe a wavelength shifter & light pipes to a low-rad zone?



Tetraphenyl butadiene, p-terphenyl, and diphenyl stilbene have been used at 58.4 nm, not that far from that filter's bandpass  
*McKinsey et.al., NIM 132 (1997) 351*

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# Photon detection

