

1 Deep Underground Neutrino Experiment (DUNE)

2 DRAFT Technical Design Report

3 **Volume 1: Executive Summary for DUNE**

4 Chapter Breakout:
5 Near Detector
6

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8
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Chapter 1

Near Detector Executive Summary

1.1 Brief Overview of the DUNE Near Detector

1.1.1 Need for the Near Detector

A primary aim of the DUNE experiment is to measure the oscillation probabilities for muon neutrino and muon antineutrinos to either remain the same flavor or oscillate to electron (anti)neutrinos. Measuring these probabilities as a function of the neutrino energy will allow definitive determination of the neutrino mass ordering, observation of leptonic charge parity (CP) violation for a significant range of δ_{CP} values, and precision measurement of PMNS parameters.

The role of the near detector (ND) is to serve as the experiment's control. The ND establishes the null hypothesis and constrains systematic errors. It measures the initial unoscillated ν_μ and ν_e energy spectra: the convolution of flux, cross section, and detector response. To first order, a "Far/Near" ratio (or migration matrix), derived from the simulation, can predict the unoscillated energy spectrum at the far detector (FD) based on the ND measurements. The energy spectra at the FD are then sensitive to the oscillation parameters, which can be extracted via a fit. Included in this, but worth emphasizing, the ND plays a critical role in establishing what the oscillation signal spectrum should look like in the FD because the expectations for the spectra in both the disappearance and appearance signals are based on the precisely measured spectra for ν_μ and $\bar{\nu}_\mu$ interactions in the ND.

To achieve the precision needed for DUNE, the experiment will have to operate beyond the first-order paradigm. With finite energy resolution and non-zero biases, the reconstructed energy spectrum is an unresolved convolution of cross section, flux, and energy response. The ND must independently constrain each of those components. The ND must provide information that can be used to model well each component. Models of the detector, beam, and interactions fill in holes and biases left by imperfect understanding and they are used to estimate the size of many systematic effects. When imperfect models are not able to match observations, the ND must provide the

1 information needed to deal with that and estimate the impact of the imperfect modeling on final
2 measurements. In general, this requires that the ND significantly outperform the FD, which is
3 limited by the need for a large, underground, mass. The ND must have multiple methods for
4 measuring neutrino fluxes as independently from cross section uncertainties as possible. With the
5 necessity of relying on models, the ND needs to measure neutrino interactions with much better
6 detail than the FD. This includes having a larger efficiency across the kinematically allowed phase
7 space of all relevant reaction channels, superior identification of charged and neutral particles,
8 better energy reconstruction, and better controls on experimental biases. The ND must also have
9 the ability to measure events in a similar way to the FD, so that it can determine the ramifications
10 of the more limited FD performance, provide corrections, and take advantage of effects cancelling
11 to some extent in the near to far extrapolation.

12 The conceptual design of the ND is based on the collective experience of the many DUNE collab-
13 orators who have had significant roles in the current generation of neutrino experiments (MINOS,
14 MiniBooNE, T2K, NO ν A, MINER ν A, and the SBN program). These programs have provided
15 (and will provide) a wealth of useful data and experience that has led to improved neutrino in-
16 teraction models, powerful new analyses and reconstruction techniques, a deep appreciation of
17 analysis pitfalls, and a better understanding of the error budget. These experiments, while simi-
18 lar to DUNE, were all either done at a lower precision, in a different energy range, or with very
19 different detector technologies. While the existing and projected experience and data from those
20 experiments provides a strong base for DUNE, it is not sufficient to enable DUNE to accomplish
21 its physics goals without a highly performing ND.

22 The DUNE ND will also have a physics program of its own measuring cross sections, nonstandard
23 interactions (NSIs), searching for sterile neutrinos, dark photons and other exotic particles. These
24 are important aims that expand the physics impact of the ND complex. Also the cross section
25 program is coupled to the oscillation measurement in so far as the cross sections will be useful as
26 input to theory and model development. (Note that many of the ND data samples are incorporated
27 into the oscillation fits directly.) The DUNE ND program of beyond the standard model physics
28 is discussed more in Appendix 1.7.

29 **1.1.2 Overview of the Near Detector**

30 The DUNE ND is formed from three primary detector components and the capability for two of
31 those components to move off the beam axis. The three detector components serve important indi-
32 vidual and overlapping functions with regard to the mission of the ND. Because these components
33 have standalone features, the DUNE ND is often discussed as a suite or complex of detectors and
34 capabilities. The movement off axis provides a valuable extra degree of freedom in the data which
35 is discussed extensively in this report. The power in the DUNE ND concept lies in the collective
36 set of capabilities. It is not unreasonable to think of the component detectors in the DUNE ND
37 as being somewhat analogous to subsystems in a collider experiment, the difference being that,
38 with one important exception (higher momentum muons), individual events are contained within
39 the subsystems. The DUNE ND is shown in the DUNE ND hall in Figure 1.1. Table 1.1 provides
40 a high-level overview of the three components of the DUNE ND along with the off-axis capability
41 that is sometimes described as a fourth component.

Table 1.1: This table gives a high-level breakdown of the three major detector components and the capability of movement for the DUNE ND along with function and primary physics goals.

Component	Essential Characteristics	Primary function	Select physics aims
LArTPC (ArgonCube)	Mass	Experimental control for the Far Detector	$\nu_\mu(\bar{\nu}_\mu)$ CC
	Target nucleus Ar	Measure unoscillated E_ν spectra	ν - e^- scattering
	Technology FD-like	Flux determination	$\nu_e + \bar{\nu}_e$ CC Interaction model
Multipurpose detector (MPD)	Magnetic field	Experimental control for the LArTPCs	$\nu_\mu(\bar{\nu}_\mu)$ CC
	Target nucleus Ar	Momentum analyze liquid Ar μ	ν_e CC, $\bar{\nu}_e$
	Low density	Measure exclusive final states with low momentum threshold	Interaction model
DUNE-PRISM	LArTPC+MPD move off-axis	Change flux spectrum	Deconvolve xsec*flux Energy reponse Provide FD-like energy spectrum at ND ID mismodeling
3D scintillator tracker spectrometer (3DSTS)	On-axis	Beam flux monitor	On-axis flux stability
	Mass	Neutrons	Interaction model
	Magnetic field CH target		A dependence ν - e^- scattering

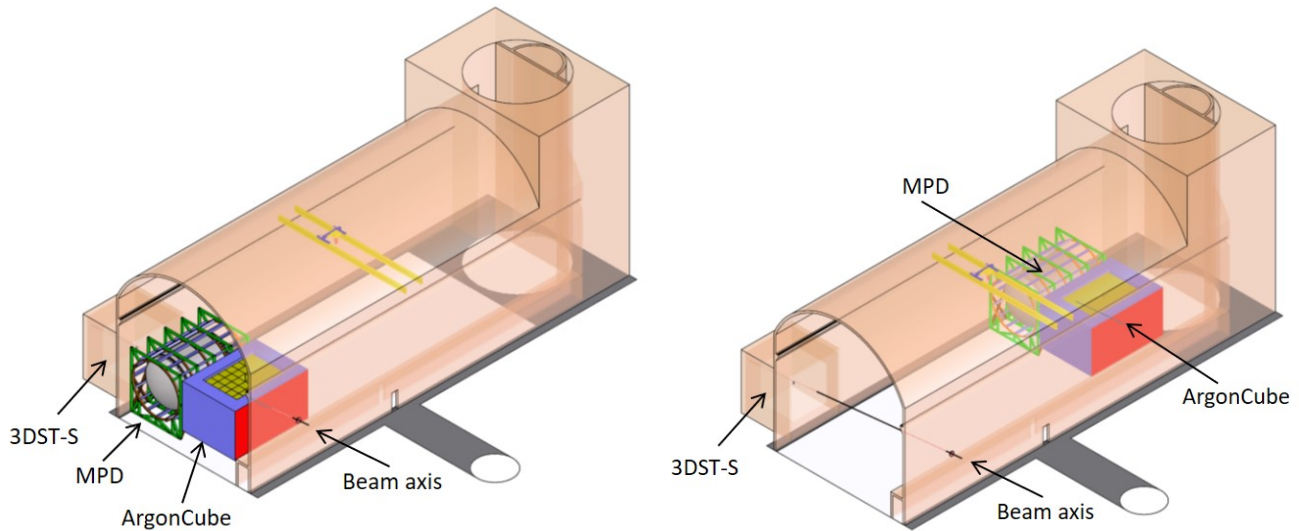


Figure 1.1: DUNE ND hall shown with component detectors all in the on-axis configuration (left) and with the liquid argon time-projection chamber (LArTPC) and MPD in an off-axis configuration (right). The 3DST-S is shown in position on the beam axis. The beam axis is shown. The beam enters the hall at the bottom of the drawings moving from right to left.

1 The core part of the DUNE ND is a LArTPC called ArgonCube. The particular implementation
 2 of the LArTPC technology in this detector is described in Section 1.3.2 below. This detector has
 3 the same target nucleus and shares some aspects of form and functionality with the FD, where the
 4 differences are necessitated by the expected intensity of the beam at the ND. This similarity in
 5 target nucleus and, to some extent, technology, reduces sensitivity to nuclear effects and detector-
 6 driven systematic errors in the extraction of the oscillation signal at the FD. The LArTPC is large
 7 enough to provide high statistics ($1 \times 10^8 \nu_\mu$ -CC events/year) and a sufficient volume to provide
 8 good hadron containment. The tracking and energy resolution, combined with the mass of the
 9 LArTPC, will allow for the measurement of the flux in the beam using several techniques, including
 10 the rare process of ν - e^- scattering.

11 The LArTPC begins to lose acceptance for muons above ~ 0.7 GeV/c momentum due to lack
 12 of containment. Because the muon momentum is a critical component of the neutrino energy
 13 determination, a magnetic spectrometer is needed downstream of the LArTPC to measure the
 14 charge sign and momentum of these muons. In the DUNE ND concept, this function is accom-
 15 plished by the multipurpose detector (MPD) which consists of a high-pressure gaseous argon time
 16 projection chamber (TPC) (HPgTPC) surrounded by an electromagnetic calorimeter (ECAL) in
 17 a 0.5 T magnetic field. The HPgTPC provides a lower density medium with excellent tracking
 18 resolution for the muons from the LArTPC. In addition, with this choice of technology for the
 19 tracker, neutrinos interacting on the argon in the gas TPC constitute a sample of ν -Ar events that
 20 can be studied with a very low charged particle tracking threshold and excellent resolution and
 21 systematic errors that differ from the liquid detector. The high pressure results in a sample of
 22 $2 \times 10^6 \nu_\mu$ -CC events/year for these studies. These events will be valuable for studying the charged
 23 particle activity near the interaction vertex since this detector can access lower momenta protons
 24 than the liquid argon (LAr) detector and has better particle identification of charged π . The lack
 25 of secondary interactions in these samples will be helpful for identifying the particles produced in

1 the primary interaction and modeling secondary interactions in denser detectors, which are known
2 to be important [3]. In addition, many neutrons with high kinetic energy produced in neutrino
3 interactions in the gaseous argon may be reconstructable via time-of-flight using the ECAL. The
4 MPD is discussed further in Section 1.3.3.

5 The LArTPC begins to lose acceptance for muons above $\sim 0.7\text{GeV}/c$ due to lack of containment.
6 Because the muon momentum is a critical component of the neutrino energy determination, a
7 magnetic spectrometer is needed downstream of the LArTPC to measure the charge sign and mo-
8 mentum of these muons. In the DUNE ND concept, this function is accomplished by the MPD
9 which consists of a HPgTPC surrounded by an ECAL in a 0.5T magnetic field. The HPgTPC pro-
10 vides a lower density medium with excellent tracking resolution for the muons from the LArTPC.
11 In addition, with this choice of technology for the tracker, neutrinos interacting on the argon in
12 the HPgTPC constitute a large (approximately $1 \times 10^6 \nu_\mu\text{-CC}$ events/year on axis) independent
13 sample of $\nu\text{-Ar}$ events that can be studied with a very low charged particle tracking threshold and
14 excellent resolution and with systematic errors that differ from the liquid detector. These events
15 will be valuable for studying the charged particle activity near the interaction vertex, since this de-
16 tector can access lower-momentum protons than the LArTPC and has better particle identification
17 of charged pions. Misidentification of pions as knock-out protons (or vice versa) causes a mistake
18 in the reconstructed neutrino energy, moving it away from its true value by the amount of a pion
19 mass. This mistake can become quite significant at the second oscillation maximum. The gas
20 detector will play an important role in mitigating this mistake, since pions are rarely misidentified
21 as protons in the HPgTPC. In addition, the lack of secondary interactions in the gas samples will
22 be helpful for identifying the particles produced in the primary interaction and modeling secondary
23 interactions in denser detectors, which are known to be important effects[3]. The high pressure
24 increases the statistics for these of studies, improves the particle identification capabilities, and
25 improves the momentum resolution. The MPD is discussed further in Section 1.3.3.

26 The LArTPC and MPD can move to take data in positions off the beam axis. This capability is
27 referred to as DUNE Precision Reaction-Independent Spectrum Measurement (DUNE-PRISM).
28 As the detectors move off-axis, the incident neutrino flux spectrum changes, with the mean energy
29 dropping and the spectrum becoming somewhat monochromatic. Though the neutrino interaction
30 rate drops off-axis, the intensity of the beam and the size of the LArTPC combine to yield ample
31 statistics even in the off-axis positions. Data taken at different off-axis angles allows for the decon-
32 volution of the neutrino flux and interaction cross section and the mapping of the reconstructed
33 versus true energy response of the detector. This latter mapping is applicable at the FD up to the
34 level to which the near and far LAr detectors are similar. Stated a different way, it is possible to
35 use information from a linear combination of the different fluxes to create a data sample at the
36 ND with an effective neutrino energy distribution that is close to that of the oscillated spectrum
37 at the FD. This data-driven technique will reduce systematic effects coming from differences in
38 the energy spectra of the oscillated signal events in the FD and the ND samples used to constrain
39 the interaction model. Finally, the off-axis degree of freedom provides a sensitivity to some forms
40 of mismodeling in the beam and/or interaction models. The DUNE-PRISM program is discussed
41 further in Section 1.3.4.

42 The final component of the DUNE ND suite is the 3D projection scintillator tracker spectrometer
43 (3DST-S), the core part of which is the 3D scintillator tracker (3DST). The 3DST is a plastic
44 scintillator detector made of 1 cm cubes that are read out along each of three orthogonal dimensions.

1 The design eliminates the typical planar-strip geometry common to detectors using scintillator,
 2 leading to improved acceptance at large angles relative to the beam direction. The 3DST is
 3 situated along the beam axis inside an envelope of high resolution, normal pressure TPCs and an
 4 ECAL. The entire structure is enclosed in a magnet. This device importantly serves as a dedicated
 5 neutrino spectrum monitor that stays on-axis when the LArTPC and MPD have moved to an off-
 6 axis position. It also provides an excellent on-axis, neutrino flux determination using many of
 7 the methods discussed in Appendix 1.8. The neutrino flux determined using this detector, with
 8 differing detector, target, and interaction systematic errors as compared to the LArTPC, can be
 9 used as an important point of comparison and systematic crosscheck for the flux as determined
 10 by the LArTPC. In addition, the 3DST has very fast timing and the ability to isolate small
 11 energy depositions from neutrons in three dimensions. This provides the capability to incorporate
 12 neutrons in the event reconstruction using energy determination via time-of-flight with a high
 13 efficiency. This capability is expected to be useful for the low- ν flux determination since it allows
 14 for tagging of events with a significant neutron energy component, or in providing a way to include
 15 that energy in the calculation. The differing A of the carbon target relative to argon may prove
 16 to be useful for developing models of nuclear effects and building confidence in the interaction
 17 model and the size of numerous systematic errors. Recent electron scattering results on C, Ti,
 18 and Ar targets are described very well by the SuSAv2-MEC superscaling framework and this is
 19 expected to be applicable to neutrinos [4]. The 3DST-S component of the ND is discussed more
 20 in Section 1.4.1.

21 Table 1.2 shows the statistics expected in the different ND components for a few processes that
 22 are important for constraining the neutrino flux. Some additional information on constraining the
 23 flux is provided in Appendix 1.8.

Table 1.2: Event rates for processes that can be used to constrain the neutrino flux. The rates are given per year for a 1 ton (FV) HPgTPC, a 25 ton (FV) LArTPC [5], and a 9 ton (FV) 3DST. The flux for the HPgTPC and LArTPC is from the simulated “2017 engineered” Long-Baseline Neutrino Facility (LBNF) beam with a primary momentum of 120 GeV/c and 1.1×10^{21} POT/year. The flux for the 3DST is the 80 GeV, three-horn, optimized beam with 1.46×10^{21} POT/year. The detectors are assumed to be on-axis. Fiducial volumes are analysis dependent and in the case of the LArTPC, it is likely the volume could be made larger by a factor of two for many analyses, and perhaps as much as three times larger for the νe scattering measurement.

Event class	HPgTPC	LArTPC	3DST
$\nu_\mu + e^-$ elastic ($E_e > 500$ MeV)	1.3×10^2	3.3×10^3	1.1×10^3
ν_μ low- ν ($\nu < 250$ MeV)	2.1×10^5	5.3×10^6	1.48×10^6
ν_μ charged current (CC) coherent	8.8×10^3	2.2×10^5	
$\bar{\nu}_\mu$ CC coherent	8.4×10^2	2.1×10^4	

24 The rest of the ND executive summary discusses the ND mission in general and provides signifi-
 25 cantly more detail on the characteristics and capabilities of the three components of the ND and
 26 the DUNE-PRISM program, and how the data from the different detectors feed into the overall
 27 DUNE physics strategy.

1.2 Role of the ND in the DUNE Oscillation Program

Oscillation experiments need to accomplish three main tasks. First, they must identify the flavor of interacting neutrinos in CC events, or identify the events as neutral current (NC) interactions. Second, they need to measure the energy of the neutrinos since oscillations occur as a function of baseline length over neutrino energy, L/E . Third, they need to compare the observed event spectrum in the FD to predictions based on differing sets of oscillation parameters, subject to constraints from the data observed in the ND. That comparison and how it varies with the oscillation parameters allows for the extraction of the measured oscillation parameters and errors.

The connection between the observations in the ND and the FD is made using a simulation that convolves models of the neutrino flux, neutrino interactions, nuclear effects, and detector response. This gives rise to a host of complicating effects that muddy the simple picture. They come from two main sources. First, the identification efficiency is not 100% and there is some background (for example, NC events with a π^0 are a background to ν_e CC interactions). Both the efficiency and the background are imperfectly known. Generally, it is helpful to have a ND that is as similar as feasible to the FD because a bias in the efficiency as a function of energy will cancel between the two detectors. Since the background level tends to be similar between the two detectors, it is helpful if the ND is more capable than the FD at characterizing backgrounds, either due to its technology, or by leveraging the much larger statistics and freedom to take data in alternative beam configuration modes (for example, different horn currents or movement off the beam axis).

The second major source of complication occurs because the FD (and the similar ND) has to be made of heavy nuclei rather than hydrogen. Neutrino interactions can be idealized as a three stage process: (1) a neutrino impinges on a nucleus with nucleons in some initial state configuration, (2) scattering occurs with one of the nucleons, perhaps creating mesons, and (3) the hadrons reinteract with the remnant nucleus on their way out (so called final-state interactions (FSIs)). The presence of the nucleus impacts all three stages in ways that ultimately drive the design of the ND complex. To better understand this it is useful to consider what would happen if the detectors were made of hydrogen.

In a detector made of hydrogen, the initial state is a proton at rest and there is no FSI. The scattering consists of a variety of processes. The simplest is quasi-elastic (QE) scattering: $\bar{\nu}_\ell p \rightarrow \ell^+ n$. The detector sees a lepton (which establishes the flavor of the neutrino), no mesons, and perhaps a neutron interaction away from the lepton's vertex. Because there are no mesons the kinematics is that of two body scattering and the neutrino energy can be reconstructed from the lepton's angle (with respect to the ν beam) and energy. This is independent of whether the neutron is observed.

For ν_ℓ interactions on hydrogen there is no QE process. The simplest scattering channel is single pion production $\nu_\ell p \rightarrow \ell^- \pi^{(+,0)}(n, p)$. In that case the neutrino energy may be reconstructed from the energy of the muon and pion, and their angles with respect to the beam¹. In both cases, the neutrino energy can be measured without bias so long as the detector itself measures lepton and

¹The nucleon does not need to be observed. This is a consequence of having four energy-momentum conservation constraints, which allows E_ν and \vec{p}_N to be computed.

1 meson momenta and angles without bias. The neutrino energy in complicated scattering channels,
2 such as ones with multiple pions or heavy baryons can be measured in a similar way (at least in
3 principle).

4 A key simplifying feature offered by a hypothetical hydrogen detector is simply that there are
5 enough constraints to measure the neutrino energy and also to neglect energy carried off by the
6 single nucleon (especially a neutron escaping the detector). Additionally, the cross sections for
7 different scattering channels (particularly the simpler ones) can be expressed in terms of leptonic
8 and hadronic currents. The leptonic current is well understood. The structural elements of the
9 hadronic current are known on general theoretical grounds. The current is often represented by
10 form factors that are constrained by electron scattering experiments, beta decay, and neutrino
11 scattering measurements that the detector can make itself (or take from other experiments).

12 The situation is significantly more complicated in a detector with heavy nuclei. The nucleons in
13 the initial state of the nucleus are mutually interacting and exhibit Fermi motion. This motion
14 ruins the the key momentum conservation constraint available in hydrogen due to the target being
15 at rest. Scattering at lower momentum transfer is suppressed because the nucleon in the final state
16 has a momentum that is excluded by the Pauli principle.

17 The nucleon momentum distribution in heavy nuclei is commonly modeled as a Fermi gas with a
18 cutoff momentum $k_F \approx 250 \text{ MeV}/c$ [6]. This picture is overly simplistic. For example, there are
19 nucleons with momenta larger than k_F due to short-range correlated nucleon-nucleon interactions
20 (SRC)[7]. Scattering on a nucleon with $p > k_F$ implies that there is a spectator nucleon recoiling
21 against the target with a significant momentum. SRC have been the subject of much investigation
22 and are not fully understood or fully implemented in neutrino event generators.

23 Additionally, there is a second multibody effect. For the few-GeV neutrinos of interest to DUNE,
24 the typical momentum transfer corresponds to a probe that has a wavelength on par with the
25 size of a nucleon. In this case the scattering can occur on two targets in the nucleus which may
26 not be closely correlated (2p2h scattering). Experiments can easily confuse this process for QE
27 scattering since there are no mesons and one or both of the two nucleons may have low energy,
28 evading detection. The presence of two nucleons in the initial and final state again ruins the
29 kinematic constraints available in hydrogen. It is now known that 2p2h scattering is a significant
30 part of the total scattering cross section at DUNE energies [8]. The 2p2h cross section is difficult
31 to compute because it cannot be expressed as the sum over cross sections on individual nucleons.
32 The dependence on atomic number and the fine details of the interaction (e.g., the final energies
33 of the two particles) are also currently unknown. Finally, it is widely expected that there are
34 components of 2p2h and SRC scattering that result in meson production. Event generators do not
35 currently include those processes.

36 Neutrino scattering on nuclei is also subject to FSIs. FSIs collectively refers to the process by
37 which nucleons and mesons produced by the neutrino interaction traverse the remnant nucleus.
38 The hadrons reinteract with a variety of consequences: additional nucleons can be liberated;
39 “thermal” energy can be imparted to the nucleus; pions can be created and absorbed; and pions
40 and nucleons can undergo charge exchange scattering (e.g., $\pi^- p \rightarrow \pi^0 n$). Event generators include
41 phenomenological models for FSI, anchoring to hadron-nucleus scattering data.

1 The heavy nuclei in a detector also act as targets for the particles that have escaped the struck
2 nucleus. Generally speaking, the denser the detector and the more crudely it samples deposited
3 energy, the more difficult it is to observe low-energy particles. Negatively and positively charged
4 pions leave different signatures in a detector since the former are readily absorbed while the latter
5 are likely to decay. Neutrons can be produced from the struck nucleus, but also from follow on
6 interactions of the neutrino's reaction products with other nuclei. The energy carried away by
7 neutrons is challenging to detect and can bias the reconstructed neutrino energy.

8 Finally, it is important to note that a significant fraction of the neutrino interactions in DUNE will
9 come from deep inelastic scattering rather than the simpler QE scattering discussed above. This
10 leads typically to a more complex morphology for events (beyond the heavy nucleus complications)
11 and greater challenges for the detector and the modeling.

12 1.2.1 Lessons from Current Experiments

13 Neutrino beams are notoriously difficult to model at the precision and accuracy required for modern
14 accelerator-based experiments. Recent long-baseline experiments make use of a ND placed close to
15 the beam source, where oscillations are not yet a significant effect. The beam model, the neutrino
16 interaction model, and perhaps the detector response model are tuned, or calibrated, by the data
17 recorded in the ND. The tuned model is used in the extraction of the oscillation signal at the FD.
18 Known effects that are not understood or modeled well must be propagated into the final results
19 as part of the systematic error budget. Unknown effects that manifest as disagreements between
20 the model and observations in the ND also must be propagated into the final results as part of
21 the systematic error budget. These kinds of disagreements have happened historically to every
22 precision accelerator oscillation experiment. When such disagreements arise, some assumption or
23 range of assumptions must be made about the source of the disagreement. Without narrowing
24 down the range of possibilities, this can become a leading systematic error.

25 Since the final results depend on the comparison of what is seen in the FD to that in the ND,
26 having functionally identical detectors (i.e., the same target nucleus and similar detector response)
27 is helpful. In a similar vein, differences between the neutrino spectrum at the ND and the oscillated
28 spectrum seen at the FD lead to increased sensitivity to systematic effects propagated from the
29 ND to the FD.

30 The past experience of the neutrino community is a driving force in the design of the DUNE
31 ND complex. The performance of current, state-of-the-art long baseline oscillation experiments
32 provides a practical guide to many of the errors and potential limitations DUNE can expect to
33 encounter, as well as case studies of issues that arose which were unanticipated at the design stage.

34 The T2K experiment uses an off-axis neutrino beam that has a narrow energy distribution peaked
35 below 1 GeV. This means, relative to DUNE, interactions in T2K are predominantly CCQE and
36 have relatively simple morphologies. The data sample has little feed-down from higher energy
37 interactions. The T2K ND (plastic scintillator and 74 TPC) technology is very different from the
38 FD (water Cerenkov), though the ND contains embedded water targets that provide samples of
39 interactions on the same target as used in the FD. The experiment relies on the flux and neutrino

1 interaction models, as well as the ND and FD response models to extrapolate the constraint from
2 the ND to the FD. In the most recent oscillation results released by T2K, the ND data constraint
3 reduces the the flux and interaction model uncertainties at the FD from 11 % to 14 % down to 2.5 %
4 to 4 % [9]. Inclusion of the water target data was responsible for a factor of two reduction in the
5 systematic uncertainties, highlighting the importance of measuring interactions on the same target
6 nucleus as the FD. *Note: these numbers are not used directly in the analysis but were extracted to*
7 *provide an indication of the power of the ND constraint.*

8 The NO ν A experiment uses an off-axis neutrino beam from NuMI that has a narrow energy
9 distribution peaked around 2 GeV. The NO ν A ND is functionally identical to the FD. Still, it is
10 significantly smaller than the FD and it sees a different neutrino spectrum due to geometry and
11 oscillations. Note that even with the functionally identical near and far detectors, NO ν A uses a
12 model to subtract NC background and relies on a model-dependent response matrix to translate
13 what is seen in the ND to the “true” spectrum, which is then extrapolated to the FD where it is
14 put through a model again to predict what is seen in the FD [10, 11]. Within the extrapolation,
15 the functional similarity of the near and far detectors reduces, but does not eliminate, many
16 systematic effects. Uncertainties arising from the neutrino cross section model dominate the NO ν A
17 ν_e appearance systematic error budget and are among the larger errors in the ν_μ disappearance
18 results. The ND constraint is significant. For the ν_e appearance signal sample in the latest NO ν A
19 results, for example, a measure of the systematic error arising from cross section uncertainties
20 without using the ND extrapolation is 12% and this drops to 5% if the ND extrapolation is used
21 [11].

22 The process of implementing the ND constraint in both T2K and NO ν A is less straightforward
23 than the typical description implies. It will not be any more straightforward for DUNE. One issue
24 is that there are unavoidable near and far differences. Even in the case of functionally identical
25 detectors, the beam spectrum and intensity are very different near to far. For DUNE, in particular,
26 ArgonCube is smaller than the FD and is divided into modular, optically isolated regions that have
27 a pixelated readout rather than the wire readout of the FD. Space charge effects will differ near to
28 far. All of this imposes model dependence on the extrapolation from near to far. This is mitigated
29 by collecting data at differing off-axis angles with DUNE-PRISM, where an analysis can be done
30 with an ND flux that is similar to the oscillated FD flux (see section 1.3.4). (It should be noted
31 data from ProtoDUNE will also be useful to understand the energy dependent detector response
32 for the FD.) Regardless, near to far differences will persist and must be accounted for through the
33 beam, detector, and neutrino interaction models.

34 Although long baseline oscillation experiments use the correlation of fluxes at the ND and the FD
35 to reduce sensitivity to flux modeling, the beam model is a critical component in understanding
36 this correlation. Recently, the MINERvA experiment used spectral information in the data to
37 diagnose a discrepancy between the expected and observed neutrino event energy distribution in
38 the NuMI medium energy beam [12]. In investigating this issue, MINERvA compared the observed
39 and simulated neutrino event energy distribution for low- ν events, as shown in Figure 1.2. Since
40 the cross section is known to be relatively flat as a function of neutrino energy for this sample,
41 the observed disagreement as a function of energy indicated a clear problem in the flux model or
42 reconstruction. MINERvA believes the observed discrepancy between the data and simulation is
43 best described by what is a mismodeling in horn focusing combined with an error in the muon
44 reconstruction via energy by range. This is notable, in part, because the two identified culprits

1 in this saga would manifest differently in the extrapolation to the far detector in an oscillation
 2 experiment. The spectral analysis provided critical information in arriving at the final conclusion.
 3 This experience illustrates the importance of good monitoring/measurements of the neutrino beam
 4 spectrum.

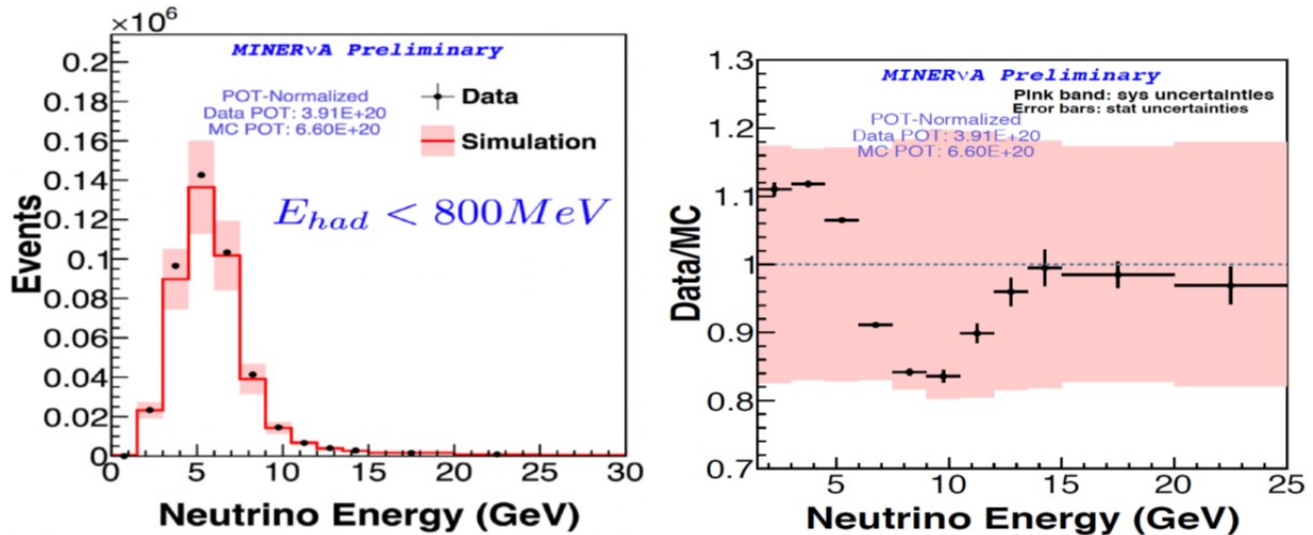


Figure 1.2: Reconstructed MINERvA medium energy NuMI neutrino event spectrum for low energy transfer events compared to simulation (left) and same comparison shown as a ratio (right). From [12].

5 Another important issue is that the neutrino interaction model is not perfect, regardless of the
 6 experiment and implementation. With an underlying model that does not describe the reality of
 7 nature, even a model tuned to ND data will have residual disagreements with that data. These
 8 disagreements must be accounted for in the systematic error budget of the ultimate oscillation
 9 measurements. Although the model(s) may improve before DUNE operation, the degree of that
 10 improvement cannot be predicted and the DUNE ND complex should have the capability to gather
 11 as much information as possible to help improve and tune the model(s) during the lifetime of the
 12 experiment. In other words, the ND needs to be capable of narrowing the range of plausible
 13 possibilities giving rise to data-model differences at the ND in order to limit the systematic error
 14 incurred in the results extracted from the FD.

15 Recent history provides illustrations of progress and continuing struggles to improve neutrino in-
 16 teraction models. The MiniBooNE collaboration published results in 2010 showing a disagreement
 17 between the data and the expected distribution of CCQE events as a function of Q^2 [13, 14]. They
 18 brought the model into agreement with the data by increasing the axial mass form factor used in
 19 the model. K2K [14] and MINOS [15] made similar measurements. It has since been shown that
 20 the observed disagreement is due to the need to include multinucleon processes and that the use
 21 of the large effective axial mass form factor used by these experiments to fit the data leads to a
 22 misreconstruction of the neutrino energy.

23 The importance of modeling multinucleon (2p2h) processes for oscillation experiments is under-
 24 scored by the fact that such interactions when reconstructed as a CCQE (1p1h) process lead to
 25 a significant low-side tail in the reconstructed neutrino energy [16]. Multinucleon processes also
 26 change the hadronic calorimetric response. The first NO ν A ν_μ disappearance oscillation results
 27 had a dominant systematic error driven by the disagreement of their model to the data in their

1 hadronic energy distribution [17]. In more recent work, the inclusion of multinucleon processes in
2 the interaction model contributed to a substantial reduction of this disagreement [10].

3 The MINER ν A experiment has compiled a significant catalog of neutrino and antineutrino results
4 and recently developed a model tune to their quasielastic-like (NuMI low energy) data [8]. The
5 tune is based on a modern neutrino interaction generator (GENIE 2.8.4 [18], using a global Fermi
6 gas model [6] with a Bodek-Ritchie tail [19] and the INTRANUKE-hA FSI model [20]). Even so,
7 MINER ν A scales down non-resonance pion production [21], includes a random phase approxima-
8 tion model (RPA) [22, 23], and incorporates a multinucleon model [24, 25, 26] with an empirical
9 enhancement in the dip region between the quasielastic and delta region that is determined by
10 a fit to the neutrino data [8]. The same tune as developed on the neutrino data also fits well
11 the MINER ν A anti-neutrino quasielastic-like data (with no additional tuning or ingredient). The
12 required enhancement of the multinucleon contribution to the model implies shortcomings in the
13 interaction model, but the decent fit to data for both neutrinos and anti-neutrinos implies that
14 the tune is effectively making up for some imperfections in the model.

15 More recent versions of GENIE include some of the modifications incorporated by MINER ν A in
16 the tune discussed above [27]. This illustrates the dynamic nature of neutrino interaction modeling
17 and the interplay between the experiments and generator developers. The evolution of the field
18 continues as illustrated with a snapshot of some of the current questions and areas of focus:

- 19 • There is a pronounced deficit of pions produced at low Q^2 in $CC1\pi^0$ events as compared to
20 expectations [28, 29, 30, 31, 32]. Current models take this into account by tuning to data
21 without any underlying physical explanation for how or why this happens.
- 22 • The MINER ν A tune that fits both neutrino and anti-neutrino CCQE data involves a sig-
23 nificant enhancement and distortion of the 2p2h contribution to the cross section. The real
24 physical origin of this cross section strength is unknown. Models of multinucleon processes
25 disagree significantly in predicted rates.
- 26 • Multinucleon processes likely contribute to resonance production. This is neither modeled
27 or well constrained.
- 28 • Cross section measurements used for comparison to models are a convolution of what the
29 models view as initial state, hard scattering, and final state physics. Measurements able to
30 deconvolve these contributions are expected to be very useful for model refinements.
- 31 • Most neutrino generators make assumptions about the form of form factors and factorize
32 nuclear effects in neutrino interactions into initial and final state effects via the impulse ap-
33 proximation. These are likely oversimplifications. The models will evolve and the systematic
34 errors will need to be evaluated in light of that evolution.
- 35 • Neutrino detectors are largely blind to neutrons and low momentum protons and pions
36 (though some π^+ are visible via Michel decay). This leads to smearing in the the recon-
37 structed energy and tranverse momentum, as well as a reduced ability to accurately identify
38 specific interaction morphologies. The closure of these holes in the reconstructed particle
39 phase space is expected to provide improved handles for model refinement.

- There may be small but significant differences in the ν_μ and ν_e CCQE cross sections which are poorly constrained [?].

Given the critical importance of neutrino interaction models and the likelihood that the process of refining these models will continue through the lifetime of DUNE, it is important the DUNE ND suite be highly capable.

1.2.2 Lessons from Past Experience

The philosophy driving the DUNE ND concept is to provide sufficient redundancy to address areas of known weaknesses in previous experiments and known issues in the interaction modeling insofar as possible, while providing a powerful suite of measurements that is likely to be sensitive to unanticipated issues and useful for continued model improvements. Anything less reduces DUNE's potential to achieve significantly improved systematic errors over previous experiments in the long-baseline analyses.

The DUNE ND incorporates many elements in response to lessons learned from previous experiments. The massive ND LArTPC has the same target nucleus and a similar technology to the FD. These characteristics reduce the detector and target systematic sensitivity in the extrapolation of flux constraints from this detector to the FD. This detector is capable of providing the primary sample of charged-current ν_μ interactions to constrain the flux at the FD, along with other important measurements of the flux from processes like ν -e⁻ scattering and low- ν . Samples taken with this detector at off-axis angles (DUNE-PRISM) will allow the deconvolution of the flux and cross section errors and provide potential sensitivity to mismodeling. The off-axis data can, in addition, be used to map out the detector response function and construct effective ND samples that mimic the energy distribution of the oscillated sample at the FD.

The DUNE ND provides access to particles produced in neutrino interactions that have been largely invisible in previous experiments, such as low-momentum protons and charged pions measured in the HPgTPC and neutrons in the 3DST and ECAL. The HPgTPC provides data on interactions that minimize the effect of secondary interactions on the produced particles. These capabilities improve the experiments ability to identify specific interaction morphologies, study samples with improved energy resolution, and extract samples potentially useful for improved tuning of model(s) of multinucleon processes. The neutron content in neutrino and antineutrino interactions is different and this will lead to differences in the detector response. For an experiment that is measuring charge-parity symmetry violation (CPV), data on neutron production in (anti)neutrino interactions is likely to be an important handle in the tuning of the interaction model and the flavor-dependent detector response function model.

The 3DST-S provides dedicated beam spectrum monitoring on axis, as well as high statistics samples (both ν_μ CC and ν -e⁻) useful for the on-axis flux determination as a crosscheck on the primary flux which has different detector and target systematics. The utility of these samples should be viewed in light of the MINERvA experience in determining the NuMI flux, the success of T2K in controlling the carbon to oxygen target difference in their oscillation analyses (recent analyses also incorporate use of data on oxygen in the ND), the fact that the nuclear densities are

1 similar in carbon and argon, recent success in describing electron scattering data on C and Ar with
2 a superscaling model [4], and the differences between the near and far LAr detectors. Relative to
3 T2K, which makes use of a rate-only beam monitor on-axis and measures the spectrum off-axis,
4 in the DUNE ND the off-axis element moves as a function of time. With this additional degree of
5 freedom complicating the modeling, the dedicated on-axis beam spectrum measurement is likely
6 to be important.

7 The large data sets that will be accumulated by the three main detectors in the ND suite will
8 allow for differential studies and the use of transverse kinematic imbalance variables, where each
9 detector brings its unique strengths to the study: the LArTPC has good tracking resolution
10 and containment and massive statistics; the HPgTPC has excellent tracking resolution, very low
11 charged particle tracking thresholds, and unambiguous track charge sign determination; and the
12 3DST has good containment and can include neutrons on an event-by-event basis. The neutrino
13 interaction samples acquired by this array of detectors will constitute a powerful laboratory for
14 deconvoluting the initial state, hard scattering, and final state physics, which, in turn, will lead to
15 improved modeling and confidence in the final results extracted from the FD.

16 **1.3 Movable components of the ND and the DUNE-PRISM** 17 **program**

18 **1.3.1 Introduction to DUNE-PRISM**

19 One of the primary challenges for DUNE will be controlling systematic uncertainties from the
20 modeling of neutrino-argon interactions. The relationship between the observable final state parti-
21 cles from a neutrino interaction and the incident neutrino energy is currently not understood with
22 sufficient precision to achieve DUNE physics goals. This is due in part to mismodeling the outgoing
23 particle composition and kinematics and due to missing energy from undetected particles, such a
24 neutrons and low energy charged pions, and misidentified particles. The latter effects tend to cause
25 a “feed-down” in reconstructed neutrino energy relative to the true energy. Since neutrino energy
26 spectra at the FD and ND have substantially different features due to the presence of oscillations
27 at the FD, these mismodeling and neutrino energy feed-down effects do not cancel in a far/near
28 ratio as a function of neutrino energy, and lead to biases in the measured oscillation parameters.

29 Understanding ND constraints on neutrino-nucleus interaction uncertainties is challenging, since
30 no complete model of neutrino-argon interactions is available. If it were possible to construct a
31 model that was known to be correct, even with a large number of unknown parameters, then the
32 task of a ND would much simpler: to build a detector that can constrain the unknown parameters
33 of the model. However, in the absence of such a model, this procedure will be subject to unknown
34 biases due to the interaction model itself, which are difficult to quantify or constrain.

35 The DUNE-PRISM ND program consists of a mobile ND that can perform measurements over
36 a range of angles off-axis from the neutrino beam direction in order to sample many different
37 neutrino energy distributions, as shown in Figure 1.3. By measuring the neutrino-interaction final

- 1 state observables over these continuously varying incident neutrino energy spectra, it is possible to
 2 experimentally determine the relationship between neutrino energy and what is measured in the
 3 detector (i.e., some observable such as reconstructed energy).

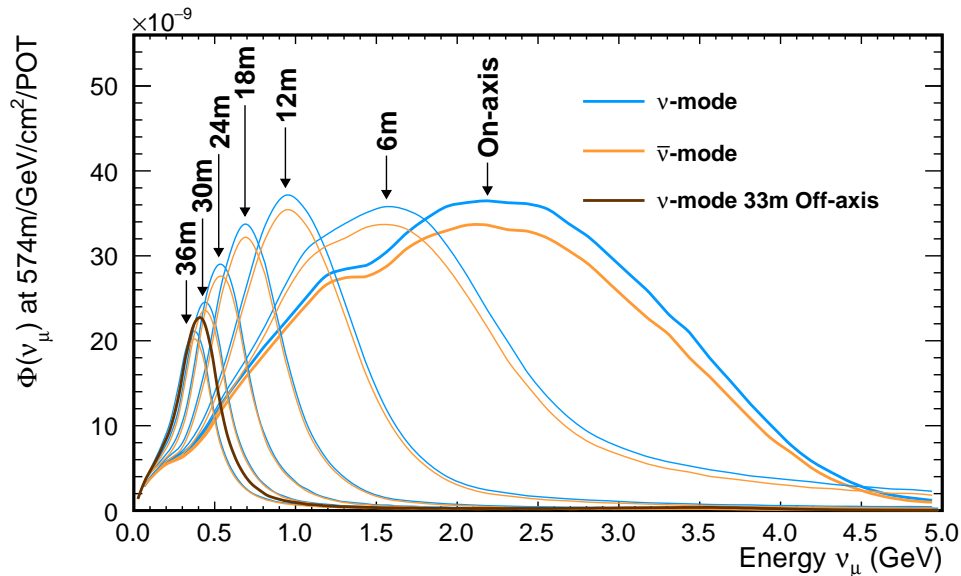


Figure 1.3: The variation in the neutrino energy spectrum is shown as a function of detector off-axis position, assuming the nominal ND location 574 m downstream from the production target.

- 4 In the DUNE ND, the movable components of the detector that are used in the DUNE-PRISM
 5 program are ArgonCube and the MPD. These components of the ND will take data both on the
 6 beam axis and off-axis. In the following sections, ArgonCube and the MPD will be described in
 7 some detail and then the DUNE-PRISM program will be described in more detail.

8 1.3.2 LArTPC Component in the DUNE ND: ArgonCube

- 9 As the DUNE FDs have LAr targets, there needs to be a major LAr component in the DUNE ND
 10 complex in order to reduce cross section and detector systematic uncertainties for oscillation anal-
 11 yses [33, 34]. However, the intense neutrino flux and high event rate at the ND makes traditional,
 12 monolithic, projective wire readout TPCs unsuitable. This has motivated a program of R&D
 13 into a new LArTPC design, suitable for such a high-rate environment, known as ArgonCube [1].
 14 ArgonCube utilizes detector modularization to improve drift field stability, reducing high voltage
 15 (HV) and the LAr purity requirements; pixelized charge readout [35, 36], which provides unam-
 16 biguous 3D imaging of particle interactions drastically simplifying the reconstruction; and new
 17 dielectric light detection techniques with ArCLight [37], which can be placed inside the field cage
 18 (FC) to increase light yield, and improve the localization of light signals. Additionally, ArgonCube
 19 uses a resistive field shell, instead of traditional field shaping rings, to maximize the active volume,
 20 and to minimize the power release in the event of a breakdown [?].

- 21 The program of ArgonCube R&D has been very successful to date, working on small component
 22 prototypes and is summarized in references [38, 39, 40, 35, 41, 36, ?, 37]. With the various techno-

1 logical developments demonstrated with small-scale TPCs, the next step in the ArgonCube pro-
2 gram is to demonstrate the scalability of the pixelized charge readout and light detection systems,
3 and to show that information from separate modules can be combined to produce high quality
4 event reconstruction for particle interactions. To that end, a mid-scale ($1.4\text{ m} \times 1.4\text{ m} \times 1.2\text{ m}$)
5 modular TPC, dubbed the ArgonCube 2×2 Demonstrator, with four independent LArTPC mod-
6 ules arranged in a 2×2 grid has been designed, and is currently under construction.

7 After a period of testing at the University of Bern, the ArgonCube 2×2 Demonstrator will be
8 placed in the MINOS ND hall at Fermilab where it will form the core of a prototype DUNE
9 ND, ProtoDUNE-ND [42]. As part of ProtoDUNE ND, the ArgonCube concept can be studied
10 and operated in an intense, few-GeV neutrino beam. This program aims to demonstrate stable
11 operation and the ability to handle backgrounds, relate energy associated with a single event
12 across ArgonCube modules, and connect tracks to detector elements outside of ArgonCube. The
13 ArgonCube 2×2 Demonstrator is described below in some detail since the DUNE ND modules are
14 anticipated to be very similar.

15 **1.3.2.1 ArgonCube in ProtoDUNE-ND**

16 The ArgonCube concept is a detector made of self-contained TPC modules sharing a common
17 cryostat. Each module is made of a rectangular box with a square footprint and a height optimized
18 to meet the physics goals and/or sensitivity constraints. The ArgonCube 2×2 Demonstrator
19 module will be housed within an existing LN_2 -cooled and vacuum-insulated cryostat, which is
20 $\sim 2.2\text{ m}$ in diameter and $\sim 2.8\text{ m}$ deep, for a total volume of $\sim 6\text{ m}^3$. The size of the cryostat sets
21 the dimensions of the modules for the demonstrator. The square base of each module will be
22 $0.67\text{ m} \times 0.67\text{ m}$, and the height will be 1.81 m . This makes the modules comparable in size to,
23 but slightly smaller than, the proposed ArgonCube DUNE ND modules, which will have a base of
24 $1\text{ m} \times 1\text{ m}$ and a 3.5 m height.

25 Individual modules can be extracted or reinserted into a common LAr bath as needed, as is
26 illustrated in Figure 1.4. This feature will be demonstrated during a commissioning run at the
27 University of Bern, but is not intended to be part of the detector engineering studies in the
28 MINOS-ND hall. The pressure inside the modules is kept close to the bath pressure, putting
29 almost no hydrostatic force on the module walls. This allows the walls to be thin, minimizing the
30 quantity of inactive material in the walls. The purity of the LAr is maintained within the modules,
31 independent of the bath, as will be described below. The argon surrounding the modules need not
32 meet as stringent purity requirements as the argon inside. Under normal operating conditions, all
33 modules are inserted with clearance distances of only 1.5 mm between modules. Cooling power to
34 the bath is supplied by liquid nitrogen circulated through lines on the outer surface of the inner
35 cryostat vessel.

36 A cutaway drawing of an individual 2×2 module is shown in Figure 1.5. The side walls of each
37 module are made from 1 cm G10 sheets, to which the resistive field shell is laminated. G10's elec-
38 tromagnetic radiation length ($X_0 = 19.4\text{ cm}$) and hadronic interaction length ($\lambda_{\text{int}} = 53.1\text{ cm}$) [43]
39 are both comparable to LAr (14.0 cm and 83.7 cm respectively). G10 provides a strong dielectric,
40 capable of 200 kV cm^{-1} when 1 cm thick [44]. This dielectric shielding eliminates the need for a

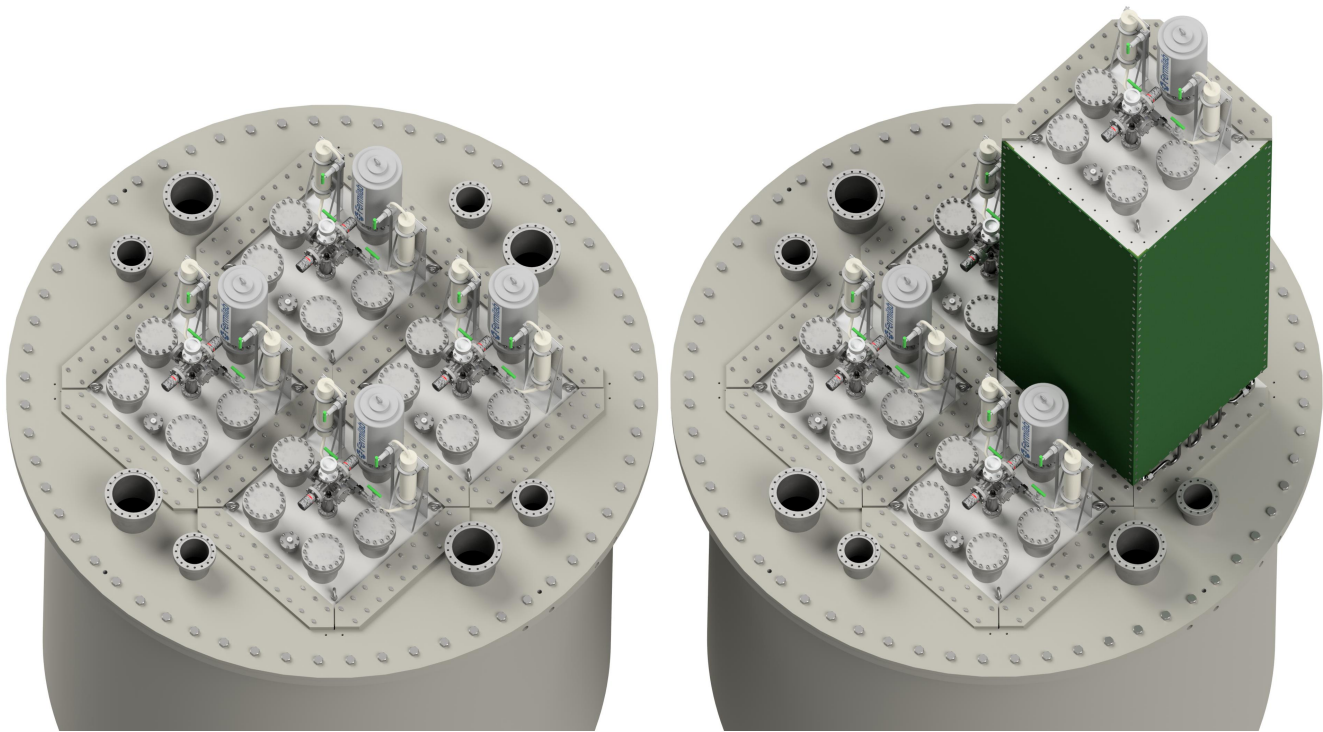


Figure 1.4: Illustration of the ArgonCube 2×2 Demonstrator module. The four modules are visible, with one of them partly extracted, on the right. This figure has been reproduced from Ref. [1].

1 clearance volume between the TPCs and the cryostat, while also shielding the TPC from field
2 breakdowns in a neighboring module.

3 Each module is split into two TPCs by a central cathode made of an additional resistive layer on a
4 G10 substrate. The segmented drift length does not require a high cathode voltage, and minimizes
5 stored energy. For the 2×2 module footprint of $0.67\text{ m} \times 0.67\text{ m}$, and an E field of 1 kV cm^{-1} ,
6 a cathode potential of only 33 kV is required. Operating a LArTPC at this voltage is feasible
7 without a prohibitive loss of active volume [39]. The high field is helpful for reducing drift time
8 and the potential for pileup, minimizing the slow component of the scintillation light, reducing
9 space charge effects, and providing robustness against loss of liquid argon purity.

10 The detector is oriented such that the cathodes are parallel to the beam. This minimizes the load
11 on the readout electronics by spreading the event over more channels and reducing the required
12 digitization rate for hit channels. In turn, this reduces the heat load generated at the charge
13 readout and prevents localized boiling.

14 During filling and emptying of the cryostat, the argon flow is controlled by hydrostatic check
15 valves located at the lower flange of the module, which require a minimal differential pressure
16 of 15 mbar to open. The purity inside each module is maintained by means of continuous LAr
17 recirculation through oxygen traps. Dirty argon is extracted from the base of the module, and
18 is then pushed through oxygen traps outside the cryostat, clean argon then re-enters the module
19 above the active volume. For optimal heat transport, the argon flow is directed along the cold

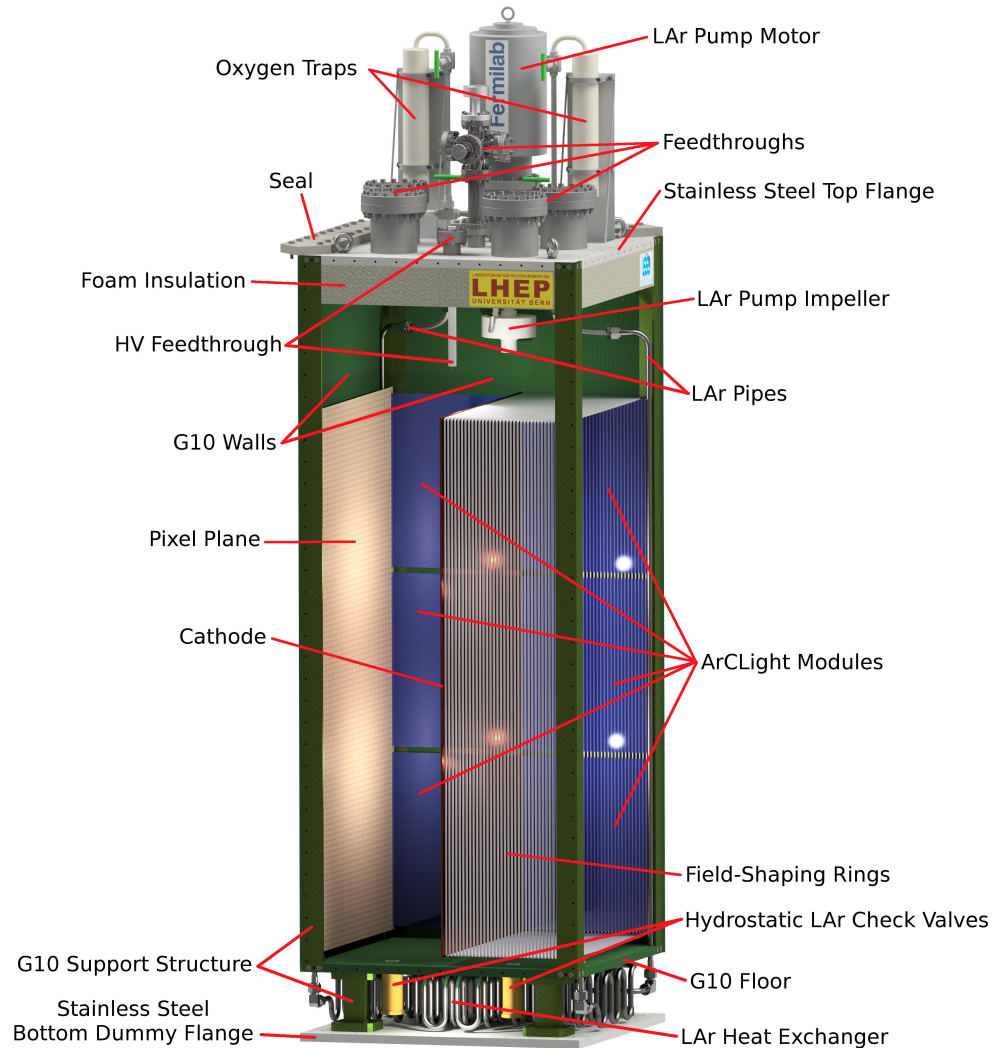
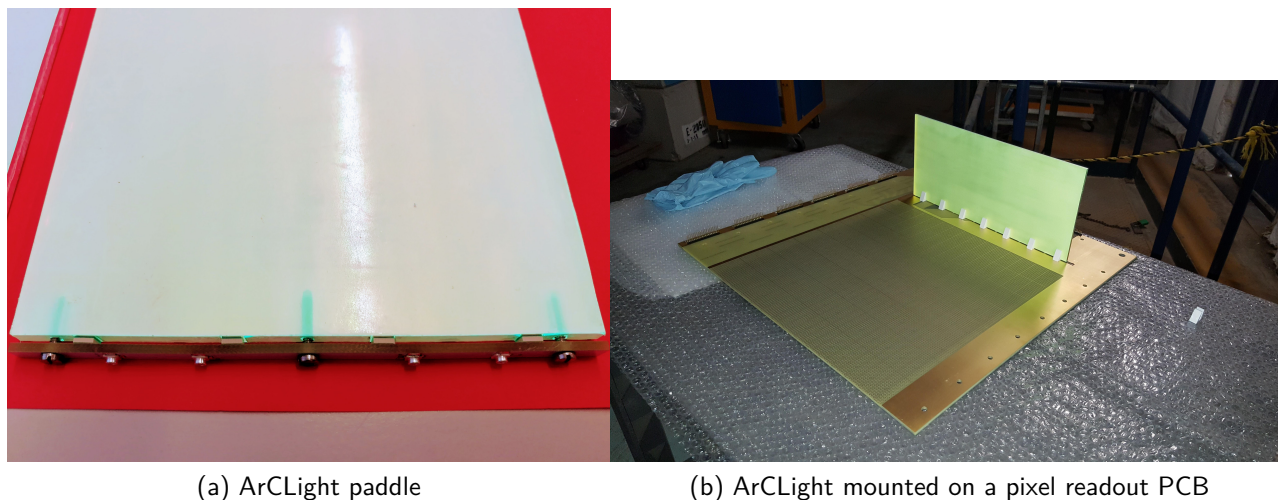


Figure 1.5: Cutaway drawing of a $0.67\text{ m} \times 0.67\text{ m} \times 1.81\text{ m}$ ArgonCube module for the 2×2 Demonstrator module. For illustrative purposes the drawing shows traditional field shaping rings instead of a resistive field shell. Note, G10 walls will completely seal the module, isolating it from the neighbouring modules and the outer LAr bath. It is also worth noting that the 2×2 modules will not have individual pumps and filters.

1 electronics. To prevent dirty argon from the bath entering the modules, their interior is held at
 2 a slight overpressure. For the 2×2 , the dirty argon from all four modules is extracted by a single
 3 pump at the base of the cryostat with a four-to-one line, and after being filtered and cooled, the
 4 clean argon is pumped back in the module via a one-to-four line. A more extensive version of the
 5 same scheme is envisaged for the DUNE ND.

6 ArgonCube offers true 3D tracking information using the LArPix cryogenic ASIC [36] pixelated
 7 charge readout. LArPix ASICs amplify and digitize the charge collected at single-pixels in the
 8 cold to mitigate the need for analogue signal multiplexing, and thus produce unambiguous 3D
 9 information. Sixty-four pixels can be connected to a single LArPix ASIC. The baseline design is
 10 for the 2×2 is a 4 mm pixel pitch, corresponding to $62.5\text{k pixels m}^{-2}$. Pixelated anode planes are
 11 located on the two module walls parallel to the cathode; each plane is $1.28\text{ m}^2 \times 0.64\text{ m}^2$. The total
 12 area across all four modules is 6.6 m^2 , which corresponds to 410k pixels. The readout electronics
 13 utilize two FPGA boards per module, connected to a single Ethernet switch. It should be noted
 14 that the pixel pitch may be reduced as prototypes develop, but this can be accommodated in the
 15 readout design.



(a) ArCLight paddle

(b) ArCLight mounted on a pixel readout PCB

Figure 1.6: (a) A prototype ArgonCube light readout paddle. The paddle is 50 cm long and 10 cm wide, with four SiPMs coupled to one end. Reproduced from Ref. [1]. (b) ArCLight paddle mounted on the PixLAR pixelated charge readout plane, as used in test beam studies at Fermilab.

16 The charge readout window (drift time) of $137\ \mu\text{s}$ is long compared to the $10\ \mu\text{s}$ [45] beam spill
 17 length in the NuMI and LBNF beams. For a 1 MW beam intensity, the expected rate of neutrino
 18 interactions at the DUNE ND is roughly 0.5 per spill per ArgonCube module. With LArPix,
 19 reconstruction issues are greatly simplified compared to a projective readout TPC. Tracks and
 20 connected energy deposits will frequently overlap in any 2D projection, but can be resolved easily
 21 with the full 3D readout. However, disconnected energy deposits, such as those from photon
 22 conversions or neutron interactions in the detector, cannot be associated easily to a specific neutrino
 23 interaction. This problem can be solved by incorporating fast timing information from the prompt
 24 scintillation light emitted in LAr. The module's opaque cathode and walls contain scintillation
 25 light within each TPC (half module), improving the detection efficiency of the prompt component
 26 of the scintillation light. Furthermore, attenuation due to Rayleigh scattering, characterized by an
 27 attenuation length of 0.66 m in LAr [46], is mitigated by the maximum photon propagation length
 28 of 0.3 m. It is desirable to have a large area photon detection system to maximize the utility of

1 scintillation light signals in the detector. To minimize any dead material within the active volume,
2 it is also desirable that the light detection be as compact as possible. The solution pursued for
3 the ArgonCube effort is ArCLight [37], which is a very compact dielectric light trap that allows
4 for light collection from a large area, inside high electric fields. An example ArCLight sheet is
5 shown in Figure 1.6. These sheets are mounted on the walls of the module, inside the field shell,
6 aligned with the drift direction, between the anode and the cathode. The additional 5 mm deep
7 dead volume is similar to the one caused by the charge readout in the perpendicular direction.

8 **1.3.2.2 Dimensions of the ArgonCube Component of the DUNE ND**

9 Since it is unrealistic to build a 25 m long LArTPC in order to contain a 5 GeV muon, the LArTPC
10 dimensions have instead been optimized for hadronic shower containment [47], relying on a down-
11 stream spectrometer to analyze crossing muons. Hadronic showers are defined as contained if a
12 reasonable efficiency across a wide range of kinematics is maintained, and there is no phase space
13 with zero acceptance. The specific metric used is that >95% of hadronic energy has to be contained
14 for interactions in the fiducial volume, excluding neutrons and their descendants.

15 To assess the efficiency, detector volumes of varying sizes were simulated in a neutrino beam.
16 This provides a good measure of the efficiency of a given volume to contain different events, but
17 it is not necessarily a good quantity to assess the required detector size. Many events are not
18 contained because of their specific location and/or orientation. Cross section coverage remedies
19 this deficiency by looking at the actual extent of the event, instead of its containment, at a random
20 position inside a realistic detector volume. However, events extending through the full detector
21 will very likely never be contained in a real detector due to the low probability of it happening in
22 exactly the right location (e.g., at the upstream edge of the detector). Therefore, the maximum
23 event size needs to be smaller than the full detector size. For the ND simulation this buffer was
24 chosen to be 0.5 m in all directions. In this way, this measure of cross section coverage allows us to
25 look for phase-space regions which are inaccessible to particular detector volume configurations.

26 To find the optimal detector size in each dimension, two are held constant at their nominal values,
27 while the third dimension is varied and the cross section coverage is plotted as a function of
28 neutrino energy. This is shown for the dimension along the beam direction in Figure 1.7. In this
29 case, Figure 1.7 shows us that 4.5 m would be sufficient, but to avoid model dependencies, 5 m has
30 been selected. Increasing the length beyond 5 m does little to improve cross section coverage, but
31 reducing to 4 m begins to limit coverage at higher energies. Note that 1 minus the cross section
32 coverage gives the fraction of events that cannot be well reconstructed no matter where their
33 vertex is, or how they are rotated within the fiducial volume. The optimized dimensions found
34 using this technique were 3 m tall, 4 m wide, and 5 m along the beam direction. There is also a
35 need to measure large angle muons that do not go into the HPgTPC. Widening the detector to
36 7 m accomplishes that goal without the added complication of a side muon detector.

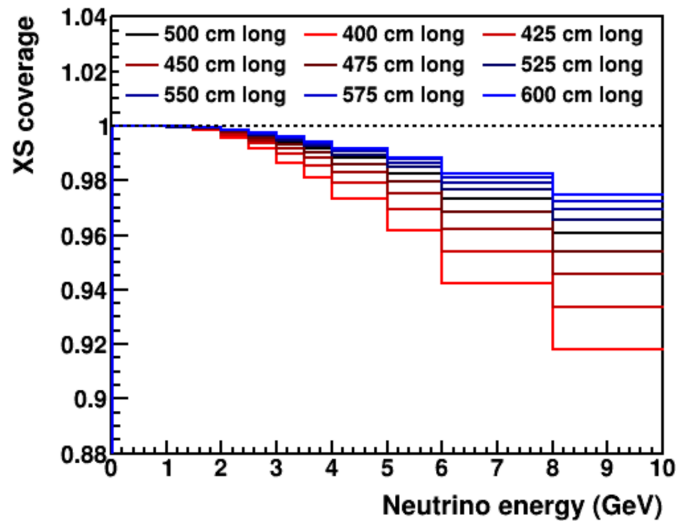


Figure 1.7: Influence of the LArTPC size on hadron containment, expressed in terms of cross section coverage as a function of neutrino energy. Two dimensions are held constant at their nominal values, while the third is varied, in this case the height is held at 2.5 m and the width at 4 m. The optimal length is found to be 5 m. See text for explanation of cross section coverage [47].

1.3.2.3 ArgonCube Module Dimensions

The DUNE ND ArgonCube module dimensions are set to maintain a high drift field, 1 kV cm^{-1} , with minimal bias voltage, and to allow for the detection of prompt scintillation light while mitigating the effects of diffusion on drifting electrons. The prompt scintillation light, $\tau < 6.2 \text{ ns}$ [48], can be efficiently measured with a dielectric light readout with $\mathcal{O}(1) \text{ ns}$ timing resolution, such as ArCLight [37]. To reduce attenuation and smearing due to Rayleigh scattering, the optical path must be kept below the 0.66 m [46] scattering length. Additionally, the slow scintillation component can be further suppressed by operating at higher E fields [49], effectively reducing the ionization density [50] required to produce excited states.

A module with a $1 \text{ m} \times 1 \text{ m}$ footprint split into two TPCs with drift lengths of 50 cm requires only a 50 kV bias. With ArCLight mounted either side of the 1 m wide TPC, the maximal optical path is only 50 cm . For a non-zero drift field, diffusion needs to be split into longitudinal and transverse components. Gushchin [51] report a transverse diffusion of $13 \text{ cm}^2 \text{ s}^{-1}$ at 1 kV cm^{-1} . This results [52] in a transverse spread of 0.8 mm for the drift time of $250 \mu\text{s}$, well below the the proposed pixel pitch of 3 mm . The longitudinal component is smaller than the transverse [52], and is therefore negligible.

1.3.2.4 ND Dimensions

Though the acceptance study discussed in Section 1.3.2.2 indicated a width of 4 m is sufficient to contain the hadronic component of most events of interest, the width has been increased to 7 m in order to mitigate the need for a side-going muon spectrometer. Figure 1.8 shows the overall dimensions of the planned ArgonCube deployment in the DUNE ND. With an active volume of

- 1 $1\text{ m} \times 1\text{ m} \times 3\text{ m}$ per module, the full ArgonCube detector corresponds to seven modules transverse
 2 to the beam direction, and five modules along it. It should be noted that the cryostat design is
 3 currently based on ProtoDUNE [53], and will be optimized for the ND pending a full engineering
 4 study.

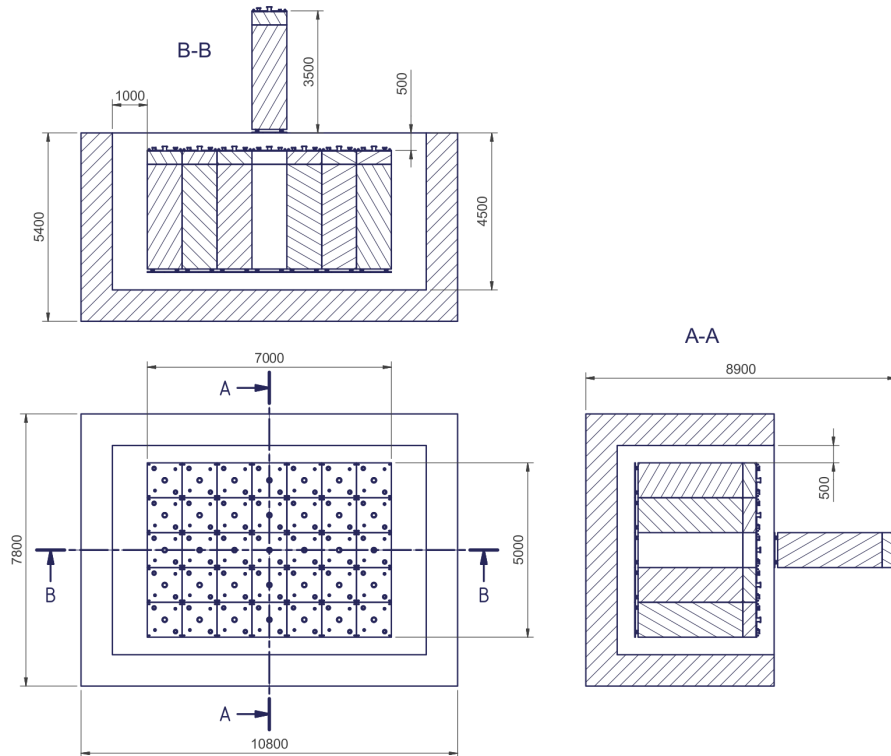


Figure 1.8: The current ArgonCube Dimensions for the DUNE ND. The cryostat is based on ProtoDUNE [53], and yet to be optimized for the DUNE ND.

5 1.3.2.4.1 Statistics in Fiducial Volume

- 6 Figure 1.9 shows 37 million total CC ν_μ neutrino events per year within a 25 t fiducial volume
 7 in FHC mode at 1.07 MW. Figure 1.10 shows only the event rate for events where the visible
 8 hadronic system is fully contained, for the same fiducial volume and beam configuration. Note
 9 that for the visible hadronic system to be contained, all energy not associated with the outgoing
 10 lepton, or outgoing neutrons, was required to be contained.

- 11 For hadronic containment, there is a 30 cm veto region upstream and on all sides of the active
 12 volume, and 50 cm veto region downstream. The fiducial volume is then defined as 50 cm from
 13 all edges, with 150 cm downstream. Within the 25 t fiducial volume in FHC mode at 1.07 MW
 14 the number of fully reconstructed (contained or matched muon, discussed below, plus contained
 15 hadrons) CC ν_μ events per year is 14 million.

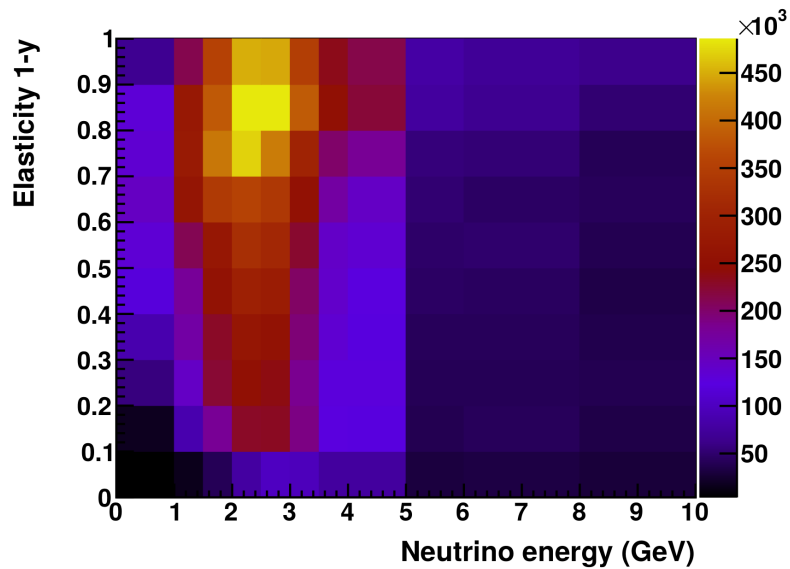


Figure 1.9: All neutrino events in the nominal 25 t fiducial volume, in FHC at 1.07 MW, per year, rates are per bin.

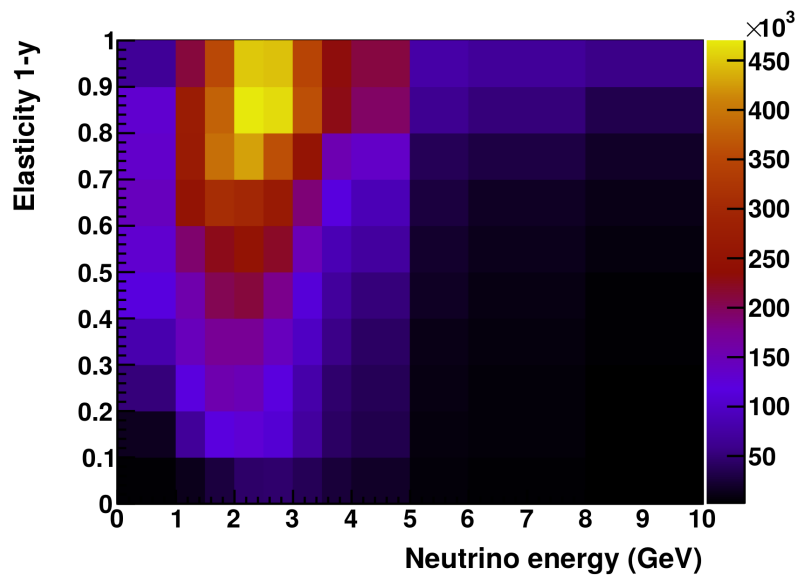


Figure 1.10: Events where the visible hadronic system is contained within the nominal 25 t fiducial volume, in FHC at 1.07 MW, per year, rates are per bin.

1.3.2.4.2 Muon Acceptance

Muons are considered as useful for physics if they stop in the active region of ArgonCube or if they leave the LAr detector and are reconstructed in a magnetic spectrometer downstream. Under the assumption that the downstream magnetic spectrometer is the multipurpose detector described in section 1.3.3, Figure 1.11 shows the muon acceptance as a function of true neutrino energy (on the left) and muon energy (on the right). The acceptance dip at 1 GeV in muon energy is from muons that exit ArgonCube and are not reconstructed in the MPD downstream.

ICARUS and MicroBooNE have used multiple Coulomb scattering to determine muon momentum [54]. This technique may prove to be useful for muons in ArgonCube and could mitigate somewhat the size of the dip in Figure 1.11.

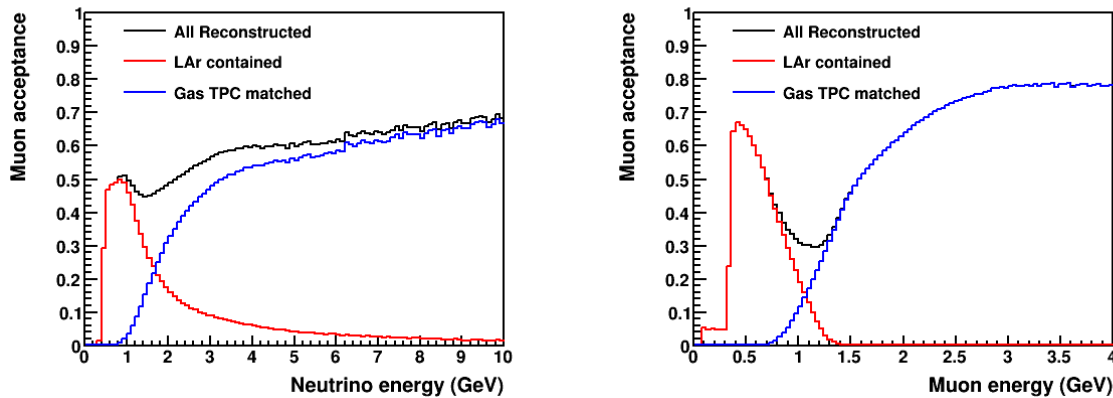


Figure 1.11: Muon acceptance shown as a function of true neutrino energy (left) and true muon energy (right). The acceptance for muons that stop in ArgonCube is shown in red and that for muons reconstructed in the downstream magnetic spectrometer is shown in blue.

1.3.2.5 Muon and Electron Momentum Resolution and Scale Error

For muons stopping in the LAr and for those with momentum measured in the downstream tracker (MPD), the energy scale uncertainty from ArgonCube is driven by the material model of the LAr and passive materials. This is expected to be known to $<1\%$.

For electrons, the energy will be measured calorimetrically, rather than by range. The minimum ionizing particle (MIP) energy scale (charge/MeV) will be set by rock muons. The scaling to more dense deposits from EM showers can give rise to uncertainties, i.e., recombination could be different. Such uncertainties can be reduced by taking data with ArgonCube modules in a test beam. Outside of this, a useful calibration sample of electrons up to 50 MeV comes from Michel electrons from stopping rock muons. The π^0 invariant mass peak is another good standard candle.

1.3.2.6 Tagging Fast Neutrons

Studies have shown that contained prompt scintillation light provides an important handle for neutron tagging, allowing for the association of detached energy deposits to the correct neutrino interaction using timing information. Such neutron tagging is important for minimizing the uncertainty on neutrino energy reconstruction, both for neutrons generated at a neutrino vertex and for hadronic showers that fluctuate to neutrons.

Figure 1.12 shows a simulated beam spill in the $5\text{ m} \times 4\text{ m} \times 3\text{ m}$ LAr component of the DUNE ND². It highlights the problem of associating fast-neutron induced energy deposits to a neutrino vertex using only collected charge.

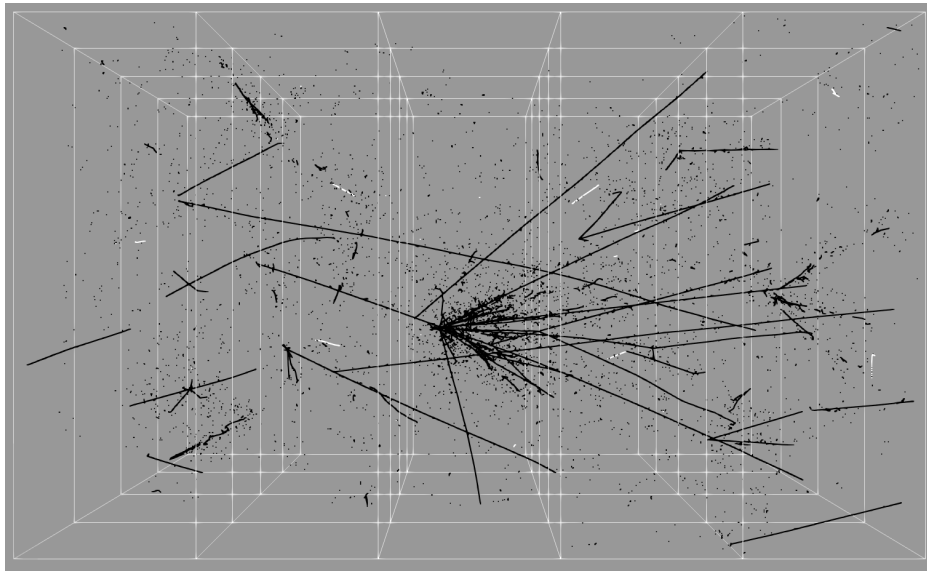


Figure 1.12: A beam spill in the LAr component of the DUNE ND. The detector volume is $5\text{ m} \times 4\text{ m} \times 3\text{ m}$. Fast-neutron induced recoiling proton tracks, with an energy threshold greater than $\sim 10\text{ MeV}$, are shown in white. The black tracks are all other energy deposits sufficient to cause charge collected at the pixel planes.

By containing scintillation light, prompt light signals can be used to associate fast-neutron induced deposits back to a neutrino vertex anywhere within the detector. Figure 1.13 shows the temporal distribution of neutrino vertices within a beam spill. The mean separation of neutrino vertices is 279 ns , with all fast-neutron induced energy deposits occurring $<10\text{ ns}$ after each neutrino interaction.

1.3.2.7 Neutrino-Electron Elastic Scattering

Neutrino scattering on atomic shell electrons, $\nu_l(\bar{\nu}_l) + e^- \rightarrow \nu_l(\bar{\nu}_l) + e^-$, is a purely electroweak process with a known cross section as function of neutrino energy, E_ν , in which all (anti)neutrino flavors participate, albeit with different cross sections. This process is not affected by nuclear

²Note that this study was performed before the detector width was increased to 7 m , as described in Section 1.3.2.4.

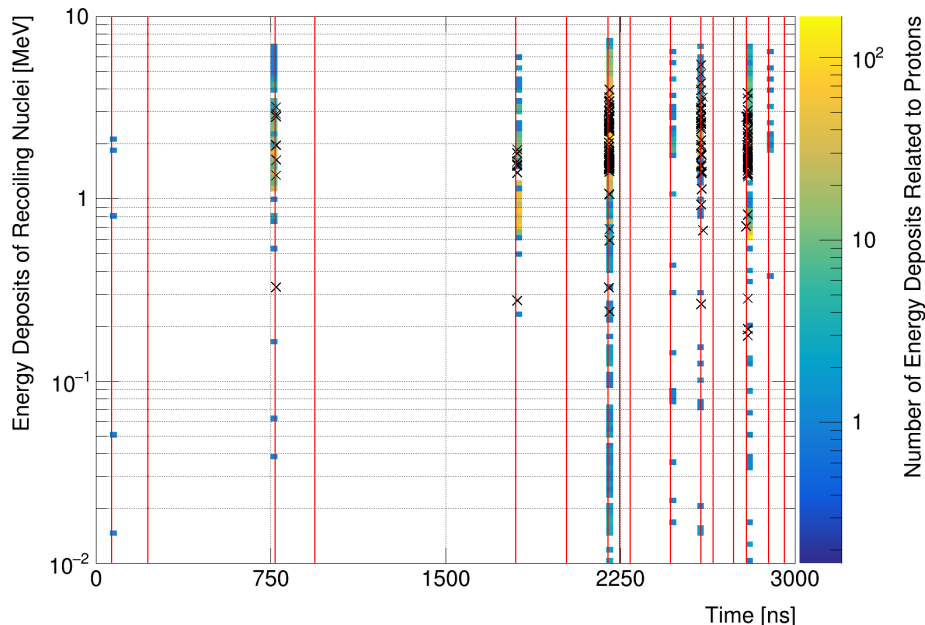


Figure 1.13: The temporal distribution of neutrino vertices (red lines) within a beam spill in the LAR component of DUNE ND. The mean separation of neutrino vertices is 279 ns. The filled bins show the number of hits due to recoiling protons, crosses indicate a hit due to a recoiling ²H, ³H, ²He or ³He nucleus. All fast-neutron induced energy deposits occur <10 ns after each neutrino interaction.

1 interactions and has a clean signal of a single very forward-going electron. MINERνA [55] has
 2 used this technique to characterize the NuMI beam flux normalization (running in the NuMI
 3 low-energy mode), although the rate and detector resolution were insufficient to make a shape
 4 constraint. It has been investigated as a cross section model-independent way to constrain the
 5 neutrino flux at the DUNE ND.

For a neutrino-electron sample, E_ν could, in principle, be reconstructed event-by-event in an ideal detector using the formula

$$E_\nu = \frac{E_e}{1 - \frac{E_e(1 - \cos\theta_e)}{m_e}}, \quad (1.1)$$

6 where m_e and E_e are the electron mass and outgoing energy, and θ_e is the angle between the
 7 outgoing electron and the incoming neutrino direction. The initial energy of the electrons are low
 8 enough to be safely neglected (~ 10 keV). It is clear from Equation 1.1 that the ability to constrain
 9 the shape of the flux is critically dependent on the energy- and, in particular, angular-resolution of
 10 electrons. For a realistic detector, the granularity of the E_ν shape constraint (the binning) depends
 11 on its performance. Additionally, the divergence of the beam (few mrad) at the DUNE ND site is
 12 a limiting factor to how well the incoming neutrino direction is known.

13 In work described in Ref. [56], the ability for various proposed DUNE ND components to con-
 14 strain the DUNE flux is shown using the latest three-horn optimized flux and including full flavor
 15 and correlation information. This was used to determine what is achievable relative to the best
 16 performance expected from hadron production target models. When producing the input flux
 17 covariance matrix, it was assumed that an NA61 [57] style replica-target experiment was already

- 1 used to provide a strong prior shape constraint. Detector reconstruction effects and potential
 2 background processes are included, and a constrained flux-covariance is produced following the
 3 method used in Ref. [55].

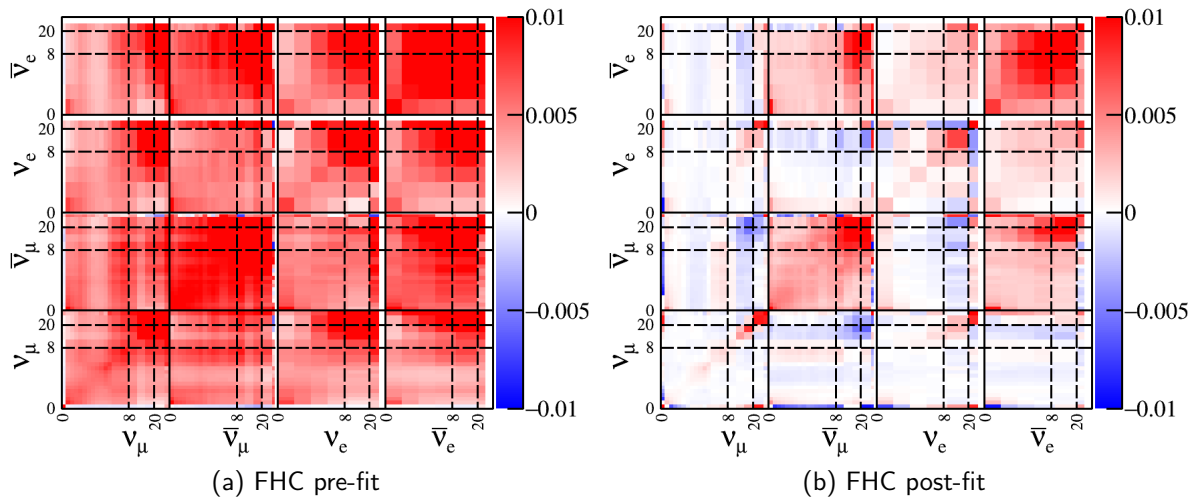


Figure 1.14: Pre- and post-fit FHC flux covariance matrices for the nominal 35 t ArgonCube LAr detector using a five-year exposure.

- 4 The impact of the neutrino-electron scattering constraint on the flux covariance is shown in Fig-
 5 ure 1.14 for FHC and a five year exposure of the nominal 35t ArgonCube LAr detector (corre-
 6 sponding to ~ 22 k neutrino-electron events). It is clear that the overall uncertainty on the flux
 7 has decreased dramatically, although, as expected, an anticorrelated component has been intro-
 8 duced between flavors (as it is not possible to tell what flavor contributed to the signal on an
 9 event-by-event basis). Similar constraints are obtained for RHC running.

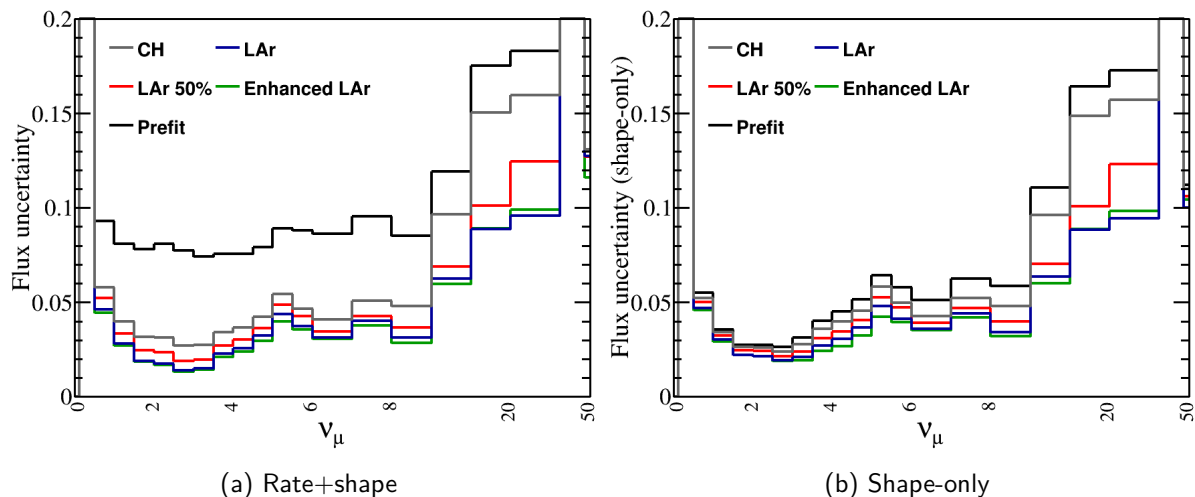


Figure 1.15: Rate+shape and shape-only bin-by-bin flux uncertainties as a function of neutrino energy for a 5 year exposure with various detector options, compared with the input flux covariance matrix before constraint.

- 10 Figure 1.15 shows the flux uncertainty as a function of E_ν for the ν_μ -FHC flux, for a variety
 11 of ND options. In each case, the constraint on the full covariance matrix is calculated (as in

1 Figure 1.14), but only the diagonal of the ν_μ portion is shown. In the flux peak of ~ 2.5 GeV, the
2 total flux uncertainty can be constrained to $\sim 2\%$ for the nominal LAr scenario, and the constraint
3 from other detector types is largely dictated by the detector mass. Clearly the neutrino-electron
4 scattering sample at the DUNE ND will be a powerful flux constraint. However, it is also clear
5 that the ability to constrain the shape of the flux is not a drastic improvement on the existing flux
6 covariance matrix, and none of the possible detectors investigated added a significantly stronger
7 constraint. That said, the neutrino-electron sample is able to make *in situ* measurements of the
8 flux prediction, and is able to diagnose problems with the flux prediction in a unique way.

9 1.3.3 Multi-Purpose Detector

10 The multi-purpose detector (MPD) extends and enhances the capabilities of the LArTPC. It does
11 this by providing a system that will measure the momentum and sign of charged particles exiting
12 the LArTPC and, for neutrino interactions taking place in the MPD, it will extend charged particle
13 measurement capabilities to lower energies than achievable in the far or near LArTPCs. This
14 capability enables further constraints of systematic uncertainties for the long-baseline oscillation
15 analysis. The MPD is a magnetized system consisting of a high-pressure gaseous argon time
16 projection chamber (HPgTPC) and a surrounding ECAL. The detector design will be discussed in
17 more detail in a later section.

18 MPD goals

19 • Measure particles that leave the LAr ND component and enter the MPD

20 The LAr component of the DUNE ND will not fully contain high-energy muons or measure
21 lepton charge. A downstream MPD will be able to determine the charge sign and measure
22 the momenta of the muons that enter its acceptance, using the curvature of the associated
23 track in the magnetic field.

24 • Constrain neutrino-nucleus interaction systematic uncertainties

25 In its 1-ton gaseous argon fiducial volume, the MPD will collect 1.5×10^6 charged-current
26 muon neutrino interactions per year (plus 5×10^5 neutral-current muon neutrino interac-
27 tions). The very low energy threshold for reconstruction of tracks in the HPgTPC gives it
28 a view of interactions that is more detailed than what is seen in the LAr, and on the same
29 target nucleus. The associated ECAL provides excellent ability to detect neutral pions, en-
30 abling the MPD to measure this important component of the total event energy while also
31 tagging the presence of these pions for interaction model studies.

32 The ability to constrain “known unknowns” is a high priority of the MPD. One example is
33 nucleon-nucleon correlation effects and meson exchange currents in neutrino-nucleus scatter-
34 ing. Although a few theoretical models that account for these effects are available in neutrino
35 event generators, no model reproduces well the observed data in NOvA, MINERvA, or T2K.
36 These experiments therefore use empirical models tuned to the limited observables in their

1 detector data. Tuning results in better agreement between simulation and data, although
2 still not perfect. In addition, this type of empirical tuning leaves some large uncertainties,
3 such as the four-momentum transfer response, the neutrino energy dependence of the cross
4 sections (where models disagree, and a “model spread” is typically used for the uncertainty),
5 and the relative fractions of final state nucleon pairs (pp vs. np).

6 Another example of a “known unknown” for which the MPD will provide a more stringent
7 cross section constraint than the LArTPC is the case of single and multiple pion production
8 in charged current neutrino interactions. An MPD-based measurement of these processes will
9 be implemented in the DUNE long-baseline oscillation analysis in the near future, making use
10 of the high-purity samples of CC- 0π , CC- 1π , and CC-multi- π events in the gaseous argon,
11 separated into π^+ and π^- subsamples and binned in neutrino energy and other variables
12 of interest. Figure 1.16 illustrates a two simple differences among the HPgTPC CC- 1π
13 subsamples; it is still to be determined which variables will be the most useful in the long-
14 baseline oscillation analysis.

15 The relative lack of secondary interactions for particles formed in neutrino interactions in the
16 gaseous argon fiducial volume will provide samples with a less model dependent connection
17 to the particles produced in the primary interaction. These secondary interactions are a
18 significant effect in denser detectors [3] and this crosscheck/validation of the reinteraction
19 model is likely to be useful in understanding the full energy response of the liquid detectors.

20 The MPD will measure ratios of inclusive, semi-exclusive, and exclusive cross sections as
21 functions of neutrino energy, where the flux cancels in the ratio. These ratios can be measured
22 separately for NC and CC events, and the NC to CC ratio itself can be measured precisely
23 with the MPD. The ratios of cross sections for different pion, proton, and kaon multiplicity
24 will help constrain interaction models used in the near and far liquid detectors.

25 The HPgTPC will have better capability than the LArTPC to distinguish among particle
26 species at low momentum using dE/dx measurements. Some muon/pion separation is possi-
27 ble via dE/dx for very low momenta, and protons are very easily distinguished from pions,
28 muons, and kaons for momenta below 2 GeV/ c , as shown in Fig. 1.28. At higher momenta,
29 the magnet makes it possible to easily distinguish π^+ from μ^- (or π^- from μ^+), as has been
30 done in T2K near-detector fits for oscillation analyses. The fact that pions will interact
31 hadronically far less often in the gas than in the liquid will give another important handle
32 for constraining uncertainties in the LArTPC. These aspects give the MPD a significant
33 complementarity to the LArTPC, which is not magnetized. Since the target nucleus in the
34 MPD is the same as that in the near and far LArTPCs this information feeds directly into
35 the interaction model constraints without complicating nuclear physics concerns.

36 Finally, having a ND that can see one level deeper than the far detector keeps open the
37 possibility to investigate “unknown unknowns” as well. Since the MPD will identify and
38 measure interactions more accurately than can be done in the LArTPC, it will provide the
39 ability to investigate more deeply our observations in the liquid argon, and the flexibility to
40 address other unexpected things we may encounter.

- 41 • **Precisely and accurately measure all components of the neutrino flux**

The magnetic field of the MPD enables the precise determination of momenta of charged particles escaping the upstream LArTPC. Because the ND is necessarily smaller than the FD, near-far differences arising from the different containment fractions are compensated by the fact that the ND has a magnetic spectrometer. Also, higher-energy particles from the neutrino interaction will be measured better in the MPD than in the liquid ND or FD (for example, non-contained muons), constraining the effects of energy feed-down in the liquid detectors.

The ability to separate charge signs will allow the MPD to measure the relative contributions of ν_μ and $\bar{\nu}_\mu$ in both the neutrino beam and the antineutrino beam, as well as distinguishing ν_e from $\bar{\nu}_e$ components. These components are important to anchor the oscillation fit. Otherwise, reliance on the beam modeling is needed to predict the small but uncertain fractions of wrong-sign neutrinos in the beams. Stopping muons' Michel signatures can be used on a statistical basis in the far detector, as the capture/decay rates differ for μ^+ and μ^- , but that is after oscillation has distorted the spectrum; no corresponding test is present for ν_e .

- **Constrain π^0 backgrounds to ν_e events**

An accurate measurement of backgrounds to the ν_e appearance measurement is a critical input for far detector oscillation analyses. In the LArTPC, the largest background to ν_e 's is NC- π^0 interactions in which one photon is not detected and the other is mistakenly identified as an electron. The HPgTPC and ECAL together provide a unique capability to constrain NC- π^0 backgrounds that are misidentified as ν_e CC in the LArTPC. The HPgTPC will collect a *background-free* sample of $\sim 20\text{k}$ ν_e CC events per year. The LArTPC detector measures ν_e + mis-ID'ed π^0 events, while the MPD measures ν_e CC events alone (by rejecting all π^0 events using the ECAL). The MPD sample will not have backgrounds from NC- π^0 events because the photon conversion length in gas is much greater than that in the liquid, and photons from π^0 decays will rarely convert in the gas volume of the HPgTPC in such a way as to fake e^\pm from ν_e interactions. The ECAL, however, will have excellent ability to detect the π^0 decays, enabling the MPD to reject π^0 events as background to ν_e 's.

The MPD measurement of ν_e CC events can be scaled to the LArTPC density and volume and corrected to the same acceptance as the LArTPC in order to provide a constraint on the π^0 -misID. The difference of the two: $(\nu_e^{\text{LAr}} + \text{mis-ID'ed } \pi^0) - (\nu_e^{\text{GAr-scaled-to-LAr}})$ yields the π^0 -misID rate in LArTPC. This measurement of the backgrounds to ν_e 's would not be possible at all if the MPD were replaced by a simple muon range detector. It would also not be easy to extrapolate to the LArTPC if the target material of the MPD were not argon.

- **Measure energetic neutrons from ν -Ar interactions via time-of-flight with the ECAL**

Neutron production in neutrino and antineutrino interactions is highly uncertain, and is a large source of neutrino energy misreconstruction. In the HPgTPC+ECAL system, a preliminary study of the time-of-flight from the HPgTPC neutrino interaction point to hits in the ECAL is encouraging, indicating that TOF can be used to detect and correctly identify neutrons. With the current ECAL design, an average neutron detection efficiency of 60% is achieved by selecting events in which an ECAL cell has one hit with more than 3 MeV. This is still very preliminary work, and further studies to understand the impact of backgrounds

and ECAL optimization (strip vs. tile, overall thickness) are underway.

- **Reconstruct neutrino energy via spectrometry and calorimetry**

Although all neutral particles from an event must be measured with the ECAL in the MPD, the HPgTPC will be able to make very precise momentum measurements of charged particle tracks with a larger acceptance than the upstream LArTPC, including tracks created by high-momentum exiting particles, which allows the measurement of the entire neutrino spectrum. In addition, short and/or stopping tracks will be measured via dE/dx . The sum of this capability provides a complementary event sample to that obtained in the LAr, whose much higher density leads to many secondary interactions for charged particles. The two methods of measurement in the MPD will help in understanding the LAr energy resolution.

- **Constrain LArTPC detector response and selection efficiency**

The MPD will collect large amounts of data in each of the exclusive neutrino interaction channels, with the exception of $\nu - e$ elastic scattering, where the HPgTPC sample will be too small to be useful. The high statistics ν -Ar interaction samples will make it possible to directly crosscheck every kinematic distribution that will be used to constrain the fluxes and cross sections. Typically these checks will be over an extended range of that variable. The high purity of the MPD samples and low detection threshold for final state particles in the HPgTPC will give a benchmark or constraint on LArTPC detector response and selection efficiencies for each of the interaction channels.

Using the events collected in the HPgTPC (where selection and energy reconstruction are easy), the performance of LAr event selection and energy reconstruction metrics can be tested by simulating the well-measured HPgTPC four-vectors in the LArTPC. This allows the validation of the LArTPC reconstruction performance on these events. This process is expected to reduce the errors in the LArTPC detector energy response model.

MPD strengths The strengths of the MPD enable it to reach the goals above and to augment the capabilities of the LArTPC. Below are a few examples of its strengths relative to the LArTPC:

- High-fidelity particle charge determination via magnetic curvature. This is the only detector that can measure electron and positron charge.
- Precise and independent measurement of particle momentum, via magnetic curvature, will allow the measurement of the momentum of higher-energy charged particles without requiring containment. This extends the utility of the ND, especially for the high-energy beam tune. The absolute momentum scale is easily calibrated in the magnetic spectrometer and provides a cross-check on energy loss through ionization measurements. Calibration strategies for the magnetic tracking include pre-assembly field maps, *in situ* NMR probes, and K_s^0 and Λ^0 reconstruction.
- Particle identification through dE/dx . The gaseous argon TPC has better separation power of particle species by dE/dx than the liquid because the momentum can be measured along

1 with energy loss.

- 2 • High-resolution imaging of particles emerging from the ν -Ar vertex (including nucleons). The
3 reconstruction threshold in the gas phase is much lower than the threshold in liquid because
4 particles travel further in the low density medium, e.g., a proton requires only 3.7 MeV kinetic
5 energy to make a 2 cm track in 10 atmospheres of gaseous argon, while a 3.7 MeV proton in
6 liquid can only travel 0.02 cm. Figure 1.17 demonstrates the difference in the thresholds for
7 reconstructing protons in the HPgTPC and the LArTPC in light of the energy spectra of
8 final state protons from a selection of types of neutrino interactions at the DUNE ND. The
9 LArTPC threshold is what has been achieved in MicroBooNE up to now, and the HPgTPC
10 threshold is what has been achieved with the tools discussed in Sec. 1.3.3.2.3.
- 11 • Separation of tracks and showers for less-ambiguous reconstruction. Particle tracks are locally
12 helical and tend to bend away from each other in the magnetic field as they travel from a
13 dense vertex. Electromagnetic showers do not occur in the gas, but in the physically separate
14 ECAL. By contrast, in a LArTPC tracks and showers frequently overlap. The measurement
15 resolution scales are comparable between the liquid and the gas, but the distance scales on
16 which interactions occur are much longer in the gas, allowing particles to be identified and
17 measured separately more easily.
- 18 • The measurement of energetic neutrons through time-of-flight with ECAL is a potential
19 game-changer for validating energy reconstruction. Preliminary studies of the HPgTPC+ECAL
20 system indicate that an average neutron detection efficiency of 60% can be achieved via a
21 time-of-flight analysis. A study of the impact of backgrounds is underway, but initial studies
22 are encouraging.
- 23 • The HPgTPC is able to measure the momentum of particles over almost the full solid angle.
24 Particles that are emitted at a large angle with respect to the beam have a high probability
25 of exiting the liquid argon without leaving a matching track in the MPD. However, events
26 collected in the HPgTPC, with its $\simeq 4\pi$ coverage, can be used in the regions of phase space
27 where the exiting fraction is high in the liquid argon to ensure that the events are accurately
28 modeled in all directions in the far detector.
- 29 • The MPD neutrino event sample, while smaller than the LArTPC sample, is a statistically
30 independent sample. Moreover, the systematic uncertainties of the MPD will be very different
31 than the LArTPC and likely smaller in many cases. This will allow the MPD to act as a
32 systematics constraint for the LArTPC.

33 1.3.3.1 MPD Technical Details

34 1.3.3.1.1 High-Pressure Gaseous Argon TPC

35 The basic geometry of the HPgTPC is a gas-filled cylinder with a high-voltage electrode at its mid-
36 plane, providing the drift field for ionization electrons. It is oriented inside the magnet such that the

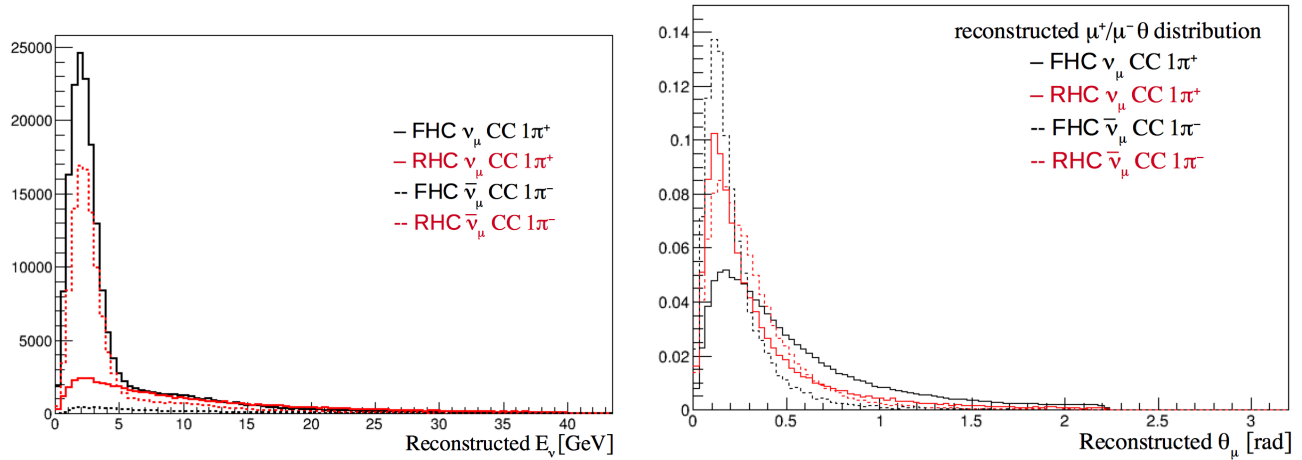


Figure 1.16: Representative differences among subsamples of charged current ν_μ interactions with one π^+ (solid lines) and charged current $\bar{\nu}_\mu$ interactions with one π^- (dashed lines). The forward- and reverse- horn current samples are shown in black and red, respectively. Left: Reconstructed neutrino energy spectra, normalized to the same number of protons on target. Right: Angle of outgoing muon relative to neutrino direction, normalized to unit area for shape comparison.

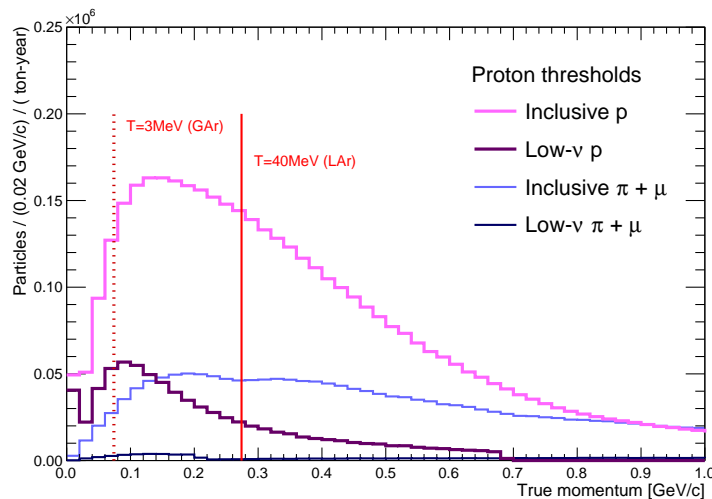


Figure 1.17: The momentum spectra of protons ejected from neutrino interactions in argon, for several categories of interaction types. The vertical lines indicate the lowest momentum protons that have been reconstructed using existing automated reconstruction tools, where the dotted line is the HPgTPC threshold, and the solid line is the LArTPC threshold (from MicroBooNE).

1 magnetic and electric fields are parallel, reducing transverse diffusion to give better point resolution.
 2 Primary ionization electrons drift to the end plates of the cylinder, which are instrumented with
 3 multi-wire proportional chambers to initiate avalanches (gas gain) at the anode wires. Signals
 4 proportional to the avalanches are induced on cathode pads situated behind the wires; readout of
 5 the induced pad signals provides the hit coordinates in two dimensions. The drift time provides
 6 the third coordinate of the hit.

7 The details of the HPgTPC design will be based closely on the design of the ALICE detector [58]
 8 shown in Figure 1.18. Two readout planes sandwich a central high-voltage electrode (25 μm of
 9 aluminized mylar) at high voltage that generates the drift field, which is parallel to a 0.5 T magnetic
 10 field. On each side of the electrode, primary ionization electrons drift up to 2.5 m to reach the
 11 endplates, which are segmented azimuthally into 18 trapezoidal regions instrumented with readout
 12 chambers (ROCs) that consist of MWPC amplification regions and cathode pad planes to read out
 13 the signals. A cross sectional view of an ALICE MWPC-based ROC is shown in Figure 1.19. The
 14 ROCs are built in two sizes: a smaller inner readout chamber (IROC) and a larger outer readout
 15 chamber (OROC). The trapezoidal segments of the endplates are divided radially into inner and
 16 outer sections, and the IROCs and OROCs are installed in those sections. The existing IROCs and
 17 OROCs will become available in 2019, when they are scheduled to be replaced by new GEM-based
 18 ROCs for upgraded pileup capability in the high rate environment of the LHC. For the DUNE
 19 HPgTPC, the existing ROCs are more than capable of providing the necessary performance in a
 20 neutrino beam.

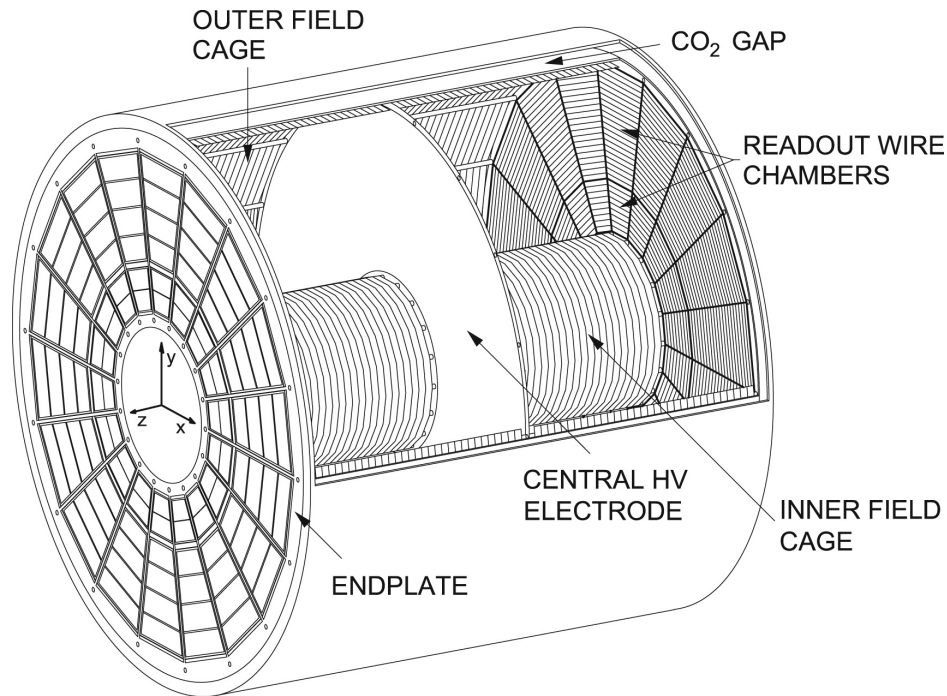


Figure 1.18: Diagram of the ALICE TPC, from Ref. [59]. The drift HV cathode is located at the center of the TPC, defining two drift volumes, each with 2.5 m of drift along the axis of the cylinder toward the endplate. The endplates are divided into 18 sectors, and each endplate holds 36 readout chambers.

21 In the ALICE design, the innermost barrel region was isolated from the TPC and instrumented with
 22 a silicon-based inner tracker; for the DUNE HPgTPC, the inner field cage labeled in Figure 1.18
 23 will be removed, and the entire inner region will be combined to make a single gas volume for

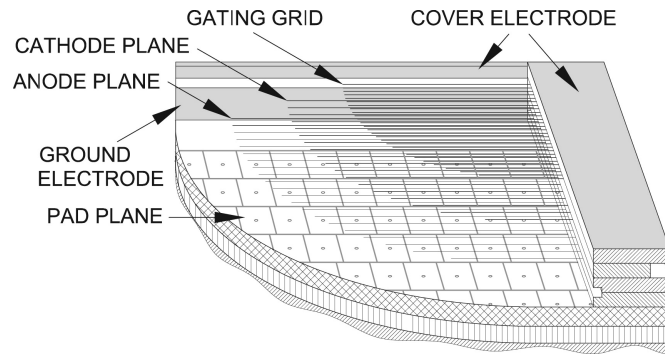


Figure 1.19: Schematic diagram of the ALICE MWPC-based ROC with pad plane readout, from Ref. [59].

1 the TPC. New ROCs will be built to fill in the central uninstrumented region, which is 1.6 m in
 2 diameter, left by reusing the existing ALICE chambers. The active dimensions of the HPgTPC
 3 will be 5.2 m in diameter and 5 m long, which yields an active mass of $\simeq 1.8$ t.

4 **MPD Pressure Vessel** The preliminary design of the pressure vessel, presented in Figure 1.20,
 5 accounts for the additional volume needed to accommodate the TPC field cage, the ROC support
 6 structure, front-end electronics, and possibly part of the ECAL.

7 The pressure vessel can be fabricated from aluminum or stainless steel, has a cylindrical section
 8 that is 6 m in diameter and 6 m long and utilizes two semi-elliptical end pieces with flanges.
 9 The walls of the cylinder barrel section are $\simeq 1.6X_0$ in thickness in the case of stainless steel or
 10 $\simeq 0.3X_0$ in the case of Al 5083. Further reduction of the thickness in radiation lengths can be
 11 accomplished with the addition of stiffening rings. This preliminary design includes two flanged
 12 endcaps. However, these large flanges are the cost driver for the pressure vessel and, therefore,
 13 vessel designs with a single flange will also be considered. As an example, DOE/NETL-2002/1169
 14 (Process Equipment Cost Estimation Final Report) indicates that a horizontal pressure vessel of
 15 the size indicated here and rated at 1034 kPag (150 psig) (approximately 10 atmospheres) is costed
 16 at \$150k (\simeq \$210k in 2019 dollars). The budgetary estimate for a vessel with two flanges was
 17 \$1.2M with the flanges driving the cost. DOE/NETL-2002/1169 also indicates that pressure is not
 18 a significant cost driver. Reducing the pressure from 1034 kPag to 103 kPag (15 psig) only reduces
 19 the basic (\$150k) vessel cost by a factor of two.

20 1.3.3.1.2 Electromagnetic Calorimeter (ECAL)

21 The MPD ECAL concept is based on a high granularity calorimeter to provide direction information
 22 in addition to the energy measurement of electromagnetic showers and an efficient rejection of
 23 background. The principal role of the ECAL is to reconstruct photons directly produced in neutrino
 24 interactions and originating from π^0 decays, providing a measurement of the photon's energy and
 25 direction to enable the association of photons to interactions observed in the HPgTPC and the
 26 determination of the decay vertex of the π^0 s. In the case of ν_e measurements in the HPgTPC, the
 27 ECAL will play an important role in rejecting events with π^0 decays, which represent a background

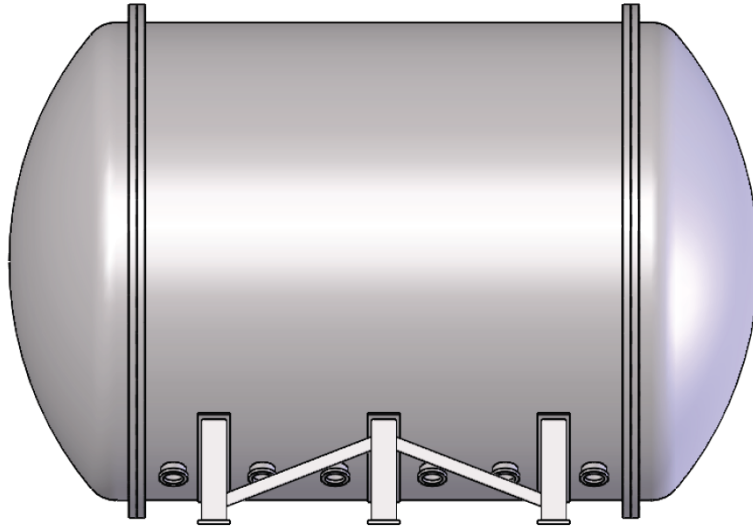


Figure 1.20: Pressure vessel preliminary design.

1 to ν_e interactions in the LArTPC. The ECAL can also be used to reject external backgrounds,
 2 such as rock neutrons and muons, providing a sub-nanosecond timestamp [60] for each hit in the
 3 detector. As the ECAL uses hydrogen-rich scintillator, it is assumed to have capabilities to provide
 4 neutron detection, and studies are underway to determine the performance of neutron detection.

5 **ECAL Design** The ECAL design is inspired by the design of the CALICE Analog Hadron
 6 Calorimeter (AHCAL) [61].

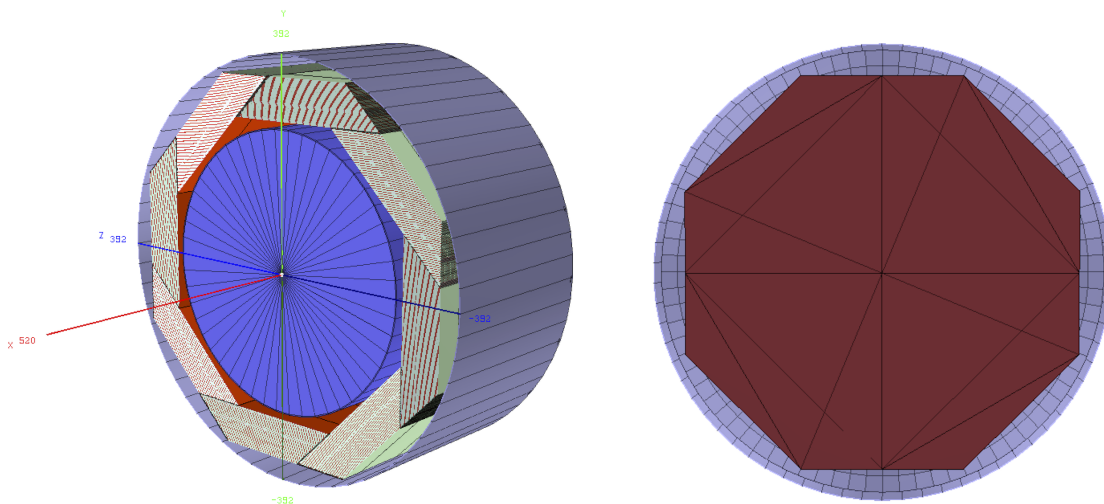


Figure 1.21: On the left, the conceptual design of the HPgTPC + ECAL barrel system for the ND. In the preliminary design, the full ECAL is located inside the HPgTPC pressure vessel. On the right, a conceptual design of the ECAL endcap system.

7 As shown in Figure 1.21, the ECAL Barrel has an octagonal shape, each quadrant composed

1 of several trapezoidal modules. The ECAL endcap has a similar design providing hermeticity
 2 and a large solid-angle coverage. Each module consists of scintillating layers of polystyrene as
 3 active material readout by silicon-photomultipliers, sandwiched between absorber sheets. The
 4 scintillating layers consist of a mix of tiles with dimensions between $2 \times 2 \text{ cm}^2$ to $3 \times 3 \text{ cm}^2$ (see
 5 Figure 1.22) and cross-strips with embedded wavelength shifting fibers to achieve a comparable
 6 effective granularity. The high granularity layers are concentrated in the front part of the detector,
 7 since that has been shown to be the most relevant factor for the angular resolution [62]. With the
 8 current design, the number of channels is of the order 2.5 to 3 million. A first design of the ECAL
 9 and the simulated performance has already been studied in [62].

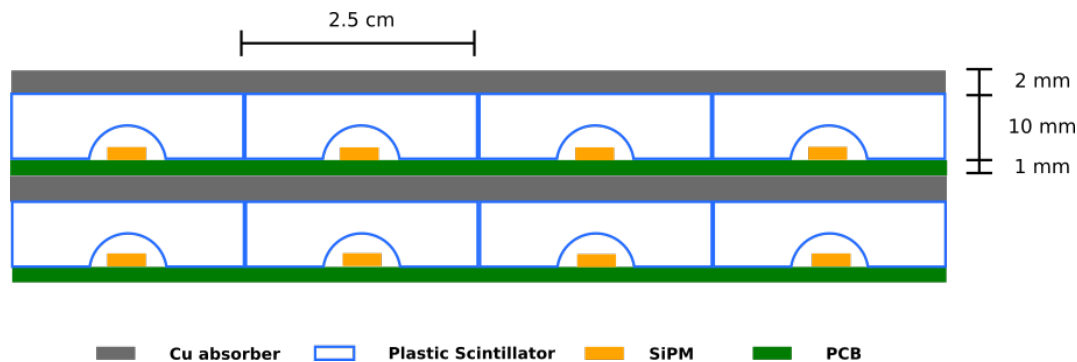


Figure 1.22: Conceptual layout of the calorimeter showing the absorber structure, scintillator tiles, SiPM and PCB.

10 One major design driver is the pressure vessel limiting the space. Therefore, a compact calorimeter
 11 is needed. In preliminary designs, it was assumed that the full ECAL barrel would be fit inside
 12 the pressure vessel. Studies have also investigated the possibility to have two segments, one inside
 13 and one outside the pressure vessel. The thickness of the pressure vessel has an impact on the
 14 calorimeter energy resolution [62], and more recent designs of the pressure vessel have reduced
 15 its thickness enough that the entire ECAL may be placed outside the pressure vessel. Currently,
 16 the ECAL design is undergoing a detailed design study in order to further optimize the detector
 17 design, cost, and performance.

18 1.3.3.1.3 Magnet

19 Two magnet designs are under consideration to house the high pressure gas TPC and the ECAL.
 20 One is a UA1-type conventional electromagnet, the other is based on a superconducting Helmholtz-
 21 coil-like design. The common requirement is a central magnetic field of 0.5 T with $\pm 20\%$ uniformity
 22 over the TPC volume (5 m long and 5 m in diameter). With the current design of the access
 23 shaft (11.8 m diameter), the clear diameter is about 7.8 m. Recent studies for the construction
 24 of an electromagnet similar to the UA1 magnet predict that the cost of the design, procurement,
 25 infrastructure (power and cooling) and assembly will be in excess of \$20 million, with operation
 26 costs of approximately \$1.6M per year of running. Because of this, the main focus has been on
 27 the superconducting design.

- 1 **Superconducting Magnet** The SC magnet design is an Helmholtz-coil-like configuration, air
 2 core, five coil magnet system. Three central coils produce the analyzing field and two outer
 3 shielding coils help contain stray field. The advantage of this design is that little or no iron is used
 4 for field containment or shaping. This eliminates background coming from neutrino interactions in
 5 the iron, which for the normal-conducting magnet case is the largest component of the background.
 6 Figure 1.23 shows the magnet concept indicating the 5-coil arrangement and support structure.

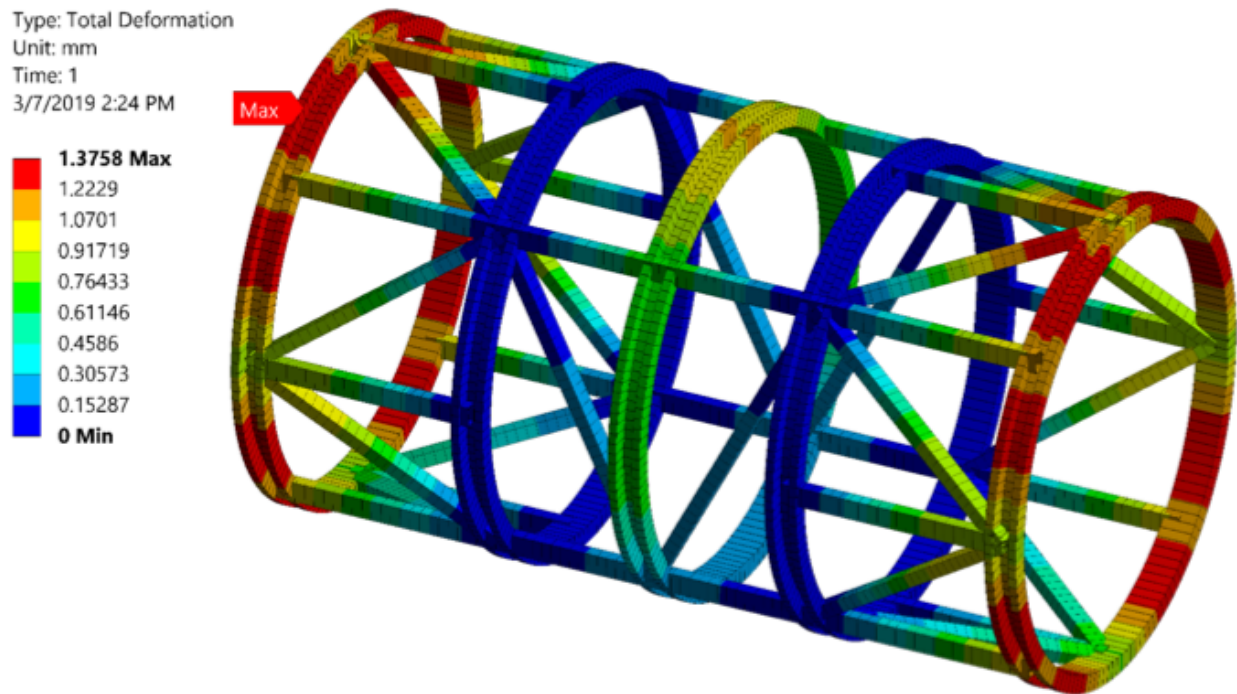


Figure 1.23: Helmholtz coil arrangement for MPD superconducting magnet.

- 7 All five coils have the same inner radius of 3.5 m. The center and shielding coils are identical with
 8 the same number of ampere-turns. The side coils are placed at 2.5 m, the shielding coils at 6 m
 9 from the magnet center along Z. The case where the shielding coils are at 5 m from the magnet
 10 center so that the magnet system would be the same width as the LAr detector is also being
 11 examined. The magnet system will have a stored energy of about 110 MJ, using a conventional
 12 NbTi superconducting cable design, a SSC-type Rutherford cable soldered in a copper channel with
 13 a 50% margin. All coils should be wired in series to reduce imbalanced forces during a possible
 14 quench. Possible small transverse centering force components are possible due to coil de-centering
 15 from mechanical errors. Shown in Figure 1.24 is the field along the Z-axis at different radii. The
 16 peak field in the coils is 2.14 T (center), 5 T (side) and 2.03 T (shield). The resulting forces are
 17 only along the Z-axis, F_z is 0.0 MN (center), -6.81 MN (side) and 2.2 MN (shield). The fringe field
 18 at the shielding coil is rather large but can be reduced further; more studies will be needed. There
 19 is a preliminary mechanical support design. A first glimpse at the radiative heat load assumes
 20 a coil and support surface of 180 m², resulting in a load of 5.4 W from 77 K to 4.5 K. The coil
 21 support and leads will likely have a much larger contribution (HTS power leads usually have 15 W
 22 for 10 kA. With a mass of 42 t the magnets are in some aspects similar to the Mu2e solenoids.

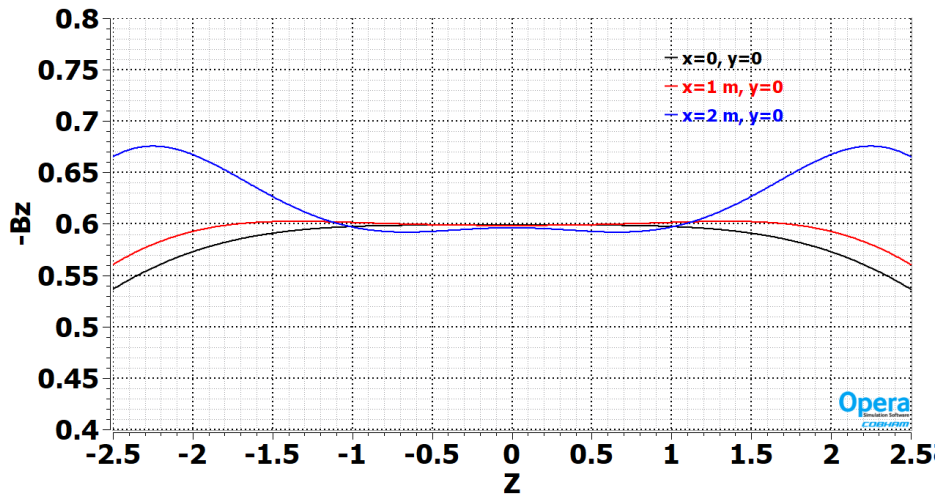


Figure 1.24: Field map of the superconducting magnet along the Z-axis. The colors represent different radii from the center line.

1 **Normal Conducting Magnet** Although the SC magnet design is the favored option, the normal
 2 conducting magnet design produced for the LBNE CDR is also being revised and studied. Due
 3 to the cylindrical geometry imposed by the HPgTPC, a cylindrical coil design for the normal
 4 conducting magnet is the baseline. The cooling requirement of the coil is approximately 3.5 MW
 5 and involves a substantial LCW flow. A thermal shield between the coils and the detector volume
 6 is required in order to minimize heat flow to the HPgTPC and the ECAL. The coil thickness
 7 becomes excessive (in order to maintain a maximum 5° C temperature in the coil) if the thermal
 8 shield is not used. The shield does take up space in the magnet volume, however. Note: the iron
 9 end-walls will most likely not be needed. The estimated magnet weight is well over 1 kt, and this
 10 mass provides a significant source of background for the high pressure gas TPC and, perhaps, the
 11 LAr. There is a significant amount of material between the LArTPC and the HPgTPC in the
 12 MPD in this configuration, which will affect the acceptance for muons emanating from events in
 13 the LAr. This option will continue to be studied as part of the optimization process.

14 1.3.3.1.4 Size optimization

15 The process of optimizing the design of the MPD is in progress. One of the more critical issues is
 16 the size of the MPD. This is an important factor in the angular acceptance of particles exiting the
 17 upstream LArTPC. A preliminary study of geometries shows that reducing the HPgTPC diameter
 18 by more than 1 meter, or reducing the length by more than 1.5 meters would have significant
 19 consequences on the acceptance. Reducing the HPgTPC diameter from its nominal 5 meters to a
 20 slightly smaller 4.5 meters while increasing its length in the direction transverse to the neutrino
 21 beam improves acceptance, since the HPgTPC would better match the 7-meter width of the
 22 LArTPC in the transverse direction. It should be noted, however, that reducing the diameter may
 23 actually result in a higher-cost MPD, since the ALICE TPC readout chambers would not be used
 24 in the configuration for which they were designed. Increasing the length of the HPgTPC is feasible,
 25 but will require additional studies of high voltage stability in the gas, since HV breakdown in gas
 26 is proportional to the pressure (in the absence of field enhancements). The HPgTPC operating

- 1 pressure will be nominally 10 times that of ALICE, so extending the drift distance from 2.5 meters
 2 to 3 meters while keeping the same drift velocity will require raising the drift HV by approximately
 3 20 kV.

4 1.3.3.2 MPD performance

- 5 The expected performance of the MPD is summarized in Table 1.3. Details of the HPgTPC
 6 performance are based upon experience from operation of the PEP-4 [63, 64, 65] and ALICE [66]
 7 time projection chambers. Performance of the ECAL is based on experience from operation of
 8 similar ECALs and on simulations.

Table 1.3: Expected MPD performance, extrapolated from ALICE

Parameter	Value	units
σ_x	250	μm
σ_y	250	μm
σ_z	1500	μm
$\sigma_{r\phi}$	<1000	μm
Two-track separation	1	cm
Angular resolution	2-4	mrad
$\sigma(dE/dx)$	5	%
σ_{p_T}/p_T	0.7	% (10-1 GeV/c)
σ_{p_T}/p_T	1-2	% (1 GeV/c to 0.1 GeV/c)
Energy scale uncertainty	$\lesssim 1$	% (dominated by δ_p/p)
Charged particle detection thresh.	5	MeV (K.E.)
ECAL energy resolution	$5-7/\sqrt{E/\text{GeV}}$	%
ECAL pointing resolution	$\simeq 6$ at 500 MeV	degrees

9 1.3.3.2.1 Track Reconstruction and Particle Identification

- 10 The combination of very high resolution magnetic analysis and superb particle identification from
 11 the HPgTPC, coupled with a high-performance ECAL will lead to excellent event reconstruction
 12 capabilities and potent tools to use in neutrino event analysis. As an example of this capability,
 13 the top panel of Figure 1.25 shows a $\nu_e + Ar \rightarrow e^- + \pi^+ + n$ in the HPgTPC with automatically-
 14 reconstructed tracks. The same event was simulated in a single-phase LAr far detector module,
 15 and is shown in the bottom panel of Figure 1.25.

- 16 Since important components of the hardware and design for the HPgTPC are taken from or
 17 duplicated from the ALICE detector, the ALICE reconstruction is a useful point of reference
 18 in this discussion. Track reconstruction in ALICE is achieved by combining hits recorded on the
 19 ROC pads into tracks following a trajectory that a charged particle traveled through the TPC drift
 20 volume. The HPgTPC is oriented so that the neutrino beam is perpendicular to the magnetic field,

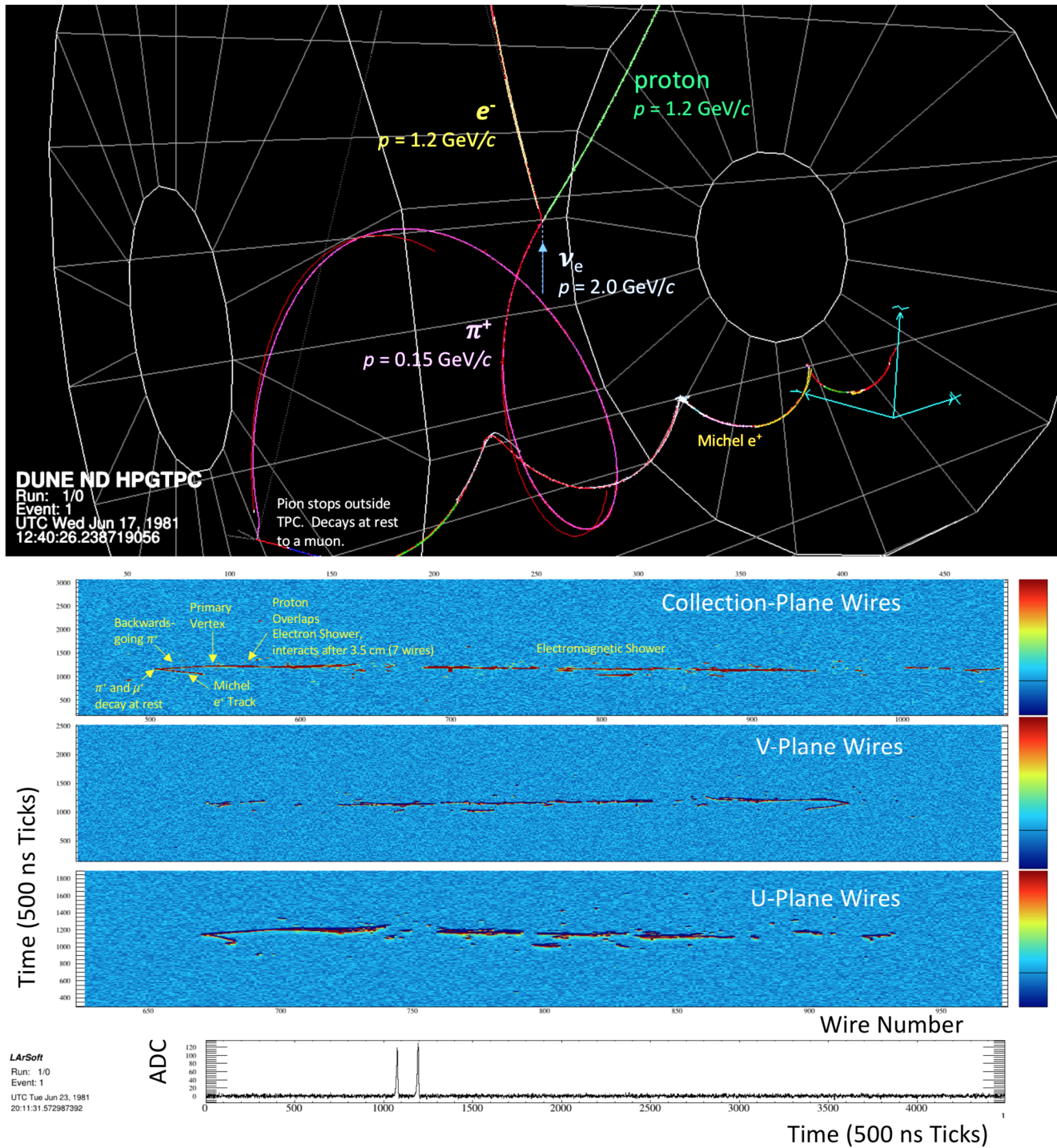


Figure 1.25: (Top) Track-reconstructed ν_e CC event in the HPgTPC, simulated and reconstructed with GARSoft. The annotations are from MC truth. (Bottom) The same ν_e CC event, but simulated in a single-phase Far Detector module using LArSoft. The topmost blue panel shows the collection-plane view, the middle blue panel shows the V view, and the bottom blue panel shows the U view. Wire number increases on the horizontal axes and sample time along the vertical axes. The wire number in the collection view is labeled on the top of the panel, while the V and U wire numbers are below their respective panels. Simulated ADC values are indicated by the colors. The curve in the bottom-most panel is a simulated waveform from a collection-plane wire. The annotations are from MC truth.

1 which is the most favorable orientation for measuring charged particles traveling along the neutrino
2 beam direction.

3 The GArSoft simulation and reconstruction package borrows heavily from LArSoft, and is based
4 on *art* and GEANT4. It is designed to be able to reconstruct tracks with a full 4π acceptance.
5 GArSoft simulates a 10 atmosphere gaseous argon detector with readout chambers filling in the
6 central holes in the ALICE geometry. GArSoft's tracking efficiency has been evaluated in a large
7 sample of GENIE ν_μ events interacting in the TPC gas at least 40 cm from the edges, generated
8 using the optimized LBNF forward horn current beam spectra. The efficiency for reconstructing
9 tracks associated with pions and muons as a function of track momentum p is shown in Figure 1.26.
10 The efficiency is above 90% for tracks with $p > 40$ MeV/ c , and it steadily rises with increasing
11 momentum.

12 Also shown is the efficiency for reconstructing all charged particles with $p > 200$ MeV/ c as a func-
13 tion of λ , the track angle with respect to the center plane. The tracking efficiency for protons is
14 shown in Figure 1.27 as a function of kinetic energy, T_p . Currently, the tracking works well down to
15 $T_p \simeq 20$ MeV. For $T_p < 20$ MeV, a machine-learning algorithm is in development, targeting short
16 tracks near the primary vertex. This algorithm, although currently in a very early stage of devel-
17 opment, is already showing good performance(see Figure 1.31), and efficiency improvements are
18 expected with more development. The machine learning algorithm is described in section 1.3.3.2.3.

19 The ALICE detector, as it runs at the LHC, typically operates with particle densities ranging from
20 2000 to 8000 charged particles per unit rapidity (dN/dy) for central Pb-Pb interactions [67]. The
21 expected particle densities in the DUNE neutrino beam will be much lower and less of a challenge
22 for the reconstruction.

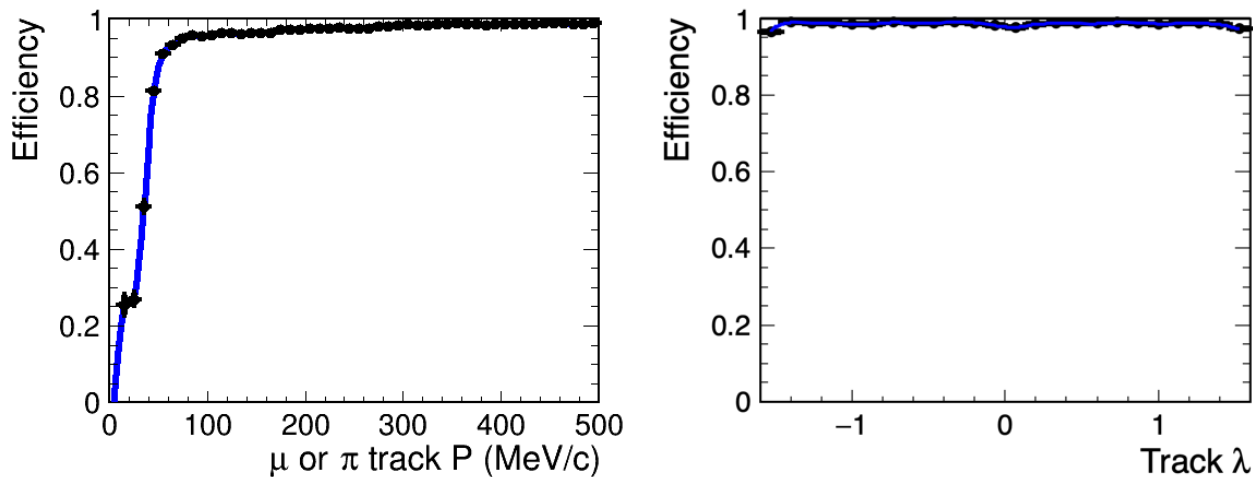


Figure 1.26: (Left) The efficiency to find tracks in the HPgTPC as a function of momentum, p , for tracks in a sample of GENIE events simulating 2 GeV and ν_μ interactions in the gas, using GArSoft. (Right) The efficiency to find tracks as a function of λ , the angle with respect to the center plane, for tracks with $p > 200$ MeV/ c .

23 ALICE chose to use neon, rather than argon, for the primary gas in their first run; the decision
24 was driven by a number of factors, but two-track separation capability was one of the primary

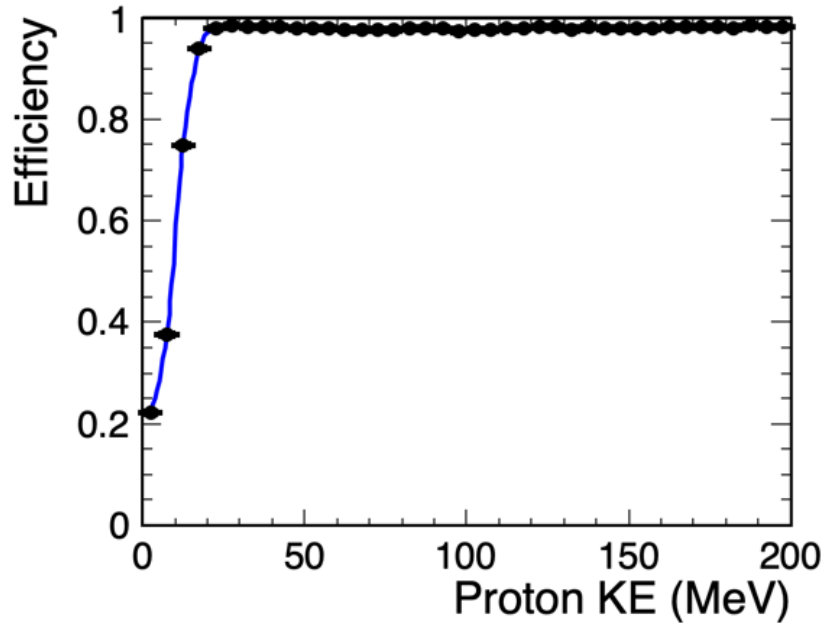


Figure 1.27: Tracking efficiency for protons as a function of kinetic energy.

1 motivations due to the extremely high track multiplicities in the experiment. Neon performs
 2 better than argon in this regard. A better comparison for the HPgTPC's operation in DUNE is
 3 the two-track separation that was obtained in PEP4 [64]. PEP4 ran an 80-20 mixture of Ar-CH₄
 4 at 8.5 atmospheres, yielding a two-track separation performance of 1 cm.

5 In ALICE, the ionization produced by charged particle tracks is sampled by the TPC pad rows
 6 (there are 159 pad rows in the TPC) and a truncated mean is used for the calculation of the
 7 PID signal. Figure 1.28 (left) shows the ionization signals of charged particle tracks in ALICE for
 8 pp collisions at $\sqrt{s} = 7$ TeV. The different characteristic bands for various particles are clearly
 9 visible and distinct at momenta below a few GeV. When repurposing ALICE as the HPgTPC
 10 component of the MPD, better performance is expected for particles leaving the active volume,
 11 since the detector will be operating at higher pressure (10 atmospheres vs. the nominal ALICE
 12 1 atmosphere operation), resulting in ten times more ionization per unit track length available for
 13 collection. Figure 1.28 (right) shows the charged particle identification for PEP-4/9 [68], a higher
 14 pressure gas TPC that operated at 8.5 atmospheres, which is very close to the baseline gas mixture
 15 and pressure of the DUNE HPgTPC.

16 1.3.3.2.2 Momentum and Angular Resolution for Charged Particles

17 The ability to determine the sign of the charge of a particle in the HPgTPC tracking volume is
 18 limited by the spatial resolution of the measured drift points in the plane perpendicular to the
 19 magnetic field, as well as multiple Coulomb scattering (MCS) in the gas. For a fixed detector
 20 configuration, the visibility of the curvature depends on the particle's p_T , its track length in the
 21 plane perpendicular to the field, and the number and proximity of nearby tracks. Because primary

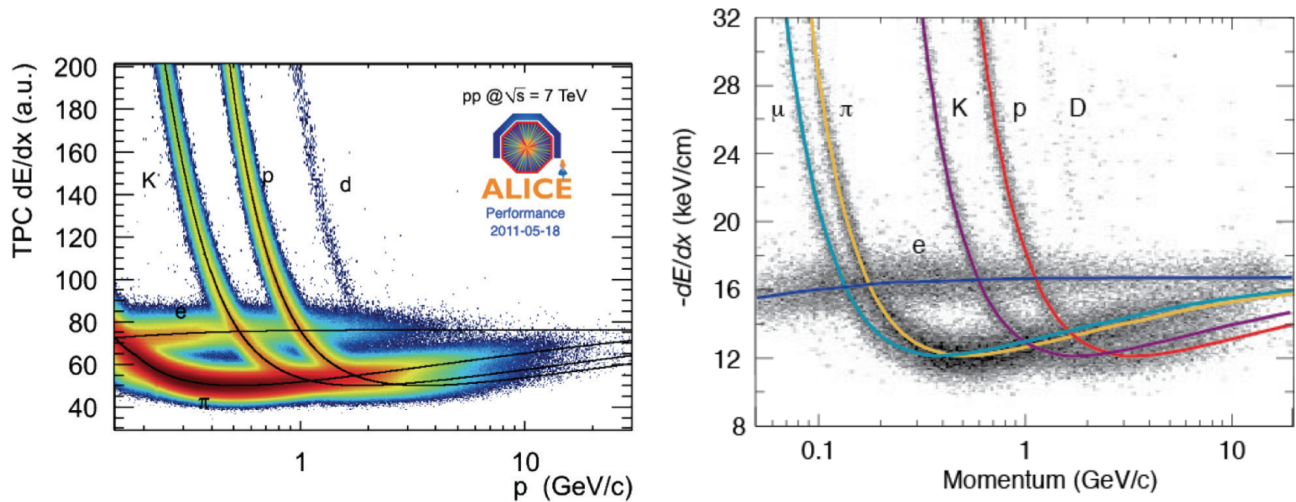


Figure 1.28: Left: ALICE TPC dE/dx -based particle identification as a function of momentum (from [69]). Right: PEP-4/9 TPC (80:20 Ar-CH₄, operated at 8.5 Atm, from [68]) dE/dx -based particle identification.

1 vertices are distributed throughout the tracking volume, the distribution of the lengths of charged-
 2 particle tracks is expected to start at very short tracks, unless sufficient fiducial volume cuts are
 3 made to ensure enough active volume remains to determine particle's track sign. The kinetic
 4 energies of particles that leave short tracks and stop in the detector will be better measured from
 5 their tracks' lengths than from their curvatures. Protons generally stop before their tracks curl
 6 around, but low-energy electrons loop many times before coming to rest in the gas.

7 Within the fiducial volume of the HPgTPC, charged particles can be tracked over the full 4π solid
 8 angle. Even near the central electrode, tracking performance will not be degraded due to the very
 9 thin ($25\ \mu\text{m}$ of mylar) nature of the central electrode. Indeed, tracks crossing the cathode provide
 10 an independent measurement of the event time, since the portions of the track on either side of
 11 the cathode will only line up with a correct event time assumed when computing drift distances.
 12 The 4π coverage is true for all charged particles. ALICE ran with a central field of 0.5 T and their
 13 momentum resolution from p -Pb data [70] is shown in Figure 1.29.

14 The momentum resolution of muons in neutrino scatters using the GARSoft simulation and recon-
 15 struction package is shown in Figure 1.30, using a sample of 2 GeV ν_μ CC events. This resolution
 16 differs from ALICE's achieved resolution due to the higher pressure, the heavier argon nucleus
 17 compared with neon, the non-centrality of muons produced throughout the detector, and the fact
 18 that the GARSoft simulation and reconstruction tools have yet to be fully optimized. The momen-
 19 tum resolution achieved for muons is $\Delta p/p = 4.2\%$, and is expected to improve with optimization
 20 of the simulation and reconstruction tools. The 3D angular resolution of muons is approximately
 21 0.8 degrees, as shown in Figure 1.30.

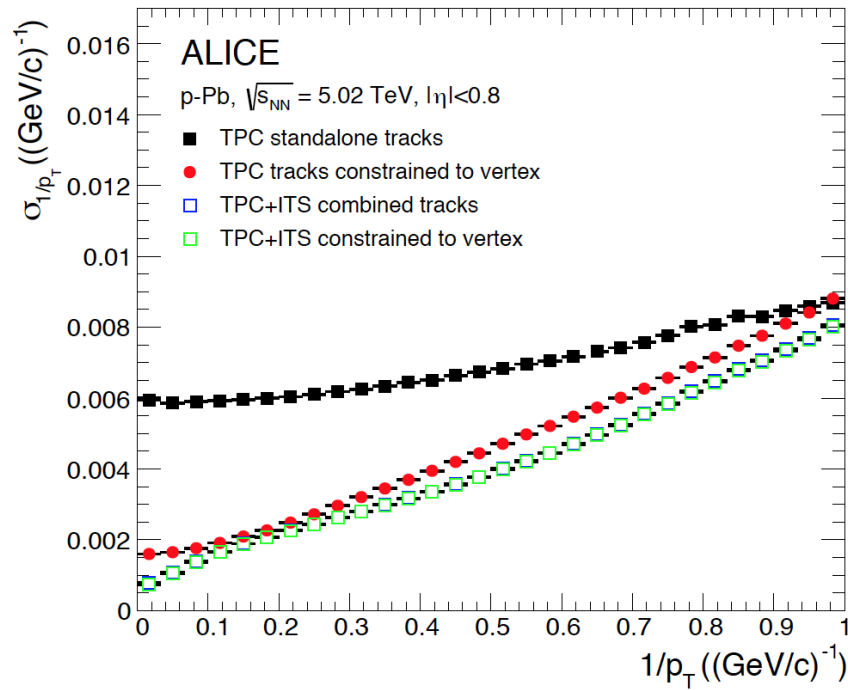


Figure 1.29: The black squares show the TPC stand-alone p_T resolution in ALICE for p -Pb collisions. From Ref. [70].

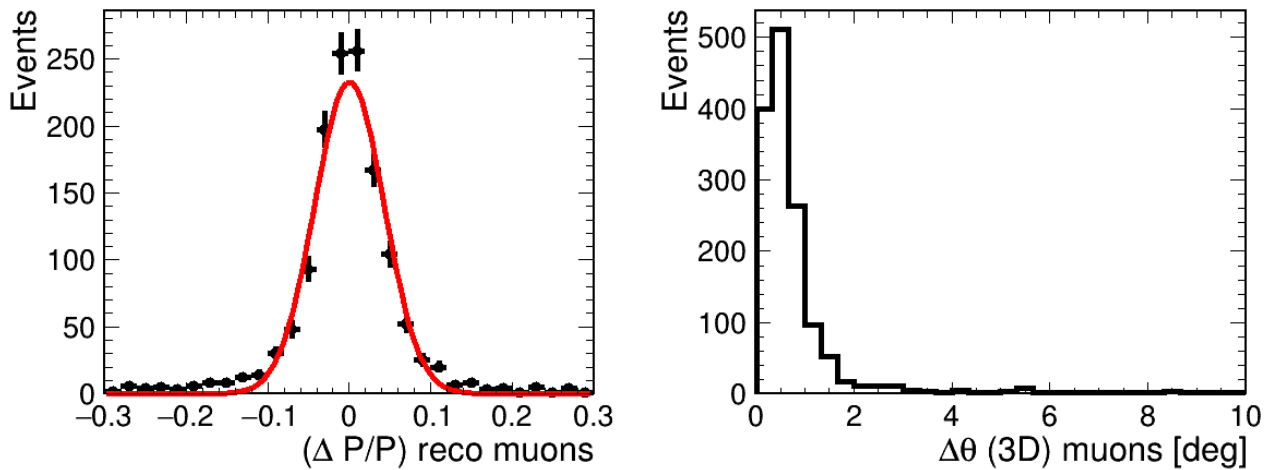


Figure 1.30: Left: the momentum resolution for reconstructed muons in GArSoft, in a sample of 2 GeV ν_μ CC events simulated with GENIE. The Gaussian fit to the $\Delta p/p$ distribution has a width of 4.2%. Right: the 3D angular resolution for the same sample of muons in GArSoft.

1.3.3.2.3 Machine Learning for Low Energy Protons

As a compliment to the existing reconstruction, an initial exploration of several machine learning methods has been performed. The main goal of this effort has been to attempt to reconstruct very low energy protons and pions where traditional tracking methods might struggle. While this study is still in very early stages, there has been success so far in using a fully connected multi-layer perceptron (MLP) to both regress the kinetic energy of and classify between protons and pions. Additionally a Random Sample Consensus (RANSAC) based clustering algorithm has been developed to group hits into short tracks for events where there are multiple particles. Together, these two algorithms can be used to measure the kinetic energy of multiple particles in a single event.

As a demonstration, a test sample of multiple proton events was generated where each event has:

- 0-4 protons, number determined randomly with equal probabilities
- all protons share a common starting point (vertex) whose position in the TPC is randomly determined
- each proton is assigned independently and randomly:
 - a direction in space (isotropically distributed)
 - a scalar momentum between 0 and 200 MeV/ c (flat distributed)

The RANSAC-based clustering algorithm assigns individual hits to proton candidate sets of hits which are passed to a MLP that was trained on a set of individual proton events in the TPC to predict kinetic energy. Figure 1.31 shows the kinetic energy residuals, the reconstruction efficiency, and a 2D scatter plot of the measured kinetic energy versus the true kinetic energy for each individual proton with kinetic energy between 3 and 15 MeV in the test sample. Additionally, the residual for the total kinetic energy in each multi-proton event is given.

1.3.3.2.4 ECAL Performance

The expected performance of the calorimeter was studied with Geant4-based [71] simulations and GARSoft [72]. In the following, a first scenario referred to as scenario A (shown by the red curve in the figures below) in which the ECAL is located inside the pressure vessel is considered. The barrel geometry consists of 55 layers with the following layout:

- 8 layers of 2 mm copper + 10 mm of 2.5×2.5 cm² tiles + 1 mm FR4
- 47 layers of 4 mm copper + 10 mm of cross-strips 4 cm wide

For the present studies, copper has been chosen as absorber material as initial studies have shown

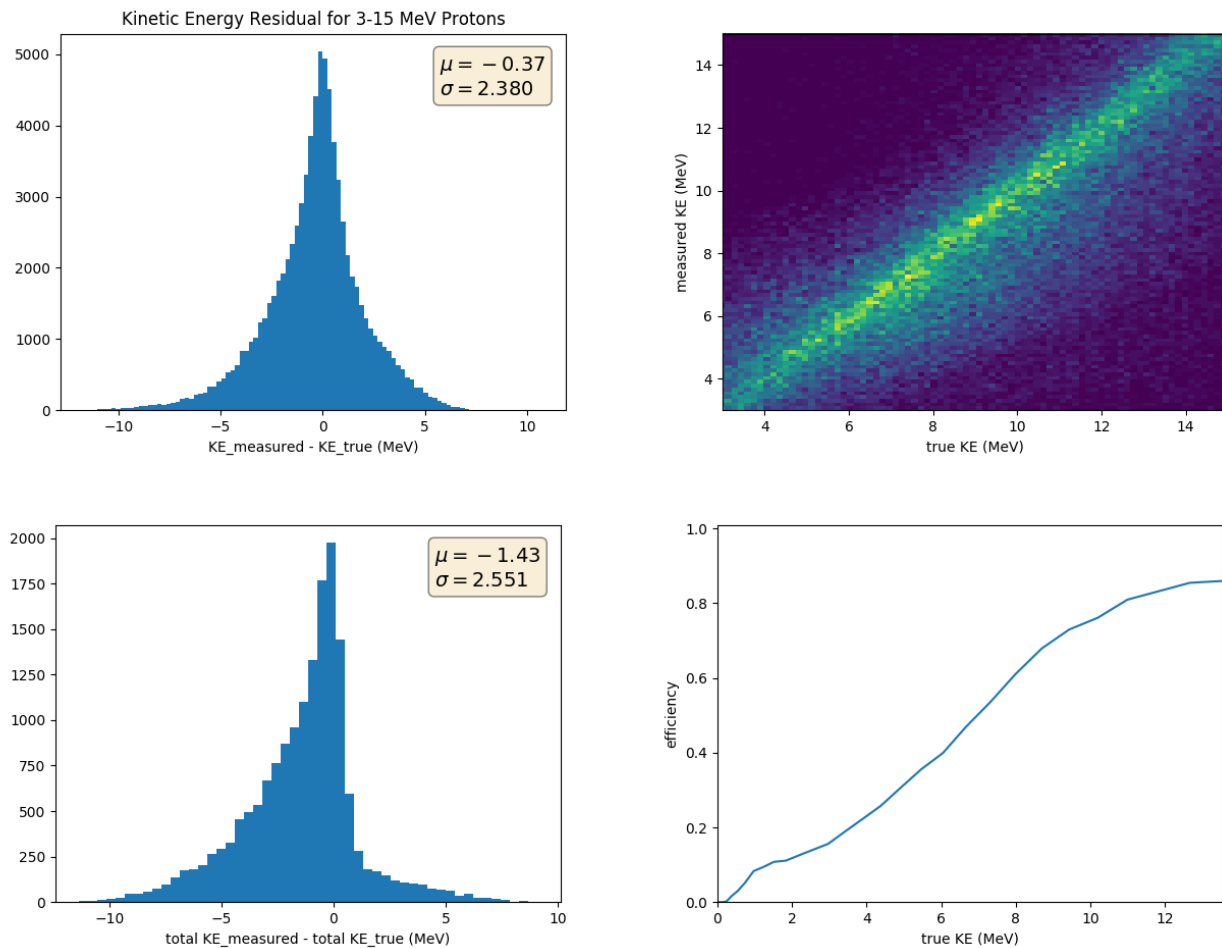


Figure 1.31: (Top left) Kinetic energy residual, (Top right) measured KE vs. true KE, and (Bottom right) reconstruction efficiency for individual protons with 3-15 MeV KE in the test set. (Bottom left) Residual of the total kinetic energy of all protons in each event in the test sample.

1 that this material provides a good compromise between calorimeter compactness, energy resolution,
 2 tion, and angular resolution. However, the choice of absorber material is still under study. The
 3 choice of granularity, scintillator thickness, and the arrangement of tiles and strips is still under
 4 optimization in order to reduce the number of readout channels while keeping the calorimeter
 5 performance. Two alternative scenarios are shown below: scenario B (black curve) has a different
 6 arrangement of the tile and strip layers, and scenario C (blue curve) has thinner absorbers in the
 7 front layers. Digitization effects are accounted for by introducing an energy threshold of 0.25 MIPs
 8 (~ 200 keV) for each detector cell/strip, a Gaussian smearing of 0.1 MeV for the electronic noise,
 9 SiPM saturation effects, and single photon statistics.

10 **Energy Resolution** The energy resolution is determined by fitting the visible energy with a
 11 Gaussian. Converted photons are rejected based on Monte-Carlo information. A fit function of
 12 the form $\frac{\sigma_E}{E} = \frac{A}{\sqrt{E}} \oplus \frac{B}{E} \oplus C$ is used, where A denotes the stochastic term, B the noise term, C
 13 the constant term, and E is in GeV. Figure 1.32 shows the energy resolution as a function of the
 14 photon energy. For scenario A, shown in red, the energy resolution is around $\frac{6.7\%}{\sqrt{E}}$. With further
 15 optimization, it is believed that an energy resolution of (or below) $\frac{6\%}{\sqrt{E}}$ is achievable. It should be
 16 noted that due to the lack of non-uniformities, dead cells, and other effects in the simulation, the
 17 energy resolution is slightly optimistic.

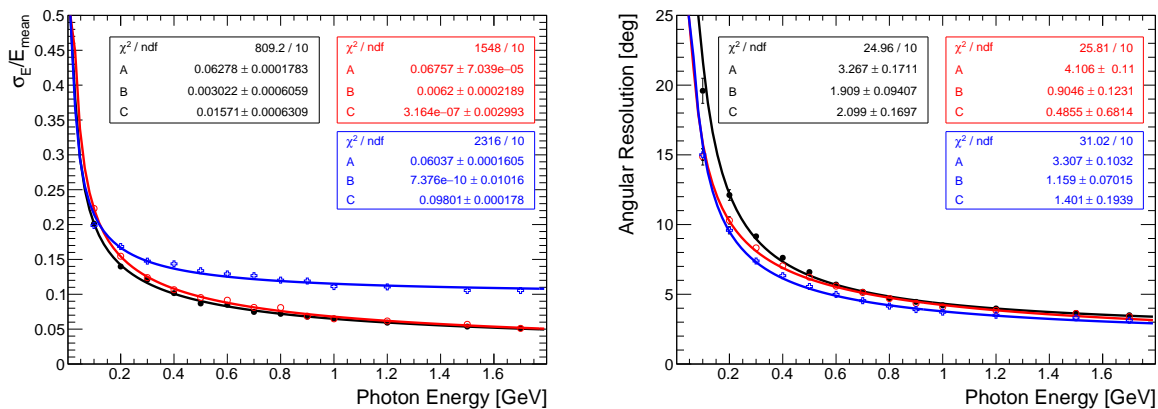


Figure 1.32: Left: energy resolution in the barrel as a function of the photon energy for three ECAL scenarios. The energy resolution is determined by a Gaussian fit to the visible energy. Right: the angular resolution in the barrel as a function of the photon energy for the three ECAL scenarios. The angular resolution is determined by a Gaussian fit to the 68% quantile distribution. For both figures, the scenario A is shown by the red curve, scenario B by the black curve and scenario C by the blue curve. The fit function is of the form $\frac{\sigma_E}{E} = \frac{A}{\sqrt{E}} \oplus \frac{B}{E} \oplus C$.

18 **Angular Resolution** The angular resolution of the calorimeter has been determined using a principal
 19 component analysis (PCA) of all reconstructed calorimeter hits. The direction is taken as
 20 the first eigenvector (main axis) of all the reconstructed hits. The angular resolution is deter-
 21 mined by taking the 68% quantile of the reconstructed angle distribution and fitting a Gaussian
 22 distribution. The mean of the Gaussian is taken as the angular resolution and the error as its
 23 variance. Figure 1.32 shows the angular resolution as a function of the photon energy. In scenario
 24 A, shown in red, an angular resolution of $\frac{3.85^\circ}{\sqrt{E}} \oplus 2.12^\circ$ can be achieved. This can potentially be

1 further improved with a different arrangement of the tile and strip layers, an optimization of the
2 absorber thickness, and an improved reconstruction method. However, the requirements will be
3 further refined and will impact the detector optimization. The angular resolution is mainly driven
4 by the energy deposits in the first layers of the ECAL. Using an absorber with a large X_0 creates
5 an elongated shower that helps in determining the direction of the shower. In general, high granu-
6 larity leads to a better angular resolution, however, studies have shown that there is no additional
7 benefit to having cell sizes below $2 \times 2 \text{ cm}^2$ [62].

8 **Neutron detection** The ECAL is sensitive to neutrons due to the scintillator containing hydro-
9 gen. Previous simulation studies showed that a detection efficiency above 60% can be achieved
10 for neutron energies greater than 50 MeV. However, the energy measurement is not very accurate
11 (around 50-60% below 600 MeV) [62]. Other methods of detection such as time of flight (ToF)
12 could be used to improve the neutron energy measurement by measuring precisely the hit time of
13 the neutron and its travel distance in the calorimeter. This is currently under study.

14 **π^0 reconstruction** For identification of neutral pions, both the energy and angular resolution
15 are relevant. In an initial study, the position of the neutral pion is determined by using a χ^2 -
16 minimization procedure taking into account the reconstructed energy of the two photons and the
17 reconstructed direction of the photon showers [62]. The location of the decay vertex of the neutral
18 pion can be determined with an accuracy between 10 cm to 40 cm, depending on the distance
19 from the downstream calorimeter and the π^0 kinetic energy. This is sufficient to associate the π^0
20 to an interaction in the HPgTPC, since the gas will have less than one neutrino interaction per
21 beam spill. The pointing accuracy to the pion decay vertex may be further improved by a more
22 sophisticated analysis technique and by using precision timing information, and is a subject of
23 current study.

24 1.3.4 DUNE-PRISM program

25 The goals of the off-axis measurements are twofold:

- 26 • **To identify problems in the cross section modeling.** By comparing ND data to MC at
27 many off-axis locations with different energy spectra, the neutrino interaction model will be
28 more tightly constrained than it would be with only on-axis measurements, and the potential
29 for biases in the measured oscillation parameters can be identified, i.e. the off-axis data
30 might be sensitive to mismodelings that are degenerate or indeterminate with only on-axis
31 measurements.
- 32 • **To overcome problems in the cross section modeling.** The most important novel feature of
33 a DUNE-PRISM detector is that measurements at different off-axis positions can be linearly
34 combined to determine any set of observables for any user-defined neutrino energy spectrum.
35 In particular, it is possible to predict the expected distribution of an observable, such as the
36 reconstructed neutrino energy, for a neutrino flux with oscillations using linear combinations

1 of ND off-axis spectra. This will greatly reduce the dependence on neutrino interaction
 2 modeling within the oscillation analysis.

3 1.3.4.1 Impact of Cross Section Modeling on Neutrino Oscillations

4 One strategy to understand the potential impact of using imperfect neutrino interaction models
 5 is to extract oscillation parameters from a “fake” data set that is different from the model used
 6 in the analysis. This fake data set represents a reality that includes effects unknown to or not
 7 accounted for properly by the model used in the analysis to fit the data. In this way, it is possible
 8 to understand potential biases in the measured oscillation parameter values extracted from a full
 9 near+far detector fit due to the use of an incorrect cross section model in the fit.

10 The fake data set considered here assumes that 20% of the kinetic energy that the interaction
 11 model originally assigned to protons was instead carried away by neutrons. The resulting model
 12 is then further modified by adjusting the differential cross section in proton energy as a function
 13 of true neutrino energy until the measured kinematic distributions in the on-axis ND match the
 14 prediction from the default model. This procedure is similar to actions that are routinely taken in
 15 actual neutrino oscillation experiments to resolve discrepancies between ND data and the Monte
 16 Carlo simulation. There are many potential modifications to the cross section model that can be
 17 chosen to resolve such disagreements. Incorrect choices can lead to biased oscillation parameter
 18 measurements because the associated incorrect particle composition and cross section model can
 19 lead to an incorrect relation between reconstructed and true energy.

20 The resulting fake data is analyzed as though it were the actual data taken by the experiment.
 21 The ND and FD data are fit simultaneously to constrain nuisance parameters in the flux and cross
 22 section models, and to extract the measured value of the neutrino oscillation parameters. The
 23 results of this fit are shown in Figure 1.33. The fit to the fake data shows a clear bias in the
 24 measured oscillation parameter values that lie outside the 95% confidence limit contours.

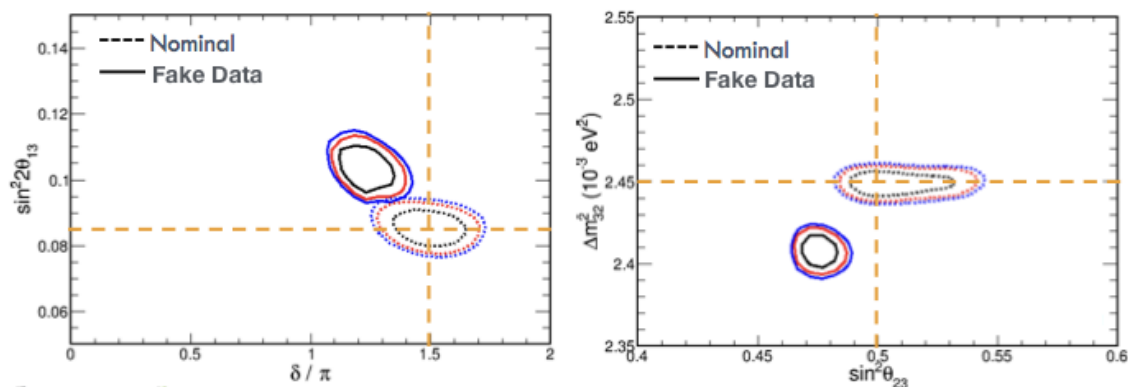


Figure 1.33: The results of a full 4-flavor near+far oscillation fit are shown for a fit to the nominal MC (dashed) and a fit to the fake dataset (solid). The true values of the oscillation parameters in each of the datasets are indicated by the dashed yellow lines. Clear biases can be seen in all oscillation parameters that are well outside the 1σ (black), 2σ (red), and 3σ (blue) contours.

25 A comparison of the fake data and the nominal Monte Carlo reconstruction energy distributions

1 is shown in Figure 1.34. In the on-axis location, good agreement is seen, as was intended in the
 2 construction of the fake data samples. This good agreement is assured since the model is tuned
 3 to the on-axis data. Conversely, clear disagreement is seen between these samples when moving
 4 off-axis. As the off-axis location is varied, this comparison can be made across a wide range of
 5 neutrino energy distributions.

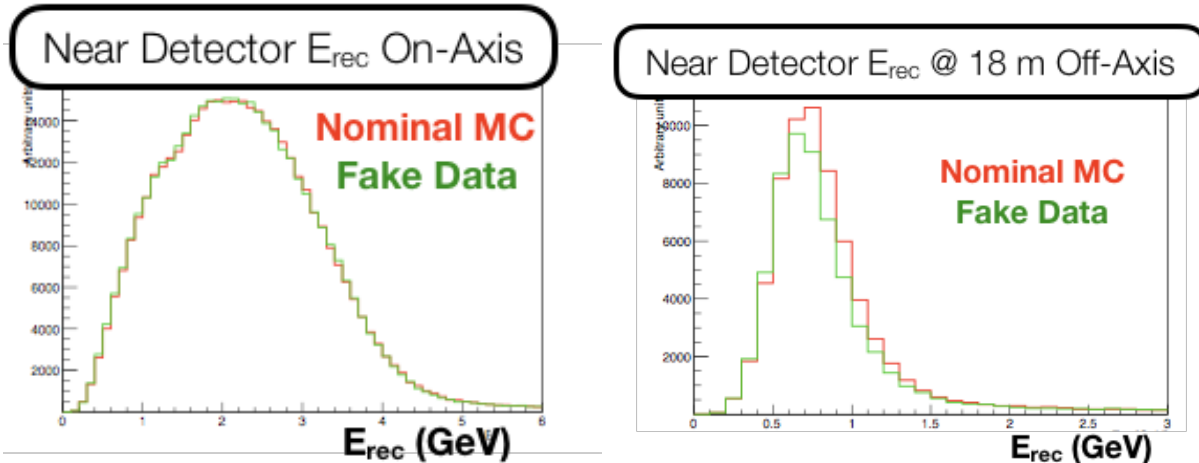


Figure 1.34: A comparison between the fake data (green) and nominal Monte Carlo (red) reconstructed neutrino energy distributions are shown for the on-axis ND location (left) and a position 18 m off-axis (right).

6 1.3.4.2 DUNE-PRISM Linear Combination Analysis

7 In addition to identifying problems in cross section modeling, DUNE-PRISM measurements provide
 8 a mechanism for creating FD predictions directly from the ND data that is largely independent
 9 of neutrino interaction modeling. By constructing linear combinations of measurements taken
 10 under exposure to different neutrino fluxes, it is possible to determine the distribution of any
 11 observable (e.g. reconstructed neutrino energy) for a different neutrino flux of interest. This
 12 means, for example, from the ND data alone it is possible to create a distribution of the expected
 13 reconstructed neutrino energy distribution at the FD in the event of oscillations with a specific set
 14 of parameters. This distribution, created using this data-driven technique, can then be compared
 15 to that seen in the FD with a reduced dependence on the flux and neutrino interaction models
 16 and their associated uncertainties.

17 A few example fits of the off-axis ND muon neutrino spectra to an oscillated FD muon neutrino
 18 energy spectrum are shown in Figure 1.35. Good agreement is seen near the first and second
 19 oscillation maxima at 2.5 GeV and 0.7 GeV, respectively. This technique can also be applied to
 20 match the off-axis muon neutrino spectra to the ND intrinsic electron neutrino spectrum, in order
 21 to make a precise measurement of $\sigma(\nu_e)/\sigma(\nu_\mu)$ with a common flux, or to the FD oscillated electron
 22 neutrino energy spectra for the measurement of δ_{CP} .

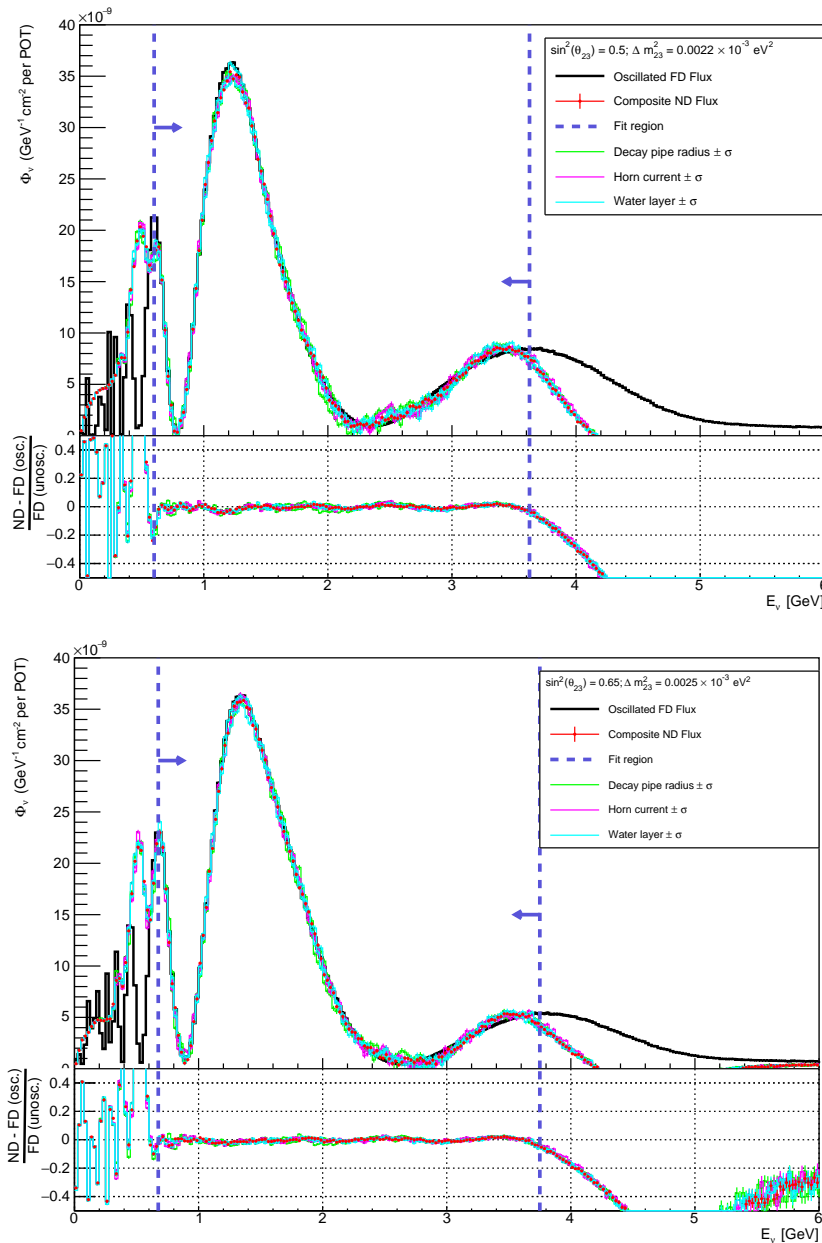


Figure 1.35: Linear combinations of off-axis fluxes giving far-detector oscillated spectra for a range of oscillation parameters. The FD oscillated flux is shown in black, the target flux is shown in green, and the linearly combined flux obtained with the nominal beam MC is shown in red. Systematic effects due to 1σ variations of the decay pipe radius (green), horn current (magenta) and horn cooling water layer thickness (teal) are shown.

1.4 Fixed on-axis component of the DUNE ND

1.4.1 Three-Dimensional Projection Scintillator Tracker Spectrometer

The 3D projection scintillator tracker spectrometer (3DST-S) consists of an active target core of scintillator called the 3DST surrounded by TPCs and an ECAL in a 0.4 T magnetic field. This system has three main goals in the context of the larger ND complex. First, the 3DST-S functions as an on-axis, high mass target and muon spectrometer that is capable of producing a statistically significant neutrino beam spectrum measurement in a short period. This dedicated, on-axis beam monitoring will be important in light of the movement of ArgonCube and the MPD within the context of the DUNE-PRISM program. Second, the 3DST-S will provide flux measurements with systematic errors and biases that differ from those in ArgonCube. This will be valuable input in debugging or developing confidence in the overall beam flux model. Third, the 3DST-S will produce measurements that are potentially useful for improving and increasing the level of confidence in the neutrino interaction model. In particular, the 3DST can measure neutrons on an event-by-event basis, including those at a lower neutron kinetic energy than those seen by the other components of the ND. (Current studies indicate both ArgonCube and the MPD detectors may be able to reconstruct neutrons on an event-by-event basis. Though still under investigation, current results show that backgrounds and the ability to associate neutron energy depositions to particular events limit this to higher kinetic energies.)

The 3DST is a fully active plastic scintillator detector made up of optically isolated 1 cm³ cubes [73]. The cubes are read out by wavelength shifting (WLS) fibers along 3 orthogonal axes providing three two-dimensional projections that yield effective three-dimensional reconstruction.

The 3DST is dense enough to provide a large statistics sample with reasonable containment of hadrons and photons from neutrino interactions. The high statistics and granularity of the 3DST will allow for timely beam monitoring, flux determination via different methods (with charge separation), and the study of many different neutrino interaction morphologies. The sub-ns timing resolution provides the capability to include neutrons in the event reconstruction via Time-of-Flight (TOF) with a reasonably high efficiency.

Neutron production plays a critical role in the interaction model since the near and far LAr detectors are largely blind to neutrons. Because the neutron content of neutrino and anti-neutrino interactions differ, the model for neutrons is particularly important for a CPV measurement. Preliminary studies show the 3DST is likely to be able to measure neutron spectrum to lower neutron KE (KE_n) than other detectors and pursue event-by-event analysis with fully reconstructed final state particles, including neutrons. GENIE and NuWro event generators both indicate neutron spectra for Ar and C are qualitatively similar. So, it is plausible that observations of neutrons produced by (anti)neutrino interactions on C can provide a higher level of confidence in the extrapolation of the Ar neutron model to lower KE_n than would otherwise be possible.

The 3DST uses the same technology as the SuperFGD detector that is being constructed for the T2K ND upgrade [74]. The two detectors are functionally identical, though somewhat different in size. The SuperFGD will be installed 2021 and will act effectively as a prototype for the larger

1 3DST in the DUNE ND.

2 1.4.1.1 Detector Configuration

3 The 3DST detector concept is shown in Figure 1.36. The scintillator composition is polystyrene
4 doped with 1.5% of paraterphenyl (PTP) and 0.01% of POPOP. After fabrication, the scintillator
5 surface of the cubes is etched with a chemical agent that results in the formation of a white,
6 reflecting polystyrene micropore deposit over the scintillator. Three orthogonal through holes
7 of 1.5 mm diameter are drilled in the cubes to accommodate WLS fibers. This novel geometry
8 provides full angular coverage for particle produced in neutrino interactions. The momentum
9 threshold for protons is about 300 MeV/c (if at least three hits are requested).

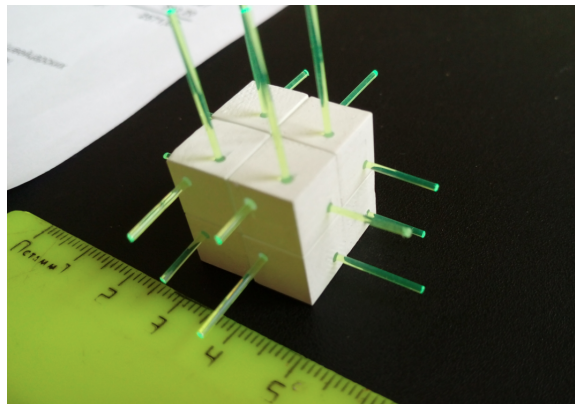


Figure 1.36: A few plastic scintillator cubes assembled with WLS fibers.

10 The 3DST and surrounding elements are shown in Figure 1.37. The size of the 3DST detector is
11 under discussion. Detectors of size $2.4 \times 2.4 \times 2.0 \text{ m}^3$, $3.0 \times 2.0 \times 2.0 \text{ m}^3$, and $2.0 \times 2.0 \times 2.0 \text{ m}^3$ have
12 been used in different studies. The primary considerations that drive the size are space, statistics,
13 and neutron containment.

14 The 3DST is surrounded by TPCs to measure the kinematics of charged particles produced but
15 not stopping in 3DST, and an ECAL to identify and reconstruct photons and electrons exiting
16 the 3DST. All the detectors will be immersed in a 0.4 T magnetic field provided by the magnet.
17 The TPCs are envisioned to be similar to those being constructed for the T2K ND280 detector
18 upgrade, described in [74]. They are approximately 80 cm thick and use a gas mixture of Ar-
19 CF₄-iC₄H₁₀ (95% - 3% - 2%). The ECAL is tentatively proposed to be a Pb-scintillator sampling
20 calorimeter. The downstream portion of the calorimeter will be thicker than the sides or the
21 upstream portions of the calorimeter in order to provide containment for electron showers in
22 neutrino-electron scattering events. The need for side and upstream TPCs and ECAL modules
23 is under study. The conceptual design of the magnet is highly preliminary. The initial concept
24 is that of a normal conducting magnet with iron flux return ~ 25 cm thick that is open on the
25 upstream and downstream sides. The outer dimension of the whole system, assuming side and
26 upstream TPCs and ECALs is approximately $5.8(\text{width}) \times 5.8(\text{height}) \times 4.6(\text{length}) \text{ m}^3$.

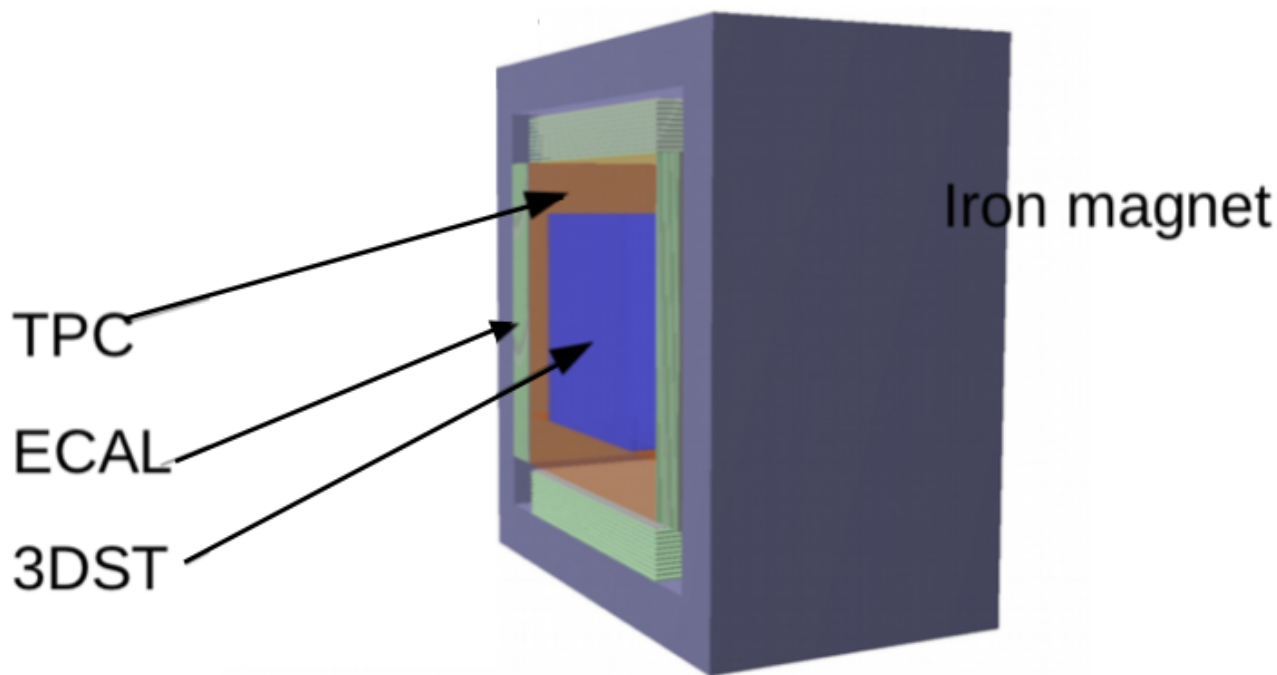


Figure 1.37: The 3DST-S detector configuration, including the 3DST (blue), TPCs (orange), ECAL (green), and the magnet (purple).

1.4.1.2 3DST Detector Performance

The performance of devices built on the 3DST concept have been tested in several test beams at CERN [75]. A small prototype of $5 \times 5 \times 5$ cubes collected data in the T10 test-beam area at CERN in 2017, with the goal of characterizing the response of the plastic scintillator cubes. The detector was instrumented with 75 WLS fibers (1 mm diameter Y11(200) Kuraray S-type of 1.3 m length). One end of the fiber was attached to a photosensor while the other end was covered by a reflective Al-based paint (Silvershine). The photosensors in the beam test were Hamamatsu MPPCs 12571-025C with a $1 \times 1 \text{ mm}^2$ active area and 1600 pixels. The data were collected with a 16-channel CAEN digitizer DT5742 with 5 GHz sampling rate and 12-bit resolution.

The average light yield was about 40 p.e./MIP in a single fiber, and the total light yield from two fibers in the same cube was measured on an event-by-event basis to be about 80 p.e., as expected. The light cross-talk probability between a cube fired by a charged particle and a neighbouring cube was studied. The light measured in the neighbouring cube was about 3.4% of the light collected from the fired cube. The timing resolution for a single fiber was measured to be ~ 0.95 ns. If the light of a cube is read out by two WLS fibers, the timing resolution becomes better than 0.7 ns and would improve further if the light collected by all the three WLS fibers is taken into account. In Figure 1.38 the light yield and the time spectra obtained from two fibers reading out the light in the same cube are shown.

In the summer of 2018, a new prototype made of 9,216 cubes with a size of $8 \times 24 \times 48 \text{ cm}^3$ collected data in the CERN T9 test-beam line. A different electronic readout was used, which was based on

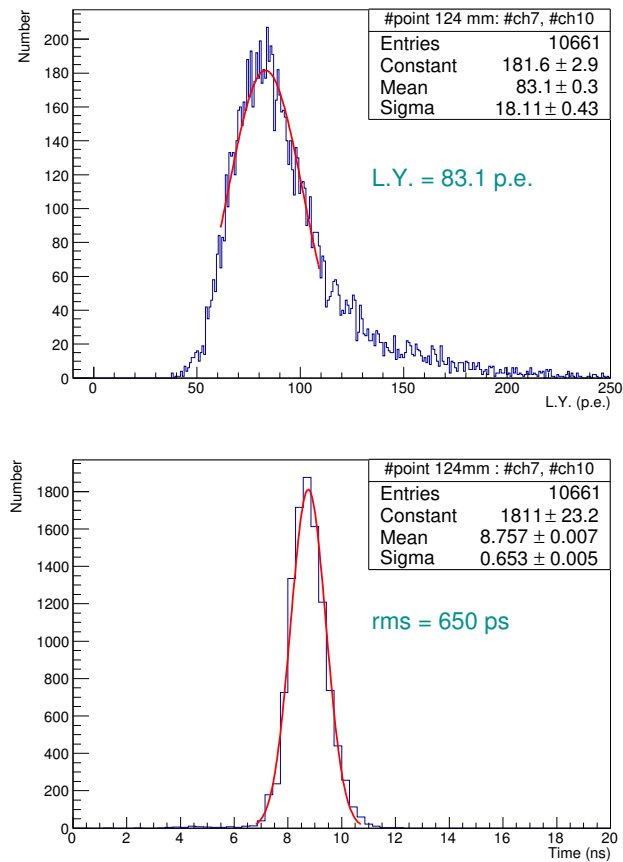


Figure 1.38: Charge and time spectra for a single cube. Charge signal is a sum from two fibers, the time is an average time between two fibers.

- 1 the CITIROC chip used in the Baby MIND experiment. Preliminary results confirmed the light
- 2 yield performances of the 2017 test-beam data. A more detailed analysis of the data is currently
- 3 ongoing. Some event displays are shown in Figure 1.39.

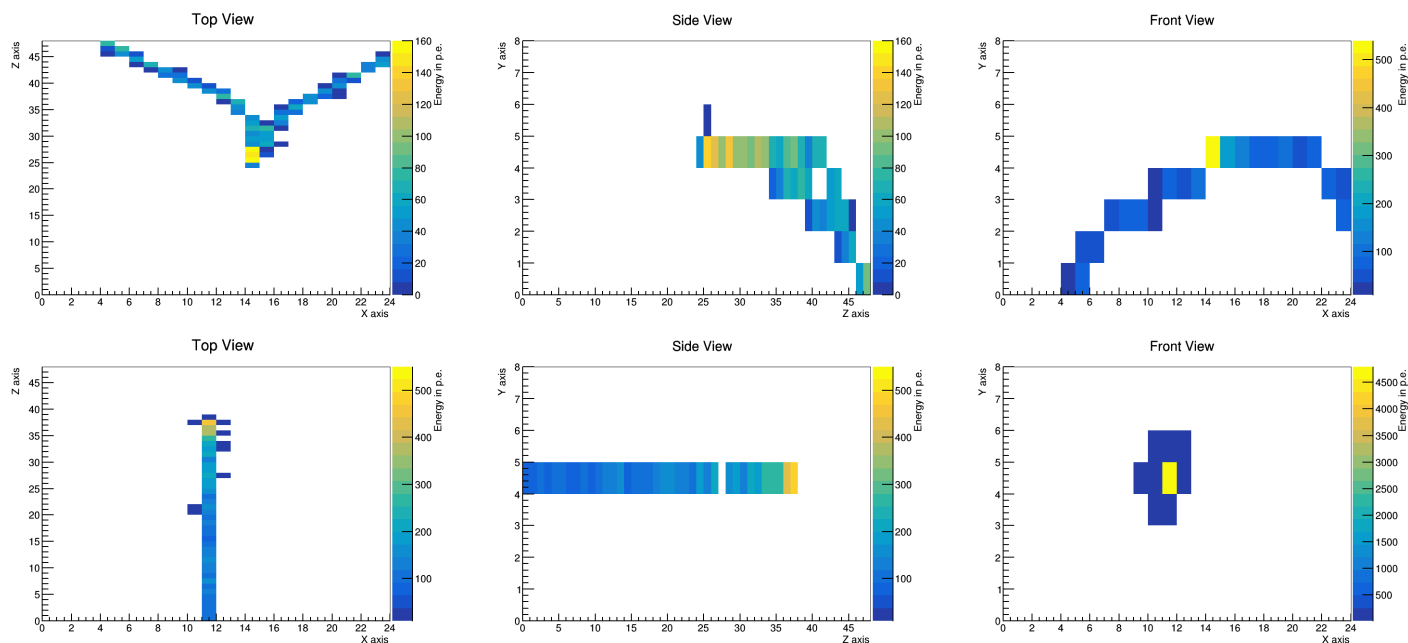


Figure 1.39: Event displays showing the three two-dimensional projections of energy from a photon conversion (top) and a stopping proton (bottom). From data collected at the 2018 test beams at the CERN T9 area.

4 1.4.1.3 Neutron Detection Performance

- 5 The MINER ν A experiment demonstrated the ability of measuring neutrons produced in neutrino
- 6 interactions with a plastic scintillator detector [76]. The 3DST should be able to do this far better
- 7 than MINER ν A because of its high granularity and exquisite timing resolution (both much better
- 8 than MINER ν A).

- 9 Neutron scattering can be seen clearly in 3DST simulations. Figure 1.40 shows an example of $\bar{\nu}_\mu$
- 10 CC single charged pion interaction. The neutron-induced energy deposition due to proton recoil
- 11 can be seen apart from the vertex region. Inspired by MINER ν A, recent studies (described below)
- 12 have shown that the 3DST can tag the presence of neutrons as well as determine the neutron energy
- 13 via time-of-flight. This capability is likely to be helpful for understanding/improving both neutrino
- 14 and antineutrino interaction models, and of potential use when faced with "unknown unknown"
- 15 sources of systematic uncertainties. The argon-based detectors in the ND complex are expected to
- 16 have some ability to detect neutrons, but studies indicate the sensitivity will be limited to regions
- 17 of relatively high neutron kinetic energy (due to backgrounds and event confusion that arise at
- 18 lower neutron kinetic energy where the considered event time window must be larger). The 3DST
- 19 will be sensitive to neutrons down to significantly lower kinetic energy. Though the measurements
- 20 are on carbon, generator truth studies show the neutron spectra in neutrino interactions on carbon

1 and argon are qualitatively similar. Thus the measurements at low neutron kinetic energy in the
 2 3DST can provide some degree of validation and additional confidence in the argon neutron model
 3 at energies lower than those accessible in the other detectors.

4 Since work to date has focused on establishing the ability and quality of the neutron detection
 5 in the 3DST (as shown below), detailed studies making use of the neutron reconstruction in
 6 simulated analyses are in an early stage. Simulations show the selection of $\bar{\nu}_\mu$ CCQE events with
 7 small missing transverse momentum, using a technique described in [77], yields a sample with
 8 a substantially improved energy resolution. This sample consists of events with relatively small
 9 nuclear effects useful for flux and studies of nuclear effects in neutrino interactions. It is also
 10 expected that neutron multiplicity can be used as an indication that multinucleon interactions
 11 or large FSI effects are present, which may be helpful for selecting events particularly useful for
 12 exploration of the interaction model.

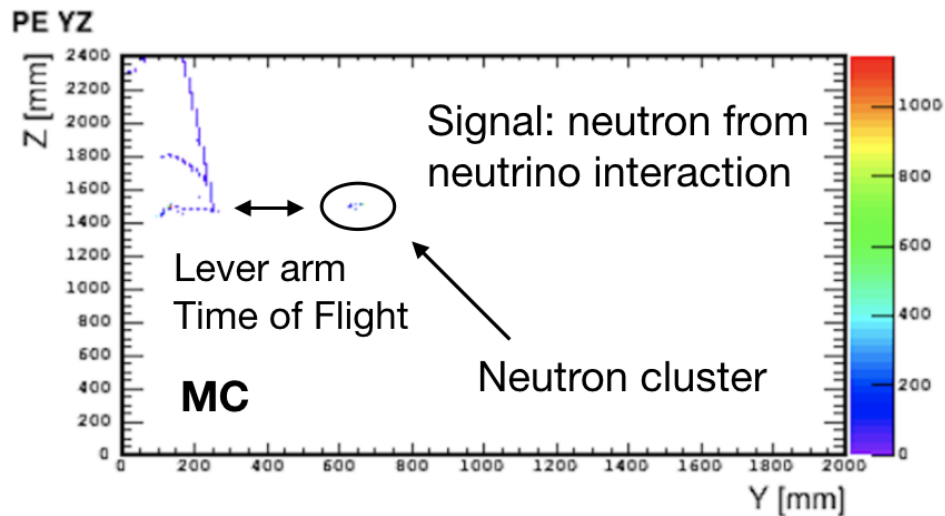


Figure 1.40: An example of the antineutrino interaction in a $2.4 \times 2.4 \times 2.0$ m³ 3DST. The number of photo-electrons (PE) is plotted. An isolated cluster of hits corresponds to a neutron indirect signature produced by the antineutrino interaction.

13 With a $2.4 \times 2.4 \times 2.0$ m³ 3DST detector, Figure 1.41 shows the reconstructed neutron energy resid-
 14 ual for 100 MeV kinetic energy neutron using time-of-flight with a lever arm (distance between
 15 neutron hit and neutrino vertex) larger than 0.5 m and smaller than 1 m. This study was con-
 16 ducted with a neutron particle gun simulation. The tail is due to both the timing resolution as
 17 well as the mis-reconstructed neutron flight distance due to non-visible interactions like elastic
 18 scattering with Carbon. The neutron energy resolution is about 18%.

19

20 Neutrons produced by neutrino interactions happening outside the 3DST fiducial volume (out-FV),
 21 such as in the ECAL, Magnet, front detector, and rock can act as a background to the neutron
 22 signal from neutrino interactions. A simulation study was performed to understand the significance
 23 of background. In this study, the 3DST-S was placed in an underground alcove and significant dead
 24 material was placed upstream. The FV was taken to be an inner core of $1.0 \times 1.0 \times 1.0$ m³
 25 of scintillator inside a 3DST of size $2.0 \times 2.0 \times 2.0$ m³. Neutrino beam spills of separation 1.3 s and

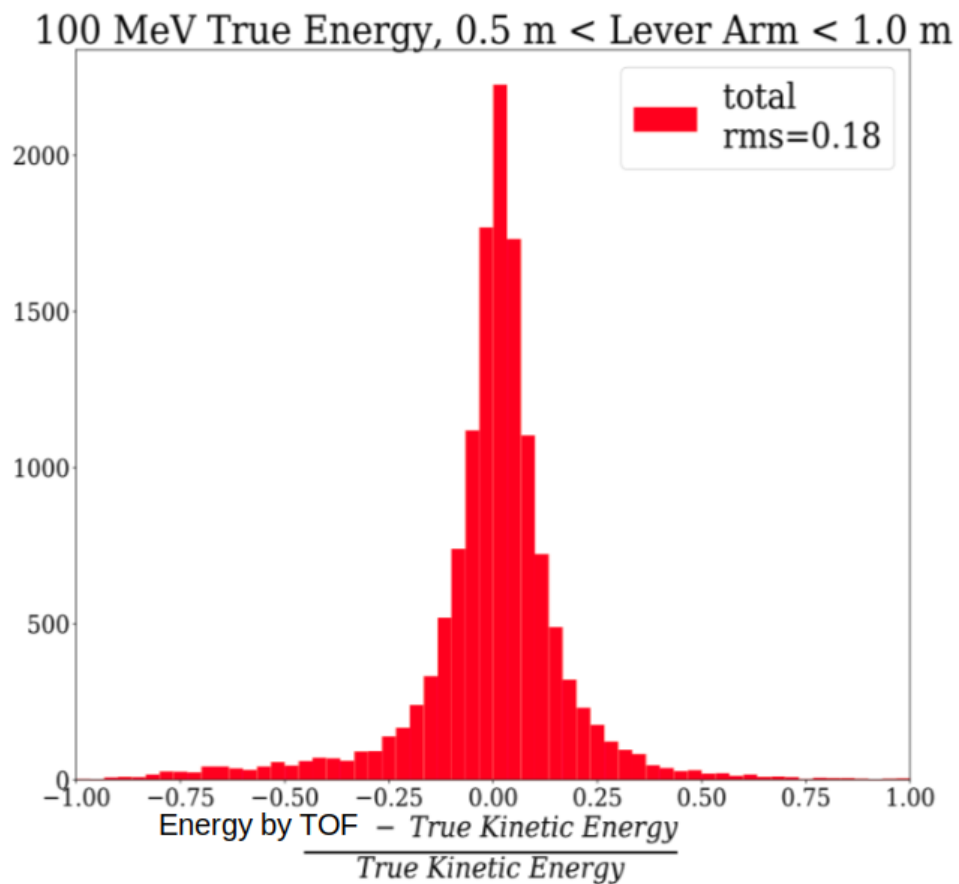


Figure 1.41: Reconstructed neutron energy residual with lever arm larger than 0.5 m and smaller than 1 m for 100 MeV for a $2.4 \times 2.4 \times 2.0$ m³ 3DST detector.

1 a uniform neutrino time distribution within each spill were used. For each neutrino interaction
 2 occurring inside the FV, the the earliest neutron-induced hit leaving an energy greater than 0.5
 3 MeV in one cube was recorded. This threshold is thought to be conservative for the 3DST system
 4 because of the large light yield expected. If that hit was from the neutrino interaction vertex, it was
 5 considered a signal neutron-induced hit. On the other hand, if that hit was created by a particle
 6 from outside the FV, it was considered a background neutron-induced hit. Figure 1.42 shows
 7 the time difference between the neutrino interaction vertex time (t_{vtx}) and the following earliest
 8 neutron-induced hit time ($t_{neutron}$). Note that a pure signal neutron sample can be obtained by
 9 cutting on $(t_{neutron} - t_{vtx})$.

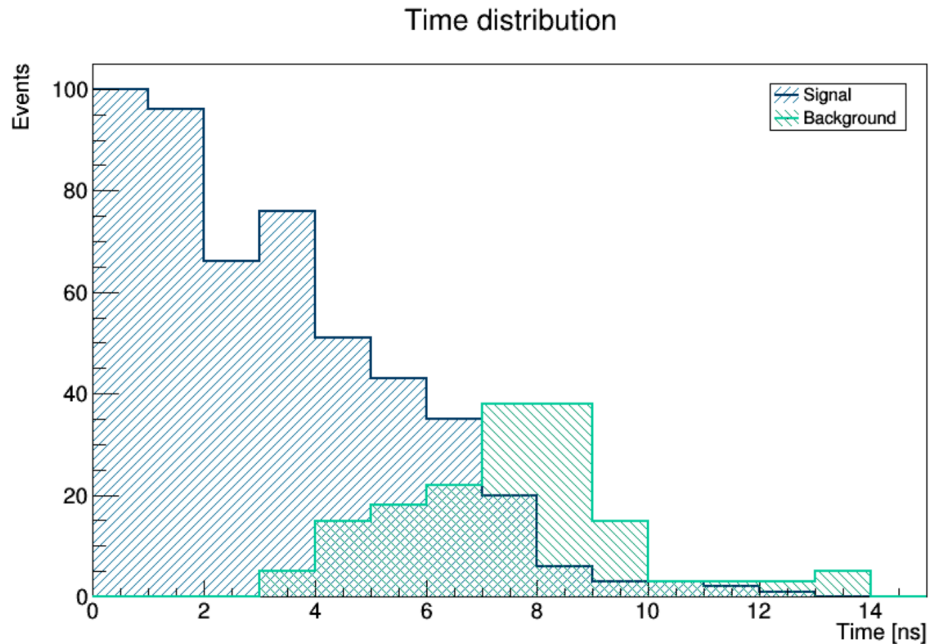


Figure 1.42: Time difference between the neutrino interaction vertex time inside the $1.0 \times 1.0 \times 1.0 \text{ m}^3$ fiducial volume core of the 3DST and the earliest neutron-induced hit time. The neutron-induced hit leaves at least 0.5 MeV in a single cube. The neutron-induced background hits arise from neutrons produced in neutrino interactions outside the FV.

10 It is likely to be possible to veto CC and NC interactions with pions in the materials surrounding
 11 the 3DST. Such a veto would reduce backgrounds from neutrons generated by these out-FV events.
 12 In this study, such a veto was not used. This will be investigated in the future.

13 To quantify the background, the purity is defined as the ratio of events where the first neutron-
 14 induced hit by time is from the signal vertex to all events which have a neutron-induced hit in the
 15 FV. Figure 1.43 shows the purity in time - lever arm space. Lines indicate regions populated by
 16 neutrons with different kinetic energies.

17 The reconstructed energy resolution in the same (time, lever arm) 2D space was studied. For
 18 this work, the time was smeared by 0.58 ns, corresponding to a per fiber time resolution of 1 ns
 19 (the documented performance in the CERN test beam is 0.9 ns). Though higher light yield can
 20 help improve the time resolution, this effect has not been taken into account. Figure 1.44 shows
 21 the reconstructed-by-ToF neutron energy resolution. In general, $\sim 20\%$ energy resolution can be
 22 reached with most of the lever arm and time windows, in the region selected by the background cut.

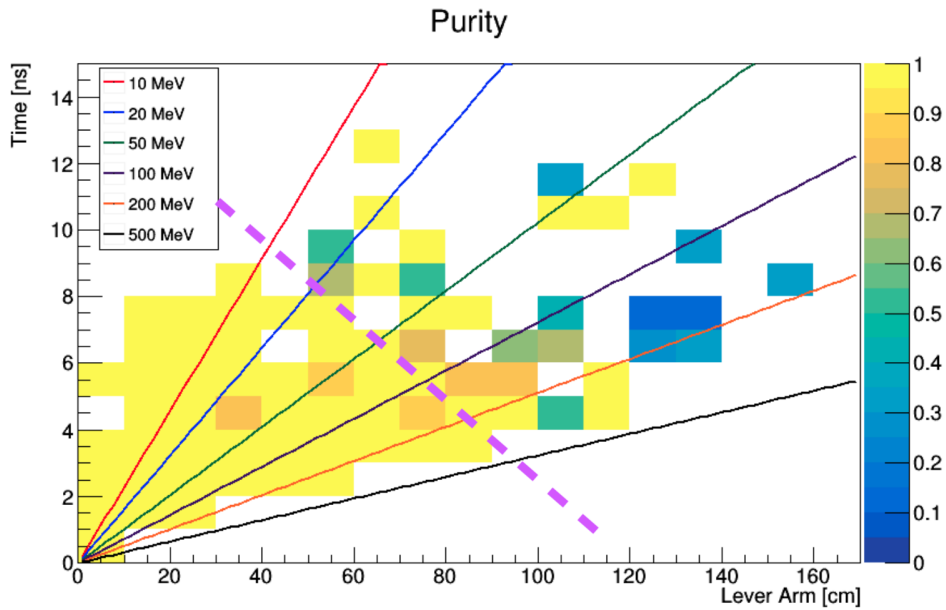


Figure 1.43: Purity of the neutron-induced hit in the (time, lever arm) space. The dashed line corresponds to the cut required to select an almost 100% pure sample of signal neutrons. The solid lines are theoretical curves for neutrons with different kinetic energies. Note that this study was performed with a total volume of $2 \times 2 \times 2 \text{ m}^3$. See text for details.

1

2 1.4.1.4 Expected Statistics

3 The default size of the 3DST is defined to be $2.4 \times 2.4 \times 2.0 \text{ m}^3$. This gives a total target mass for
 4 the 3DST of 12 metric tons. Implementing a generic veto region around each side of the detector
 5 of 10 cm, gives a fiducial mass of 8.7 tons. Table 1.4 gives the number of events expected per year
 6 in the fiducial volume of such a 3DST detector. The numbers given in the table are assuming the
 7 80 GeV, 3 horn, optimized LBNF beam flux and 1.46×10^{21} POT/year.

Table 1.4: This table summarizes the projected event rates per year for a $2.4 \times 2.4 \times 2.0 \text{ m}^3$ 3DST detector, assuming the 80 GeV, three horn, optimized LBNF beam. A 10 cm veto region at each side was required.

Channel	ν mode	$\bar{\nu}$ mode
ν_μ CC inclusive	13.6×10^6	5.1×10^6
CCQE	2.9×10^6	1.6×10^6
CC π^0 inclusive	3.8×10^6	0.97×10^6
NC total	4.9×10^6	2.1×10^6
ν_μ - e^- scattering	1067	1008
ν_e CC inclusive	2.5×10^5	0.56×10^5

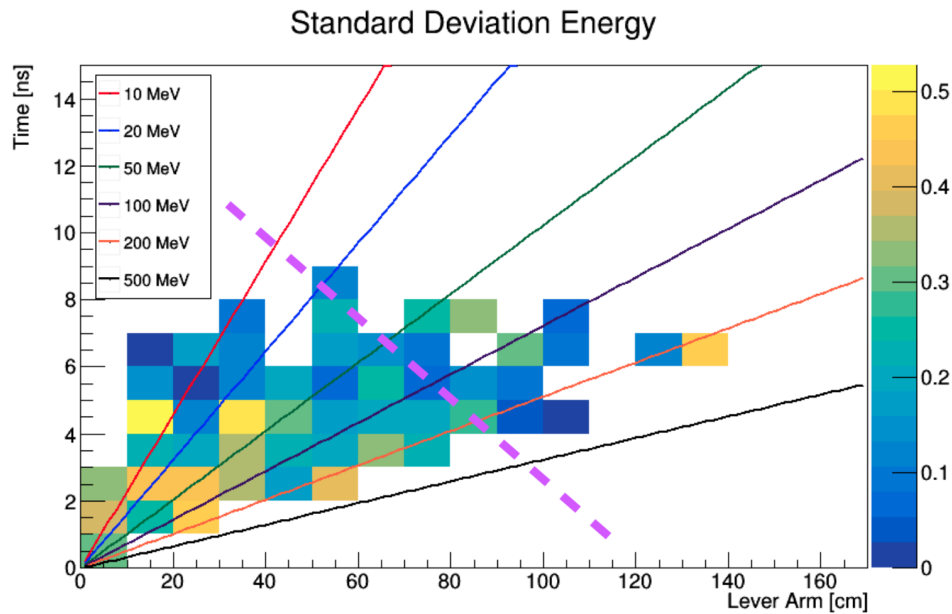


Figure 1.44: Energy resolution of the neutron candidates in the (time, lever arm) space. The dashed line corresponds to the cut required to select an almost 100% pure sample of signal neutrons. The solid lines are theoretical curves for neutrons with different kinetic energies. Note that this study was performed with a total volume of $2.0 \times 2.0 \times 2.0 \text{ m}^3$. See text for details.

1.5 ND Hall and Construction

Figure 1.45 shows the current design of the underground hall as required for the ND construction concept. The underground hall must house the detector components and allow for the required movement. The layout shows the space required for the detector, safety, and egress. This is work in progress.

The overall construction method places requirements on the conventional facilities. The primary access shaft is large enough for lowering the pressure vessel and the magnet coils. The LAr cryostat is shown in its construction position near the main shaft. The multipurpose detector and the LAr detector are also shown in the on-axis position. Since the 3DST detector does not need to move for DUNE-PRISM, it is shown in a dedicated alcove downstream of the LAr and multipurpose detectors.

The overall method of detector construction must be consistent with the construction concepts of each of the elements as outlined in previous sections. The construction method must also allow for parallel activities on major components and reduce demand on individual facilities. The underground hall will be the last facility to be completed. Therefore, insofar as possible, components must be constructed elsewhere and lowered as large assemblies.

The current assumptions for the overall construction involving the major components are listed below. Only the major components are considered, as they place the main constraints on the conventional facilities.

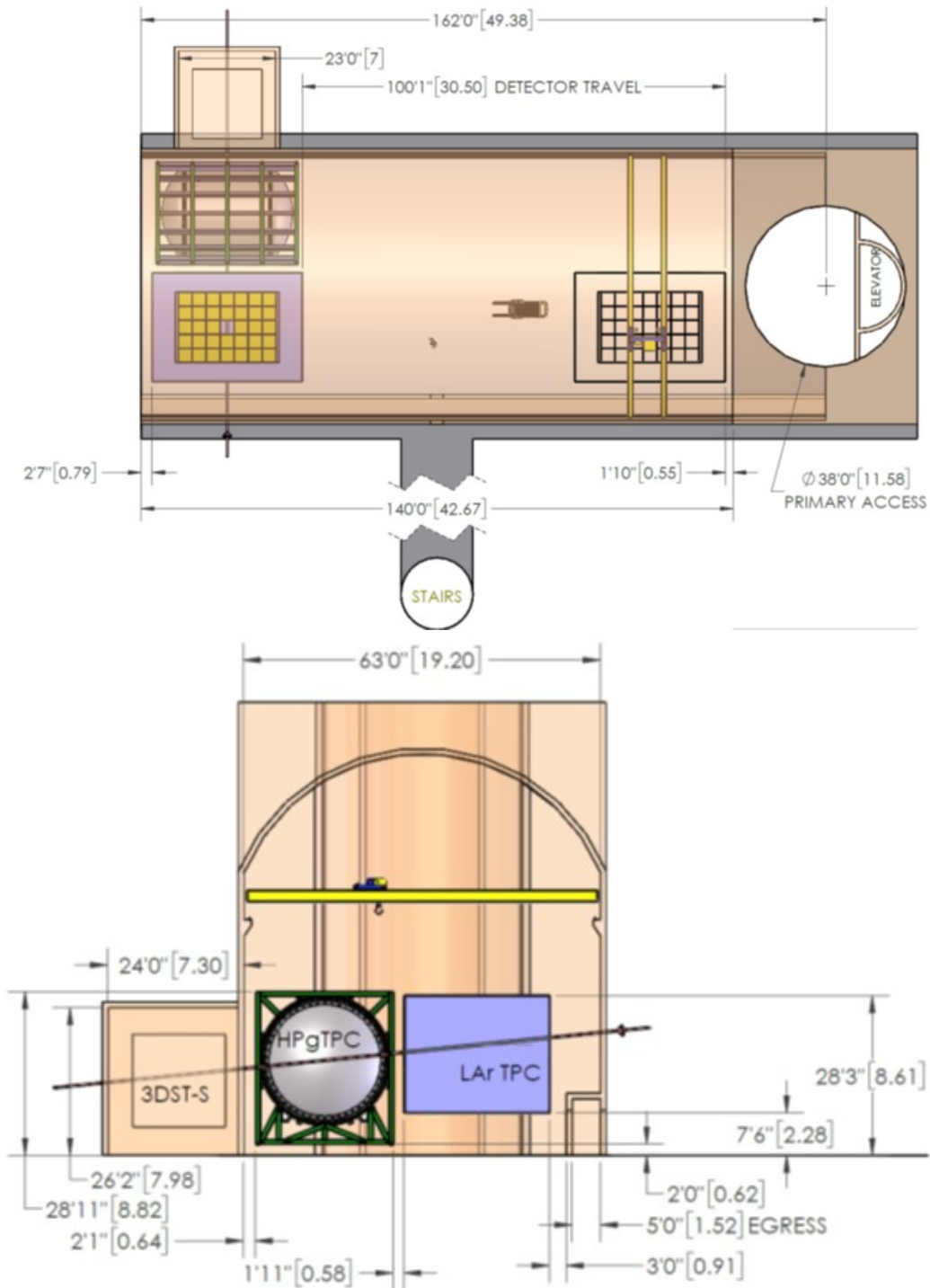


Figure 1.45: DUNE ND hall shown from above (left) and from the side transverse to the beam (right). The LArTPC, MPD, and 3DST detectors are shown in position on the beam axis in both drawings. On the left, the LArTPC is also shown in an off-axis position, suitable for module installation.

- 1 • The primary access shaft diameter must be sufficient to accommodate the lowering of the
2 pressure vessel and magnet coils separately. The coils and pressure vessel are constructed on
3 the surface and lowered. This allows for remote and parallel construction.
- 4 • Two transport frames and moving systems are built in the cavern. These frames will be used
5 to support and move the multipurpose detector and the LAr detector.
- 6 • Articulated carriers are built to carry services supporting the moving detectors.
- 7 • The components of the 3DST-S are constructed remotely and assembled in the underground
8 alcove cavern.
- 9 • The five magnet coils are constructed and integrated with cryostats remotely. The magnet
10 coils and cryostats are lowered into the ND hall and assembled together to form the magnet
11 system on one of the transport frames.
- 12 • The pressure vessel is fabricated remotely and transported to the ND hall surface building.
13 The TPC and some of the ECAL components are constructed in the surface building and
14 installed in the pressure vessel. The pressure vessel must be fabricated and certified by a
15 qualified fabricator.
- 16 • The LAr cryostat is constructed in the cavern on the second transport frame near the main
17 shaft. The construction of the cryostat starts by erecting and assembling the warm exoskele-
18 ton from pre-fabricated structural steel members. The warm membrane is then installed and
19 welded in situ. Insulation is then installed inside the warm vessel. The cold membrane is
20 the last component and is welded in situ to form the final containment vessel for the LAr. A
21 thin window is installed on the side facing the multipurpose detector.
- 22 • The LAr modules are constructed remotely and lowered down the shaft and installed in the
23 cryostat. The modules are inserted/extracted from the top of the cryostat using an overhead
24 lifting device.
- 25 • The LAr services are installed.
- 26 • The HPgTPC inside the pressure vessel is lowered in the hall and then mounted inside the
27 magnet system.
- 28 • The ECAL segments are lowered into the hall and mounted around the pressure vessel.
- 29 • Services are installed.

30 The basic requirement for DUNE-PRISM is that both the multipurpose detector and LAr detector
31 can move horizontally to a position off the beam axis. The direction of the motion is to one side
32 of the beam and the total motion is approximately 30.5 meter.

33 Though the multipurpose detector and the LAr detector will be moved together to different po-
34 sitions for operations, they will be able to move independently for engineering, construction, and

1 maintenance reasons. The exact method of movement is not determined at this time. However, it
2 is anticipated that tracks and rollers will be used in a fashion similar to what has been done for
3 other large particle physics detectors. The driving mechanism may be a rack and pinion drive, or a
4 similar system, which also allows for continuous motion. It is planned that the speed of movement
5 will allow for the entire motion to be completed in one 8-hour shift. This requires a speed of
6 approximately 6 cm/min. If it is desired that data can be taken during the movement, a slower
7 speed may be utilized. A speed of about 0.6 cm/min will result in the entire round trip to take
8 about one week.

9 Services for both the LAr and the multipurpose detector will need to be connected while moving or
10 disconnected and reconnected at intermediate positions. Ideally, no services will be disconnected
11 and reconnected and articulated service carriers will be used to maintain the connections during
12 movement. In the case of the LAr this presents particular challenges and will require flexible
13 conduits.

14 1.6 Meeting the Near Detector Requirements

15 As discussed in sections 1.2 and 1.1, the DUNE ND complex has many missions, and the compo-
16 nents of the ND address these missions in a complementary fashion. In this section we first discuss
17 the key overarching requirements driving the ND complex. We then discuss some thought exper-
18 iments and case studies that illustrate how different parts of the complex work together. These
19 case studies naturally suggest more detailed capabilities, performance statistics, and technical re-
20 quirements that we are in the process of tabulating. These are noted as [RT:SomeRequirement].
21 Most of these require additional studies before numerical values can be tabulated.

22 1.6.1 Overarching Requirements

- 23 • **O0: Predict the neutrino spectrum at the FD** The ND must provide a prediction for
24 the energy spectrum of ν_μ , $\bar{\nu}_\mu$, ν_e and $\bar{\nu}_e$ at the FD. The prediction must be provided as a
25 function of the oscillation parameters and systematic uncertainties must be small enough to
26 achieve the required CP coverage. This is the primary requirement of the DUNE ND.
- 27 • **O0.1: Measure interactions on argon** The ND must measure neutrino interactions on argon
28 to reduce uncertainties due to nuclear modelling. The ND must be able to determine the
29 neutrino flavor and measure the full kinematic range of the interactions that will be seen at
30 the FD.
- 31 • **O0.2: Measure the neutrino energy** The ND must be able to reconstruct the neutrino en-
32 ergy in CC events and control for any biases in energy scale or resolution, keeping them small
33 enough to achieve the required CP coverage. These measurements must also be transferable
34 to the FD.

- 1 • **O0.3: Constrain the cross-section model** The ND must measure neutrino cross-sections
2 in order to constrain the cross-section model used in the oscillation analysis. In particular,
3 cross-section mismodelling that causes incorrect FD predictions as a function of neutrino
4 flavor and true or reconstructed energy must be constrained well enough to achieve the
5 required CP coverage.
- 6 • **O0.4: Measure neutrino fluxes** The ND must measure neutrino fluxes as a function of
7 flavor and neutrino energy. This allows for neutrino cross-section measurements to be made
8 and constrains the beam model and the extrapolation of neutrino energy spectra from the
9 ND to the FD.
- 10 • **O0.5: Obtain data with different fluxes** The ND must measure neutrino interactions in
11 different beam fluxes (especially ones with different mean energies) to disentangle flux and
12 cross-sections, verify the beam model, and guard against systematic uncertainties on the
13 neutrino energy reconstruction.
- 14 • **O0.6: Monitor the neutrino beam** The ND must monitor the neutrino beam energy spec-
15 trum with sufficient statistics to be sensitive to intentional or accidental changes in the beam
16 on short timescales. The precise requirement will be informed by the run plan as well as
17 experience from previous experiments.

18 1.6.2 Flux Measurements

19 The three most straightforward flux measurements are described here. Other powerful but more
20 complex techniques are described in Section 1.8.

21 1.6.2.1 CC ν_μ and $\bar{\nu}_\mu$ Interactions

22 Each core component of the ND complex will have large data samples with which to constrain the
23 flux model: ArgonCube will accumulate 3.7×10^7 CC ν_μ events per year (on axis, less when off
24 axis); The 3DST will see 1.4×10^7 CC ν_μ per year on axis; and the HPgTPC will see 1.6×10^6
25 CC ν_μ events per year (on axis, less when off axis).

26 1.6.2.2 Intrinsic Electron (Anti)Neutrino Flux

27 The intrinsic ν_e and $\bar{\nu}_e$ component of the beam is discussed in section 1.8.0.4. This is an important
28 component to quantify in the ND since it represents an irreducible background for the appearance
29 oscillation analysis at the FD. The number of CC ν_e events expected in the ND per year (on axis)
30 are 6.7×10^5 , 2.5×10^5 , and 2.5×10^4 for ArgonCube, the 3DST, and the HPgTPC, respectively.
31 The primary background comes from NC π^0 production. The systematics are dominated by the
32 flux model and the interaction model (which enters in the background subtraction). In the past,

1 statistics has been a limitation. That will not be the case for DUNE ND. With large samples,
2 ArgonCube and the 3DST each will measure this component of the beam fairly quickly with
3 somewhat different systematic errors. Although accumulating statistics more slowly, the HPgTPC
4 will provide the best overall measurements the ν_e and $\bar{\nu}_e$ components of the beam. Photons mostly
5 do not convert in the gas. This eliminates the primary background to electron (anti-)neutrino
6 identification and the accompanying interaction model uncertainty in the background subtraction.
7 In addition, the HPgTPC has a magnetic field that allows for the sign separation of ν_e and $\bar{\nu}_e$.

8 **1.6.2.3 Neutrino-Electron Scattering**

9 This process and estimates of the ND performance measuring the flux using this technique is
10 discussed in sections 1.8.0.1 and 1.3.2.7. Measuring the flux using this process is a critical ND
11 mission because it is independent of nuclear effects. This is a rare process that can be used by both
12 ArgonCube and the 3DST-S components of the ND to measure the neutrino flux. The A of the
13 target nucleus is irrelevant for neutrino-electron scattering. The measurement places a premium
14 on the overall target mass (for statistics) as well as electron energy and angular resolutions. The
15 primary backgrounds are CC interactions of intrinsic beam ν_e and NC π^0 interactions. ArgonCube
16 will do this measurement well as indicated by the results of a study shown in Figure 1.15. Also,
17 that study shows a MINER ν A-like scintillator detector can do the measurement fairly well. The
18 3DST-S will have better angular resolution than the detector used in the study. Note that the
19 detector and reconstruction systematic errors will be different for the two very different detectors.
20 For such an important measurement, the duplication is good, and with many uncorrelated errors
21 it may be possible to combine the ArgonCube data set with that from the 3DST-S for a somewhat
22 improved constraint.

23 **1.6.3 Control of Systematic Errors**

24 **1.6.3.1 PRISM program**

25 The PRISM program of on- and off-axis measurements is an essential component of the ND complex
26 and is plays a key role in reducing systematic uncertainties[RT:PRISMProgram].

27 **1.6.3.2 Absolute muon energy scale**

28 The ND complex must provide knowledge of the absolute muon energy scale in the LAr ND with
29 sufficient accuracy to meet the oscillation physics requirement [RT:OscEmu] and the ancillary low-
30 nu capability [RT:LowNuEMu]. The complex will utilize MPD magnetic field survey information
31 [RT:BFieldSurvey], as well as Ks and Lambda decays to charged hadrons [RT:MPDKsRate] within
32 the MPD, to establish the charged particle momentum scale in the MPD. The measurement will be
33 translated to the LAr ND by tracking and momentum analyzing, in the MPD, cosmic ray muons go-
34 ing through the MPD into the LAr ND [RT:BackwardCosmicRate,CosmicTrigger,MaterialAssay].

- 1 The LAr ND will then measure the momentum of the muons (particularly stopping muons) by
- 2 range and multiple coulomb scattering (MCS) to establish its muon momentum scale and verify
- 3 the reconstruction and detector simulation.

- 4 The momentum scale will be translated to the FD by measuring and comparing the range and
- 5 MCS of stopping tracks in the data and the simulation.

6 **1.6.3.3 Hadronic energy scale; Hadronic response of the LArTPC**

7 The ND complex must calibrate the response (energy measured vs true energy) of the LAr ND and
8 FD to the hadronic system in neutrino interactions with an accuracy described in [RT:OscEHadUncertainty]
9 to meet the oscillation physics goals. The complex will utilize a simulation benchmarked against
10 the single particle response measured in protoDUNE as a baseline. It is expected that the response
11 will differ for different particle species and will not be constant as a function of energy. Additional
12 calibration is needed due to the imperfectly known particle spectra in neutrino interactions, and
13 the confusion that the LAr ND and FD will experience in identifying the composition of particles
14 in the final state. The ND complex will accomplish the calibration by first observing ν_μ -CC inter-
15 actions in the MPD. The MPD will identify the protons, charged pions and photons in the final
16 state[RT:MPDPID], and precisely reconstruct their energy via curvature in the magnetic field and,
17 for photons, energy deposition in the calorimeter[RT:MPDEReco]. The MPD will also observe,
18 statistically, the energy going into neutrons using time of flight[RT:MPDneutron].

19 The muon kinematics will then be used to select one or more semi-inclusive set of events occurring in
20 the LAr ND and MPD that have identical hadronic final states. A comparison of the reconstructed
21 hadronic energy in the LAr ND with the precisely measured MPD calibrates the response and
22 establishes the hadronic energy scale in LAr. It will also be possible to take the reconstructed
23 data events from the MPD and simulate them in the LAr to compare to the actual events in the
24 LAr. This will serve as a cross-check and as a way of studying the response. For example, one
25 could drop any primary neutron information from the MPD reconstructed events and simulate the
26 rest in LAr ND to compare to the LAr ND data as a way of understanding the amount of neutron
27 energy seen by the LAr ND.

28 **1.6.3.4 Beam Monitoring**

29 Previous experience with the NUMI, JPARC, and Booster neutrino beams has shown that DUNE
30 must prepare for changes in the beam that occur on timescales as short as a few days but that
31 are not readily detected by primary and secondary beam monitors. Sudden changes in the beam
32 have the potential of polluting the PRISM program if they occur when the LAr ND and MPD are
33 taking data at off-axis locations. The ND complex addresses this problem with the 3DSTS which
34 will continuously measure the ν_μ and $\bar{\nu}_\mu$ energy spectrum on-axis[RT:3DSTSMonitoring].

1.6.3.5 LArTPC energy resolution

The PRISM measurement program demands that the energy resolution of the LAr ND and FD be as similar as possible, and that any differences can be understood and corrected for³. The resolution can be constrained using a similar approach as followed for the hadronic response.

1.6.3.6 LAr ND acceptance

Neutrino flavor change due to oscillation occurs over a broad energy range which demands that the ND complex is able to achieve a broad energy coverage. The LAr ND has acceptance limits (relative to the FD) at large muon angle ($\theta > 25$ deg) and high hadronic energy due to its limited size. Moreover, there is a strongly energy dependent and deep acceptance dip for $1.0 < E_\mu < 1.5$ GeV/c due to dead material between the the active portions of the LAr ND and MPD TPC. The experiment must understand these features of the LAr ND event sample in order to use it to make predictions for the FD[RT:LArNDAcceptance].

To address this shortcoming the MPD must have a nearly 4π acceptance for charged tracks and photons, a high tracking efficiency, and the ability to reconstruct events with high E_ν [RT:MPDAcceptance]. A comparison of MPD and LAr ND event rates as a function of kinematic variables will verify the LAr ND acceptance model.

1.6.3.7 LAr detection thresholds

The energy threshold for detecting charged particles in LAr will be verified using the MPD in a way that is similar to the more general LAr ND acceptance study described above.

1.6.3.8 Managing pile-up

At the location of the ND, the LBNF beam is expected to generate approximately 1 neutrino interaction per 10 tons per 10 μ s spill. Each of the core detectors in the ND complex can eliminate most of the pile-up background with timing from optical elements. For ArgonCube, it is estimated that there will be approximately 0.5 neutrino interactions per spill per ArgonCube module. Prompt scintillation light from the argon detected in ArCLight detectors or something similar is used to provide t_o and separate events (both connected and disconnected parts of the event). For the HPgTPC, an estimate based on a significantly (15x) more massive magnet than the superconducting option presented in section 1.3.3.1.3 suggests there will be approximately 75 tracks per 10 μ s spill from interactions in surrounding materials passing through the TPC. The excellent <10 ns timing resolution of the ECAL surrounding the TPC will be used to provide a t_o and to define a time window for pileup rejection. Similarly the exquisite (sub-ns) time resolution of the cubes

³In fact this is true if the experiment only makes on-axis measurements.

- 1 in the 3DST can be used to generate a narrow window in time around neutrino interactions and
- 2 limit the potential for overlapping events.

3 **1.6.3.9 Energy Carried by Neutrons**

4 The ND complex must measure or otherwise account for the neutrino energy that goes into neutrons
5 because much of that energy ends up being undetectable by a LArTPC. The fraction of hadronic
6 energy carried by neutrons is sizable and it also differs between ν_μ and $\bar{\nu}_\mu$: 20% vs 40% for the
7 flux peak. To meet the oscillation physics goals these fractions must be known sufficiently well
8 [RT:OscNeutron].

9 The complex will approach this challenge in multiple ways. First, the PRISM measurement pro-
10 gram is required to map the relationship between true and reconstructed energy using inclusive
11 CC scattering[RT:PRISM]. These measurements are sensitive to cumulative mismodelings but
12 may have trouble pinning down their origin. To augment the PRISM program, the MPD and
13 3DSTS are required to measure the energy carried by neutrons using calorimetry and time of
14 flight[RT:MPDNeutron,3DSTSNeutron]. The MPD has the advantage of measuring neutron pro-
15 duction off of an Ar target, but the disadvantage of doing so with lower statistics than the 3DSTS.
16 The MPD measurement is also challenging due to the interaction rate in its calorimeter and the
17 composition of the calorimeter driven by its multi-role nature. The advantage of the 3DSTS is in
18 finer granularity and a better ability to reconstruct neutron energy on an event by event basis.

19 **1.6.3.10 ν_e -CC rate and π^0/γ background**

20 Neutral current events with a final state π^0 , or a single γ , are a potentially problematic background
21 to the ν_e and $\bar{\nu}_e$ appearance measurements. The ND complex will address this background by using
22 the MPD to precisely measure the rate of ν_e (and $\bar{\nu}_e$) charged current interactions as a function
23 of energy and other kinematic variables[RT:MPDNueRate,MPDNueEff]. The measurement will
24 be free of π^0 and γ backgrounds because photons do not convert in the MPD TPC. The LAr
25 ND will make a similar measurement. The efficiency and background of that measurement will
26 be benchmarked using the data from the MPD.

27 **1.7 Appendix: Beyond standard model opportunities with the** 28 **ND**

29 The unique combination of the high-intensity LBNF proton beams with a highly-capable precision
30 DUNE ND, and massive LArTPC far detector modules at a 1300 km baseline, enables a variety of
31 opportunities for Beyond the Standard Model (BSM) physics, either novel or with unprecedented
32 sensitivity. The ND plays an essential role in taking full advantage of the LBNF beam in most of
33 the BSM physics topics. Of the many BSM opportunities, we describe a handful of representative

1 topics and briefly summarize how DUNE can make leading contributions in this arena, taking
2 advantage of the capable ND.

3 **1.7.1 Search for low-mass dark matter**

4 Various cosmological and astrophysical observations strongly support the existence of dark matter
5 (DM) representing $\sim 27\%$ of the mass-energy of the universe, but its nature and potential non-
6 gravitational interactions with regular matter remain undetermined. The lack of evidence for
7 weakly interacting massive particles (WIMP) at direct detection and the LHC experiments has
8 resulted in a reconsideration of the WIMP paradigm. For instance, if dark matter has a mass
9 which is much lighter than the electroweak scale (e.g., below GeV level), it motivates theories for
10 dark matter candidates that interact with ordinary matter through a new vector portal mediator.
11 High flux neutrino beam experiments, such as DUNE, have been shown to provide coverage of
12 DM+mediator parameter space which cannot be covered by either direct detection or collider
13 experiments. In LBNF, low-mass dark matter may be produced through proton interactions in
14 the target, and can be detected in the ND through neutral-current-like interactions either with
15 electrons or nucleons in the detector material via elastic scattering. Since these experimental
16 signatures are virtually identical to those of neutrinos, neutrinos are a significant background that
17 can be suppressed using timing and kinematics of the final-state electron or nucleons in the ND.
18 Therefore, it is essential for the ND to be able to differentiate arrival time differences of the order
19 a few ns or smaller, which determines the reachable range of the dark matter, and to measure
20 precisely the kinematic parameters of the recoil electrons, such as the scattering angle and the
21 energy. These capabilities will enable DUNE's search for light dark matter to be competitive and
22 complementary to other experiments at mass range below 1 $\hat{\text{A}}\text{S}$ 2 GeV. In addition, a recent
23 study for off-axis data taking in the context of DUNE-PRISM [?] shows a significant improvement
24 in search sensitivity compared to the on-axis data taking, thanks to the improvement in signal to
25 background ratio resulting from the faster reduction of the neutrino flux than the dark matter.
26 The various running conditions for the different combinations of data taking with the maximum
27 off-axis range to 36 m or 24 m show little difference in the sensitivity. This is due primarily to the
28 fact that the shape of the neutrino flux as a function of the lateral position at this range changes
29 slowly compared to the position near the beam center. A sensitivity plot to reflect both the on-axis
30 and DUNE-PRISM scenario are presented in Volume 2, DUNE Physics, of this technical design
31 report (TDR).

32 **1.7.2 Sterile neutrino search**

33 Experimental results in tension with the three-neutrino-flavor paradigm, which may be interpreted
34 as mixing between the known active neutrinos and one or more sterile states, have led to a rich
35 and diverse program of searches for oscillations into sterile neutrinos. DUNE is sensitive over a
36 broad range of values of the sterile neutrino mass splitting by looking for disappearance of CC
37 and NC interactions over the long distance separating the near and far detectors, as well as over
38 the short baseline of the ND. The ND provides most of the sensitivity at values larger than 1 eV^2 ,
39 which where the LSND best-fit and the regions still allowed by fits to global data reside. The

1 combination of the intense LBNF beam, and a highly-capable ND provide DUNE with strong sen-
2 sitivity in those regions, specifically in probing sterile-driven electron neutrino appearance and/or
3 tau neutrino appearance (along with associated muon neutrino disappearance). The combination
4 of high-resolution ND components, capable of high-efficiency particle ID, combined with a precise
5 muon monitor system for the LBNF beam would further enhance this sensitivity.

6 **1.7.3 Neutrino tridents**

7 Neutrino trident production is a rare weak process in which a neutrino, scattering off the Coulomb
8 field of a heavy nucleus, generates a pair of charged leptons. The typical final state of a neutrino
9 trident interaction contains two leptons of opposite charge. Measurements of muonic neutrino
10 tridents were carried out at the CHARM-II, CCFR, and NuTeV experiments, and yielded results
11 consistent with Standard Model predictions, but those measurements leave ample room for po-
12 tential searches for New Physics. As an example, a class of models that modify the trident cross
13 section are those that contain an additional neutral gauge boson, Z'_0 , that couples to neutrinos and
14 charged leptons. This Z'_0 boson can be introduced by gauging an anomaly-free global symmetry
15 of the Standard Model, with a particular interesting case realized by gauging $L_\mu - L_\tau$. Such a Z'_0
16 is not very tightly constrained and could address the observed discrepancy between the Standard
17 Model prediction and measurements of the anomalous magnetic moment of the muon, $(g-2)_\mu$.
18 The DUNE ND offers an excellent environment to generate a sizable number of trident events
19 (100/year), offering very promising prospects to both improve the above measurements, and to
20 look for an excess of events above the Standard Model (SM) prediction, which would be an in-
21 dication of new physics. In particular, DUNE can potentially discover or rule out the complete
22 parameter space allowed for the Z'_0 to explain the $g-2$ anomaly. Another category of BSM Physics
23 models that can be probed through neutrino trident measurements are dark neutrino sectors. In
24 these scenarios, SM neutrinos mix with heavier SM singlet fermions (dark neutrinos) with their
25 own new interactions. Due to this mixing, neutrinos inherit some of this new interaction and may
26 up-scatter to dark neutrinos. These heavy states in turn decay back to SM fermions, giving rise
27 to trident signatures. These scenarios can explain the smallness of neutrino masses and possibly
28 the MiniBooNE low energy excess of events. Reconstruction of neutrino trident events would be
29 strongly enhanced by a gaseous argon component in the DUNE ND, and the inclusion of a mag-
30 netic field could dramatically improve background removal capability by providing sign-selection
31 of the opposite charge leptons in the final state of the trident interaction.

32 **1.7.4 Heavy Neutral Leptons**

33 The DUNE ND can be used to search topologies of rare event interactions and decays that originate
34 from very weakly-interacting long-lived particles, including heavy neutral leptons – right-handed
35 partners of the active neutrinos, vector, scalar, or axion portals to the hidden sector, and light
36 supersymmetric particles. The high intensity of the NuMI source and the capability of production
37 of charm mesons in the beam allow accessing a wide variety of lightweight long-lived, exotic,
38 particles. Competitive sensitivity is expected for the case of searches for decay-in-flight of sub-GeV
39 particles that are also candidates for dark matter, and may provide an explanation for leptogenesis

1 in the case of CPV indications. DUNE would probe the lighter particles of their hidden sector,
 2 which can only decay in SM particles in the form of pairs like e^+e^- , $\mu^+\mu^-$, $q\bar{q}$. The parameter space
 3 explored by the DUNE ND extends to the cosmologically relevant region that is complementary
 4 to the LCH heavy-mass dark-matter searches through missing energy and mono-jets. The DUNE
 5 ND can therefore compete with, and complement, the measurements to be carried out by the SHiP
 6 experiment, expected to be starting operations at CERN on a time scale similar to DUNE's, as
 7 well as extend or confirm results from searches presently being carried out at MicroBooNE, or in
 8 the near future with the SBN program.

9 1.7.5 Non-standard Neutrino Interactions

10 While the role of ND in this measurement is less significant, DUNE can also search for deviations
 11 from the PMNS neutrino mixing paradigm arising due to NSIs, in particular those occurring at
 12 neutrino production, which would leave subtle imprints in the neutrino flux. Sensitivity to these
 13 effects would require a very well characterized flux for it to be competitive with probes of the same
 14 phenomenon in coherent electron neutrino scattering experiments. Further, the more common
 15 search for NSI affecting neutrino propagation through the Earth benefits from constraints on cross
 16 section and flux provided by a highly-capable ND in the same way as the CPV probe would. If
 17 the DUNE data are consistent with standard oscillations for three massive neutrinos, interaction
 18 effects of order 0.1 GF can be ruled out at DUNE. DUNE could improve current constraints on
 19 $\epsilon_{\tau e}$ and $\epsilon_{\mu e}$ by a factor 2 to 5.

20 1.8 Appendix: Constraining the Flux

21 The DUNE FD will not measure the neutrino oscillation probability directly. Instead, it will
 22 measure the neutrino interaction rate for different neutrino flavors as a function of the reconstructed
 23 neutrino energy. It is useful to formalize the measurements that are performed in the near and far
 24 detector modules in the following equations:

$$\frac{dN_x^{FD}}{dE_{rec}}(E_{rec}) = \int \Phi_{\nu_\mu}^{FD}(E_\nu) P_{\nu_\mu \rightarrow x}(E_\nu) \sigma_x^{Ar}(E_\nu) T_x^{FD,Ar}(E_\nu, E_{rec}) dE_\nu \quad (1.2)$$

$$\frac{dN_x^{ND}}{dE_{rec}}(E_{rec}) = \int \Phi_x^{ND}(E_\nu) \sigma_x^m(E_\nu) T_x^{d,m}(E_\nu, E_{rec}) dE_\nu \quad (1.3)$$

$$(1.4)$$

25 with

- 26 • $x = \nu_e, \nu_\mu$
- 27 • $d = \text{detector index(ND,FD)}$

- 1 • $m =$ interaction target/material, (e.g., H, C, or Ar)
- 2 • $E_\nu =$ true neutrino energy
- 3 • $E_{rec} =$ reconstructed neutrino energy
- 4 • $T_x^{d,m}(E_\nu, E_{rec}) =$ true to reconstruction transfer function
- 5 • $\sigma_x^m(E_\nu) =$ neutrino interaction cross section
- 6 • $\Phi_x^d(E_\mu) =$ un-oscillated neutrino flux
- 7 • $\frac{dN_x^d}{dE_{rec}}(E_{rec}) =$ measured differential event rate per target (nucleus/electron)

8 There are equivalent formulae for antineutrinos. For simplicity, the instrumental backgrounds
 9 (wrongly selected events) and the intrinsic beam contaminations (ν_e interactions in case of the
 10 appearance measurement) have been ignored. But an important function of the ND is also to
 11 quantify and characterize those backgrounds.

It is not possible to constrain the FD neutrino flux directly, but the near-to-far flux ratio is believed to be tightly constrained by existing hadron production data and the beamline optics. As such Equation 1.2 can be rewritten as

$$\frac{dN_x^{FD}}{dE_{rec}}(E_{rec}) = \int \Phi_{\nu_\mu}^{ND}(E_\nu) R(E_\nu) P_{\nu_\mu \rightarrow x}(E_\nu) \sigma_x^{Ar}(E_\nu) T_x^{d,Ar}(E_\nu, E_{rec}) dE_\nu \quad (1.5)$$

with $R(E_\nu) = \frac{\Phi_{\nu_\mu}^{FD}(E_\nu)}{\Phi_{\nu_\mu}^{ND}(E_\nu)}$ taken from the beam simulation

It is not possible to measure only a near-to-far event ratio and extract the oscillation probability since many effects do not cancel trivially. This is due to the non-diagonal true-to-reconstruction matrix, which not only depends on the underlying differential cross section, but also on the detector used to measure a specific reaction.

$$\frac{dN_x^{FD}}{dE_{rec}}(E_{rec}) / \frac{dN_{\nu_\mu}^{ND}}{dE_{rec}}(E_{rec}) \neq R(E_\nu) P_{\nu_\mu \rightarrow x}(E_\nu) \frac{\sigma_x^{Ar}(E_\nu)}{\sigma_{\nu_\mu}^m(E_\nu)} \quad (1.6)$$

12 It is therefore important that the DUNE ND suite constrain as many components as possible.

13 While the near-to-far flux ratio is tightly constrained to the level of 1% to 2%, the same is not
 14 true for the absolute flux itself. T2K, using hadron production data obtained from a replica target,
 15 can constrain the absolute flux at the ND to 5% to 6% in the peak region and to around 10%
 16 in most of its energy range. The NuMI beam has been constrained to 8% using a suite of thin
 17 target hadron production data. The better the ND flux is known, the easier it is to constrain
 18 modeling uncertainties by measuring flux-integrated cross sections. Predicting the event rate at
 19 the FD to a few percent will require additional constraints to be placed with the ND or substantial
 20 improvements in our understanding of the hadron production and focusing uncertainties.

1 Several handles to constrain the flux are addressed below. Briefly they offer the following con-
 2 straints:

- 3 • The overall flux normalization and spectrum can be constrained by measuring neutrino scat-
 4 tering off of atomic electrons.
- 5 • The energy dependence (“shape”) of the ν_μ and $\bar{\nu}_\mu$ flux can be constrained using the “low- ν ”
 6 scattering process.
- 7 • The flux ratio $\bar{\nu}_\mu/\nu_\mu$ can be constrained using charged current coherent neutrino scattering.
- 8 • The ν_e/ν_μ flux ratio in the energy region where standard oscillations occur is well-constrained
 9 by the beam simulation. The experiment can also measure the ν_e/ν_μ interaction ratio and
 10 constrain the flux ratio using cross section universality.

11 1.8.0.1 Neutrino-Electron Elastic Scattering

12 Neutrino-electron scattering ($\nu e \rightarrow \nu e$) is a pure electroweak process with calculable cross section
 13 at tree level. The final state consists of a single electron, subject to the kinematic constraint

$$1 - \cos \theta = \frac{m_e(1 - y)}{E_e}, \quad (1.7)$$

14 where θ is the angle between the electron and incoming neutrino, E_e and m_e are the electron mass
 15 and total energy, respectively, and $y = T_e/E_\nu$ is the fraction of the neutrino energy transferred to
 16 the electron. For DUNE energies, $E_e \gg m_e$, and the angle θ is very small, such that $E_e\theta^2 < 2m_e$.

17 The overall flux normalization can be determined by counting $\nu e \rightarrow \nu e$ events. Events can be iden-
 18 tified by searching for a single electromagnetic shower with no other visible particles. Backgrounds
 19 from ν_e CC scattering can be rejected by looking for large energy deposits near the interaction
 20 vertex, which are evidence of nuclear breakup. Photon-induced showers from neutral-current π^0
 21 events can be distinguished from electrons by the energy profile at the start of the track. The
 22 dominant background is expected to be ν_e CC scattering at very low Q^2 , where final-state hadrons
 23 are below threshold, and $E_e\theta^2$ happens to be small. The background rate can be constrained with
 24 a control sample at higher $E_e\theta^2$, but the shape extrapolation to $E_e\theta^2 \rightarrow 0$ is uncertain at the 10%
 25 to 20% level.

26 For the DUNE flux, approximately 100 events per year per ton of fiducial mass are expected with
 27 electron energy above 0.5 GeV. For a LArTPC mass of 25 tons, this corresponds to 3300 events
 28 per year. The statistical uncertainty on the flux normalization from this technique is expected to
 29 be $\sim 1\%$. MINER ν A has achieved a systematic uncertainty just under 2% and it seems plausible
 30 that DUNE could do at least as well[78]. The 3DST can also do this measurement with significant
 31 statistics and with detector and reconstruction systematics largely uncorrelated with ArgonCube.
 32 The signal is independent of A and the background is small; so, it seems plausible the samples can

1 be combined to good effect.

2 **1.8.0.2 The Low- ν Method**

3 The inclusive cross section for CC scattering ($\nu_l + N \rightarrow l^- + X$) does not depend on the neutrino
4 energy in the limit where the energy transferred to the nucleus $\nu = E_\nu - E_l$ is zero [79]. In that
5 limit, the event rate is proportional to the flux, and by measuring the rate as a function of energy,
6 one can get the flux “shape.” This measurement has been used in previous experiments and has
7 the potential to provide a constraint in DUNE with a statistical uncertainty $< 1\%$.

8 In practice, one cannot measure the rate at $\nu = 0$. Instead it is necessary to restrict ν to be less
9 than a few 100 MeV. This introduces a relatively small E_ν dependence into the cross section that
10 must be accounted for to obtain the flux shape. Thus the measurement technique depends on the
11 cross section model but the uncertainty is manageable [80]. This is particularly true if low-energy
12 protons and neutrons produced in the neutrino interaction can be detected.

13 **1.8.0.3 Coherent Neutrino-Nucleus Scattering**

14 The interactions $\nu_l + A \rightarrow l^- + \pi^+ + A$ and occur with very low three momentum transfer to the
15 target nucleus (A). As such, the interactions proceed coherently with the entire nucleus, and do not
16 suffer from nuclear effects (though background channels certainly do). These coherent interactions
17 are most useful as a constraint on the $\bar{\nu}_\mu/\nu_\mu$ flux ratio. Identifying with high efficiency and purity
18 requires a detector with excellent momentum and angular resolution.

19 **1.8.0.4 Beam ν_e Content**

20 Electron neutrinos in a wideband beam come from two primary sources: kaon decays and muon
21 decays. These “beam” ν_e are an irreducible background in $\nu_\mu \rightarrow \nu_e$ oscillation searches. As such,
22 the LBNF beam was optimized to make the ν_e flux as small as possible while maximizing the ν_μ
23 flux. In the energy range relevant for oscillations (0.5 GeV - 4.0 GeV) the predicted ν_e/ν_μ ratio
24 varies between 0.5% and 1.2% as a function of energy. The beam ν_e flux in the same energy range
25 is strongly correlated with the ν_μ flux due to the decay chain $\pi^+ \rightarrow \mu^+ \nu_\mu$ followed by $\mu^+ \rightarrow \bar{\nu}_\mu e^+ \nu_e$
26 (and likewise for $\bar{\nu}_e$). As a result, the LBNF beam simulation predicts that the uncertainty on the
27 ν_e/ν_μ ratio varies from 2.0% to 4.5%. At the FD, in a 3.5 year run, the statistical uncertainty on
28 the beam ν_e component is expected to be 7% for the ν mode beam and 10% for the $\bar{\nu}$ mode beam.
29 The systematic uncertainty on the beam ν_e flux is therefore subdominant, but not negligible.

1 Glossary

- 2 **2p2h** two particle, two hole. 9, 12
- 3 **3D scintillator tracker spectrometer (3DST-S)** The 3D projection scintillator tracker spectrom-
4 eter. iv, 4, 6, 7, 14, 54, 56, 59, 65, 68
- 5 **3D scintillator tracker (3DST)** The core part of the 3D projection scintillator tracker spectrom-
6 eter. iv, 6, 7, 14, 15, 54–56, 58–64, 67, 68, 71, 76
- 7 **ALICE** A Large Ion Collider Experiment, at CERN. 41
- 8 **ArCLight** a light detector ArgonCube effort. 16, 20–22, 70
- 9 **ArgonCube** The name of the core part of the DUNE near detector (ND), a liquid argon time-
10 projection chamber (LArTPC). iii, 5, 11, 16–23, 25, 28, 54, 67, 68, 70, 76, 78
- 11 **ASIC** application-specific integrated circuit. 20, 79
- 12 **CCQE** charged current quasielastic interaction. 10, 12–14, 59, 62
- 13 **charged current (CC)** Refers to an interaction between elementary particles where a charged
14 weak force carrier (W^+ or W^-) is exchanged. 7, 8, 13, 14, 23, 58, 61, 62, 66–68, 76, 77
- 15 **charge-parity symmetry violation (CPV)** Lack of symmetry in a system before and after charge
16 and parity transformations are applied. 14, 54, 74
- 17 **charge parity (CP)** Product of charge and parity transformations. 2, 66, 67
- 18 **detector module** The entire DUNE far detector is segmented into four modules, each with a
19 nominal 10 kt fiducial mass. 74, 81
- 20 **DUNE Precision Reaction-Independent Spectrum Measurement (DUNE-PRISM)** a mobile near
21 detector that can perform measurements over a range of angles off-axis from the neutrino
22 beam direction in order to sample many different neutrino energy distributions. iv, 6, 7, 11,
23 14–16, 50, 52, 54, 63, 65, 72

- 1 **DUNE** Deep Underground Neutrino Experiment. iii, iv, 2, 3, 5–7, 9–12, 14–17, 20, 22, 23, 26, 27,
2 29, 55, 66, 68, 74–76, 78, 80
- 3 **electromagnetic calorimeter (ECAL)** A detector component that measures energy deposition of
4 traversing particles. 5–7, 29, 31–33, 36–38, 40, 41, 47, 49, 50, 54–56, 59, 65, 70
- 5 **field cage (FC)** The component of a LArTPC that contains and shapes the applied E field. 16
- 6 **far detector (FD)** Refers to the 40 kt fiducial mass DUNE detector to be installed at the far site
7 at SURF in Lead, SD, to be composed of four 10 kt modules. 2, 3, 5, 6, 8, 10–12, 14–16, 31,
8 51–53, 66, 67, 74, 75, 77
- 9 **FHC** forward horn current (ν_μ mode). iii, 23, 24, 28
- 10 **final-state interactions (FSI)** Refers to interactions between elementary or composite particles
11 subsequent to the initial, fundamental particle interaction, such as may occur as the products
12 exit a nucleus. 8, 9
- 13 **HPgTPC** high-pressure gaseous argon time projection chamber (TPC). 5–7, 14, 15, 21, 29–31,
14 33, 36, 41, 44, 65, 67, 68, 70
- 15 **high voltage (HV)** Generally describes a voltage applied to drive the motion of free electrons
16 through some media. 16
- 17 **ICARUS** add def. 25
- 18 **inner readout chamber (IROC)** inner (radial) readout chamber for gaseous argon TPC. 35
- 19 **L/E** length-to-energy ratio. 8
- 20 **LArPix** ASIC pixelated charge readout for a TPC . 20
- 21 **liquid argon time-projection chamber (LArTPC)** A class of detector technology that forms the
22 basis for the DUNE far detector modules. It typically entails observation of ionization activity
23 by electrical signals and of scintillation by optical signals. iii, 5–7, 14–18, 21, 22, 29–34, 37,
24 40, 64, 71, 76, 78
- 25 **liquid argon (LAr)** The liquid phase of argon. iii, 5, 6, 15–20, 25–29, 54, 63, 65, 66
- 26 **Long-Baseline Neutrino Facility (LBNF)** The organizational entity responsible for developing
27 the neutrino beam, the cryostats and cryogenics systems, and the conventional facilities for
28 DUNE. 7, 20, 73, 77
- 29 **MicroBooNE** The LArTPC-based MicroBooNE neutrino oscillation experiment at Fermilab. 25
- 30 **MINER ν A** The MINER ν A neutrino cross sections experiment at Fermilab. 3, 13, 27, 58, 68, 76

- 1 **MINOS** add def. 3, 12, 17
- 2 **minimum ionizing particle (MIP)** Refers to a momentum traversing some medium such that the
3 particle is losing near the minimum amount of energy per distance traversed. 25
- 4 **MPD** multi-purpose detector. i, iii, iv, 5–7, 16, 25, 29–33, 36, 39–41, 44, 54, 64, 68–71
- 5 **MWPC** multi-wire proportional chamber. 35
- 6 **NA61** CERN hadron production experiment. 27
- 7 **neutral current (NC)** Refers to an interaction between elementary particles where a neutrally
8 charged weak force carrier (Z^0) is exchanged. 8, 11
- 9 **near detector (ND)** Refers to the detector(s) installed close to the neutrino source at Fermilab.
10 iii, 2, 3, 5–8, 10–12, 14–17, 20–23, 26–31, 33, 37, 50–52, 54, 55, 58, 63–68, 70–75, 78, 80
- 11 **NO ν A** The NO ν A off-axis neutrino oscillation experiment at Fermilab. 3, 11, 12
- 12 **nonstandard interactions (NSI)** A general class of theory of elementary particles other than the
13 Standard Model. 3, 74
- 14 **NuMI** a set of facilities, collectively called “Neutrinos at the Main Injector.” The NuMI neutrino
15 beamline target system converts an intense proton beam into a focused neutrino beam. iii,
16 12, 14, 20, 27, 73
- 17 **outer readout chamber (OROC)** outer (radial) readout chamber for gaseous argon TPC. 35
- 18 **ProtoDUNE-ND** a prototype DUNE ND. 17
- 19 **ProtoDUNE** Either of the two DUNE prototype detectors constructed at CERN. One prototype
20 implements SP and the other DP technology. . 11, 23
- 21 **quasi-elastic (QE)** Refers to interaction between elementary particles and a nucleus in an energy
22 range where the interaction can be modeled as occurring between constituent quarks of one
23 nucleon and resulting in no bulk recoil of the resulting nucleus. 8–10
- 24 **RHC** reverse horn current ($\bar{\nu}_\mu$ mode). 28
- 25 **readout chamber (ROC)** readout chamber for gaseous argon TPC. 35
- 26 **SBN** Short-Baseline Neutrino program (at Fermilab). 3
- 27 **Standard Model (SM)** Refers to a theory describing the interaction of elementary particles. 73,
28 74
- 29 **SRC** short-range correlated nucleon-nucleon interactions. 9

- 1 **T2K** T2K (Tokai to Kamioka) is a long-baseline neutrino experiment in Japan studying neutrino
2 oscillations . 3, 10, 11, 14, 54, 75
- 3 **technical design report (TDR)** A formal project document that describes the experiment at a
4 technical level. 72
- 5 **time projection chamber (TPC)** The portion of each DUNE detector module that records ion-
6 ization electrons after they drift away from a cathode through the LAr, and also through
7 gaseous argon in a DP module. The activity is recorded by digitizing the waveforms of cur-
8 rent induced on the anode as the distribution of ionization charge passes by or is collected
9 on the electrode. 5, 7, 10, 16–18, 20, 22, 54–56, 79, 80

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