

Supernova T0 Updates

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August 28, 2019

SNB/LE Working Group Meeting

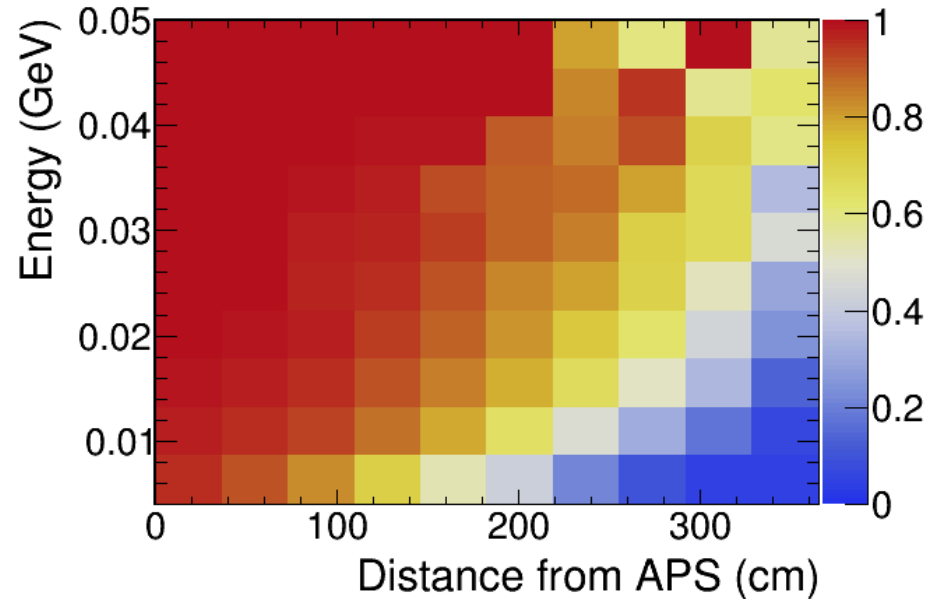
Outline

- Updates from the [last talk](#):
 - Updated efficiency matrix
 - Studied alternate drift correction method
 - Studied different figures of merit from fractional difference from neutrino energy:
 - Fraction of events contained in largest peak
 - "Intrinsic" resolution of largest peak
- Figures of merit vs energy
- Takeaways

Studying PD Effects: Reminder

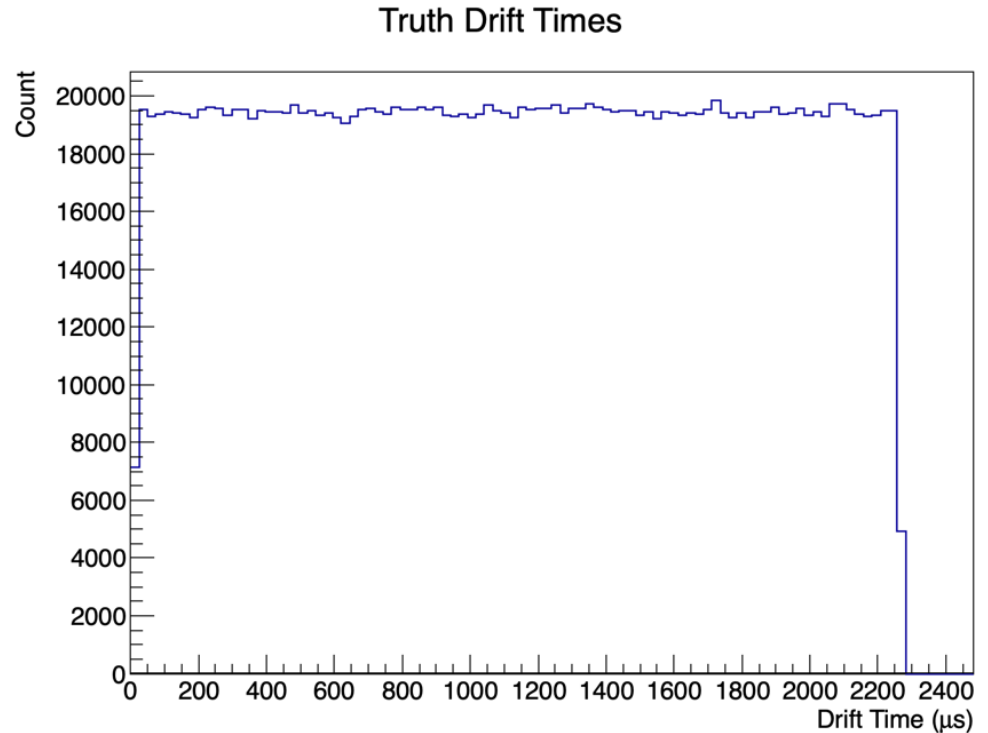
- In 2018, developed toy drift correction scheme to understand effects of PD system on supernova events
 - Use efficiency matrices corresponding to different PD performances
- Using MCC11 MARLEY events and updated PD performance types, study effects of PD systems for supernova event energy reconstruction

2018 efficiency matrix for ARAPUCA design



Reminder: “Random T0” Method

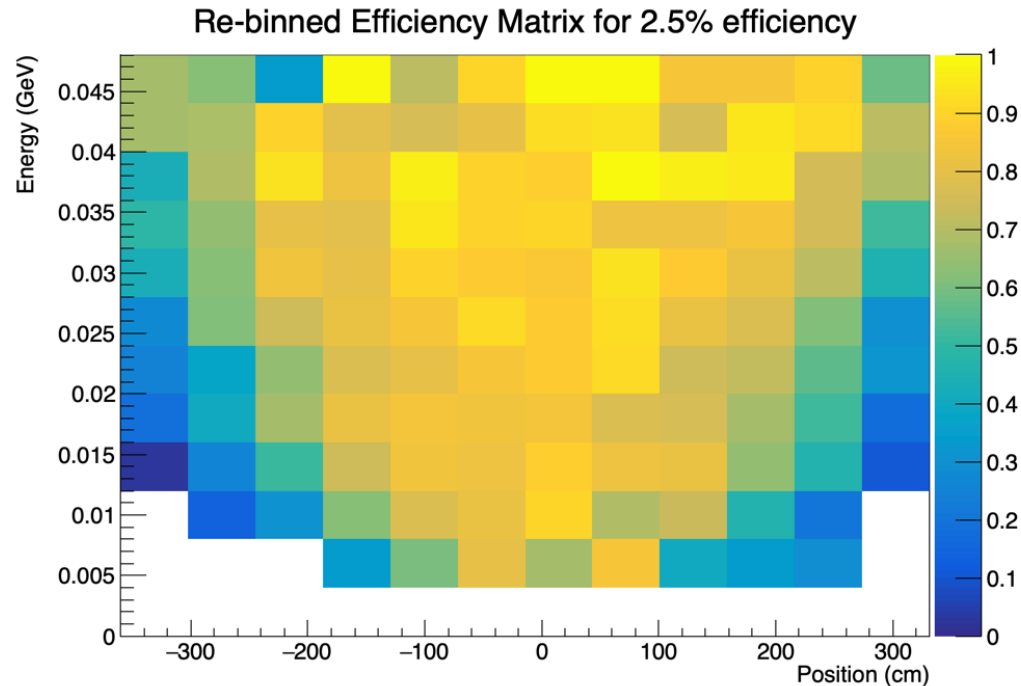
- Given MARLEY neutrino energy and distance from APA, find probability in efficiency matrix (different PD performances)
- Throw a random number $[0.0, 1.0]$ to determine what correction will take place:
 - If less than efficiency, drift correct with MC truth T0
 - If greater than efficiency, correct with a random T0



Distribution used for random correction

Updated Efficiency Matrices

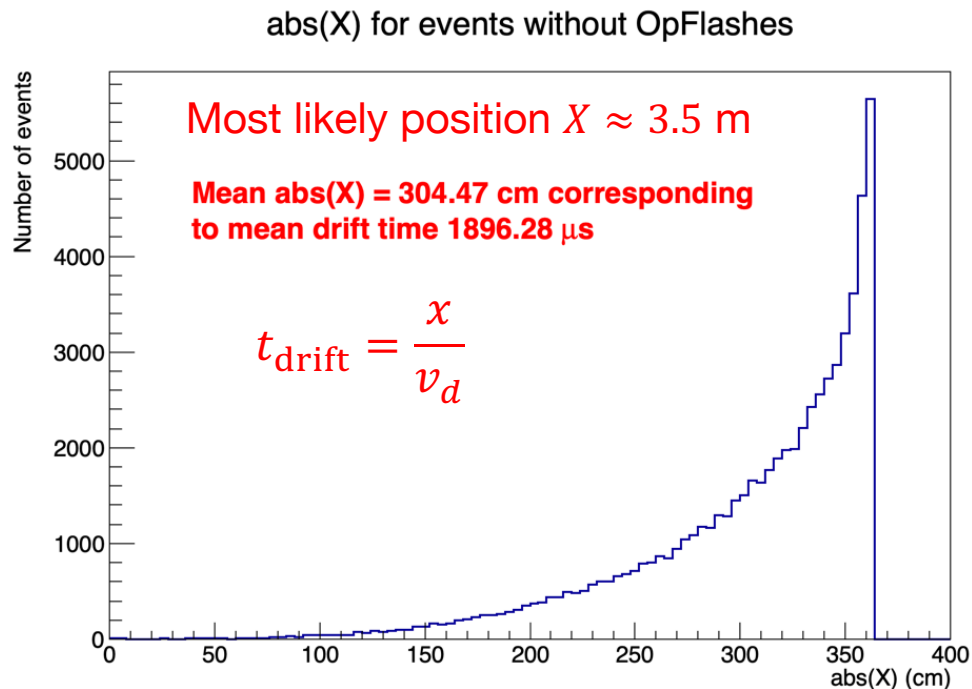
- Probability of successful flash matching as a function of energy and distance from APS
 - Re-binned; see [backup](#)
- Stringent efficiency definition (finding largest flash with distance cut associated with event)
 - Example matrix shown here; events farther from APA less likely to find photon flash



Example efficiency matrix

Other Toy Correction Schemes

- For events that don't find flash in toy method, drift correct with specific MC truth T0
 - Use mean, most likely position of events that don't have OpFlash's
 - Essentially making the assumption that we can identify bad flash matches



Fractional Energy from Truth

- See effects of toy drift correction by looking at fractional energy from truth: $(E_{\text{reco}} - E_{\text{true}})/E_{\text{true}}$
- Right: fractional energy distributions for MARLEY MCC11 clean events (weighted by GVKM supernova energy spectrum)

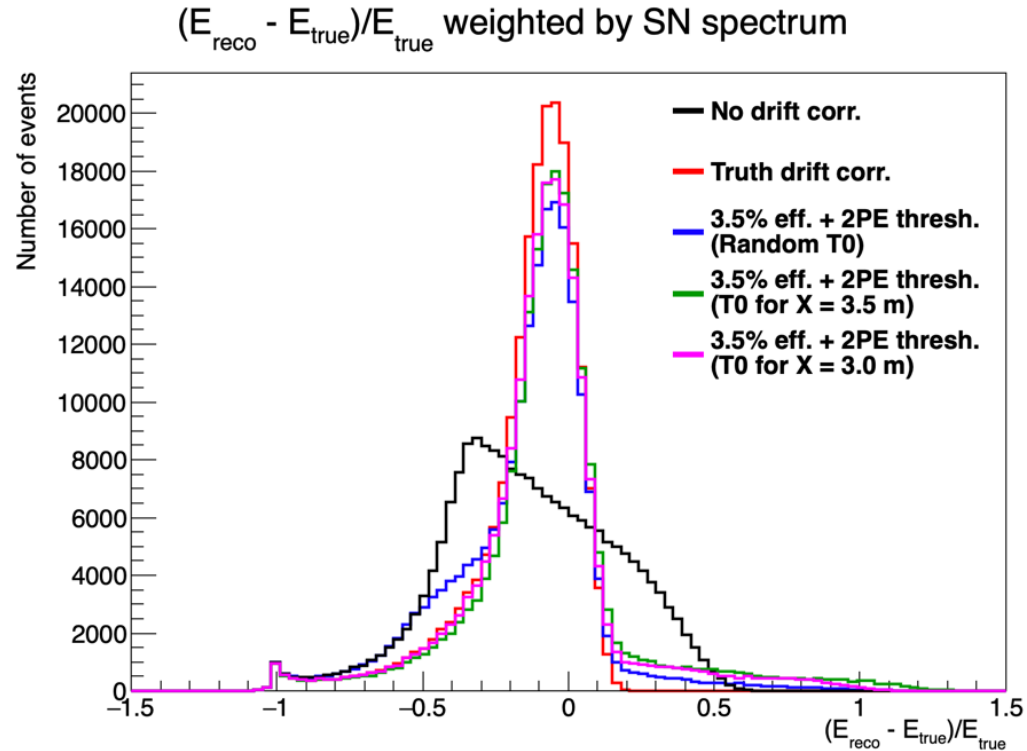
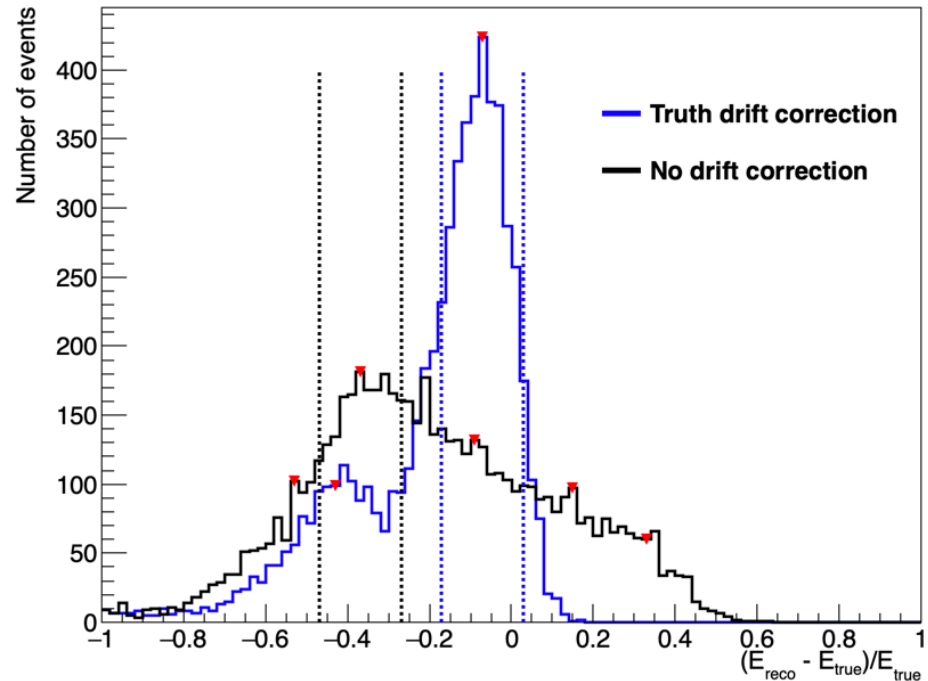


Figure of Merit for MARLEY Events

- Previous metric (σ of fractional differences) not suitable for this study; see [backup](#)
- New metric considered events close to peak
 - Found largest peak using ROOT peak-finding tools
 - Found number of events in region (peak - 0.1, peak + 0.1)
 - Metric:

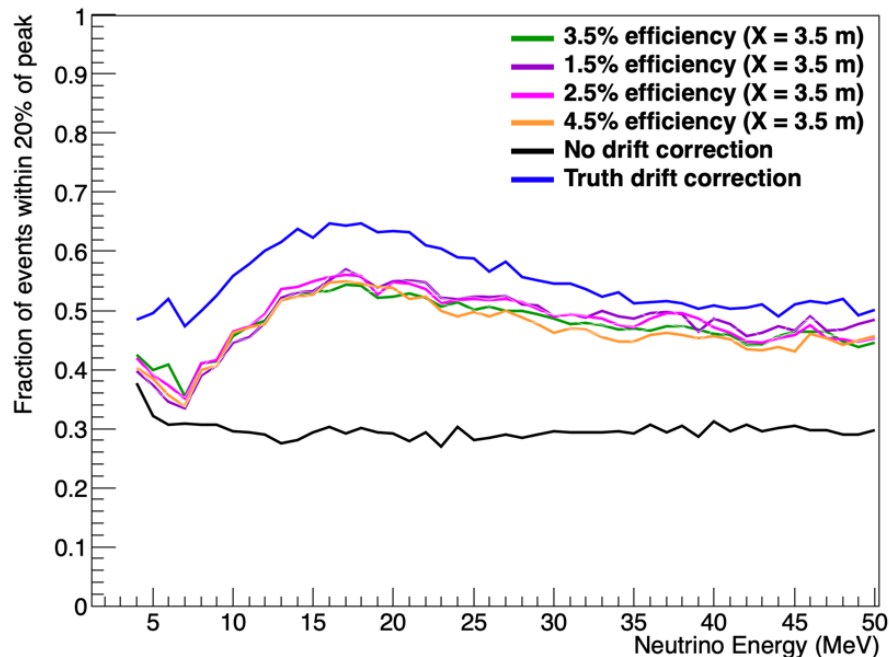
$$\frac{\# \text{ of events in region}}{\text{Total \# events}}$$

$(E_{\text{reco}} - E_{\text{true}})/E_{\text{true}}$ for 25.25 MeV MCC11 events

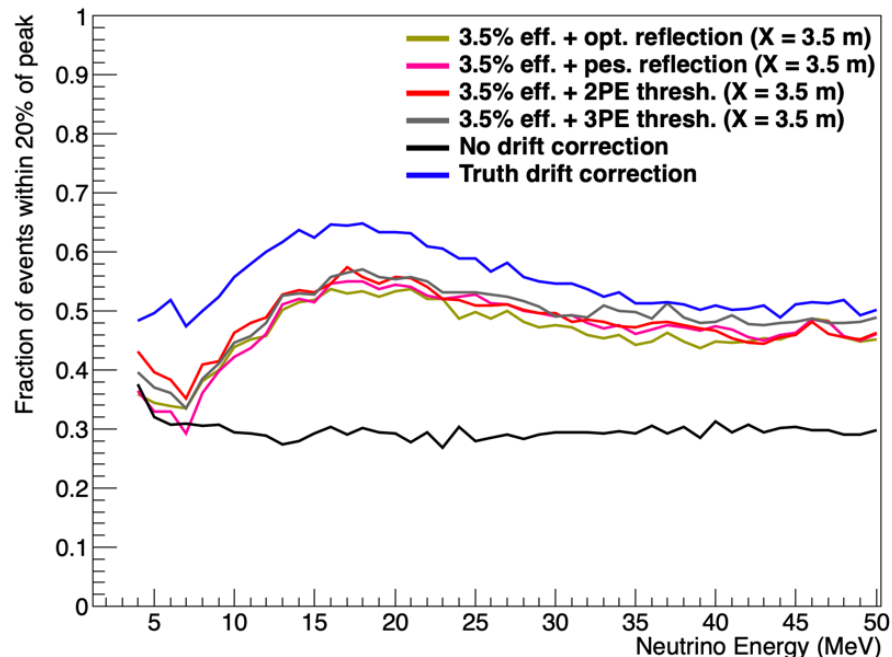


Metric vs Energy: $X = 3.5$ m correction

Fraction of events within 20% of peak vs Neutrino Energy



Fraction of events within 20% of peak vs Neutrino Energy

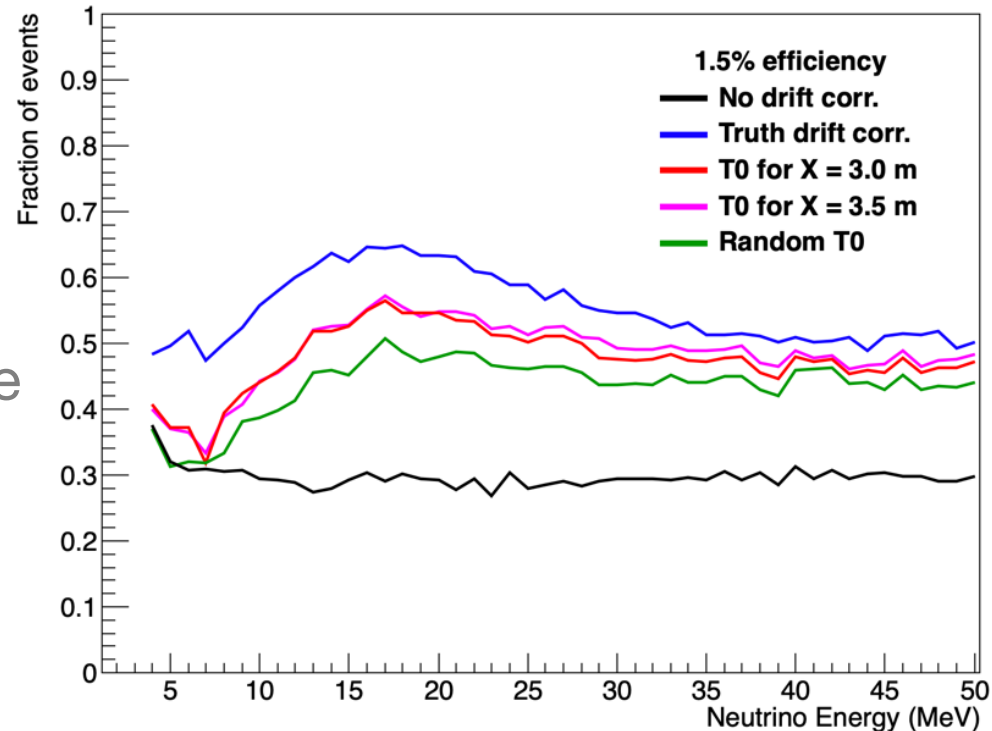


Re-binned 2 MCC11 energy levels \rightarrow 1 bin to reduce the noisy behavior

Comparing Drift Correction Methods

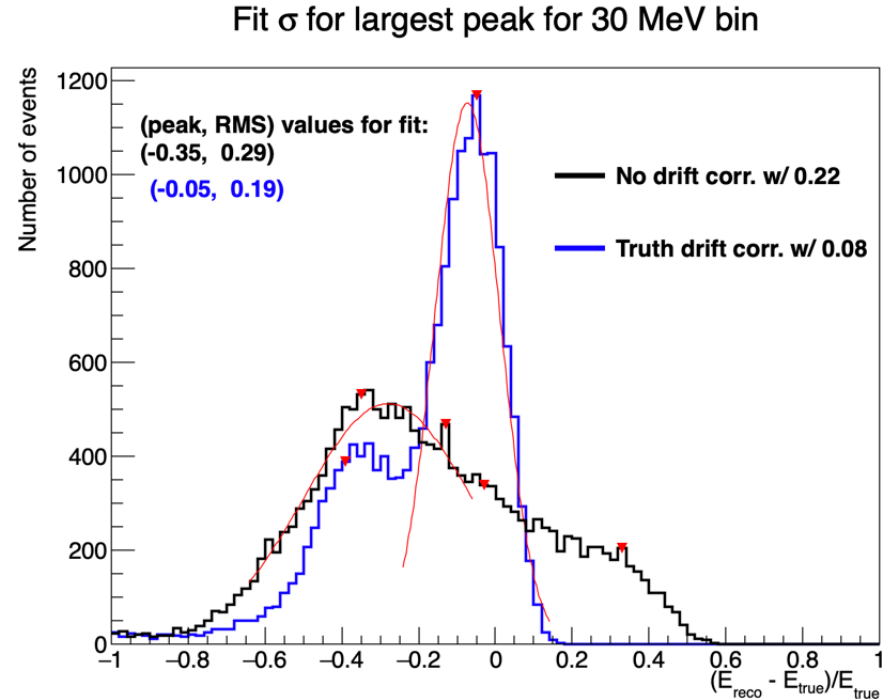
- Toy method performances (different ways of correcting events w/o flash matches) between no drift correction and truth correction
- Random T0 method performs the worst – makes sense since the correction was random
 - Correction using $X = 3.5$ m tends to perform the best – makes sense since it's the largest correction

Fraction of events close to largest peak: 1.5% efficiency



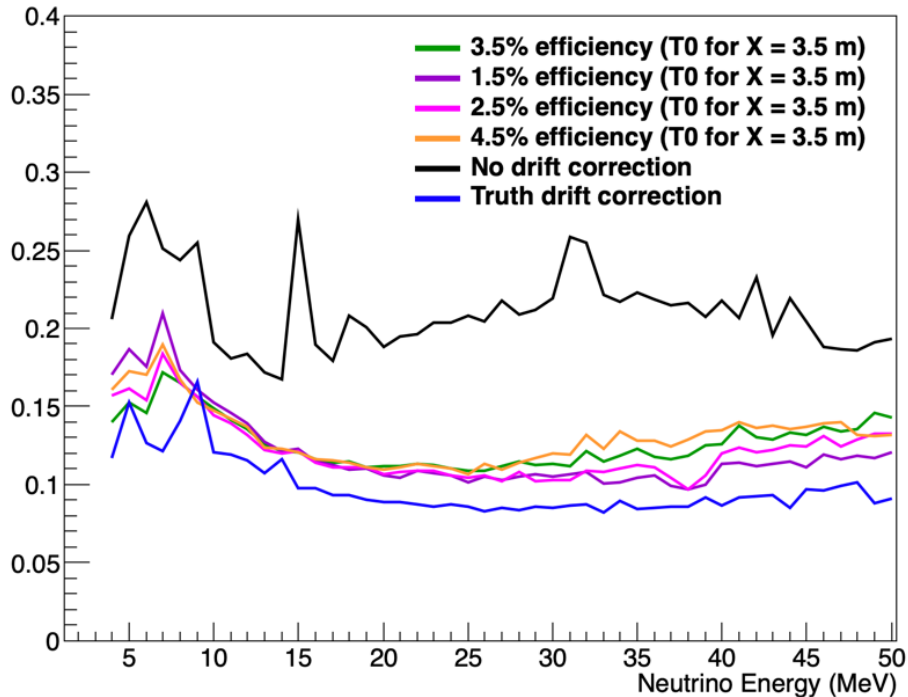
Alternate Figure of Merit: “Intrinsic” Resolution

- Try to find metric that captures “true” resolution (resolution of events without nucleon emission)
- From histogram of fractional differences, find largest peak
 - Then fit the distribution from (largestPeak – RMS, largestPeak + RMS)
 - The σ from that Gaussian fit is the metric

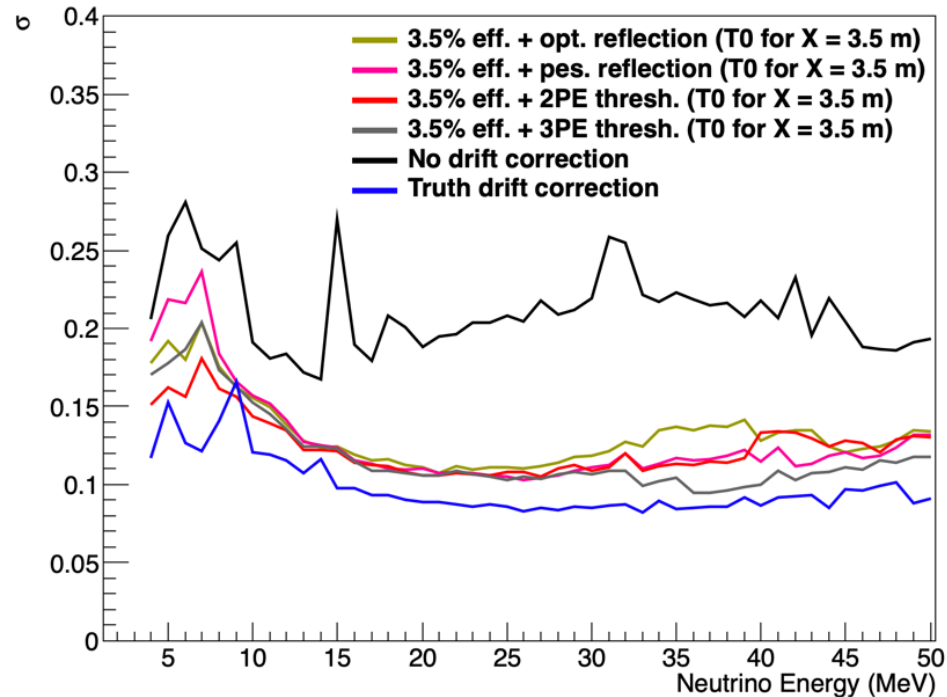


Metric vs Energy: $X = 3.5$ m correction

σ of closest peak vs Neutrino Energy



σ of closest peak vs Neutrino Energy



Takeaways

- PDs improve energy reconstruction, but resolution not strongly tied to detector performance
 - Both metrics capture quality of energy reconstruction, but do not distinguish different PD performance types
- Note: these results might only be true for MCC11 MARLEY sample

Backup Slides

Drift Correction Reminder

True drift correction

- $Q = Q_0 \exp\left(\frac{x}{v_d \tau_e}\right)$
 - Q : Truth charge
 - Q_0 : Observed charge
 - x : Distance from electron vertex to APA (MC Truth)
 - v_d : Electron drift velocity
 - τ_e : Electron lifetime

Reco drift correction

- $Q = Q_0 \exp\left(\frac{t_0}{\tau_e}\right)$
 - Q : Truth charge
 - Q_0 : Observed charge
 - τ_e : Electron lifetime
 - t_0 : Reco interaction start time
- Find t_0 using photon flash, reco hit information (used longest track as reco electron track)

PD Performance Types: Reminder

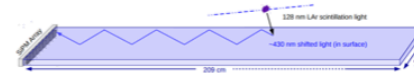
- Motivation: evaluate photon detector requirements for SN physics; coupling physics to PD performance
- Distinguish photon detector performance variations based on “effective area”
 - Right: slide from a [talk](#) by Logan Rice

Effective Areas Values for Possible Designs

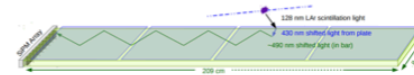


- Effective Area = (Ave. prob. of a photon reaching the detector surface to be recorded) × (Total area)

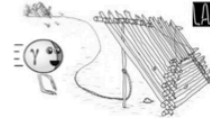
- ▶ Dip-Coated Designs in protoDUNE: 3.84 cm^2



- ▶ Double-Shifted Designs in protoDUNE 4.1 cm^2

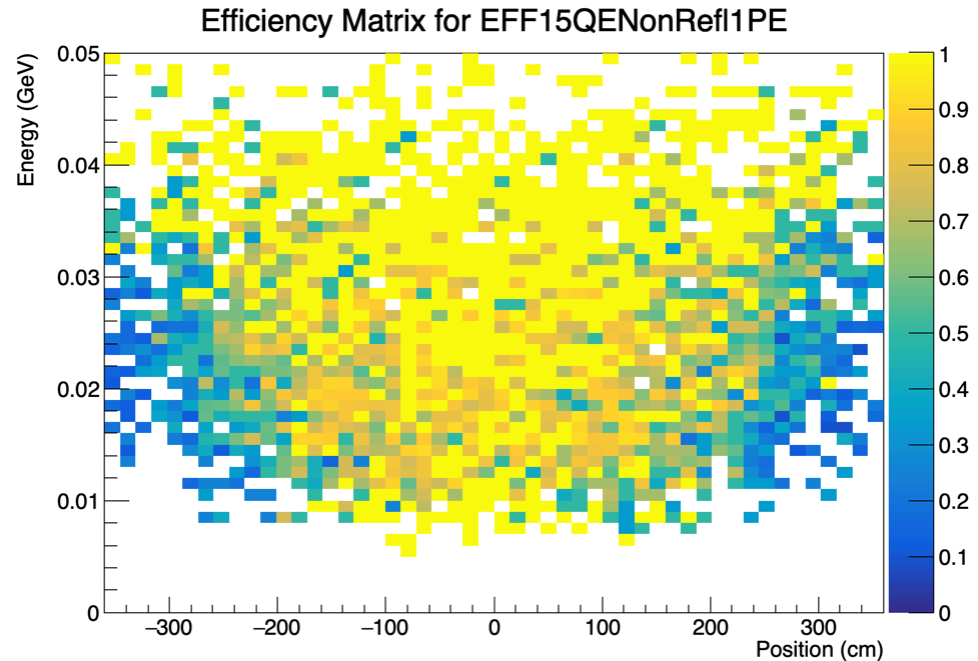


- ▶ Various Arapuca Designs: 5.12 cm^2 , 12.80 cm^2 , 23 cm^2



Un-binned Efficiency Matrices

- Efficiency matrix:
Probability of successful flash matching given true neutrino energy, distance from APA
- Less statistics compared to previous efficiency matrices; re-binned to reduce number of “holes”
 - Merged 4 bins into 1 for both axes



1.5% QE (before re-binning)

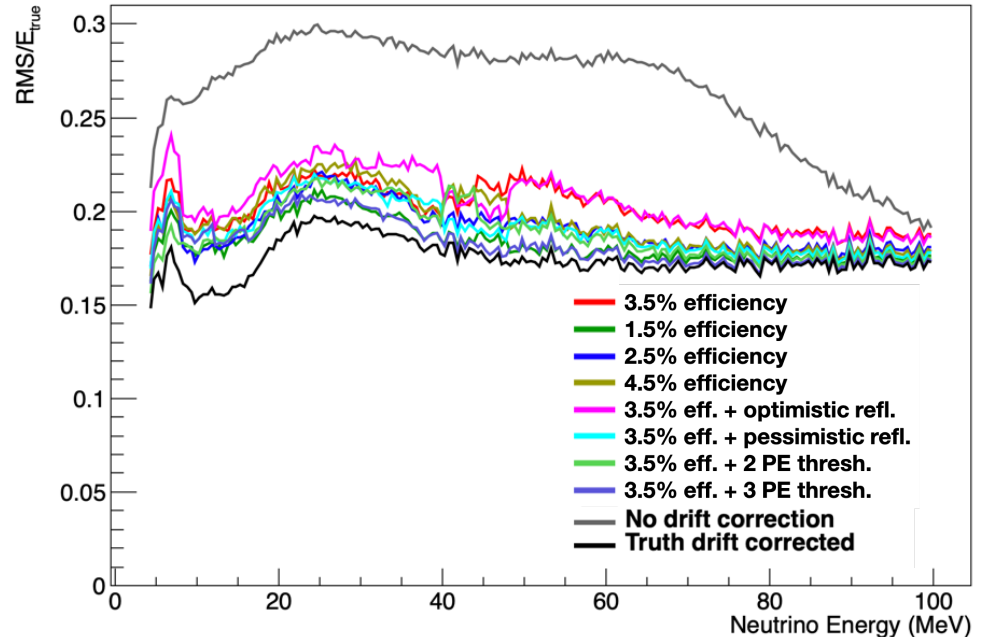
Resolution Plots: Random T0 Method

- Using fractional energy values, calculate standard deviation:

$$\sigma_{\text{frac}} = \sqrt{\frac{\sum_{i=1}^N (E_{\text{frac},i} - \bar{E}_{\text{frac}})^2}{N - 1}}$$

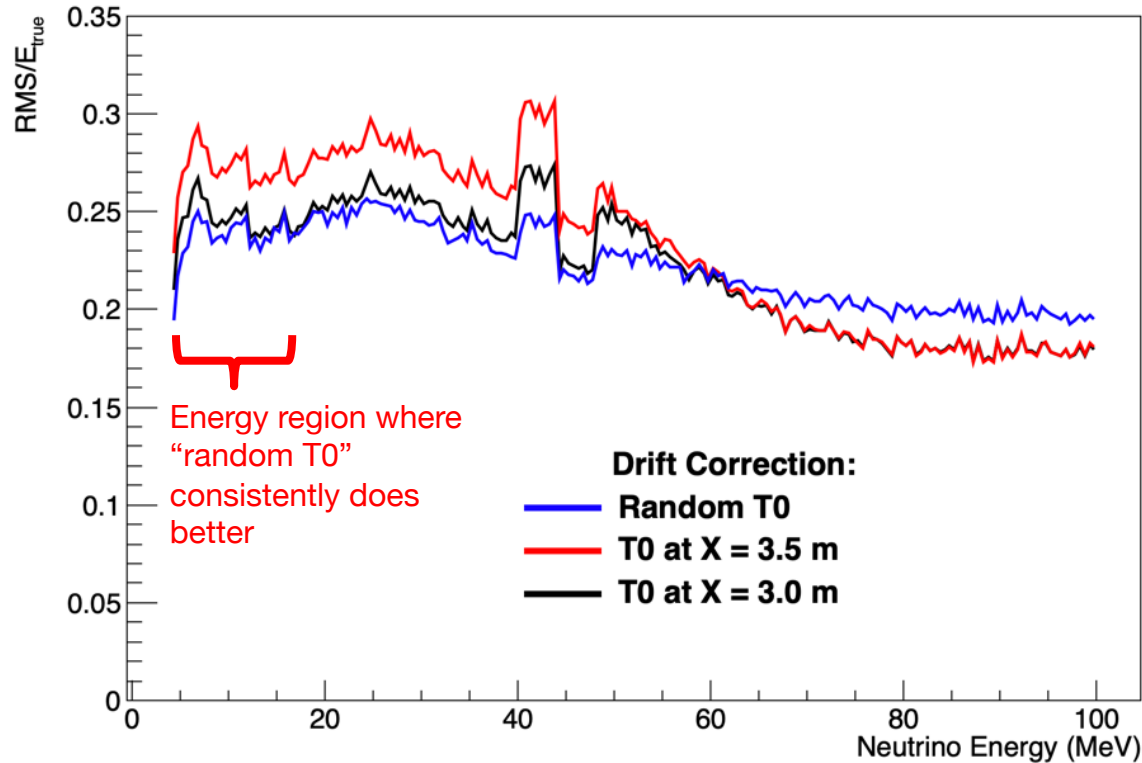
- Comparable to truth drift correction!
- This metric failed to capture differences in drift correction methods for each PD performance type

σ vs Neutrino Energy: Charge Reco Energy



Comparing Drift Correction Methods

σ vs Neutrino Energy: 3.5% eff. + 2PE thresh.



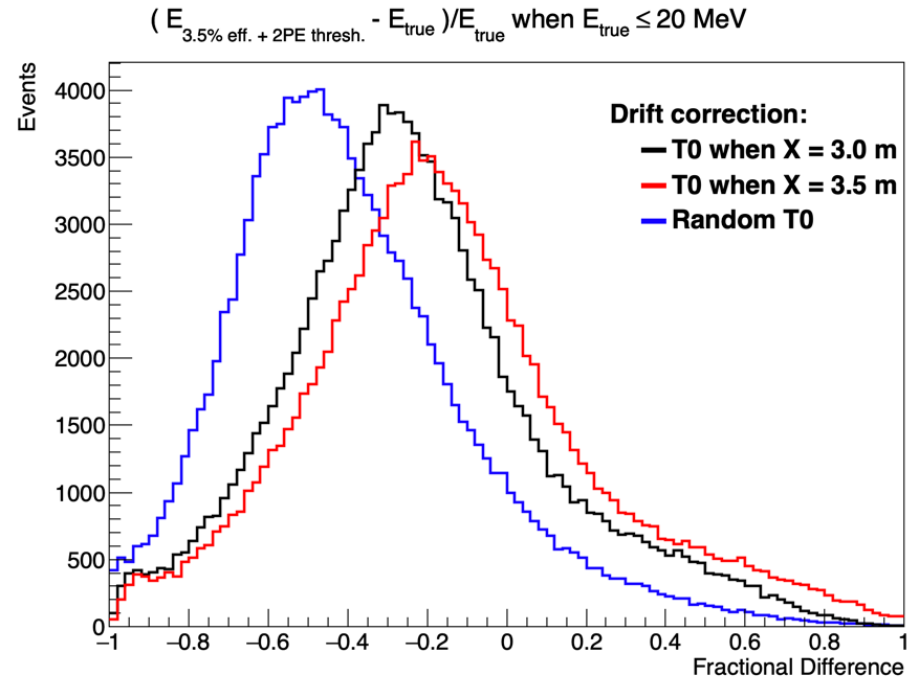
Using fractional energy values, calculate standard deviation:

$$\sigma_{\text{frac}} = \sqrt{\frac{\sum_{i=1}^N (E_{\text{frac},i} - \bar{E}_{\text{frac}})^2}{N - 1}}$$

However, this figure of merit implies that the "random T0" method is better for events under 50 MeV, which doesn't make sense...

Comparing Drift Correction Methods

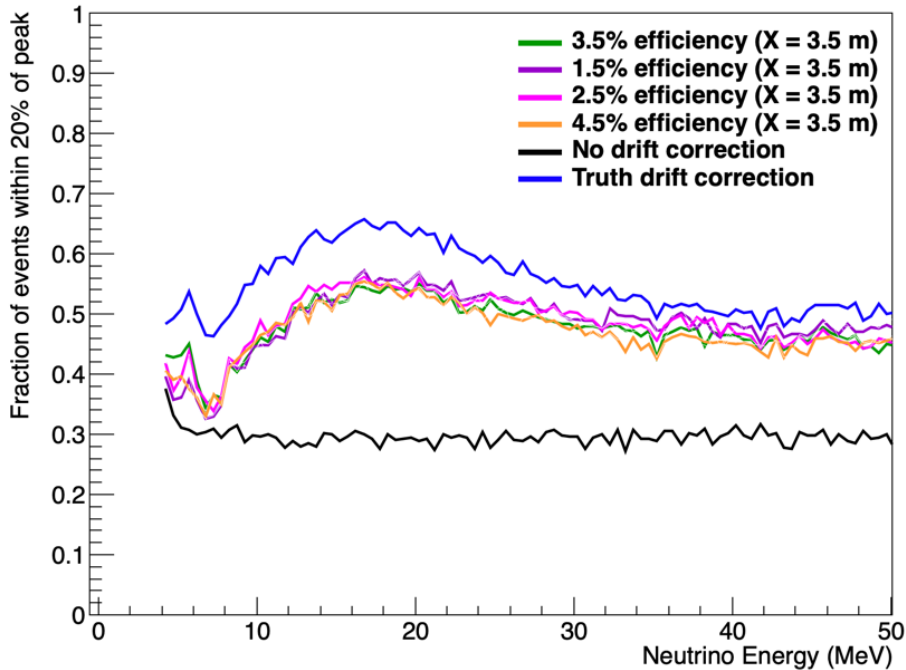
- Consider histograms of fractional difference, $(E_{\text{reco}} - E_{\text{true}})/E_{\text{true}}$
- Random T0 method reconstructs less energy on average, but σ /width looks smaller by eye
 - Motivated search for new figure of merit



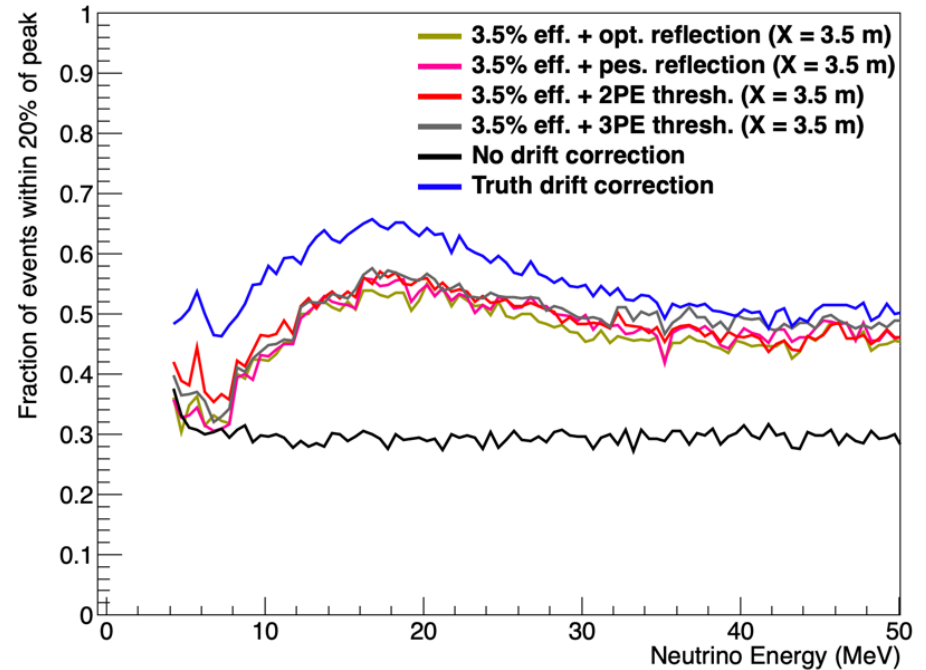
Fractional difference histograms for events without flash and energy cut

Updated Metric vs Un-binned Energy

Fraction of events within 20% of peak vs Neutrino Energy

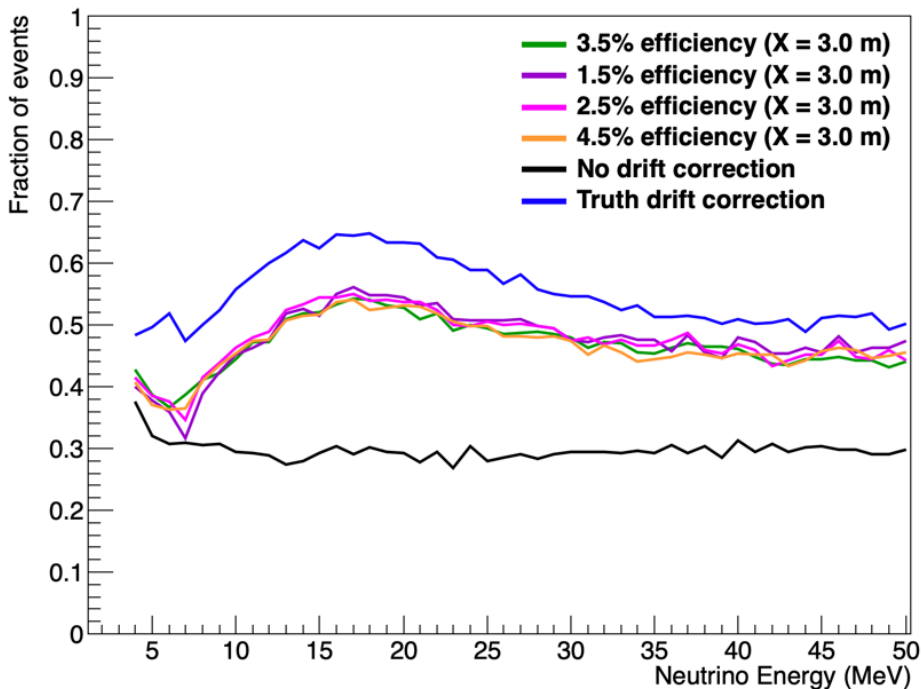


Fraction of events within 20% of peak vs Neutrino Energy



Metric vs Energy: $X = 3.0$ m correction

Fraction of events close to largest peak



Fraction of events close to largest peak

